Can primordial black holes form without fine-tuning?

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Overview



PBHs from inflation

Fine-tuning



Overdensities seeded by inflation

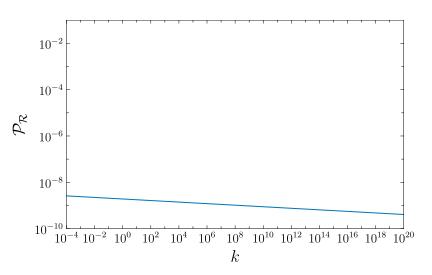


- Overdensities seeded by inflation
- ► Slow-roll inflation explains CMB

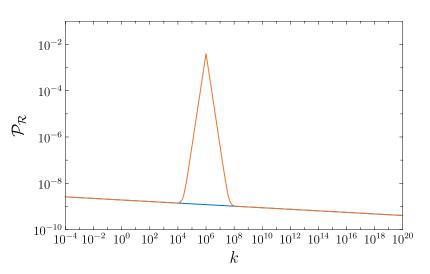


- Overdensities seeded by inflation
- ► Slow-roll inflation explains CMB
- Nearly scale-invariant power spectrum









Inflaton potential



▶ Inflation driven by scalar field ϕ

Inflaton potential



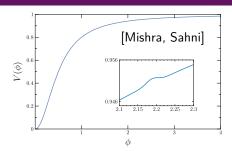
- lacktriangle Inflation driven by scalar field ϕ
- ▶ Governed by potential $V(\phi)$

Inflaton potential

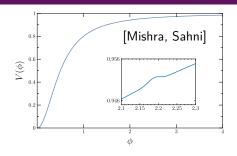


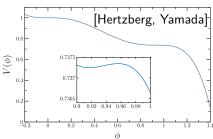
- ▶ Inflation driven by scalar field ϕ
- lacktriangle Governed by potential $V(\phi)$
- ► Need some feature in potential



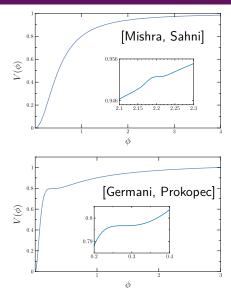


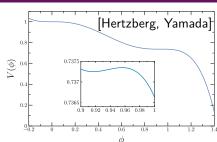




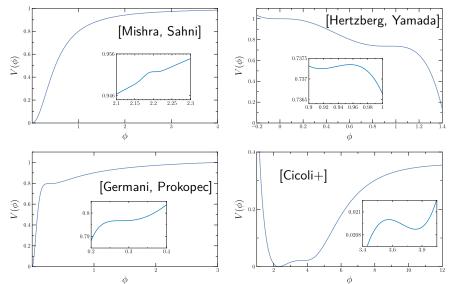














	n_s	r
Mishra & Sahni	0.9648	0.0026
Hertzberg & Yamada	0.9820	4.8×10^{-7}
Hertzberg & Yamada Germani & Prokopec	0.9567	0.0063
Cicoli et al.	0.9400	0.018

Planck constraint:

- $n_s = 0.9649 \pm 0.0042$
- r < 0.032

Overview



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Fine-tuning

Power spectrum fine-tuning



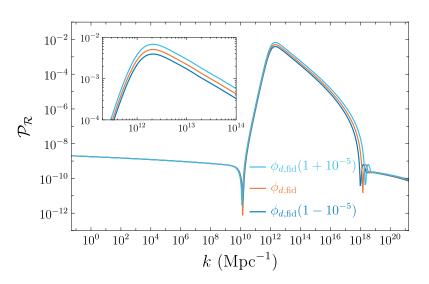
► Solve Mukhanov–Sasaki equation for each potential

Power spectrum fine-tuning

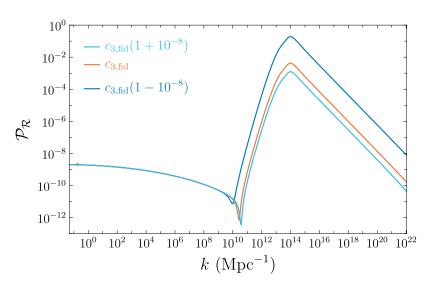


- ► Solve Mukhanov–Sasaki equation for each potential
- Perturb potential parameters and compare power spectra

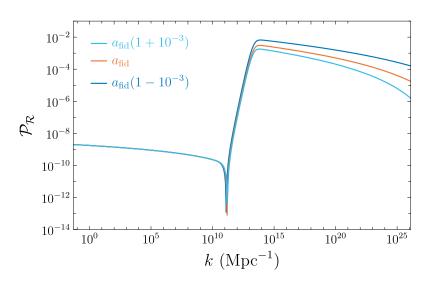






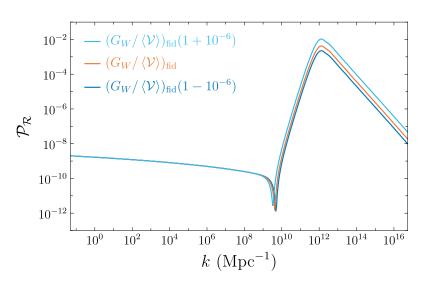






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Fine-tuning measure



Quantify fine-tuning using differential measure [Azhar, Loeb]

$$\epsilon_{\mathcal{O}} = \frac{\mathsf{d}\log\mathcal{O}}{\mathsf{d}\log p}$$



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Model	$\epsilon_{\mathcal{P}_{peak}}$	$\epsilon_{f_{PBH}}$	ρ
Mishra & Sahni	2.7×10^{4}	6.2×10^{5}	23
Hertzberg & Yamada	-1.8×10^{8}	-4.7×10^9	27
Germani & Prokopec	-6.0×10^2	-2.2×10^{4}	37
Cicoli et al.	7.5×10^5	2.2×10^7	29



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- Typical potentials seem to be quite fine-tuned
- Total PBH fine-tuning actually dominated by power spectrum fine-tuning
- Could we be saved by non-Gaussianity?

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