# Atom interferometry and gravitational wave detection

**Quantum Sensors for Fundamental Physics** 

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Jason Hogan Stanford University September 2021

# Outline

#### Lecture 1

- General introduction/motivation
- Non-relativistic atom interferometer phase theory
- Example applications
- Tutorial: Accelerometer phase response

Lecture 2

- General relativistic phase shift theory
- GR effects and clock interferometry
- Gravitational wave detection and MAGIS
- Aharonov-Bohm phase shifts
- Tutorial: Gravitational wave phase response

Lecture 3

- Advanced atom optics (large momentum transfer techniques)
- Systematic errors, backgrounds, and mitigations
- Supporting tools: matter wave lensing, optical lattices, phase shear readout

# Lecture 2

# Interference at long interrogation time





Dickerson, et al., PRL **111**, 083001 (2013).

# GR Back of the Envelope

Phase resolution target of some proposed experiments:

$$\frac{\delta\phi}{\phi}\approx 10^{-15}$$



At this level of sensitivity, does relativity start to affect results?

• Gravitational red shift of light:

$$\frac{\delta\nu}{\nu} = \frac{gL}{c^2} \approx 10^{-15}$$

• Special relativistic corrections:

$$\frac{v^2}{c^2} \approx 10^{-15}$$

### Summary: Non-relativistic phase shift calculation

The atom interferometer phase shift can be decomposed as

$$\Delta\phi_{\rm tot} = \Delta\phi_{\rm propagation} + \Delta\phi_{\rm separation} + \Delta\phi_{\rm laser}$$

$$\Delta \phi_{\text{propagation}} = \sum_{\text{upper}} \left( \int_{t_{I}}^{t_{F}} (L_{c} - E_{i}) dt \right) - \sum_{\text{lower}} \left( \int_{t_{I}}^{t_{F}} (L_{c} - E_{i}) dt \right)$$
$$\Delta \phi_{\text{laser}} = \left( \sum_{j} \pm \phi_{L}(t_{j}, \mathbf{x}_{u}(t_{j})) \right)_{\text{upper}} - \left( \sum_{j} \pm \phi_{L}(t_{j}, \mathbf{x}_{l}(t_{j})) \right)_{\text{lower}}$$
$$\Delta \phi_{\text{separation}} = \bar{\mathbf{p}} \cdot \Delta \mathbf{x}$$

# General relativistic phase shifts

Light-pulse interferometer phase shifts in GR:

- Geodesic propagation for atoms and light.
- Path integral formulation to obtain quantum phases.
- Atom-field interaction at intersection of laser and atom geodesics.



Prior work, de Broglie interferometry: Post-Newtonian effects of gravity on quantum interferometry, Shigeru Wajima, Masumi Kasai, Toshifumi Futamase, Phys. Rev. D, 55, 1997; Bordé, et al.

# Propagation Phase: The Atom's Clock

The atom's propagation phase measures proper time.

- Feynman path integral propagator:
- The action is modified in GR:

$$S_{\rm NR} = \int Ldt \qquad \Longrightarrow \qquad S_{\rm GR} = \int mc^2 \, d\tau$$

• Propagation phase for atom:

$$\Delta \phi_{\text{propagation}} = \frac{1}{\hbar} \int L \, dt = \int \omega_{\text{C}} \, d\tau$$
Compton frequency (Cs):  

$$\omega_{\text{C}} = mc^2/\hbar$$

$$= 3.2 \times 10^{25} \text{ Hz}$$

 $\Sigma e^{iS/\hbar}$ 

Proper time

paths

The action is proportional to the length of the word line (proper time).

$$\phi_{\text{propagation}} = \int Ldt = \int md\tau = \int p_{\mu}dx^{\mu} \qquad p^{\mu} \equiv m\frac{dx^{\mu}}{d\tau} \qquad d\tau^{2} = g_{\mu\nu}dx^{\mu}dx^{\nu}$$

$$(c=1) \qquad p_{\mu}dx^{\mu} = mg_{\mu\nu}\frac{dx^{\mu}}{d\tau}dx^{\nu} = m\frac{d\tau}{d\tau}d\tau$$

$$(39)$$

# Proper time, NR limit

Consider the proper time of a clock in the Schwarzschild spacetime:

$$ds^{2} = (1 + 2\phi/c^{2})c^{2}dt^{2} - \frac{1}{1 + 2\phi/c^{2}}dr^{2} - r^{2}d\Omega^{2} \qquad \phi(r) = -\frac{GM}{r} \approx -\frac{GM}{R_{e}} + gz$$
  
Earth radius

Proper time:  $c^2 d\tau^2 = g_{\mu\nu} dx^{\mu} dx^{\nu}$ 

$$c \, d\tau = \sqrt{g_{tt} + g_{rr} \frac{dr}{dt} \frac{dr}{dt}} \, dt$$

To first order in  $\phi(r)$ :

$$d\tau \approx \sqrt{1 + 2\phi(r)/c^2 - (v/c)^2} \, dt$$
  
$$\tau \approx \int \left(1 + \frac{\phi(r)}{c^2} - \frac{v^2}{2c^2}\right) \, dt$$
  
Gravitational redshift Time dilation

# Propagation Phase: Two Limits

$$S_{\rm GR} = \int mc^2 d\tau \qquad \qquad \Delta \phi_{\rm propagation} = \frac{1}{\hbar} \int L \, dt = \int \omega_{\rm C} \, d\tau$$

1. Non-relativistic limit:

(NR Lagrangian)  $S_{\rm GR} = mc^2 \int \frac{d\tau}{dt} dt \approx mc^2 \int \left(1 + \frac{\phi(r)}{c^2} - \frac{v^2}{2c^2}\right) dt = \int \left(mc^2 + m\phi(r) - \frac{1}{2}mv^2\right) dt$ 

 $\rightarrow$  Proper time reduces to the non-relativistic Lagrangian.

- 2. Classical limit (Principle of Least Action):
  - $\rightarrow$  QM: particles travel along the **extremal path** when  $~S \gg \hbar$
  - $\rightarrow$  GR: particles travel along **geodesics**, paths that maximize proper time.

Atoms in an AI travel along the classical, geodesic path.

# Atom-light interaction in LLF

Geodesic equation

$$\frac{d^2x^{\mu}}{d\tau^2} + \Gamma^{\mu}_{\alpha\beta}\frac{dx^{\alpha}}{d\tau}\frac{dx^{\beta}}{d\tau} = 0$$

Local Lorentz Frame (LLF) is the freely-falling frame; Christoffels transformed to zero

$$g_{\mu'\nu'}(x') = \eta_{\mu'\nu'} + \mathcal{O}({x'}^2)$$

Locally flat

Metric *locally flat* in LLF

The atomic physics dynamics in the LLF are the *normal rules of non-relativistic quantum mechanics*. For example, the photon recoil is

$$m_{\rm in} \frac{dx'_{\rm atom}^{i}}{d\tau} \bigg|_{\rm after} = m_{\rm fi} \frac{dx'_{\rm atom}^{i}}{d\tau} \bigg|_{\rm before} + \frac{dx'_{\rm light}^{i}}{d\lambda} \quad ({\rm in \ LLF}) \qquad \qquad m_{\rm in} = m_{\rm fi} \frac{d\tau}{d\tau} \bigg|_{\rm before} + \frac{dx'_{\rm light}^{i}}{d\tau} = m_{\rm fi} \frac{d\tau}{d\tau} \bigg|_{\rm before} = m_{\rm fi} \frac{d\tau}{d\tau} \bigg|_{\rm before} + \frac{d\tau}{d\tau} \frac{d\tau}{d\tau} \bigg|_{\rm fight} = m_{\rm fi} \frac{d\tau}{d\tau} = m_{\rm fi} \frac{d\tau}{d\tau} \bigg|_{\rm before} = m_{\rm fi} \frac{d\tau}{d\tau} \bigg|_{\rm fight} = m_{\rm fight} = m$$

 $m_{\rm in} = m_{\rm fi} \pm \omega_{\rm a}$ 

atom mass change (internal energy of excited state)

There are small corrections O(R) due to Riemann curvature R, but these can be safely neglected since the atom is small.

Which LLF? The atom's recoil velocity says there is no unique frame! True, but corrections are small,  $\mathcal{O}(v_r^2)$ 

### Laser phase

$$\Delta \phi_{\text{laser}} = \left( \sum_{j} \pm \phi_L(t_j, \mathbf{x}_u(t_j)) \right)_{\text{upper}} - \left( \sum_{j} \pm \phi_L(t_j, \mathbf{x}_l(t_j)) \right)_{\text{lower}}$$

#### A null geodesic is a 'path of constant phase'



Height

- Phase is **constant** along the null geodesic
- Phase imprinted on atom is the phase of light at (spacetime) point of emission

# Single photon vs two photon atom optics



#### **Consequential differences:**

- For 2-photon, light pulses must leave at different times
- By convention, "passive" laser is on before "control" pulse arrives
- Arrival of control pulse determines passive null geodesic (and its laser phase)
- No passive laser for single photon atom optics

### Separation phase



 $\rightarrow$  Phase to move from one wavepacket to the other in spacetime.

A very useful GR argument:

- The expression is the right answer in one frame (here, the LLF)
- It's written in a covariant manner
- Therefore, it's the correct relativistic result in all frames:

$$\phi_{\text{separation}} = \int_{E}^{D} \overline{p}_{\mu} dx^{\mu}$$
 (in any frame)

### Example: Schwarzschild and PPN

Near the Earth,

$$ds^2 = (1+2\phi) dt^2 - \frac{1}{1+2\phi} dr^2 - r^2 d\Omega^2$$
 Schwarzschild

$$\phi = -\frac{GM}{r}$$
 Very small (~1e-9)

$$ds^{2} = (1 + 2\phi + 2\beta\phi^{2})dt^{2} - (1 - 2\gamma\phi)dr^{2} - r^{2}d\Omega^{2}$$

Parameterized post-Newtonian (PPN)

 $\beta$  and  $\gamma$  parameters describe deviations from GR

$$\beta = \gamma = 1$$
 in GR



# PPN geodesics are complicated

$$\begin{split} r(\tau) &= r_0 + v_{r0}\tau + \frac{\left(-v_{r0}^2(-1+\gamma) + \eta + 2\left(v_{r0}^2(\beta-2\gamma) + \beta\eta\right)\phi(r_0) - 6v_{r0}^2\beta\gamma\phi(r_0)^2\right)\partial_r\phi(r_0)}{2(-1+2\gamma\phi(r_0))\left(1+2\phi(r_0) + 2\beta\phi(r_0)^2\right)}\tau^2 + \\ &\quad \frac{1}{6}v_{r0}\left(\frac{2\gamma\left(v_{r0}^2(-1+2\gamma) - \eta - 2\left(v_{r0}^2(\beta-3\gamma) + \beta\eta\right)\phi(r_0) + 8v_{r0}^2\beta\gamma\phi(r_0)^2\right)\partial_r\phi(r_0)^2}{(1-2\gamma\phi(r_0))^2\left(1+2\phi(r_0) + 2\beta\phi(r_0)^2\right)} + \\ &\quad \frac{2\left(-v_{r0}^2 - \eta + 2v_{r0}^2\gamma\phi(r_0)\right)\left(-2 + \beta + \gamma - 6(\beta - \gamma)\phi(r_0) - 6\beta(\beta - 3\gamma)\phi(r_0)^2 + 16\beta^2\gamma\phi(r_0)^3\right)\partial_r\phi(r_0)^2}{(1-2\gamma\phi(r_0))^2\left(1+2\phi(r_0) + 2\beta\phi(r_0)\right)^2\right)}\tau^3 + \mathcal{O}(\tau^4) \\ &\quad \frac{v_{r0}^2\gamma\partial_r^2\phi(r_0)}{1-2\gamma\phi(r_0)} + \frac{(1+2\beta\phi(r_0))\left(v_{r0}^2 + \eta - 2v_{r0}^2\gamma\phi(r_0)\right)\partial_r^2\phi(r_0)}{(-1+2\gamma\phi(r_0))\left(1+2\phi(r_0) + 2\beta\phi(r_0)^2\right)}\right)\tau^3 + \mathcal{O}(\tau^4) \\ &\quad t(\tau) = t_0 + \sqrt{\frac{v_{r0}^2 + \eta - 2v_{r0}^2\gamma\phi(r_0)}{1+2\phi(r_0) + 2\beta\phi(r_0)^2}}\tau - \frac{v_{r0}(1+2\beta\phi(r_0))\sqrt{\frac{v_{r0}^2 + \eta - 2v_{r0}^2\gamma\phi(r_0)}{1+2\phi(r_0) + 2\beta\phi(r_0)^2}}}{1+2\phi(r_0) + 2\beta\phi(r_0)^2}\tau^2 - \\ &\quad \left(\left(\sqrt{\frac{v_{r0}^2 + \eta - 2v_{r0}^2\gamma\phi(r_0)}{1+2\phi(r_0) + 2\beta\phi(r_0)^2}}\left(\left(-v_{r0}^2(-5 + 2\beta + \gamma) + \eta + 2\left(v_{r0}^2(-6\gamma + \beta(8 + \gamma)) + 2\beta\eta\right)\phi(r_0) + \right)\right)\right)\tau^2 \right)\tau^2 \right)\tau^2 \right)\tau^2 \tau^2 - \\ &\quad t(\tau) = t_0 + v_0 + v$$

$$\begin{aligned} f(r) &= t_0 + \sqrt{\frac{v_{r_0}^2 + \eta - 2v_{r_0}^2\gamma\phi(r_0)}{1 + 2\phi(r_0) + 2\beta\phi(r_0)^2}} \tau - \frac{\frac{6r_0(1 + 2\beta\phi(r_0))\sqrt{1 + 2\phi(r_0) + 2\beta\phi(r_0)^2}}{1 + 2\phi(r_0) + 2\beta\phi(r_0)^2} \tau^2 - \\ & \left( \left( \sqrt{\frac{v_{r_0}^2 + \eta - 2v_{r_0}^2\gamma\phi(r_0)}{1 + 2\phi(r_0) + 2\beta\phi(r_0)^2}} \left( \left( -v_{r_0}^2 (-5 + 2\beta + \gamma) + \eta + 2\left( v_{r_0}^2 (-6\gamma + \beta(8 + \gamma)) + 2\beta\eta \right) \phi(r_0) + \right. \right. \\ & \left. 2\beta\left( v_{r_0}^2 (8\beta - 19\gamma) + 2\beta\eta \right) \phi(r_0)^2 - 36v_{r_0}^2\beta^2\gamma\phi(r_0)^3 \right) \partial_r\phi(r_0)^2 + \\ & \left. v_{r_0}^2 (-1 + 2\gamma\phi(r_0)) \left( 1 + 2(1 + \beta)\phi(r_0) + 6\beta\phi(r_0)^2 + 4\beta^2\phi(r_0)^3 \right) \partial_r^2\phi(r_0) \right) \right) \tau^3 \right) / \\ & \left( 3\left( \left( -1 + 2\gamma\phi(r_0) \right) \left( 1 + 2\phi(r_0) + 2\beta\phi(r_0)^2 \right)^2 \right) \right) + \mathcal{O}(\tau^4) \end{aligned}$$

Dimopoulos et al., PRD 78 (2008)

- This expression can be used for both the light and atom geodesics (choose  $\eta$ )  $\eta = g_{\mu\nu} \frac{dx^{\mu}}{d\tau} \frac{dx^{\nu}}{d\tau}$
- Need to find intersection points!
- Can solve perturbatively (small velocity, weak gravity)
- Consistently Taylor expand in all small variables simultaneously: "physical series"

# PPN phase shift results

		GR Phase Shift	Size (rad)	Interpretation	NR Phase Shift	
	1.	$-k_{ m eff}gT^2$	$3. \times 10^8$	Newtonian gravity	$-k_{ m eff}gT^2$	
	2.	$-k_{ m eff}(\partial_r g) v_L T^3$	$-2. \times 10^3$	1st gradient	$-k_{\rm eff}(\partial_r g)v_L T^3$	
	3.	$-rac{7}{12}k_{ ext{eff}}(\partial_r g)gT^4$	$9. \times 10^2$		$-\frac{7}{12}k_{\rm eff}(\partial_r g)gT^4$	
	4.	$-3k_{ m eff}g^2T^3$	$-4. \times 10^1$	finite speed of light and		
	5.	$-3k_{\mathrm{eff}}gv_LT^2$	$4. \times 10^1$	Doppler shift corrections	2	
	6.	$-rac{k_{ m eff}^2}{2m}(\partial_r g)T^3$	$-7.  imes 10^{-1}$	1st gradient recoil	$-rac{k_{ m eff}^2}{2m}(\partial_r g)T^3$	
Clock AI	7.	$(\omega_{ m eff} - \omega_a)  gT^2$	$-4. \times 10^{-1}$	detuning		Missing from
response	8.	$(2-2eta-\gamma)k_{ m eff}g\phi T^2$	$-2. \times 10^{-1}$	GR (non-linearity)		NR calc
response	9.	$-rac{3k_{ m eff}^2}{2m}gT^2$	$2. \times 10^{-2}$			
	10.	$-\frac{7}{12}k_{\mathrm{eff}}v_L^2(\partial_r^2 g)T^4$	$8.  imes 10^{-3}$	2nd gradient	$-\frac{7}{12}k_{\text{eff}}v_L^2(\partial_r^2 g)T^4$	
	11.	$-rac{35}{4}k_{ ext{eff}}(\partial_r g)gv_LT^4$	$6. \times 10^{-4}$		12	
	12.	$-4k_{ m eff}(\partial_r g)v_L^2T^3$	$-3. \times 10^{-4}$			
	13.	$2\omega_a g^2 T^3$	$2. \times 10^{-4}$			
	14.	$2\omega_a g v_L T^2$	$-2. \times 10^{-4}$		201	
	15.	$-rac{7k_{ m eff}^2}{12m}v_L(\partial_r^2g)T^4$	$7.  imes 10^{-6}$	2nd gradient recoil	$-\frac{7k_{\mathrm{eff}}^2}{12m}v_L(\partial_r^2 g)T^4$	
	16.	$-12k_{\mathrm{eff}}g^2v_LT^3$	$-7. \times 10^{-6}$			
	17.	$-7k_{ m eff}g^3T^4$	$4. \times 10^{-6}$			
	18.	$-5k_{ m eff}gv_L^2T^2$	$3. \times 10^{-6}$	GR (velocity-dependent force)		
	19.	$(2-2\beta-\gamma)k_{\rm eff}\partial_r(g\phi)v_LT^3$	$2. \times 10^{-6}$	GR 1st gradient	N	·
	20.	$\frac{7}{12}(4-4\beta-3\gamma)k_{\rm eff}\phi(\partial_r g)gT^4$	$-2. \times 10^{-6}$	GR	not previc	busiy known because
	21.	$\left(\omega_{ ext{eff}}-\omega_{a} ight)(\partial_{r}g)v_{L}T^{3}$	$2. \times 10^{-6}$		their calcul	ation requires a fully
	22.	$rac{7}{12} \left( \omega_{\mathrm{eff}} - \omega_a  ight) \left( \partial_r g  ight) g T^4$	$-1. \times 10^{-6}$		relativistic	calculation"
	23.	$-\frac{7}{12}(2-2\beta-\gamma)k_{\rm eff}g^{3}T^{4}$	$-3. \times 10^{-7}$	$\operatorname{GR}$		
	24.	$-rac{7k_{ m eff}^2}{2m_{ m o}}(\partial_r g)v_LT^3$	$-2. \times 10^{-7}$			
	25.	$-rac{27k_{ m eff}^2}{8m}(\partial_r g)gT^4$	$2. \times 10^{-7}$			
	26.	$rac{k_{ m eff}\omega_a}{m}gT^2$	$-1. \times 10^{-7}$			
	27.	$6(2-2\beta-\gamma)k_{\mathrm{eff}}\phi g^2T^3$	$5.  imes 10^{-8}$	$\operatorname{GR}$		
	28.	$3\left(\omega_{ ext{eff}}-\omega_{a} ight)g^{2}T^{3}$	$4. \times 10^{-8}$			
	29.	$3\left(\omega_{\mathrm{eff}}-\omega_{a} ight)gv_{L}T^{2}$	$-4. \times 10^{-8}$			
	30.	$6(1-\beta)k_{ m eff}\phi g v_L T^2$	$3.  imes 10^{-8}$	$\operatorname{GR}$		4

# Single photon vs. two photon AI

Breakdown the origin of various phase terms in the relativistic calculation (for Raman transitions)

	-				-	-
	Parameter dependence	Total phase shift coefficient	Propagation phase coefficient	Separation phase coefficient	Laser phase coefficient	Size (rad)
1.	$k_{\rm eff}T^3(\partial_r g)$	0	1	-1	0	$4  imes 10^{10}$
2.	$k_{\rm eff}gT^2$	-1	-1	1	-1	$3 \times 10^{8}$
3.	$\omega_{\rm eff}gT^2$	1	2	-2	1	$3 \times 10^{3}$
4.	$\omega_a g T^2$	-1	-1	0	0	$3 \times 10^{3}$
5.	$k_{\rm eff}(\partial_r g)T^3 v_L$	-1	2	-2	-1	$2 \times 10^3$
6.	$k_{\rm eff}(\partial_r g)\phi T^3$	0	$2\gamma + 2\beta - 2$	$-2\gamma - 2\beta + 2$	0	$3 \times 10^{1}$
7.	$k_{\rm eff}gT^2v_L$	-3	-5	5	-3	$1 \times 10^{1}$
8.	$k_{\rm eff}g\phi T^2$	$2-2\beta-\gamma$	$2-2\beta-\gamma$	$-2 + 2\beta + \gamma$	$2-2\beta-\gamma$	$2 \times 10^{-1}$
9.	$k_{\rm eff}g^2T^3v_L$	-12	-17	17	-12	$7 \times 10^{-6}$
10.	$k_{\rm eff}\partial_r(g\phi)T^3v_L$	$2-2\beta-\gamma$	$-4 + 4\beta + 2\gamma$	$4-4eta-2\gamma$	$2-2\beta-\gamma$	$2 \times 10^{-6}$
11.	$k_{\rm eff}gT^2v_L^2$	-5	-9	9	-5	$5 \times 10^{-7}$

- GR calculation gives terms arising from **change of rest mass** in different atomic states.
- In the case of a single photon interferometer, this is the dominate effect.
- Phase shift comes from propagation phase, not laser phase.
- Leading single photon AI phase shift **does not depend on the laser wavevector.**
- Non-relativistic single photon calculation incorrectly predicts kgT<sup>2</sup>

# **General Relativity Effects**

Schwarzschild metric, PPN expansion:



*Coordinate* acceleration, from geodesic equation

Corresponding AI phase shifts:

	Phase Shift	Size (rad)	Interpretation
1.	$-k_{\text{eff}}gT^2$	$3 \times 10^8$	gravity
2.	$-k_{\text{eff}}(\partial_r g)T^3v_L$	$-2 \times 10^3$	1st gradient
3.	$-3k_{\rm eff}gT^2v_L$	$4 \times 10^{1}$	Doppler shift
4.	$(2 - 2\beta - \gamma)k_{\text{eff}}g\phi T^2$	$2 \times 10^{-1}$	$\operatorname{GR}$
5.	$-\frac{7}{12}k_{\text{eff}}(\partial_r^2 g)T^4 v_L^2$	$8 \times 10^{-3}$	2nd gradient
6.	$-5k_{\text{eff}}gT^2v_L^2$	$3 \times 10^{-6}$	$\operatorname{GR}$
7.	$(2-2\beta-\gamma)k_{\text{eff}}\partial_r(g\phi)T^3v_L$	$2 \times 10^{-6}$	GR 1st grad
8.	$-12k_{\text{eff}}g^2T^3v_L$	$-6 \times 10^{-7}$	$\operatorname{GR}$

#### Projected experimental limits:

Tested	current	AI	AI	AI	AI far
Effect	limit	initial	upgrade	future	future
PoE	$3\times 10^{-13}$	$10^{-15}$	$10^{-16}$	$10^{-17}$	$10^{-19}$
PPN $(\beta, \gamma)$	$10^{-4} - 10^{-5}$	$10^{-1}$	$10^{-2}$	$10^{-4}$	$10^{-6}$

(Dimopoulos, et al., PRL 2007; PRD 2008)

### **Measurement Strategies**

Can these effects be distinguished from backgrounds?

1. Velocity dependent gravity (Kinetic Energy Gravitates)

- $\rightarrow$  Phase shift  $5k_{\rm eff}gT^2v_L^2$  has unique scaling with  $v_L$  , T
- $\rightarrow$  Compare simultaneous interferometers with different  $v_L$ .
- 2. Non-linear gravity (Gravity Gravitates)

So, in GR in vacuum:

$$\nabla \cdot \frac{d\mathbf{v}}{dt} \neq 0$$

Can discriminate from Newtonian gravity using three axis measurement

Violates "Gauss's law" for gravity

Gravitational field

# Example: Gravitational waves

Metric in TT gauge

$$ds^{2} = dt^{2} - (1 + h\sin(\omega(t - z)))dx^{2} - (1 - h\sin(\omega(t - z)))dy^{2} - dz^{2}$$

- Geodesics much simpler, since h is small
- For MAGIS-like configurations, can ignore motion of atom (recoil effects are corrections)
- GW signal arises from affect on null geodesics (in TT gauge)
- Arrival time of light determines atom propagation phase ('clock' phase)
- Laser phase is suppressed in a gradiometer configuration



# Atomic sensors for gravitational wave detection

Atomic clocks and atom interferometry offer the potential for gravitational wave detection in an *unexplored frequency range* ("mid-band")



#### Mid-band science

- LIGO sources before they reach LIGO band
- Sky localization: predict when and where events will occur (multi-messenger astronomy)
- Cosmological sources
- Wave-like dark matter (dilaton, ALP, ...)

# Sky position determination

Sky localization precision:

$$\sqrt{\Omega_s} \sim \left( \text{SNR} \cdot \frac{R}{\lambda} \right)^{-1}$$

#### Mid-band advantages

- Small wavelength  $\boldsymbol{\lambda}$
- Long source lifetime (~months) maximizes effective R

Benchmark	$\sqrt{\Omega_s}  [\mathrm{deg}]$
GW150914	0.16
GW151226	0.20
NS-NS (140 Mpc)	0.19



Graham et al., **PRD** 024052 (2018).

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# International efforts in atomic sensors for mid-band GW

MIGA: Matter Wave laser Interferometric Gravitation Antenna (France)



# AION: Atom Interferometer Observatory and Network (UK)



ELGAR: European Laboratory for Gravitation and Atom-interferometric Research



ZAIGA: Zhaoshan Long-baseline Atom Interferometer Gravitation Antenna (China)



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# **Gravitational Wave Detection**

$$ds^{2} = dt^{2} - (1 + h\sin(\omega(t - z)))dx^{2} - (1 - h\sin(\omega(t - z)))dy^{2} - dz^{2}$$







- LIGO and other optical interferometers **use two baselines**
- In principle, only one is required
- Second baseline needed to reject laser technical noise

# MAGIS concept

### Matter wave Atomic Gradiometer Interferometric Sensor

Passing gravitational waves cause a small modulation in the distance between objects. Detecting this modulation requires two ingredients:

#### **1. Inertial references**

- Freely-falling objects, separated by some baseline
- Must be insensitive to perturbations from non-gravitational forces
- 2. Clock
  - Used to monitor the separation between the inertial references
  - Typically measures the time for light to cross the baseline, via comparison to a precise phase reference (e.g. a clock).

In MAGIS, atoms play both roles.

Atom as "active" proof mass: Atomic coherence records laser phase, avoiding the need of a reference baseline – **single baseline** gravitational wave detector.

## Simple Example: Two Atomic Clocks







### Simple Example: Two Atomic Clocks



# MAGIS-100: Detector prototype at Fermilab

#### Matter wave Atomic Gradiometer Interferometric Sensor



- 100-meter baseline atom interferometry in existing shaft at Fermilab
- Intermediate step to full-scale (km) detector for gravitational waves
- Clock atom sources (Sr) at three positions to realize a gradiometer
- Probes for ultralight scalar dark matter beyond current limits (Hz range)
- Extreme quantum superposition states: >meter wavepacket separation, up to 9 seconds duration







LIVERPOOL **Fermilab** 



# MAGIS-100 design



# Clock atom interferometry

New kind of atom interferometry using **singlephoton transitions** between long-lived **clock states** 







#### Excited state phase evolution:

 $\Delta\phi\sim\omega_A\left(2L/c\right)$ 

(variations over time T)

Two ways for phase to vary: $\delta \omega_A$ Dark matter $\delta L = hL$ Gravitational wave

Graham et al., PRL **110**, 171102 (2013). Arvanitaki et al., PRD **97**, 075020 (2018).

# Ultralight (wave-like) dark matter

*Ultralight dilaton DM* acts as a background field (e.g., mass  $\sim 10^{-15}$  eV)

$$\mathcal{L} = +\frac{1}{2}\partial_{\mu}\phi\partial^{\mu}\phi - \frac{1}{2}m_{\phi}^{2}\phi^{2} - \sqrt{4\pi G_{N}}\phi \begin{bmatrix} d_{m_{e}}m_{e}\bar{e}e - \frac{d_{e}}{4}F_{\mu\nu}F^{\mu\nu} \end{bmatrix} + \dots$$

$$\begin{array}{c} \text{Electron} & \text{Photon} & e.g., \\ \text{coupling} & \text{coupling} & \text{QCD} \end{bmatrix}$$

$$\phi(t, \mathbf{x}) = \phi_{0}\cos\left[m_{\phi}(t - \mathbf{v} \cdot \mathbf{x}) + \beta\right] + \mathcal{O}\left(|\mathbf{v}|^{2}\right) \qquad \phi_{0} \propto \sqrt{\rho_{\text{DM}}} \quad \begin{array}{c} \text{DM mass} \\ \text{density} \end{bmatrix}$$

DM coupling causes time-varying atomic energy levels:



# MAGIS-100 projected sensitivity



# MAGIS-style satellite detector



Satellite detector concept

- Two spacecraft, MEO orbit
- Atom source in each
- Heterodyne laser link
- Resonant/LMT sequences

#### Example design

- $L = 4 \times 10^7$  meters
  - $10^{-4} \text{ rad}/\sqrt{\text{Hz}}$

 $\frac{n\hbar k}{m}T < 1 \text{ m}$ 2TQ < 300 s $n_p < 10^3$ 



- Heterodyne link concept analogous to LISA (synthesize ranging between two test masses)
- Decouples atom-laser interaction strength from baseline length (diffraction limit)

# Full scale MAGIS projected GW sensitivity



- Mid-band GW sources detectable from ground and space
- Gravity gradient noise (GGN) likely limits any terrestrial detector at low frequencies
- Longer baselines available in space reduce requirements (e.g., LMT), but can impact frequency response at high frequencies
- Flexible detection strategies possible (broadband vs resonant) with different tradeoffs in sensitivity/bandwidth

# Development path

#### MAGIS detector development

Experiment	(Proposed) Site	Baseline $L$ (m)	$\begin{array}{c} \text{LMT Atom} \\ \text{Optics } n \end{array}$	Atom Sources	Phase Noise $\delta \phi \ (rad/\sqrt{Hz})$	
Sr prototype tower	Stanford	10	$10^{2}$	2	$10^{-3}$	State of
MAGIS-100 (initial)	Fermilab (MINOS shaft)	100	$10^{2}$	3	$10^{-3}$	the art
MAGIS-100 (final)	Fermilab (MINOS shaft)	100	$4 \times 10^4$	3	$10^{-5}$	
MAGIS-km	Homestake mine (SURF)	2000	$4 \times 10^4$	40	$10^{-5}$	
MAGIS-Space	Medium Earth orbit (MEO)	$4 \times 10^7$	$10^{3}$	2	$10^{-4}$	

Reaching required sensitivity requires extensive technology development in three key areas:

Sensor technology	State of the art	Target	GW sensitivity improvement
LMT atom optics	$10^{2}$	$10^{4}$	100
Spin squeezing	20  dB (Rb), 0  dB (Sr)	20  dB (Sr)	10
Atom flux	$\sim 10^6 \text{ atoms/s}$	$10^8$ atoms/s	10

- Phase noise improvement strategy is a combination of increasing atom flux and using quantum entanglement (spin squeezing).
- LMT requirement is reduced in space proposals (longer baselines)