<span id="page-0-1"></span><span id="page-0-0"></span>Equation of state of isospin asymmetric QCD with small baryon chemical potentials

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this session:

focus on isospin chemical potential  $\mu_I$ important part of QCD parameter space

• no sign problem





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- no sign problem
- relevant for many physical systems
- dominant for some physical systems











Convenient chemical potential basis for simulations: ("isospin" basis)

$$
\mu_u = \mu_L + \mu_I \qquad \mu_d = \mu_L - \mu_I \qquad \mu_s
$$

 $\mu_I \neq 0$ ,  $\mu_L = \mu_s = 0$  pure isospin chemical pot. – no sign problem





**I** Taylor expansion and  $\lambda$ -extrapolations

#### II Improved computation of expansion coefficients

#### III An application: EoS at non-zero charge chemical potential



#### **I** Taylor expansion and  $\lambda$ -extrapolations

# Taylor expansion around  $\mu_I \neq 0$



Extension to  $\mu_L$ ,  $\mu_s \neq 0$ :

$$
p(T, \vec{\mu}) = p(T, \mu_I, 0, 0) + \sum_{n,m=1}^{\infty} \frac{1}{n! \, m!} \left. \frac{\partial^n \partial^m [p(T, \vec{\mu})]}{\partial \mu_L^n \partial \mu_s^m} \right|_{\mu_L, \mu_s = 0} (\mu_L)^n (\mu_s)^m
$$

Eleading order: only non-zero coefficients are

$$
\chi_2^L(T,\mu_I) \equiv \left. \frac{\partial^2 [p(T,\vec{\mu})]}{\partial \mu_L^2} \right|_{\mu_L,\mu_s=0} \& \quad \chi_2^s(T,\mu_I) \equiv \left. \frac{\partial^2 [p(T,\vec{\mu})]}{\partial \mu_s^2} \right|_{\mu_L,\mu_s=0}
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$$

$$
\blacktriangleright \text{ Generally:} \quad \chi_2^X = \frac{T}{V} \left[ \underbrace{\langle c_{XX} \rangle}_{\text{connected}} + \underbrace{\langle (c_X)^2 \rangle - \langle c_X \rangle^2}_{\text{disconnected}} \right]
$$

with 
$$
c_X = \text{Tr}\Big[M^{-1} \frac{\partial M}{\partial \mu_X}\Big]
$$
  
\n
$$
c_{XX} = \frac{\partial c_X}{\partial \mu_X} = \text{Tr}\Big[M^{-1} \frac{\partial^2 M}{(\partial \mu_X)^2}\Big] + \text{Tr}\Big[M^{-1} \frac{\partial M}{\partial \mu_X} M^{-1} \frac{\partial M}{\partial \mu_X}\Big]
$$



$$
D = \gamma_{\mu}D_{\mu} + m_{ud} + \gamma_0 \tau_3 \mu_I
$$
  
\n
$$
SU_V(2) \longrightarrow U_Q(1) \longrightarrow \emptyset
$$
  
\nexplicit  
\n
$$
\mu_I \neq 0 \qquad \mu_I \geq m_{\pi}/2
$$

- cannot observe spontaneous symmetry breaking in finite  $V$
- low mode in simulations







- cannot observe spontaneous symmetry breaking in finite  $V$
- low mode in simulations







- $\blacktriangleright$  facilitated by improvement program: [ Brandt, Endrődi, Schmalzbauer '18 ]
	- leading order reweighting
	- valence quark improvement for light quarks
- cannot observe spontaneous symmetry breaking in finite  $V$
- low mode in simulations





Up to now: consider 1st derivatives – densities (e.g.:  $n_I = c_I$ )  $O \equiv \textsf{Tr}\Big[ M^{-1} \hat{O} \Big] \qquad c_X \colon \; \hat{O} = \frac{\partial M}{\partial u_N}$  $\frac{\partial M}{\partial \mu_X}$  with  $M = D^{\dagger}(\mu)D(\mu) + \lambda^2$ 

 $\blacktriangleright$  typically: compute the trace stochastically



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ightharpoonup alternative: use singular values  $D^{\dagger}(\mu)D(\mu)\varphi_n = \xi_n^2\varphi_n$ 

$$
O = \sum_{n=0}^{N_{\text{lat}}} \frac{\varphi_n^{\dagger} \hat{O} \varphi_n}{\xi_n^2 + \lambda^2} \approx \sum_{n=0}^{N} \frac{\mathcal{O}_{nn}}{\xi_n^2 + \lambda^2}
$$

with  $\mathcal{O}_{nm} = \varphi_n^{\dagger} \hat{O} \varphi_m$ 

here: can formally set  $\lambda = 0$ 



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 $O =$  $\sum^{\text{N}_{\text{lat}}}$  $n=0$  $\varphi_n^\dagger \hat{O} \varphi_n$  $\frac{\varphi_n^\dagger\hat{O}\varphi_n}{\xi_n^2+\lambda^2} \approx \sum_{n=0}^N$  $n=0$  $\mathcal{O}_{nn}$  $\xi_n^2 + \lambda^2$ 

with 
$$
\mathcal{O}_{nm} = \varphi_n^{\dagger} \hat{O} \varphi_m
$$

 $\rightarrow$  eliminate leading  $\lambda$ -dependence via

here: can formally set  $\lambda = 0$ 

$$
\lim_{\lambda \to 0} \left\langle O \right\rangle = \lim_{\lambda \to 0} \left\langle O - \delta_O^N \right\rangle + \lim_{\lambda \to 0} \left\langle \delta_O^N \right\rangle
$$

$$
\text{with}\quad \delta_O^N = \sum_{n=0}^N \mathcal{O}_{nn}\Big(\frac{1}{\xi_n^2+\lambda^2}-\frac{1}{\xi_n^2}\Big)
$$



Up to now: consider 1st derivatives – densities (e.g.: 
$$
n_I = c_I
$$
)

\n
$$
O \equiv \text{Tr}\left[M^{-1}\hat{O}\right] \qquad c_X: \ \hat{O} = \frac{\partial M}{\partial \mu_X} \quad \text{with} \quad M = D^{\dagger}(\mu)D(\mu) + \lambda^2
$$
\nby the two equations, we have:

\n
$$
V = \sum_{i=1}^{n} \sum_{j=1}^{n} \mu_i \left(\frac{\partial M}{\partial \mu_i}\right) \qquad \text{with} \quad M = D^{\dagger}(\mu)D(\mu) + \lambda^2
$$

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$$

 $\overline{M}$ 

here: can formally set  $\lambda = 0$ 

eliminate leading  $\lambda$ -dependence via ►

$$
\lim_{\lambda \to 0} \langle O \rangle = \lim_{\lambda \to 0} \langle O - \delta_O^N \rangle + 0
$$

$$
\text{with}\quad \delta_O^N = \sum_{n=0}^N \mathcal{O}_{nn}\Big(\frac{1}{\xi_n^2+\lambda^2}-\frac{1}{\xi_n^2}\Big)
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with 
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O_{nm} = \varphi_n^{\dagger} \hat{O} \varphi_m
$$



 $O =$ 

 $\sum^{\text{N}_{\text{lat}}}$  $n=0$ 



Up to now: consider 1st derivatives – densities

\n
$$
(e.g.: n_I = c_I)
$$
\n
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O \equiv \text{Tr}\left[M^{-1}\hat{O}\right] \qquad c_X: \ \hat{O} = \frac{\partial M}{\partial \mu_X} \quad \text{with} \quad M = D^{\dagger}(\mu)D(\mu) + \lambda^2
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with 
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$$
\chi_2^X = \frac{T}{V} \big[ \langle c_{XX} \rangle + \langle (c_X)^2 \rangle - \langle c_X \rangle^2 \big]
$$

► disconnected terms: as above  $\langle (c_X - \delta_{c_X}^N)^2 \rangle - \langle c_X - \delta_{c_X}^N \rangle^2$ 



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two connected terms:

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c_{XX} = \text{Tr}\Big[M^{-1} \frac{\partial^2 M}{(\partial \mu_X)^2}\Big] + \text{Tr}\Big[M^{-1} \frac{\partial M}{\partial \mu_X} M^{-1} \frac{\partial M}{\partial \mu_X}\Big]
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• treat as above



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$$

• treat as above

$$
\begin{array}{lcl} C_{12} & \equiv & \operatorname{Tr}\Big[M^{-1}\hat{O}^{(1)}M^{-1}\hat{O}^{(2)}\Big] \\ & \approx & \left\langle \sum_{n,m=0}^N \frac{\mathcal{O}^{(1)}_{nm}}{\xi_m^2+\lambda^2}\frac{\mathcal{O}^{(2)}_{nm}}{\xi_n^2+\lambda^2} \right\rangle \end{array}
$$



$$
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$$

• treat as above

$$
C_{12} = Tr \left[ M^{-1} \hat{O}^{(1)} M^{-1} \hat{O}^{(2)} \right] \qquad \qquad \frac{\hat{S}}{\hat{S}}_{32} \longrightarrow 0.5
$$
\n
$$
\approx \left\langle \sum_{n,m=0}^{N} \frac{\mathcal{O}_{nm}^{(1)}}{\xi_m^2 + \lambda^2} \frac{\mathcal{O}_{nm}^{(2)}}{\xi_n^2 + \lambda^2} \right\rangle \qquad \qquad \sum_{n=1}^{N} \left\{ \sum_{n=1}^{N} \left[ \oint_{\alpha} \left[ \oint_{\alpha} \phi_n \right] \right] \left[ \oint_{\alpha} \phi_n \right] \right\} \left[ \oint_{\alpha} \phi_n \right] \left[ \oint_{\alpha} \phi
$$

 $0<sub>0</sub>$ 

improvement term:

$$
\delta_{C_{12}}^N = \sum_{n,m=1}^N \mathcal{O}^{(1)}_{nm} \mathcal{O}^{(2)}_{mn} \Big( \frac{1}{\xi_n^2 + \lambda^2} \frac{1}{\xi_m^2 + \lambda^2} - \frac{1}{\xi_n^2} \frac{1}{\xi_m^2} \Big)
$$



$$
\chi_2^X = \frac{T}{V} \big[ \langle c_{XX} \rangle + \langle (c_X)^2 \rangle - \langle c_X \rangle^2 \big]
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$$

• treat as above

$$
C_{12} \equiv \text{Tr}\left[M^{-1}\hat{O}^{(1)}M^{-1}\hat{O}^{(2)}\right] \approx \left\{\sum_{n,m=0}^{N} \frac{\mathcal{O}_{nm}^{(1)}}{\xi_n^2 + \lambda^2} \frac{\mathcal{O}_{nm}^{(2)}}{\xi_n^2 + \lambda^2}\right\} \approx \left\{\begin{array}{c} \frac{\mu_1/m_\pi = 0.77}{T = 132 \text{ MeV}} \\ \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{3}} \\ \frac{1}{\omega} & \frac{1}{\omega} \\ \frac{1}{\omega} & \frac{1}{\omega} \\ \frac{1}{\omega \text{ without improvement}} \\ \frac{1}{\omega} & \frac{1}{\omega} \end{array}\right\}
$$
\n
$$
\approx \left\{\sum_{n,m=0}^{N} \frac{\mathcal{O}_{nm}^{(1)}}{\xi_n^2 + \lambda^2} \right\}
$$

 $10 -$ 

$$
\delta_{C_{12}}^N = \sum_{n,m=1}^N \mathcal{O}^{(1)}_{nm} \mathcal{O}^{(2)}_{mn} \Big( \frac{1}{\xi_n^2 + \lambda^2} \frac{1}{\xi_m^2 + \lambda^2} - \frac{1}{\xi_n^2} \frac{1}{\xi_m^2} \Big)
$$

# $|\lambda$ -extrapolated T<u>aylor coefficents</u> @  $N_t=8$



results for  $\chi_2^L(T,\mu_I)$  using standard  $\lambda$ -extrapolations:



Large uncertainties for  $\chi_2^L(T,\mu_I)$  in the BEC phase!

 $\blacktriangleright \ \chi_2^s(T,\mu_I)$  not affected (no source parameter)





#### II Improved computation of  $\chi_2^L$  $\frac{\iota}{2}(T,\mu_I)$

#### Improved computation method



Observation: equal connected parts in  $\mu_L$  and  $\mu_I$  derivatives

 $c_{LL} = c_{II}$ 

 $\blacktriangleright$  Can we use that somehow?



#### Observation: equal connected parts in  $\mu_L$  and  $\mu_I$  derivatives

 $c_{LL} = c_{II}$ 

**► Can we use that somehow?** 

 $\triangleright$  for EoS computation: model independent spline interpolation of  $n_I$ [ Brandt, Cuteri, Endrődi '22 ]

 $\longrightarrow$  know  $\mu_I$  dependence of  $n_I$ 

compute  $\frac{\partial n_I}{\partial n}$  $\frac{\partial n_I}{\partial \mu_I} = \chi_2^I(T, \mu_I)$  analytically from spline interpolation



#### Observation: equal connected parts in  $\mu_L$  and  $\mu_I$  derivatives

 $c_{II} = c_{II}$ 

 $\rightarrow$  Can we use that somehow?

 $\triangleright$  for EoS computation: model independent spline interpolation of  $n_I$ [ Brandt, Cuteri, Endrődi '22 ]

 $\longrightarrow$  know  $\mu_I$  dependence of  $n_I$ 

compute  $\frac{\partial n_I}{\partial n}$  $\frac{\partial n_I}{\partial \mu_I} = \chi_2^I(T, \mu_I)$  analytically from spline interpolation

**Finally compute**  $\chi_2^L(T, \mu_I)$  using

$$
\chi_2^L(T,\mu_I) = \underbrace{\chi_2^I(T,\mu_I)}_{\lambda=0} + \frac{T}{V} \left[ \left\langle (c_L)^2 \right\rangle - \left\langle c_L \right\rangle^2 - \left\{ \left\langle (c_I)^2 \right\rangle - \left\langle c_I \right\rangle^2 \right\} \right]
$$

 $(\lambda$ -extrapolation for disconnected typically better behaved)



**D** comparison of the methods for  $\chi_2^L(T, \mu_I)$  in BEC:



reduced uncertainties for  $\chi_2^L(T,\mu_I)$  in the BEC phase!

#### Results for Taylor coefficients @  $N_t = 8$



To extend the EoS: interpolate Taylor expansion coefficients as for EoS: [ Brandt, Cuteri, Endrődi '22 ] [ Brandt, Endrődi '16 ] use spline interpolations with Monte-Carlo generated nodepoints

 $\triangleright$  constraint:  $\frac{L}{2}(T,\mu_I)$  &  $\chi_2^s(T,\mu_I)$  vanish at  $T=0$ 



#### Results for Taylor coefficients  $\mathcal{O} N_t = 8$



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#### III An application: EoS at non-zero charge chemical potential

# EoS for early Universe  $\Phi$  large  $l_{\ell}$



early Universe: sum of lepton flavour asymmetries is constrained  $|l_e + l_u + l_\tau| < 0.012$  [ Oldengott, Schwarz '17 ]

but: for large individual  $l_{\ell}$ 

 $\rightarrow$  large  $\mu_Q$  and small  $\mu_B$  &  $\mu_S$  along trajectories



model equation of state – matched to lattice at pure  $\mu_I$ 

[ Vovchenko, et al '20 ]

▶ Taylor expansion from  $\vec{\mu} = 0$  – no pion condensation

[ Middeldorf-Wygas, et al '20 ]

# EoS for early Universe  $\overline{Q}$  large  $l_{\ell}$



early Universe: sum of lepton flavour asymmetries is constrained

 $|l_e + l_u + l_\tau| < 0.012$  [ Oldengott, Schwarz '17 ]

but: for large individual  $l_{\ell}$ 

 $\blacktriangleright$  large  $\mu_Q$  and small  $\mu_B$  &  $\mu_S$  along trajectories

can provide a full lattice EoS:



assumption: BEC phase boundary does not change drastically

#### **Conclusions**



- Simulations at  $\mu_I \neq 0$ : offer a novel expansion point to explore  $(\mu_B, \mu_Q, \mu_S)$  space
- $\blacktriangleright$   $\lambda$ -extrapolations necessary:
	- large uncertainties for  $\chi_2^L(T,\mu_I)$ in BEC phase
- ightharpoonly alternative: computation via  $\chi_2^I(T,\mu_I)$ obtained from spline interpolation of  $n_I$ 
	- $\rightarrow$  improves uncertainties some details still to be understod
- $\blacktriangleright$  application and outlook: compute EoS at non-zero  $\mu_Q$ and in its vicinity

