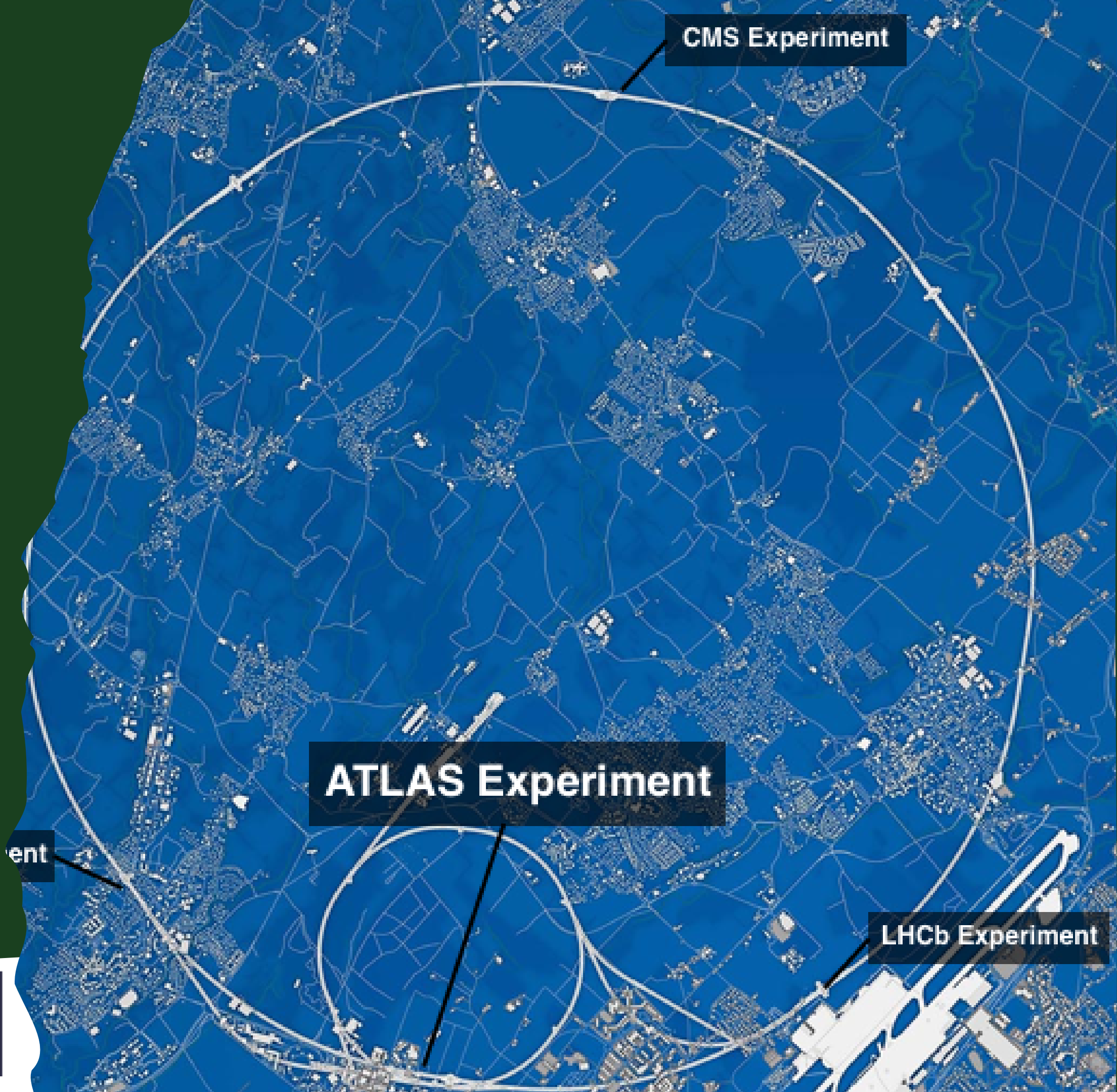


Features of new physics in LLP searches

at LHC-based LLP detectors

for the ANUBIS Experiment

Anna Mullin
University of Cambridge



FIPs & LLP models

Hidden Valley / Sector / FIP revolution reshaped the theoretical and experimental particle physics landscape

- Shift away from **focus only on states at ever higher energies**

- Hidden sectors appear generically in many **UV complete theories** (e.g. strings)

- Heavy mediators can **mediate rare decays** / flavour violating processes

Generic physics features:

- Light DS **hidden by high barrier** to interactions with SM
- Explains DM using some symmetry to **stabilise at least one hidden sector state**
- Results in **neutral** bound states with **low masses and long lifetimes**
- DS states decay to SM through **heavy mediators**
- Mechanisms to set the **DM relic abundance**

Classes of models:

- **Common benchmark models** for comparison across experiments
- Simplicity: **minimal vs. non-minimal** models

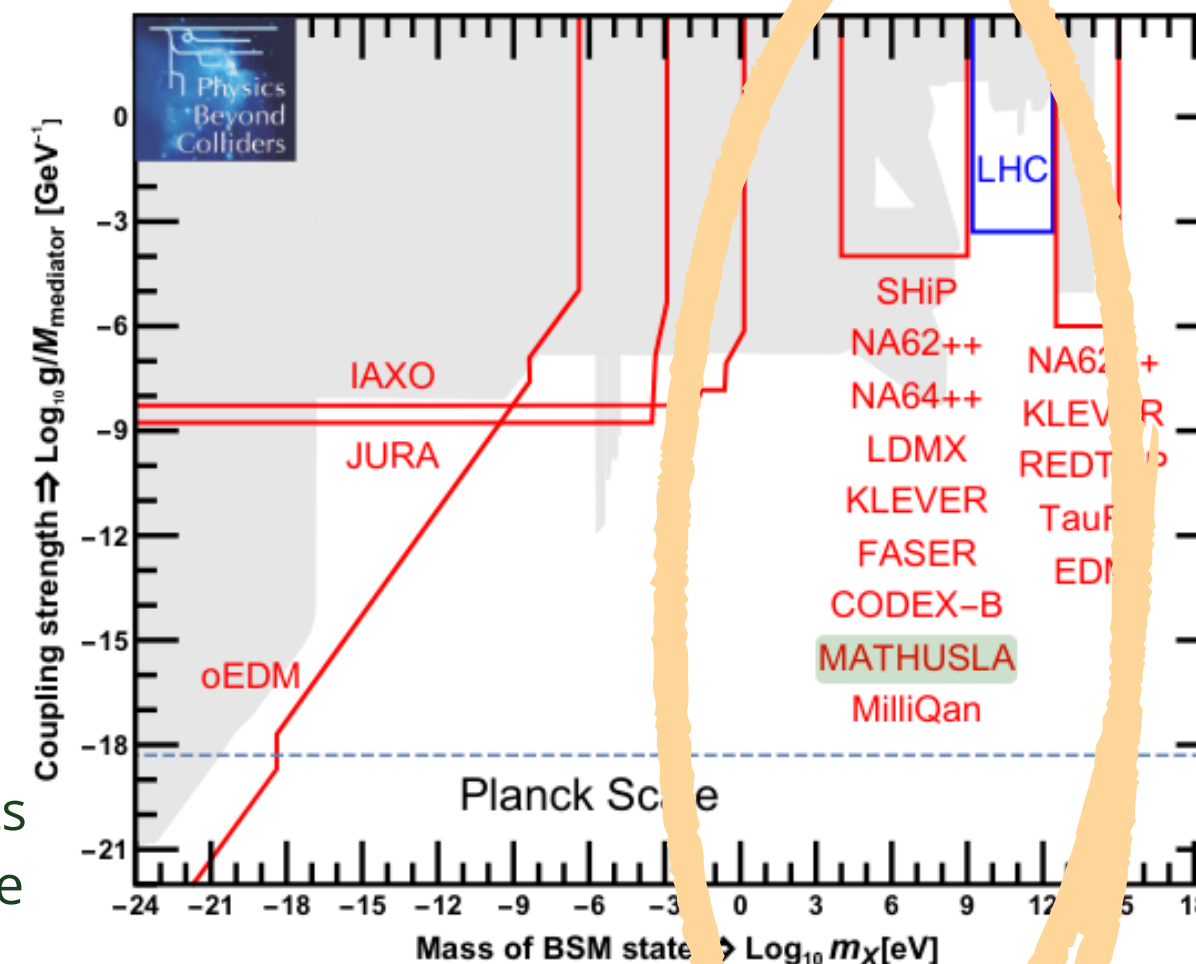
Flexible structure of hidden sector

wide range of models. How should we organise our search at ANUBIS?

Experiments & search strategies

Strategies for covering remaining LLP parameter space
Target mass & lifetime of LLP

- Filling gaps in parameter space
 - E.g. lifetimes longer than 10^8 seconds are less constrained by LHC experiments
 - Lifetimes shorter than a few minutes are less constrained by BBN
 - Sensitivity in MeV-GeV range



Transverse vs forward LHC LLP detectors

- Forward:
 - Light, weakly interacting particles produced dominantly along beam axis
- Transverse:
 - Heavier / more strongly interacting LLPs

Concentrating on ANUBIS:

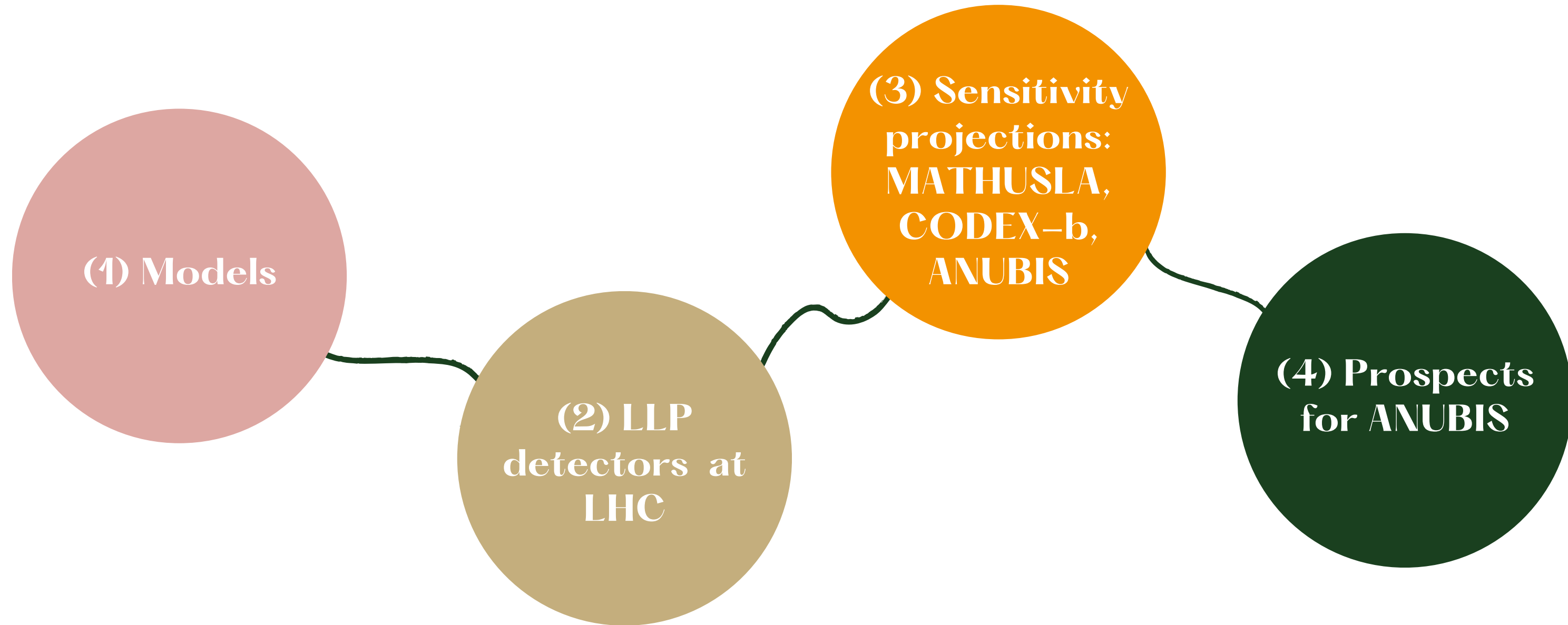
- Focus on scenarios predicting LLPs boosted in transverse direction
- Signatures from production of unstable portal/DS particle with displaced decay to SM
- Signatures from extended DS with additional Higgs/Z boson couplings



CODEX-b

MATHUSLA

Contents



Models – generic features

Aims:

- Organise our search for specific models
 - by comparing sensitivity projections with other LLP detectors
 - and identifying interesting model parameter space

Link to dark matter:

- If **thermal / freeze-out** → small set of highly predictive portal models
(direct thermal contact with SM)
- If **non-thermal / freeze-in** → many complicated & less constrained models
(DM never reached equilibrium with thermal bath)

Focus on predictive thermal DM models

PBC benchmark models:

- Motivated by Dark Sector / Hidden Valleys
- Categorising the dimension-4 and -5 portal interactions allowing a total SM singlet to interact with SM fields
- 11 benchmarks

Generic search strategies:

- **Direct freeze-out:** DM freezes out directly to SM final states
 - Mediator decays into DM which is lightest state
 - **Mainly invisible decays**
 - Missing energy, sharply-characterised production targets
- **Secluded freeze-out:** DM freezes out into mediator particles
 - Mediator/at least one DS particle is lighter than DM
 - **Mediator decays visibly**

The search for thermal relic DM so far:

- GeV-TeV scale: null results in direct/indirect/LHC DM experiments
- MeV-GeV scale: thermal relic DM is much less constrained

Hidden Valleys – PBC

General classification for LLP scenarios

- Hidden sector states with mass scales below the weak scale are hidden from the visible sector due to a high barrier (i.e. sub-weak interactions between sectors)
- Flexible structure: strongly or weakly coupled, dark photons or Higgses, a dark SU(N) with N_f flavours, sterile neutrinos or asymmetric dark matter, wide range of dynamics

Lowest-dimensional renormalisable portals

Portal	Coupling
Dark Photon, A_μ	$-\frac{\epsilon}{2\cos\theta_W} F'_{\mu\nu} B^{\mu\nu}$
Dark Higgs, S	$(\mu S + \lambda S^2) H^\dagger H$
Axion, a	$\frac{a}{f_a} F_{\mu\nu} \tilde{F}^{\mu\nu}, \frac{a}{f_a} G_{i,\mu\nu} \tilde{G}_i^{\mu\nu}, \frac{\partial_\mu a}{f_a} \bar{\psi} \gamma^\mu \gamma^5 \psi$
Sterile Neutrino, N	$y_N L H N$

Vector portal (dark photon/ Z)

Scalar portal (dark Higgs)

Axion portal (ALPs)

Neutrino portal (HNLs)

PBC benchmarks

- Motivated by simplicity
 - E.g. 3 cases identified for HNL benchmarks, all with only 1 HNL state and single-flavour couplings
- BC1-BC11 describe cases for all 4 above portals

Dark scalar & dark photon/Z'

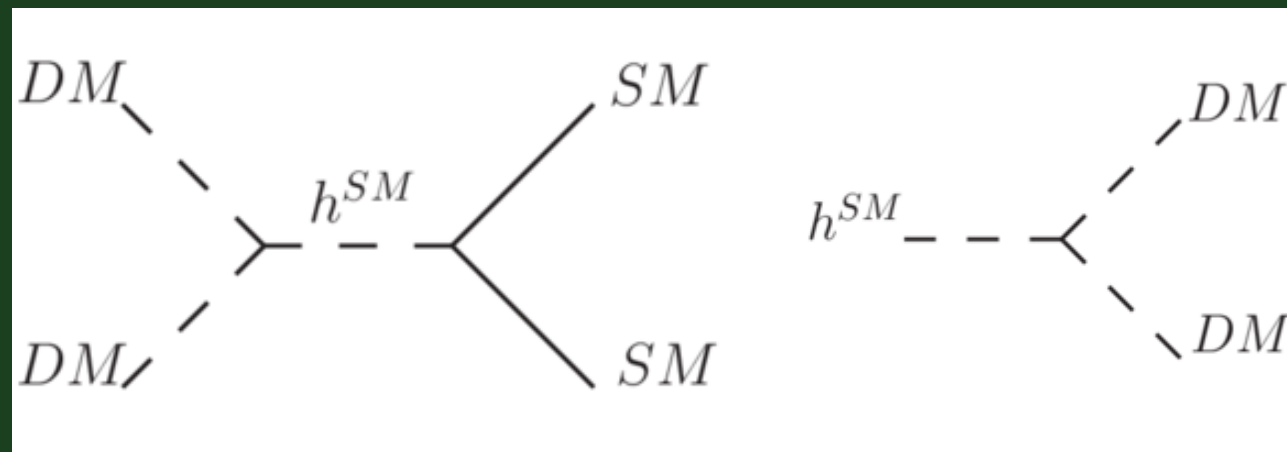
- Renormalizable vector portals mediated by a dark vector boson; scalar portals mediated by a new scalar mixing with the SM Higgs boson
- New vector provides simplest possible thermal freeze-out scenario producing correct relic abundance with MeV-GeV DM masses

Dark scalar/Higgs

$$\mathcal{L}_{\text{scalar}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DS}} - (\mu S + \lambda S^2) H^\dagger H.$$

$$\mathcal{L}_{\text{DS}} = S \bar{\chi} \chi + \dots$$

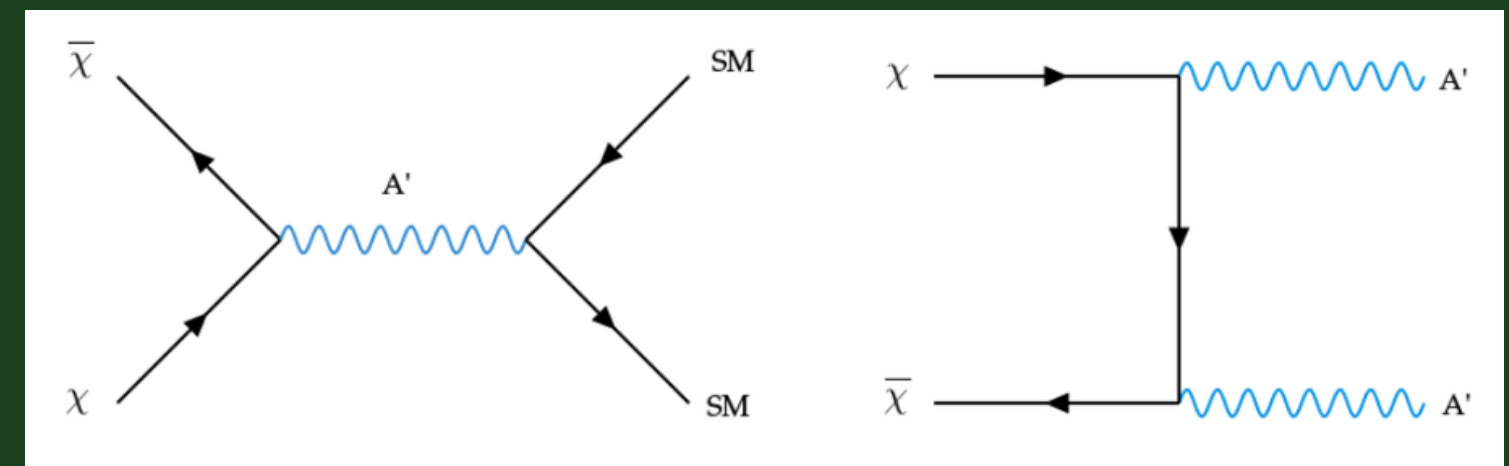
- E.g. if scalar is Higgs boson



Dark vector/photon

$$\mathcal{L}_{\text{vector}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DS}} - \frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} B_{\mu\nu}$$

$$\mathcal{L}_{\text{DS}} = -\frac{1}{4} (F'_{\mu\nu})^2 + \frac{1}{2} m_{A'}^2 (A'_\mu)^2 + |(\partial_\mu + ig_D A'_\mu) \chi|^2 + \dots$$

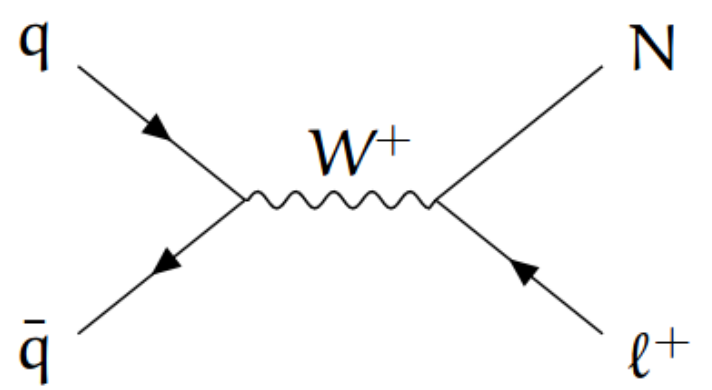


Heavy Neutral Leptons

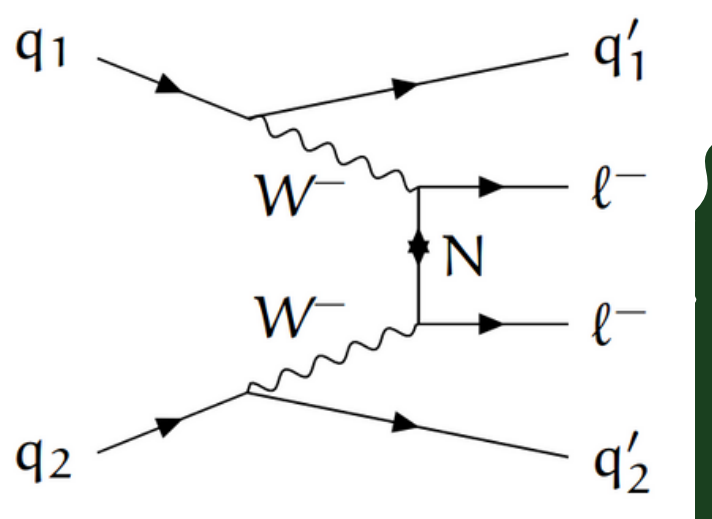
Minimal & non-minimal HNL scenarios

- Explain nonzero neutrino masses via the seesaw mechanism
- Mechanism for generating matter-antimatter asymmetry
- Object of ATLAS and CMS searches covering mass range from a few GeV to several TeV
- Take into account different production channels: **D-mesons, B-mesons, W-bosons, Z-bosons, SM Higgs bosons, and the top quark**
- For meson decays, include three-body and two-body decays

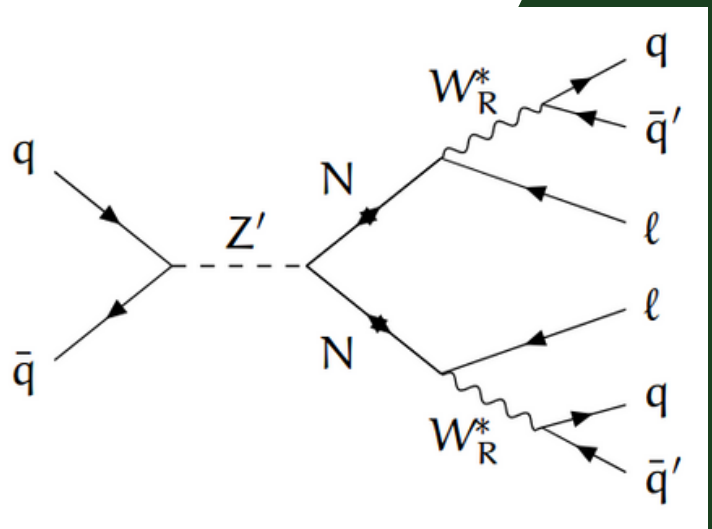
Drell-Yann process in type-I seesaw models



VBF t-channel process in type-I seesaw models



Z' decays



E.g. gauge group with simple SM extension

RH sector neutrino mixing matrix

CC and NC interactions relevant to later analysis

Interesting scenario:

- Targeting scenarios with large mass gap between HNL and Z' such that HNLs and decay products are highly boosted

$$SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$\mathcal{L} = \frac{g_R}{\sqrt{2}} (\bar{d}\gamma^\mu P_R u + V_{\alpha N}^R \cdot \bar{l}_\alpha \gamma^\mu P_R N) W_{R\mu}^-$$

$$+ \frac{g_R}{\sqrt{1 - \tan^2 \theta_W (g_L/g_R)^2}}$$

$$\times Z_{LR}^\mu \bar{f} \gamma_\mu [T_{3R} + \tan^2 \theta_W (g_L/g_R)^2 (T_{3L} - Q)] f$$

Non-minimal LLP models

Rich structure and phenomenology

Inelastic DM

- DM + mediator + **additional particles**
- Best probed at high intensity experiments

The only relevant interaction is inelastic:

$$\mathcal{L} \supset \frac{ie_D m_D}{\sqrt{m_D^2 + (\delta_\xi - \delta_\eta)^2/4}} A'_\mu (\bar{\chi}_1 \gamma^\mu \chi_2 - \bar{\chi}_2 \gamma^\mu \chi_1)$$

The elastic piece is very small ($\delta_{\eta,\xi} \ll m_D$):

$$\mathcal{L} \supset \frac{e_D (\delta_\xi - \delta_\eta)}{\sqrt{4m_D^2 + (\delta_\xi - \delta_\eta)^2}} A'_\mu (\bar{\chi}_2 \gamma^\mu \chi_2 - \bar{\chi}_1 \gamma^\mu \chi_1)$$

SIMPs

- Standard thermal freeze-out via 2 --> 2 annihilation
 - Predicts a TeV scale mass range

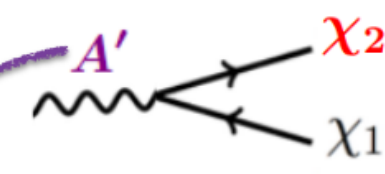
$$m_{\text{DM}} \sim \alpha_{\text{ann}} (T_{\text{eq}} M_{\text{Pl}})^{1/2} \sim \text{TeV}$$

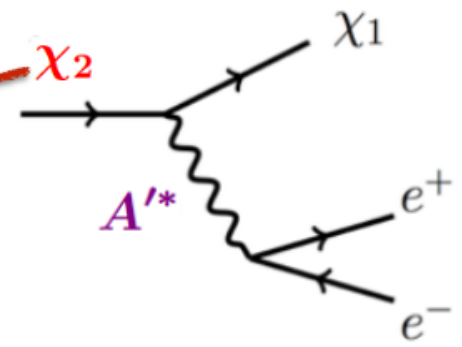
matter-radiation equality temperature

- In **SIMP models**, freeze-out via 3 --> 2 annihilation
 - Points to sub-GeV scale DM

$$m_{\text{DM}} \sim \alpha_{\text{eff}} (T_{\text{eq}}^2 M_{\text{Pl}})^{1/3} \sim 100 \text{ MeV}$$

effective strength of SIMP self-interaction (order 1)


$m_X < m_{A'}$

 Copiously produced at high intensity experiments

with


$$\Gamma(\chi_2 \rightarrow \chi_1 e^+ e^-) \simeq \frac{4\epsilon^2 \alpha_{\text{em}} \alpha_D \Delta^5 m_1^5}{15\pi m_{A'}^4}$$

E.g: **BC_iDM**

Complex DM of BC2 split into 2 non-degenerate real states

 [1402.5143.pdf](#)

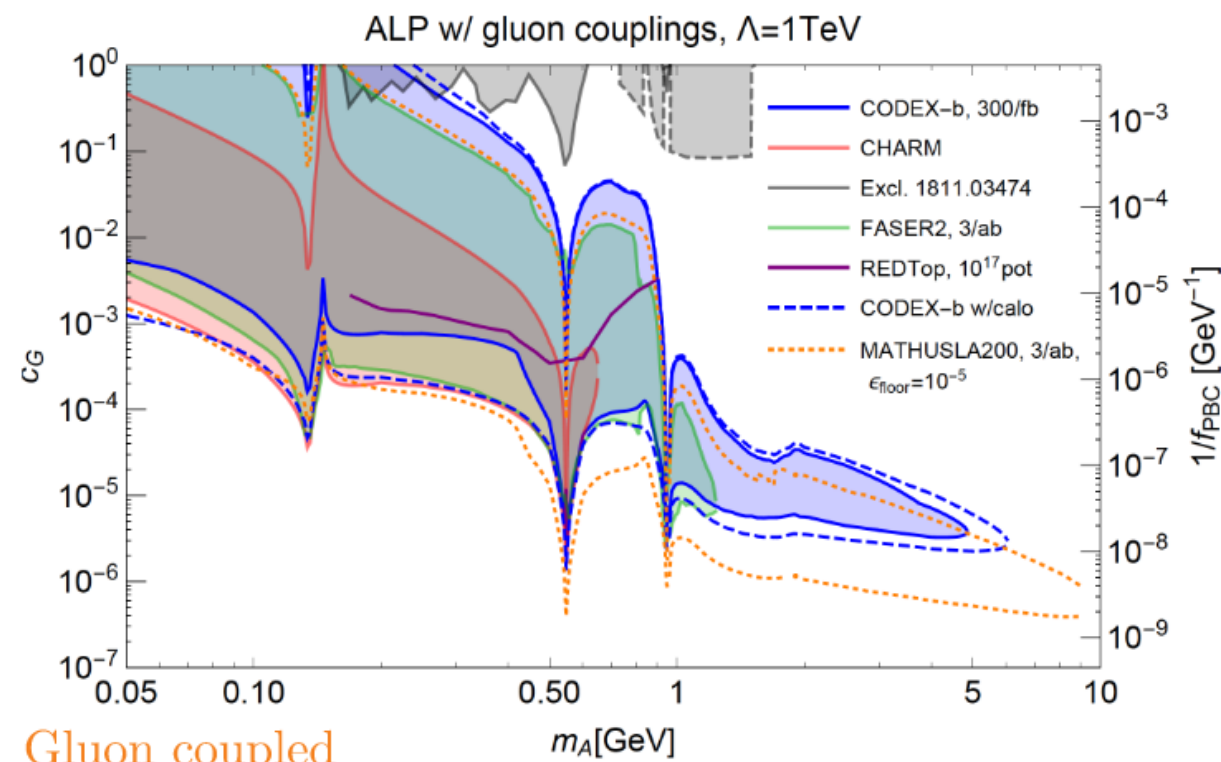
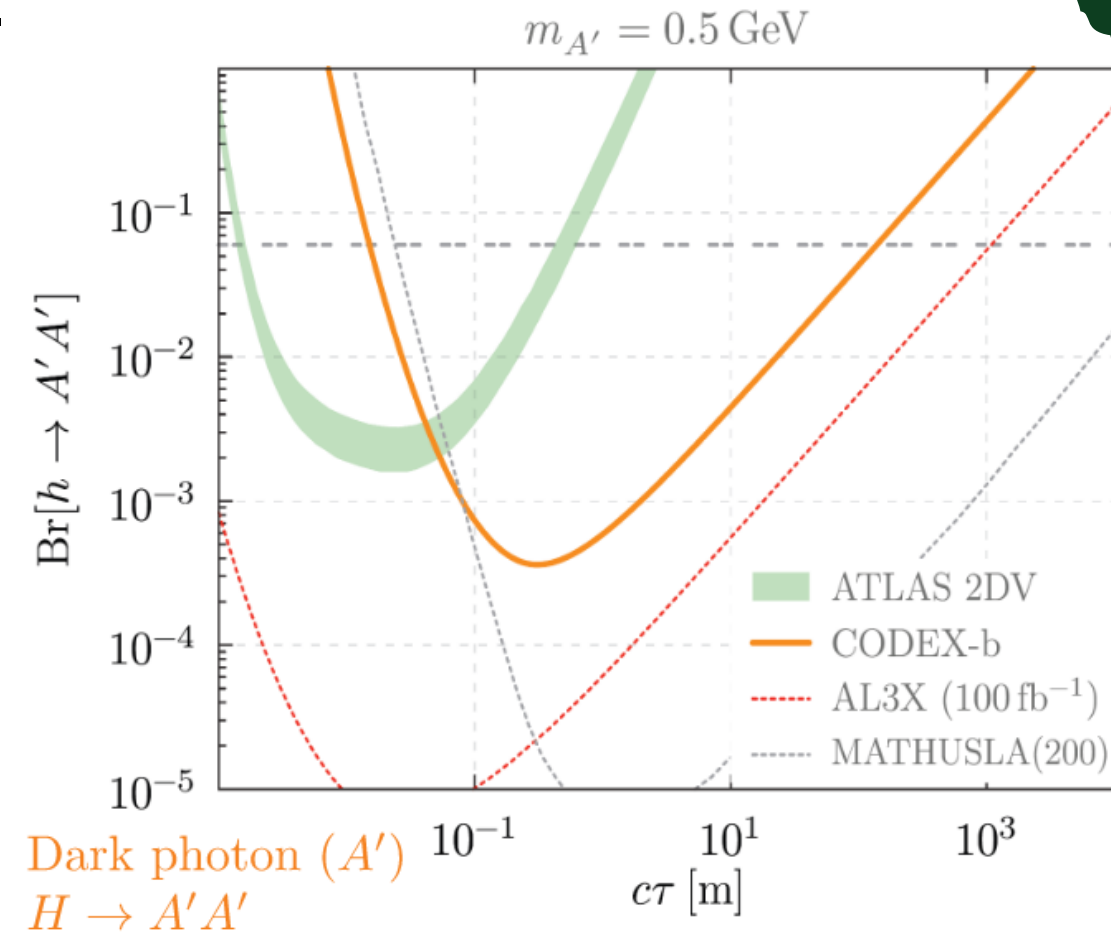
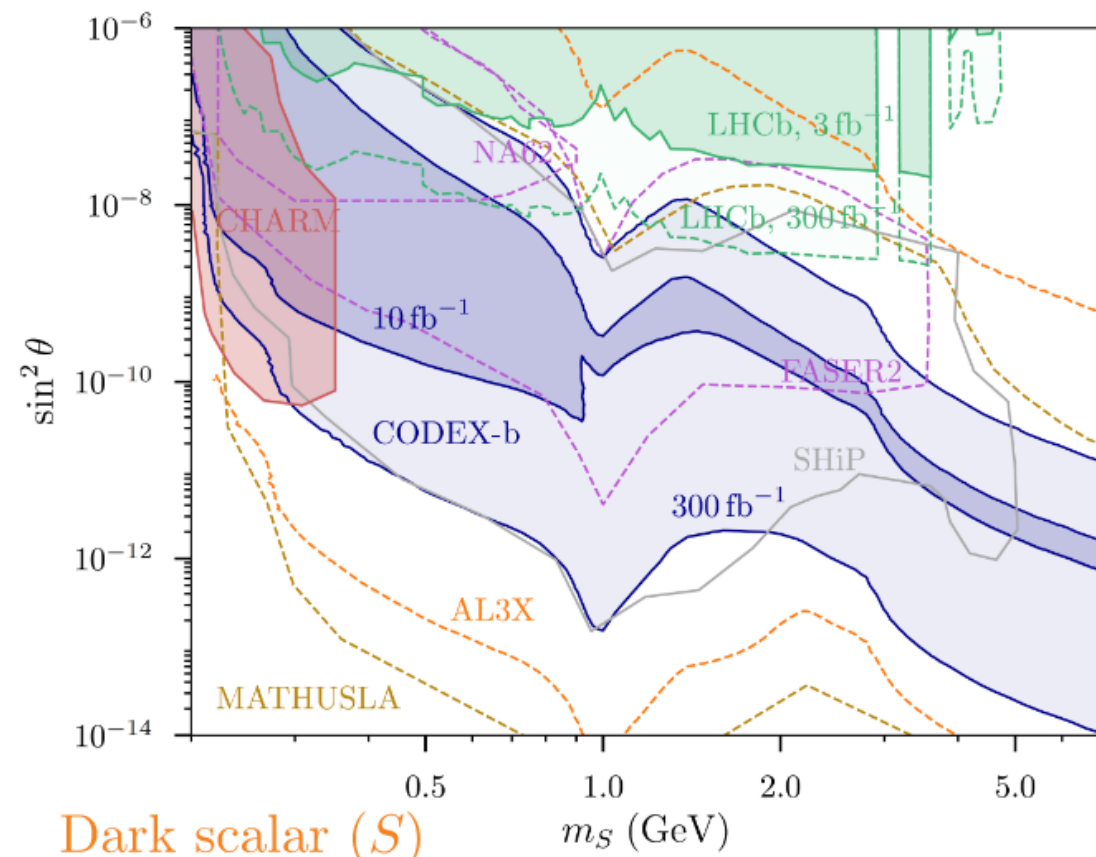
LHC transverse LLP detectors

	MATHUSLA Massive Timing Hodoscope for Ultra-Stable neutral particles	CODEX-b A Compact Detector for Exotics at LHCb
Experimental setup	<ul style="list-style-type: none">• large decay volume 200 m x 200 m x 20 m near CMS• RPCs as sensor technology	<ul style="list-style-type: none">• smaller acceptance and luminosity than MATHUSLA but closer proximity to IP so overall competitive in low-lifetime regime• 10 m x 10 m x 10 m tracker box
Combined analysis	<ul style="list-style-type: none">• supply an L1 trigger to ensure retention of relevant info at CMS• allows to determine production mode, decay mode, parent particle mass (if applicable)	<ul style="list-style-type: none">• proximity to LHCb permits interface with LHCb's planned triggerless readout, allowing identification and partial reconstruction of LLP event
Shielding / backgrounds	<ul style="list-style-type: none">• located above ground but partly excavated, 100m horizontally and vertically away from IP• relying on shielding to deliver small/zero backgrounds	<ul style="list-style-type: none">• behind a 3.2 m thick shield• largest gains in reach for relatively light LLPs (<10 GeV)<ul style="list-style-type: none">◦ where backgrounds in ATLAS and CMS are prohibitive

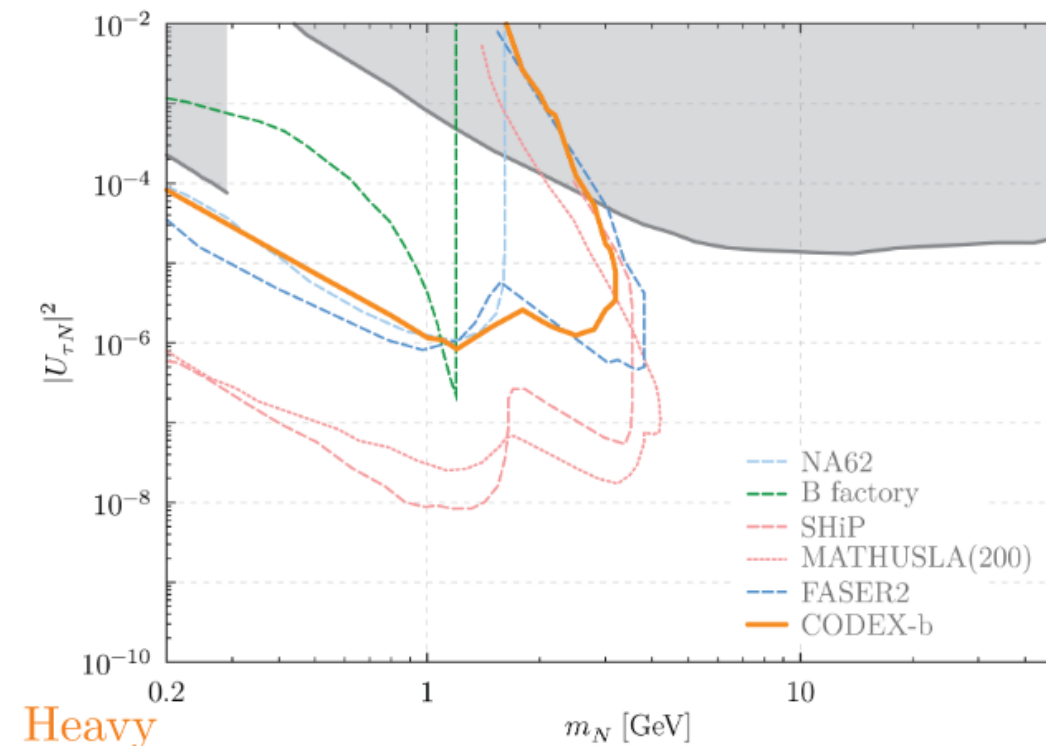
Portal models – projections

MATHUSLA, CODEX-b

[2251269/3819001/CODEX-b](#)



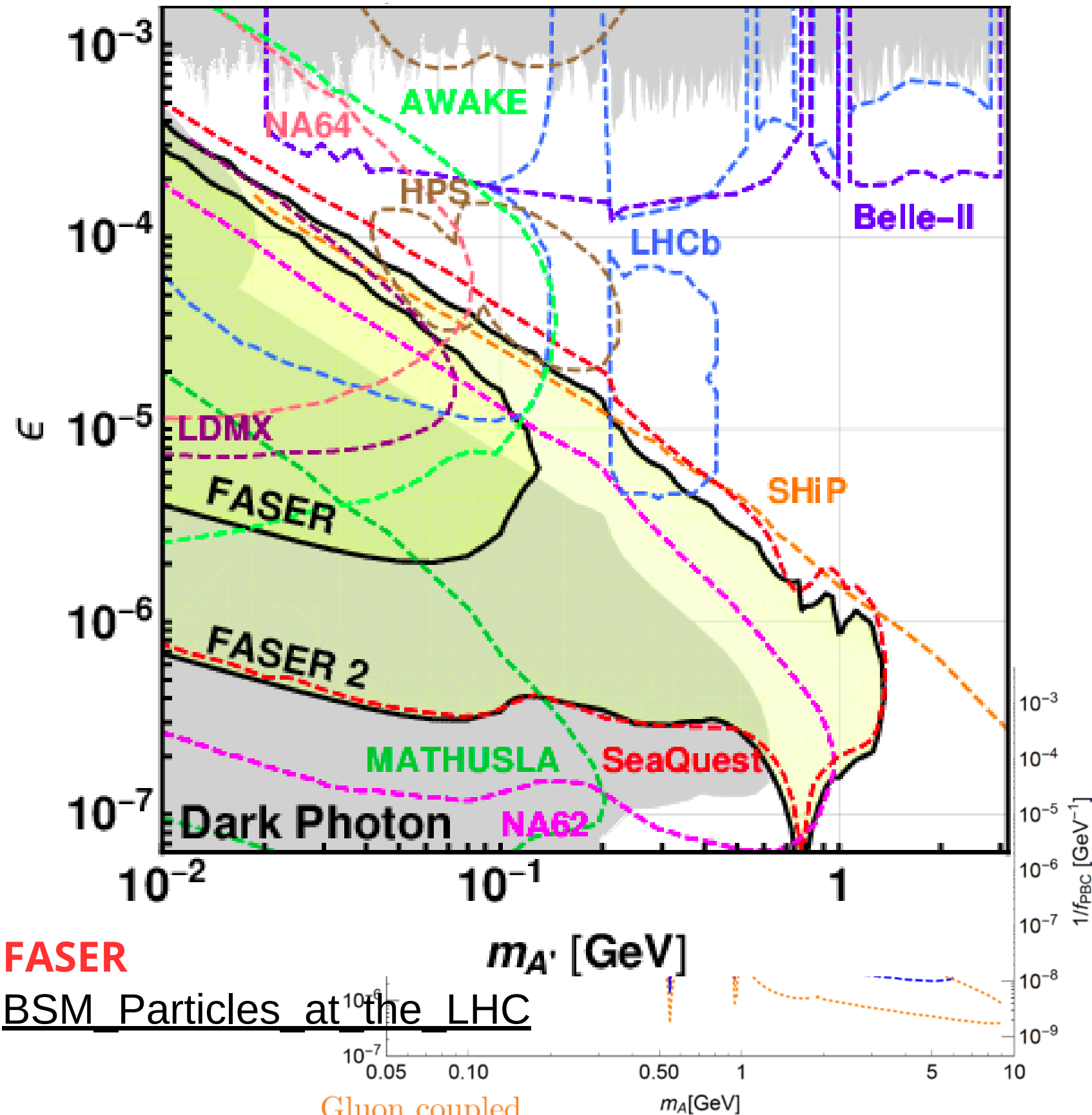
Gluon coupled
ALPs



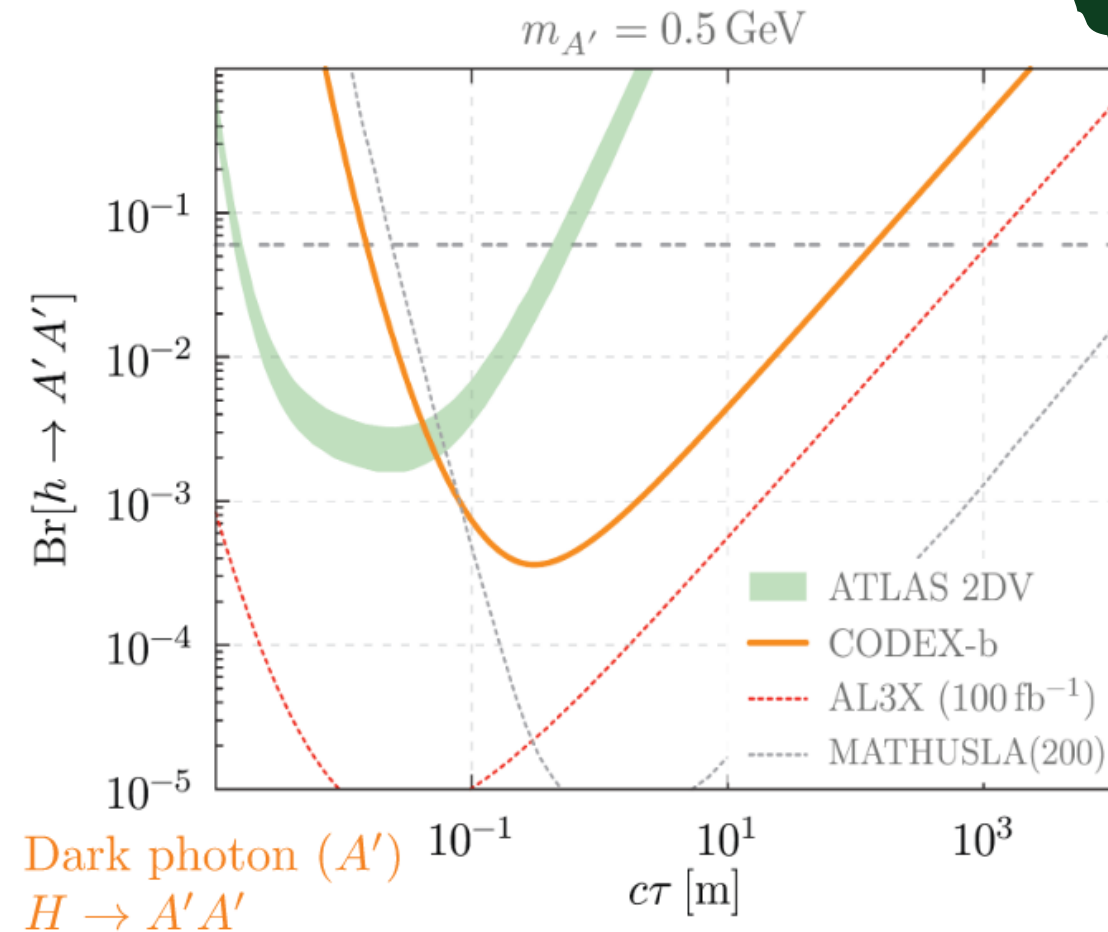
Portal models – projections

MATHUSLA, CODEX-b

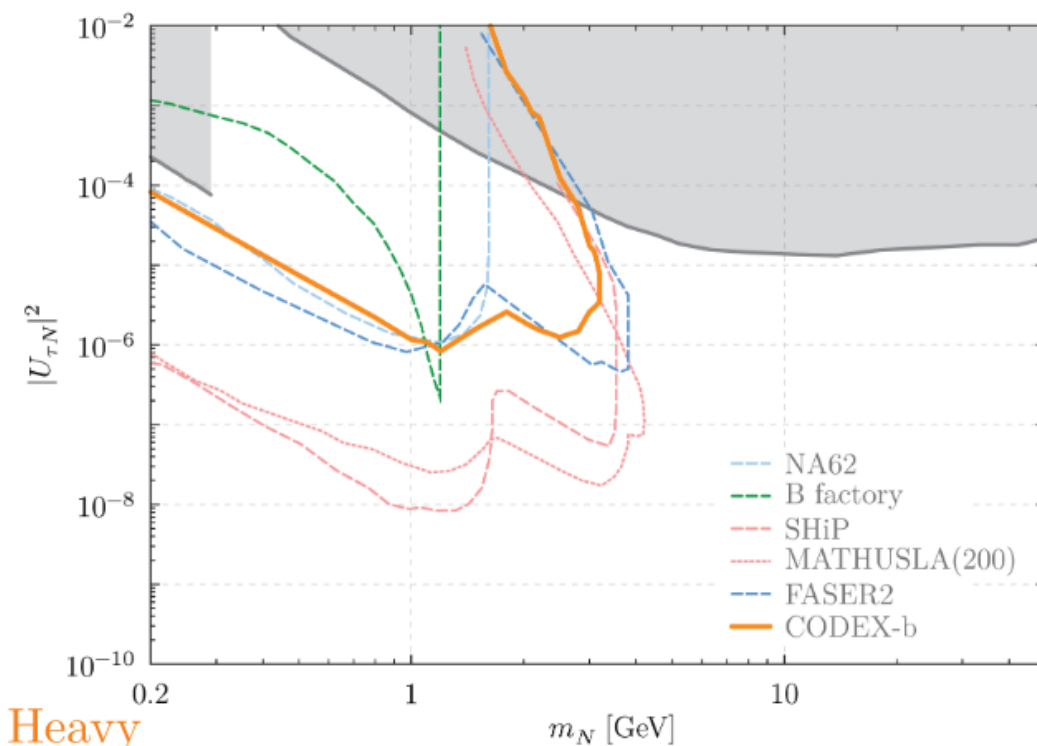
[2251269/3819001/CODEX-b](#)



Gluon coupled
ALPs



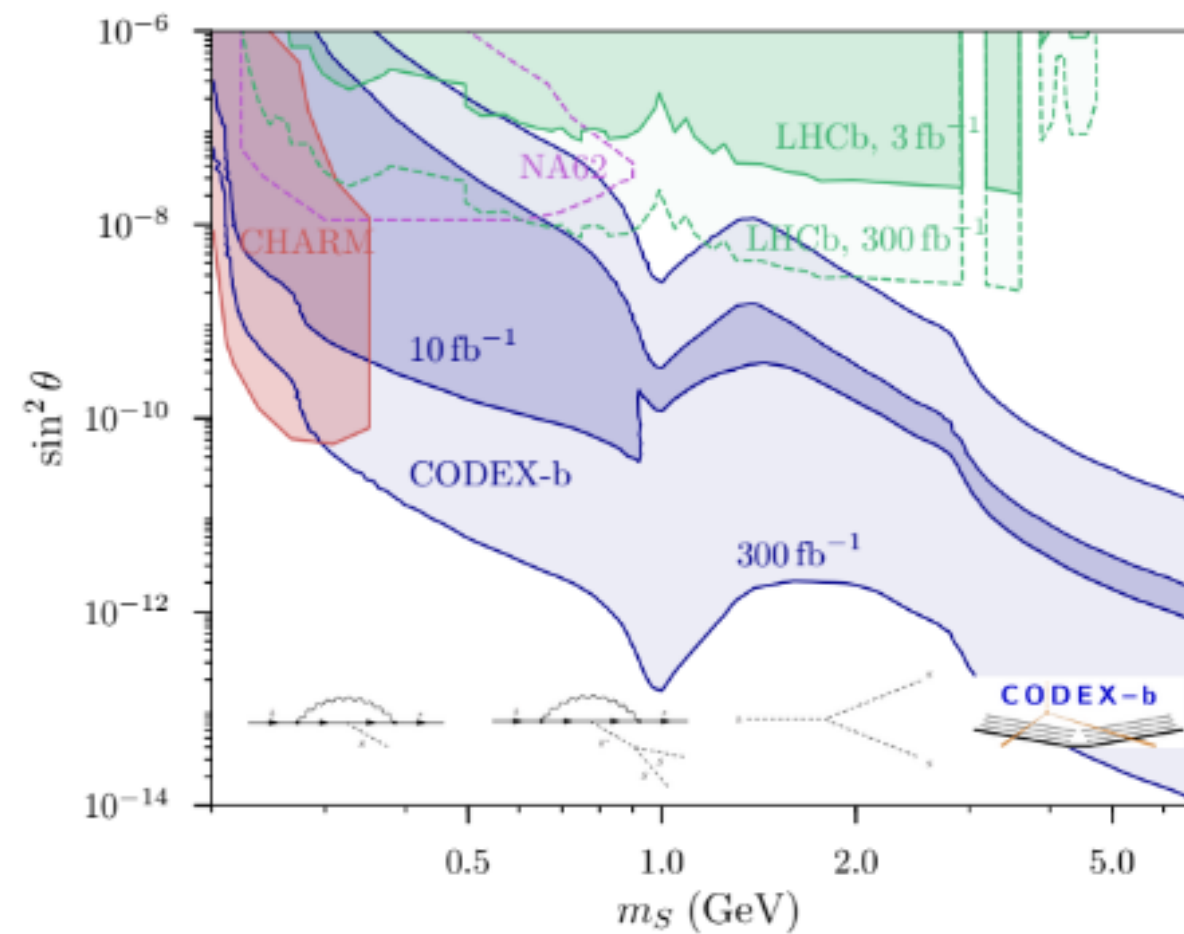
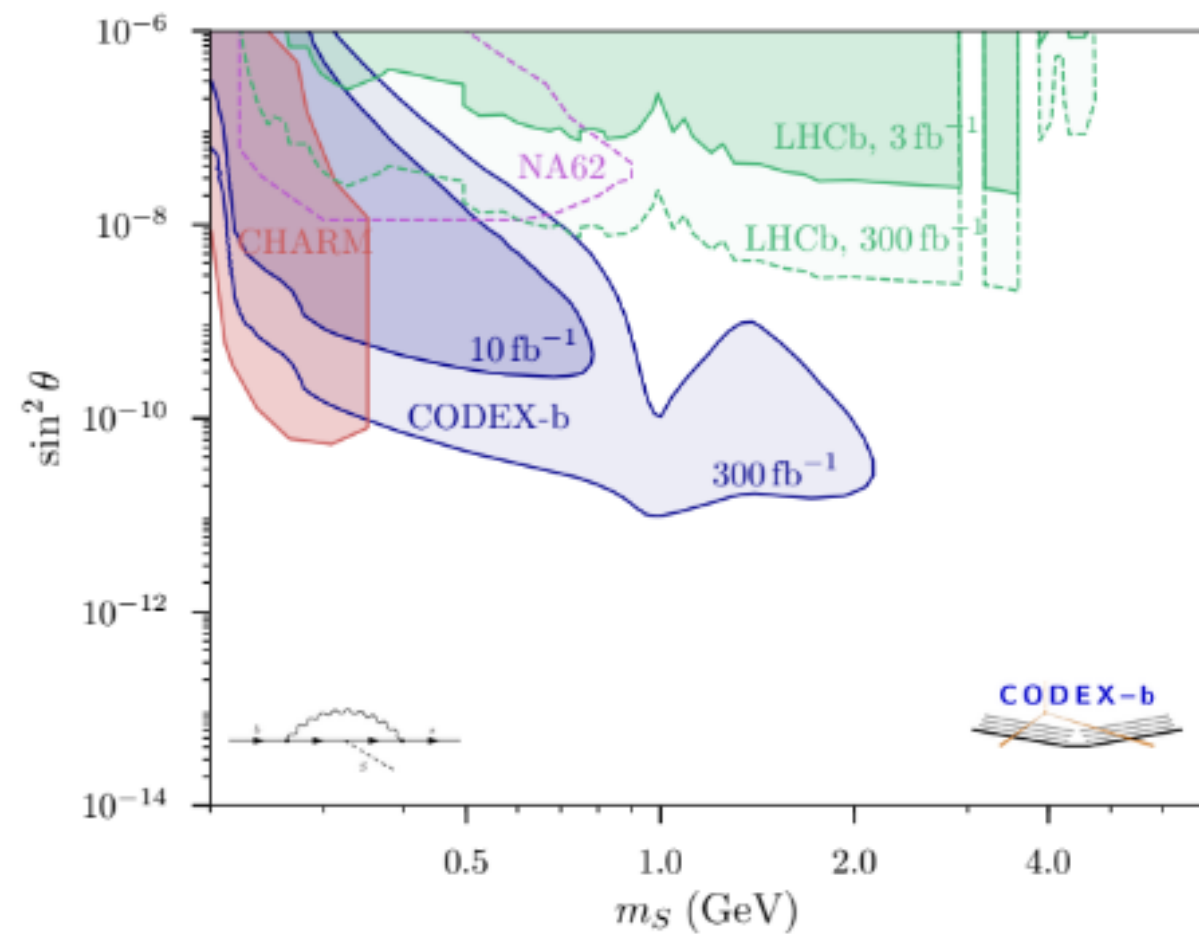
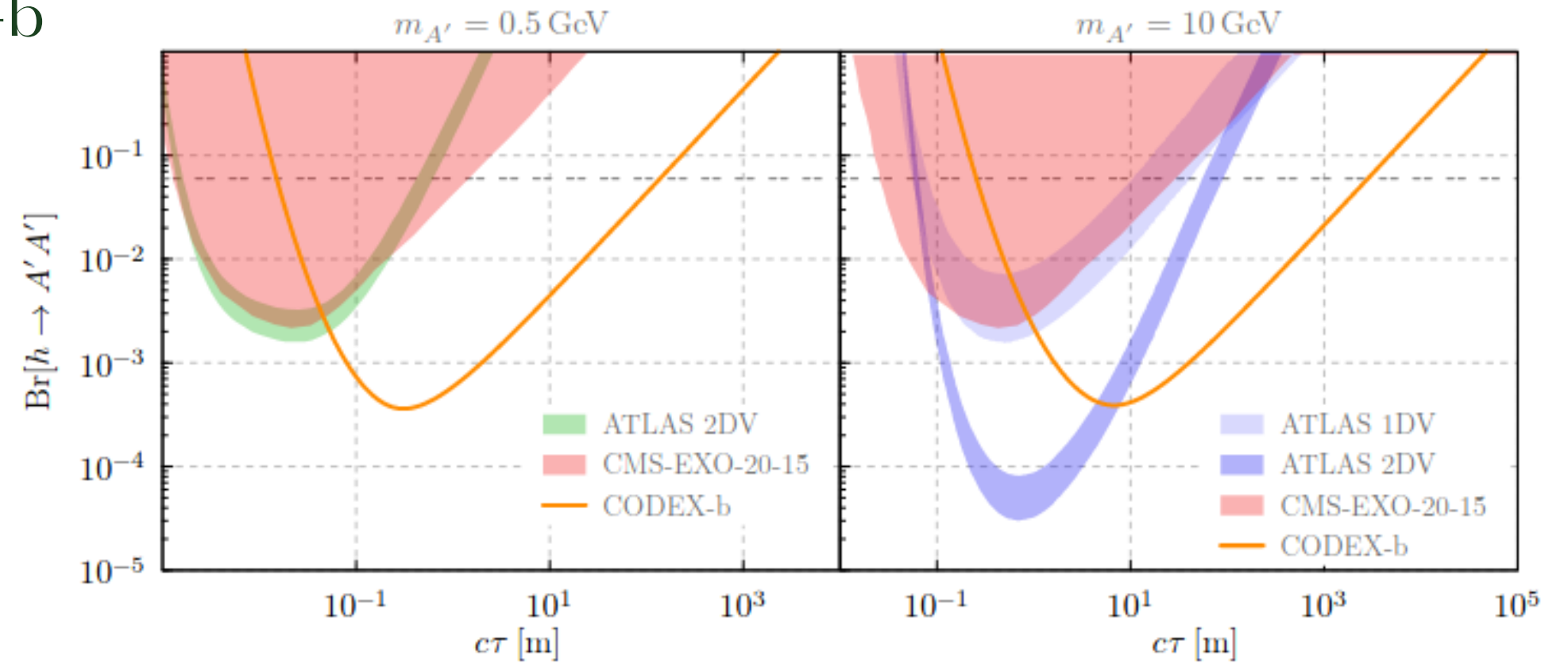
Heavy
neutral leptons



Scalar & vector portal bounds

More detail on CODEX-b

Dark Higgs



Dark photon

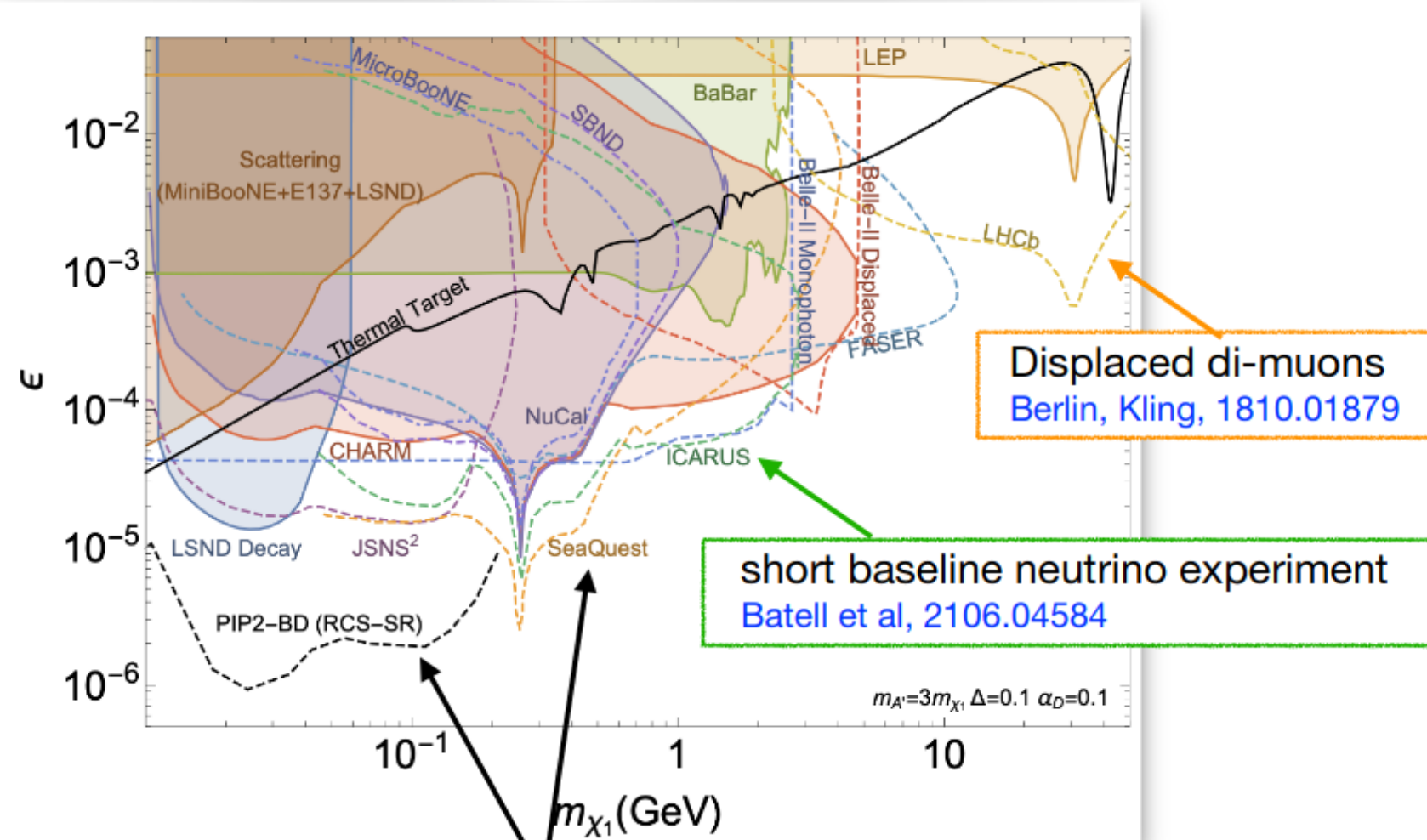
Non-minimal model – projections

MATHUSLA, CODEX-b

Fermionic inelastic DM:

- small mass splitting between 2 DM states -> require high intensity collider searches
- off-diagonal coupling of dark photon to 2 DM states motivates semi-visible search strategy

Snowmass whitepaper Krnjaic, Toro et al, 2207.00597



proton beam-dumps,
8 GeV, Toups et al, 2203.08079
120 GeV, Berlin, SG, Schuster, Toro, 1804.00661

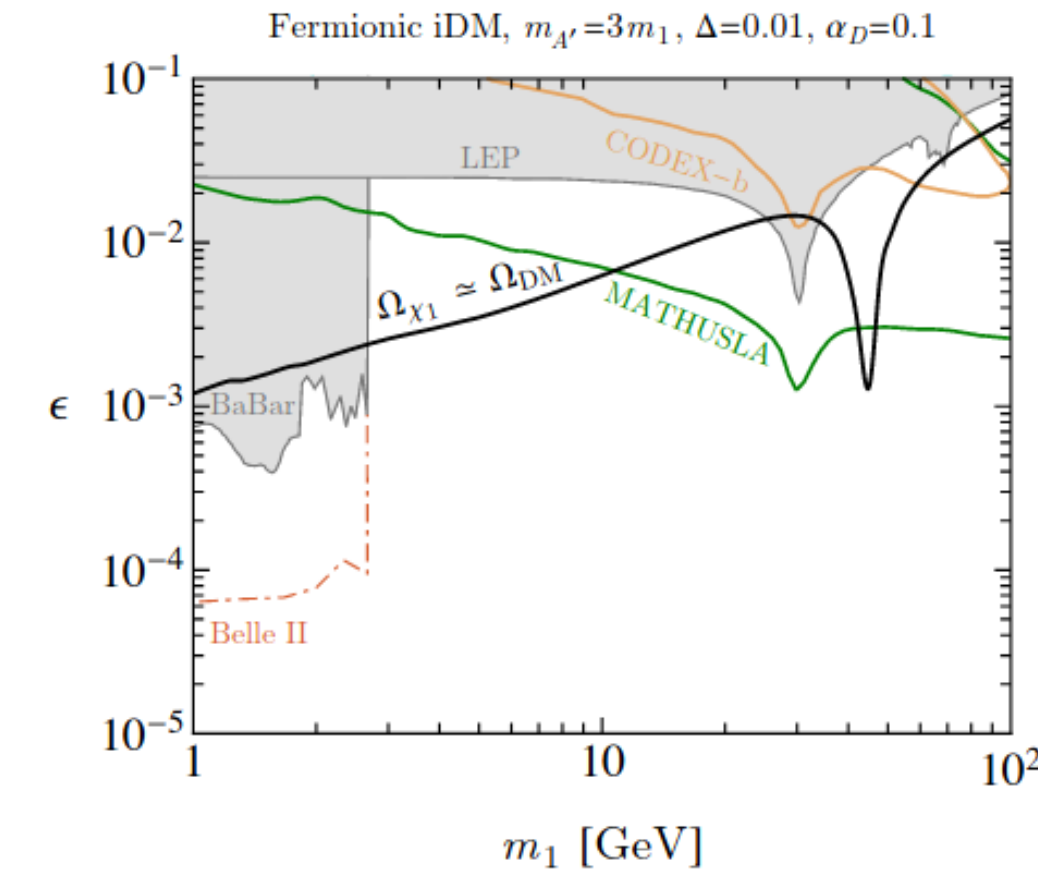
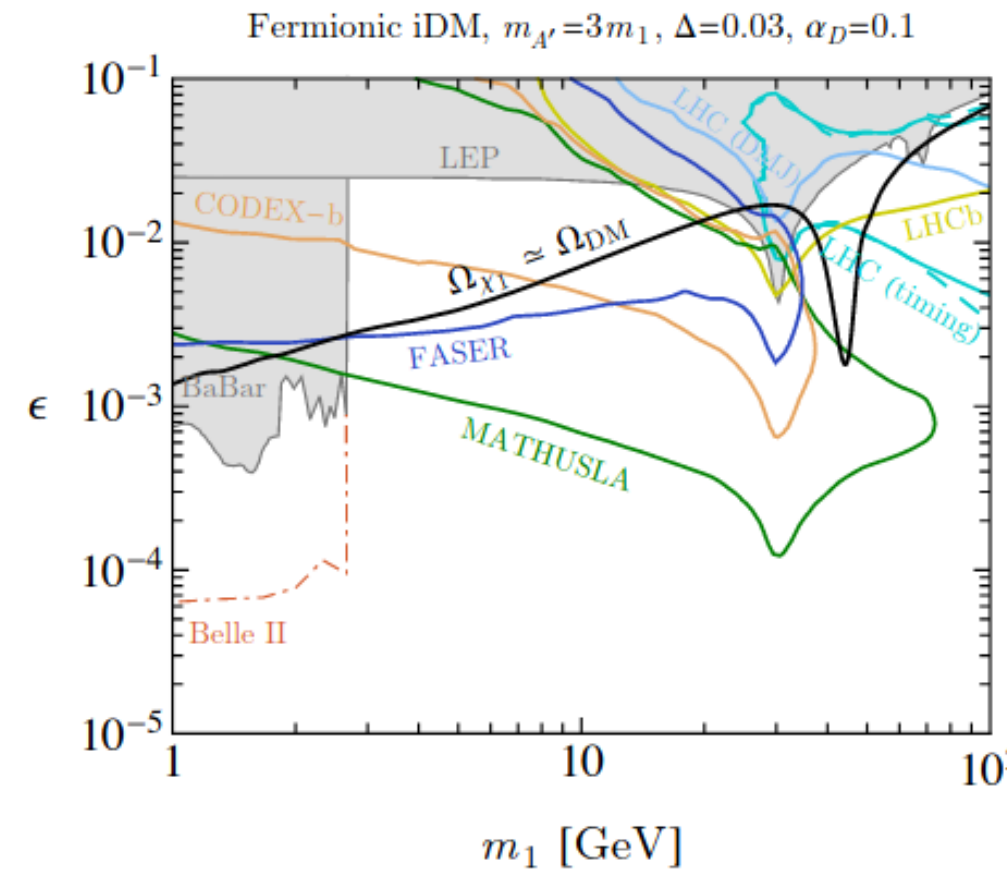


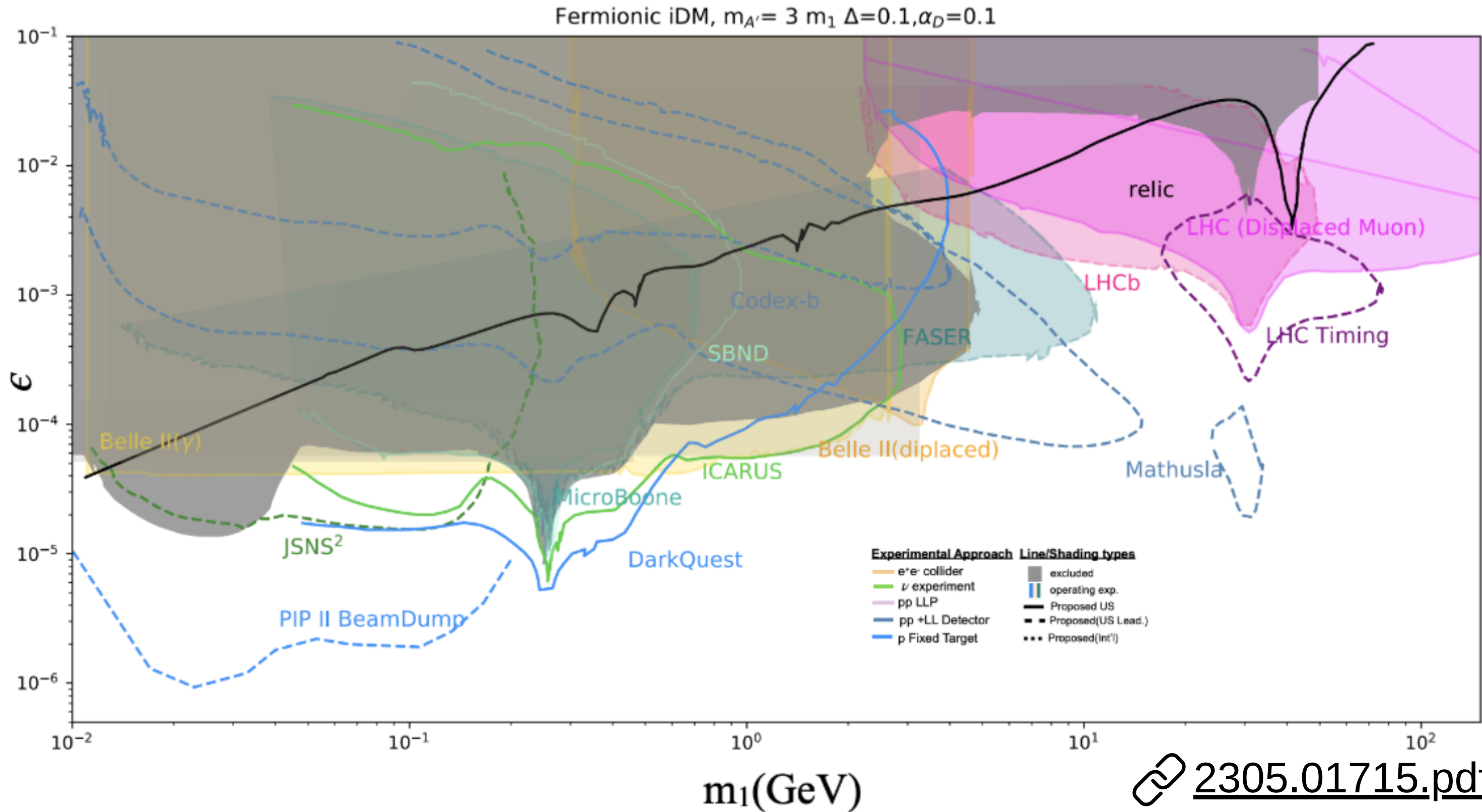
Figure 72: Reach of MATHUSLA and other LHC experiments and searches for inelastic Dark Matter (iDM) with a dark photon of mass $m_{A'}$ that has kinetic mixing ϵ with the SM photon, and mass splittings Δ in the percent range. The black curve indicates where thermal co-annihilations $\chi_2\chi_1 \rightarrow A' \rightarrow f\bar{f}$ SM fermions give the observed DM relic density. Figure taken from [851].



N

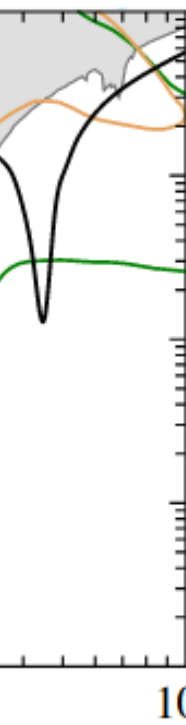
Ferm

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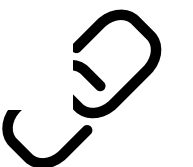
DEX-b

$\alpha_D = 0.1$



k Matter splittings

→ $f\bar{f}$



proton be
8 GeV, To
120 GeV,

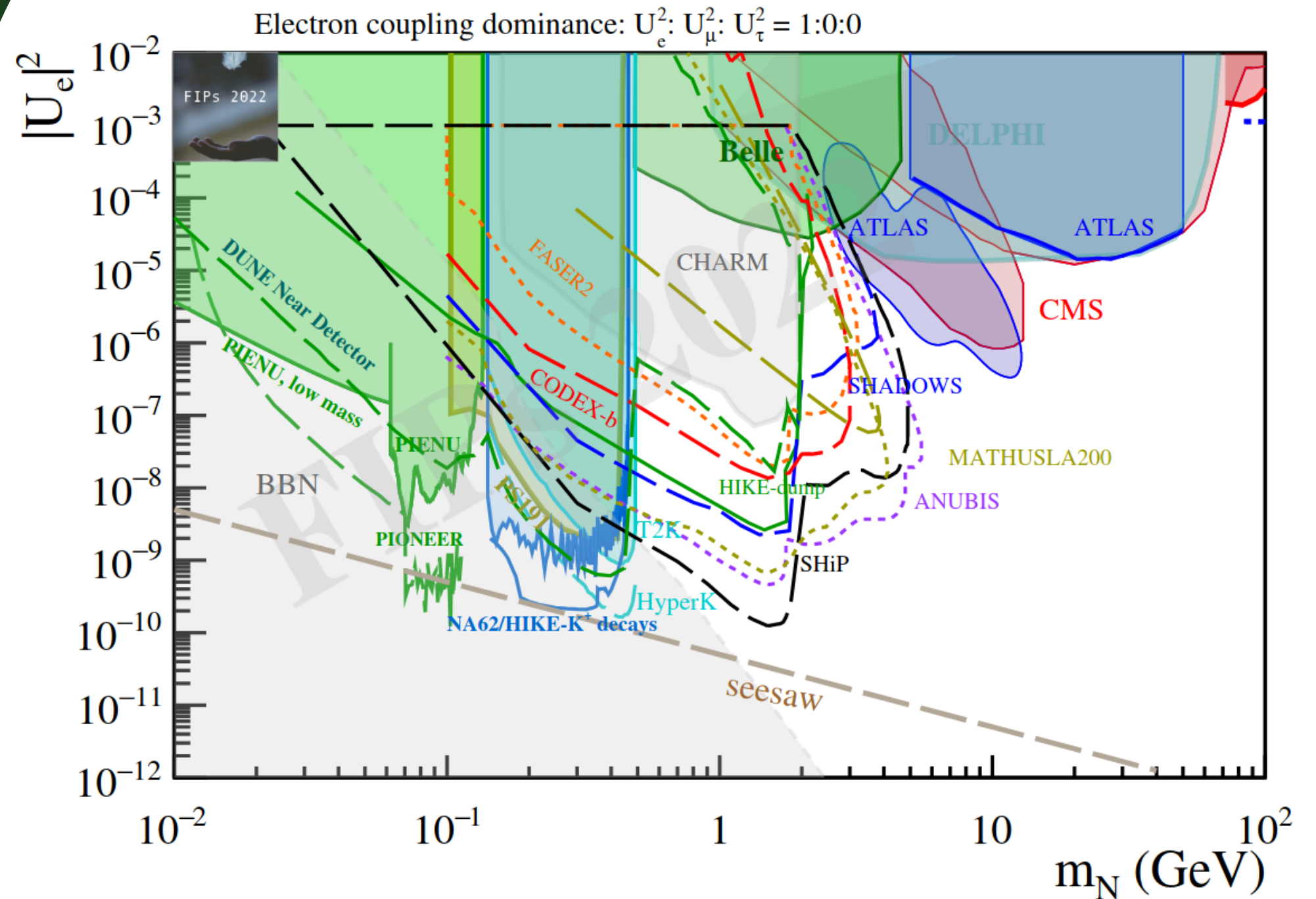
the dark photon decays to an unstable particle with a mass splitting between the excited particle (m_2) and dark matter particle (m_1) given by $\Delta = \frac{m_2 - m_1}{2(m_2 + m_1)} = 0.1$ and a dark matter coupling given by $\alpha_D = 0.1$.

SIMP: future aims to add MATHUSLA, CODEX-b (& ANUBIS?) to SIMP model plots

HNLs – ANUBIS

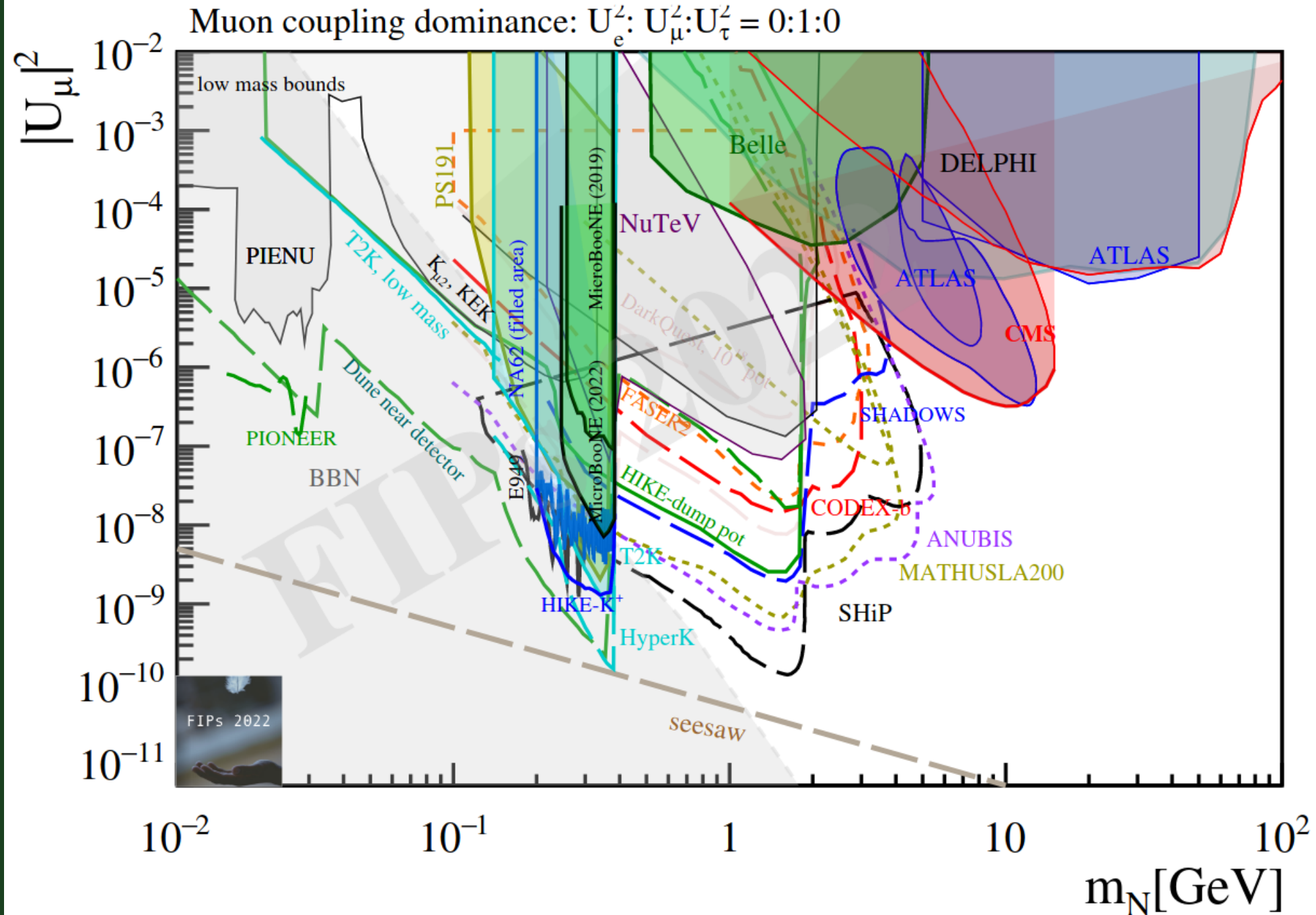
- For the minimal HNL scenario, the contributions from W 's decaying to HNLs are more important at ANUBIS than at MATHUSLA, extending the sensitivity to slightly larger HNL masses at ANUBIS
- Plots assume previous (shaft not cavern) geometry of ANUBIS so must be recalculated
 - Cavern configuration: sensitive to the products of neutral LLP decays occurring between the ATLAS muon spectrometer and the cavern ceiling
 - Shaft configuration (outdated): sensitive to decays which occur within the PX14 service shaft

Sensitivity to HNL with electron coupling (BC6)



Filled grey areas: bounds from interpretation of old data sets or astrophysical data,
 Filled coloured areas: bounds set by experiment,
 Solid lines: projections based on existing data sets,
 Dashed coloured lines: projections based on full MC simulations,
 dotted coloured lines: projections based on toy MC simulations.

HNLs – ANUBIS



Filled grey areas: bounds from interpretation of old data sets or astrophysical data,

Filled coloured areas: bounds set by experiment,

Solid lines: projections based on existing data sets,

Dashed coloured lines: projections based on full MC simulations, dotted coloured lines: projections based on toy MC simulations.

- ANUBIS has comparable sensitivity with MATHUSLA for both muon and electron coupling HNLs
- Stronger exclusion potential than forward detectors e.g. FASER2
- Note that these simplified benchmarks are not consistent with neutrino oscillation data in minimal seesaw scenarios (until flavour mixing pattern is expanded to include mixing with >1 flavour - **proposed additional benchmarks**)

Sensitivity to HNL with muon coupling (BC7)

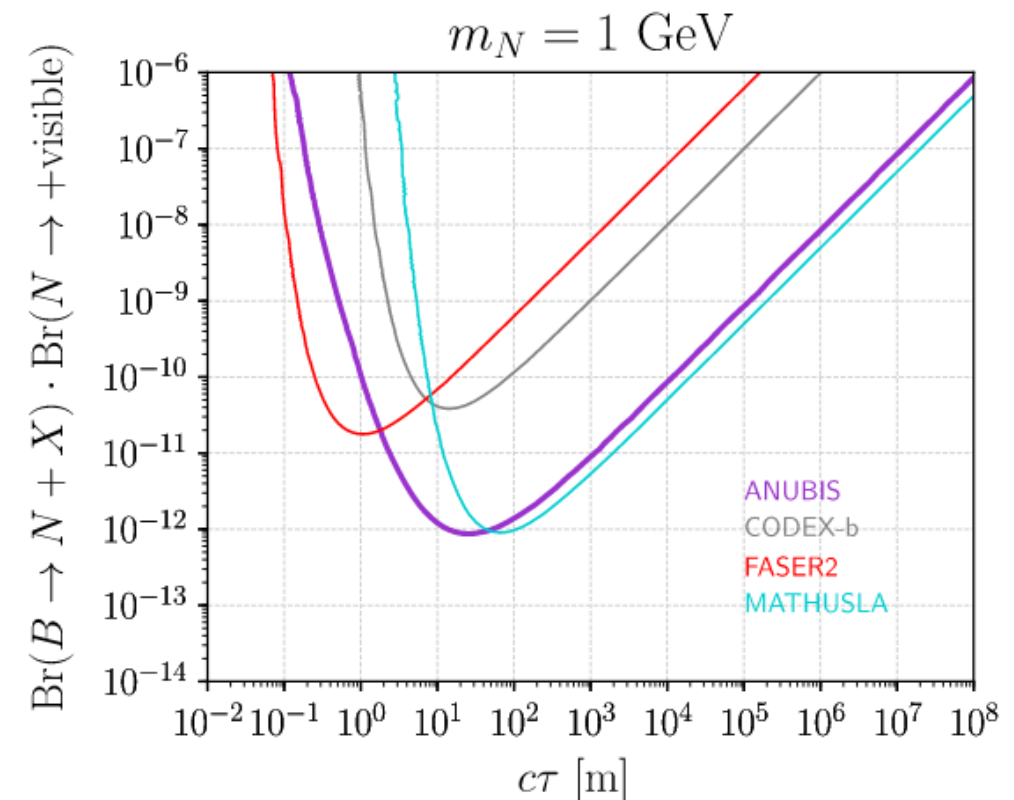
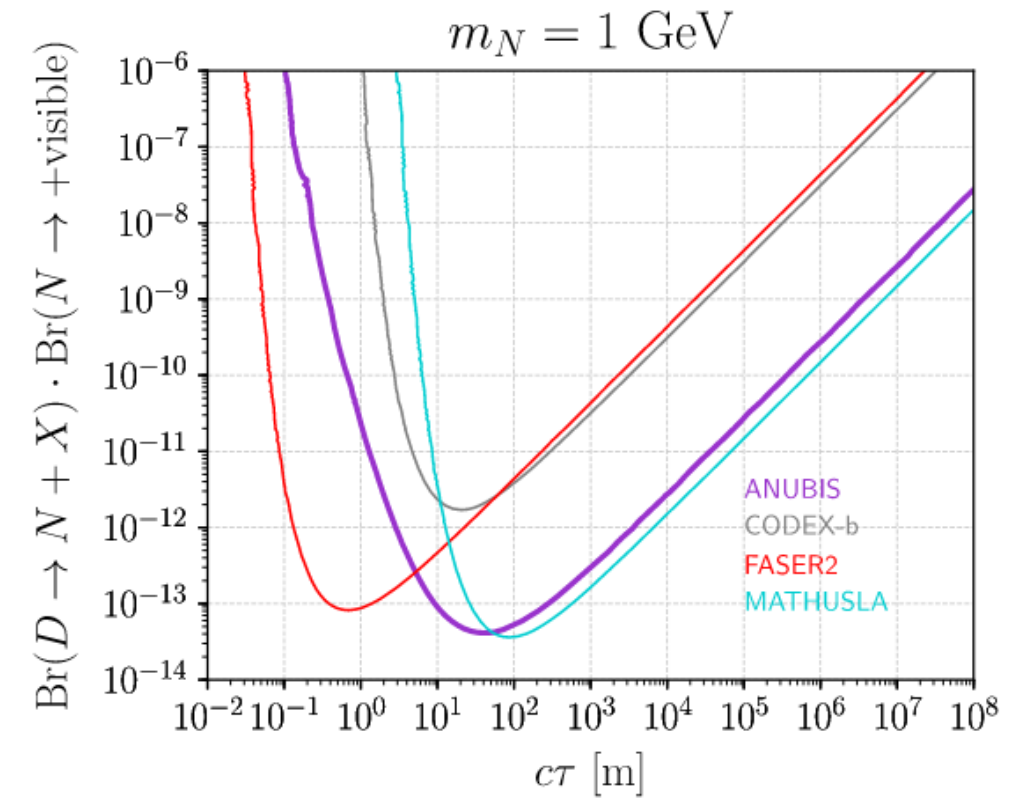
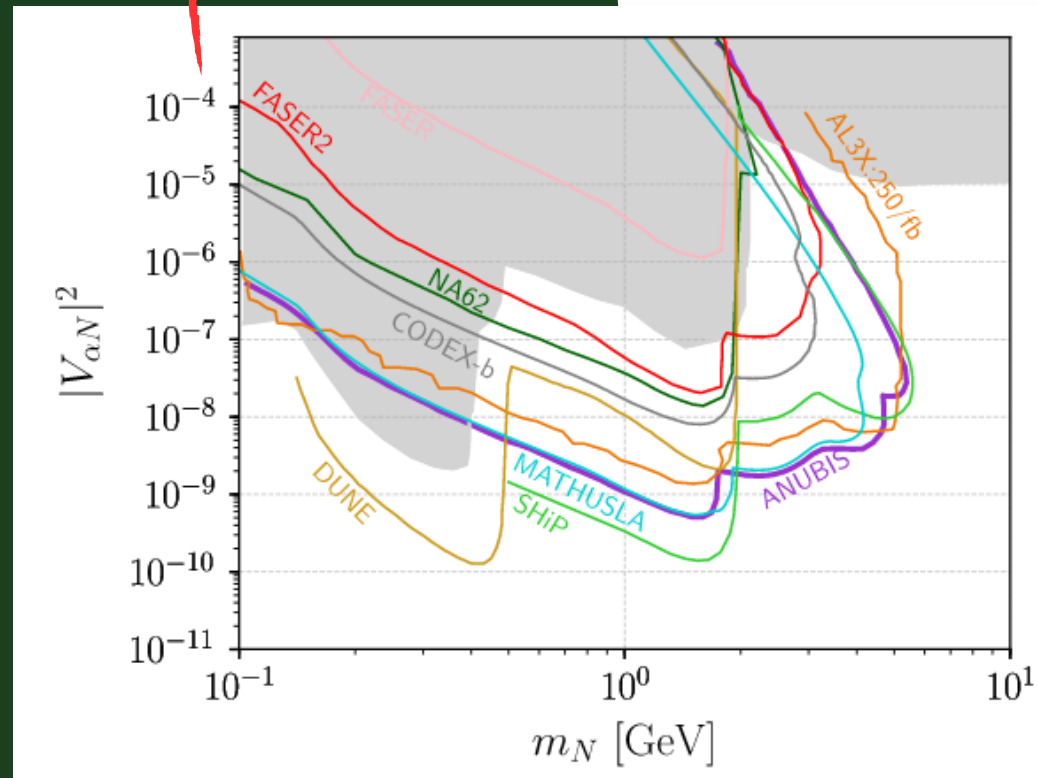
HNLs – ANUBIS

HNL bounds: minimal model

- Despite much smaller instrument volume of ANUBIS, see similar minimum branching ratios to MATHUSLA for B- and D-decays to HNLs
 - Due to smaller distance to IP
- MATHUSLA has max sensitivity at larger lifetimes for HNLs from both B- and D-decays
 - Due to distance to IP
 - & due to how HNLs of mass 1 GeV travelling inside MATHUSLA typically have boost factors larger than HNLs travelling towards ANUBIS (by factor <2)
- FASER in forward position detects lighter particles, has vastly different sensitivity here

Minimal scenario where production and decay of HNLs controlled by active-sterile neutrino mixing

Includes dominant production modes: B-, D-mesons and W-bosons

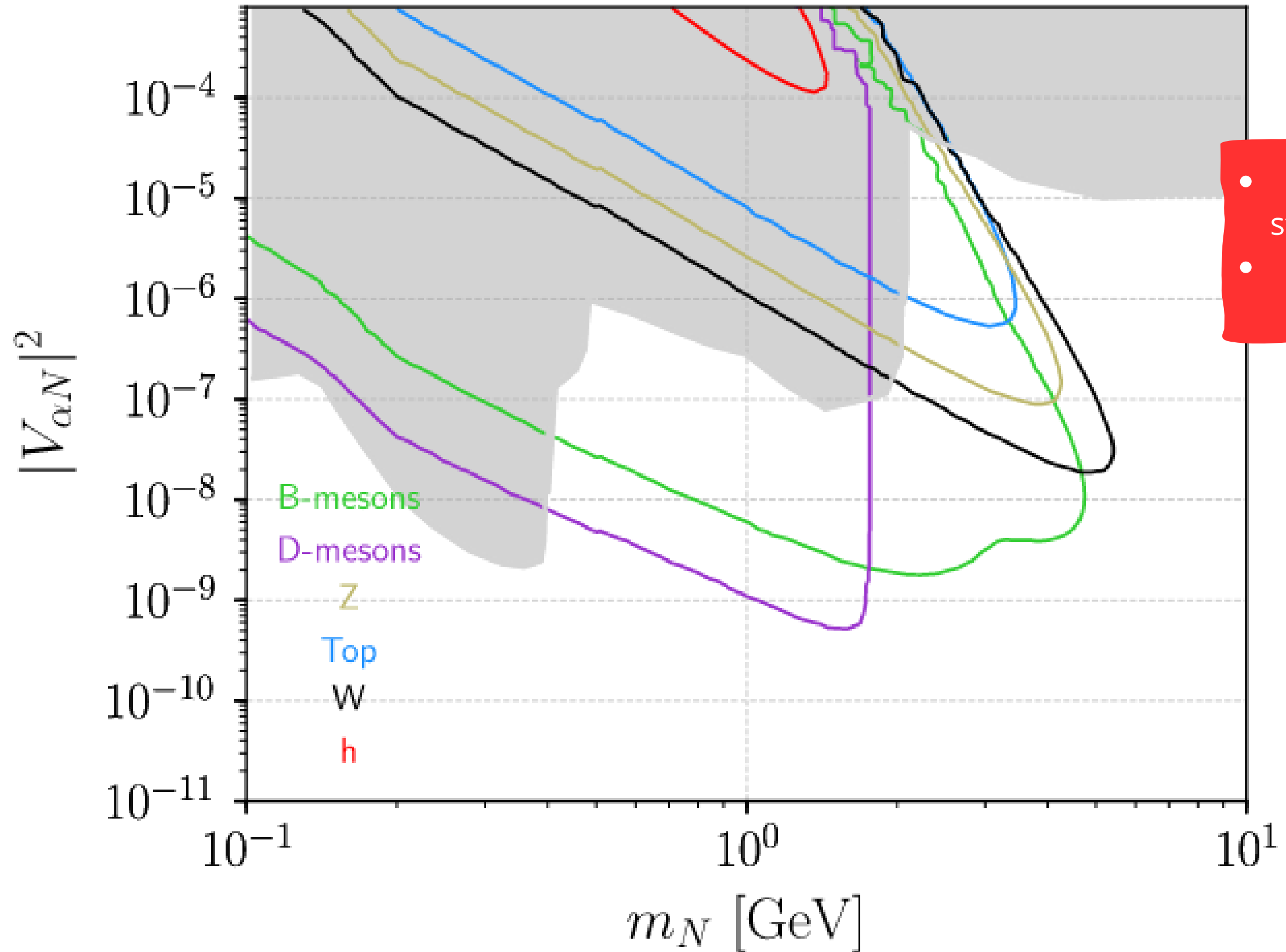


HNLs from D-decays (top) and B-decays (bottom right) in the minimal HNL scenario. HNLs from combined channels (bottom left). HNLs with one generation of N mixing with one of either electron or muon neutrino for combined sensitivities of dominant production modes.

HNLs – ANUBIS

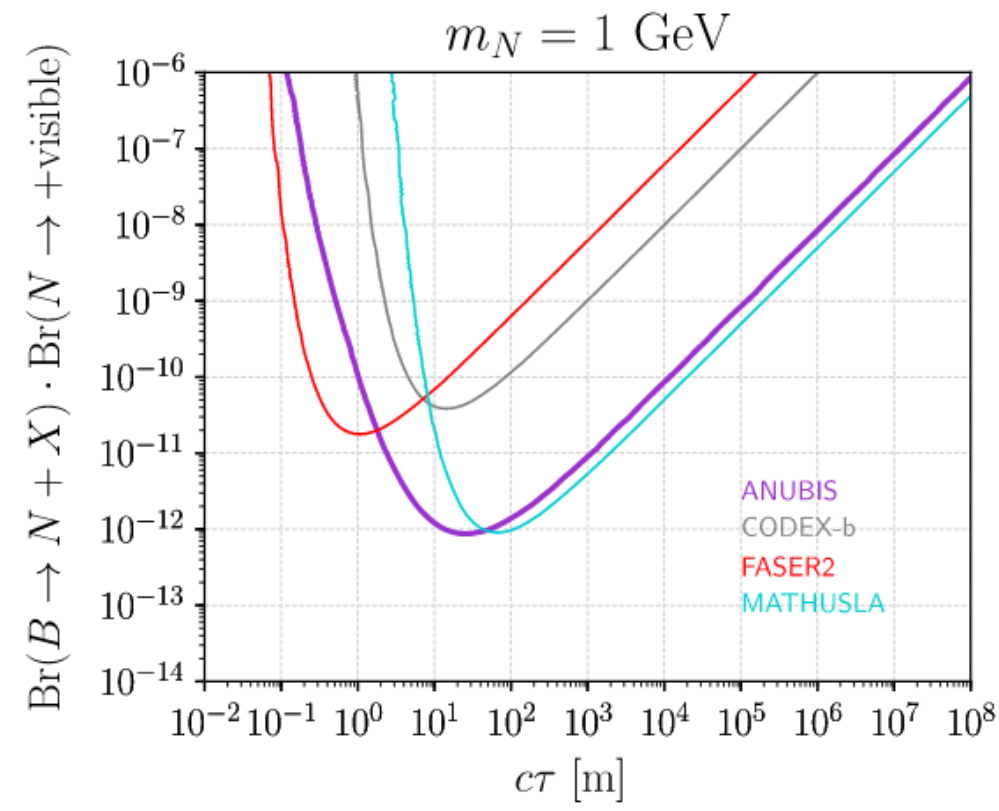
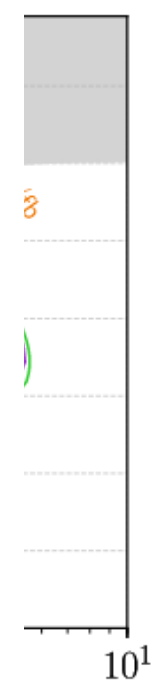
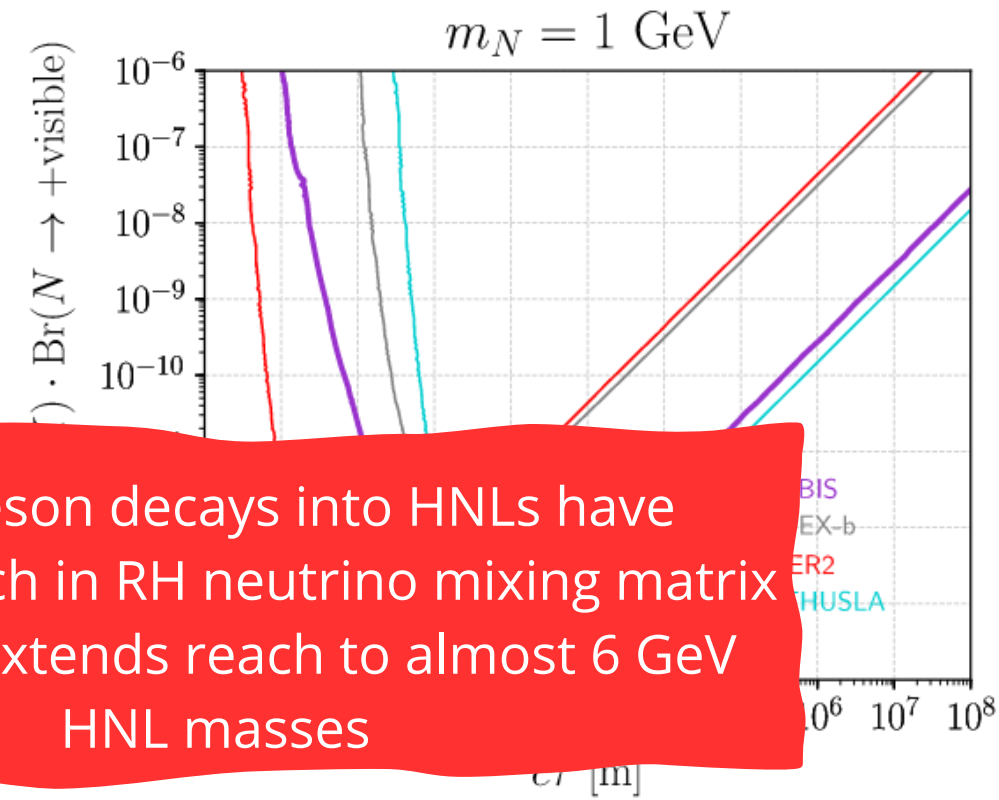
HNL

- D
- M
- F



• Heavy meson decays into HNLs have strongest reach in RH neutrino mixing matrix

• W channel extends reach to almost 6 GeV HNL masses



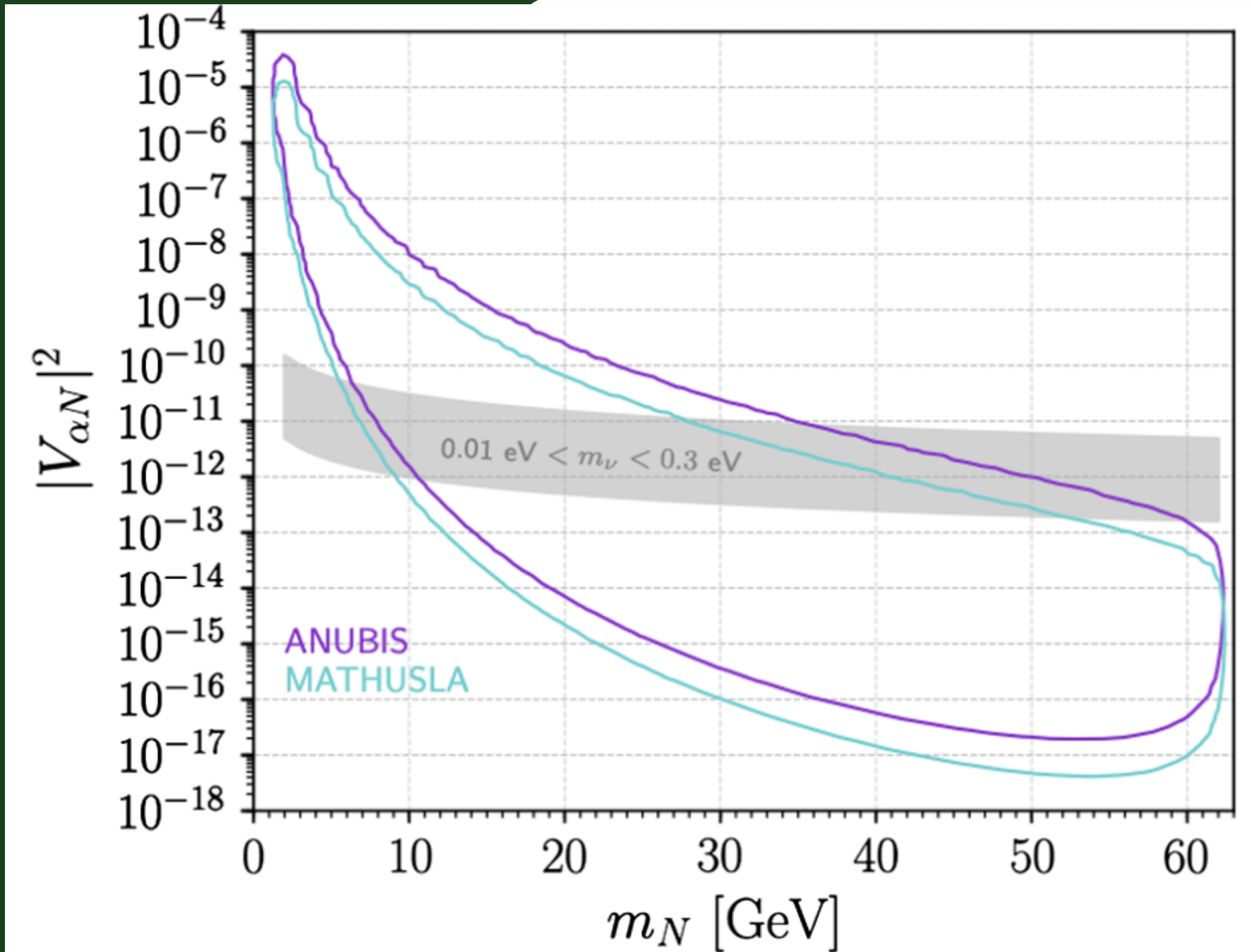
... (top) and B-decays (bottom right) in the minimal HNL combined channels (bottom left). HNLs with one generation of mixing with one of either electron or muon neutrino for combined sensitivities of dominant production modes.

HNLs – ANUBIS

Non-minimal extended scenario
where gauge group extended by
new U(1) B-L

Non-minimal scenario

- SM-like Higgs decay into a pair of HNLs which mix with electron or muon neutrino
- Grey band: interesting parameter space where **type-I seesaw limits** for light neutrino masses are between 0.01 eV and 0.3 eV
- ANUBIS sensitive to slightly **larger values of the neutrino mixing matrix squared** because closer to IP



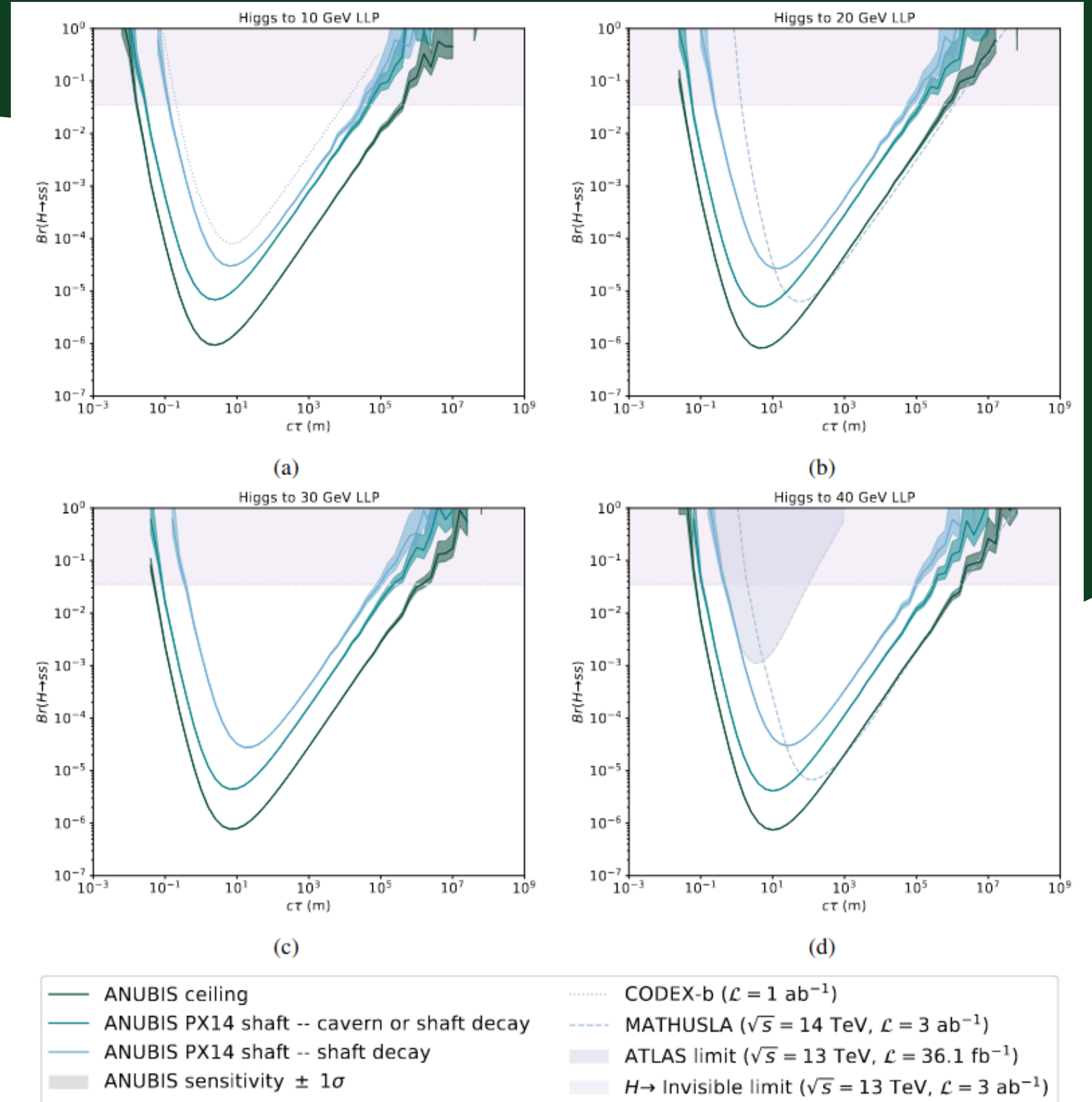
HNLs produced from SM-like Higgs decays

Our initial steps

- towards further develop these sensitivity studies,
- identifying most promising regions of parameter space
- and including simulations of the updated ANUBIS design

Review of Toby's calculations

- Sensitivity to LLPs produced at electroweak scale
 - benchmark model **neutral LLPs from Higgs decays**
- **Gluon fusion and VBF Higgs boson** production modes at NLO with Powheg
 - Higgs decays into LLPs with masses 10, 20, 30 and 40 GeV
 - subsequent decay into $b\bar{b}$, showered and hadronised with Pythia
- Investigate what fraction of decays would occur between ATLAS detectors and ANUBIS, and what fraction of events would produce **≥ 2 final-state, charged-jet particles with trajectories permitting detection with ANUBIS**
- Ceiling configuration of ANUBIS sensitive to branching ratios of LLPs with **masses of 10, 20, 30 and 40 GeV** reaching a limit of order 10^{-6} for particles with **ctau of 2, 4, 6 or 10 m** respectively



Outlook

Currently working towards...

- Best ways to simulate meson decays for HNL production
 - Pythia, MadSpin
- FeynRules parameters for producing UFOs consistent with those used by other experiments
 - Starting with simplest configurations of scalar and vector portal models
- Understanding axion portal model of interest to ANUBIS

An aerial night photograph of a city, showing a dense network of streets and illuminated buildings. The image is overlaid with a white circular graphic that is partially cut off at the top and bottom. The background is a dark green gradient.

Thanks

Backup

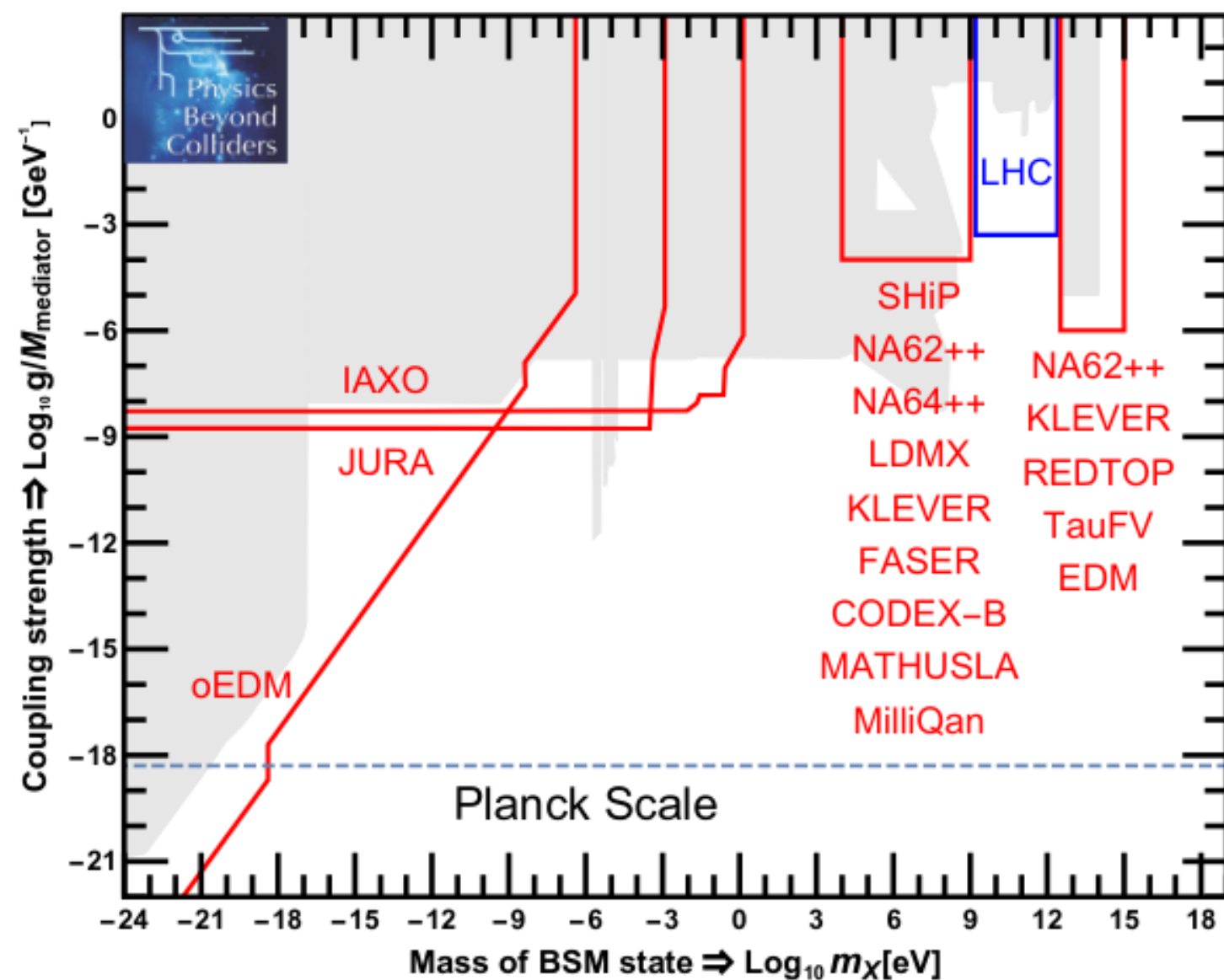


Figure 1: Schematic overview of the BSM landscape, based on a selection of specific models, with a rough outline of the areas targeted by the PBC experiments. The x -axis corresponds to the mass m_X of the lightest BSM state, and the y -axis to the scale of the effective new interaction $f = M_{\text{Mediator}}/g$, where M_{Mediator} is the mass of a heavy mediator and g its (dimensionless) coupling constant to the Standard Model. The grey shaded area outlines the currently excluded regions for a class of models corresponding to the benchmarks BC9 and BC11 (see Refs [27, 37, 38]).

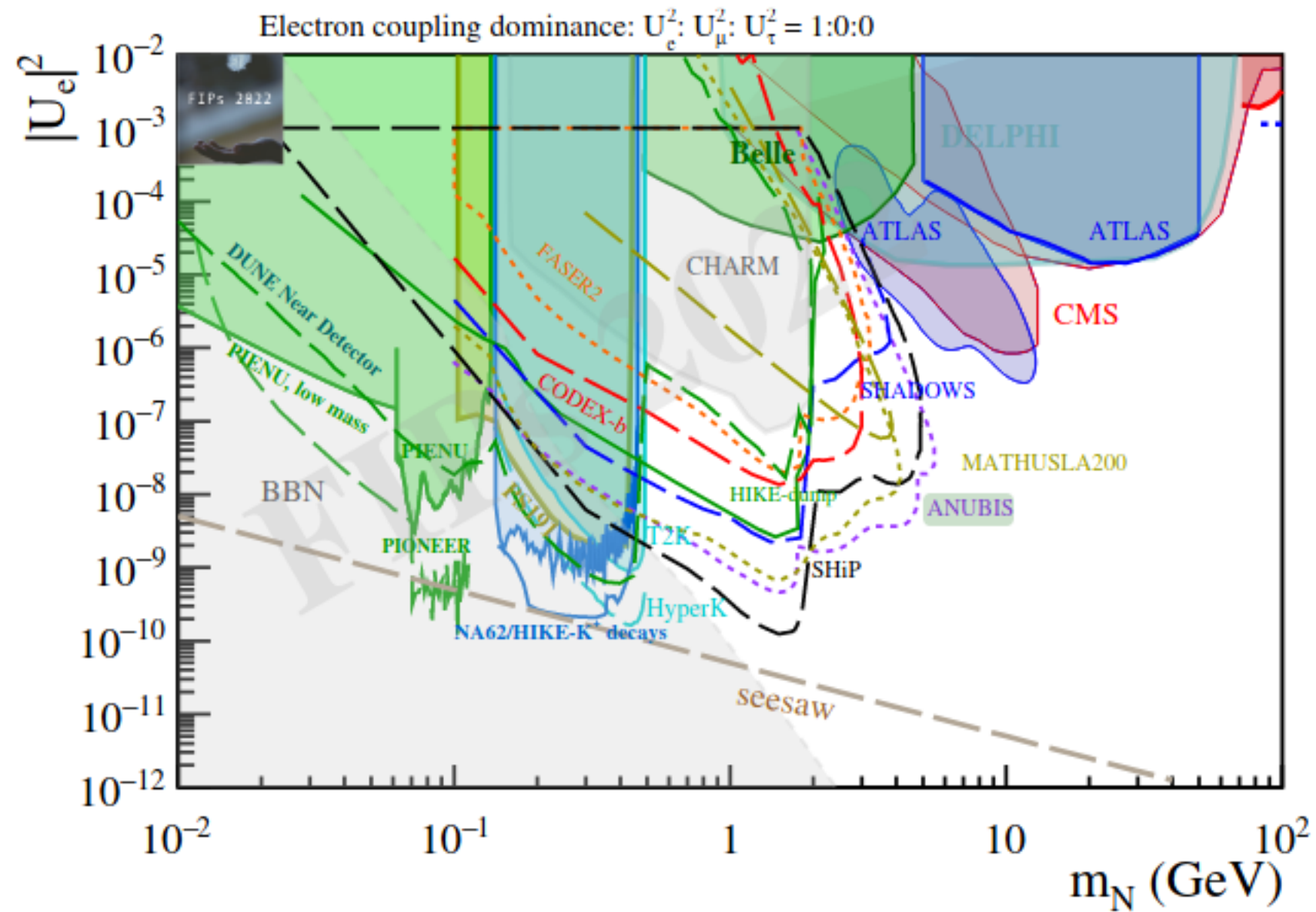


Figure 174: **Sensitivity to HNL with electron coupling (BC6).** Current bounds and future projections for 90% CL exclusion limits. **Legend:** filled gray areas are bounds coming from interpretation of old data sets or astrophysical data; filled coloured areas are bounds set by experimental collaborations; Solid coloured lines are projections based on existing data sets; Dashed coloured lines are projections based on full Monte Carlo simulations; Dotted coloured lines are projections based on toy Monte Carlo simulations. Filled areas are existing bounds from: PS191 [1561], CHARM [1432], PIENU [1785], NA62 (K_{eN}) [1177], T2K [1786], Belle [1787], DELPHI [1788], ATLAS [1627, 1621], and CMS [1625, 1620]. Coloured curves are projections from: PIONEER [1789], HIKE- K^+ [1392, 1191]; HIKE-dump [1392, 1191]; DarkQuest [1258], Belle II [1782], FASER2 [1416]; DUNE near detector [1172], Hyper-K (projections based on [1562]), CODEX-b [1030], SHiP [1402], SHADOWS [1401] and MATHUSLA200 [1441]. The BBN bounds are from [1686]. The dashed seesaw line is given by $|U_\alpha|^2 = \sqrt{\Delta m_{atm}^2}/m_N$ corresponding to the naive seesaw scaling and should be considered only as indicative.