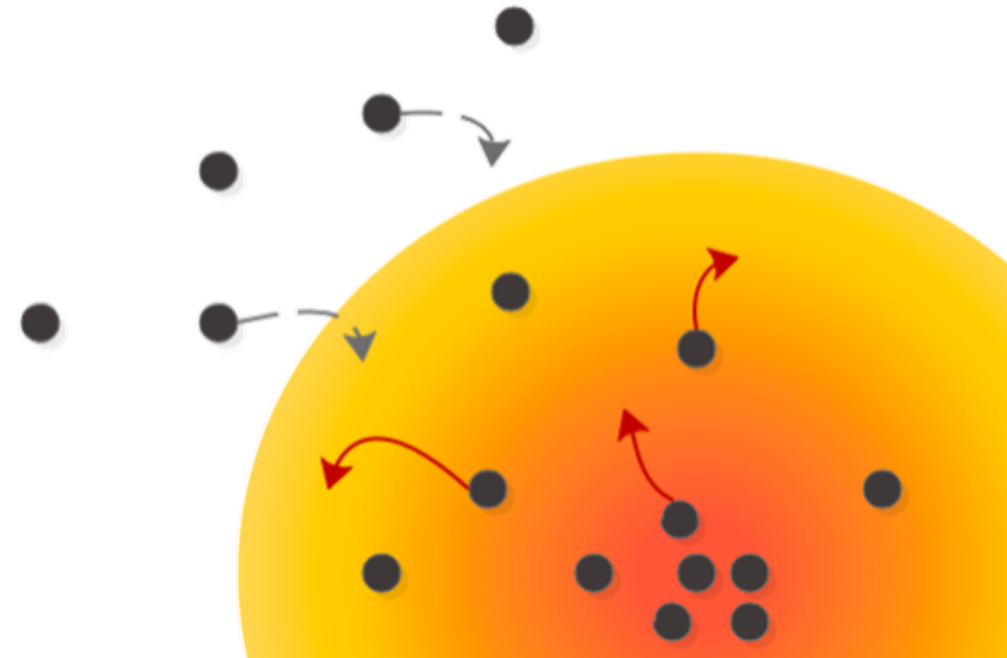


# Simulation of energy transport by dark matter scattering in the Sun

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Based on 2111.06895 with S. Ansari, A.C Vincent and P. Scott



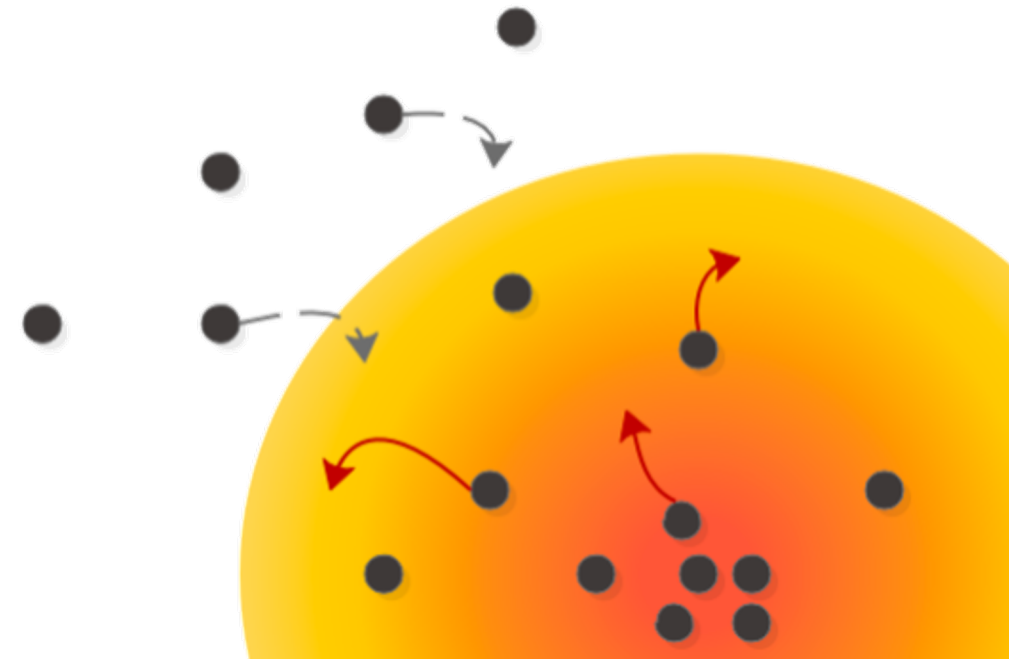
# The Idea

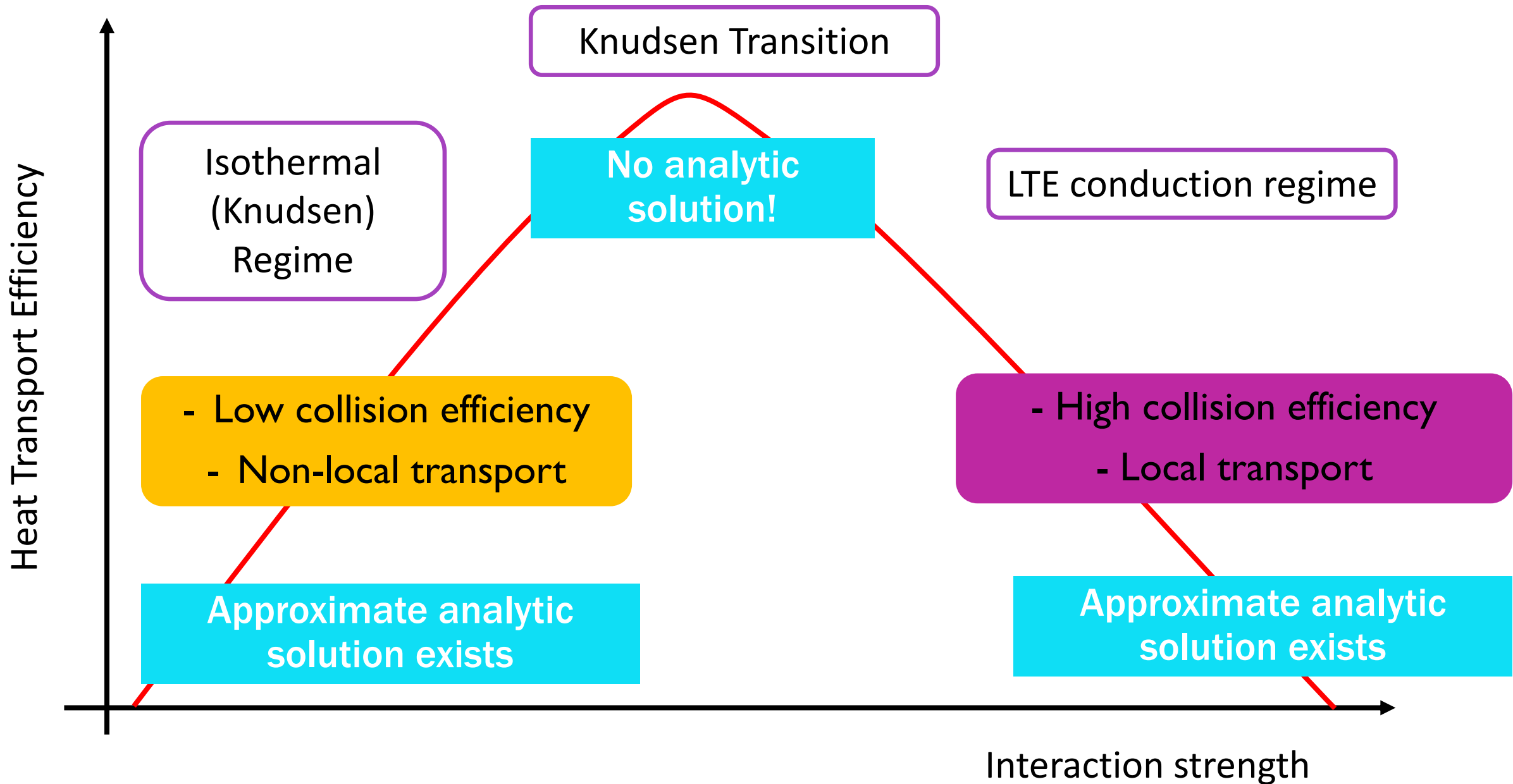
- Can use the Sun as a “direct detection experiment”
- (Asymmetric) DM particles can become trapped in the Sun
- Interactions with nuclei conducts heat, modifying internal structure and properties

Probe by



1. Helioseismology
2. Neutrino fluxes





# The existing formalisms:


## Spergel & Press

Based on Isothermal  
“solutions”

VS

## Gould & Raffelt

Based on modified LTE  
“solutions”

- 
1. Use assumptions only valid for one regime
  2. Calibrated using a single, simplified simulation done > 30 years ago
  3. Only tested for constant interactions

*So, is either accurate?*

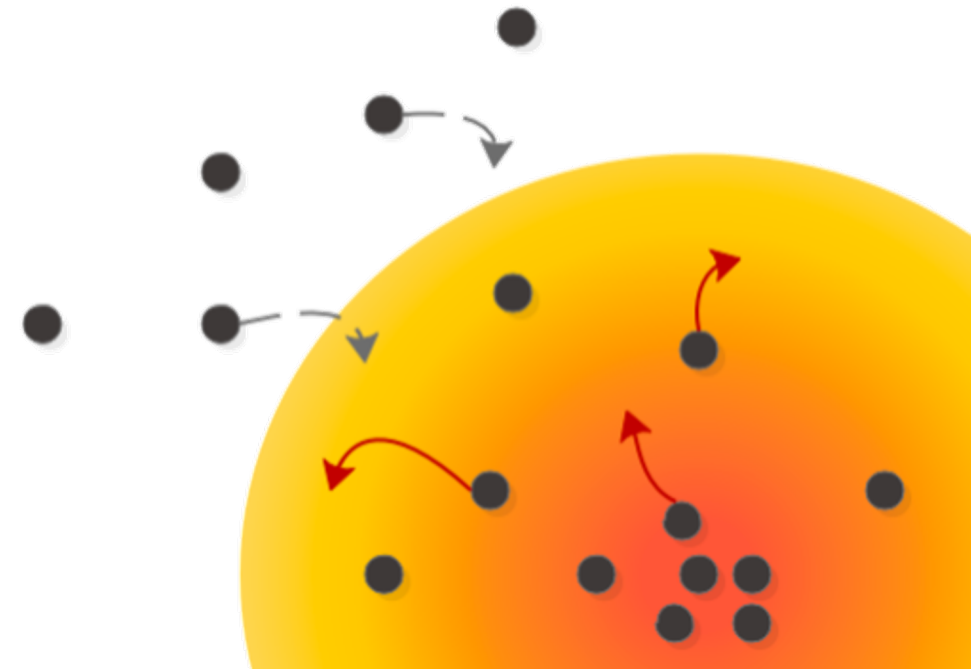
- Simulate chains of  $10^7$ - $10^8$  DM-nucleon collisions
- Track DM velocity, position and energy transfer at collision points
- Directly extract thermal properties from distributions
- Consider cross sections that depend on relative velocity of DM and nucleon,  $v_{\text{rel}}$  or momentum transferred,  $q$ :

$$\sigma = \sigma_0 \left( \frac{v_{\text{rel}}}{v_0} \right)^{2n}$$

$$\sigma = \sigma_0 \left( \frac{q}{q_0} \right)^{2n}$$

... Monte Carlo!

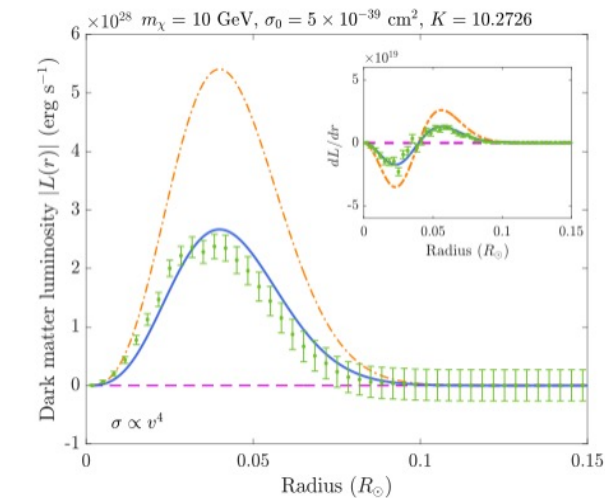
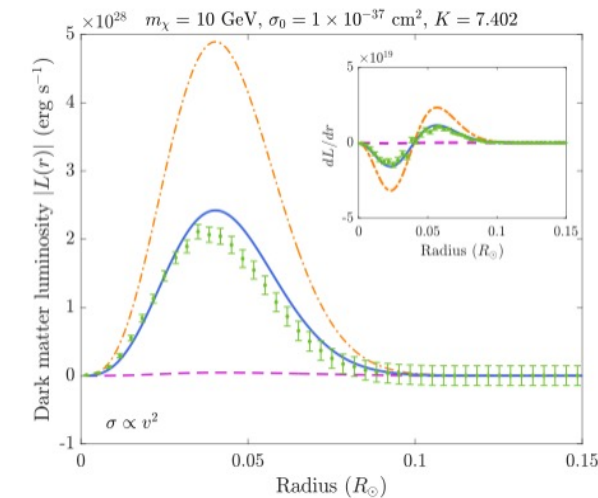
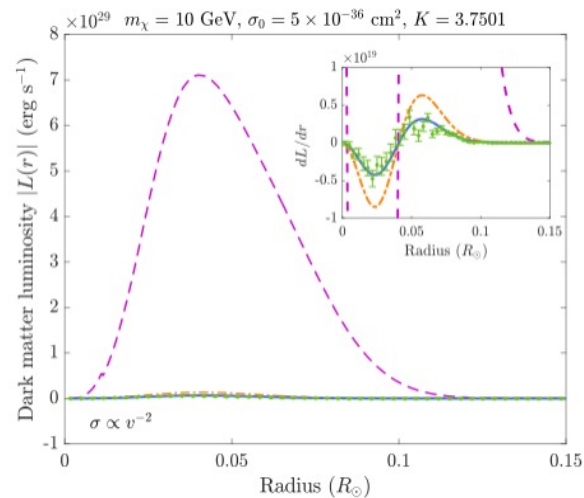
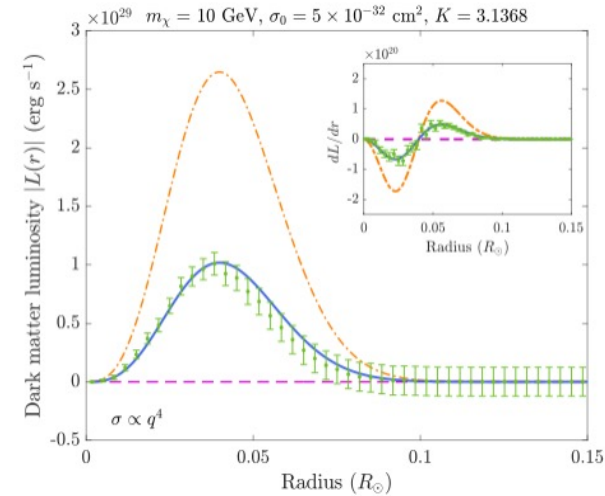
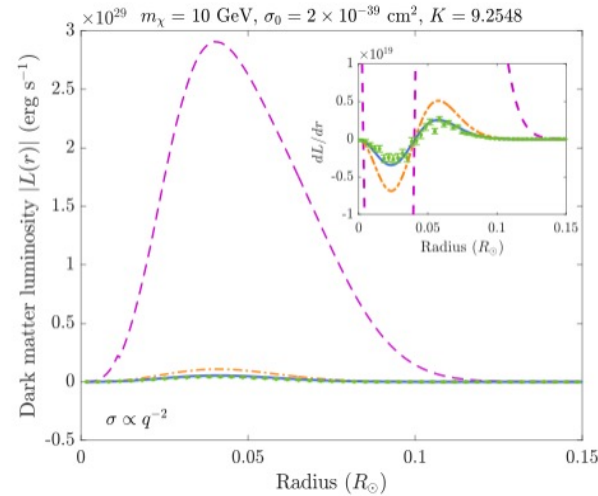
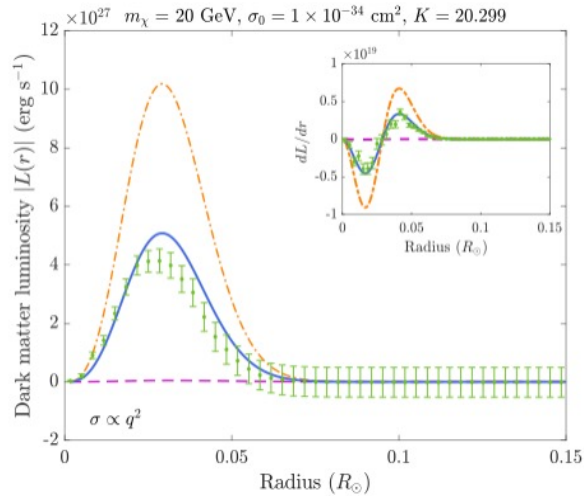
Billions of CPU hrs later....



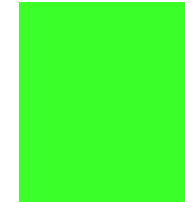
# Neither G&R not S&P give a good description!



DM Luminosity and Energy Transport (inset) distributions for various interaction types and strengths:



= S & P



= Monte-Carlo

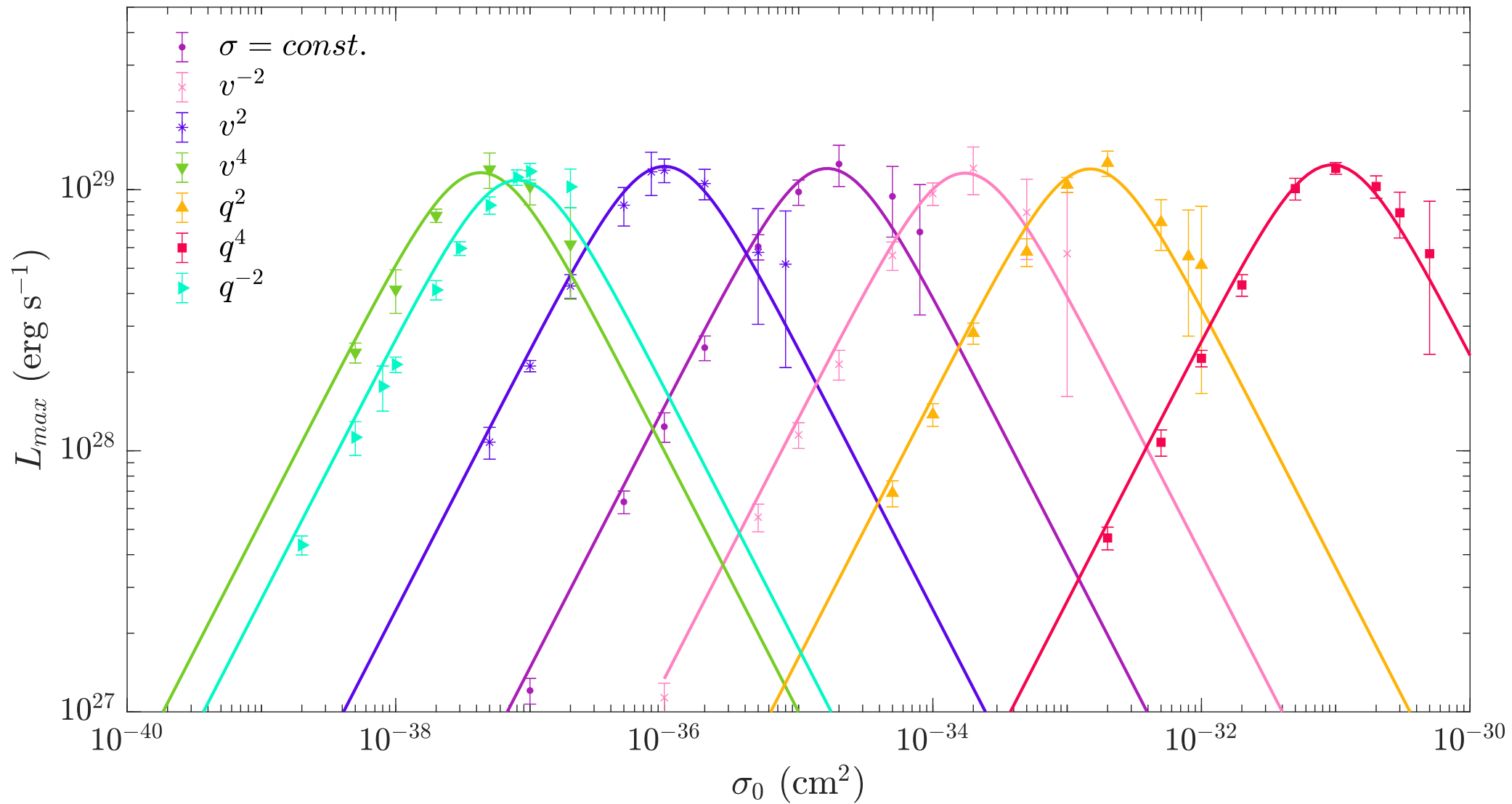


= G&R



= Our scheme

But a simple universal rescaling of S&P scheme works for all interaction types and across all regimes!



# Summary

- First direct simulation of DM mediated energy transport in stars performed for non-constant cross-sections
- Neither of the conventional approaches accurately describe DM mediated conduction in realistic simulations
- Simple, universal, rescaling of Spergel & Press formalism accurate for all interaction types and regimes



Can now find new bounds on properties of DM!

