

Thoughts on combining the experimental data

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Including content from

Isidori, Lancierini, Owen, Serra, [arXiv:2104.05631](https://arxiv.org/abs/2104.05631)

Isidori, Lancierini, Mathad, Owen, Serra, Coutinho, [arXiv:2110.09882](https://arxiv.org/abs/2110.09882)

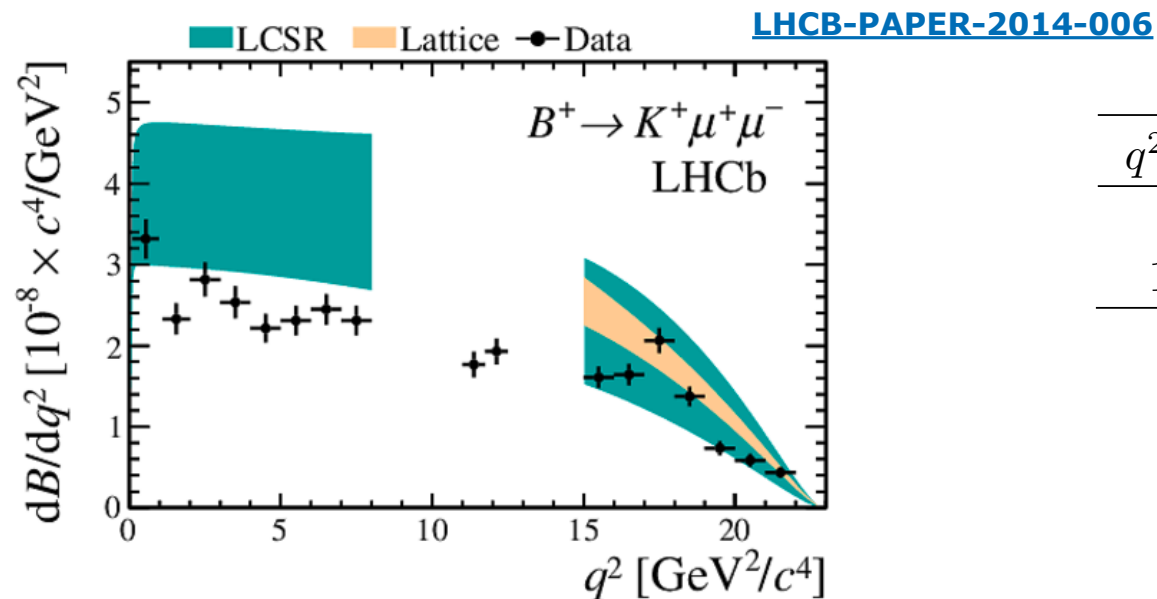


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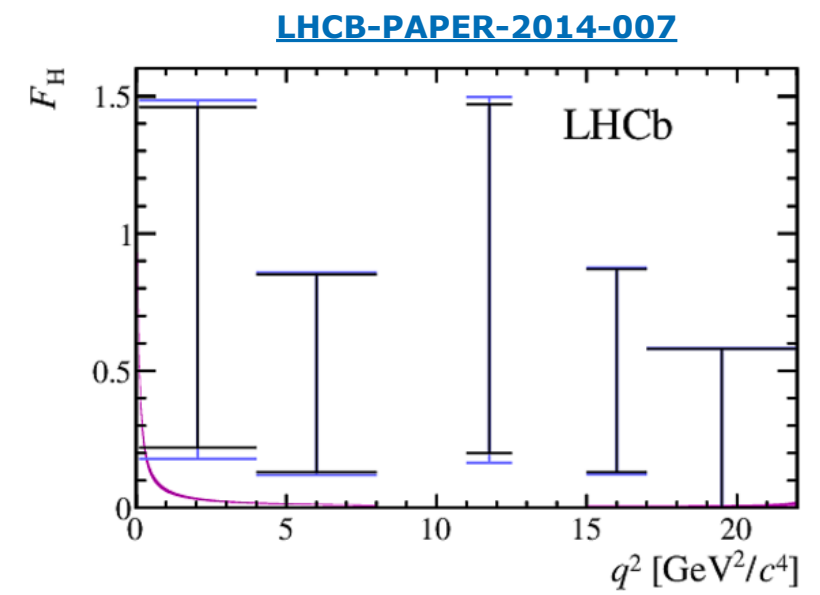
Some general considerations

- Statistical correlations between measurements, if large, will be given in the paper.
- Systematic correlations are instead often implicit.
 - e.g. $b \rightarrow s \mu \mu$ branching fraction.



q^2 range (GeV^2/c^4)	central value	stat	syst
$1.1 < q^2 < 6.0$	24.2	0.7	1.2
$15.0 < q^2 < 22.0$	12.1	0.4	0.6

- Less obvious is the correlation across papers. E.g. $B_s^0 \rightarrow \phi \mu \mu$.
- What to do if one is given an 68% CL, but nothing else?
 - Full FC scans now included in CDS material.



Questions the data can answer

- Two questions to ask of the data:
 - What are values of Wilson Coefficients? and do they deviate from the SM?
 - What is the global statistical significance of the new physics hypothesis?
- 1st question involves disentangling hadronic effects from new physics, requires theoretical calculations, models, amplitude fits etc.
 - While very important, this is not the focus of this talk.
- Instead we focus on the second question, and construct a highly general alternative hypothesis to produce a conservative answer.
 - We use C_9 as a SM nuisance parameter.
 - A case for making the hypothesis general (look-elsewhere-effect).
 - Inclusion of so-called non-exclusive R ratios (connection to Yasmine and Gianluca's talk).

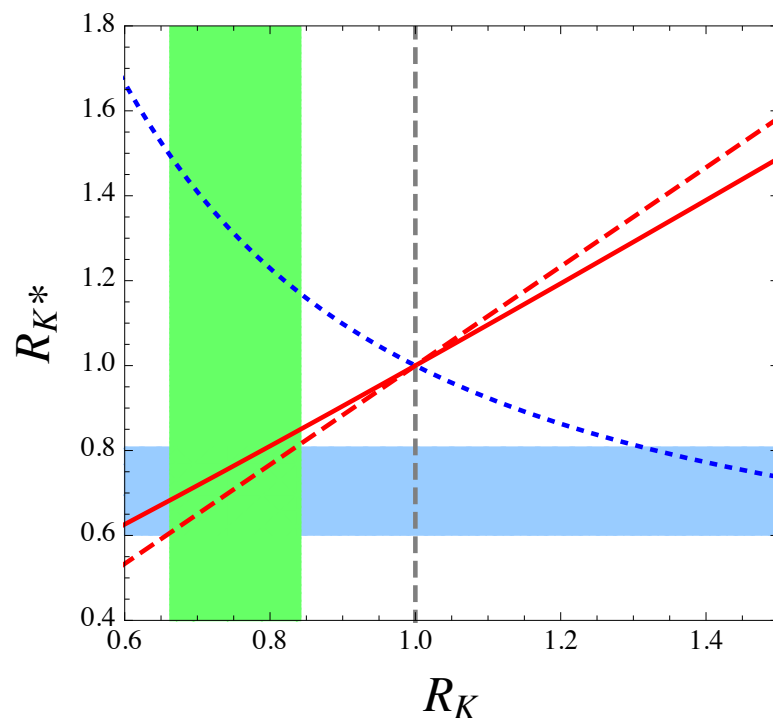
A case for generality

- Want to provide significance which can stand up to the skeptical.
- Original idea: Only combine observables for which there is wide consensus on the SM prediction.
 - Concretely: Combine $B(B_s^0 \rightarrow \mu\mu)$ and LFU ratios and fit for $C_{9,10}^\mu$.
- However, such a fit smuggles in information about the rest of the system:
 - No new physics in electrons (understandably justified from BF measurements).
 - No right-handed currents or scalar new physics (also reasonable).
- As we argued in [1], constraining the observables and alternative hypothesis in way leads to an overestimation of the significance via the look-elsewhere effect.
 - Only combining the observables that deviate.
 - Fitting with a restricted set of operators.

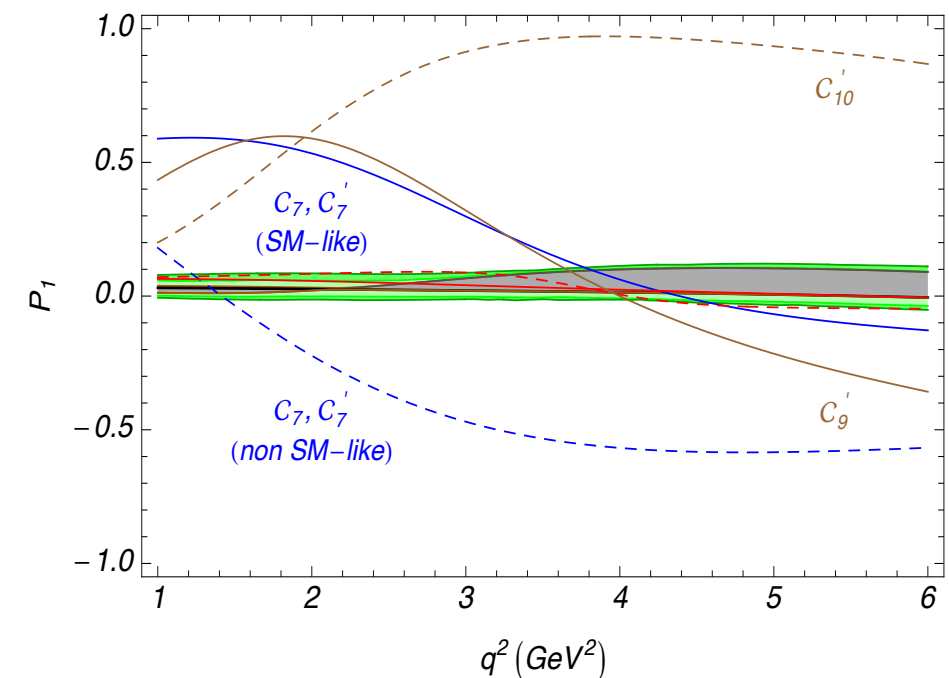
The look elsewhere effect

- The look elsewhere effect (LEE) occurs when the alternative hypothesis test implicitly uses the central values of the data.
 - Example is testing the presence of a resonance of the mass seen in data.
- In our case, floating only left-handed LFUV Wilson Coefficients ignores alternate universes where we saw deviations in $K^*\mu\mu$ and/or different values for R_K and R_{K^*} .

Hiller, Nisandzic, 2017



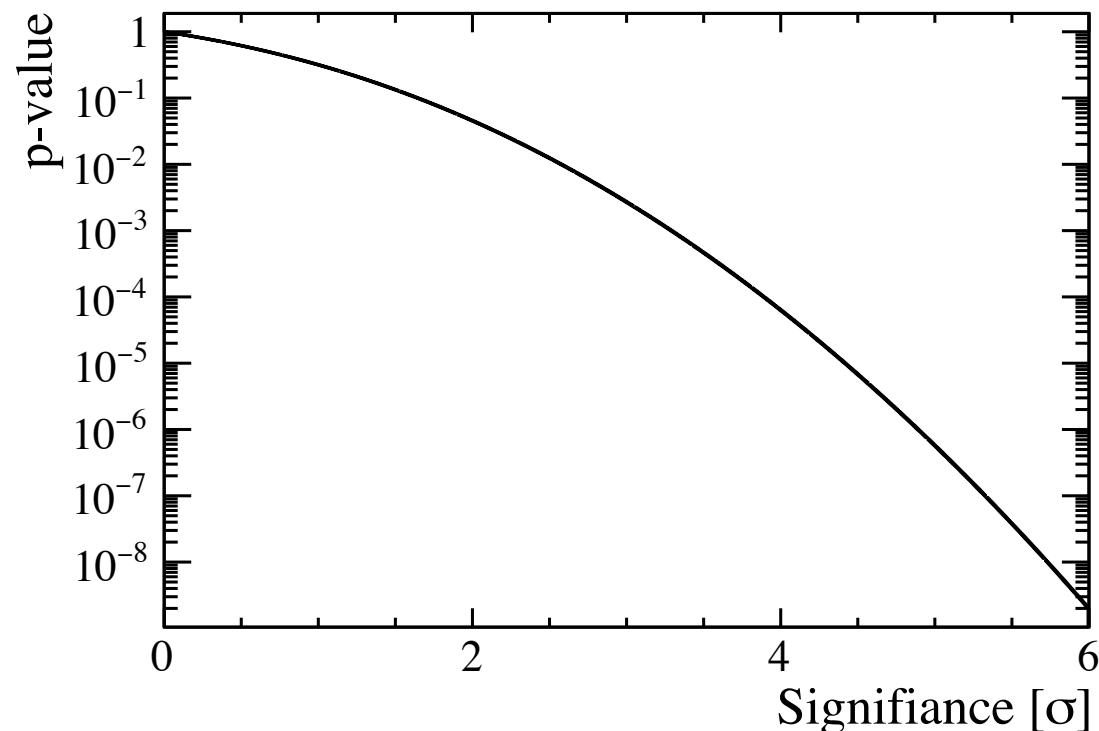
Matias et al, 2012



- Fitting only LFU/clean observables are therefore **local** significances, in analogy with local p-value at a particular mass point in a bump hunt.

Common talking points on the LEE

1. **The LEE only occurs in mass peak searches.** The look-elsewhere effect originates from the 1950s, known as the *problem of multiple comparisons*, and can effect any situation whereby the data is implicitly used twice. Also known as the post-hoc analysis.
2. **The question can always be bigger, meaning that the LEE is subjective.** The main purpose of the LEE is to take you back to a hypothesis test which is a-priori, it does not always need to be bigger. (Otherwise it would be impossible to claim anything).
3. **The LEE has a smaller effect for large significances.** A trial factor of 10 will dilute a significance of 3.0σ to 2.2σ whereas a 5.0σ will only move to 4.4σ . This is simply due to the non-linear relationship between a Gaussian significance and the p-value.



Approach in arXiv:2104.05631

- Write down all operators to which LHCb measurements were sensitive to.

$$\begin{aligned}\mathcal{O}_9^\ell &= (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \ell), & \mathcal{O}_{10}^\ell &= (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \gamma_5 \ell) \\ \mathcal{O}_9^{\ell'} &= (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \ell), & \mathcal{O}_{10}^{\ell'} &= (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \gamma_5 \ell), \\ \mathcal{O}_{\hat{S}}^\ell &= (\bar{s}_L b_R)(\bar{\ell}_R \ell_L), & \mathcal{O}_{\hat{S}}^{\ell'} &= (\bar{s}_R b_L)(\bar{\ell}_L \ell_R).\end{aligned}$$

- Use amplitudes and form factors from Flavio[1] to translate observables into WC space.
- Include $C7^{(\ell)}$ with a-priori constraint from $b \rightarrow s \gamma$ B-factory results (0.2σ impact).
- Consider one scalar contributes to $B_s \rightarrow \mu\mu$ (only one independent contribution assuming $\Lambda_{NP} \gg \text{VEV}$) [Alonso, 2014].
- Assume no scalars in Kee.
- Assume that WC are real.
- End up with 9 WC to which the measurements are sensitive.
 - 4 muonic WC: $C_9, C_{10}, C_9', C_{10}'$
 - 4 LFUV WC: $\Delta C_9, \Delta C_{10}, \Delta C_9', \Delta C_{10}'$ [Difference in muonic/electronic WC]
 - 1 scalar contribution in muons: $C_S - C_S'$

Approach in arXiv:2104.05631

- Observables included: R_K , R_{K^*} , $B_s \rightarrow \mu\mu$ and $B^0 \rightarrow K^{*0}\mu\mu$ angular analysis.
- Generate toys based from SM predictions and experimental uncertainties.
 - R_K and R_{K^*} uses full likelihood but assumed to be independent.
 - $K^*\mu\mu$ observables generated/fit with full experimental correlations.

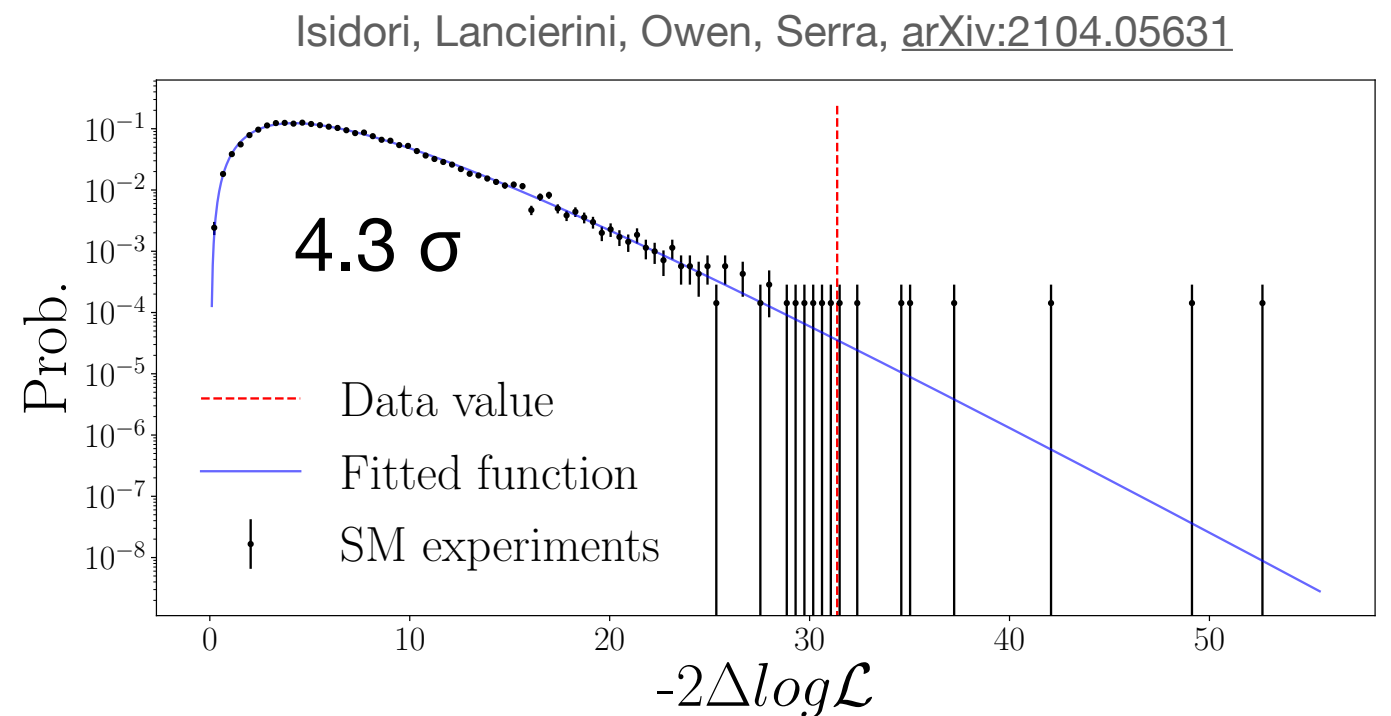
- For each toy calculate test statistic.

$$\Delta\chi^2 = -2 \log \frac{\mathcal{L}(X|\Delta\hat{C}_9^U, C_i^{\text{SM}})}{\mathcal{L}(X|\hat{C}_i)}$$

- Compare distribution of SM toys to one seen in data. Integrate to get significance.

- Why use toys and not Wilk's theorem?

- Can get flat directions in WC space.
- Effective degrees of freedom in system not necessarily integer number.
- E.g. $B_s \rightarrow \mu\mu$ and C_{10} vs C_s .

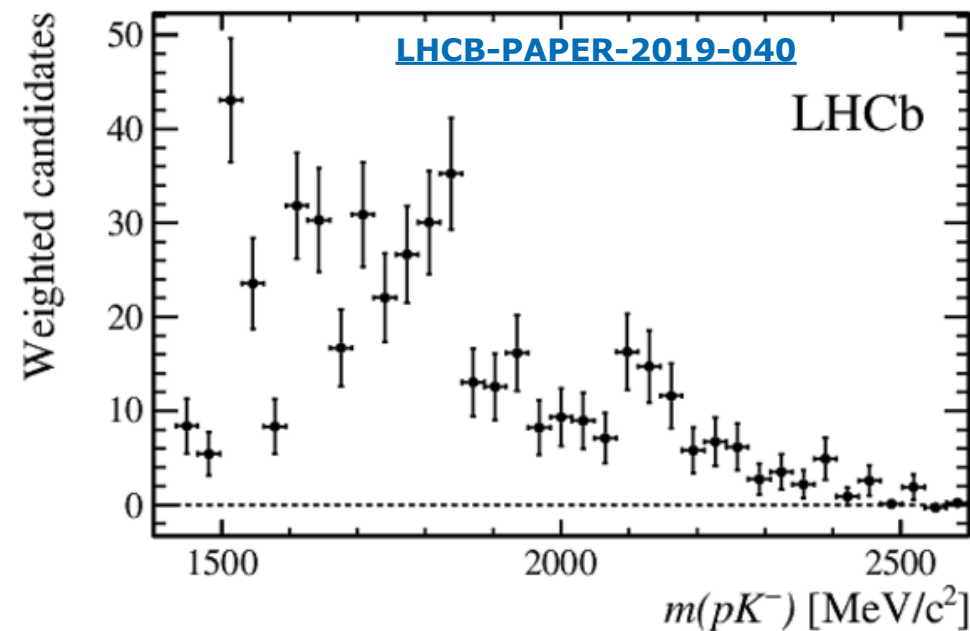


Importance of non-exclusive LFU ratios

- Several LFU ratios contain decays with broad, overlapping resonances whose hadronic structure is unknown.

- Most prominent example: R_{pK}

$$R_{pK} = \frac{\mathcal{B}(\Lambda_b^0 \rightarrow pK^- \mu^+ \mu^-)}{\mathcal{B}(\Lambda_b^0 \rightarrow pK^- e^+ e^-)}$$



- If R_{pK} deviates significantly from unity then it's NP. However, translating that into WC space is tricky for such a large mass range.
 - Experimental/theoretical progress would help in this particular case.
- However, the LFU ratio is likely to stay inclusive, as it maximises the precision.
 - Can we include these measurements already in a combination?

Inclusion of non-exclusive LFU ratios

Isidori et al, [arXiv:2110.09882](https://arxiv.org/abs/2110.09882)

- Neglecting lepton masses ($q^2 \gg m_l^2$), no interference between left and right handed lepton currents. [Hiller, Schmaltz, 2014]

$$\frac{d\Gamma_X^\ell}{dq^2} = \frac{d\Gamma_{X,L}^\ell}{dq^2} + \frac{d\Gamma_{X,R}^\ell}{dq^2}$$

$$\frac{d\Gamma_{X,R}^\ell}{dq^2} = \frac{d\Gamma_{X,L}^\ell}{dq^2} \Big|_{\{C_L^\ell \rightarrow C_R^\ell, C_L^{\ell'} \rightarrow C_R^{\ell'}\}}$$

$$\begin{aligned} C_L^\ell &= C_9^\ell - C_{10}^\ell, & C_L^{\ell'} &= C_9^{\ell'} - C_{10}^{\ell'} \\ C_R^\ell &= C_9^\ell + C_{10}^\ell, & C_R^{\ell'} &= C_9^{\ell'} + C_{10}^{\ell'} \end{aligned}$$

Get general formula, applicable to any LFU ratio, but only used for non-exclusive modes.

$$R_X = \frac{\left\{ (C_L^\mu)^2 + (C_L^{\prime\mu})^2 + \langle \eta_X^0 \rangle C_L^\mu C_L^{\prime\mu} + C_7 \cdot (\langle \eta_X^{77} \rangle C_7 + \langle \eta_X^{79} \rangle C_L^\mu + \langle \eta_X^{\prime 79} \rangle C_L^{\prime\mu}) \right\} + (L \rightarrow R)}{\left\{ (C_L^e)^2 + (C_L^{\prime e})^2 + \langle \eta_X^0 \rangle C_L^e C_L^{\prime e} + C_7 \cdot (\langle \eta_X^{77} \rangle C_7 + \langle \eta_X^{79} \rangle C_L^e + \langle \eta_X^{\prime 79} \rangle C_L^{\prime e}) \right\} + (L \rightarrow R)}$$

The η parameters encode hadronic information

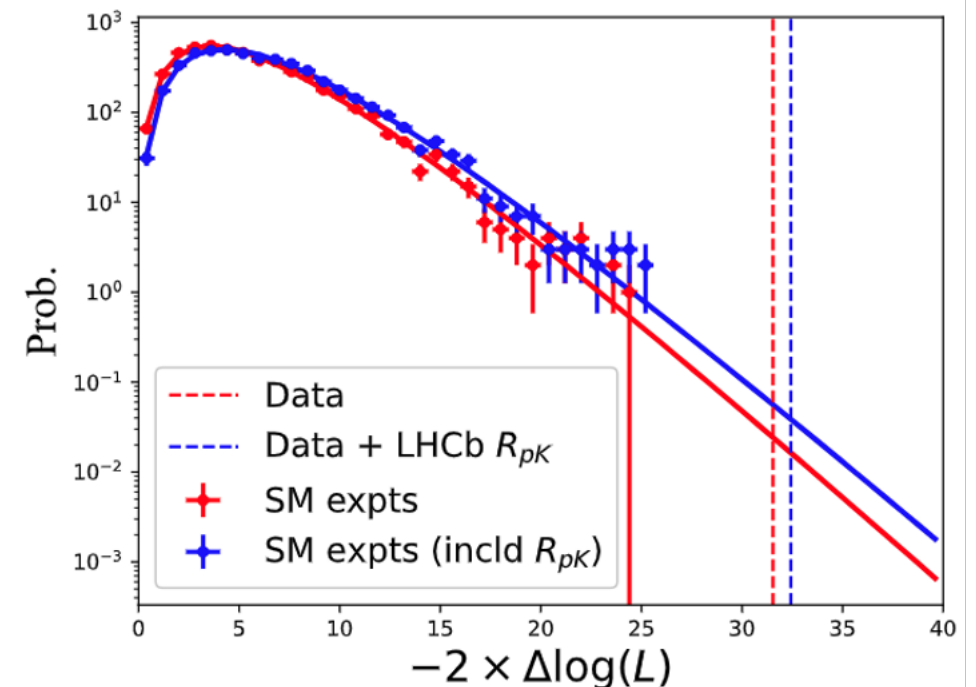
$$\begin{aligned} F_X^\ell &= \int_{q_{\min}^2}^{q_{\max}^2} f_X^\ell(q^2) dq^2 \\ \langle \eta_X^{i,\ell} \rangle &= \frac{1}{F_X^\ell} \int_{q_{\min}^2}^{q_{\max}^2} f_X^\ell(q^2) \eta_X^i(q^2) dq^2 \end{aligned}$$

Parameter	Limits	
	default	4q-ops
$\langle \eta_X^0 \rangle$		[-2,2]
$\langle \eta_X^{79} \rangle$	[-12,12]	[-20,20]
$\langle \eta_X^{79'} \rangle$	[-4,4]	[-10,10]
$\langle \eta_{pK}^{77} \rangle$	[0,120]	[0, 160]
$\langle \eta_{K\pi, K\pi\pi}^{77} \rangle$	[0,60]	[0,100]

Impact on the significance

Isidori, Lancierini, Mathad, Owen,
Serra, Coutinho, [arXiv:2110.09882](https://arxiv.org/abs/2110.09882)

- Inclusion of R_{pK} slightly decreases significance.
- Why?
 - Not perfectly aligned with other R ratios.
 - Quite consistent with SM.
 - Lots of hadronic uncertainty.
- Projecting precision from full run II shows substantial improvement in discovery potential.
 - Non-exclusive ratios can have a large impact here.
- Fixing the hadronic parameters has reasonably small impact on significance ($<0.5\sigma$).



Linearising the expression w.r.t. $\Delta(C_i)$ and neglecting interference with suppressed amplitudes (e.g. C_i').

$$R_X - 1 \approx \frac{\text{Re} \left(2 \frac{\Delta C_L}{C_L^{\text{SM}}} + \langle \eta_X^0 \rangle \frac{\Delta C_L'}{C_L^{\text{SM}}} \right)}{1 + \langle \eta_X^{77} \rangle \left| \frac{C_7^{\text{SM}}}{C_L^{\text{SM}}} \right|^2 + \text{Re} \left[\langle \eta_X^{79} \rangle \frac{C_7^{\text{SM}}}{C_L^{\text{SM}}} \right]}$$

Discussion points

- Inclusion of upper limits on $B_q \rightarrow ee$ and $b \rightarrow s\tau\tau$
- How to treat C_9 ?
 - Currently included as SM nuisance parameter.
 - Comments to make it q^2 /helicity dependent.
- Finally, a more general one...

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