(Particle) dark matter theory

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The problem isn't discovering dark matter...







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Existence of dark matter on astrophysical and cosmological scales is known and well characterised



Simulated galaxy

Extended dark matter halo

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The problem is characterising dark matter...



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Dark Matter Particle (X^0)

$$X^0$$
 mass: $m = ?$
 X^0 spin: $J = ?$
 X^0 parity: $P = ?$
 X^0 lifetime: $\tau = ?$
 X^0 interactions with normal matter?



Cosmology and Particle Physics give clues

Cosmology

$\Omega_{\rm DM} h^2 = 0.120 \pm 0.001$

Explaining this value suggests dark and visible matter interactions are generic

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Particle Physics



Informs and limits the possible interactions





candidates)



Particle DM landscape: many possibilities



Cooley et al arXiv:2209.07426



Cosmic Frontier's recommendation:

- **Delve deep** (cover high priority targets e.g., WIMPs)
- Search wide (explore as much DM parameter space as possible)





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High priority target: WIMPs

- (Probably) the most studied DM candidate
- 'Natural' mechanism to produce the observed relic abundance 'WIMP-miracle'
- One-to-one correspondence with a single particle physics input:

Annihilation cross-section with SM particles

- Embedded in theories that alleviate the 'hierarchy problem' (SUSY, etc)
 - Idea of 'Natural WIMPs' but 'Unnatural WIMPs' also work







High priority target: WIMPs

- The 'single particle physics' input focussed minds
- Produced many complementary searches:

Direct detection









'Delve deep'

Indirect detection

Break it



Collider



Make it









High priority target: WIMPs

- Also somewhat of an industry for theorists to probe models from many directions

Direct detection





Indirect detection



Break it









WIMPs: What has direct detection taught us so far?



'Delve deep'

@~10⁻³⁹ cm²: Z-mediated (spin-independent) interactions: excluded

@~10⁻⁴⁴ cm² and below: Higgsmediated interactions: putting pressure on this channel





WIMPs: What has direct detection taught us so far?



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'Delve deep'

@~10⁻³⁹ cm²: *Z*-mediated (spin-independent) interactions: excluded

 $@~10^{-44}$ cm² and below: Higgsmediated interactions: putting pressure on this channel

...and complementary to collider probes of the Higgs coupling





WIMPs: to come?



'Delve deep'

- XLZD (LXe): discovery machine for natural WIMPs down to the 'neutrino fog'
- Liquid Argon experiments also on track to cover similar parameter space

(see dedicated talks)





To come: non-natural WIMPs

 10^{-44}

 10^{-45}

 $\left[\frac{2}{\sigma^{2}} \right]^{2}$ 10^{-46}

 10^{-47}

- XLZD(ARWIN): will finally probe the 'minimal dark matter' family of models

Cirelli et al, arXiv:0512090

- Minimal dark matter: introduce a stable, weakly interacting particle as part of an SU(2) multiplet



'Delve deep'



Aalbers, JPhyD arXiv:2203.02309





Beyond Standard Model Mediators

Isotopes with unpaired spin give sensitivity to interactions that couple to spin [e.g, 129Xe and 131Xe isotopes @~50% abundance]



Axial-vector & Axi

LXe-TPC



Can test a menagerie of dark matter interactions that arise through more exotic mediators

	Abbrev.	Operator	Dimension	Coherent	Coefficients
		(\mathcal{Q})		enhancement	
ipole	_	$\bar{\chi}\sigma^{\mu u}\chi F_{\mu u}$	5	Partial	C_F
pole	-	$ar{\chi}\sigma^{\mu u}\chi ilde{F}_{\mu u}$	5	Yes	$ ilde{C}_F$
ctor	VV	$ar{\chi}\gamma^\mu\chiar{q}\gamma_\mu q$	6	Yes	$C_{u,d,s}^{VV}$
Vector	AV	$ar{\chi}\gamma^{\mu}\gamma_5\chiar{q}\gamma_{\mu}q$	6	Yes	$C^{AV}_{u,d}$
nsor	TT	$\bar{\chi}\sigma^{\mu u}\chi\bar{q}\sigma_{\mu u}q$	6	Yes	$C_{u,d,s}^{TT}$
Tensor	\widetilde{TT}	$\bar{\chi}\sigma^{\mu u}i\gamma_5\chi\bar{q}\sigma_{\mu u}q$	6	Yes	$ ilde{C}_{u,d,s}^{TT}$
alar	SS	$ar{\chi}\chi m_qar{q}q$	7	Yes	$C_{u,d,s}^{SS}$
on	S_g	$lpha_s \bar{\chi} \chi G^a_{\mu u} G^{\mu u}_a$	7	Yes	C_g^S
- gluon	$ ilde{S}_g$	$\alpha_s \bar{\chi} i \gamma_5 \chi G^a_{\mu u} G^{\mu u}_a$	7	Yes	$ ilde{C}_g^S$
\Im Scalar	PS	$ar{\chi}i\gamma_5\chi m_qar{q}q$	7	Yes	$C^{PS}_{u,d,s}$
	-	$ar{\chi}\gamma_\mu i\partial_ u\chiar{ heta}^{\mu u}_{q(g)}$	8	Yes	$C_{u,d,s,g}^{(2)}$
ial-vector	AA	${ar \chi} \gamma^\mu \gamma_5 \chi {ar q} \gamma_\mu \gamma_5 q$	6	No	$C_{u,d,s}^{AA}$





More radical ideas: ultra-heavy dark matter

- Hidden sectors enrich the possibilities for dark matter
- Consider dark particles that self interact through an attractive force:

particles may bind together to form large composite states

- Composite size on the strength and range of the force, and on the presence or absence of bottlenecks

McCullough et al, arXiv: 1406.2276 Wise et al, arXiv:1411.1772 Hardy et al, arXiv:1411.3739 Zurek et al, arXiv:1707.02313 + many, many more

'Search wide'







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Direct searches for Planck scale candidates



- Hidden sectors enrich the possibilities for sub-GeV dark matter
- Relic density from:
 - variations on the WIMP mechanism afforded by portal interactions (SIMP, ELDER, ...)
 - Freeze-in mechanism
- Exciting developments using existing detectors and proposing new technologies



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new technologies	10-
See talks tomorrow morning	

Landscape of sub-GeV dark matter

'Search wide'



Essig et al arXiv: 2203.08297





The search for particle dark matter continues unabated...

Current strategy adopted by the international community summarised as 'delve deep and search wide'

Direct detection searches allow for the test of the full mass range of *particle* dark matter: Now: in the regime where experiments are probing Higgs-DM interactions Future: definitively probes 'natural-WIMPs' to the neutrino floor, start to probe minimal DM models (loop-induced interactions)

As we will hear: UK activities are at the forefront of the field with proposals for 'broadband' searches (LZ, DarkSide, XLZD) and more focussed searches (DarkSPHERE, SOLAIRE, UltraDark...)

Summary

• Testing richer hidden sector extensions with candidates from the keV to Planck scale

