

ML-Wave

Tilman Plehn

LHC theory

ML introduction

Examples

Regression

Generative AI

Transformation

Ride the ML-Wave or Drown? (How To Not Become a Physics Museum)

Tilman Plehn

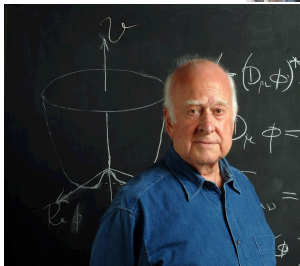
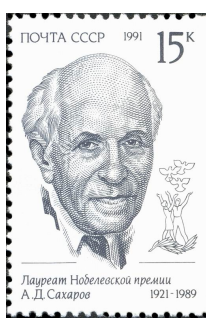
Universität Heidelberg

Annual Theory Meeting, IPPP, 2024



Classic motivation

- dark matter?
- matter vs antimatter?
- origin of Higgs boson?



All about LHC physics

Classic motivation

- dark matter?
- matter vs antimatter?
- origin of Higgs boson?

LHC physics

- fundamental questions
- huge data set
- first-principle, precision simulations
- complete uncertainty control

Successful past

- measurements of total rates
- analyses inspired by simulation
- model-driven Higgs discovery



Classic motivation

- dark matter?
- matter vs antimatter?
- origin of Higgs boson?

LHC physics

- fundamental questions
- huge data set
- first-principle, precision simulations
- complete uncertainty control

Successful past

- measurements of total rates
- analyses inspired by simulation
- model-driven Higgs discovery

First-principle, precision simulations

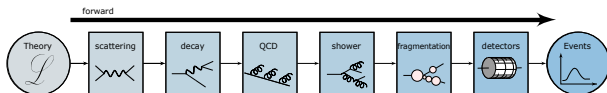
- start with Lagrangian
- calculate scattering using QFT
- simulate collisions
- simulate detectors

→ LHC collisions in virtual worlds

(Instead of) BSM searches

- compare simulations and data
- infer underlying theory [SM or BSM]
- publish results/data meaningfully

→ Understand LHC data systematically

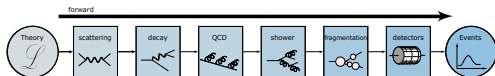


LHC Theory

Turning data to knowledge

- Quantum Field Theory
start with Lagrangian
- compute hard scattering
compute decays
compute jet radiation
- parton densities [NNPDF]
hadron-level QCD

→ First-principle simulations, with help from ML



Turning data to knowledge

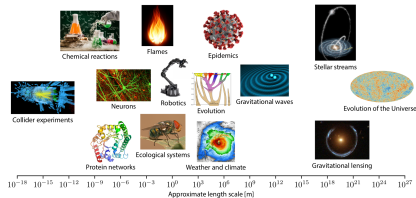
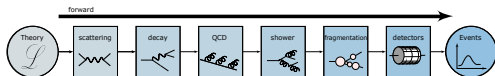
- Quantum Field Theory start with Lagrangian
- compute hard scattering
compute decays
compute jet radiation
- parton densities [NNPDF]
hadron-level QCD

→ First-principle simulations, with help from ML

HL-LHC: inference with 10× more data

- SBI starts with Simulation...
- statistical improvement $\sqrt{10} = 3$
- rate over phase space to $< 0.1\%$
- theory to follow
- precision = QFT × Compute

→ Everything, faster and better...



Fit-like approximation

- approximate $f_\theta(x) \approx f(x)$
- no parametrization, just very many θ
- new representation/latent space θ

Applications

- applications all over experiment
- regression $x \rightarrow f_\theta(x)$
- classification $x \rightarrow f_\theta(x) \in [0, 1]$
- generation $r \sim \mathcal{N} \rightarrow f_\theta(r)$
- conditional generation $r \sim \mathcal{N} \rightarrow f_\theta(r|x)$

Architectures

- physics-aware questions and data representation
 - symmetries, locality, etc
 - accuracy, control, error bars?
 - is LHC data images or language?
- Complexity a feature, not a problem



Encoding a transition amplitude-squared

- expectation value from probability

$$\langle A \rangle(x) = \int dA A p(A|x)$$

- internal representation θ

$$\langle A \rangle = \int dA A \int d\theta p(A|\theta) p(\theta|A_{\text{train}})$$

- training a generalization of θ -probability

$$\int d\theta p(A|\theta) p(\theta|A_{\text{train}}) \approx \int d\theta p(A|\theta) q(\theta)$$



Encoding a transition amplitude-squared

- expectation value from probability

$$\langle A \rangle(x) = \int dA A p(A|x)$$

- internal representation θ

$$\langle A \rangle = \int dA A \int d\theta p(A|\theta) p(\theta|A_{\text{train}})$$

- training a generalization of θ -probability

$$\int d\theta p(A|\theta) p(\theta|A_{\text{train}}) \approx \int d\theta p(A|\theta) q(\theta)$$

- similarity from minimal KL-divergence

$$\begin{aligned} D_{\text{KL}}[q(\theta), p(\theta|A_{\text{train}})] &\equiv \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta|A_{\text{train}})} \\ &= \int d\theta q(\theta) \log \frac{q(\theta)p(A_{\text{train}})}{p(A_{\text{train}}|\theta)p(\theta)} \\ &= - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta)} + \dots \end{aligned}$$

→ Simplification: likelihood + regularization + dropout

$$\mathcal{L}_{\text{BNN}} = - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + D_{\text{KL}}[q(\theta), p(\theta)]$$

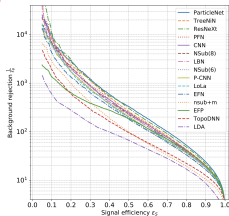
$$\rightarrow (A_{\theta} - A_{\text{train}})^2 + c(\theta - \theta_0)^2$$



Top tagging [classification, 2016-today]

- 'hello world' of LHC-ML
- end of QCD-taggers
- ever-improving [Huilin Qu]

→ **Driving NN-architectures**



SciPost Physics

Submission

The Machine Learning Landscape of Top Taggers

G. Kasieczko (ed)¹, T. Plehn (ed)², A. Butter³, K. Cranmer⁴, D. Dikou⁵, B. M. Dolan⁶, M. Fairbairn⁷, D. A. Faroughy⁷, W. Fisher⁸, C. Gay⁹, L. Grigoras⁹, J. F. Kanieta¹⁰, P. T. Komarek¹¹, S. Laha¹², A. Lauer¹³, S. Mariani¹⁴, E. M. Metodiev¹⁵, L. Moore¹⁶, B. Nachman^{17,18}, K. Nishitani^{19,20}, J. D. Potter²¹, H. Qiu²², Y. Rauh²³, M. Ripken²⁴, D. Shi²⁵, J. M. Thompson²⁶, and S. Varma²⁷

¹ Institut für Experimentalphysik, Universität Hamburg, Germany

² Institut für Theoretische Physik, Universität Heidelberg, Germany

³ Center for Cosmology and Particle Physics and Center for Data Science, NYU, USA

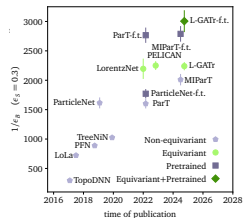
⁴ NHECT, Dept. of Physics and Astronomy, Rutgers, The State University of NJ, USA

⁵ Jozef Stefan Institute, Ljubljana, Slovenia

⁶ Theoretical Particle Physics and Cosmology, King's College London, United Kingdom

⁷ Department of Physics and Astronomy, The University of British Columbia, Canada

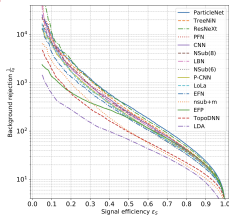
⁸ Department of Physics, University of California, Santa Barbara, USA



ML in experiment

Top tagging [classification, 2016-today]

- 'hello world' of LHC-ML
 - end of QCD-taggers
 - ever-improving [Huilin Qu]
- **Driving NN-architectures**



Particle flow [2020-today]

- mother of jet analyses
 - combining detectors with different resolution
 - optimality the key
- **Modern jet analysis basics**

Towards a Computer Vision Particle Flow *

Francesco Armando Di Bello^{1,2}, Samay Ganguly^{3,4}, Elian Gross¹, Marumi Kado^{5,6},
Michael Pitt¹, Lorenzo Santini⁷, Jonathan Shlomo⁸

¹Wizman Institute of Science, Rehovot 7610, Israel

²CEBN, CB 1211, Geneva 23, Switzerland

³Università di Roma Sapienza, Piazza Aldo Moro, 2, 00185 Roma, Italy e INFN, Italy

⁴Université Paris-Saclay, CNRS/IN2P3, UCLM, 91195, Orsay, France

Farek Mokhtar⁹, Jonny Patel⁹, Javier Duarte⁹, Eric Wu¹⁰,
Maurizio Piatek¹¹ and Jean-Roch Villain¹²
(on behalf of the CMS Collaboration)

⁹University of California San Diego, La Jolla, CA 92093, USA

¹⁰UCPS, Geneva 04, 1213, Switzerland

¹¹European Organization for Nuclear Research (CERN), CH 1211, Geneva 23, Switzerland

¹²California Institute of Technology, Pasadena, CA 91125, USA

*Email: francescoarmando@ictp.ac, jonny@particle.ox.ac, jvillain@ucsd.edu

Progress towards an improved particle flow algorithm at CMS with machine learning

SciPost Physics

Submission

The Machine Learning Landscape of Top Taggers

G. Kasieczko (ed)¹, T. Plehn (ed)², A. Butter³, K. Cranzer⁴, D. Datta⁵, B. M. Dillon⁶,
M. Fairhead⁷, D. A. Faroughy⁸, W. Fisher⁹, C. Gay¹⁰, L. Goodson¹¹, J. F. Kaniatis¹²,
P. T. Kennedey¹³, S. Laha¹⁴, A. Lister¹⁵, S. Mariani¹⁶, E. M. Metodiev¹⁷, E. Moon¹⁸,
B. Nachman^{19,20}, K. Nourbakhsh^{21,22}, J. Pfaender²³, H. Qi²⁴, Y. Rath²⁵, M. Rinke²⁶, D. Shih²⁷,
J. M. Thompson²⁸, and S. Viretta²⁹

¹ Institut für Experimentalphysik, Universität Hamburg, Germany

² Institut für Theoretische Physik, Universität Heidelberg, Germany

³ Center for Cosmology and Particle Physics and Center for Data Science, NYU, USA

⁴ INFNCT, Dep. of Physics and Astronomy, Rutgers The State University of NJ, USA

⁵ Josef Stefan Institute, Ljubljana, Slovenia

⁶ Theoretical Particle Physics and Cosmology, King's College London, United Kingdom

⁷ Department of Physics and Astronomy, The University of British Columbia, Canada

⁸ Department of Physics, University of California, Santa Barbara, USA

⁹ Department of Mathematics and Physics, University of Ljubljana, Ljubljana, Slovenia

¹⁰ Center for Theoretical Physics, MIT, Cambridge, USA

¹¹ CP3, Université Catholique de Louvain, Louvain-la-Neuve, Belgium

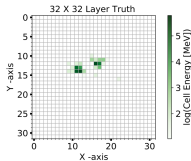
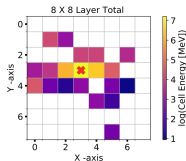
¹² Physics Division, Lawrence Berkeley National Laboratory, Berkeley, USA

¹³ Simons Inst. for the Theory of Computing, University of California, Berkeley, USA

¹⁴ National Institute for Subatomic Physics (NIKHEF), Amsterdam, Netherlands

¹⁵ LPTHE, CNRS & Sorbonne Université, Paris, France

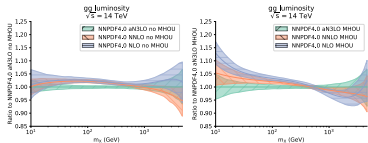
¹⁶ III. Physikalisches Institut A, RWTH Aachen University, Germany



Parton densities [NNPDF, 2002-today]

- pdfs without functional bias and full uncertainties
- precision and calibrated uncertainties

→ Drivers of ML-theory

The Path to N²LO Parton Distributions

The NNPDF Collaboration:

Richard D. Ball¹, Andrea Bharucha¹, Alessandro Cacciari^{2,3}, Stefano Carrazza², Juan Cruz-Martinez⁴, Luigi Del Debbio⁵, Stefano Forte⁶, Tommaso Gehrmann⁷, Hilke Hehner^{8,9}, Zakari Karakochev⁹, Niccolò Leonardi², Giacomo Magni^{4,5}, Emanuele M. Nicosia⁴, Tanjaana H. Reheisenhauer^{4,5}, Juan Rojo^{4,5}, Christopher Schmidt¹⁰, Roy Steegmans¹¹, and Martin Ubach⁸

¹The Hugh Downs Centre for Theoretical Physics, University of Edinburgh, JCMB, KB, Hughdown Rd, Edinburgh EH9 1JZ, Scotland

²INFN, Dipartimento di Fisica, Università di Milano and INFN, Sezione di Milano, Via Celoria 16, I-20133 Milano, Italy

³CERN, Theoretical Physics Department, CH-1211 Geneva 23, Switzerland

⁴Department of Physics and Astronomy,rije University, NL-1091 BT Amsterdam

⁵Mathij Theory Group, Science Park 105, 1098 XJ Amsterdam, The Netherlands

⁶University of Jyväskylä, Department of Physics, P.O. Box 35, FI-00014 University of Jyväskylä, Finland

⁷Helsinki Institute of Physics, P.O. Box 64, FI-00014 University of Helsinki, Finland

⁸DAMTP, University of Cambridge, Wilberforce Road, Cambridge, CB3 0BB, United Kingdom

⁹Dipartimento di Fisica, Università degli Studi di Teramo and INFN, Sezione di Teramo, Via Porto Garibaldi, I-66100 Teramo, Italy

¹⁰Universität Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany

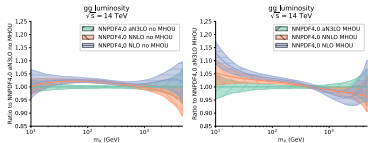
This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being



Parton densities [NNPDF, 2002-today]

- pdfs without functional bias and full uncertainties
- precision and calibrated uncertainties

→ Drivers of ML-theory

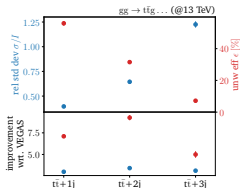


Ultra-fast event generators [Sherpa, MadNIS, MLHad]

- event generation modular
- improve and replace by ML-modules

→ Beat state of the art

| | | | |
|------------------|----------------------------------|-------------------------------|--------------------------------|
| Triple-W | $u\bar{d} \rightarrow W^+W^+W^-$ | | |
| VBS | $uc \rightarrow W^+W^+ds$ | | |
| W+jets | $gg \rightarrow W^+d\bar{u}$ | $gg \rightarrow W^+d\bar{u}g$ | $gg \rightarrow W^+d\bar{u}gg$ |
| $t\bar{t}$ +jets | $gg \rightarrow t\bar{t}+g$ | $gg \rightarrow t\bar{t}+gg$ | $gg \rightarrow t\bar{t}+ggg$ |

The Path to N²LO Parton Distributions

The NNPDF Collaboration:

Richard D. Ball¹, Andrea Bharucha², Alessandro Ciaffaglia^{3,4}, Stefano Carrazza⁵, Juan Cruz-Martinez⁶, Luigi Del Debbio⁷, Stefano Forte⁸, Tommaso Gehrmann⁹, Jolite Heikkinen^{10,11}, Zakari Karimov¹², Niccolò Lauretti¹³, Giacomo Maga^{14,15}, Emanuele M. Nicosia¹⁶, Tapani H. Reunanen^{17,18}, Jean Roy^{19,20}, Christopher Schmidt²¹, Roy Stegmann²², and Maria Ubachs²³

¹The High Energy Physics Theory Group, University of Edinburgh, UK²INFN, Sezione di Milano, Via Celoria 16, I-20133 Milano, Italy³INFN, Sezione di Milano, Via Celoria 16, I-20133 Milano, Italy⁴CEBN, Theoretical Physics Department, CN-LSI Geneva 23, Switzerland⁵Department of Physics and Astronomy, Piip University, EE-10111 Tartu, Estonia⁶Mathematical Physics Group, Science Park 107, 1098 XE Amsterdam, The Netherlands⁷University of Jyväskylä, Department of Physics, P.O. Box 35, FI-40014 University of Jyväskylä, Finland⁸HEP, Institute of Physics, P.O. Box 61, FI-00014 University of Helsinki, Finland⁹DMTP, University of Cambridge, Wilberforce Road, Cambridge, CB3 0WA, United Kingdom¹⁰Department of Physics, University of Jyväskylä, Finland¹¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy¹²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany¹³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy¹⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany¹⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy¹⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany¹⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy¹⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany¹⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy²⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany²¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy²²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany²³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy²⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany²⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy²⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany²⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy²⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany²⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy³⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany³¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy³²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany³³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy³⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany³⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy³⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany³⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy³⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany³⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁴⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁴¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁴²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁴³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁴⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁴⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁴⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁴⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁴⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁴⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁵⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁵¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁵²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁵³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁵⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁵⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁵⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁵⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁵⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁵⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁶⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁶¹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁶²University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁶³INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁶⁴University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁶⁵INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁶⁶University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁶⁷INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁶⁸University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany⁶⁹INFN, Sezione di Trieste, Via Padua 16, I-34127 Trieste, Italy⁷⁰University of Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany

SciPost Physics

Submission

The MANNIS Reloaded

Théo Heijmaal¹, Nathan Bharucha², Fabio Maltoni^{3,4},
Olivier Mattelaer⁵, Tilman Plehn⁶, and Ramon Winterhalder⁷

¹ Institut für Theoretische Physik, Universität Heidelberg, Germany
² CP3, Université catholique de Louvain, Louvain-la-Neuve, Belgium
³ Dipartimento di Fisica e Astronomia, Università di Bologna, Italy

December 17, 2024

Abstract

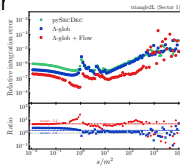
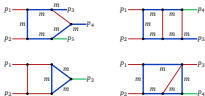
In pursuit of precise and fast theory predictions for the LHC, we present an implementation of the MANNIS method in the MADGRAPH5 event generator. A series of improvements in MADGRAPH5 further enhance its efficiency and speed. We validate this implementation for realistic partonic processes and find significant gains from using modern machine learning in event generators.



Optimizing integration paths [invertible networks]

- find optimal integration paths
- learn variable transformation

→ Theory-integrator



SciPost

SciPost Phys. 12, 129 (2022)

Targeting multi-loop integrals with neural networks

Ramon Winterhalder^{1,2,3}, Vitaly Magrya⁴, Emilia Villa⁵, Stephen P. Jones², Matthias Kerler⁶, Anja Baier¹, Gudrun Heinrich^{3,4} and Tilman Plehn^{1,2}

¹ Institut für Theoretische Physik, Universität Heidelberg, Germany

² HEPA - Heidelberg Karlsruhe Strategic Partnership, Heidelberg University, Karlsruhe Institute of Technology (KIT), Germany

³ Centre for Cosmology, Particle Physics and Phenomenology (CP3), Université catholique de Louvain, Belgium

⁴ Institut für Theoretische Physik, Karlsruher Institut für Technologie, Germany

⁵ Institute for Particle Physics Phenomenology, Durham University, UK

⁶ Institut für Astroteilchenphysik, Karlsruher Institut für Technologie, Germany

Abstract

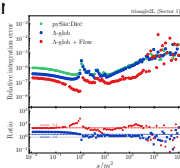
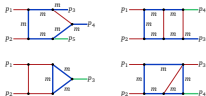
Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct, the numerical precision for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of optimized, global complex shifts and a normalizing flow. They can lead to a significant gain in precision.



Optimizing integration paths [invertible networks]

- find optimal integration paths
- learn variable transformation

→ Theory-integrator



Navigating string landscape [reinforcement learning]

- searching for viable vacua
- high dimensions, unknown global structure

→ Model space sampling

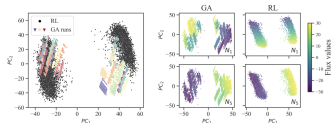


Figure 1: *Left:* Cluster structure in dimensionally reduced flux samples for RL and 25 GA runs (PCA on all samples of GA and RL). The colors indicate individual GA runs. *Right:* Dependence on flux (input) values (N_1 and N_2 respectively) in relation to principal components for a PCA fit of the individual output of GA and RL.

Targeting multi-loop integrals with neural networks

Ramon Winterhalder^{1,2,3}, Vitaly Magyari⁴, Emilio Villa⁵, Stephen P. Jones², Matthias Kerzer⁶, Anja Baier⁷, Gudrun Heinrich^{3,4} and Tilman Plehn^{1,2}

- Institut für Theoretische Physik, Universität Heidelberg, Germany
- HEiKA - Heidelberg Karlsruhe Strategic Partnership, Heidelberg University, Karlsruhe Institute of Technology (KIT), Germany
- Centre for Cosmology, Particle Physics and Phenomenology (CP3), Université catholique de Louvain, Belgium
- Institut für Theoretische Physik, Karlsruher Institut für Technologie, Germany
- Institute for Particle Physics Phenomenology, Durham University, UK
- Institut für Astroteilchenphysik, Karlsruher Institut für Technologie, Germany

Abstract

Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct, the numerical precision for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of optimized, global complex shifts and a normalizing flow. They can lead to a significant gain in precision.

Probing the Structure of String Theory Vacua with Genetic Algorithms and Reinforcement Learning

Alex Cude
University of Amsterdam
a.a.c.cude@uva.nl

Seon Krippendorfer
Arnold Sommerfeld Center for Theoretical Physics
LMU Munich
seon.krippendorfer@physik.uni-muenchen.de

Andreas Schachner
Centre for Mathematical Sciences
University of Cambridge
as273@cam.ac.uk

Gary Shiu
University of Wisconsin-Madison
shiu@physics.wisc.edu

Abstract

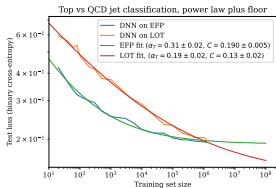
Identifying string theory vacua with desired physical properties at low energies requires searching through high-dimensional solution spaces – collectively referred to as the string landscape. We highlight that this search problem is amenable to reinforcement learning and genetic algorithms. In the context of this vacua, we are able to reveal novel features (suggesting previously unobserved symmetries) in the string theory solutions required for properties such as the string coupling. In order to identify these features robustly, we combine results from both search methods, which we argue is imperative for reducing sampling bias.



Scaling laws for classification networks (statistical learning)

- networks are complex systems
- training as statistical process

→ Now solving problems



Julian Bates^{*}
Independent Researcher
Oakland, CA 94612
jrbates_bates@gmail.com

Yusufan Kuhn
Center for Artificial Intelligence Innovations and
Department of Physics
University of Illinois Urbana-Champaign
Urbana, IL 61801
ykun@illinois.edu

ABSTRACT

We demonstrate the emergence of scaling laws in the benchmark top versus QCD jet classification problem in collider physics. Six distinct physically-motivated classifiers exhibit power-law scaling of the binary cross-entropy test loss as a function of training set size, with distinct power-law indices. This result highlights the importance of comparing classifiers as a function of dataset size rather than for a fixed training set, as the optimal classifier may change considerably as the dataset is scaled up. We speculate on the interpretation of our results in terms of previous models of scaling laws observed in natural language and image datasets.

Collective variables of neural networks: empirical time evolution and scaling laws

Svenja Toyoy
Institute for Computational Physics
University of Stuttgart
Stuttgart, Germany, 70569
stoyoy@ip.uni-stuttgart.de

Sven Krüppendorf
Cavendish Laboratory and DAMTP
University of Cambridge
Cambridge, United Kingdom, CB3 0WA
s.kruepp@cam.ac.uk

Michael Spannowsky
Institute for Particle Physics Phenomenology
Department of Physics
Durham University
Durham, DH1 1TA, U.K.

Konstantin Nikolov
Institute for Computational Physics
University of Stuttgart
Stuttgart, Germany, 70569

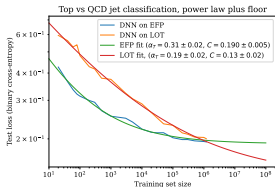
Christian Helm
Institute for Computational Physics
University of Stuttgart
Stuttgart, Germany, 70569



Scaling laws for classification networks [statistical learning]

- networks are complex systems
- training as statistical process

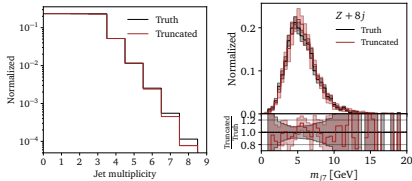
→ Now solving problems



Extrapolating transformers

- train on QCD jet radiation
- learn to generate universal patterns

→ Extrapolation at work



Julian Bates¹
 Independent Researcher
 Oakland, CA 94617
 julian_bates@protonmail.com

Vincent Kuhn
 Center for Artificial Intelligence Innovation and
 Department of Physics
 University of Illinois Urbana-Champaign
 Urbana, IL 61801
 vkuhn@illinois.edu

ABSTRACT

We demonstrate the emergence of scaling laws in the benchmark top versus QCD jet classification problem in collider physics. We derive physically-motivated conditions which power-law scaling of the binary cross-entropy test loss as a function of training set size, with distinct power-law indices. This reveals the importance of computing conditions as a function of dataset size rather than for a fixed training set, as the optimal condition may change considerably as the dataset is scaled up. We speculate on the interpretation of our results in terms of previous models of scaling laws observed in natural language and image datasets.

Collective variables of neural networks: empirical time evolution and scaling laws

Stefan Ewerth
 Institute for Computational Physics
 University of Stuttgart
 Stuttgart, Germany, 70569
 steyerth@ipc.uni-stuttgart.de

Sven Krippendorff
 Cavendish Laboratory and DAMTP
 University of Cambridge
 Cambridge, United Kingdom, CB3 0HA
 s.k13@cam.ac.uk

Michael Spannowsky
 Institute for Particle Physics Phenomenology
 Department of Physics
 Durham University
 Durham, DH1 1TA, U.K.

Konstantin Nikolov
 Institute for Computational Physics
 University of Stuttgart
 Stuttgart, Germany, 70569

Christian Heitsch
 Institute for Computational Physics
 University of Stuttgart
 Stuttgart, Germany, 70569

Extrapolating Jet Radiation with Autoregressive Transformers

Anja Bayle^{1,2}, François Chardon³, Javier Marín Villadamiñg³,
 Aydoğdu On¹, Tilman Plehn^{1,4}, and Jonas Spitzer¹

¹ Institut für Theoretische Physik, Universität Heidelberg, Germany

² LPNHE, Sorbonne Université, Université Paris Cité, CNRS/IN2P3, Paris, France

³ Meta FAIR, CERMICS - Ecole des Ponts

⁴ Interdisciplinary Center for Scientific Computing (WZR), Universität Heidelberg, Germany

December 17, 2024

Abstract

Generative networks are an exciting tool for fast LHC event generation. Usually, they are used to generate configurations with a fixed number of particles. Autoregressive transformers allow us to generate events with variable numbers of particles, very much in line with the physics of QCD jet radiation. We show how they can learn a factorized likelihood for jet radiation and extrapolation in terms of the number of generated jets. For this extrapolation, bootstrapping training data and training with modifications of the likelihood loss can be used.



An IPPP story...

- $gg \rightarrow ZZ$ [Bishara & Montull (2019)] \rightarrow BDTs as a start
- $e^+e^- \rightarrow 5 \text{ jets}$ [Badger, Bullock (2020)] \rightarrow ensembles and K -factors
- $gg \rightarrow \gamma\gamma g(g)$ [Aylett-Bullock, Badger, Moodie (2021)] \rightarrow speed gain 10^4
- $e^+e^- \rightarrow 5 \text{ jets}$ [Maitre & Truong (2021)] \rightarrow Catani-Seymour coefficients
- $e^+e^- \rightarrow 5$ [Maitre & Truong (2023)] \rightarrow antenna functions
- $gg \rightarrow \gamma\gamma g(g)$ [Badger+Heidelberg (2024)] \rightarrow boosted training
- $t\bar{t}H$ NNLO tests [Breso, Heinrich, Magerya, Olsson] \rightarrow race for best performance

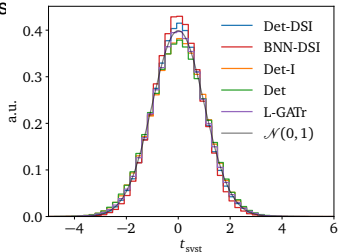


An IPPP story...

- $gg \rightarrow ZZ$ [Bishara & Montull (2019)] \rightarrow BDTs as a start
- $e^+e^- \rightarrow 5 \text{ jets}$ [Badger, Bullock (2020)] \rightarrow ensembles and K -factors
- $gg \rightarrow \gamma\gamma g(g)$ [Aylett-Bullock, Badger, Moodie (2021)] \rightarrow speed gain 10^4
- $e^+e^- \rightarrow 5 \text{ jets}$ [Maitre & Truong (2021)] \rightarrow Catani-Seymour coefficients
- $e^+e^- \rightarrow 5$ [Maitre & Truong (2023)] \rightarrow antenna functions
- $gg \rightarrow \gamma\gamma g(g)$ [Badger+Heidelberg (2024)] \rightarrow boosted training
- $t\bar{t}H$ NNLO tests [Breso, Heinrich, Magerya, Olsson] \rightarrow race for best performance

Learned uncertainties [Bahl, Elmer, Favaro, Haussmann, TP, Winterhalder]

- systematic: (added) noise, expressivity, data representation learned by heteroscedastic loss and BNNs
- statistical: too little training data learned by BNN or repulsive ensembles
- calibration of learned uncertainties?

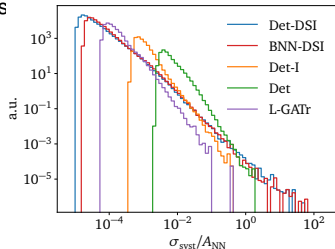


An IPPP story...

- $gg \rightarrow ZZ$ [Bishara & Montull (2019)] \rightarrow BDTs as a start
- $e^+e^- \rightarrow 5 \text{ jets}$ [Badger, Bullock (2020)] \rightarrow ensembles and K -factors
- $gg \rightarrow \gamma\gamma g(g)$ [Aylett-Bullock, Badger, Moodie (2021)] \rightarrow speed gain 10^4
- $e^+e^- \rightarrow 5 \text{ jets}$ [Maitre & Truong (2021)] \rightarrow Catani-Seymour coefficients
- $e^+e^- \rightarrow 5$ [Maitre & Truong (2023)] \rightarrow antenna functions
- $gg \rightarrow \gamma\gamma g(g)$ [Badger+Heidelberg (2024)] \rightarrow boosted training
- $t\bar{t}H$ NNLO tests [Breso, Heinrich, Magerya, Olsson] \rightarrow race for best performance

Learned uncertainties [Bahl, Elmer, Favaro, Haussmann, TP, Winterhalder]

- systematic: (added) noise, expressivity, data representation learned by heteroscedastic loss and BNNs
 - statistical: too little training data learned by BNN or repulsive ensembles
 - calibration of learned uncertainties?
- \rightarrow Path to 10^{-5} accuracy



Energy calibration with uncertainties [ATLAS + Vogel, 2412.04370]

- interpretable calorimeter phase space x
- learned calibration function

$$\mathcal{R}^{\text{BNN}}(x) \pm \Delta \mathcal{R}^{\text{BNN}}(x) \approx \frac{E^{\text{obs}}(x)}{E^{\text{dep}}(x)}$$

- **uncertainties:** noise in data
network expressivity
data representation ...



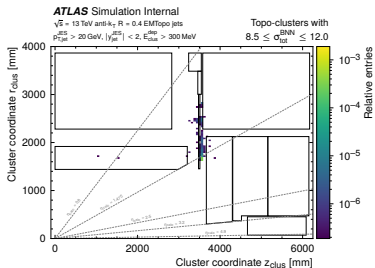
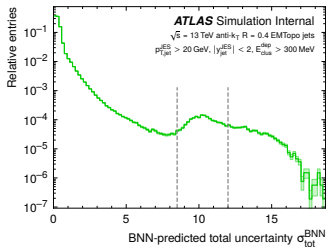
Energy calibration with uncertainties [ATLAS + Vogel, 2412.04370]

- interpretable calorimeter phase space x
- learned calibration function

$$\mathcal{R}^{\text{BNN}}(x) \pm \Delta \mathcal{R}^{\text{BNN}}(x) \approx \frac{E^{\text{obs}}(x)}{E^{\text{dep}}(x)}$$

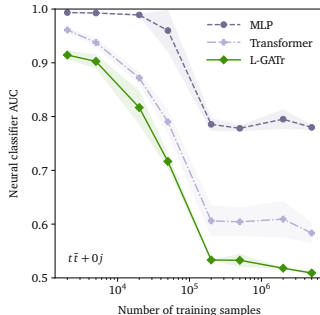
- **uncertainties:** noise in data
network expressivity
data representation ...

→ Understand (simulated) detector

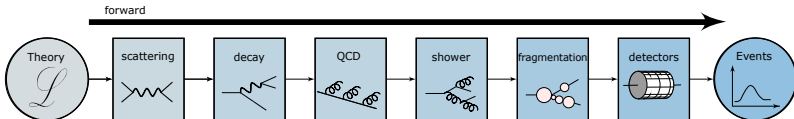


Simulations, MadNIS, calorimeters,...

- learn phase space density
fast sampling Gaussian \rightarrow phase space
- Variational Autoencoder
 \rightarrow low-dimensional physics
- Generative Adversarial Network
 \rightarrow generator trained by classifier
- Normalizing Flow/Diffusion
 \rightarrow (bijective) mapping
- JetGPT, ViT
 \rightarrow non-local structures
- Equivariant L-GATr
 \rightarrow Lorentz symmetry for efficiency



\rightarrow [Equivariant transformer CFM...](#) [Maitre, Ngairangbam, Spannowsky,...]



Bayesian generative networks [Bellagente, Haussmann, Luchmann, TP]

- network weight distributions for density
- sampling phase space events with error bars on weights
- learned density & uncertainty reflecting network learning?

→ **Generative networks like fitted densities**



Bayesian generative networks [Bellagente, Haussmann, Luchmann, TP]

- network weight distributions for density
- sampling phase space events with error bars on weights
- learned density & uncertainty reflecting network learning?

→ **Generative networks like fitted densities**

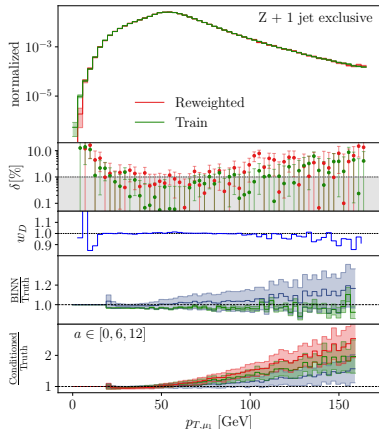
Z+jets events [Heimel, Vent...]

- per-cent accuracy on density
- statistical uncertainty from BNN
- systematics in training data

$$w = 1 + a \left(\frac{p_{T,j_1} - 15 \text{ GeV}}{100 \text{ GeV}} \right)^2$$

sampling a conditionally

→ **Precision and uncertainty control**



Compare generated with training data

- regression accuracy $\Delta = (A_{\text{data}} - A_{\theta})/A_{\text{data}}$
- harder for generation, unsupervised density
classify training vs generated events $D(x)$
learned density ratio [Neyman-Pearson]

$$w(x_i) = \frac{D(x_i)}{1 - D(x_i)} = \frac{\rho_{\text{data}}(x_i)}{\rho_{\text{model}}(x_i)}$$

→ Test ratio over phase space



Compare generated with training data

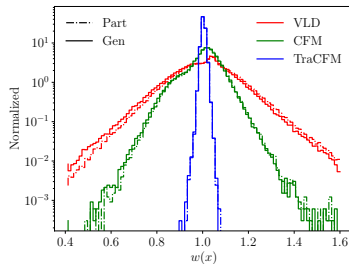
- regression accuracy $\Delta = (A_{\text{data}} - A_{\theta}) / A_{\text{data}}$
- harder for generation, unsupervised density classify training vs generated events $D(x)$
learned density ratio [Neyman-Pearson]

$$w(x_i) = \frac{D(x_i)}{1 - D(x_i)} = \frac{\rho_{\text{data}}(x_i)}{\rho_{\text{model}}(x_i)}$$

→ Test ratio over phase space

Progress in NN-generators

- any generative AI task
 - compare different architectures
 - accuracy from width of weight distribution
 - tails indicating failure mode
- Systematic performance test



Transforming LHC physics

Number of searches

- optimal inference: signal and background simulations
- CPU-limitation for many signals?

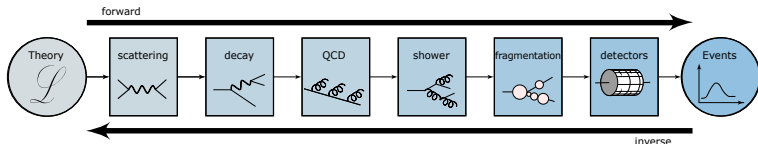
Optimal analyses

- theory limiting many analyses
- include predictions not in event generators

Public LHC data

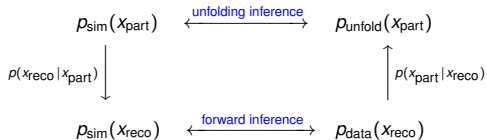
- common lore:
LHC data too complicated for amateurs
- in truth:
hard scattering and decay simulations public
BSM physics not in hadronization and detector

→ **Unfold to suitable level**



Basic structure [Butter, Köthe, TP, Winterhalder]

- four phase space distributions



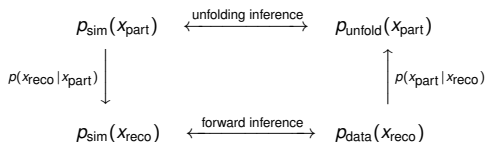
- learn conditional probabilities from $(x_{\text{part}}, x_{\text{reco}})$ [forward-inverse symmetric]

→ ML for unbinned and high-dimensional unfolding?



Basic structure [Butter, Köthe, TP, Winterhalder]

- four phase space distributions

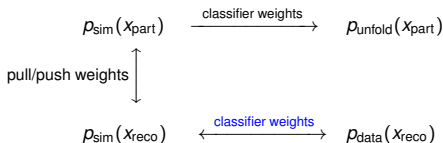


- learn conditional probabilities from $(x_{\text{part}}, x_{\text{reco}})$ [forward-inverse symmetric]

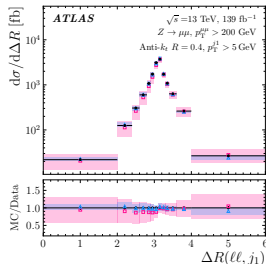
→ ML for unbinned and high-dimensional unfolding?

OmniFold [Andreassen, Komiske, Metodiev, Nachman, Thaler + ATLAS]

- learn $\rho_{\text{sim}}(x_{\text{reco}}) \leftrightarrow \rho_{\text{data}}(x_{\text{reco}})$ [Neyman-Pearson]
- reweight $\rho_{\text{sim}}(x_{\text{part}}) \rightarrow \rho_{\text{unfold}}(x_{\text{part}})$

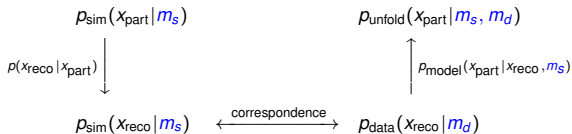


→ Z+jets in 24D [ATLAS]



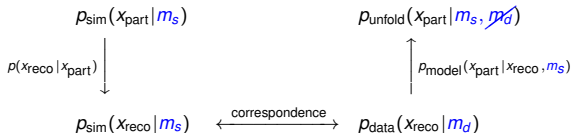
A challenge [Favaro, Kogler, Paasch, Palacios Schweitzer, TP, Schwarz]

- first measure m_t in unfolded data
then unfold full kinematics
- model dependence: simulation m_s vs data m_d



A challenge [Favaro, Kogler, Paasch, Palacios Schweitzer, TP, Schwarz]

- first measure m_t in unfolded data
then unfold full kinematics
- complete training bias $m_d \rightarrow m_s$ [too bad to reweight]

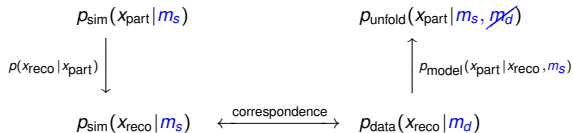


- 1 weaken bias by training on m_s -range
- 2 strengthen data by including batch-wise $m_d \sim M_{jjj} \in x_{\text{reco}}$



A challenge [Favaro, Kogler, Paasch, Palacios Schweitzer, TP, Schwarz]

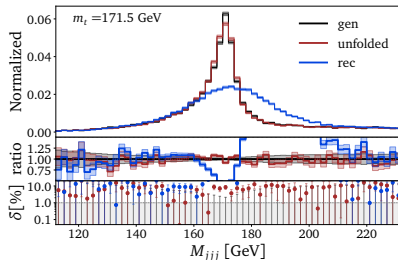
- first measure m_t in unfolded data
then unfold full kinematics
- complete training bias $m_d \rightarrow m_S$ [too bad to reweight]



- 1 weaken bias by training on m_S -range
- 2 strengthen data by including batch-wise $m_d \sim M_{jjj} \in X_{\text{reco}}$

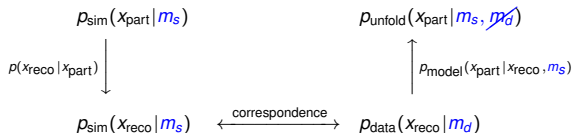
Preliminary unfolding results [TraCFM]

- 4D for calibrated mass measurement



A challenge [Favaro, Kogler, Paasch, Palacios Schweitzer, TP, Schwarz]

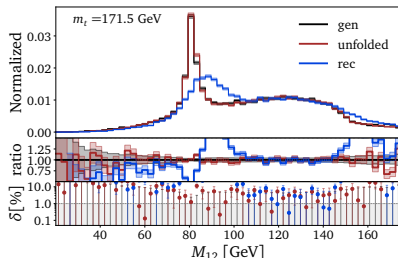
- first measure m_t in unfolded data
then unfold full kinematics
- complete training bias $m_d \rightarrow m_S$ [too bad to reweight]



- 1 weaken bias by training on m_S -range
- 2 strengthen data by including batch-wise $m_d \sim M_{jjj} \in X_{\text{reco}}$

Preliminary unfolding results [TraCFM]

- 4D for calibrated mass measurement
 - 12D published data
- CMS data next



Developing ML for the best science

- 1 just another numerical tool for a numerical field
- 2 completely transformative new language
 - driven by money from data science and medical research
 - physics should be leading scientific AI
 - 1000 Einsteins...
 - ...improving established tools
 - ...developing new tools for established tasks
 - ...transforming through new ideas

→ You can be the golden generation!

Modern Machine Learning for LHC Physicists

Tilman Plehn^a, Anja Butter^{a,b}, Barry Dillon^a,
 Theo Heime^a, Claudius Krause^c, and Ramon Winterhalder^d

^a Institut für Theoretische Physik, Universität Heidelberg, Germany

^b LPNHE, Sorbonne Université, Université Paris Cité, CNRS/IN2P3, Paris, France

^c HEPHY, Austrian Academy of Sciences, Vienna, Austria

^d CP3, Université catholique de Louvain, Louvain-la-Neuve, Belgium

March 19, 2024

Abstract

Modern machine learning is transforming particle physics fast, bullying its way into our numerical tool box. For young researchers it is crucial to stay on top of this development, which means applying cutting-edge methods and tools to the full range of LHC physics problems. These lecture notes lead students with basic knowledge of particle physics and significant enthusiasm for machine learning to relevant applications. They start with an LHC-specific motivation and a non-standard introduction to neural networks and then cover classification, unsupervised classification, generative networks, and inverse problems. Two themes defining much of the discussion are well-defined loss functions and uncertainty-aware networks. As part of the applications, the notes include some aspects of theoretical LHC physics. All examples are chosen from particle physics publications of the last few years.¹

:2211.01421v2 [hep-ph] 17 Mar 2024

