Flavonstrahlung at Current and Future Colliders

Eetu Loisa

DAMTP, University of Cambridge

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Based on arXiv:2212.07440 with Ben Allanach 💿 🔊

Background

Neutral current B-anomalies: disagreement between SM and experiment in processes involving $b \to s \ell \ell$ decays

Example:



Tree-level explanations



Another key motivation for flavour non-universal Z's: fermion mass puzzle¹

¹Allanach and Davighi, 1809.01158

 $B_3 - L_2 \text{ model}^2$

Gauge group:

$$\mathcal{G} = SU(3) \times SU(2) \times U(1) \times U(1)_{B_3-L_2}$$

Field content:

$$SM + Z' + \theta$$
 (SM singlet scalar) + $3\nu_R$

Spontaneous symmetry breaking:

$$\langle \theta \rangle = rac{v_{\theta}}{\sqrt{2}} \sim \mathcal{O}(\text{TeV}) \Rightarrow Z' \text{ becomes massive}$$

²Alonso et al., 1705.03858; Bonilla et al, 1705.00915; Allanach, 2009.02197

Fermion sector

$$\mathcal{L}_{Z'\psi} = -g_{Z'} \Big(\overline{Q'_{3L}} \mathbf{Z}' Q'_{3L} + \overline{u'_{3R}} \mathbf{Z}' u'_{3R} + \overline{d'_{3R}} \mathbf{Z}' d'_{3R} \\ - 3\overline{L'_{2L}} \mathbf{Z}' L'_{2L} - 3\overline{e'_{2R}} \mathbf{Z}' e'_{2R} - 3\overline{\nu'_{2R}} \mathbf{Z}' \nu'_{2R} \Big)$$

How to connect to $b \rightarrow s\mu^+\mu^-$? We need to specify the fermion mixing matrices

$$\mathbf{P}' = V_I \mathbf{P}$$

for $I \in \{u_L, d_L, e_L, \nu_L, u_R, d_R, e_R, \nu_R\}$.

Simple mixing ansatz

Use simplicity, ease of passing bounds and ability to explain B-anomalies as a guiding principle:

$$V_{d_L} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{sb} & -\sin \theta_{sb} \\ 0 & \sin \theta_{sb} & \cos \theta_{sb} \end{pmatrix},$$

$$V_{d_R} = 1, V_{e_R} = 1, V_{e_L} = 1$$
 and $V_{u_R} = 1$. These imply $V_{u_L} = V_{d_L} V_{\mathsf{CKM}}^{\dagger}$ and $V_{\nu_L} = U_{\mathsf{PMNS}}^{\dagger}$

Now, in the mass eigenbasis:

$$\mathcal{L}_{Z'\psi} \supset -g_{Z'} \left[\left(\frac{1}{2} \sin 2\theta_{sb} \overline{s} \overline{Z}' P_L b + \text{H.c.} \right) - 3\overline{\mu} \overline{Z}' \mu \right]$$

$\begin{aligned} & \mathsf{Scalar potential} \\ V(H,\theta) &= -\mu_H^2 H^{\dagger} H + \lambda_H (H^{\dagger} H)^2 - \mu_{\theta}^2 \theta^* \theta \\ &\quad + \lambda_{\theta} (\theta^* \theta)^2 + \lambda_{\theta H} \theta^* \theta H^{\dagger} H. \end{aligned}$

After symmetry breaking:

$$H = igg(egin{array}{c} 0 \ rac{m{v}_{H}+m{h}'}{\sqrt{2}} igg), \qquad heta = rac{m{v}_{ heta}+artheta'}{\sqrt{2}} \end{array}$$

 $V(H, \theta) \supset -\lambda_{\theta H} v_{\theta} v_H h' \vartheta' \Rightarrow$ non-diagonal mass matrix

Rotate into mass eigenbasis:

$$\begin{pmatrix} h \\ \vartheta \end{pmatrix} = \begin{pmatrix} \cos \phi & -\sin \phi \\ \sin \phi & \cos \phi \end{pmatrix} \begin{pmatrix} h' \\ \vartheta' \end{pmatrix} \stackrel{\text{(h)}}{\underset{7/16}{}}$$

Constraints on Higgs-flavon mixing

Higgs signal strength: $h' = \cos \phi h + \sin \phi \vartheta$ \Rightarrow SM Higgs interactions scaled by $\cos \phi$

Direct flavon searches: Can use null results from SM Higgs searches at colliders to rule out light flavons

More constraints

Perturbativity: Impose $|\lambda_i| < 4\pi$

W boson mass: Take M_Z , G_F and α as experimental inputs. Obtain a (recursive) prediction for M_W :

$$M_{W}^{2} = \frac{1}{2}M_{Z}^{2}\left[1 + \sqrt{1 - \frac{4\pi\alpha}{\sqrt{2}G_{F}M_{Z}^{2}}[1 + \Delta r(M_{W}^{2})]}\right]$$

Putting it all together



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How to produce the flavon?

The flavonstrahlung process:



Upon further decay, $Z' \rightarrow \mu^- \mu^+$ and $\vartheta \rightarrow hh$, get resonances at $M_{Z'}$ and m_{ϑ}

Both hadron and muon colliders of the future should have good sensitivity

Collider simulations

Choose benchmark points in the $M_{Z'} - g_{Z'}$ plane, compute flavonstrahlung cross-sections as a function of flavon mass



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Flavonstrahlung at FCC-hh



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Flavonstrahlung at muon colliders



Summary

Family non-universal Z's, such as $B_3 - L_2$, well-motivated by $b \rightarrow s \mu^+ \mu^-$ and fermion mass puzzle

We have studied the phenomenology of the scalar field θ and the flavonstrahlung process

Unlikely to be observed at the HL-LHC, but a 100 TeV FCC-hh or a 10 TeV muon collider would have excellent discovery prospects

The End

Thank you for listening!