

Probing Long-Lived Heavy Neutral Leptons in Cosmology









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Outlines

- 1. Heavy Neutral Leptons & Seesaw Mechanism
- 2. HNL-ALP Interaction
- 3. Big Bang Nucleosynthesis
- 4. HNL and ALP Populations
- 5. Conclusion



Heavy Neutral Lepton & Seesaw Mechanism

Motivation: Massless SM neutrinos HNLs give mass to light active neutrinos

Dirac mass by **Right-Handed Neutrinos**:

$$\mathcal{L}_{Dirac} = -Y_{\nu}\bar{L} \cdot H\nu_R + \text{h.c.}$$

Majorana Right-Handed Neutrinos:

$$\mathcal{L}_{Majorana} = -Y_{\nu}\bar{N}L \cdot H - 1/2\bar{N}^{c}M_{R}N + \text{h.c.}$$

$$\mathcal{M}_{\nu} = \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} = U \begin{pmatrix} m_{\nu} & 0 \\ 0 & m_N \end{pmatrix} U^T$$



Light Neutrino Mass:

$$m_{\nu} \simeq \frac{m_D^2}{M_R} \simeq \frac{(1)^2}{10^{10}} \text{GeV} = 0.1 \text{eV}$$

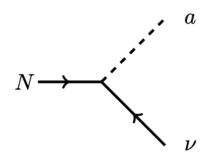
$$|V_{eN}|^2 \simeq \frac{m_{\nu}}{M_R}$$
 Active-Sterile mixing

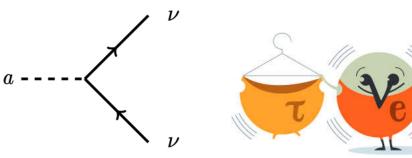


HNL and Axion Like Particle (ALP)

$$\mathcal{L}_{aN} = \frac{C_{aN}}{f_a} \left(\partial_{\mu} a \right) \bar{N} \gamma^{\mu} \gamma_5 N = -\frac{2iC_{aN}}{f_a} m_N a \bar{N} \gamma_5 N = -\frac{2iC_{aN}U_{\nu N}}{f_a} m_N a \bar{N} \gamma_5 \nu$$

HNL-ALP interaction



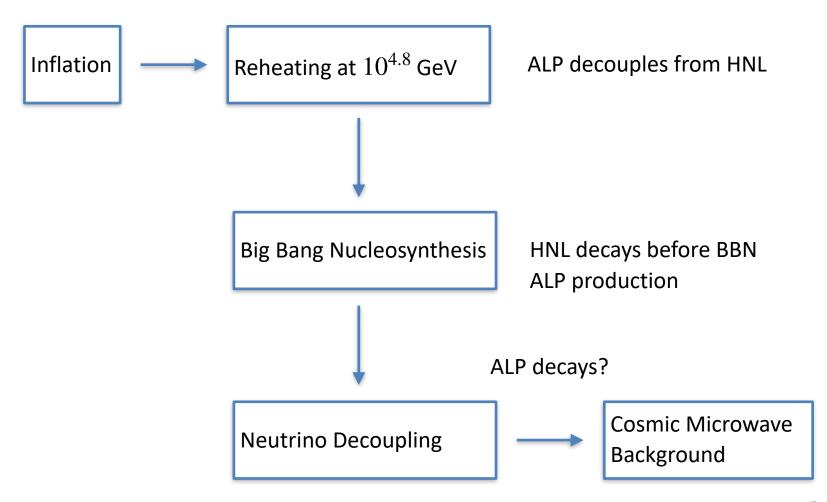


$$\tau_N = \frac{1}{\Gamma_{Na\nu}} \simeq \frac{8\pi f_a^2}{U_{\nu N}^2 m_N^3} = 1 \sec \times 8\pi \times \left(\frac{f_a}{1 \text{TeV}}\right)^2 \times \left(\frac{10^{-7}}{U_{\nu N}}\right)^2 \times \left(\frac{100 \text{MeV}}{m_N}\right)^3$$

GeV HNL can survive for 0.1s
$$\left| f_a = 10^3 \, \mathrm{GeV}, \, \left| V_{eN} \right|^2 = 10^{-9} \right|$$

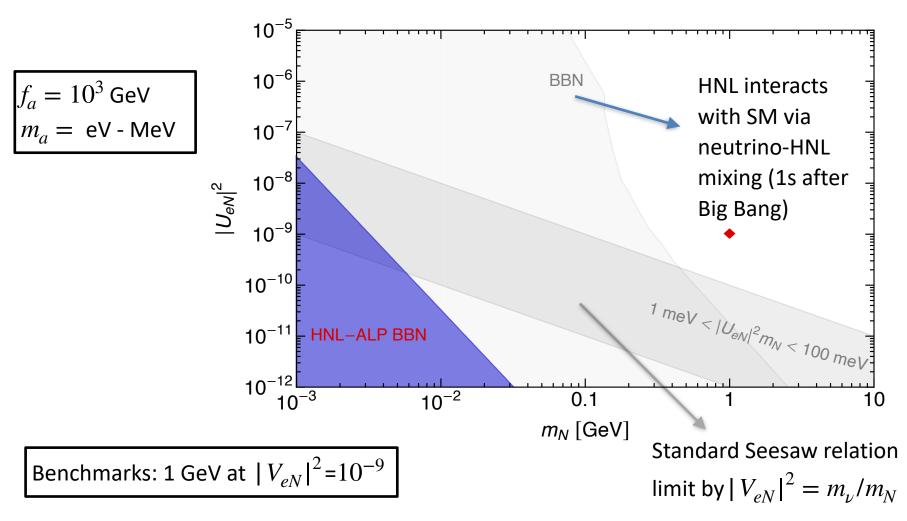


Thermal History





Big Bang Nucleosynthesis





Conclusion

- 1. Neutrino mass from light HNL can be achieved by non-standard seesaw mechanisms
- 2. BBN limits on HNL mass can be relaxed by exotic interactions
- 3. ALPs can be produced by "Freeze-In" from HNL decay
- 4. Long-Lived HNL can be probed by beam dump experiments

STANDARD MODEL &



Thanks for listening Any Questions?

