

Liverpool, Lattice2024

Extracting scattering amplitudes from Euclidean correlators

work in collaboration with Nazario Tantalo
based on [arXiv:2407.02069](https://arxiv.org/abs/2407.02069)

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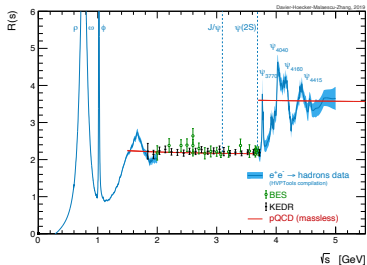
- ▶ Scattering amplitudes are inherently Minkowskian observables. Only Euclidean correlators are calculable in lattice QCD. Analytic continuation is needed to real time. Numerically ill-posed problem.
- ▶ Scattering amplitudes can be extracted from energy levels in large but finite volume. Energy levels can be calculated from Euclidean correlators. More theory needs to be developed every time a new multi-particle threshold is opened.
 - M. Luscher, *Commun. Math. Phys.* **105** (1986), 153-188
 - M. Luscher, *Nucl. Phys. B* **354** (1991), 531-578
 - C. h. Kim, C. T. Sachrajda and S. R. Sharpe, *Nucl. Phys. B* **727** (2005), 218-243
 - M. T. Hansen and S. R. Sharpe, *Phys. Rev. D* **90** (2014) no.11, 116003
 - [...]
- ▶ Approximate scattering amplitudes as a linear combination of Euclidean correlators sampled at discrete times.
 - J. C. A. Barata and K. Fredenhagen, *Commun. Math. Phys.* **138** (1991), 507-520
 - J. Bulava and M. T. Hansen, *Phys. Rev. D* **100** (2019) no.3, 034521

An analogy: spectral densities

M. Hansen, A. Lupo and N. Tantalo, Phys. Rev. D **99**, no.9, 094508 (2019)

William Jay, Lattice24 talk, Mon 17:30 (see also references therein)

Matteo Saccardi, Lattice24 talk, Wed 11:55



M. Davier, A. Hoecker, B. Malaescu and Z. Zhang, EurPhysJ. C80, no.3, 241 (2020).

$$\begin{array}{ccc} \boxed{C(t)} = \int d^3\mathbf{x} \langle j_k(t, \mathbf{x}) j_k(0) \rangle & = & \int_0^\infty dE e^{-tE} \boxed{\rho(E)} \\ \downarrow & & \downarrow \\ \text{Euclidean correlator} & & \text{Spectral density } (\propto R\text{-ratio}) \end{array}$$

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$$\rho(E) = \lim_{\sigma \rightarrow 0^+} \int dE' K_\sigma(E' - E) \rho(E') .$$

The smearing kernel must be smooth with

$$\lim_{\sigma \rightarrow 0} K_\sigma(E) = \delta(E) .$$

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Euclidean correlator
Spectral density (\propto R-ratio)

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2. Approximate of smearing kernel

$$K_\sigma(E) \simeq P_{\sigma, \epsilon}(e^{-\tau E}) = \sum_{n=1}^N w_n^{\sigma, \epsilon} e^{-n\tau E} ,$$

with given precision

$$\|K_\sigma(E) - P_{\sigma, \epsilon}(e^{-\tau E})\| < \epsilon .$$

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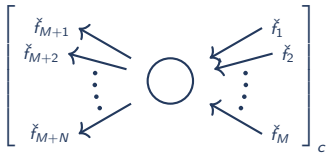
3. Approximation formula

$$\rho(E) = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{n=1}^N w_n^{\sigma, \epsilon} e^{\tau E} C(n\tau)$$

Can we do anything similar for scattering amplitudes?

Scattering amplitudes

Haag-Ruelle
scattering theory



Polynomial approx
of smearing kernel

$$= \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \hat{f}_A^{(*)}(\mathbf{p}_A) \right] \int \left[\prod_A \frac{d\omega_A}{2\pi} \hat{K}_\sigma(\omega, \mathbf{p}) \rho_c(\omega, \mathbf{p}) \right]$$

wave functions of in/out particles
smearing kernel with radius σ
spectral density of $(M+N)$ -pt function

$$= \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} \hat{w}_{n,b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \hat{T}_b(n\tau, \mathbf{p}) C_c(n\tau, \mathbf{p}) \right]$$

coefficients of polynomial approx
known kinematical functions
time-momentum Euclidean $(M+N)$ -pt function

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \right] \hat{\Upsilon}_b(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

Euclidean correlator:

$$\hat{C}_c(s; \mathbf{p}) = \langle \Omega | \hat{\phi}(\mathbf{p}_{M+1}) e^{-s_{M+N} H} \dots \hat{\phi}(\mathbf{p}_{M+N}) e^{-s_M H} \hat{\phi}(\mathbf{p}_M)^\dagger \dots e^{-s_1 H} \hat{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle_c$$

Kinematic function:

$$\hat{\Upsilon}_b(s; \mathbf{p}) = [\Delta(\mathbf{p})]^b \tilde{h}(\Delta(\mathbf{p})) \exp \left\{ \sum_{A=1}^M s_A \sum_{B=1}^A E(\mathbf{p}_B) + \sum_{A=M+1}^{M+N-1} s_A \sum_{B=M+1}^A E(\mathbf{p}_B) \right\}$$

Violation of asympt. energy conservation: $\Delta(\mathbf{p}) = \left\{ \sum_{A=M+1}^{M+N} - \sum_{A=1}^M \right\} E(\mathbf{p}_A)$.

$\tilde{h}(\Delta)$ auxiliary function: smooth, compact support, $\tilde{h}(0) = 1$.

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \right] \hat{\Upsilon}_b(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

Coefficients of polynomial approximation of smearing kernels:

$$K_\sigma(\omega, \Delta) \simeq P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta) = \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \left[\prod_A (e^{-\tau\omega_A})^{n_A} \right] \Delta^b$$

$$\|K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)\| < \epsilon$$

Theorem. For every $r > 0$, two constants A, B_r (independent of ϵ and σ) exist such that

$$\left| \left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c - \text{approx}(\sigma, \epsilon) \right| < A\epsilon + B_r \sigma^r$$

assuming that the wave functions have non-overlapping velocities [not essential].

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \right] \hat{\Upsilon}_b(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

■ Coefficients of polynomial approximation of smearing kernels:

$$K_\sigma(\omega, \Delta) \simeq P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta) = \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \left[\prod_A (e^{-\tau\omega_A})^{n_A} \right] \Delta^b$$

$$\|K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)\| < \epsilon$$

What I am not telling you:

- ▶ What does the smearing kernel look like?
- ▶ What norm do we need to choose?

See paper or backup slides.

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n,b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \right] \hat{\Upsilon}_b(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

- ▶ Smaller $\epsilon \Rightarrow$ better approximation of Haag-Ruelle kernel \Rightarrow larger values of $n \Rightarrow$ larger statistical noise.
- ▶ Smaller $\sigma \Rightarrow$ Haag-Ruelle kernel more peaked \Rightarrow harder to approximate \Rightarrow larger values of $n \Rightarrow$ larger statistical noise.
- ▶ Also recall: $\Upsilon_h(n\tau; \mathbf{p})$ increases exponentially with n .
- ▶ Optimization problem: smaller ϵ and σ means larger statistical errors, larger ϵ and σ means larger systematic error. One could design a strategy based on HLT to minimize total error:

$$A[w] = \|K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)\|^2$$

$$B[w] = \sum_{n, b, n', b'} w_{n,b}^{\sigma, \epsilon} \langle\langle C_{n,b} C_{n',b'} \rangle\rangle_c w_{n',b'}^{\sigma, \epsilon}$$

Approximation formula

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- ▶ A finite-volume estimator is obtained trivially by replacing $\int \frac{d^3 \mathbf{p}_A}{(2\pi)^3}$ with $\frac{1}{L^3} \sum_{\mathbf{p}_A}$.
If coefficients $w_{n,b}$ are kept fixed as the volume is varied, then the $L \rightarrow +\infty$ limit is approached exponentially fast. Having Schwartz wave functions is essential for this step.
- ▶ The continuum limit of the estimator can be understood in terms of Symanzik effective theory.
- ▶ In this approach, the $L \rightarrow \infty$ and $a \rightarrow 0$ limits must be taken before the $\epsilon \rightarrow 0$ and $\sigma \rightarrow 0$ limits. In particular τ cannot be identified with the lattice spacing. For the opposite approach, see Barata and Fredenhagen.

Conclusions and outlook

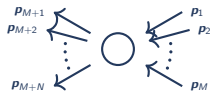
- ▶ We have derived an approximation for scattering amplitudes as a linear combination of Euclidean correlators sampled at discrete times.
- ▶ This formula provides the blueprints for a potentially viable numerical strategy.
- ▶ Our approximation can be calculated from finite-volume correlators and the infinite-volume limit is approached exponentially fast.
- ▶ Whether statistical and systematic errors are under control in typical QCD simulations remains to be seen.
- ▶ Recent algorithmic methods (e.g. Hansen-Lupo-Tantalo), which have been successful in approximations of spectral densities, can be adapted to this problem.
- ▶ The class of operators used to approximate asymptotic states can be generalized by relaxing the constraint that $\tilde{f}^t(\rho)$ must have compact support. This may make the numerics easier.

Backup slides

Talk given at CERN workshop, July 2024

Introduction

- ▶ How do we calculate hadron scattering amplitude from Quantum Chromodynamics? In principle...



A Feynman diagram showing a central circle representing a scattering process. On the left side, there are three outgoing lines labeled p_{M+1} , p_{M+2} , and p_{M+N} , with vertical dots between p_{M+2} and p_{M+N} . On the right side, there are three incoming lines labeled p_1 , p_2 , and p_M , with vertical dots between p_2 and p_M .

$$\propto \lim_{p_0 \rightarrow \pm E(\mathbf{p})} \left[\prod_A (p_A^2 - m^2) \right] \langle \Omega | T \tilde{\phi}(p_{M+1}) \cdots \tilde{\phi}(p_{M+N}) \tilde{\phi}(p_M)^\dagger \cdots \tilde{\phi}(p_1)^\dagger | \Omega \rangle$$

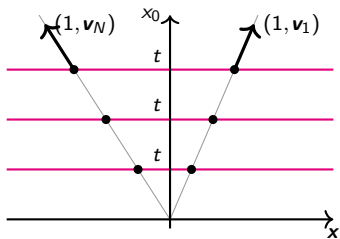
- ▶ Numerical lattice QCD is the only known tool which allows the calculation of observables in QCD at the nonperturbative level.
- ▶ Only Euclidean correlators are calculable in lattice QCD. Analytic continuation is needed to real time. Numerically ill-posed problem.
- ▶ Find another way...

- ▶ Scattering amplitudes can be extracted from energy levels in large but finite volume. Energy levels can be calculated from Eucliden correlators. More theory needs to be developed every time a new multi-particle threshold is opened.
M. Luscher, *Commun. Math. Phys.* **105** (1986), 153-188
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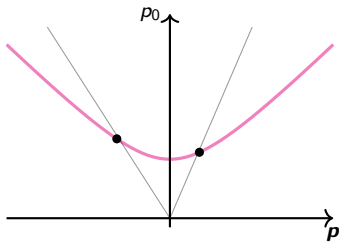
Outlook

Theoretical background

Haag-Ruelle scattering theory

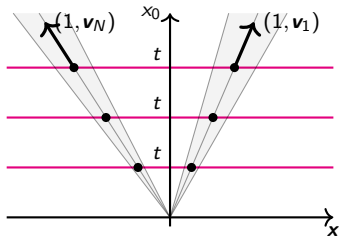


► Black lines = classical trajectories.

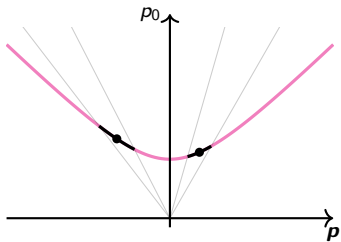


► Pink line: $p^2 = m^2$

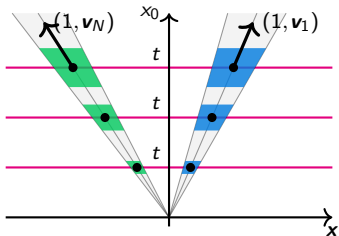
► Black dots: energy-momentum of particle.



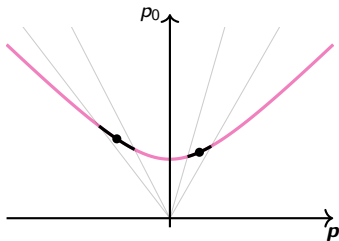
- ▶ Black lines = classical trajectories.
- ▶ Allow velocity indetermination.
- ▶ Gray regions = cones of classical trajectories.



- ▶ Pink line: $p^2 = m^2$
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- ▶ Allow momentum indetermination.
- ▶ Black lines: allowed values for energy-momentum.

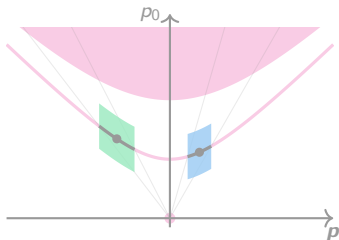
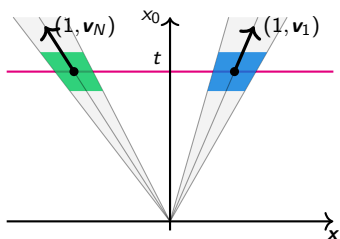


- ▶ Black lines = classical trajectories.
- ▶ Allow velocity indetermination.
- ▶ Gray regions = cones of classical trajectories.
- ▶ Green/blue regions = position of particle at time t .



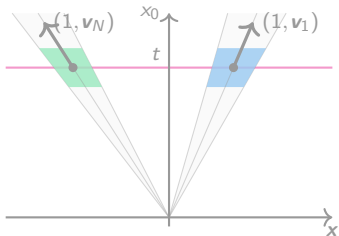
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$$|\Psi_{\text{out}}(t)\rangle = \int d^4 x_N f_N^t(x_N) \phi(x_N)^\dagger \cdots \int d^4 x_1 f_1^t(x_1) \phi(x_1)^\dagger |\Omega\rangle$$

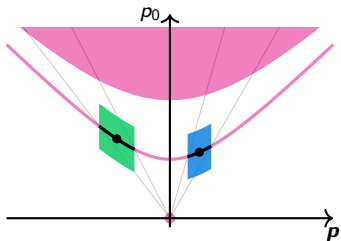


- ▶ Gray regions = cones of classical trajectories.
- ▶ Green/blue regions scale with t .
- ▶ $f_A^t(x)$ is localized in green/blue regions.

$$|\Psi_{\text{out}}(t)\rangle = \int \frac{d^4 p_N}{(2\pi)^4} \tilde{f}_N^t(p_N) \tilde{\phi}(p_N)^\dagger \cdots \int \frac{d^4 p_1}{(2\pi)^4} \tilde{f}_1^t(p_1) \tilde{\phi}(p_1)^\dagger |\Omega\rangle$$

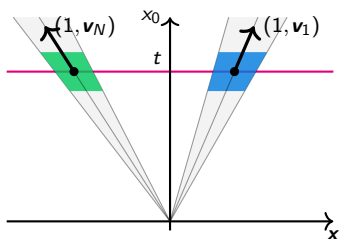


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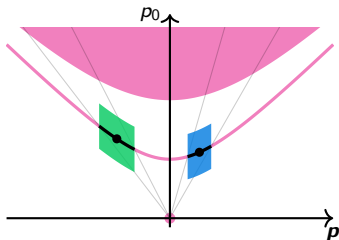


- ▶ Pink regions = spectrum of P .
- ▶ Green/blue regions intersect spectrum of P on 1-particle mass shell.
- ▶ $\tilde{f}_A^t(p)$ has support in green/blue regions.

$$|\Psi_{\text{out}}(t)\rangle = \int \frac{d^4 p_N}{(2\pi)^4} \tilde{f}_N^t(p_N) \tilde{\phi}(p_N)^\dagger \cdots \int \frac{d^4 p_1}{(2\pi)^4} \tilde{f}_1^t(p_1) \tilde{\phi}(p_1)^\dagger |\Omega\rangle$$

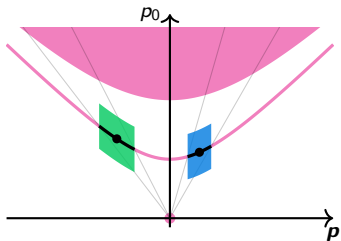
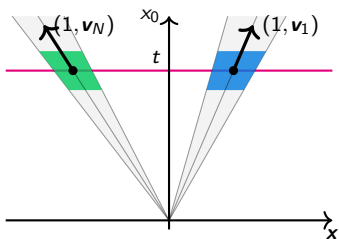


- ▶ Gray regions = cones of classical trajectories.
- ▶ Green/blue regions scale with t .
- ▶ $f_A^t(x)$ is localized in green/blue regions.
- ▶ Interaction between particles decreases with t .



- ▶ Pink regions = spectrum of P .
- ▶ Green/blue regions intersect spectrum of P on 1-particle mass shell.
- ▶ $\tilde{f}_A^t(p)$ has support in green/blue regions.

$$|\Psi_{\text{out}}(t)\rangle = \int \frac{d^4 p_N}{(2\pi)^4} \tilde{f}_N^t(p_N) \tilde{\phi}(p_N)^\dagger \cdots \int \frac{d^4 p_1}{(2\pi)^4} \tilde{f}_1^t(p_1) \tilde{\phi}(p_1)^\dagger |\Omega\rangle$$



$$\tilde{f}_A^t(\mathbf{p}) = e^{it[p_0 - E(\mathbf{p})]} \zeta_A(p_0 - E(\mathbf{p})) \check{f}_A(\mathbf{p})$$

- ▶ $\check{f}_A(\mathbf{p})$ = asymptotic particle wave function and $E(\mathbf{p}) = \sqrt{m^2 + \mathbf{p}^2}$.
- ▶ $\zeta_A(\omega)$ cuts off multi-particle states. $\zeta_A(\omega)$ smooth and compact support, $\zeta_A(0) = 1$.
- ▶ Support of $\tilde{f}_A^t(\mathbf{p})$ intersects spectrum of P only on 1-particle mass shell.

Haag-Ruelle scattering theory

$$|\Psi_{\text{out}}(t)\rangle = \prod_A \int \frac{d^4 p_A}{(2\pi)^4} \tilde{f}_A^t(p_A) \tilde{\phi}(p_A)^\dagger |\Omega\rangle$$

$$\stackrel{t \rightarrow \pm\infty}{=} \prod_A \int \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A(\mathbf{p}_A) a_{\text{out}}^\dagger(\mathbf{p}_A) |\Omega\rangle + O(|t|^{-\infty})$$

- ▶ $\tilde{f}_A^t(\mathbf{p}) = e^{it[p_0 - E(\mathbf{p})]} \zeta_A(p_0 - E(\mathbf{p})) \check{f}_A(\mathbf{p})$
- ▶ Error is $O(|t|^{-\infty})$ for non-overlapping velocities, otherwise $O(|t|^{-1/2})$.
- ▶ $a_{\text{out}}^\dagger(\mathbf{p})$ are standard creation operators:

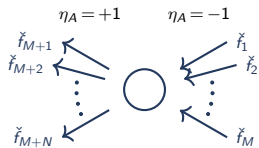
$$[a_{\text{out}}(\mathbf{p}), a_{\text{out}}^\dagger(\mathbf{p}')] = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') \quad [a_{\text{out}}(\mathbf{p}), a_{\text{out}}(\mathbf{p}')] = 0$$

$$[\mathbf{P}, a_{\text{out}}^\dagger(\mathbf{p})] = \mathbf{p} a_{\text{out}}^\dagger(\mathbf{p}) \quad [H, a_{\text{out}}^\dagger(\mathbf{p})] = E(\mathbf{p}) a_{\text{out}}^\dagger(\mathbf{p})$$

Approximation formula for scattering amplitudes

Rough sketch of derivation

Scattering amplitude



$$= \lim_{t \rightarrow +\infty} \langle \Psi_{\text{out}}(t) | \Psi_{\text{in}}(-t) \rangle$$

Scattering amplitude

$$= \lim_{t \rightarrow +\infty} \langle \Psi_{\text{out}}(t) | \Psi_{\text{in}}(-t) \rangle$$

$$= \lim_{t \rightarrow +\infty} \int \left[\prod_A \frac{d^4 p_A}{(2\pi)^4} \check{f}_A^{(*)}(\mathbf{p}_A) \zeta_A^{(*)}(p_A^0 - E(\mathbf{p}_A)) \right] e^{it \sum_A \eta_A [p_A^0 - E(\mathbf{p}_A)]}$$

$$\times \langle \Omega | \tilde{\phi}(\mathbf{p}_{M+1}) \cdots \tilde{\phi}(\mathbf{p}_{M+N}) \tilde{\phi}(\mathbf{p}_M)^\dagger \cdots \tilde{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle$$

Scattering amplitude

$$= \lim_{t \rightarrow +\infty} \langle \Psi_{\text{out}}(t) | \Psi_{\text{in}}(-t) \rangle$$

$$= \lim_{t \rightarrow +\infty} \int \left[\prod_A \frac{d^4 p_A}{(2\pi)^4} \check{f}_A^{(*)}(\mathbf{p}_A) \zeta_A^{(*)}(\mathbf{p}_A^0 - E(\mathbf{p}_A)) \right] e^{it \sum_A \eta_A [p_A^0 - E(\mathbf{p}_A)]}$$

$$\times \langle \Omega | \tilde{\phi}(\mathbf{p}_{M+1}) \cdots \tilde{\phi}(\mathbf{p}_{M+N}) \tilde{\phi}(\mathbf{p}_M)^\dagger \cdots \tilde{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle$$

- ▶ Wildly oscillating phase for $t \rightarrow +\infty$.
- ▶ Not good for numerics.
- ▶ Cancellation of regions with $\sum_A \eta_A [p_A^0 - E(\mathbf{p}_A)] \neq 0$.
- ▶ Can we achieve the same effect in a different way? Some mathematical trickery...

Scattering amplitude

$$= \lim_{t \rightarrow +\infty} \langle \Psi_{\text{out}}(t) | \Psi_{\text{in}}(-t) \rangle$$

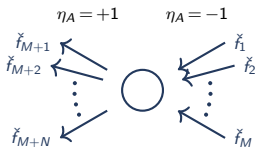
Introduce two auxiliary functions:

- ▶ $\Phi(t)$ Schwartz with unit integral and closed support in $(0, +\infty)$;
- ▶ $h(t)$ Schwartz with unit integral.

$$\lim_{\sigma \rightarrow 0^+} \int dt ds \Phi(t) h(s) \langle \Psi_{\text{out}}\left(\frac{t}{2\sigma} - s\right) | \Psi_{\text{in}}\left(-\frac{t}{2\sigma} - s\right) \rangle$$

$$= \int ds h(s) \int_0^{+\infty} dt \Phi(t) \lim_{\sigma \rightarrow 0^+} \langle \Psi_{\text{out}}\left(\frac{t}{2\sigma} - s\right) | \Psi_{\text{in}}\left(-\frac{t}{2\sigma} - s\right) \rangle = \langle \Psi_{\text{out}}(+\infty) | \Psi_{\text{in}}(-\infty) \rangle$$

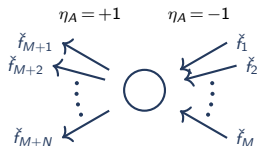
Scattering amplitude (2)



$$= \lim_{\sigma \rightarrow 0^+} \int dt ds \Phi(t) h(s) \langle \Psi_{\text{out}} \left(\frac{t}{2\sigma} - s \right) | \Psi_{\text{in}} \left(-\frac{t}{2\sigma} - s \right) \rangle$$

$$= \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^4 p_A}{(2\pi)^4} \check{f}_A^{(*)}(\mathbf{p}_A) \zeta_A^{(*)}(p_A^0 - E(\mathbf{p}_A)) \right] \tilde{h} \left(\sum_A \eta_A E(\mathbf{p}_A) \right) \tilde{\Phi} \left(\frac{1}{\sigma} \sum_A \eta_A [p_A^0 - E(\mathbf{p}_A)] \right) \\ \times \langle \Omega | \tilde{\phi}(\mathbf{p}_{M+1}) \cdots \tilde{\phi}(\mathbf{p}_{M+N}) \tilde{\phi}(\mathbf{p}_M)^\dagger \cdots \tilde{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle$$

Scattering amplitude (2)



$$\begin{aligned}
 &= \lim_{\sigma \rightarrow 0^+} \int dt ds \Phi(t) h(s) \langle \Psi_{\text{out}} \left(\frac{t}{2\sigma} - s \right) | \Psi_{\text{in}} \left(-\frac{t}{2\sigma} - s \right) \rangle \\
 &= \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^4 p_A}{(2\pi)^4} \check{f}_A^{(*)}(\mathbf{p}_A) \zeta_A^{(*)}(p_A^0 - E(\mathbf{p}_A)) \right] \tilde{h} \left(\sum_A \eta_A E(\mathbf{p}_A) \right) \tilde{\Phi} \left(\frac{1}{\sigma} \sum_A \eta_A [p_A^0 - E(\mathbf{p}_A)] \right) \\
 &\quad \times \langle \Omega | \tilde{\phi}(\mathbf{p}_{M+1}) \cdots \tilde{\phi}(\mathbf{p}_{M+N}) \tilde{\phi}(\mathbf{p}_M)^\dagger \cdots \tilde{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle
 \end{aligned}$$

$\tilde{\Phi}$ regularizes the wildly-oscillating phase factor and selects the desired time-ordering. **It must be complex!**

$\tilde{h}(\Delta)$ can be chosen with compact and arbitrarily narrow support around $\Delta = 0$. It cuts away contributions characterized by non-zero violations of the asymptotic energy conservation.

Wightman function in momentum space \simeq spectral density.

Wightman function \simeq spectral density

$$\langle \Omega | \tilde{\phi}(p_{M+1}) \tilde{\phi}(p_{M+2}) \cdots \tilde{\phi}(p_{M+N}) \tilde{\phi}(p_M)^\dagger \cdots \tilde{\phi}(p_2)^\dagger \tilde{\phi}(p_1)^\dagger | \Omega \rangle$$

Wightman function \simeq spectral density

$$\langle \Omega | \tilde{\phi}(p_{M+1}) \uparrow^{\varepsilon_{M+1}=p_{M+1}^0} \tilde{\phi}(p_{M+2}) \uparrow^{\varepsilon_{M+2}=p_{M+1}^0+p_{M+2}^0} \cdots \tilde{\phi}(p_{M+N}) \uparrow^{\varepsilon_M=p_1^0+\cdots+p_M^0} \tilde{\phi}(p_M)^\dagger \uparrow^{\varepsilon_{M-1}=p_1^0+\cdots+p_{M-1}^0} \cdots \tilde{\phi}(p_2)^\dagger \uparrow^{\varepsilon_2=p_1^0+p_2^0} \tilde{\phi}(p_1)^\dagger | \Omega \rangle$$

Wightman function \simeq spectral density

$$\langle \Omega | \tilde{\phi}(\mathbf{p}_{M+1}) \uparrow^{\mathcal{E}_{M+1} = p_{M+1}^0} \tilde{\phi}(\mathbf{p}_{M+2}) \uparrow^{\mathcal{E}_{M+2} = p_{M+1}^0 + p_{M+2}^0} \cdots \tilde{\phi}(\mathbf{p}_{M+N}) \uparrow^{\mathcal{E}_M = p_1^0 + \cdots + p_M^0} \tilde{\phi}(\mathbf{p}_M)^\dagger \cdots \tilde{\phi}(\mathbf{p}_2)^\dagger \uparrow^{\mathcal{E}_2 = p_1^0 + p_2^0} \tilde{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle$$

$$= 2\pi\delta(\mathcal{E}_{M+N} - \mathcal{E}_M)$$

$$\times \langle \Omega | \hat{\phi}(\mathbf{p}_{M+1}) 2\pi\delta(H - \mathcal{E}_{M+1}) \cdots \hat{\phi}(\mathbf{p}_{M+N}) 2\pi\delta(H - \mathcal{E}_M) \hat{\phi}(\mathbf{p}_M) \cdots 2\pi\delta(H - \mathcal{E}_1) \hat{\phi}(\mathbf{p}_1) | \Omega \rangle$$

$$\text{definitions: } \hat{\phi}(\mathbf{p}) = \int d^3\mathbf{x} e^{-i\mathbf{p}\mathbf{x}} \phi(0, \mathbf{x})$$

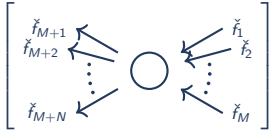
$$\omega_A = \mathcal{E}_A - [\mathcal{E}_A]_{\text{on-shell}}$$

$$= 2\pi\delta(\mathcal{E}_{M+N} - \mathcal{E}_M) \rho(\omega, \mathbf{p})$$

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] \tilde{h}(\Delta(\mathbf{p})) \\
 \times \int \left[\prod_A \frac{d\omega_A}{2\pi} \right] K_\sigma(\omega, \Delta(\mathbf{p})) \rho_c(\omega, \mathbf{p})$$

Approximation formula



$$\left[\begin{array}{c} \check{f}_{M+1}^* \\ \check{f}_{M+2}^* \\ \vdots \\ \check{f}_{M+N}^* \end{array} \leftarrow \bigcirc \rightarrow \begin{array}{c} \check{f}_1^* \\ \check{f}_2^* \\ \vdots \\ \check{f}_M^* \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] \tilde{h}(\Delta(\mathbf{p})) \\ \times \int \left[\prod_A \frac{d\omega_A}{2\pi} \right] K_\sigma(\omega, \Delta(\mathbf{p})) \rho_c(\omega, \mathbf{p})$$

Haag-Ruelle kernel $K_\sigma(\omega, \Delta)$ smears the spectral density in the energy variable ω . The parameter σ plays the role of the smearing radius.

$$K_\sigma(\omega, \Delta) = \tilde{\Phi} \left(\frac{2\omega_M - \Delta}{2\sigma} \right) \zeta_1(\omega_1) \left[\prod_{A=2}^{M-1} \zeta_A(\omega_A - \omega_{A-1}) \right] \zeta_M(\omega_M - \omega_{M-1}) \\ \times \zeta_{M+1}^*(\omega_{M+1}) \left[\prod_{A=M+2}^{M+N-1} \zeta_A^*(\omega_A - \omega_{A-1}) \right] \zeta_{M+N}^*(\omega_M - \omega_{M+N-1} - \Delta)$$

Violation of asymptotic energy conservation: $\Delta(\mathbf{p}) = \sum_A \eta_A E(\mathbf{p}_A)$.

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] \tilde{h}(\Delta(\mathbf{p})) \\
 \times \int \left[\prod_A \frac{d\omega_A}{2\pi} \right] K_\sigma(\omega, \Delta(\mathbf{p})) \rho_c(\omega, \mathbf{p})$$

Approximation is obtained by replacing the Haag-Ruelle kernel with a polynomial in the variables $e^{-\tau\omega}$ and Δ :

$$K_\sigma(\omega, \Delta) \rightarrow P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta) = \sum_{n_1, n_2, \dots \geq 1} \sum_{b \geq 0} w_{n, b}^{\sigma, \epsilon} \left[\prod_A (e^{-\tau\omega_A})^{n_A} \right] \Delta^b$$

$$\|K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)\|_{???} < \epsilon$$

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] \tilde{h}(\Delta(\mathbf{p})) \\
 \times \int \left[\prod_A \frac{d\omega_A}{2\pi} \right] K_\sigma(\omega, \Delta(\mathbf{p})) \rho_c(\omega, \mathbf{p})$$

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$$\|K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)\|_{???} < \epsilon$$

Integrating $P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)$ against the spectral density yields the Euclidean correlator!

Approximation formula

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] \\
 \times \sum_{n_1, n_2, \dots \geq 1} \sum_{b \geq 0} w_{n, b}^{\sigma, \epsilon} [\Delta(\mathbf{p})]^b \Upsilon_h(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

Euclidean correlator:

$$\hat{C}_c(s; \mathbf{p}) = \langle \Omega | \hat{\phi}(\mathbf{p}_{M+1}) e^{-s_{M+N} H} \dots \hat{\phi}(\mathbf{p}_{M+N}) e^{-s_M H} \hat{\phi}(\mathbf{p}_M)^\dagger \dots e^{-s_1 H} \hat{\phi}(\mathbf{p}_1)^\dagger | \Omega \rangle_c$$

Kinematic function:

$$\Upsilon_h(s; \mathbf{p}) = \tilde{h}(\Delta(\mathbf{p})) \exp \left\{ \sum_{A=1}^M s_A \sum_{B=1}^A E(\mathbf{p}_B) + \sum_{A=M+1}^{M+N-1} s_A \sum_{B=M+1}^A E(\mathbf{p}_B) \right\}$$

Which norm?

$$\sum_{\substack{\|\alpha\|_1 = \mathfrak{N}_\omega \\ 0 \leq b \leq \mathfrak{N}_p}} \bar{\Delta}^b \int_{\mathbb{K}} \left[\prod_{A=1}^{M+N-1} \frac{d\omega_A}{2\pi} \right] d\Delta e^{\tau \sum_A \omega_A} \left| D_\omega^\alpha \partial_\Delta^b [K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)] \right|^2 < \epsilon^2$$


- ▶ One can choose some linear combinations of weighted L^2 norm for various derivatives.
- ▶ The integration domain \mathbb{K} is completely determined by kinematics.
- ▶ The number of derivatives that one needs to control (\mathfrak{N}_ω , \mathfrak{N}_p) depend on how singular the spectral density is.
- ▶ The l.h.s. is a quadratic function of the polynomial coefficients $w_{n,b}^{\sigma, \epsilon}$. Minimizing the l.h.s. can be done by solving a system of linear equations.
- ▶ Some speculative argument suggests $\mathfrak{N}_\omega = M + N$ and $\mathfrak{N}_p = 0$. We need to understand this better...

Summary

$$\sum_{\substack{\|\alpha\|_1 = \mathfrak{N}_\omega \\ 0 \leq b \leq \mathfrak{N}_p}} \bar{\Delta}^b \int_{\mathbb{K}} \left[\prod_{A=1}^{M+N-1} \frac{d\omega_A}{2\pi} \right] d\Delta e^{\tau \sum_A \omega_A} \left| D_\omega^\alpha \partial_\Delta^b [K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta)] \right|^2 < \epsilon^2$$

$$\text{approx}(\sigma, \epsilon) = \sum_{\substack{n_1, n_2, \dots \geq 1 \\ b \geq 0}} w_{n, b}^{\sigma, \epsilon} \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] [\Delta(\mathbf{p})]^b \Upsilon_h(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

Theorem. For every $r > 0$, two constants A, B_r (independent of ϵ and σ) exist such that

$$\left| \left(\text{Diagram} \right) - \text{approx}(\sigma, \epsilon) \right| < A\epsilon + B_r \sigma^r$$


assuming that the wave functions have non-overlapping velocities [not essential].

Approximation formula for scattering amplitudes

How can we use it?

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{n_1, n_2, \dots \geq 1} \sum_{b \geq 0} w_{n,b}^{\sigma, \epsilon} C_{n,b}$$

$$C_{n,b} = \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] [\Delta(\mathbf{p})]^b \Upsilon_h(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

- ▶ Smaller $\epsilon \Rightarrow$ better approximation of Haag-Ruelle kernel \Rightarrow larger values of $n \Rightarrow$ larger statistical noise.
- ▶ Smaller $\sigma \Rightarrow$ Haag-Ruelle kernel more peaked \Rightarrow harder to approximate \Rightarrow larger values of $n \Rightarrow$ larger statistical noise.
- ▶ Also recall: $\Upsilon_h(n\tau; \mathbf{p})$ increases exponentially with n .
- ▶ Optimization problem: smaller ϵ and σ means larger statistical errors, larger ϵ and σ means larger systematic error. One could design a strategy based on HLT to minimize total error:

$$A[w] = \left\| K_\sigma(\omega, \Delta) - P_{\sigma, \epsilon}(e^{-\tau\omega}, \Delta) \right\|_{???}^2$$

$$B[w] = \sum_{n, b, n', b'} w_{n,b}^{\sigma, \epsilon} \langle \langle C_{n,b} C_{n', b'} \rangle \rangle_c w_{n', b'}^{\sigma, \epsilon}$$

$$\left[\begin{array}{c} \check{f}_{M+1} \\ \check{f}_{M+2} \\ \vdots \\ \check{f}_{M+N} \end{array} \leftarrow \bigcirc \leftarrow \begin{array}{c} \check{f}_1 \\ \check{f}_2 \\ \vdots \\ \check{f}_M \end{array} \right]_c = \lim_{\sigma \rightarrow 0^+} \lim_{\epsilon \rightarrow 0^+} \sum_{n_1, n_2, \dots \geq 1} \sum_{b \geq 0} w_{n,b}^{\sigma, \epsilon} C_{n,b}$$

$$C_{n,b} = \int \left[\prod_A \frac{d^3 \mathbf{p}_A}{(2\pi)^3} \check{f}_A^{(*)}(\mathbf{p}_A) \right] [\Delta(\mathbf{p})]^b \gamma_h(n\tau; \mathbf{p}) \hat{C}_c(n\tau; \mathbf{p})$$

- ▶ A finite-volume estimator is obtained trivially by replacing $\int \frac{d^3 \mathbf{p}_A}{(2\pi)^3}$ with $\frac{1}{L^3} \sum_{\mathbf{p}_A}$.
If coefficients $w_{n,b}$ are kept fixed as the volume is varied, then the $L \rightarrow +\infty$ limit is approached exponentially fast. Having Schwartz wave functions is essential for this step.
- ▶ The continuum limit of the estimator can be understood in terms of Symanzik effective theory.
- ▶ In this approach, the $L \rightarrow \infty$ and $a \rightarrow 0$ limits must be taken before the $\epsilon \rightarrow 0$ and $\sigma \rightarrow 0$ limits. In particular τ cannot be identified with the lattice spacing. For the opposite approach, see Barata and Fredenhagen.

Conclusions and outlook

- ▶ We have derived an approximation for scattering amplitudes as a linear combination of Euclidean correlators sampled at discrete times.
- ▶ This formula provides the blueprints for a potentially viable numerical strategy.
- ▶ Our approximation can be calculated from finite-volume correlators and the infinite-volume limit is approached exponentially fast.
- ▶ Whether statistical and systematic errors are under control in typical QCD simulations remains to be seen.
- ▶ Recent algorithmic methods (e.g. Hansen-Lupo-Tantalo), which have been successful in approximations of spectral densities, can be adapted to this problem.
- ▶ The class of operators used to approximate asymptotic states can be generalized by relaxing the constraint that $\tilde{f}^t(\rho)$ must have compact support. This may make the numerics easier.