

# Listening for dark matter

---

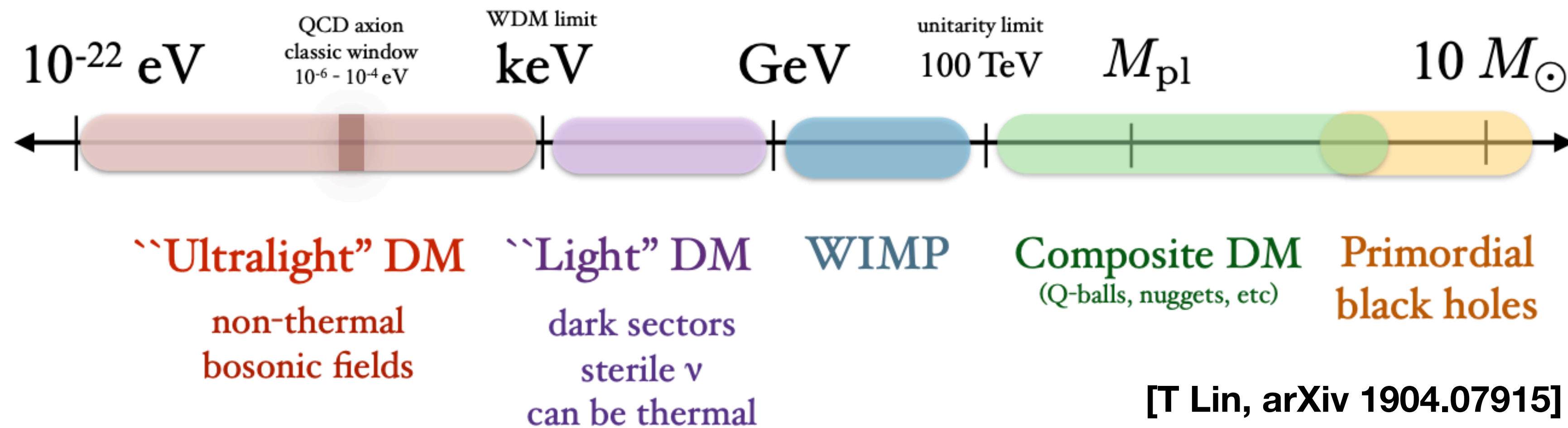
**Damon Cleaver** and Christopher McCabe  
King's College London

*Gong Talk, YTF 2023, IPPP Durham*

# What do we know?

$$\Omega_{\text{CDM}} h^2 \sim 0.120 \pm 0.001$$


[Planck, 2018]



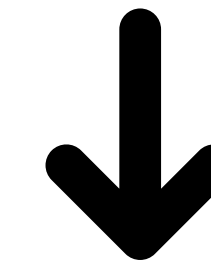
# What are Direct Detection (DD) experiments sensitive to?

---

$\sigma_\chi$  

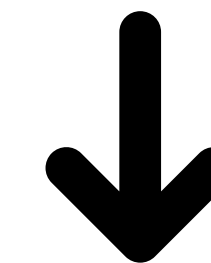
$\frac{\sigma_\chi}{m_\chi}$  

Why? Density of DM is  
**fixed**



Higher mass,  
lower number  
density

$$n_\chi \sim 1/m_\chi$$



Experiment with **higher exposure**  
**required** to constrain **same cross**  
**section**

# Macroscopic Dark Matter

---

- **General class** of DM with:
  - Planck-mass (or above)  $m_\chi \gtrsim M_{pl}$  - often parameterised in grams (!)
  - DM nucleon cross section is roughly that of the geometric cross section  $\sigma_{\chi N} \approx \sigma_{\chi,geo} \equiv \pi R_\chi^2$ .
  - $R_\chi$  encodes the macro geometry (from being constituent DM) and absorbs any short range interaction - correction to the geometric radius

# Who has large exposure experiments?

---

- Proposition for **acoustic neutrino experiments**.
- Designed to find low number density UHE Neutrinos
- Showers created from PeV-EeV ( $10^{15}$  -  $10^{18}$  eV) neutrinos could create a detectable pressure wave from **thermo-acoustic heating**.

$$\nabla^2 \underbrace{P(\mathbf{r}, t)}_{\text{Acoustic pressure}} - \frac{1}{c_s^2} \frac{\partial^2 P(\mathbf{r}, t)}{\partial t^2} = - \frac{\alpha}{C_p} \frac{\partial^2 \underbrace{q(\mathbf{r}, t)}_{\text{Energy Deposition Density}}}{\partial t^2}$$

# What experiment can be done?

---

- Idea: Instrument **up to 100km<sup>3</sup> of water** (albeit ambitiously) with sensitive hydrophones to detect the resulting pressure wave
- **Why not also look for macro dark matter?**

$$\left(\frac{dE}{dx}\right)_{\text{water}} = \rho_w \sigma_\chi v_\chi^2 \exp\left(-2\frac{\sigma_\chi}{m_\chi} \rho_w L\right) \xrightarrow{\text{Thermoacoustic Heating (meat of calculation)}} P(\mathbf{r}, t)$$

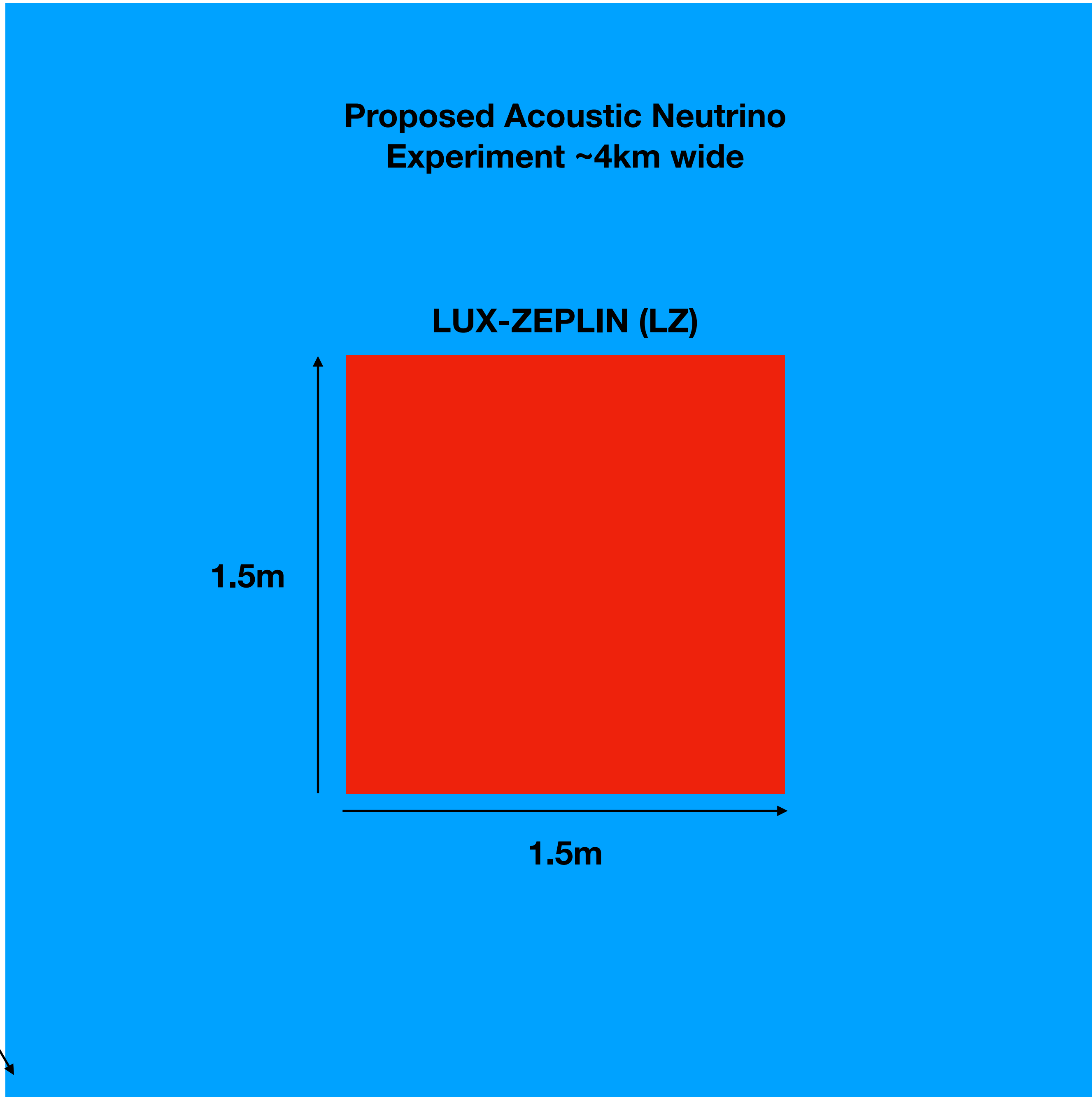
**Proposed Acoustic Neutrino  
Experiment ~4km wide**

**LUX-ZEPLIN (LZ)**

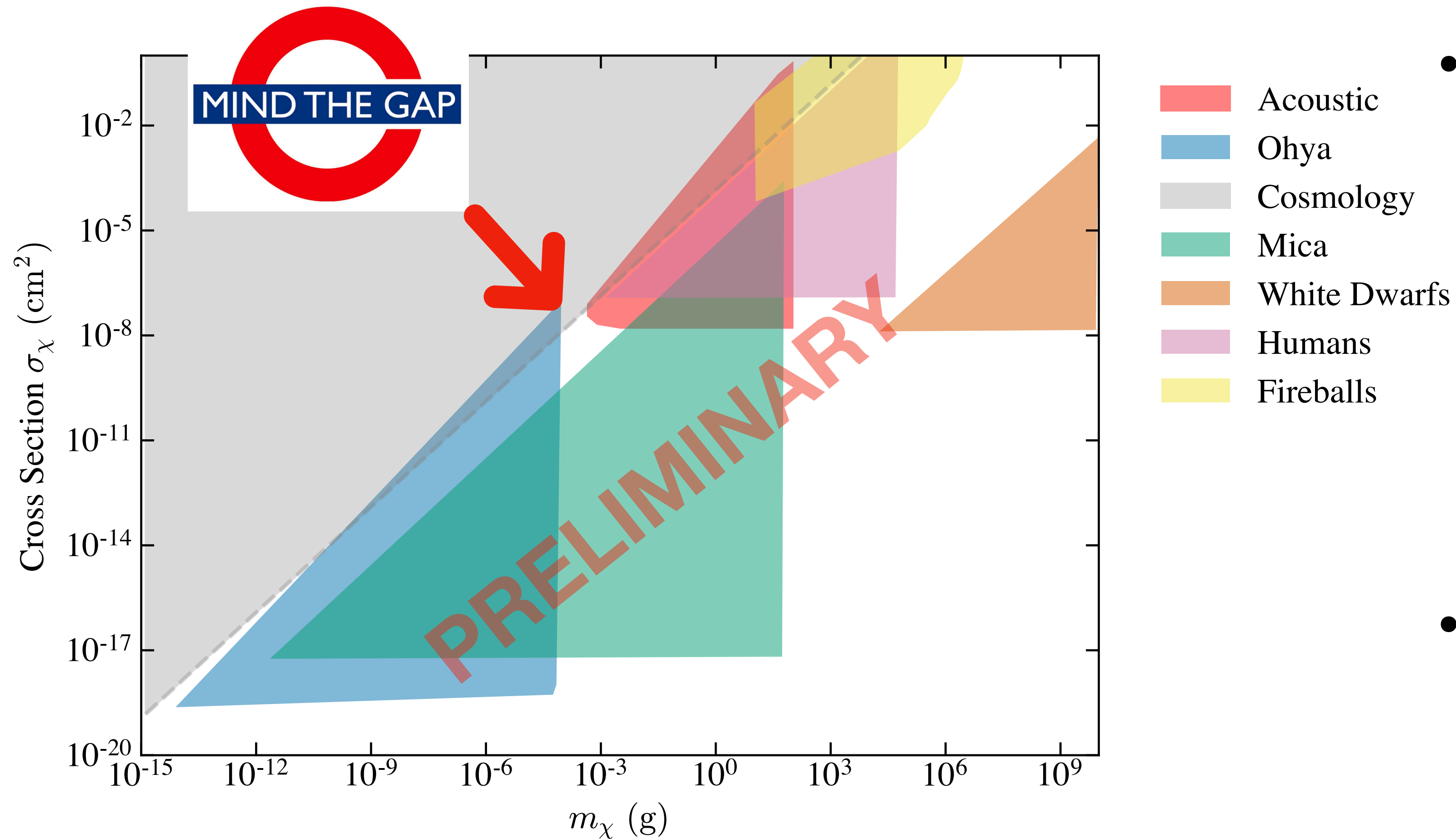
**1.5m**

**1.5m**

**LZ**



# Sensitivities



- With hydrophones available **right now**, bounds lie mostly within already excluded space
- Mica [1] and Humans [2] bounds overlap
- **Better hydrophone sensitivity = better cross section sensitivity - could plug the gap!**

[1] D. M. Jacobs, G. D. Starkman, and B. W. Lynn, Macro dark matter (2015).

[2] J. Singh Sidhu, R. J. Scherrer, and G. Starkman, Death and serious injury from dark matter (2020)



# Punchline

---

Macro dark matter, despite its low number density, could be detected in proposed acoustic neutrino experiments!

**Thank you for listening**

---