#### Computation of window quantities in a LO-HVP  $\boldsymbol{\mu}$

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> which includes both statistical and systematic variations. The weight  $w_{ia}$  corresponds to a systematics variation, whose flat weighted component is labelled by  $i$ , while its AIC weighted component is labelled by  $a$ . The statistical variations are assumed to follow a normal distribution, i.e.  $\mathcal{N}(y, \sigma; Y)$  is a normal PDF with mean y and standard deviation  $\sigma$ .

**Central value and error.** The central value of  $Y$  is defined by the median of the constructed PDF. The lower and upper total errors are defined by quantiles of the corresponding cumulative distribution function (CDF)

• Standard  $1\sigma$ :  $15.87\%$  and  $84.13\%$ .

• Half of standard  $2\sigma$ :  $2.28\%$  and  $97.72\%$ .

Combining distributions. In order to obtain the physical value of an observable Y, the physical values of  $w_0$  and  $M_{ss}$ are included as inputs using the procedure<sup> $a$ </sup>:

- 1. First we perform the analysis for  $Y$  at two fixed values of  $w_0$  and two fixed values of  $M_{ss}$ , given by the edges of the central one-sigma bands of their distributions:  $Y_k$  with  $k=1,\ldots 4.$
- 2. Take a random selection for the systematic ingredients shared by  $Y, w_0$  and  $M_{ss}$ .
- 3. Make a random selection for the remaining independent ingredients with a probability given by

 $P(i,a) = \frac{1}{\sum_{i=1}^{n} a_i}$  $_b$   $w_{ib}\sum_j 1$ ,

independently for  $w_0$  and  $M_{ss}$ .

- 4. Perform a bilinear interpolation of the fit values and their corresponding  $\chi^2$  values from the  $Y_k$  obtained at fixed values of  $w_0$  and  $M_{ss}$  to the sampled values.
- 5. Compute the weights corresponding to the interpolated  $\chi^2$  values and perform the above importance-sampling once more to obtain the desired sample  $Y^r$ .

After repeating the above procedure  $N_R = 10^6$  times, we obtain the distribution of Y with the uncertainties of the physical values of  $w_0$  and  $M_{ss}$  included.

<sup>a</sup>A similar sampling technique was recently proposed in Ref. [9].

Conservative estimate: take larger of these.

Lattice spacing dependence. The naive  $a^2$ -like behaviour is modified to  $\alpha_s^n$  $s<sup>n</sup>a<sup>2</sup>$  by anomalous dimensions of operators in the Symanzik effective theory [8], where  $\alpha_s$  is the strong coupling at the scale of the lattice spacing. To account for this uncertainty, we use two types of fits:

- $A(a^2)$ : Conventional polynomial in  $a^2/w_0^2$ .
- $A'(\Delta_{KS})$ : Scaling of staggered taste violation parameter  $\Delta_{KS}$  is compatible to  $\alpha_s^n$  $\int_s^n a^2$  with  $n=3$  $\longrightarrow$  use polynomials of  $\Delta_{KS}w_0^2$  $\frac{2}{0}$ .

where  $\hat{M}^2 = \frac{1}{2}$ 2  $\left(M_{uu}^2 + M_{dd}^2\right)$  and "phys" denotes the physical values taken from [7].

Putting together, the global fit function is

 $Y = A(a^2) + A'(\Delta_{KS}) + B(a^2)X_l + C(a^2)X_s$ 

 $i,a$ b  $\displaystyle j$ 

where  $Y$  is one of our dimensionless target observables. In the fit function we have the variations:

 $\bullet$  A and  $A'$ : Two scenarios: Either A can be linear, quadratic or cubic in  $a^2$  and  $A'$ is set to zero,

or  $A'$  can be linear, quadratic or cubic in  $\Delta_{KS}$  and A is set to zero.

- B and C: both constant and linear polynomials in  $a^2$ .
- Omit between zero and four of the coarsest lattice spacings (out of a total of seven) from the fits. Include at least one more lattice spacings than number of coefficients in A or  $A'$ , and two more than in B or C.

$$
D(i, c) = w_{ia}
$$

### Intermediate window





95% of our result for  $a_{\mu}$ .

- Beyond reducing the uncertainty on  $a_{\mu,28-\infty}$  by an order of magnitude, the use of a data-driven tail reduces the finite-volume correction that must be applied to the lattice result by a factor of 2 and the associated uncertainty by even more.
- For these large times the data-driven determinations of  $a_{\mu,28-\infty}$  agree very well.

Figure 1: Representative continuum extrapolations of the intermediate window between 0.4 fm and 1.0 fm, as a function of  $a^2$  (left), and  $\Delta_{KS}$  (right). The data points are shifted to the physical point using the  $B$  and  $C$  coefficients. Figure 2: Left: PDF including both statistical and systematic variations. The median is given by the blue vertical line, the  $1\sigma/2\sigma$  error band is shown with green/yellow color. Right: A comparison with other lattice results in the literature [7, 10–16], and a recent pure data-driven computation from Benton et al [17].

## Continuum extrapolation

We perform global fits to the lattice spacing and quark mass dependence.

> • We have checked that the data-driven contribution to  $a_{\mu}$ from  $t \geq 2.8$  fm is entirely compatible with our lattice calculation.



Quark mass dependence. The deviation from the physical light and strange quark mass is described by the variables

 $X_l = \hat M^2 w_0^2$  $\hat{O}^2 - [\hat{M}^2 w_0^2]$  $\left[ \begin{smallmatrix} 2 \ 0 \end{smallmatrix} \right]_\text{phys} \quad \text{and} \quad X_s = M_{ss}^2 w_0^2$  $_0^2$   $\left[ M_{ss}^2 w_0^2 \right.$  $^2_0\big]_{\rm phys}$  ,

> • Choosing to start the data-driven tail above  $t = 2.8$  fm guarantees that the lattice contribution accounts for over

#### References







• The total uncertainty on our average of the data-driven  $a_{\mu,28-\infty}$  is 0.26 in our  $10^{-10}$  units, a number that must be compared to our total uncertainty of 3.3 on  $a_{\mu}$ . Thus its impact on the uncertainty of our final result for  $a_{\mu}$  is completely negligible.

### Reasons for the choice of  $t_{\text{cut}} = 2.8$  fm:

The computation is performed following the approach of Ref. [18]. The measurements of the  $\pi^+\pi^-$  spectrum by BaBar [19, 20], KLOE [21–24], CMD-3 [25] and via hadronic  $\tau$ decays [26, 27] are considered separately. Outside their centerof-mass energy ranges and for other hadronic channels, the data from each experiment are complemented by the combined experimental and perturbative QCD results compiled in Ref. [27, 28], with a full treatment of uncertainties and correlations. Then the HVPTools framework [28–31] is used to Laplace transform these four spectra into the corresponding Euclidean-time correlators, which are subsequently integrated to give the four data-driven results for  $a_{\mu,28-\infty}$ . Altogether, we obtain

# $a_{\mu,28-\infty}=27.59(17)(9)[26] \quad \text{and} \quad a_{\mu,28-35}=18.12(11)(5)[16]$

for our tail-related window results in the data-driven approach. The first error comes from the weighted average procedure of different experiments and includes a PDG-style error rescaling. The second error is the additional uncertainty from including or not including the  $\tau$  data set. The third, conservative total error is the first two added linearly.