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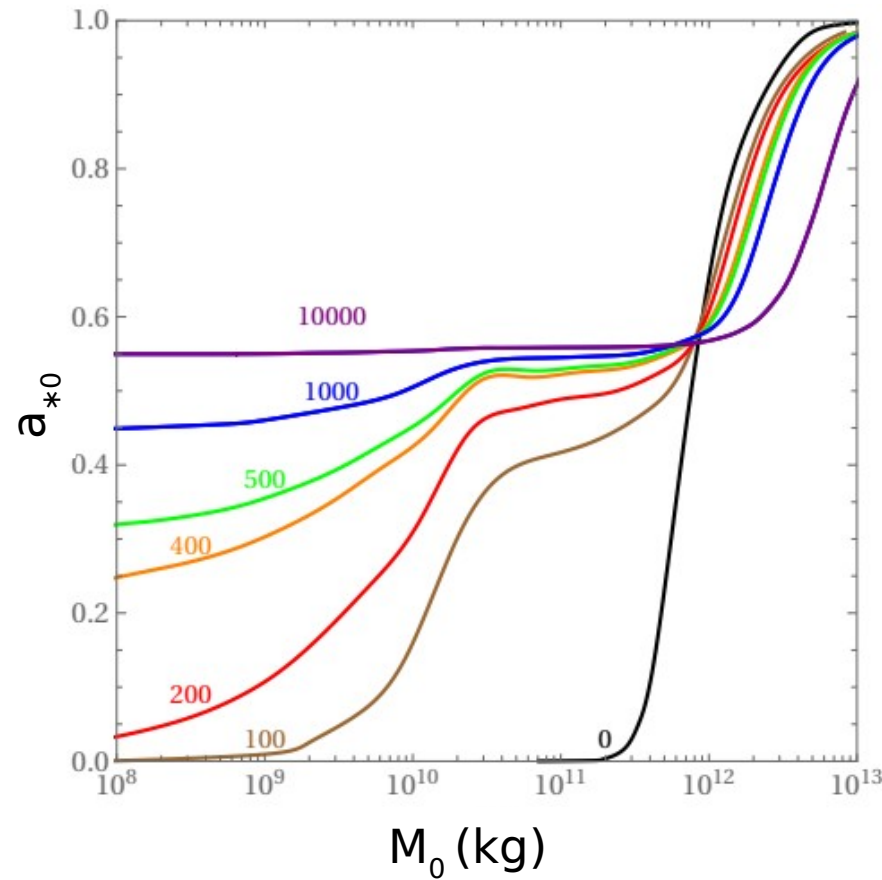
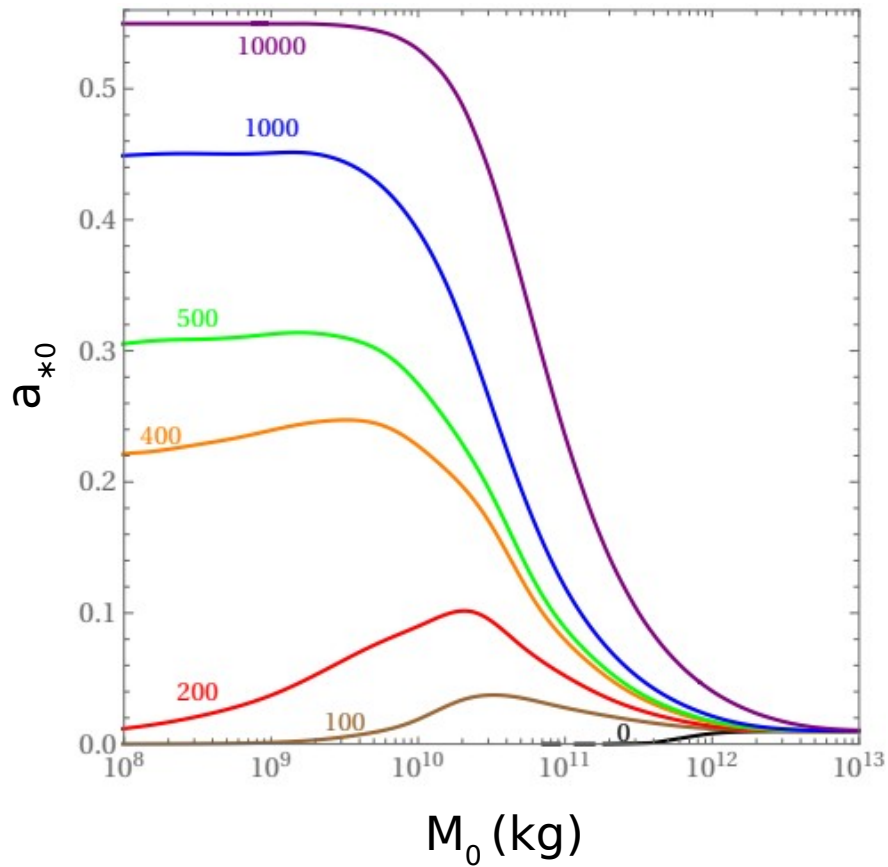
Determining mass and spin of light PBHs

Marco Calzà.

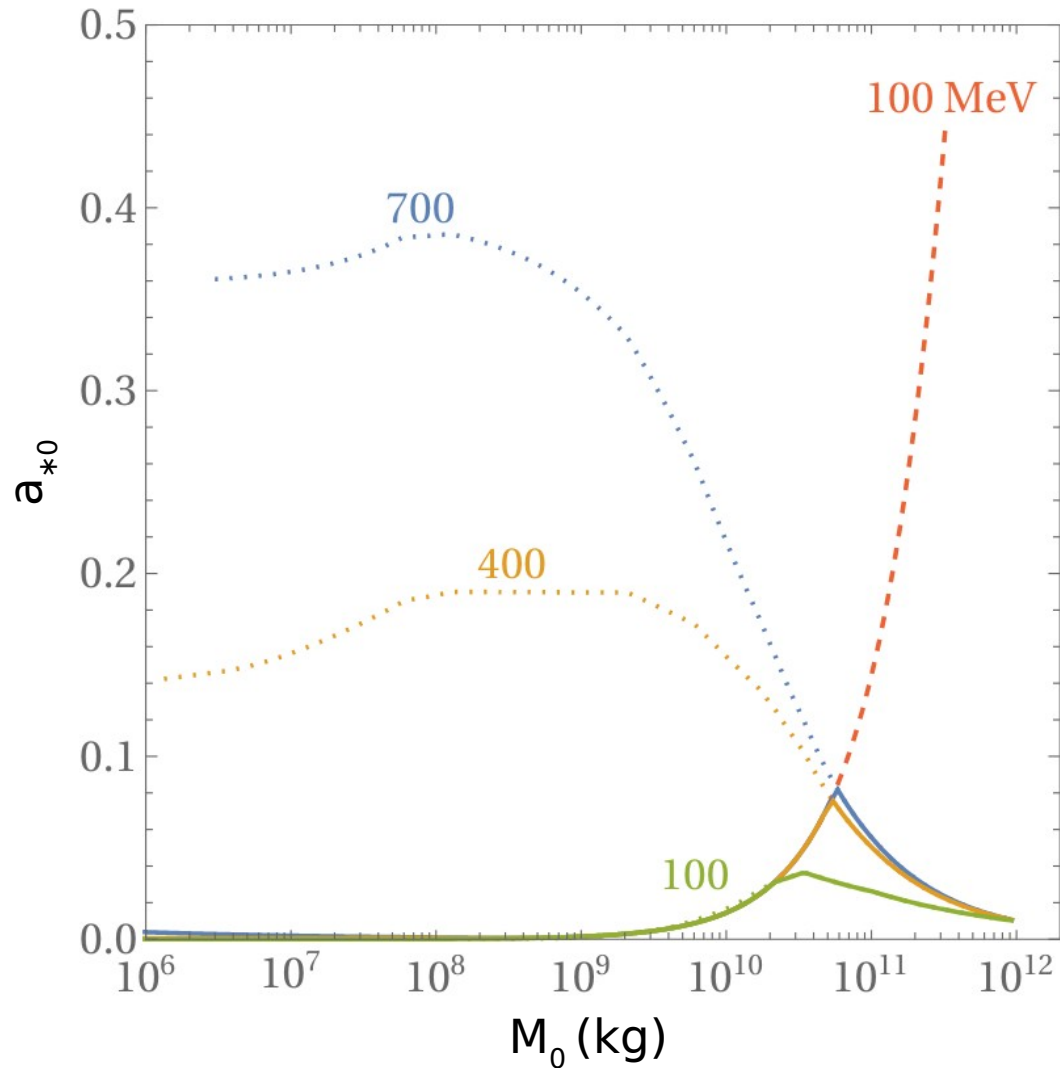
University of Trento.

In collaboration with João G. Rosa.

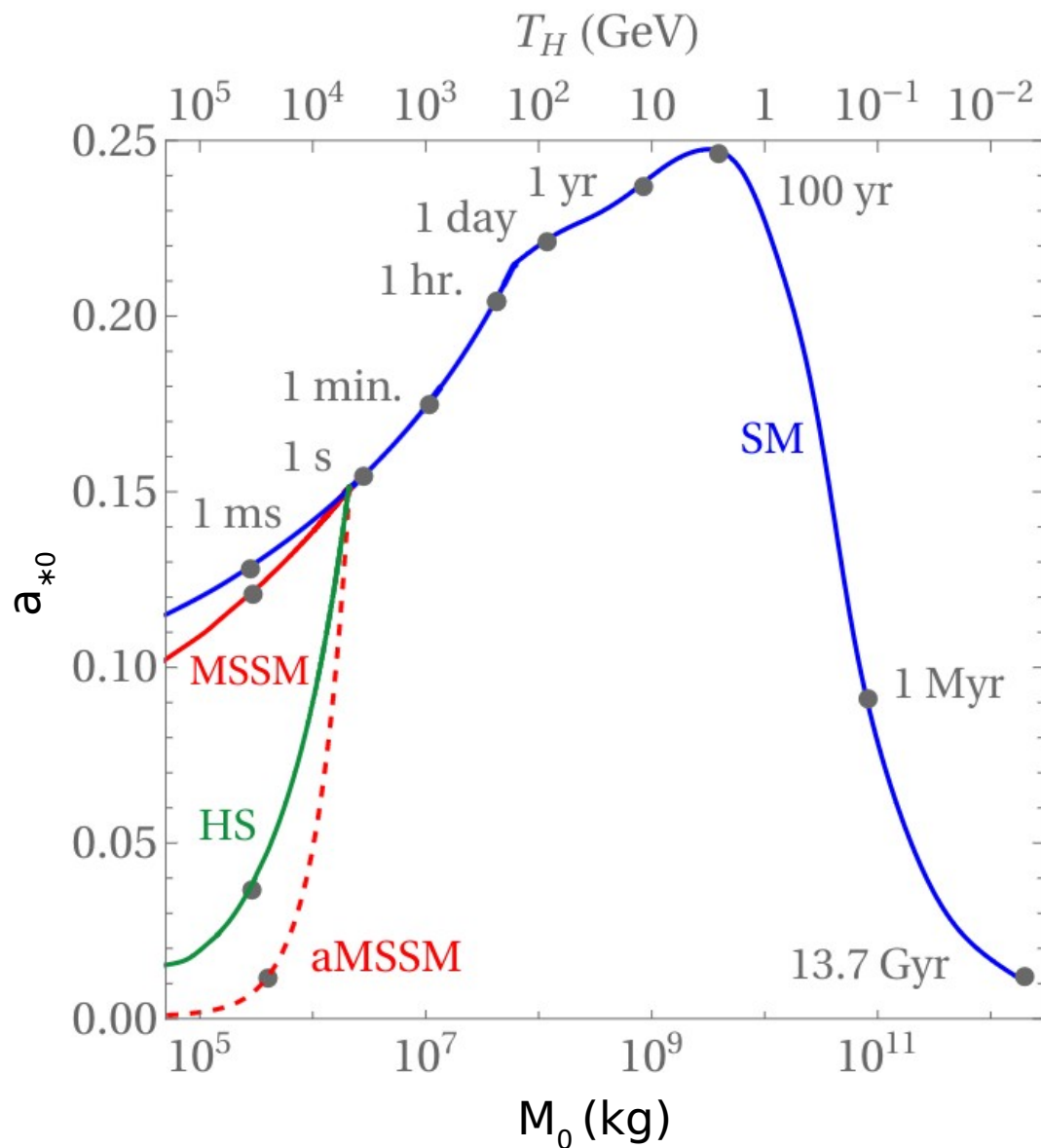
Axiverse fingerprint in PBH evaporation



Superradiance + ALPs

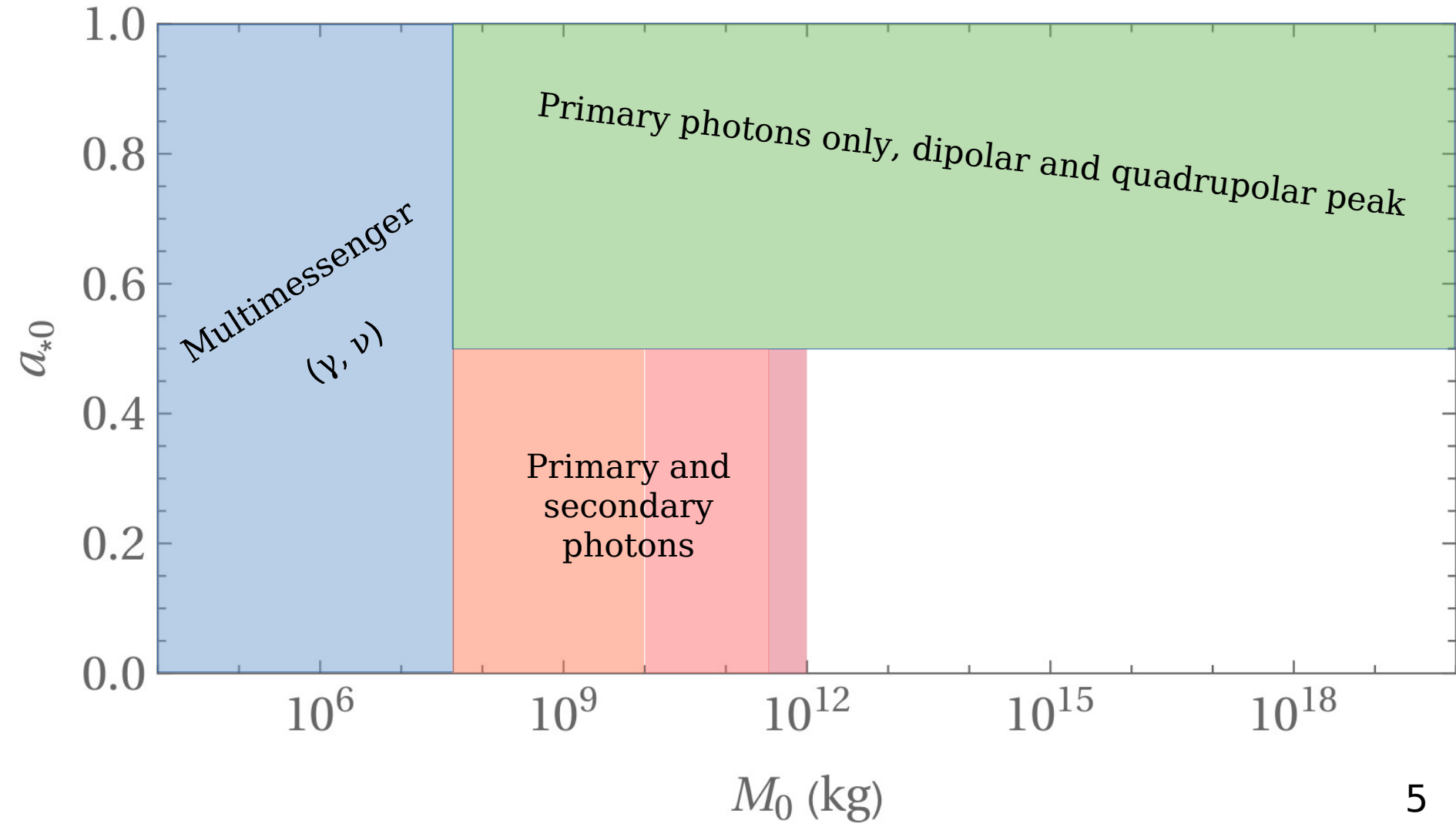


High Energy BSM

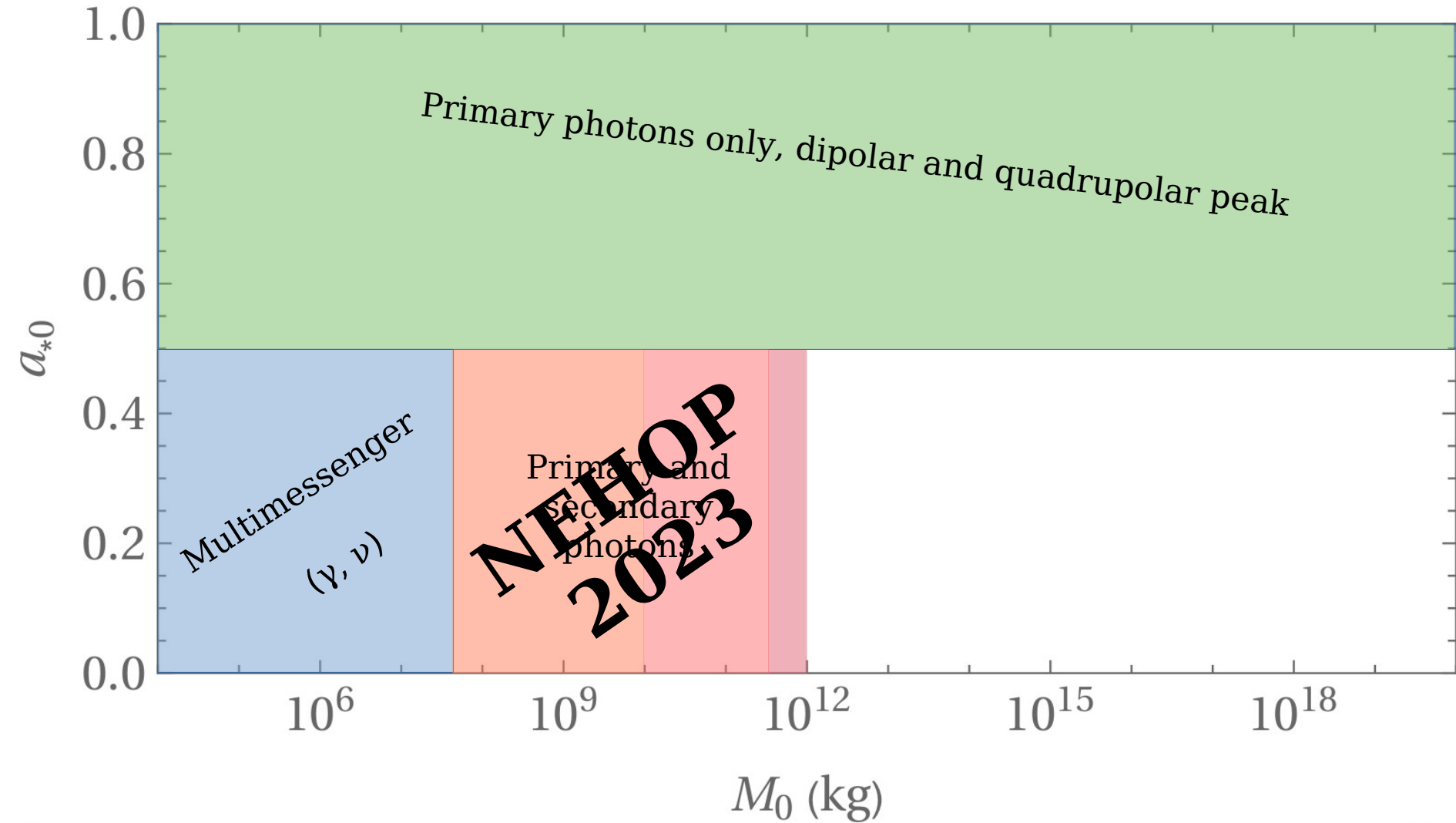


Take home message:

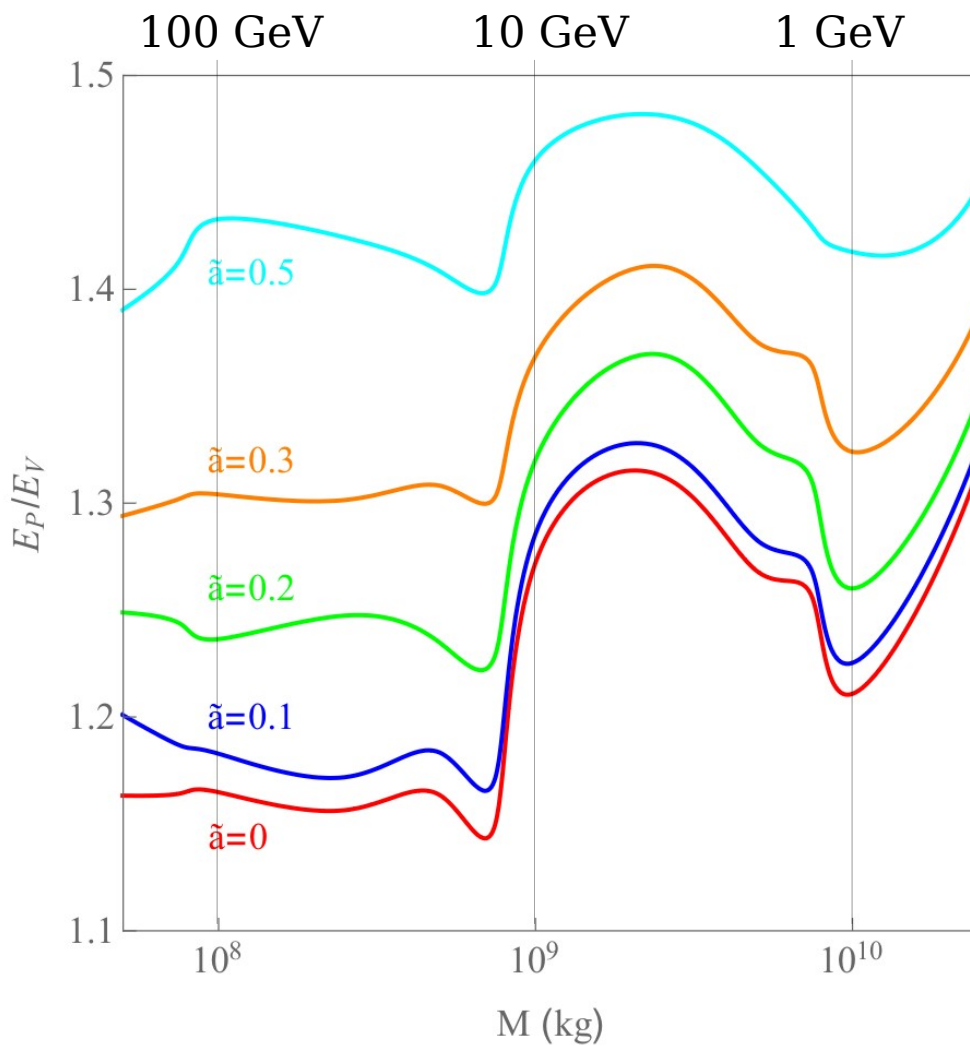
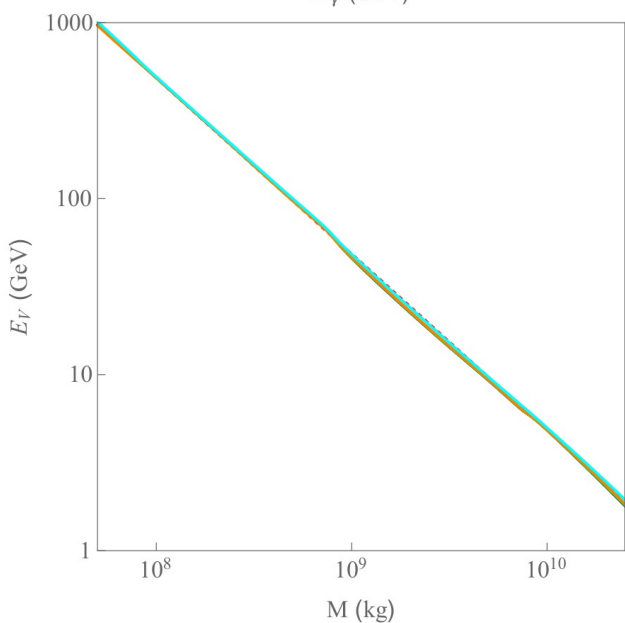
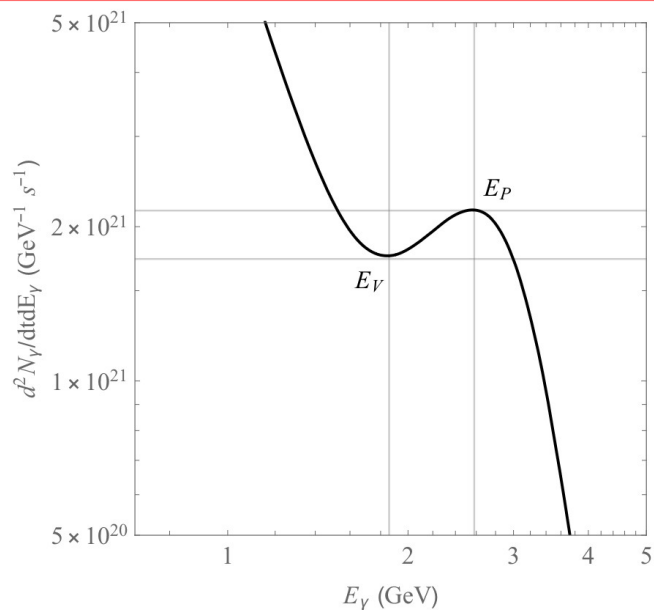
New Physics with PBHs \rightarrow know the position on the plane (a_{*0}, M)



Prequel

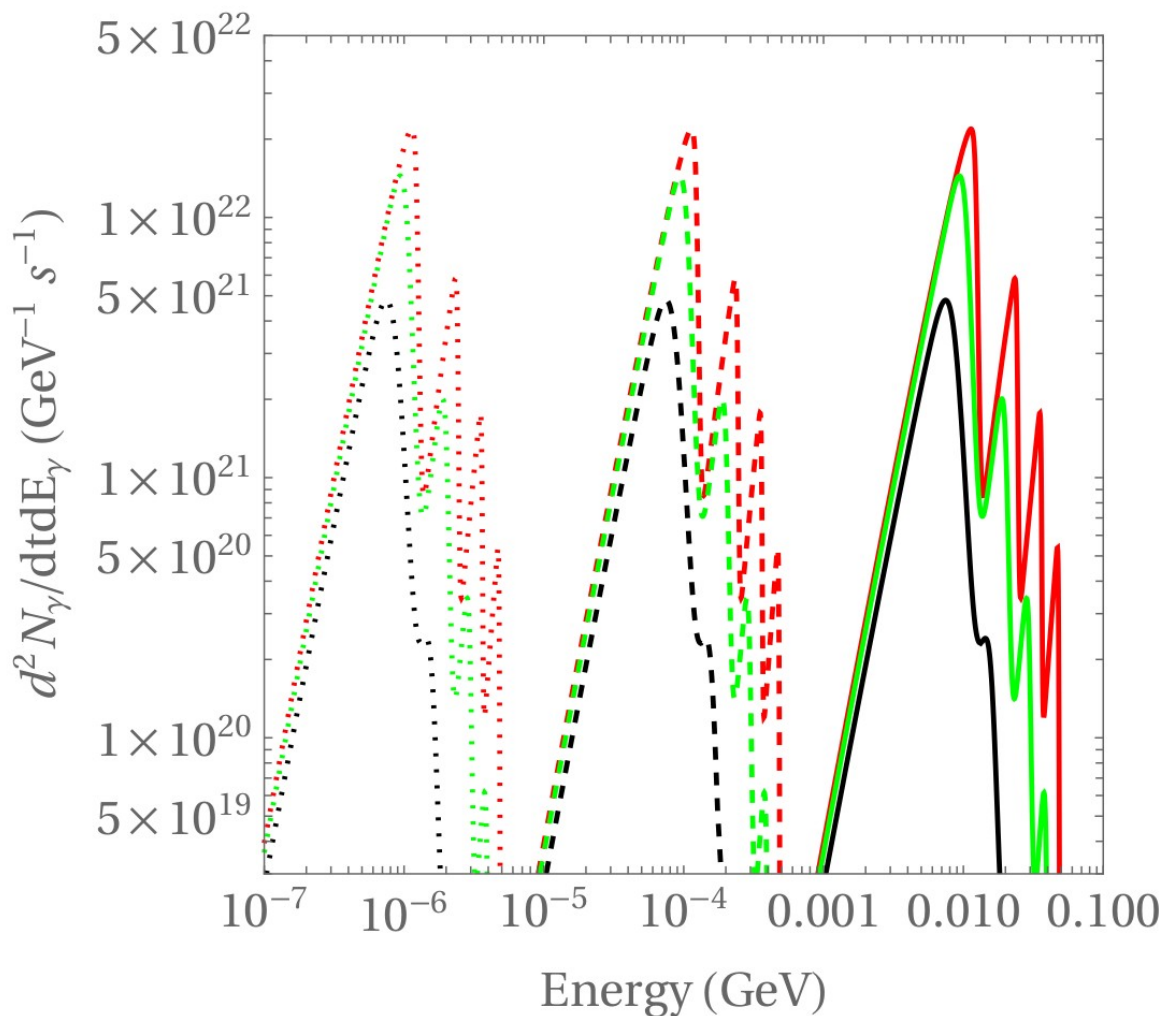


In previous episod: recap NEHOP 2023

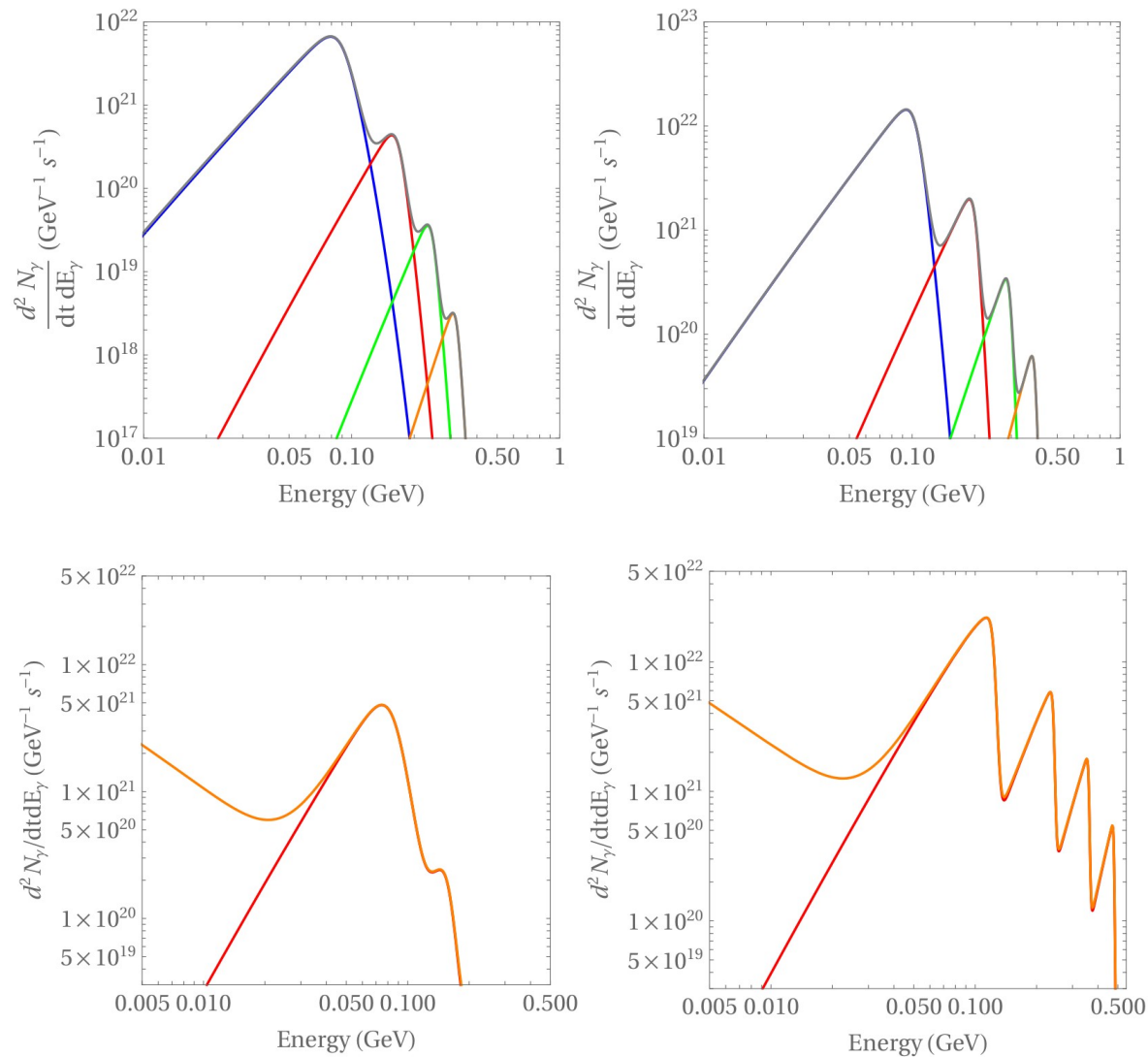


What happen increasing the spin?

$$\frac{d^2 N_{P,i}}{dt dE_i} = \frac{1}{2\pi} \sum_{l,m} \frac{\Gamma_{l,m}^s(\omega)}{e^{2\pi k/\kappa} \pm 1},$$

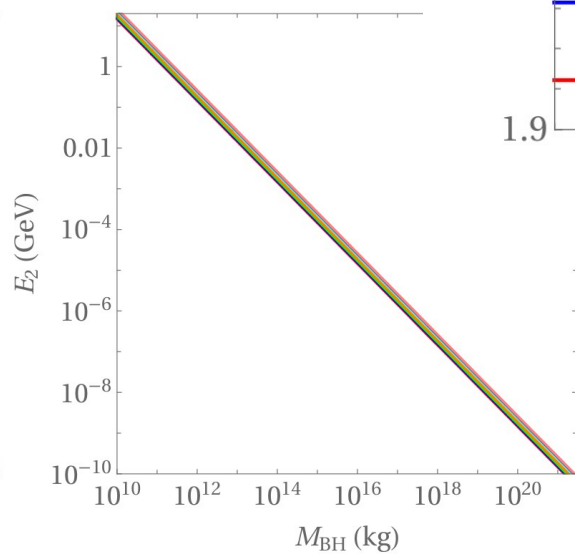
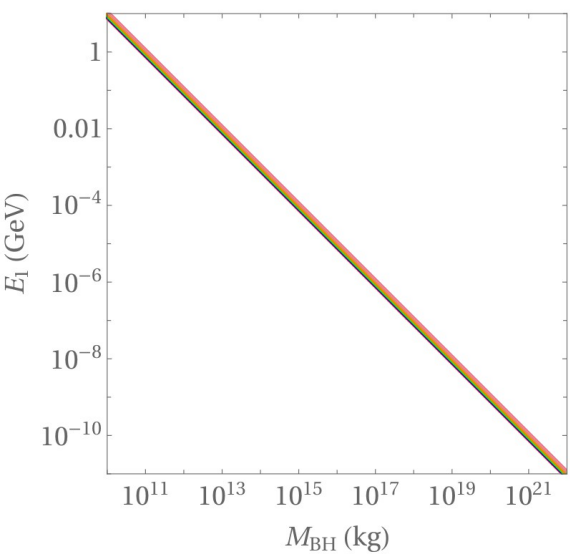
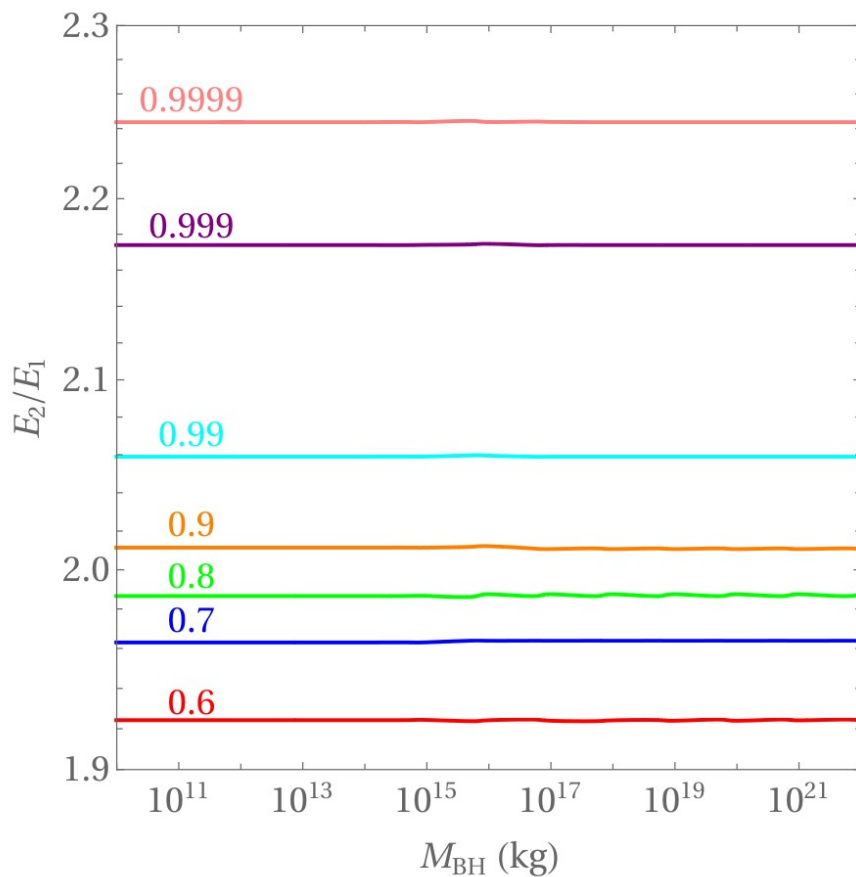
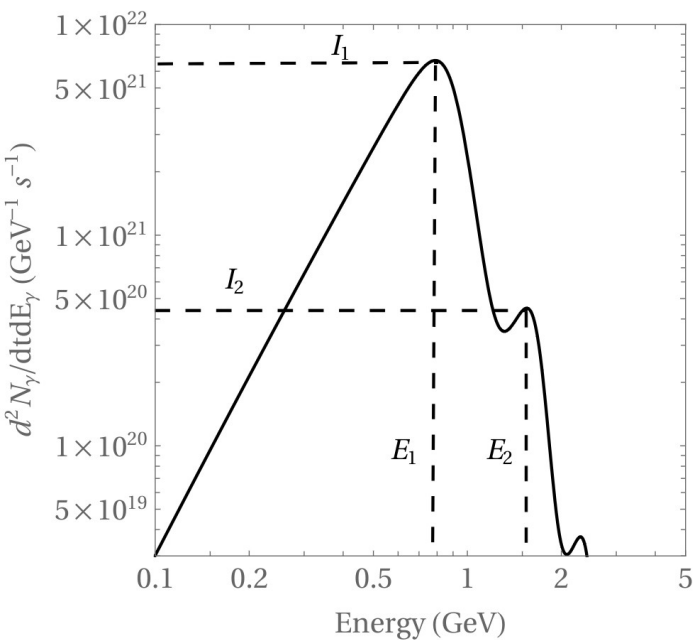


What happen increasing the spin?



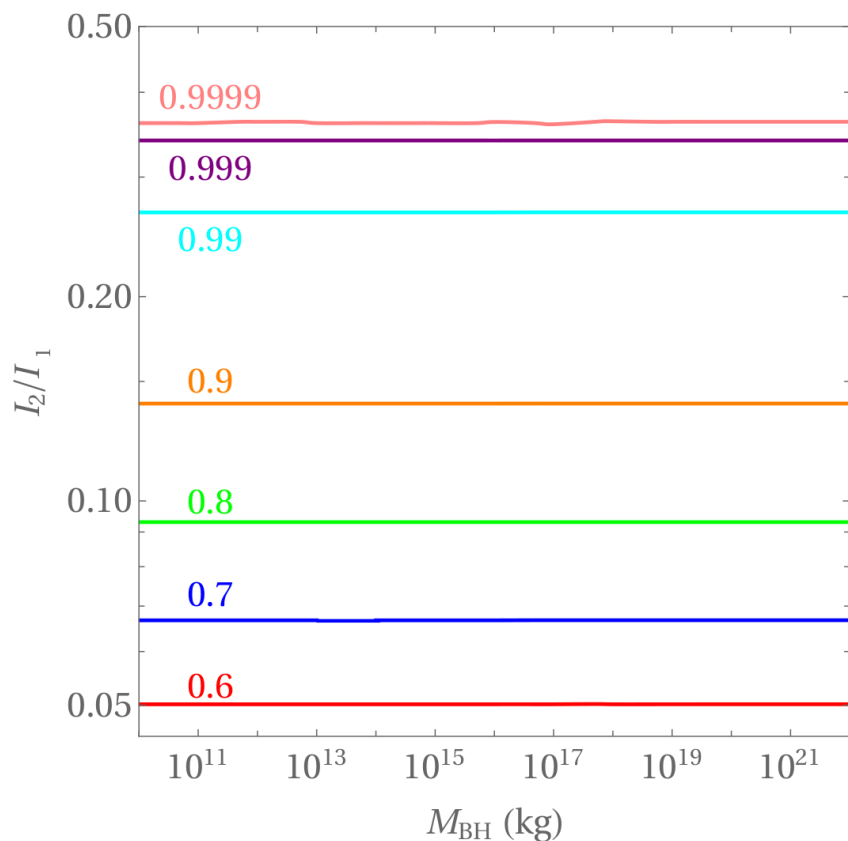
PBH with $M = 10^{12}$ kg and (left) $\tilde{a} = 0.6$ and (right) $\tilde{a} = 0.99$.

What happen increasing the spin?



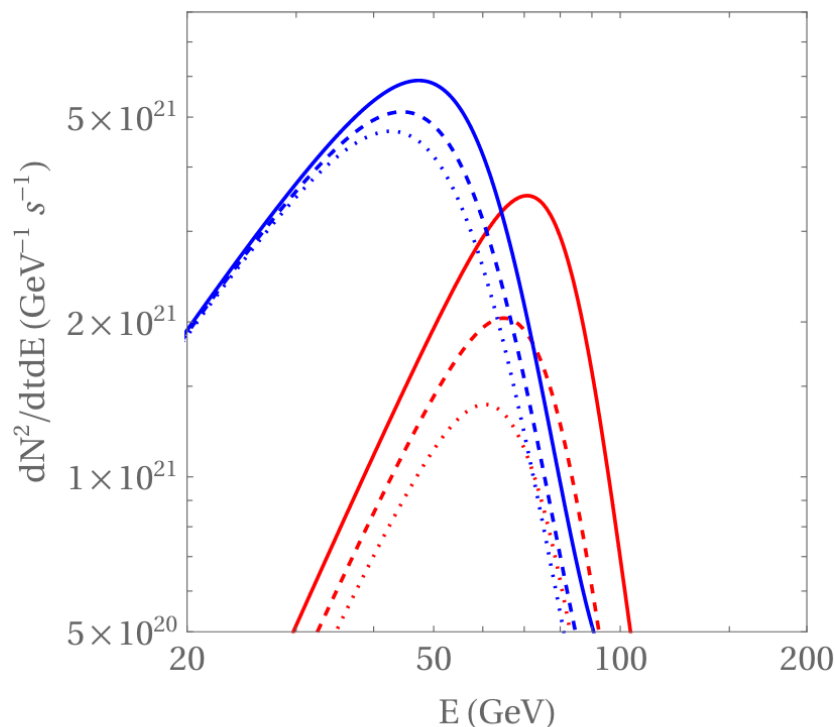
What happen increasing the spin?

$$\frac{d^2 N_{P,i}}{dt dE_i d\Omega} = \frac{1}{4\pi} \sum_{l,m} \frac{\Gamma_{l,m}^s}{e^{2\pi k/\kappa} \pm 1} (|S_{l,m}^{-s}(\theta)|^2 + |S_{l,m}^s(\theta)|^2) ,$$



$$\frac{\mathcal{F}_2}{\mathcal{F}_1} = \frac{I_2}{I_1} \frac{|S_{2,2}^1(\theta)|^2}{|S_{1,1}^1(\theta)|^2} \simeq \frac{I_2}{I_1} \frac{|Y_{2,2}^1(\theta)|^2}{|Y_{1,1}^1(\theta)|^2} \simeq \frac{5}{3} \frac{I_2}{I_1} \sin^2 \theta ,$$

How to infer mass and spin?



Primary photon (red) and single neutrino (blue) Hawking emission spectra for a PBH with $M = 10^9$ kg and $\tilde{a} = 0$ (dotted), $\tilde{a} = 0.2$ (dashed) and $\tilde{a} = 0.5$ (solid).

$$E_{\gamma,\nu} \propto T_H \propto M^{-1}$$

Maximum emission rates $I_{\nu,\nu}$ independent of the PBH mass.

The ratios between these quantities for ν and γ depend only on \tilde{a}

$$\frac{E_\nu}{E_\gamma} = 0.705 - \frac{0.559\tilde{a}^2}{1 + 5.18\tilde{a}}, \quad \frac{I_\nu}{I_\gamma} = 3.423 - \frac{31.05\tilde{a}^2}{1 + 7.05\tilde{a}}$$

Conclusions

- Use PBHs to investigate fundamental building blocks of Physics
 - Number of light ALPs.
 - Possible heavy ALPs.
 - High energy BSM content.
- The information is encoded in the Regge plots $(M, \tilde{\alpha})$.
- It is possible to investigate the $(M, \tilde{\alpha})$ parameter space using different methods:
 - Primary and secondary photons (low spin).
 - Dipolar and Quadrupolar photon peak (high spin).
 - Multimessenger: primary ν , γ (high temperature).

References

- [1] M. Calzá, J. March-Russell & J. G. Rosa, [arXiv:2110.13602] (sub. to PRL)
- [2] M. Calzá & J. G. Rosa, JHEP 12 (2022) 090
- [3] M. Calzá, Phys. Rev. D 107 (2023) no.4, 044067
- [4] M. Calzá, J. G. Rosa & F. Serrano, JHEP 05 (2024) 140
- [5] M. Calzá, J. G. Rosa, [arXiv:2311.12930] (sub. to JHEP)
- [6] M. Calzá, J. G. Rosa, [arXiv:2312.09261] (sub. to PRL)

Thanks for your attention!!!



“That’s all Folks!”