Cosmological and Quantum Gravity applications of QRG **On polygons**

Julio Narciso Argota Quiroz, QMUL

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Spacetime at Plank scale



Quantum spacetime hypothesis: space-time coordinates better modelled as noncommutative due to quantum gravity effects

Quantum Riemannian Geometry Differential calculus

- An algebra A over a field K
- An A-bimodule Ω^1
- A linear map $d: A \to \Omega^1$
- $\Omega^1 = \operatorname{span}\{a \cdot db \mid a, b \in A\}$
- Ker $d = k \cdot 1_A$, $k \in K$, $1_A \in A$
- $d: \Omega^n \to \Omega^{n+1}$

$$(a \cdot \omega) \cdot b = a \cdot (\omega \cdot b), \ a, b \in A, \ \omega \in A$$

$$d(ab) = (da) \cdot b + a \cdot (db)$$

Surjectivity

Connectedness (Optional)

 $d^2 = 0$



Geometrical Structures on a differential calculus (A, Ω^1 , d) $g \in \Omega^1 \otimes_A \Omega^1$ Metric $(,): \Omega^1 \otimes_A \Omega^1 \to A$ Inverse metric $\nabla: \Omega^1 \to \Omega^1 \otimes_{\scriptscriptstyle A} \Omega^1$ Connection $\sigma: \Omega^1 \otimes_A \Omega^1 \to \Omega^1 \otimes_A \Omega^1$ Braiding map Ricci $\in \Omega^1 \otimes_A \Omega^1$ S = (,) Ricci $i: \Omega^2 \to \Omega^1 \otimes \Omega^1$ $\nabla(f\omega) = \mathrm{d}f \otimes \omega + f\nabla\omega \qquad \nabla(\omega f) = \sigma(\omega \otimes \mathrm{d}f) + (\nabla\omega)f$ $\nabla(\omega \otimes \eta) = \nabla \omega \otimes \eta + (\sigma \otimes \mathrm{id})(\omega \otimes \nabla \eta)$ $\nabla g = 0$ $T_{\nabla} = \wedge \nabla - \mathrm{d}$ $T_{\nabla}: \Omega^1 \to \Omega^2$ QLC

Quantization of \mathbb{Z}_n

FLRW model, Class. Quantum Grav. (2020) 245001 (43pp)



J. N. Argota-Quiroz and S. Majid, <u>Quantum gravity on polygons and $\mathbb{R} \times \mathbb{Z}_{p}$ </u>

Euclideanized Quantum Gravity

Algebra
$$A = \mathbb{C}(\mathbb{Z}_n)$$

$$e^{\pm}f = R_{\pm}fe^{\pm}, \quad df = \sum_{\pm} \partial_{\pm}fe^{\pm}; \quad (R_{\pm}f)(i)$$

Metric $g = ae^+ \otimes e^- + be^- \otimes e^+$ $b = R_{(a)}$

Inverse metric

$$(e^+, e^-) = 1/R_+(b)$$
 $(e^-, e^+) = 1/R_-(a)$ $(e^\pm, e^\pm) = 0$
 Ω^2 calculus
 $(e^\pm)^2 = 0, e^+ \wedge e^- + e^- \wedge e^+ = 0, de^\pm = 0$

 $= f(i \pm 1), \quad \partial_{+} = R_{+} - id$

Edge symmetry condition

Calculus Ω N-gon $a(i) = g_{i \to i+1}$ $b(i+1) = g_{i+1 \rightarrow i}$ -1n - 12

0 $a(0) = g_{0 \rightarrow 1}$ $b(1) = g_{1 \to 0}$

Geometrical structures for \mathbb{Z}_n

•
$$g = ae^+ \otimes e^- + R_-(a)e^- \otimes e^+$$

• $\nabla e^+ = (1 - \rho)e^+ \otimes e^+, \quad \nabla e^- = (1 - R_-^2 \rho^{-1})e^+$
• $R_{\nabla}e^+ = \partial_-\rho e^+ \wedge e^- \otimes e^+, \quad R_{\nabla}e^- = -\partial_+(R_-^2 \rho^{-1})e^+$
• Ricci $= \frac{1}{2} \left(\partial_-(R_-\rho)e^- \otimes e^+ - \partial_-\rho^{-1}e^+ \otimes e^- \right), \quad S = \frac{1}{2} \left(-\frac{\partial_-\rho^{-1}}{a} + \frac{\partial_-(R_-\rho)}{R_-a} \right), \quad \Delta f = -\frac{R_-(\rho) + 1}{a} (\partial_+ + \partial_-)f$
• $\Delta f = -\frac{R_-(\rho) + 1}{a} (\partial_+ + \partial_-)f$
 $\int S = \sum_{\mathbb{Z}_n} aS = \frac{1}{4} \sum_{\mathbb{Z}_n} \rho(\partial_+ + \partial_-)\rho$

Metric

 $e^{-1})e^{-}\otimes e^{-1}$

 $(r^{-1})e^+ \wedge e^- \otimes e^-$

QLC

Curvature

Ricci tensor

Ricci scalar $\partial_{\pm} f(i) = f(i+1) \pm f(i)$

 $\rho = \frac{R_+(a)}{}$

a

Laplacian

Hilbert-Einstein Action



Quantum Gravity Applications



Euclideanized Quantum Gravity

Closed boundary condition $\rho_0 \rho_1 \cdots \rho_{n-1} = 1$ constraint manifold with metric \mathfrak{g}_{ρ}

Measure for integration over
$$\rho$$

Taking n = 3, and ρ_0, ρ_1 as coordinates

$$D\rho = (\prod_{i=0}^{n-2} d\rho_i) \sqrt{\det(\mathfrak{g}_{\rho})}$$

$$det(\mathfrak{g}_{\rho}) = 1 + \frac{1}{\rho_0^4 \rho_1^2} + \frac{1}{\rho_0^2 \rho_1^4}$$

$$Z = \int_0^\infty d\rho_0 \int_0^\infty d\rho_1 \sqrt{\det(\mathfrak{g}_{\rho})} e^{-\frac{1}{2G}(\rho_0^2 + \rho_1^2 + \rho_2^2 - \rho_0 \rho_1 - \rho_1 \rho_2 - \rho_2 \rho_0)}$$

For large G

$$\frac{\Delta \rho_i}{\langle \rho_i \rangle} \sim 1.1$$

$$\frac{\langle \rho_i \rho_j \rangle}{\langle \rho_i \rangle \langle \rho_j \rangle} \ge .808$$

Uniform uncertainty in metric

For small G







Quantize fluctuations relative to the geometrical mean for n = 3

With the variables
$$b_i = a_i/A$$

Partition function





And the restriction

 $A = (\prod_i a_i)^{\frac{1}{n}}$ $b_0 b_1 \dots b_n = 1$

Mesure $da_0 da_1 da_2 = \frac{3A^2}{b_0 b_1} db_0 db_1 dA$. Action $S_g = \frac{1}{2} \left(\frac{3}{b_1} + \frac{3}{b_2} + \frac{3}{b_2} - (\frac{3}{b_1})^2 - (\frac{3}{b_1})^2 - (\frac{3}{b_2})^2 \right)$



Cosmological applications

Geometrical structures of FLRW model

$$g = -\mathrm{d}t \otimes \mathrm{d}t - \mathrm{R}^{2}(t)e^{+} \otimes_{s} e^{-}$$
$$\nabla \mathrm{d}t = R\dot{R}e^{+} \otimes_{s} e^{-}, \quad \nabla e^{\pm} = -\frac{\dot{R}}{R}e^{\pm} \otimes_{s} \mathrm{d}t,$$
$$R_{\nabla}e^{\pm} = -\frac{\ddot{R}}{R}\mathrm{d}t \wedge e^{\pm} \otimes \mathrm{d}t \pm \left(\frac{\dot{R}}{R}\right)^{2} R^{2}e^{+} \wedge e^{-} \otimes e^{\pm},$$

$$\operatorname{Ricci} = \frac{\ddot{R}}{R} dt \otimes dt + \frac{1}{2} \left(\frac{\dot{R}^2}{R^2} + \frac{\ddot{R}}{R} \right) R^2 e^+ \otimes_s e^-, \quad S$$

$$\operatorname{Eins} = \operatorname{Ricci} - \frac{1}{2}Sg = -\frac{1}{2}\left(\frac{\dot{R}}{R}\right)^2 \mathrm{d}t \otimes \mathrm{d}t - \frac{R\ddot{R}}{2}e^+ \otimes_s e^- \qquad \nabla \cdot \operatorname{Eins} = 0$$

$$T = f dt \otimes dt - pR^2 e^+ \otimes_s e^- \qquad p = -\frac{1}{8\pi G} \left(\frac{\ddot{R}}{R}\right), \quad f = \frac{1}{8\pi G} \left(\frac{\dot{R}}{R}\right)^2$$

$$R(t) = R_0 \left(1 + \sqrt{8\pi G f_0} (1+w)t \right)^{\frac{1}{1+w}}$$

Metric

QLC

Ricci

Curvature

$$R_{\nabla} \mathrm{d}t = \ddot{R}R\mathrm{d}t \wedge e^+ \otimes_s e^-,$$

$$= -2\frac{\ddot{R}}{R} - \left(\frac{\dot{R}}{R}\right)^2$$

Naive Einstein Tensor

Fluid tensor for a pressure p and density f

1+2 rate of expansion

$< 0 \mid T[q]$ With *R* being constantan (QFT)

Particle creation in the adiabatic limit \dot{R}/R –

$$\Delta = -\partial_t^2 - 2\frac{\dot{R}}{R}\partial_t + \frac{2}{R^2}(\partial_+ + \partial_-) \qquad (-\Delta + n)$$

No particle creation

 $\mathbb{R} \times \mathbb{Z}_n$

$$\mathbb{R} \times S^{1}$$

$$R \propto t^{2}$$

$$\frac{1}{2}\frac{\ddot{R}}{R} = \frac{1}{4}\left(\frac{\dot{R}}{R}\right)^{2} \quad (m \to \infty)$$

$$\ddot{R} = 0 \ (m \to \infty)$$

$$\frac{\ddot{R}}{R} = -\frac{1}{2} \left(\frac{\dot{R}}{R} \right)$$
$$R \propto t^{\frac{2}{3}}$$

$$\phi_i(t_a)\phi_j(t_b)] \mid 0 > = \sum_{k=0}^{n-1} \frac{1}{w_k} \cos\left(\frac{2\pi}{n}k(i-j)\right) e^{-iw_k|t_a-t_b|}$$

$$\rightarrow 0, \ \ddot{R}/R \rightarrow 0$$
 $w_k(t) = \sqrt{m^2 + \frac{8}{R^2(t)}} \sin^2\left(\frac{\pi}{n}k\right)$

 $m^2)\phi = 0$





Thank you for your attention