



Learning Hadron Interactions From Lattice QCD

<u>Lingxiao Wang</u> (王凌霄) RIKEN-iTHEMS

In preparation within HAL QCD collaboration (Takumi Doi, Tetsuo Hatsuda, Yan Lyu)

August 2, LATTICE 2024, the University of Liverpool, UK



Hadron Forces



slides@T.Hatsuda









Importance Sampling

Hybrid MC = MD + Metropolis

Continuum & Thermodynamic Limits

 $a \to 0, L \to \infty$

HAL QCD



K(2011-2019)

Hadrons to Atomic nuclei from Lattice QCD (HAL QCD Collaboration)





Fugaku(2020-)

- S. Aoki, T. M. Doi, E. Itou (Kyoto U.)
- T. Aoyama (ISSP)
- T. Doi, T. Hatsuda, Y. Lyu, L. Wang,
 - R. Yamada, L. Zhang (RIKEN)
- F. Etminan (U. of Birjand)
- Y. Ikeda, N. Ishii, P. Junnarkar, H. Nemura, K. Sasaki(Osaka U.)
- T. Inoue (Nihon U.)
- K. Murakami (TITech)
- K. Murase (Tokyo Metropolitan U.)
- T. Sugiura (Rissho U.)
- H. Tong (U. of Bonn)

HAL QCD Method

Rebuild Potential

N. Ishii, S. Aoki, and T. Hatsuda, Phys. Rev. Lett. 99, 022001 (2007) S. Aoki, T. Hatsuda, and N. Ishii, Prog. Theor. Phys. 123, 89 (2010) Aoki, S., Doi, T., Front. in Phys. 8, 307 (2020) S. Aoki and T. Doi, in Handbook of Nuclear Physics(2023), pp. 1–31



Learning Potential

Inverse Problem



Non-local Potential

Inverse Problem



Toy Model

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Toy Model Separable Potential

Aoki, S., Doi, T., Front. in Phys. 8, 307 (2020)

$$U(\mathbf{r},\mathbf{r}') \equiv \omega \nu(\mathbf{r})\nu(\mathbf{r}'), \quad \nu(\mathbf{r}) \equiv e^{-\mu r}$$

The S-wave solution of the Schrodinger equation with this potential is given exactly by,

$$\phi_k^0(r) = \frac{e^{i\delta_0(k)}}{kr} \left[\sin\{kr + \delta_0(k)\} - \sin\delta_0(k)e^{-\mu r} \left(1 + \frac{r(\mu^2 + k^2)}{2\mu}\right) \right],$$

where,

$$k \cot \delta_0(k) = -\frac{1}{4\mu^2} \left[2\mu(\mu^2 - k^2) - \frac{3\mu^2 + k^2}{4\mu^3} (\mu^2 + k^2)^2 + \frac{(\mu^2 + k^2)^4}{8\pi m\omega} \right]$$

As a numerical example, we take $\mu = 1.0, \omega = -0.017 \mu^4, m = 3.30 \mu, R = 2.5/\mu$

Toy Model Separable Potential



As a numerical example, we take $\mu = 1.0$, $\omega = -0.017\mu^4$, $m = 3.30\mu$, $R = 2.5/\mu$

Warm-Up

Two parameters

$$U_{\theta}(\mathbf{r}, \mathbf{r}') = \omega \exp(-\theta_1 r) \exp(-\theta_2 r')$$



$$\phi_k^0(r) = \frac{e^{i\delta_0(k)}}{kr} \left[\sin\{kr + \delta_0(k)\} - \sin\delta_0(k)e^{-\mu r} \left(1 + \frac{r(\mu^2 + k^2)}{2\mu}\right) \right]$$

A practical set-up for training, $k = [0.01, 1.0], N_k = 20, r = R.$

$$\min_{\theta} \mathscr{L} = \sum_{k} \left[\left(E_{k} - H_{0} \right) \phi_{\mathbf{k}}(r = R) - \int 4\pi dr' r' U_{\theta}(R, r') \phi_{k}(r') \right]^{2}$$

Epoch 0, Loss: 4.375710964202881, Theta1: 0.5099999904632568, Theta2: 0.49000000953674316 Epoch 200, Loss: 0.1396370679140091, Theta1: 0.950462281703949, Theta2: 0.39753425121307373 Epoch 400, Loss: 0.04474024847149849, Theta1: 1.14754319190979, Theta2: 0.510127604007721 Epoch 600, Loss: 0.01694628968834877, Theta1: 1.184393048286438, Theta2: 0.6280280947685242 Epoch 800, Loss: 0.009242076426744461, Theta1: 1.1597602367401123, Theta2: 0.6997004151344299 Epoch 1000, Loss: 0.005497976206243038, Theta1: 1.133785605430603, Theta2: 0.752680778503418 Epoch 1200, Loss: 0.0034021069295704365, Theta1: 1.1116809844970703, Theta2: 0.7952543497085571 Epoch 1400, Loss: 0.0021468503400683403, Theta1: 1.0929081439971924, Theta2: 0.8305171728134155 Epoch 1600, Loss: 0.0013651829212903976, Theta1: 1.0768871307373047, Theta2: 0.8601892590522766 Epoch 1800, Loss: 0.0008674096898175776, Theta1: 1.063176155090332, Theta2: 0.8853644728660583 Epoch 2000, Loss: 0.0005469319876283407, Theta1: 1.0514411926269531, Theta2: 0.9067927598953247 Epoch 2200, Loss: 0.00034017590223811567, Theta1: 1.041424036026001, Theta2: 0.9250178933143616 Epoch 2400, Loss: 0.00020751418196596205, Theta1: 1.0329188108444214, Theta2: 0.9404546618461609 Epoch 2600, Loss: 0.00012344479910098016, Theta1: 1.025755763053894, Theta2: 0.9534343481063843 Epoch 2800, Loss: 7.118881330825388e-05, Theta1: 1.019789695739746, Theta2: 0.9642329216003418 Epoch 3000, Loss: 3.954790372517891e-05, Theta1: 1.0148907899856567, Theta2: 0.9730932712554932 Epoch 3200, Loss: 2.1019024643464945e-05, Theta1: 1.0109381675720215, Theta2: 0.980238676071167 Epoch 3400, Loss: 1.0608729098748881e-05, Theta1: 1.0078167915344238, Theta2: 0.9858796000480652 Epoch 3600, Loss: 5.043078999733552e-06, Theta1: 1.0054134130477905, Theta2: 0.9902217388153076 Enach 3800 Loss: 2.2376170818461105e-06. Theta1: 1.003617763519287, Theta2: 0.993465781211853 Optimised Theta1: 1.0023272037506104 Optimised Theta2: 0.9957966208457947

Non-Local Potential

 $U_{\theta}(\mathbf{r},\mathbf{r}') = \omega f_{\theta}(\mathbf{r},\mathbf{r}'), U_{\mathbf{NN}}(\mathbf{r},\mathbf{r}') \equiv f_{\theta}(\mathbf{r},\mathbf{r}')$



Symmetricly Sharing Parameters

Non-Local Potential

 $U_{\theta}(\mathbf{r},\mathbf{r}') = \omega f_{\theta}(r,r'), U_{\mathbf{NN}}(r,r') \equiv f_{\theta}(r,r')$

$$\min_{\theta} \mathscr{L} = \sum_{r} \sum_{k} \left[\left(E_{k} - H_{0} \right) \phi_{k}(r) - \int 4\pi d \, r' r' U_{\theta}(r, r') \phi_{k}(r') \right]^{2}$$



Non-Local Potential: More Physics Priors

$$U_{\theta}(\mathbf{r},\mathbf{r}') = \omega f_{\theta}(r,r'), U_{\mathsf{NN}}(r,r') \equiv f_{\theta}(r,r')$$

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Non-Local Potential: More Physics Priors





$\Omega_{ccc}\Omega_{ccc}({}^{1}S_{0})$

In preparation within HAL QCD collaboration(Takumi Doi, Tetsuo Hatsuda, Yan Lyu)

Real World

Time-Dependent HAL QCD

N. Ishii, etc., Phys. Lett. B 712, 437 (2012)

Normalized NN correlation function

$$R(t,\vec{r}) \equiv C_{NN}(\vec{r},t) / \left(e^{-m_N t}\right)^2$$



"Time-Dependent" Schrödinger-like Equation

$$\left\{\frac{1}{4m_N}\frac{\partial^2}{\partial t^2} - \frac{\partial}{\partial t} - H_0\right\}R(t,\vec{r}) = \int d^3r' U(\vec{r},\vec{r}')R(t,\vec{r}')$$

Alleviate the Ground State Saturation



Time-Dependent HAL QCD

$$\left\{\frac{1}{4m_N}\frac{\partial^2}{\partial t^2} - \frac{\partial}{\partial t} - H_0\right\}R(t,\vec{r}) = \int d^3r' U(\vec{r},\vec{r}')R(t,\vec{r}')$$

Maximize Likelihood Estimation

$$\min_{\theta} \mathscr{L} = \sum_{t} \left\{ \frac{1}{4m_{N}} R_{tt}(t,r) - R_{t}(t,r) + \frac{1}{m_{N}} R_{r}(t,r) - \int 4\pi r'^{2} dr' U_{\theta}(r,r') R(t,r') \right\}^{2}$$

$$R_{tt}(t,r) \equiv \partial_t^2 R(t,r), R_t(t,r) \equiv \partial_t R(t,r), R_r(t,r) \equiv \nabla^2 R(t,r)$$

Time-Dependent HAL QCD

$$\min_{\theta} \mathscr{L} = \sum_{t} \left\{ \frac{1}{4m_{N}} R_{tt}(t,r) - R_{l}(t,r) + \frac{1}{m_{N}} R_{r}(t,r) - \int 4\pi r'^{2} dr' U_{\theta}(r,r') R(t,r') \right\}^{2}$$

$$\underset{t}{\overset{\mathsf{U}_{\theta}}(r,r') \xrightarrow{\mathsf{BP}}} \operatorname{Residual of}_{\mathsf{Schrödinger Eq.}} \left\{ \begin{array}{c} & & \\$$







$$\begin{aligned} \mathscr{L}_{t} &= \sum_{t} \left\{ \frac{1}{4m_{N}} R2(t,r) - R1(t,r) + \frac{1}{m_{N}} Rr(t,r) - \int dr' U_{\theta}(r,r') R(t,r') \right\} \\ R2 &= R_{t+1} - 2R_{t} + R_{t-1}, R1 = (R_{t+1} - R_{t-1})/2, Rr = \nabla^{2} R(t,r) \\ \int dr' U_{NN}(r,r') R(t,r') &\approx \sum_{r'} \Delta r' U_{\theta}(r,r') R(t,r') \end{aligned}$$

Asymptotic Behaviour as Regulator

$$U_{\theta}(r > 3\mathbf{fm}, r' > 3\mathbf{fm}) \rightarrow 0$$

$$\mathcal{L}_r = \sum_{r=3fm}^{r_{max}} \sum_{r'=3fm}^{r_{max}} U_{\theta}(r, r')^2$$

$$\mathscr{L} \equiv \mathscr{L}_t + \lambda \mathscr{L}_r$$

 $\lambda = 10^8$









$$V_{\theta}(r) \equiv \frac{\sum_{r'} \Delta r' U_{\theta}(r, r') R(t, r')}{R(t, r)}$$

Case study: $\Omega_{ccc}\Omega_{ccc}(^{1}S_{0})$



Phase Shifts

To be calculated...



continuous function



• phase shift (black: using local potential)



same phase shifts up to ~100 MeV for I=1 KN while ~10 MeV for I=0 KN

e.g., Separable Potential

K. Murakami presents applications to Lambda(1405) @Lattice 2024, Aug 2, 2024, 12:55PM

Non-local potential matters!

Summary

Take-home messages

- k-independent and non-local potential!
- Physics prior as regularization!

 $U_{\theta}(r,r')[MeV]$







Outlook

Roadmap

- Rebuild Separable Potential
 - Neural Network Non-Local Potential
 - Exchange symmetry
 - Asympotoic behaviour
- t-HAL QCD method
 - Omega-Omega(s-channel) ✓
 - Non-local potential \checkmark
 - Phase Shifts 🦾

• Next Steps

- Full-t joint learning
- More real cases (B-B, N-B, N-M, N-N, elastic scattering...)



Symmetricly Sharing Parameters

+

Asymptotic Behaviour as Regulator

 $\lim_{\mathbf{r}>R,\mathbf{r}'>R} U(\mathbf{r},\mathbf{r}')\to 0$

Thank you!



Physics-Driven Deep Learning

DEEP-IN

iTHEM。 ² 理化学研究所 数理創造プログラム RIKEN Interdisciplinary Theoretical and Mathematical Sciences Program



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DEEP-IN Working Group

"DEEP learning for INverse problems (DEEP-IN) in Sciences" working group (April 1st, 2024 -)

Lattice Computations

Gert Aarts, Swansea U. Takumi Doi, iTHEMS Andreas Ipp, TU Wien Tetsuo Hatsuda, iTHEMS Yan Lyu, iTHEMS

Heavy-Ion Collisions

Long-Gang Pang, CCNU Shuzhe Shi, THU Kai Zhou, CUHK-ShenZhen

Now mostly physicists -> Future more diverse scientists

BioPhysics: Catherine Beauchemin, iTHEMS Condensed Matter Physics: Steffen Backes, iTHEMS QCD Physics: Kenji Fukushima, UTokyo Nuclear Physics: Haozhao Liang, UTokyo Quantum Computing: Enrico Rinaldi, Quantinuum K.K./iTHEMS

Astrophysics

Márcio Ferreira, Coimbra U. Yuki Fujimoto, INT->iTHEMS Akira Harada, NIT-Ibaraki Zhenyu Zhu, TDLI->RIT



Akinori Tanaka, AIP/iTHEMS Lingxiao Wang, iTHEMS

Physics-Driven Learning for Solving Inverse Problems towards QCD Physics

in preparation [Review]

Gert Aarts¹, Kenji Fukushima², Tetsuo Hatsuda³, Andreas Ipp⁴, Shuzhe Shi⁵, Lingxiao Wang^{3,*}, and Kai Zhou^{6,7}

¹Department of Physics, Swansea University, SA2 8PP, Swansea, United Kingdom
 ²Department of Physics, The University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo, 113-0033, Japan
 ³Interdisciplinary Theoretical and Mathematical Sciences Program (ITHEMS), RIKEN, Wako, Saitama 351-0198, Japan
 ⁴Institute for Theoretical Physics, TU Wien, Wiedner Hauptstraße 8-10/136, A-1040 Vienna, Austria
 ⁵Department of Physics, Tsinghua University, Beijing 100084, China

⁶School of Science and Engineering, The Chinese University of Hong Kong, Shenzhen (CUHK-Shenzhen), Guangdong, 518172, China

⁷Frankfurt Institute for Advanced Studies, Ruth Moufang Strasse 1, D-60438, Frankfurt am Main, Germany *e-mail: lingxiao.wang@riken.jp

ABSTRACT

The integration of deep learning techniques and physics-driven designs is reforming the way we address inverse problems, in which accurate physical properties are extracted from complex data sets. This is particularly relevant for quantum chromodynamics (QCD), the theory of strong interactions, with its inherent limitations in observational data and demanding computational approaches. This perspective highlights advances and potential of physics-driven learning methods, focusing on predictions of physical quantities towards QCD physics, and drawing connections to machine learning(ML). It is shown that the fusion of ML and physics can lead to more efficient and reliable problem-solving strategies. Key ideas of ML, methodology of embedding physics priors, and generative models as inverse modelling of physical probability distributions are introduced. Specific applications cover first-principle lattice calculations, and QCD physics of hadrons, neutron stars, and heavy-ion collisions. These examples provide a structured and concise overview of how incorporating prior knowledge such as symmetry, continuity and equations into deep learning designs can address diverse inverse problems across different physical sciences.

https://sites.google.com/view/deep-in-wg/homepage

Contact at lingxiao.wang@riken.jp

Backups

Scattering



Imaginary Time

N. Ishii, S. Aoki, and T. Hatsuda, Phys. Rev. Lett. 99, 022001 (2007), S. Aoki, T. Hatsuda, and N. Ishii, Prog. Theor. Phys. 123, 89 (2010). N. Ishii, etc.(HAL QCD), Phys. Lett. B 712, 437 (2012)

$$\langle N_1(\mathbf{x}, t) N_2(\mathbf{y}, t) \mathcal{J}_1^{\dagger}(0) \mathcal{J}_2^{\dagger}(0) \rangle$$

$$= \sum_n \langle 0 | N_1(\mathbf{x}) N_2(\mathbf{y}) | n \rangle a_n e^{-E_n t}$$

$$\xrightarrow{t > t^*} \phi(\mathbf{r}, t) = \sum_{n < n^*} b_n \phi_n(\mathbf{r}) e^{-E_n t}$$

 $\phi(\mathbf{r},t)
ightarrow$ 2 PI Kernel

$$(E_k - H_0)\phi_{\mathbf{k}}(\mathbf{r}) = \int d^3 r' U(\mathbf{r}, \mathbf{r}')\phi_{\mathbf{k}}(\mathbf{r}'), \quad r < R$$

Consider the wave function at "interacting region" \longrightarrow Phase shift, Binding energy



Toy Model-II Yukawa Potential

Local potential approximation will give a Schordinger equation,

on,
$$\left(-\frac{\nabla^2}{2m} + V(r)\right)\psi(r) = E\psi(r),$$

where $V(r) = -\alpha \frac{e^{-\mu r}}{r}$,

and α is the coupling(interaction) constant and μ is the mass of the exchanged particle.



Lingxiao Wang(王凌霄)

Toy Model-II



 $V_{\mathbf{NN}}(r) \equiv f_{\theta}(r)$ $\min_{\theta} \mathscr{L} = \sum_{k} \sum_{k} \left[\left(E_{k} - H_{0} \right) \phi_{\mathbf{k}}(r) - V_{\mathbf{NN}}(r) \phi_{\mathbf{k}}(r) \right]^{2}$

