

Tutorial 2: Neutrino Oscillations

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Derivation of oscillation probability

$$\mathcal{L}_{\text{CC}} = \sum_{\alpha=e,\mu,\tau} \left[\frac{g}{\sqrt{2}} \bar{\nu}_{\alpha L} \gamma^\mu \ell_{\alpha L} W_\mu^+ + \frac{g}{\sqrt{2}} \bar{\ell}_{\alpha L} \gamma^\mu \nu_{\alpha L} W_\mu^- \right]$$

$$\nu_{\alpha L} = \sum_j U_{\alpha j} \nu_{jL} \leftarrow \text{Fields with definite masses } m_j$$

PMNS mixing matrix

$$|\nu_\alpha\rangle = \sum_j U_{\alpha j}^* |\nu_j\rangle \leftarrow \text{States with definite masses } m_j$$

Let's treat $|\nu_j\rangle$ as plane waves. Then, at a distance L and time t after production

$$|\nu_\alpha(t, L)\rangle = \sum_j U_{\alpha j}^* e^{-iE_j t + ip_j L} |\nu_j\rangle$$

A neutrino detector measures the neutrino flavour β

$$\langle \nu_\beta | = \sum_j U_{\beta j} \langle \nu_j |$$

Derivation of oscillation probability

The amplitude for transition $\nu_\alpha \rightarrow \nu_\beta$

$$\mathcal{M}_{\alpha\beta} = \langle \nu_\beta | \nu_\alpha(t, L) \rangle = \sum_{j,k} U_{\alpha j}^* U_{\beta k} e^{-iE_j t + ip_j L} \underbrace{\langle \nu_k | \nu_j \rangle}_{\delta_{jk}} = \sum_j U_{\alpha j}^* U_{\beta j} e^{-iE_j t + ip_j L}$$

The oscillation probability

$$P_{\alpha\beta} = |\mathcal{M}_{\alpha\beta}|^2 = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* e^{-i(E_j - E_k)t + i(p_j - p_k)L}$$

In all practical cases neutrinos travel almost at the speed of light, so

$$p_j \approx p_k \equiv p \approx E \quad \text{and} \quad t \approx L$$

$$E_j = \sqrt{p^2 + m_j^2} \approx p + \frac{m_j^2}{2E}$$

$$P_{\alpha\beta} = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* e^{-i \frac{\Delta m_{jk}^2 L}{2E}} \quad \Delta m_{jk}^2 \equiv m_j^2 - m_k^2$$

2-flavour case

Exercise 1. Using the obtained formula, *i.e.*

$$P_{\alpha\beta} = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* e^{-i \frac{\Delta m_{jk}^2 L}{2E}} \quad \Delta m_{jk}^2 \equiv m_j^2 - m_k^2$$

and taking into account that in the case of two neutrino mixing

$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

derive the expression for the oscillation probability $P_{e\mu}$ of $\nu_e \rightarrow \nu_\mu$

(Define $\Delta m^2 \equiv m_2^2 - m_1^2$)

Hint: Use the trigonometric identities

$$\sin(2\theta) = 2 \sin(\theta) \cos(\theta)$$

$$\cos(2\theta) = 2 \cos^2(\theta) - 1$$

Experimentally relevant units

Exercise 2. Express the argument of the second sine in terms of experimentally relevant units, *i.e.* find # in

$$\frac{\Delta m^2 L}{4E} = \# \frac{\Delta m^2 (\text{eV}^2) L (\text{m})}{E (\text{MeV})}$$

Hint: Use the relation

$$\hbar c = 1 \approx 197 \text{ fm} \cdot \text{MeV}$$
$$(1 \text{ fm} = 10^{-15} \text{ m} \quad \text{and} \quad 1 \text{ MeV} = 10^6 \text{ eV})$$

Oscillation length

Exercise 3.

- Taking into account that the dependence of $P_{e\mu}$ on L/E is periodic, derive the expression for the **oscillation length** L_{osc} associated with Δm^2
- Express L_{osc} in meters for Δm^2 in eV^2 and E in MeV

$$P_{e\mu} = \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E}\right) \approx \sin^2(2\theta) \sin^2\left(1.27 \frac{\Delta m^2 (\text{eV}^2) L (\text{m})}{E (\text{MeV})}\right)$$

Typical oscillation lengths

Exercise 4. Using the obtained formula, *i.e.*

$$L_{\text{osc}} \approx 2.48 \text{ m} \frac{E (\text{MeV})}{|\Delta m^2 (\text{eV}^2)|}$$

compute the oscillation length for

- reactor neutrinos with $E = 1 \text{ MeV}$, assuming $\Delta m^2 = 7.4 \cdot 10^{-5} \text{ eV}^2$
- accelerator neutrinos with $E = 1 \text{ GeV}$, assuming $\Delta m^2 = 2.5 \cdot 10^{-3} \text{ eV}^2$

What are representative neutrino oscillation experiments?

Where would you place the detector to measure oscillations driven by $\Delta m^2 = 10 \text{ eV}^2$ with $E \sim 2.5 \text{ GeV}$?

Oscillation maxima

Exercise 5. Assuming $\Delta m^2 = 2.5 \cdot 10^{-3} \text{ eV}^2$ and $L = 1300 \text{ km}$, derive the energies of the first and second **oscillation maxima**

$$P_{e\mu} = \sin^2(2\theta) \sin^2\left(1.27 \frac{\Delta m^2 (\text{eV}^2) L (\text{km})}{E (\text{GeV})}\right)$$