First Demonstration of Event-by-Event Directional Reconstruction in the SNO+ Scintillator Phase

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Motivation

- Directional information can be a very powerful tool for background rejection in neutrino experiments
  - This information is often obtained through Cherenkov rings

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- Scintillator experiments benefit from high light yield
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  - Scintillation light is isotropic – no directional information
- If Cherenkov light can be isolated from the scintillation signal, directional reconstruction in scintillator could be possible
  - This can be done using timing as scintillation is non-instantaneous
  - Slow scintillators have shown promise on a bench-top scale (Nucl. Instrum. Meth. A 972, 164106)

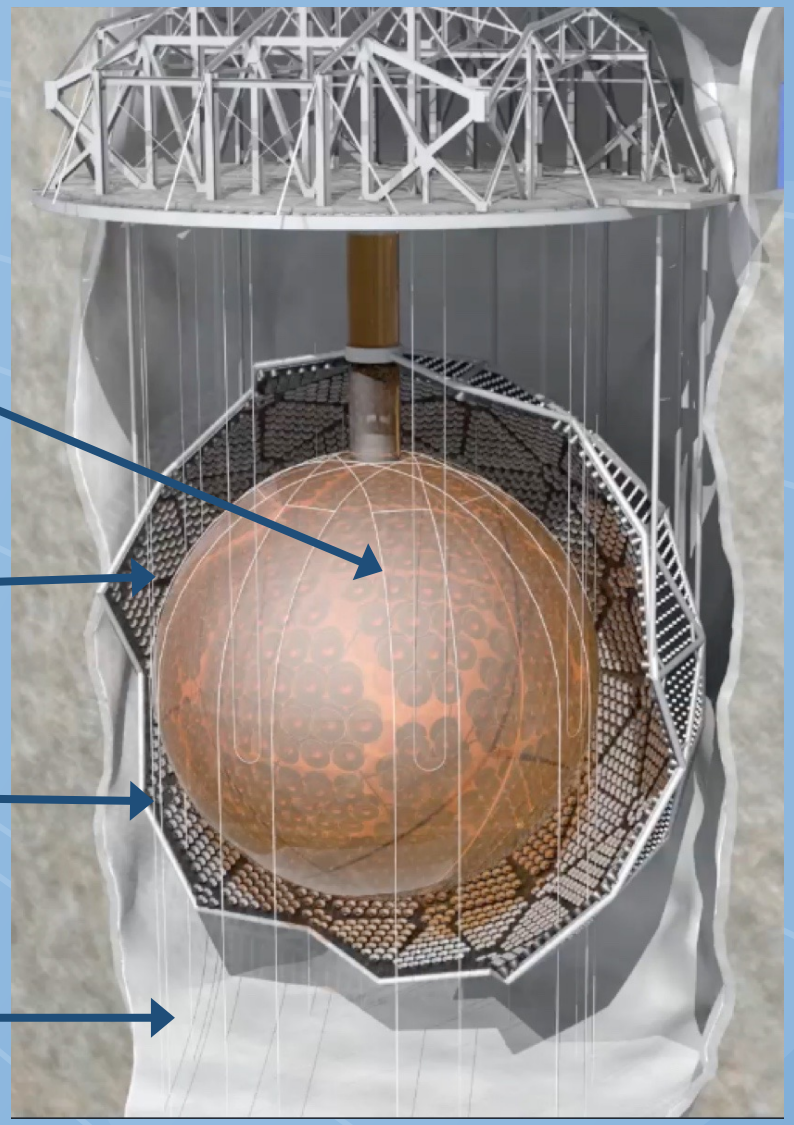
The SNO+ Detector

12 m diameter Acrylic Vessel (AV)
Partial Fill Phase: 365 t Scintillator
Interface ~75 cm above the equator
Lower concentration of primary fluor

9,300 inward facing
PMTs for ~50% effective
coverage

PMT Support Structure
(PSUP)

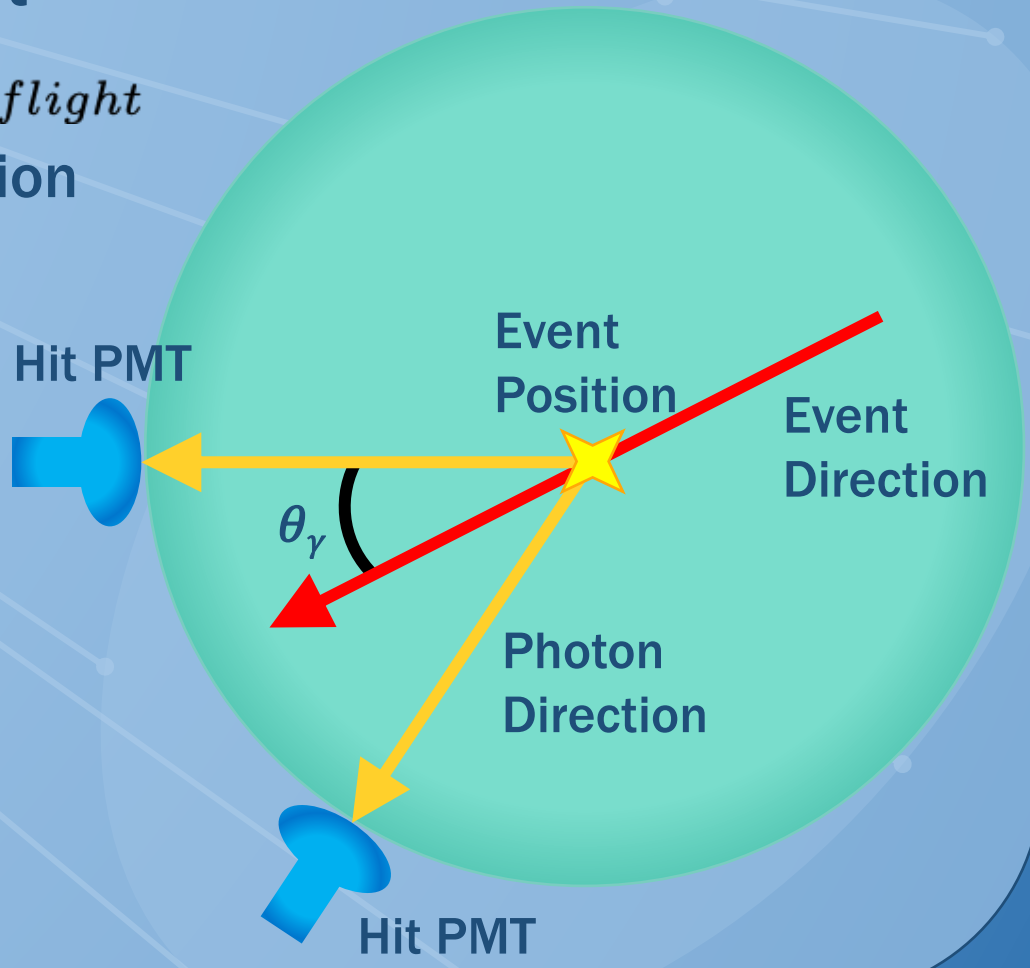
Water shielding
1700 T inside PSUP
5300 T outside PSUP



The SNO+ Experiment
2021 JINST 16 P08059

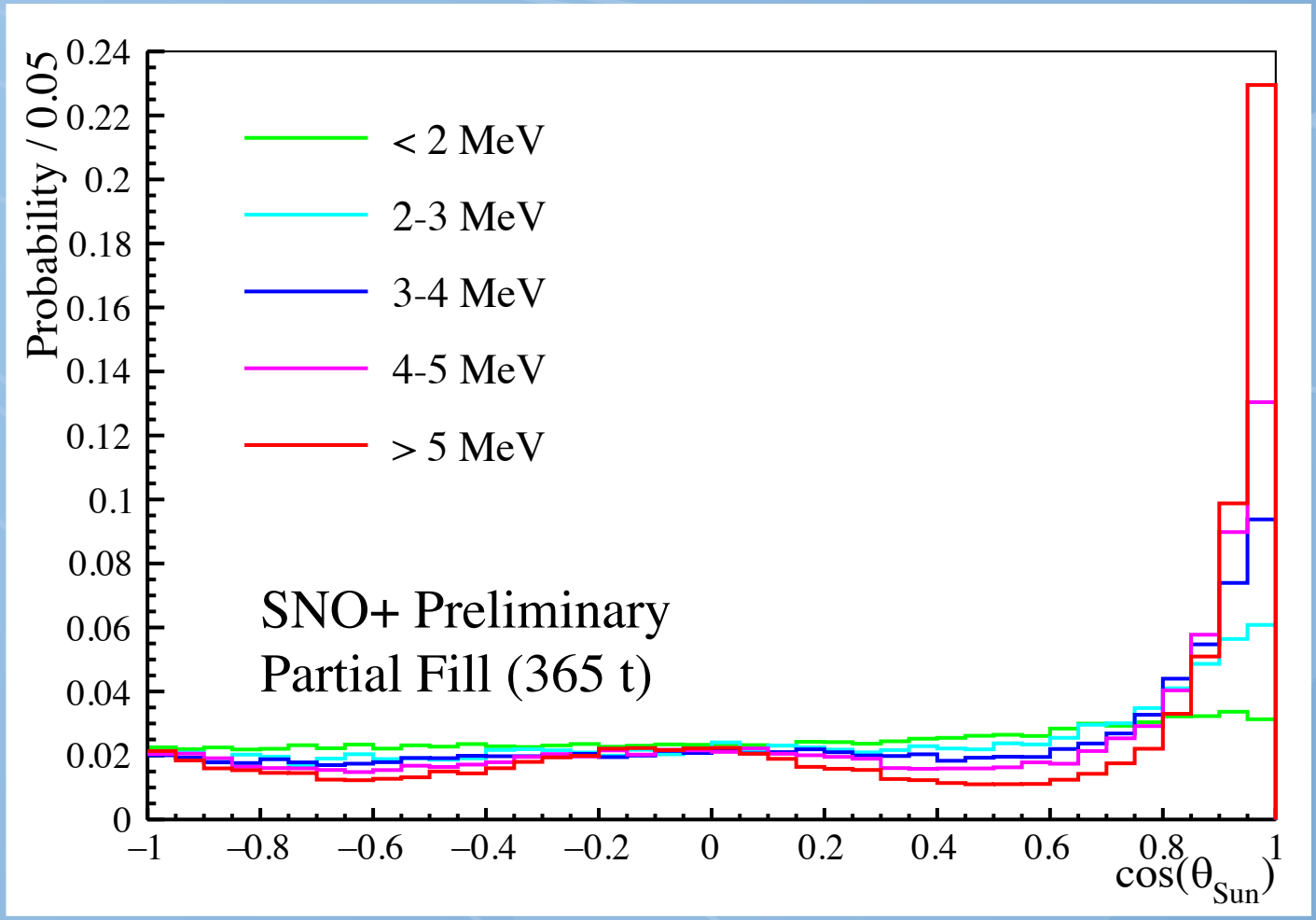
Cherenkov Peak Isolation

- Two ways to separate Cherenkov light
- Time Residual: $t_{res} = t_{hit} - t_{event} - t_{flight}$
  - Cherenkov will be earlier than scintillation
- Photon Angle:
  - Cherenkov will have a peak in θ_γ
- Cherenkov peak can be isolated using these parameters
- Can use this information to reconstruct the electron direction

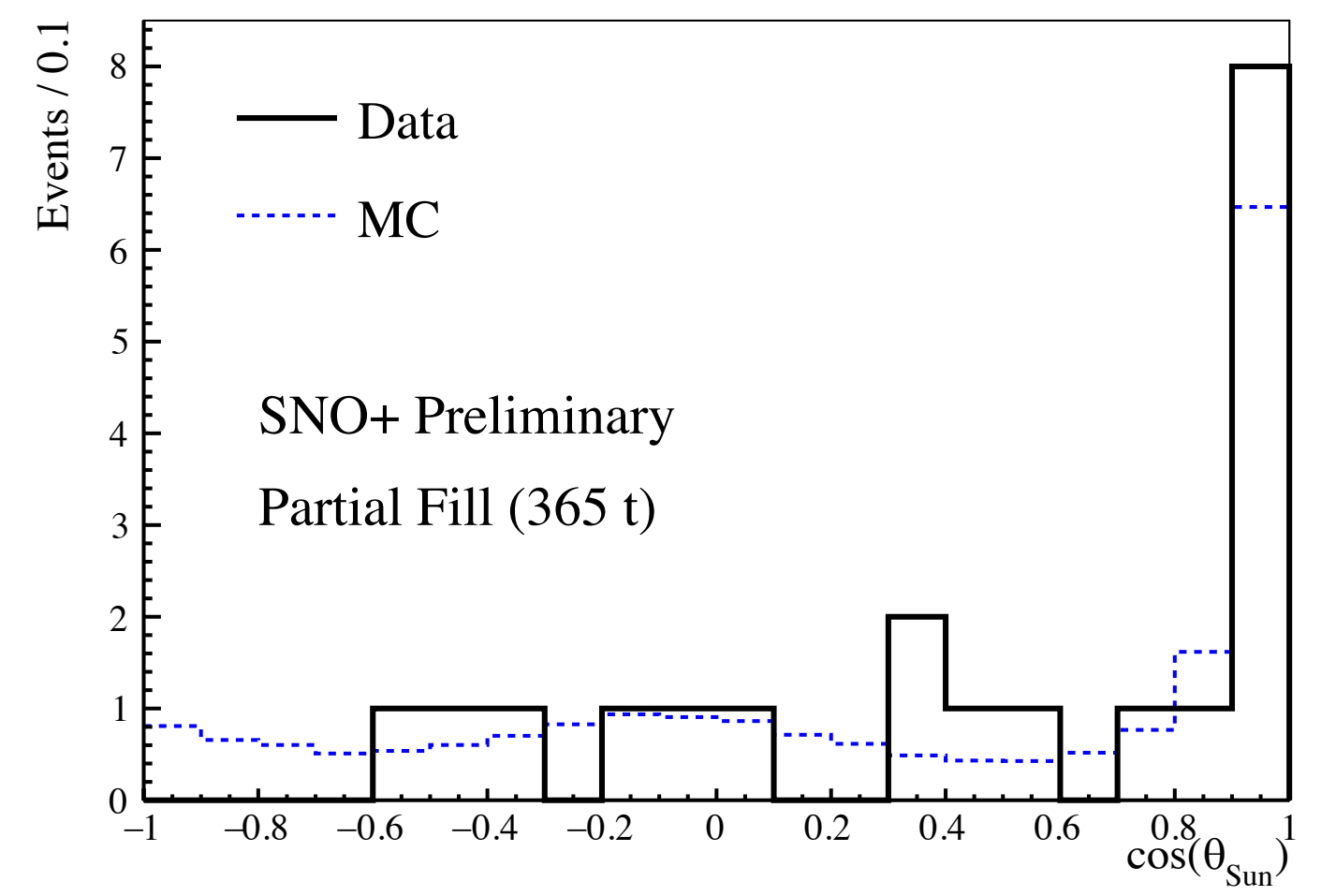


Directional Reconstruction - MC

- ^{8}B solar ν_e MC
  - Run conditions of the Partial Fill detector
  - Cuts: z > 1 m above equator, r < 5.5 m from centre of AV
- Reconstruction with a likelihood fit over time and angle distribution of photons
- $\cos(\theta_{Sun})$ - angle between reconstructed and solar direction



Data Results



- Data taken from the Partial Fill Phase
  - Livetime of 92 days
  - PMT Hit cut of >1500 used
    - ~ 5 MeV
  - 20 events extracted
  - Equivalent MC plot included
- 20 ^{8}B events extracted
- First event-by-event directional reconstruction a high-yield large-scale liquid scintillator experiment!

Summary

- **First demonstration of event-by-event* direction reconstruction in a high-yield** liquid scintillator detector!**
  - Direction has been reconstructed for Solar events > 5 MeV in the SNO+ Partial Fill Phase
- Even more improvements to come!
  - Position-reconstruction effects
  - Scintillator cocktail tuning
  - And more!

- *Statistical separation of solar directionality has been shown by Borexino using ~20,000 events [1]
- ** LSND[2] MiniBooNE[3] have previously used directional reconstruction at higher energy scales using lower yield scintillators.

Thank you for listening

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References

1. Borexino Collaboration. Correlated and Integrated Directionality for sub-MeV solar neutrinos in Borexino. arXiv:2109.04770 [hep-ex], 2021.
2. C Athanassopoulos et al. The liquid scintillator neutrino detector and LAMPF neutrino source. Nucl. Instrum. Meth. A, 388(1):149–172, 1997.
3. R.B. Patterson and E.M. Laird and Y. Liu and P.D. Meyers and I. Stancu and H.A. Tanaka. The extended-track event reconstruction for MiniBooNE. Nucl. Instrum. Meth. A, 608(1):206–224, 2009.

Back-up Slides

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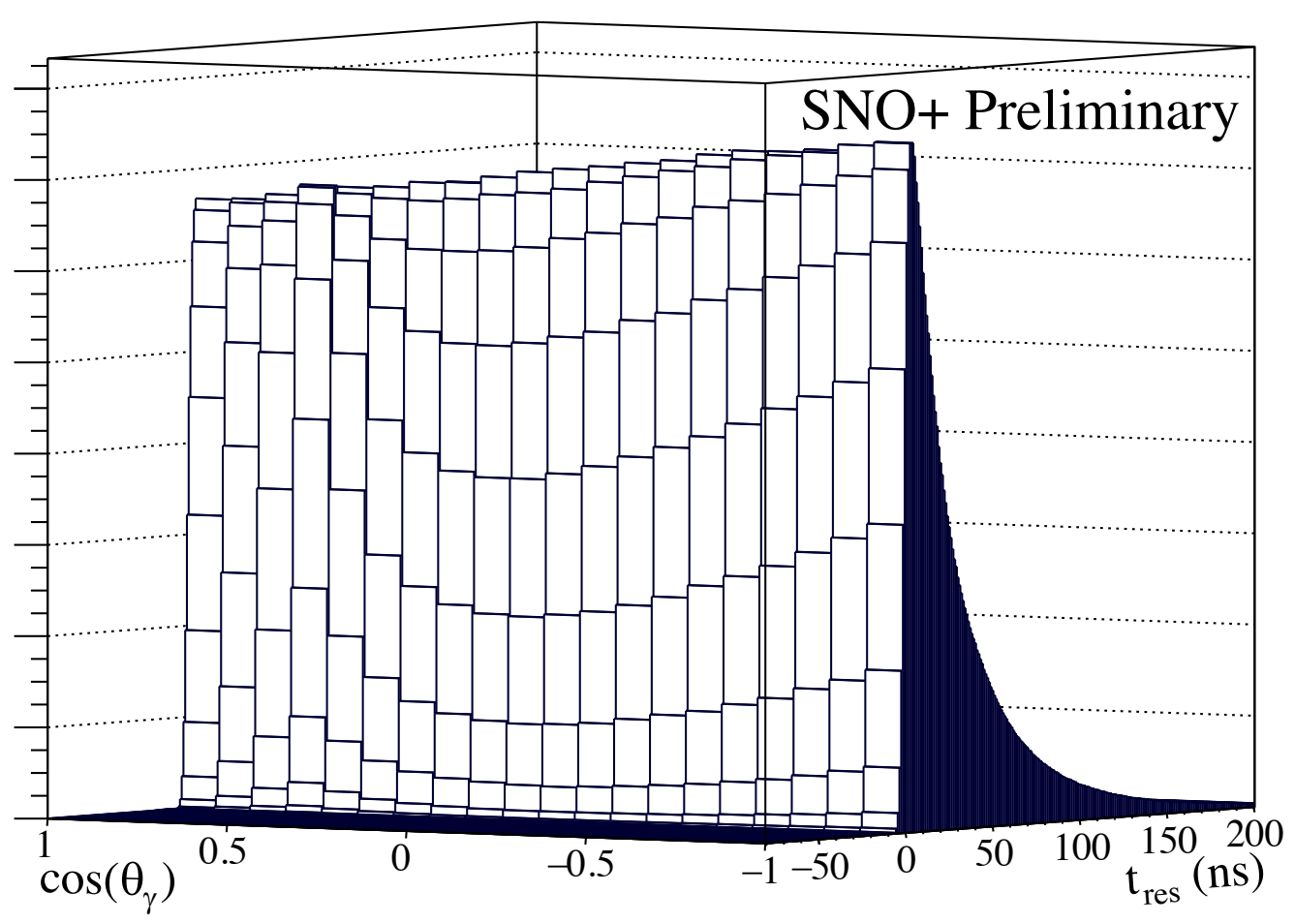
Direction Reconstruction PDF



- 6 MeV electrons simulated
  - Perfect-state full-fill detector
  - Full volume, isotropic directions
  - 0.6 g/L PPO scintillator
- Clear Cherenkov peak seen
- “Backwards” peak also seen
  - Caused by bias in positional reconstruction
- Can be used as a PDF for a directional likelihood fitter