



UNIVERSITÀ  
DEGLI STUDI  
DI PADOVA



Dipartimento  
di Fisica  
e Astronomia  
Galileo Galilei



# Solving the Strong CP Problem in String-Inspired Theories with Modular Invariance

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Based on **2505.20395** with **F. Feruglio**, **A. Marrone** and **A. Strumia**  
also **2305.08908** with **FF** and **AS** and **2406.01689** with **FF**, **AS** and **M. Parriciatu**

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PIANO NAZIONALE  
DI RIPRESA E RESILIENZA

# Outline

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1. Strong CP problem
2. Existing solutions
3. Modular invariance and global supersymmetry
4. Example model
5. Conclusions

# The strong CP problem

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$$\mathcal{L}_{\text{QCD}} = \bar{q}i\not{D}q - \left[ \bar{q}_R M_q q_L + \text{h.c.} \right] - \frac{1}{4g_3^2} G_{\mu\nu}^a G^{a,\mu\nu} + \frac{\theta_{\text{QCD}}}{32\pi^2} G_{\mu\nu}^a \tilde{G}^{a,\mu\nu}$$

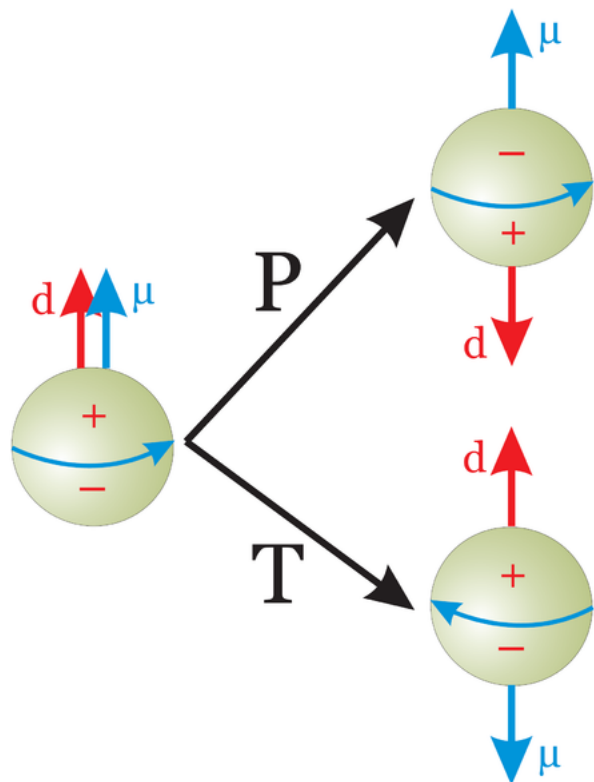
$$\bar{\theta} = \theta_{\text{QCD}} + \arg \det M_q \quad \text{CPV parameter}$$

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Neutron EDM  $d$



$$d = 2.4 \times 10^{-16} \bar{\theta} e \cdot \text{cm} \quad \text{Pospelov, Ritz, hep-ph/9908508v4}$$

$$|d| \leq 1.8 \times 10^{-26} e \cdot \text{cm} \quad (90\% \text{ C.L.}) \quad \text{Abel et al., 2001.11966}$$

$$|\bar{\theta}| \lesssim 10^{-10}$$

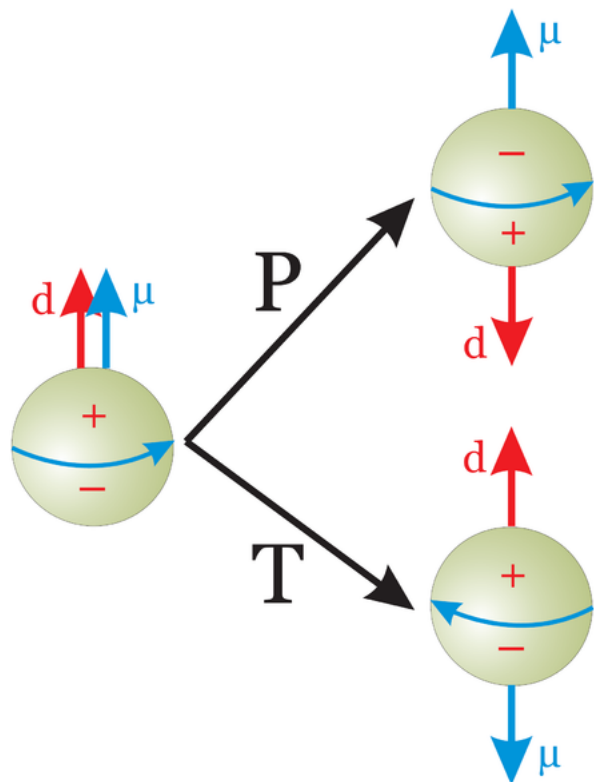
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Why so small???

... and the CPV phase in the CKM matrix  $\delta_{\text{CKM}} \approx 1.2$

# Solution 1: the Axion

---

Promote  $\bar{\theta}$  to a dynamical scalar field  $a$ , the **axion**, which washes out CP violation in QCD

$$\mathcal{L}_a = \frac{1}{2}(\partial_\mu a)^2 + \frac{1}{32\pi^2} \frac{a}{f_a} G\tilde{G} + \dots$$

$$\bar{\theta} = \frac{\langle a \rangle}{f_a} \quad \text{with} \quad \langle a \rangle = 0$$

New global  $U(1)_{PQ}$

Peccei, Quinn, PRL **38** (1977) 1440; PRD **16** (1977) 1791

- ▶ spontaneously broken  $\Rightarrow$  the axion is a NGB
- ▶ anomalous under QCD  $(\partial_\mu J_{PQ}^\mu \propto G\tilde{G}) \Rightarrow$  the **axion is a pNGB**

Quality problem

- ▶ Corrections of order  $(f_a/M_{Pl})^\#$  from higher-dimensional operators
- ▶  $U(1)_{PQ}$  should be an accidental symmetry in a complete model

# Solution 2: CP (P) is symmetry of UV

- ▶ CP (P) is a symmetry of the UV
- ▶ It is broken spontaneously in such a way that  $\bar{\theta} = 0$  and  $\delta_{\text{CKM}} = \mathcal{O}(1)$

## Nelson—Barr models

Nelson, PLB **136** (1984) 387; Barr, PRL **53** (1984) 329

New heavy vector-like quarks  $Q$  and scalars  $\eta$  with CPV complex VEVs  $\langle \eta \rangle$

$$(\overline{q_R} \ \overline{Q_R}) M_q \begin{pmatrix} q_L \\ Q_L \end{pmatrix} = (\overline{q_R} \ \overline{Q_R}) \begin{pmatrix} y v_H & y' \langle \eta \rangle \\ 0 & \mu \end{pmatrix} \begin{pmatrix} q_L \\ Q_L \end{pmatrix}$$

- ▶ CP is a symmetry  $\Rightarrow \theta_{\text{QCD}} = 0$  and the couplings  $(y, y', \mu)$  are real
- ▶  $\det M_q = y v_H \mu$  is real (and positive)  $\Rightarrow \arg \det M_q = 0$
- ▶ Effective light quark mass matrix depends on multiple  $\langle \eta \rangle \Rightarrow \delta_{\text{CKM}} \neq 0$

## Additional matter, tuning, loop corrections...

Dine, Draper, 1506.05433

# Our solution

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Supersymmetry + Modular invariance + CP

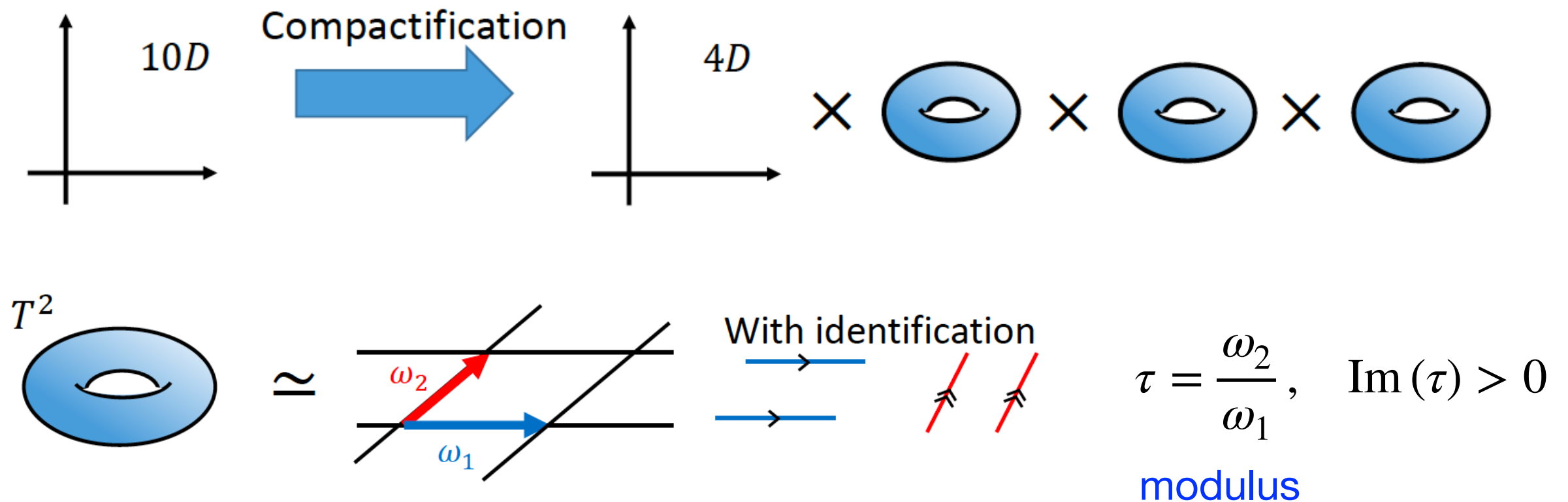


$$\bar{\theta} = 0$$

# Modular invariance

String theory requires extra dimensions

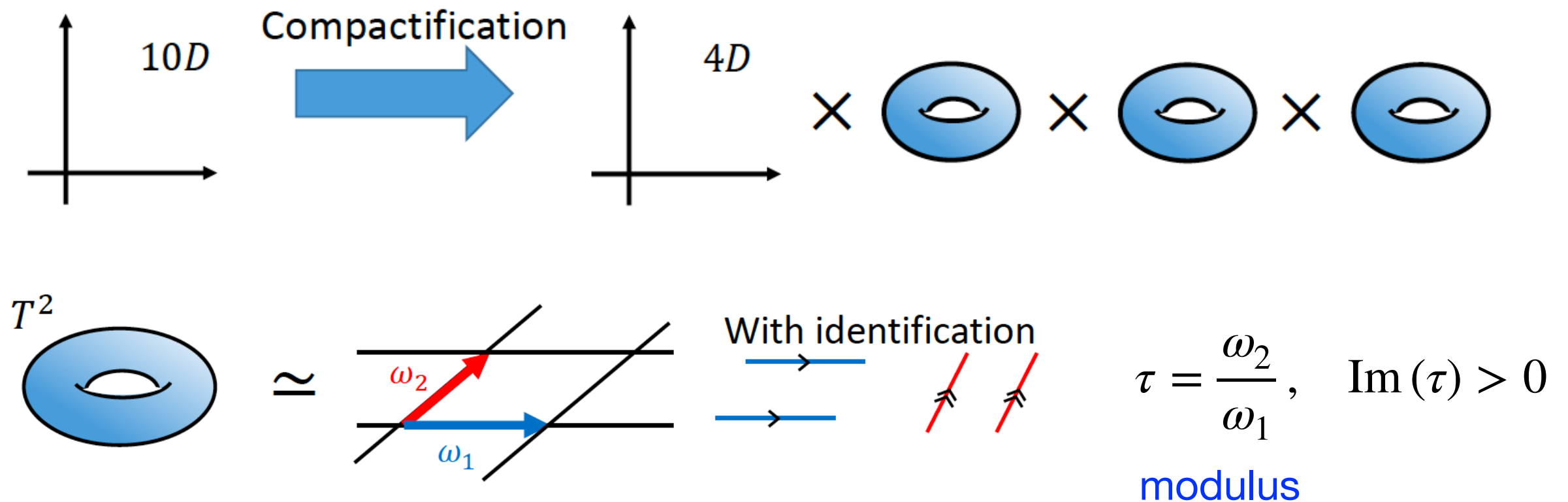
Images: [Takuya H. Tatsuishi](#)



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Lattice is invariant under basis transformations

$$\begin{cases} \omega_2 \rightarrow \omega'_2 = a\omega_2 + b\omega_1 \\ \omega_1 \rightarrow \omega'_1 = c\omega_2 + d\omega_1 \end{cases} \quad a, b, c, d \in \mathbb{Z} \quad ad - bc = 1$$

$$\tau \rightarrow \tau' = \frac{a\tau + b}{c\tau + d}$$

modular transformations

$\tau$  and  $\tau'$  describe the same torus

# Modular group

Homogeneous modular group

$$\Gamma = \langle S, T \mid S^4 = (ST)^3 = I \rangle \cong \mathrm{SL}(2, \mathbb{Z})$$

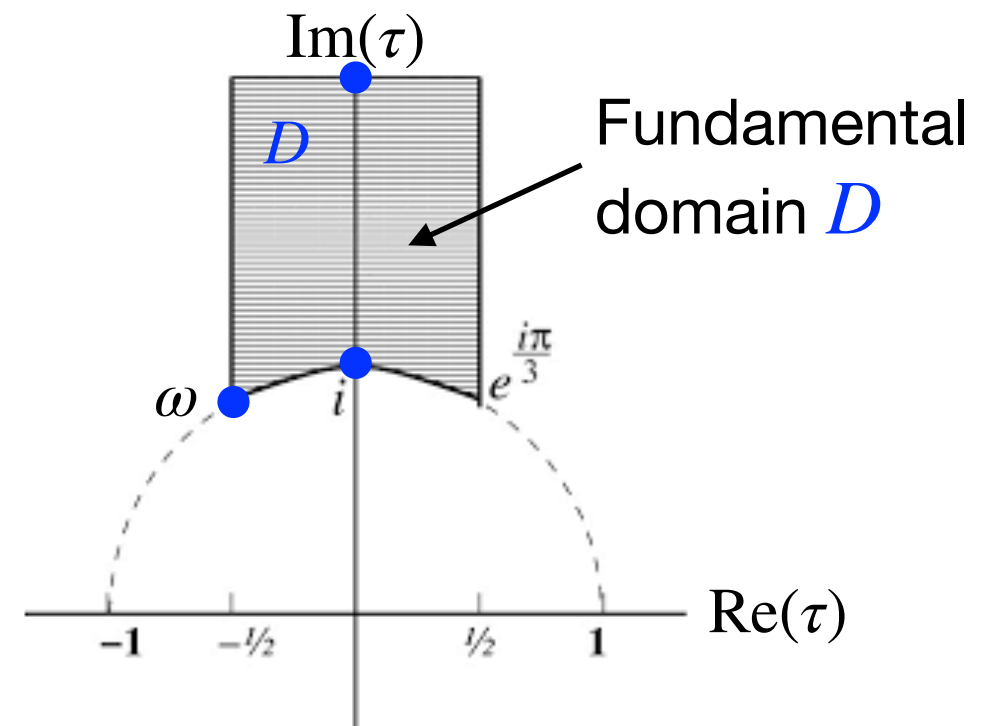
$$S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \quad T = \begin{pmatrix} 1 & 1 \\ 0 & 1 \end{pmatrix}$$

$$\tau \xrightarrow{S} -\frac{1}{\tau}$$

duality

$$\tau \xrightarrow{T} \tau + 1$$

discrete shift symmetry



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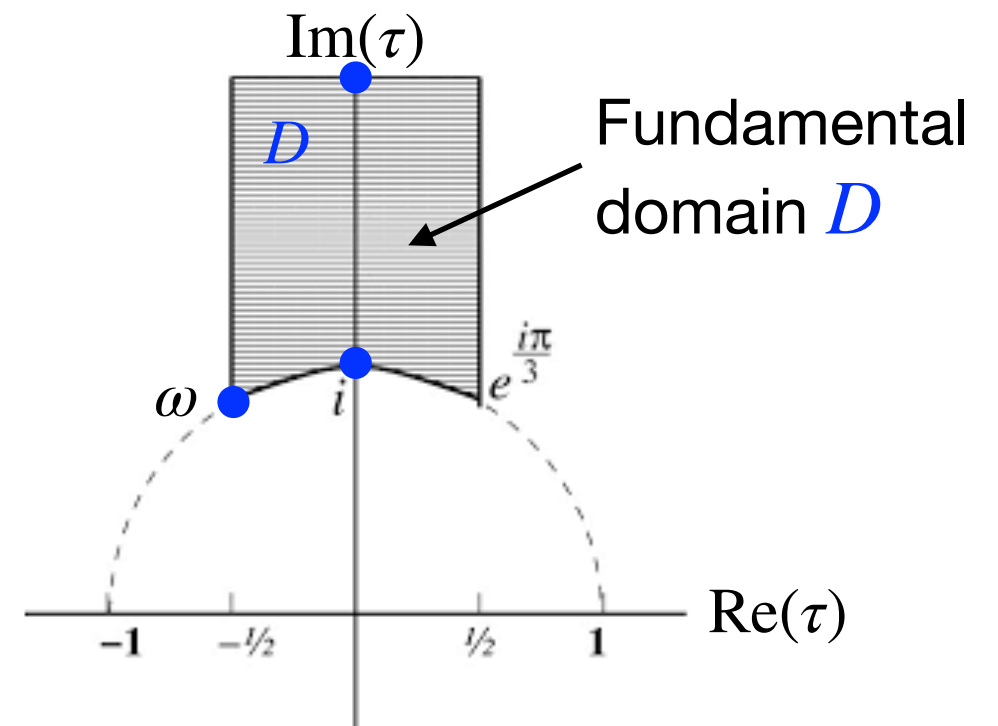
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## Fixed points

$$\triangleright \tau = i : \quad i \xrightarrow{S} -\frac{1}{i} = i \quad \Rightarrow \quad Z_4^S$$

$$\triangleright \tau = \omega \equiv e^{\frac{2\pi i}{3}} : \quad \omega \xrightarrow{ST} -\frac{1}{\omega + 1} = \omega \quad \Rightarrow \quad Z_3^{ST} \times Z_2^{S^2}$$

$$\triangleright \tau = i\infty : \quad i\infty \xrightarrow{T} i\infty + 1 = i\infty \quad \Rightarrow \quad Z^T \times Z_2^{S^2}$$

# Modular forms

---

Holomorphic functions on  $\mathcal{H} = \{\tau \in \mathbb{C} : \operatorname{Im}(\tau) > 0\}$  transforming under  $\Gamma$  as

$$f(\gamma\tau) = (c\tau + d)^k f(\tau), \quad \gamma \in \Gamma$$

$k$  is **weight**, a non-negative even integer

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Normalised Eisenstein series

$$E_k(\tau) = \frac{1}{2\zeta(k)} \sum_{(m,n) \neq (0,0)} \frac{1}{(m + n\tau)^k}$$

Each modular form of weight  $k$  can be written as a polynomial in  $E_4$  and  $E_6$

$$f(\tau) = \sum_{a,b \geq 0} c_{ab} E_4^a(\tau) E_6^b(\tau) \quad \text{with} \quad 4a + 6b = k$$

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|                    |   |   |       |       |               |                    |                |                      |
|--------------------|---|---|-------|-------|---------------|--------------------|----------------|----------------------|
| Modular weight $k$ | 0 | 2 | 4     | 6     | 8             | 10                 | 12             | 14                   |
| Modular forms      | 1 | – | $E_4$ | $E_6$ | $E_8 = E_4^2$ | $E_{10} = E_4 E_6$ | $E_4^3, E_6^2$ | $E_{14} = E_4^2 E_6$ |

# Supersymmetry

---

$\mathcal{N} = 1$  global SUSY action

$$\mathcal{L} = \int d^2\theta d^2\bar{\theta} K(e^{2V}\Phi, \Phi^\dagger) + \left[ \int d^2\theta W(\Phi) + \frac{1}{16} \int d^2\theta f_a \mathcal{W}^a \mathcal{W}^a + \text{h.c.} \right]$$

Kähler potential  $K$   
(kinetic terms,  
gauge interactions)

Superpotential  $W$   
(Yukawa interactions)

Gauge kinetic functions  $f_a$   
$$f_3 = \frac{1}{g_3^2} - i \frac{\theta_{\text{QCD}}}{8\pi^2}$$

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$$\bar{\theta} = -8\pi^2 \text{Im} f_3 + \arg \det(Y^u Y^d) = \arg \left[ e^{-8\pi^2 f_3} \det(Y^u Y^d) \right]$$

- $M_q = M_u \oplus M_d = v_u Y^u \oplus v_d Y^d$  with  $v_{u,d} = \langle H_{u,d} \rangle$  real
- $\bar{\theta}$  does not depend on  $K$  Hiller, Schmaltz, hep-ph/0105254

# Modular-invariant SUSY theories

Under modular transformations  $\gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in \Gamma$  Ferrara et al., PLB **225** (1989) 363  
Ferrara et al., PLB **233** (1989) 147  
Feruglio, 1706.08749

$$\left\{ \begin{array}{l} \tau \rightarrow \frac{a\tau + b}{c\tau + d} \\ S \rightarrow S \\ \Phi \rightarrow (c\tau + d)^{-k_\Phi} \Phi \\ V \rightarrow V \end{array} \right.$$

modulus is promoted to a (dimensionless) superfield  
 dilaton  
matter supermultiplets  $k_\Phi$  is modular weight of  $\Phi$   
 vector supermultiplets

Modular symmetry acts non-linearly on field space

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Modular symmetry acts **non-linearly** on field space

**Modular invariance** of the action requires

$$\left\{ \begin{array}{l} K(\tau, \Phi, \tau^\dagger, \Phi^\dagger) \rightarrow K(\tau, \Phi, \tau^\dagger, \Phi^\dagger) + f_K(\tau, \Phi) + \bar{f}_K(\tau^\dagger, \Phi^\dagger) \\ W(\tau, \Phi) \rightarrow W(\tau, \Phi) \\ f_a(S, \tau) \rightarrow f_a(S, \tau) + \frac{1}{8\pi^2} \sum_{\Phi} 2T_a(\Phi) k_\Phi \ln(c\tau + d) \end{array} \right. \quad T_a(\Phi) \text{ is Dynkin index of } \Phi \text{ gauge rep}$$

Konishi, Shizuya, Nuovo Cim. A **90** (1985) 111

anomaly term

Arkani-Hamed, Murayama, hep-th/9707133

# Modular-invariant SUSY theories

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Minimal Kähler potential

$$K = -h^2 \ln(-i\tau + i\tau^\dagger) + \sum_{\Phi} \frac{\Phi^\dagger e^{2V} \Phi}{(-i\tau + i\tau^\dagger)^{k_\Phi}}$$

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Superpotential

$$W = U_i^c Y_{ij}^u(\tau) Q_j H_u + D_i^c Y_{ij}^d(\tau) Q_j H_d + E_i^c Y_{ij}^e(\tau) L_j H_d + \frac{1}{2\Lambda_L} L_i C_{ij}^\nu(\tau) L_j H_u H_u$$

$\tau$ -dependent Yukawa couplings (and Majorana mass)

$$Y_{ij}^u(\tau) \rightarrow (c\tau + d)^{k_{Y_{ij}^u}} Y_{ij}^u(\tau) \quad \text{with} \quad k_{Y_{ij}^u} = k_{U_i^c} + k_{Q_j} + k_{H_u}$$

(similarly for  $Y^d$ ,  $Y^e$ ,  $C^\nu$ )

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are modular forms!

$$\det(Y^u Y^d) \rightarrow (c\tau + d)^{k_{\det}} \det(Y^u Y^d)$$

$$k_{\det} = \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right) + 3 \left( k_{H_u} + k_{H_d} \right)$$

# Modular-invariant SUSY theories

---

Gauge kinetic functions

$$e^{-8\pi^2 f_a(S,\tau)} \rightarrow (c\tau + d)^{k_{f_a}} e^{-8\pi^2 f_a(S,\tau)} \quad \text{with} \quad k_{f_a} = - \sum_{\Phi} 2T_a(\Phi) k_{\Phi}$$

In particular,

$$k_{f_3} = - \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right)$$

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In particular,

$$k_{f_3} = - \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right)$$

$\bar{\theta}$  becomes a field-dependent quantity

$$\bar{\theta} = \arg A(S, \tau)$$

$$A(S, \tau) \equiv e^{-8\pi^2 f_3(S, \tau)} \det [Y^u(\tau) Y^d(\tau)]$$

$$A(S, \tau) \rightarrow (c\tau + d)^{k_A} A(S, \tau) \quad \text{with} \quad k_A = 3 \left( k_{H_u} + k_{H_d} \right)$$

# Conditions for $\bar{\theta} = 0$

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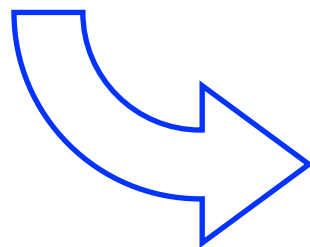
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$$\bar{\theta} = \arg A(S, \tau) = 0$$

independently of the vacuum  $\langle \tau \rangle$

# Structure of $A(S, \tau)$

---

$$A(S, \tau) \equiv e^{-8\pi^2 f_3(S, \tau)} \det [Y^u(\tau) Y^d(\tau)]$$

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- $k_{f_3} = k_{\text{det}} = 0$  (and no singularities in either **gauge** or **Yukawa** factors)

$\theta_{\text{QCD}} = 0$  while  $f_3 = \frac{1}{g_3^2}$  and  $\det(Y^u Y^d)$  are real  $\tau$ -independent constants

[Solution proposed in [Feruglio, Strumia, AT, 2305.08908](#)]

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- $k_{f_3} < 0$  and  $k_{\det} > 0$  (or vice versa)

►  $e^{-8\pi^2 f_3(S, \tau)}$  is singular at some point(s) in the moduli space

►  $\det[Y^u(\tau) Y^d(\tau)]$  is holomorphic everywhere

For  $A(S, \tau)$  to be  $\tau$ -independent, singularities of  $e^{-8\pi^2 f_3(S, \tau)}$  must be cancelled by zeroes of  $\det[Y^u(\tau) Y^d(\tau)]$  occurring at the same points in the moduli space

[Solution proposed in [Feruglio, Marrone, Strumia, AT, 2505.20395](#)]

# Singularities in the moduli space

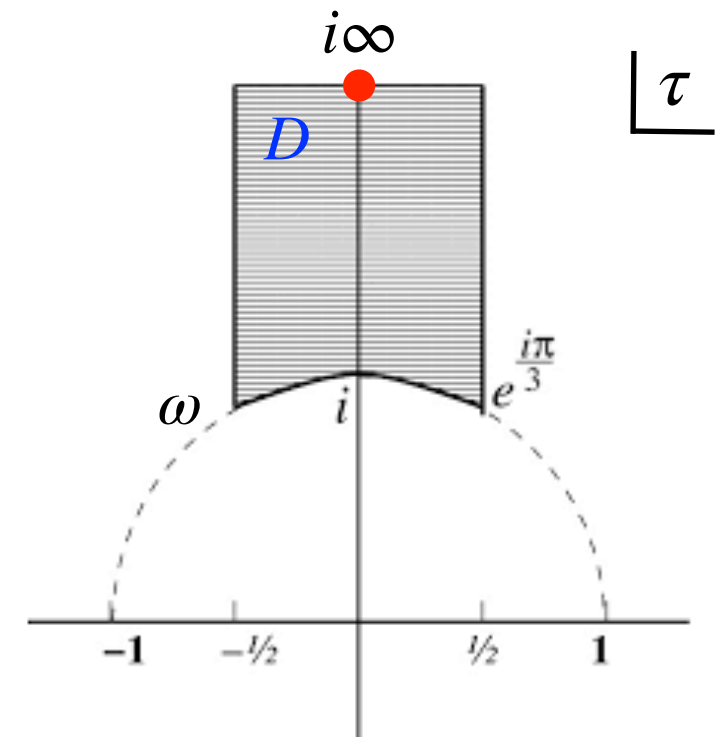
$e^{-8\pi^2 f_3(S, \tau)}$  is expected to be **singular**  
**at the cusp  $\tau = i\infty$**  (decompactification limit)  
 Gonzalo, Ibanez, Uranga, 1812.06520

**Distance Conjecture** Ooguri, Vafa, hep-th/0605264

$\tau = i\infty$  is at infinite distance from any point in  $D$

Explicit computations in string compactifications

$$f_3(S, \tau) = \kappa_3 S - \frac{k_{f_3}}{8\pi^2} \ln \eta^2(\tau) \quad \text{Kaplunovsky, Louis, hep-th/9502077}$$



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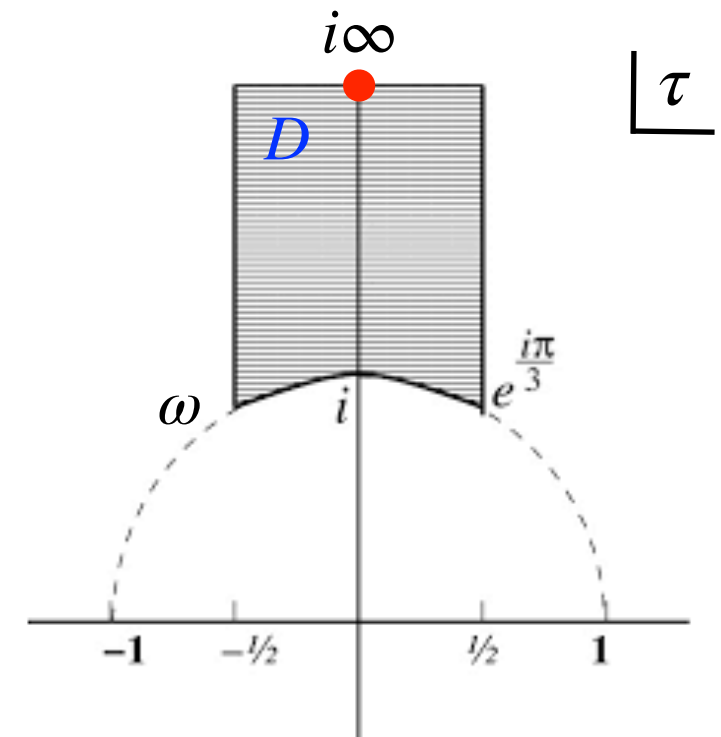
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Explicit computations in string compactifications

$$f_3(S, \tau) = \kappa_3 S - \frac{k_{f_3}}{8\pi^2} \ln \eta^2(\tau) \quad \text{Kaplunovsky, Louis, hep-th/9502077}$$



We assume that  $\det [Y^u(\tau) Y^d(\tau)]$  has a **unique zero at  $\tau = i\infty$**  (*quark masses never vanish except in this limit*), i.e., it is a cusp form of weight  $k_{\det} = 12m$ :

$$\det [Y^u(\tau) Y^d(\tau)] \propto \Delta(\tau)^m$$

$$\Delta(\tau) = \frac{E_4(\tau)^3 - E_6(\tau)^2}{1728} = \eta(\tau)^{24} = q \prod_{n=1}^{\infty} (1 - q^n)^{24} \quad \text{with} \quad q = e^{2\pi i \tau} \quad (k_{\Delta} = 12)$$

# Modular invariance and CP

---

Fields

$$\tau \xrightarrow{\text{CP}} -\tau^\dagger \quad \text{and} \quad \Phi \xrightarrow{\text{CP}} \Phi^\dagger$$

Modular forms

$$Y(\tau) \xrightarrow{\text{CP}} Y(-\tau^*) = Y(\tau)^*$$

Novichkov, Penedo, Petcov, **AT**, 1905.11970; Baur, Nilles, Trautner, Vaudrevange, 1901.03251

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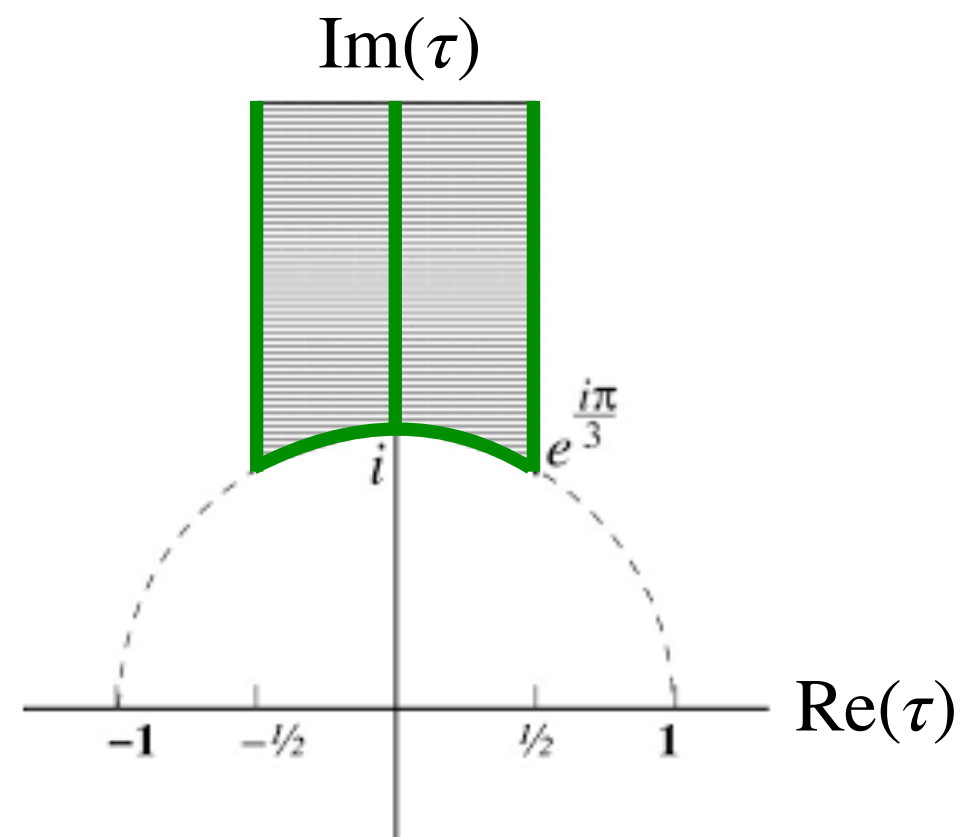
## CP-conserving values of $\tau$

$$\tau \xrightarrow{\text{CP}} -\tau^* = \gamma\tau \quad (\text{goes to itself up to } \gamma)$$

$$1. \quad \tau = iy \xrightarrow{\text{CP}} iy$$

$$2. \quad \tau = -\frac{1}{2} + iy \xrightarrow{\text{CP}} \frac{1}{2} + iy = T\tau$$

$$3. \quad \tau = e^{i\varphi} \xrightarrow{\text{CP}} -e^{-i\varphi} = S\tau$$



Novichkov, Penedo, Petcov, **AT**, 1905.11970; Baur, Nilles, Trautner, Vaudrevange, 1901.03251

# Example model

---

$$k_{H_u} = k_{H_d} = 0 \quad k_{Q_i} = k_{U_i^c} = k_{D_i^c} = (2, 4, 6) \quad \Rightarrow \quad k_{\text{det}} = 48 \quad (m = 4)$$

$$Y^{u,d}(\tau) = \begin{pmatrix} E_4 & E_6 & E_8 \\ E_6 & E_8 & E_{10} \\ E_8 & E_{10} & E_{12} \end{pmatrix} \quad \Rightarrow \quad \det Y^{u,d} \propto \Delta(\tau)^2$$

$$E_8 = E_4^2, \quad E_{10} = E_4 E_6, \quad E_{12} = a E_4^3 + (1 - a) E_6^2$$

# Example model

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$$E_8 = E_4^2, \quad E_{10} = E_4 E_6, \quad E_{12} = a E_4^3 + (1 - a) E_6^2$$

Slightly non-minimal Kähler potential

$$K = \sum_{i=1}^3 \left[ c_{Q_i}^{-2} y^{-k_{Q_i}} |Q_i|^2 + c_{U_i^c}^{-2} y^{-k_{U_i^c}} |U_i^c|^2 + c_{D_i^c}^{-2} y^{-k_{D_i^c}} |D_i^c|^2 \right] \quad \text{with} \quad y \equiv 2 \operatorname{Im} \tau$$

In the basis where kinetic terms of quarks are canonical

$$Y_{\text{can}}^q = \begin{pmatrix} c_{Q_1^c} c_{Q_1} y^2 E_4 & c_{Q_1^c} c_{Q_2} y^3 E_6 & c_{Q_1^c} c_{Q_3} y^4 E_8 \\ c_{Q_2^c} c_{Q_1} y^3 E_6 & c_{Q_2^c} c_{Q_2} y^4 E_8 & c_{Q_2^c} c_{Q_3} y^5 E_{10} \\ c_{Q_3^c} c_{Q_1} y^4 E_8 & c_{Q_3^c} c_{Q_2} y^5 E_{10} & c_{Q_3^c} c_{Q_3} y^6 E_{12} \end{pmatrix} \quad \text{with} \quad Q^c = \begin{cases} U^c & \text{for } q = u \\ D^c & \text{for } q = d \end{cases}$$

# Example model: quarks

| Observable       | Central value $\pm 1\sigma$ |
|------------------|-----------------------------|
| $m_u/m_c$        | $(1.93 \pm 0.60) 10^{-3}$   |
| $m_c/m_t$        | $(2.82 \pm 0.12) 10^{-3}$   |
| $m_d/m_s$        | $(5.05 \pm 0.62) 10^{-2}$   |
| $m_s/m_b$        | $(1.82 \pm 0.10) 10^{-2}$   |
| $m_t/\text{GeV}$ | $87.5 \pm 2.1$              |
| $m_b/\text{GeV}$ | $0.97 \pm 0.01$             |

| Observable                | Central value $\pm 1\sigma$ |
|---------------------------|-----------------------------|
| $\sin^2 \theta_{12}^q$    | $(5.08 \pm 0.03) 10^{-2}$   |
| $\sin^2 \theta_{13}^q$    | $(1.22 \pm 0.09) 10^{-5}$   |
| $\sin^2 \theta_{23}^q$    | $(1.61 \pm 0.05) 10^{-3}$   |
| $\delta_{\text{CKM}}/\pi$ | $0.385 \pm 0.017$           |

Quark masses and mixing parameters renormalised at  $2 \times 10^{16}$  GeV assuming  $M_{\text{SUSY}} = 10$  TeV and  $\tan \beta = 10$

Antusch, Maurer, 1306.6879

Best-fit point ( $\chi^2 \approx 0$ ):

$$\tau = -0.286 + 1.096i$$

$$q_{13} = 0.037, \quad q_{23} = 0.075, \quad u_{13} = 0.035, \quad u_{23} = 19.98, \quad d_{13} = 3.44, \quad d_{23} = 0.203$$

$$c_{U_3^c} c_{Q_3} = 4.15 \times 10^{-4} \quad \text{and} \quad c_{D_3^c} c_{Q_3} = 4.76 \times 10^{-4} \quad (\text{overall scales})$$

$$(q_{i3} \equiv c_{Q_i}/c_{Q_3}, \quad u_{i3} \equiv c_{U_i^c}/c_{U_3^c}, \quad d_{i3} \equiv c_{D_i^c}/c_{D_3^c}, \quad i = 1, 2)$$

8 real parameters +  $\tau$  vs. 10 observables

# Example model: leptons

| Observable                 | Central value $\pm 1\sigma$                                |
|----------------------------|------------------------------------------------------------|
| $m_e/m_\mu$                | $(4.74 \pm 0.04) 10^{-3}$                                  |
| $m_\mu/m_\tau$             | $(5.88 \pm 0.05) 10^{-2}$                                  |
| $\delta m^2/ \Delta m^2 $  | $(2.95 \pm 0.06) 10^{-2}$<br>$(2.99 \pm 0.07) 10^{-2}$     |
| $m_\tau/\text{GeV}$        | $1.293 \pm 0.007$                                          |
| $\delta m^2/\text{eV}^2$   | $(7.37 \pm 0.17) 10^{-5}$                                  |
| $ \Delta m^2 /\text{eV}^2$ | $(2.495 \pm 0.020) 10^{-3}$<br>$(2.465 \pm 0.020) 10^{-3}$ |

| Observable                 | Central value $\pm 1\sigma$                            |
|----------------------------|--------------------------------------------------------|
| $\sin^2 \theta_{12}^\ell$  | $(3.03 \pm 0.14) 10^{-1}$                              |
| $\sin^2 \theta_{13}^\ell$  | $(2.23 \pm 0.05) 10^{-2}$                              |
| $\sin^2 \theta_{23}^\ell$  | $(4.73 \pm 0.24) 10^{-1}$<br>$(5.45 \pm 0.23) 10^{-1}$ |
| $\delta_{\text{PMNS}}/\pi$ | $1.20 \pm 0.22$<br>$1.48 \pm 0.12$                     |

Lepton masses and mixing parameters

Antusch, Maurer, 1306.6879

Capozzi et al., 2503.07752

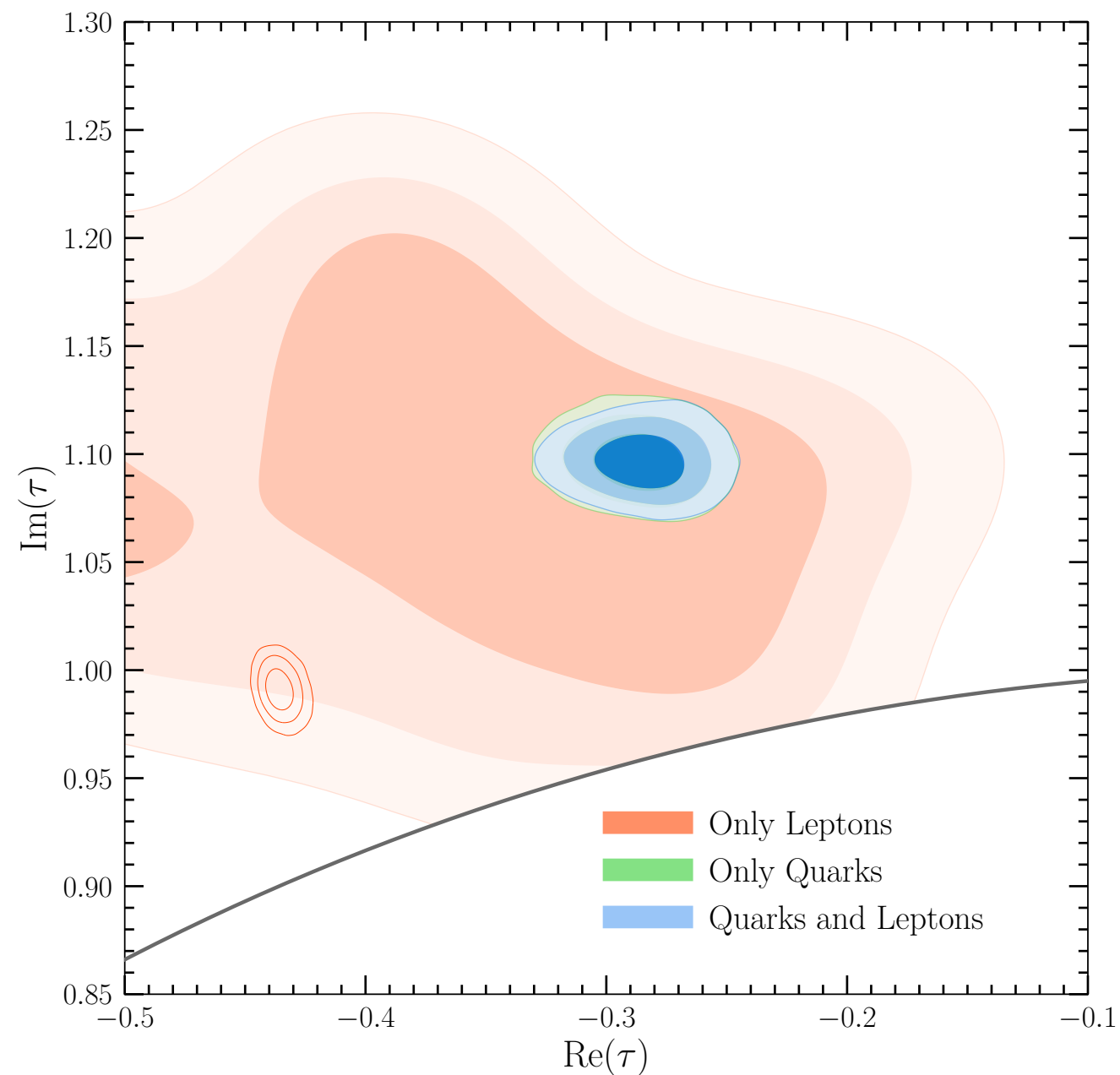
Best-fit point ( $\chi^2 \approx 0$ ) for the same value of  $\tau$ :

$$l_{13} = 2.51, l_{23} = 1.94, e_{13} = 2.21, e_{23} = 0.0079, c_{33}^e = 5.61, c_{33}^\nu = 0.076$$

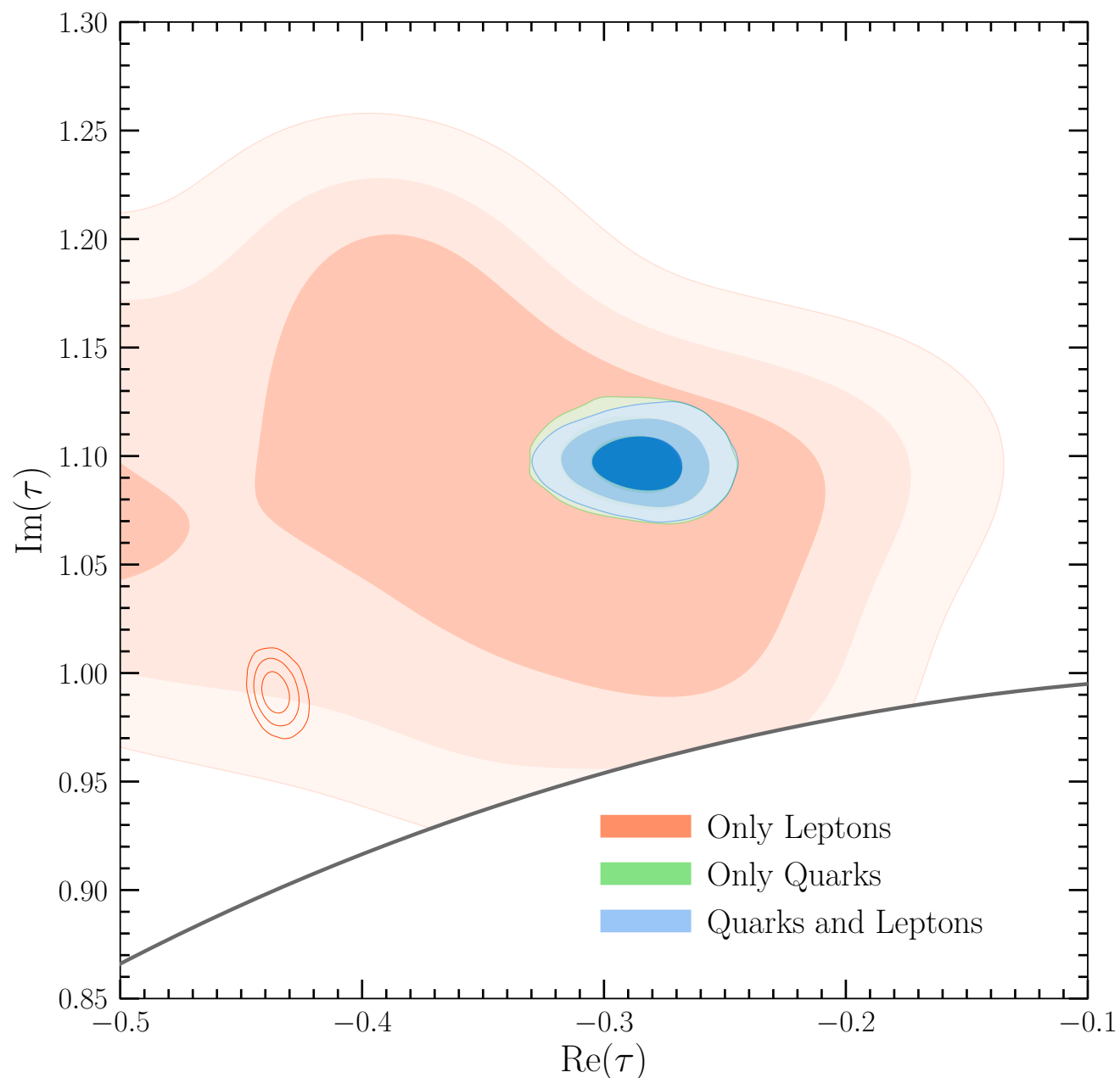
$$c_{E_3^c} c_{L_3} = 7.66 \cdot 10^{-5} \quad \text{and} \quad \frac{c_{L_3}^2}{2\Lambda_L} = \frac{6.60 \cdot 10^{-2}}{10^{16} \text{ GeV}} \quad (\text{overall scales})$$

8 parameters vs. 8 observables + 4 predictions

# Example model: combined fit



# Example model: combined fit



Predictions in the neutrino sector

## Masses

- Normal ordering with  $m_1 \approx 10$  meV

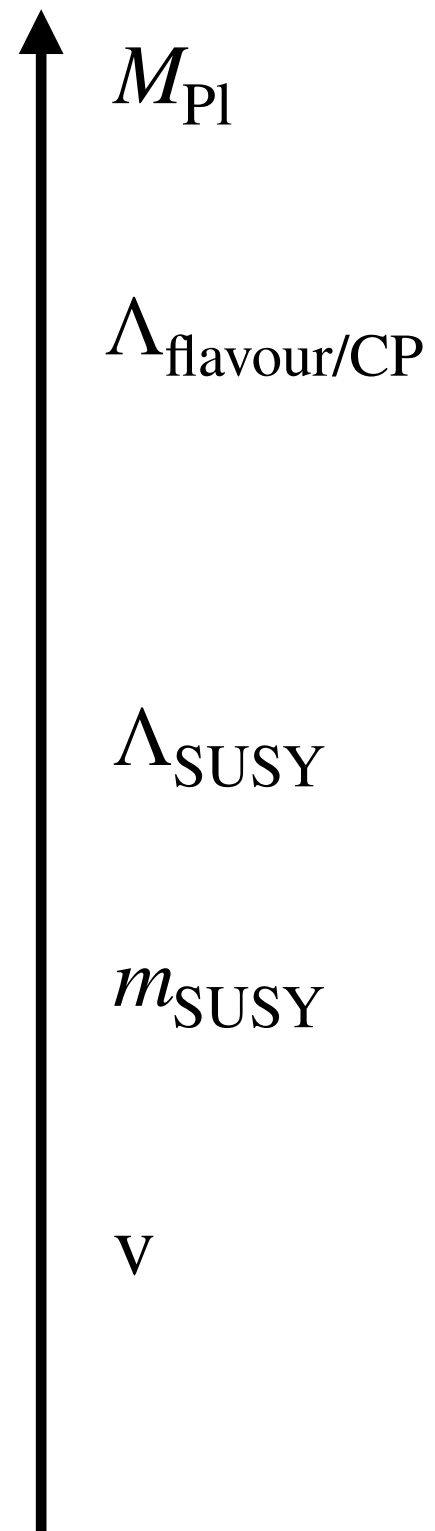
$$m_{\beta\beta} \approx 6 \text{ meV} \quad \sum_{i=1}^3 m_i \approx 74 \text{ meV}$$

## CPV phases

- $\delta_{\text{PMNS}} \approx 0.94 \pi$
- $\alpha_{21} \approx 1.29 \pi \quad \alpha_{31} \approx 0.14 \pi$

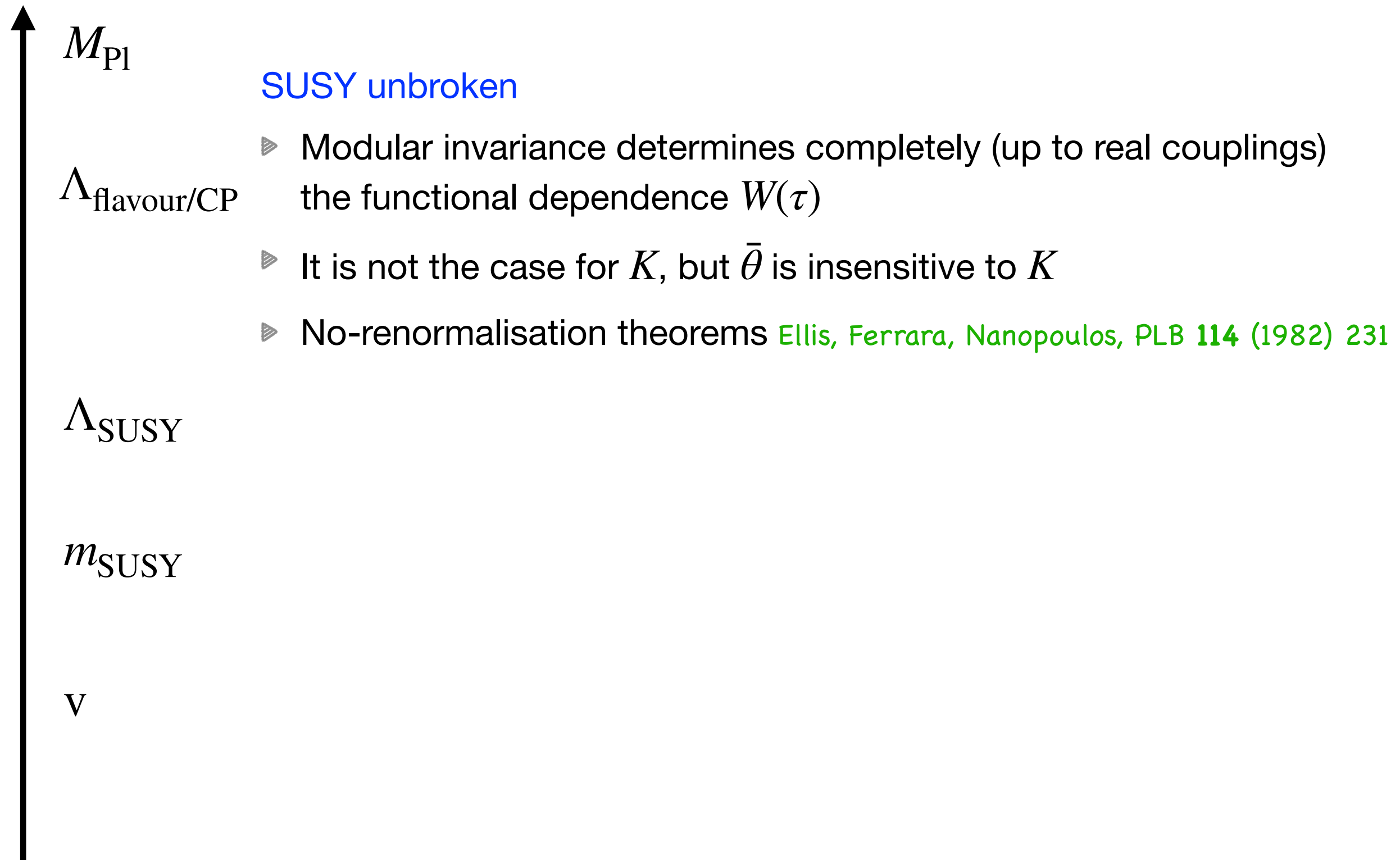
# Corrections to $\bar{\theta} = 0$

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# Corrections to $\bar{\theta} = 0$

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# Corrections to $\bar{\theta} = 0$

|                               |                                                                                                                                                                                                                                                                                                                                                                                  |
|-------------------------------|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| $M_{\text{Pl}}$               | SUSY unbroken                                                                                                                                                                                                                                                                                                                                                                    |
| $\Lambda_{\text{flavour/CP}}$ | <ul style="list-style-type: none"> <li>► Modular invariance determines completely (up to real couplings) the functional dependence <math>W(\tau)</math></li> <li>► It is not the case for <math>K</math>, but <math>\bar{\theta}</math> is insensitive to <math>K</math></li> <li>► No-renormalisation theorems Ellis, Ferrara, Nanopoulos, PLB <b>114</b> (1982) 231</li> </ul> |
| $\Lambda_{\text{SUSY}}$       | SUSY breaking corrections                                                                                                                                                                                                                                                                                                                                                        |
| $m_{\text{SUSY}}$             | <ul style="list-style-type: none"> <li>► In general, can be large</li> <li>► Small if <math>\Lambda_{\text{flavour/CP}} \gg \Lambda_{\text{SUSY}}</math> (as e.g. in gauge mediation) and soft breaking terms respect the flavour structure of the SM</li> </ul>                                                                                                                 |
| $v$                           | $\bar{\theta} \lesssim \frac{m_t^4 m_b^4 m_c^2 m_s^2}{v^{12}} J_{\text{CP}} \tan^6 \beta \sim 10^{-28} \tan^6 \beta$                                                                                                                                                                                                                                                             |

# Corrections to $\bar{\theta} = 0$

|                               |                                                                                                                                                                                                                                                                                                                                                                                  |
|-------------------------------|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| $M_{\text{Pl}}$               | SUSY unbroken                                                                                                                                                                                                                                                                                                                                                                    |
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| $\Lambda_{\text{SUSY}}$       | SUSY breaking corrections                                                                                                                                                                                                                                                                                                                                                        |
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| $v$                           | $\bar{\theta} \lesssim \frac{m_t^4 m_b^4 m_c^2 m_s^2}{v^{12}} J_{\text{CP}} \tan^6 \beta \sim 10^{-28} \tan^6 \beta$ <p>SM corrections are negligible</p> <ul style="list-style-type: none"> <li>► <math>\bar{\theta} \sim 10^{-18}</math> at four loops Khriplovich, PLB <b>173</b> (1986) 193<br/>Ellis, Gaillard, NPB <b>150</b> (1979) 141</li> </ul>                        |

# Conclusions

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- ▶ **Modular invariance** is inherent to **toroidal compactifications** in string theory and can be consistently implemented in a **supersymmetric QFT**
- ▶ Quarks with **positive weights** and **non-trivial gauge kinetic functions**
- ▶ The VEV of the modulus  $\tau$  is the only source of **spontaneous CP violation**, while the dilaton  $S$  has a CP-conserving VEV

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- ▶ The VEV of the modulus  $\tau$  is the only source of **spontaneous CP violation**, while the dilaton  $S$  has a CP-conserving VEV

$$\bar{\theta} = \arg A(S, \tau) \quad \text{with} \quad A(S, \tau) \equiv e^{-8\pi^2 f_3(S, \tau)} \det [Y^u(\tau) Y^d(\tau)]$$

- ▶  $f_3$  is singular at  $\tau = i\infty$ , while  $\det (Y^u Y^d)$  has the only zero at this point, such that  $A(S, \tau)$  is a real constant and hence  **$\bar{\theta} = 0$**
- ▶ **Quark and lepton masses and mixings** are accommodated for the same  $\langle \tau \rangle$
- ▶ Corrections to  $\bar{\theta} = 0$  are small under certain assumptions on SUSY breaking

# Back-up slides

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# What can solve the strong CP problem?

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Kaplan, Melia, Rajendran, 2505.08358

- ▶  $\theta_{\text{QCD}}$  cannot be set to zero by CP/P, since it is not a parameter (like other couplings), but a label of a vacuum state
- ▶ Viable solutions should be dynamical, like the axion solution

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Kaplan, Melia, Rajendran, 2505.08358

- ▶  $\theta_{\text{QCD}}$  cannot be set to zero by CP/P, since it is not a parameter (like other couplings), but a label of a vacuum state
- ▶ Viable solutions should be dynamical, like the axion solution

In our framework, CP is conserved, but  $\bar{\theta}$  is a dynamical field-dependent quantity, i.e., we do not set  $\theta_{\text{QCD}} = 0$  by CP

$$\bar{\theta} = \arg A(S, \tau) \quad \text{with} \quad A(S, \tau) \equiv e^{-8\pi^2 f_3(S, \tau)} \det [Y^u(\tau) Y^d(\tau)]$$

- CP invariance  $\Rightarrow f_3(S^*, -\tau^*) = f_3(S, \tau)^*$
- $\theta_{\text{QCD}}$  is inside  $S$   $\Rightarrow f_3(S, \tau) = S + \dots$

# Phenomenology and cosmology

---

- ▶ Couplings to matter are suppressed by  $1/h$  ( $1/M_{\text{Pl}}$  in SUGRA)
- ▶ No couplings to gauge bosons in the exact SUSY limit
- ▶  $m_\tau \gtrsim 10$  TeV not to spoil BBN
- ▶ Fermionic component of  $\tau$  could be LSP and maybe DM
- ▶ Scalar potential  $V(\tau) = V(-\tau^*) \Rightarrow$  CP-conjugated minima  
(domain walls are inflated away if CP breaking occurs before inflation)

# Matter fields and canonical normalisation

Gauge quantum numbers

|           | $Q$           | $U^c$                       | $D^c$                       | $L$            | $E^c$    | $H_u$         | $H_d$          |
|-----------|---------------|-----------------------------|-----------------------------|----------------|----------|---------------|----------------|
| $SU(3)_c$ | <b>3</b>      | <b><math>\bar{3}</math></b> | <b><math>\bar{3}</math></b> | <b>1</b>       | <b>1</b> | <b>1</b>      | <b>1</b>       |
| $SU(2)_L$ | <b>2</b>      | <b>1</b>                    | <b>1</b>                    | <b>2</b>       | <b>1</b> | <b>2</b>      | <b>2</b>       |
| $U(1)_Y$  | $\frac{1}{6}$ | $-\frac{2}{3}$              | $\frac{1}{3}$               | $-\frac{1}{2}$ | 1        | $\frac{1}{2}$ | $-\frac{1}{2}$ |

Canonical normalisation

$$K \supset \frac{\Phi^\dagger \Phi}{(-i\tau + i\tau^\dagger)^{k_\Phi}} = \Phi_{\text{can}}^\dagger \Phi_{\text{can}} \quad \Phi_{\text{can}} = \{\phi_{\text{can}}, \psi_{\text{can}}\}$$

$$\psi_{\text{can}} \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{-\frac{k_\Phi}{2}} \psi_{\text{can}} = e^{-ik_\Phi \alpha(\tau)} \psi_{\text{can}} \quad \alpha(\tau) = \arg(c\tau + d)$$

Modular symmetry acts on canonically normalised fields  
as a  **$\tau$ -dependent phase rotation** (with  $\tau = \tau(x)$ )

# Modular-SM anomalies

Conditions for modular-gauge anomaly cancellation

$$\text{SU}(3)_c : \quad A \equiv \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right) = 0$$

$$\text{SU}(2)_L : \quad \sum_{i=1}^3 \left( 3k_{Q_i} + k_{L_i} \right) + k_{H_u} + k_{H_d} = 0$$

$$\text{U}(1)_Y : \quad \sum_{i=1}^3 \left( k_{Q_i} + 8k_{U_i^c} + 2k_{D_i^c} + 3k_{L_i} + 6k_{E_i^c} \right) + 3 \left( k_{H_u} + k_{H_d} \right) = 0$$

Simplest solution

$$k_Q = k_{U^c} = k_{D^c} = k_L = k_{E^c} = (-k, 0, k) \quad \text{and} \quad k_{H_u} + k_{H_d} = 0$$

Cancellation of modular-QCD anomaly along with  $k_{H_u} + k_{H_d} = 0$  implies

$$k_{\text{det}} = \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right) + 3 \left( k_{H_u} + k_{H_d} \right) = 0$$

# More on modular-gauge anomalies

$$\Phi \rightarrow \Phi' = (c\tau + d)^{-k_\Phi} \Phi$$

$$\text{Jacobian } J: \mathcal{D}\Phi' = J \mathcal{D}\Phi$$

Arkani-Hamed, Murayama, hep-th/9707133

$$\log J = -\frac{i}{64\pi^2} \int d^4x d^2\theta \left[ \sum_{\Phi} T(\Phi) k_\Phi \right] W^a W^a \log(c\tau + d)$$

$T(\Phi)$  is the Dynkin index of the rep of  $\Phi$ :  $\text{tr}(t_a t_b) = T(\Phi) \delta_{ab}$

$$\sum_{\Phi} T(\Phi) k_\Phi = 0$$

$$\text{SU}(3)_c : \sum_i \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} \right) = 0$$

$$\text{SU}(2)_L : \sum_i \left( 3k_{Q_i} + k_{L_i} \right) + k_{H_u} + k_{H_d} = 0$$

$$\text{U}(1)_Y : \sum_i \left( k_{Q_i} + 8k_{U_i^c} + 2k_{D_i^c} + 3k_{L_i} + 6k_{E_i^c} \right) + 3 \left( k_{H_u} + k_{H_d} \right) = 0$$

# Simplest example: quarks

Simplest non-trivial example giving  $k_{\text{det}} = 0$  and  $A = 0$

$$k_Q = k_{U^c} = k_{D^c} = (-6, 0, 6) \quad \text{and} \quad k_{H_u} = k_{H_d} = 0$$

Yukawa matrices

$$Y^q = \begin{pmatrix} 0 & 0 & c_{13}^q \\ 0 & c_{22}^q & c_{23}^q E_6 \\ c_{31}^q & c_{32}^q E_6 & c_{33}^q E_4^3 + c_{33}'^q E_6^2 \end{pmatrix} \Rightarrow Y^q|_{\text{can}} = \begin{pmatrix} 0 & 0 & c_{13}^q \\ 0 & c_{22}^q & c_{23}^q (2\text{Im}\tau)^3 E_6 \\ c_{31}^q & c_{32}^q (2\text{Im}\tau)^3 E_6 & (2\text{Im}\tau)^6 [c_{33}^q E_4^3 + c_{33}'^q E_6^2] \end{pmatrix}$$

$$\det Y^q|_{\text{can}} = -c_{13}^q c_{22}^q c_{31}^q \in \mathbb{R}$$

Fixing  $\tau = 1/8 + i$  and  $\tan \beta = 10$

$$c_{ij}^u \approx 10^{-3} \begin{pmatrix} 0 & 0 & 1.56 \\ 0 & -1.86 & 0.87 \\ 1.29 & 4.14 & 3.51, 1.40 \end{pmatrix} \quad c_{ij}^d \approx 10^{-3} \begin{pmatrix} 0 & 0 & 1.55 \\ 0 & -2.59 & 4.59 \\ 0.378 & 0.710 & 0.734, 1.76 \end{pmatrix}$$

reproduce the quark masses, mixing angles and  $\delta_{\text{CKM}}$  at the GUT scale of  $2 \times 10^{16}$  GeV

# Simplest example: leptons

---

$$k_L = k_{E^c} = (-6, 0, 6)$$

Weinberg operator  $\mathcal{C}_{ij}^\nu (L_i H_u)(L_j H_u)$  for neutrino masses

Charged lepton Yukawa matrix and coefficient of the Weinberg operator

$$Y^e = \begin{pmatrix} 0 & 0 & c_{13}^e \\ 0 & c_{22}^e & c_{23}^e E_6 \\ c_{31}^e & c_{32}^e E_6 & c_{33}^e E_4^3 + c'_{33} E_6^2 \end{pmatrix} \quad \mathcal{C}^\nu = \begin{pmatrix} 0 & 0 & c_{13}^\nu \\ 0 & c_{22}^\nu & c_{23}^\nu E_6 \\ c_{13}^\nu & c_{23}^\nu E_6 & c_{33}^\nu E_4^3 + c'_{33} E_6^2 \end{pmatrix}$$

Fixing  $\tau = 1/8 + i$  and  $\tan \beta = 10$

$$c_{ij}^e = 10^{-3} \begin{pmatrix} 0 & 0 & 1.29 \\ 0 & 5.95 & 0.35 \\ -2.56 & 1.47 & 1.01, 1.32 \end{pmatrix} \quad c_{ij}^\nu = \frac{1}{10^{16} \text{ GeV}} \begin{pmatrix} 0 & 0 & 3.4 \\ 0 & 7.1 & 1.2 \\ 3.4 & 1.2 & 0.19, 0.95 \end{pmatrix}$$

reproduce the lepton masses and mixings, including  $\delta_{\text{PMNS}}$

# Models with larger modular weights

| Yukawa matrices<br>$Y_{u,d}$                                                                       | Modular weights<br>$(u_L, d_L)_{1,2,3}$ $u_{R1,2,3}$ $d_{R1,2,3}$ |                                              |                                              | Alternative bigger weights<br>$(u_L, d_L)_{1,2,3}$ $u_{R1,2,3}$ $d_{R1,2,3}$ |                                               |                                               |
|----------------------------------------------------------------------------------------------------|-------------------------------------------------------------------|----------------------------------------------|----------------------------------------------|------------------------------------------------------------------------------|-----------------------------------------------|-----------------------------------------------|
| $\begin{pmatrix} 0 & 0 & 1 \\ 0 & 1 & E_6 \\ 1 & E_6 & E_4^3 + E_6^2 \end{pmatrix}$                | $\begin{pmatrix} -6 \\ 0 \\ 6 \end{pmatrix}$                      | $\begin{pmatrix} -6 \\ 0 \\ 6 \end{pmatrix}$ | $\begin{pmatrix} -6 \\ 0 \\ 6 \end{pmatrix}$ | $\begin{pmatrix} -4 \\ 2 \\ 8 \end{pmatrix}$                                 | $\begin{pmatrix} -8 \\ -2 \\ 4 \end{pmatrix}$ | $\begin{pmatrix} -8 \\ -2 \\ 4 \end{pmatrix}$ |
| $\begin{pmatrix} 0 & 0 & 1 \\ 0 & 1 & E_4^2 \\ 1 & E_4 & E_4^3 + E_6^2 \end{pmatrix}$              | $\begin{pmatrix} -6 \\ -2 \\ 6 \end{pmatrix}$                     | $\begin{pmatrix} -6 \\ 2 \\ 6 \end{pmatrix}$ | $\begin{pmatrix} -6 \\ 2 \\ 6 \end{pmatrix}$ | $\begin{pmatrix} -4 \\ -4 \\ 8 \end{pmatrix}$                                | $\begin{pmatrix} -8 \\ 4 \\ 4 \end{pmatrix}$  | $\begin{pmatrix} -8 \\ 4 \\ 4 \end{pmatrix}$  |
| $\begin{pmatrix} 0 & 0 & 1 \\ 0 & 1 & E_4^2 \\ 1 & E_4^2 & E_4(E_4^3 + E_6^2) \end{pmatrix}$       |                                                                   |                                              |                                              | $\begin{pmatrix} -8 \\ 0 \\ 8 \end{pmatrix}$                                 | $\begin{pmatrix} -8 \\ 0 \\ 8 \end{pmatrix}$  | $\begin{pmatrix} -8 \\ 0 \\ 8 \end{pmatrix}$  |
| $\begin{pmatrix} 0 & 0 & 1 \\ 0 & 1 & E_4 E_6 \\ 1 & E_6 & E_4(E_4^3 + E_6^2) \end{pmatrix}$       |                                                                   |                                              |                                              | $\begin{pmatrix} -8 \\ -2 \\ 8 \end{pmatrix}$                                | $\begin{pmatrix} -8 \\ 2 \\ 8 \end{pmatrix}$  | $\begin{pmatrix} -8 \\ 2 \\ 8 \end{pmatrix}$  |
| $\begin{pmatrix} 0 & 0 & 1 \\ 0 & 1 & E_4^3 + E_6^2 \\ 1 & E_4 & E_4(E_4^3 + E_6^2) \end{pmatrix}$ |                                                                   |                                              |                                              | $\begin{pmatrix} -8 \\ -4 \\ 8 \end{pmatrix}$                                | $\begin{pmatrix} -8 \\ 4 \\ 8 \end{pmatrix}$  | $\begin{pmatrix} -8 \\ 4 \\ 8 \end{pmatrix}$  |

# Heavy quarks and singularities

- Heavy quarks are not needed for the mechanism to work, but assume they exist

$$k_q = (-6, -2, 0, +2, +6) \quad \text{and} \quad k_{H_u} = k_{H_d} = 0$$

Light chiral quarks    Heavy vector-like quarks

- In the full theory  $f_{UV} \in \mathbb{R}$  and  $\det M_{\text{all}} \in \mathbb{R} \Rightarrow \bar{\theta}_{UV} = 0$

$$M_{\text{all}} = \begin{pmatrix} M_{LL} & M_{LH} \\ M_{HL} & M_{HH} \end{pmatrix}$$

$$M_{\text{light}} \approx M_{LL} - M_{LH} M_{HH}^{-1} M_{HL} \quad M_{\text{heavy}} \approx M_{HH}$$

Singularities where  $\det M_{\text{heavy}}(\tau) = 0$   
(breakdown of EFT)

$$\det M_{\text{all}} = \det M_{\text{light}} \det M_{\text{heavy}}$$

$$\det M_{\text{light}} \rightarrow (c\tau + d)^{k_{\text{light}}} \det M_{\text{light}} \quad \det M_{\text{heavy}} \rightarrow (c\tau + d)^{k_{\text{heavy}}} \det M_{\text{heavy}}$$

$$k_{\text{all}} = k_{\text{light}} + k_{\text{heavy}} = 0$$

# EFT of light quarks

---

In the EFT of light quarks

$$\bar{\theta}_{\text{IR}} = \theta_{\text{QCD}} + \arg \det M_{\text{light}} = -8\pi^2 \text{Im} f_{\text{IR}} + \arg \det M_{\text{light}}$$

The EFT has anomalous field content with  $k_q = (-6, -2, 0)$

Anomaly is cancelled by a new contribution to the gauge kinetic function arising from the integration over the heavy quarks

$$f_{\text{IR}} = f_{\text{UV}} - \frac{1}{8\pi^2} \log \det M_{\text{heavy}}$$

Thus

$$\bar{\theta}_{\text{IR}} = \arg \det M_{\text{heavy}} + \arg \det M_{\text{light}} = \arg \det M_{\text{all}} = 0$$

# Model with VLQs based on $\Gamma_2 \cong S_3$

Feruglio, Parriciatu, Strumia, **AT**, 2406.01689

|                            | SM quarks                        |                                  |                                  | Extra vector-like quarks         |                                  |                                  |                                  |
|----------------------------|----------------------------------|----------------------------------|----------------------------------|----------------------------------|----------------------------------|----------------------------------|----------------------------------|
|                            | $Q$                              | $D^c$                            | $U^c$                            | $D'^c$                           | $D'$                             | $U'^c$                           | $U'$                             |
| $SU(2)_L \otimes U(1)_Y$   | $2_{1/6}$                        | $1_{1/3}$                        | $1_{-2/3}$                       | $1_{1/3}$                        | $1_{-1/3}$                       | $1_{-2/3}$                       | $1_{2/3}$                        |
| Flavor symmetry $\Gamma_2$ | $\mathbf{2} \oplus \mathbf{1}_0$ | $\mathbf{2} \oplus \mathbf{1}_1$ | $\mathbf{2} \oplus \mathbf{1}_0$ | $\mathbf{2} \oplus \mathbf{1}_0$ | $\mathbf{2} \oplus \mathbf{1}_1$ | $\mathbf{2} \oplus \mathbf{1}_0$ | $\mathbf{2} \oplus \mathbf{1}_0$ |
| Modular weights $k_\Phi$   | $-2$                             | $-2$                             | $-2$                             | $+2$                             | $+2$                             | $+2$                             | $+2$                             |

$$W_{UV} \supset Q^T m^d D^c + Q^T n^d D'^c + D'^T N^d D^c + D'^T M^d D'^c$$

$$\mathcal{M}_d = \begin{pmatrix} m^d & n^d \\ N^d & M^d \end{pmatrix} \quad m^d = v_d Y^d \quad n^d = v_d Y'^d$$

$$m^d = 0_{3 \times 3} \quad n^d = n_d \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & \alpha_d \end{pmatrix} \quad N^d = N_d \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & \beta_d \end{pmatrix} \quad M^d = M_d \begin{pmatrix} -Z_1^{(4)} + \gamma_{d1} Z_3^{(4)} & Z_2^{(4)} & \gamma_{d2} Z_1^{(4)} \\ Z_2^{(4)} & Z_1^{(4)} + \gamma_{d1} Z_3^{(4)} & \gamma_{d2} Z_2^{(4)} \\ \gamma_{d3} Z_2^{(4)} & -\gamma_{d3} Z_1^{(4)} & 0 \end{pmatrix}$$

$$(Z_1^{(4)}(\tau), Z_2^{(4)}(\tau))^T \sim \mathbf{2} \quad Z_3^{(4)}(\tau) \sim \mathbf{1}_0 \quad \text{level } N=2 \text{ modular forms of weight } k=4$$

$$\det \mathcal{M}_d = \det \left[ m^d - n^d (M^d)^{-1} N_d \right] \det M^d = -\det [n_d N_d] = -n_d^3 N_d^3 \alpha_d \beta_d \in \mathbb{R}$$

# Modular group

---

## Homogeneous modular group

$$\Gamma = \langle S, T \mid S^4 = (ST)^3 = I \rangle \cong \mathrm{SL}(2, \mathbb{Z})$$

$$S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \quad T = \begin{pmatrix} 1 & 1 \\ 0 & 1 \end{pmatrix}$$

$$\tau \xrightarrow{S} -\frac{1}{\tau}$$

duality

$$\tau \xrightarrow{T} \tau + 1$$

discrete shift symmetry

## Inhomogeneous modular group

$$\bar{\Gamma} = \langle S, T \mid S^2 = (ST)^3 = I \rangle \cong \mathrm{PSL}(2, \mathbb{Z}) = \mathrm{SL}(2, \mathbb{Z}) / \{I, -I\}$$

In other words,  $\mathrm{SL}(2, \mathbb{Z})$  matrices  $\gamma$  and  $-\gamma$  are identified

$$\tau \xrightarrow{\gamma} \gamma\tau = \frac{a\tau + b}{c\tau + d} \quad \tau \xrightarrow{-\gamma} (-\gamma)\tau = \frac{-a\tau - b}{-c\tau - d} = \gamma\tau$$

# Modular forms

---

Holomorphic functions on  $\mathcal{H} = \{\tau \in \mathbb{C} : \operatorname{Im}(\tau) > 0\}$  transforming under  $\Gamma$  as

$$f(\gamma\tau) = (c\tau + d)^k f(\tau), \quad \gamma \in \Gamma$$

$k$  is **weight**, a non-negative even integer

$$\gamma = -I \Rightarrow f(\tau) = (-1)^k f(\tau) \Rightarrow k \text{ is even}$$

Modular forms are periodic and admit  **$q$ -expansions**

$$\gamma = T = \begin{pmatrix} 1 & 1 \\ 0 & 1 \end{pmatrix} \Rightarrow f(\tau + 1) = f(\tau) \Rightarrow f(\tau) = \sum_{n=0}^{\infty} a_n q^n, \quad q = e^{2\pi i \tau}$$

Modular forms of weight  $k$  form a **linear space  $\mathcal{M}_k$  of finite dimension**

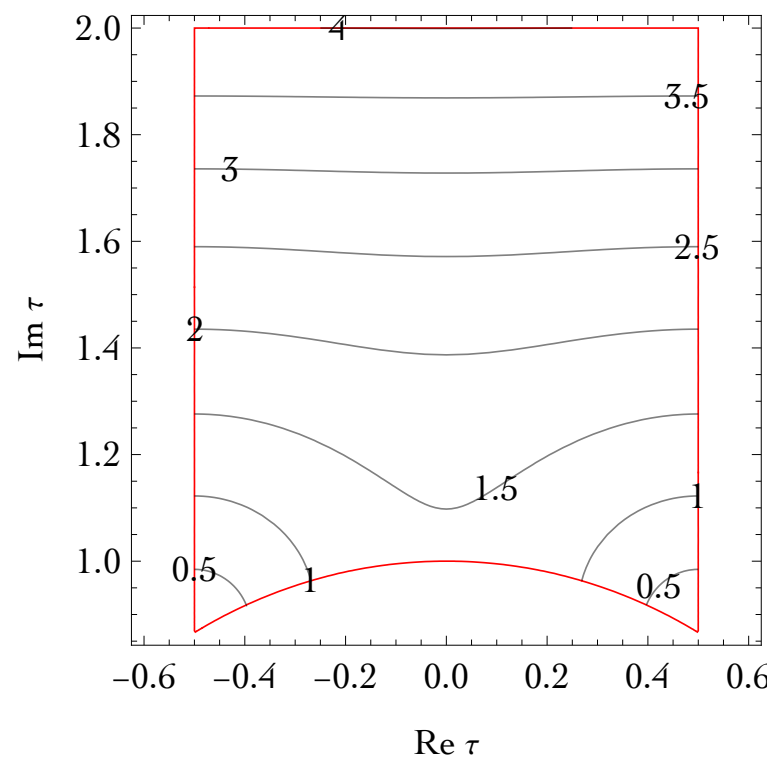
$$\dim \mathcal{M}_k = \begin{cases} 0 & \text{if } k \text{ is negative or odd} \\ \lfloor k/12 \rfloor & \text{if } k \equiv 2 \pmod{12} \\ \lfloor k/12 \rfloor + 1 & \text{if } k \not\equiv 2 \pmod{12} \end{cases}$$

# E4 and E6

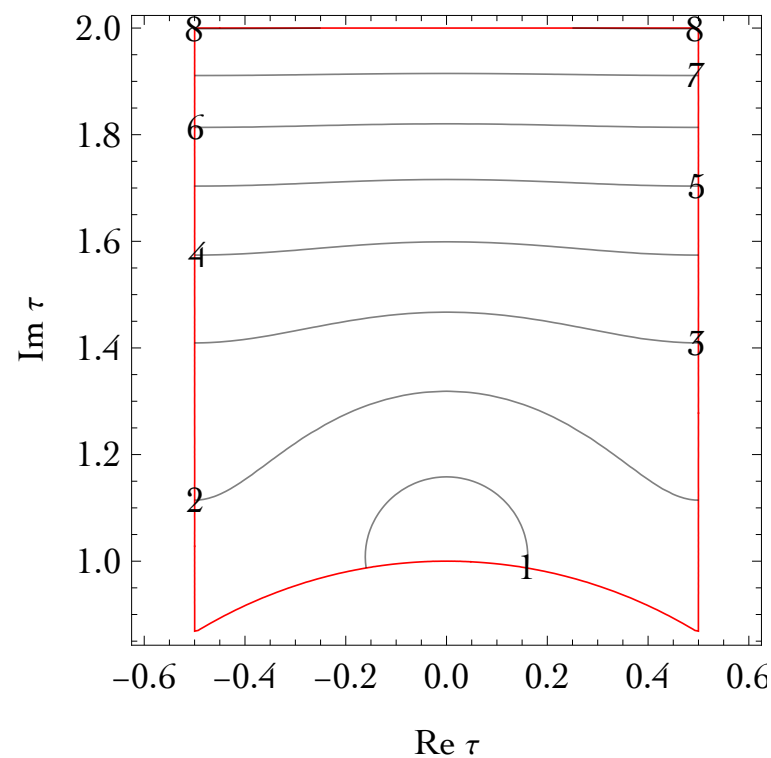
$$E_4(\tau) = 1 + 240 \sum_{n=1}^{\infty} \frac{n^3 q^n}{1 - q^n} = 1 + 240q + 2160q^2 + 6720q^3 + 17520q^4 + \mathcal{O}(q^5)$$

$$E_6(\tau) = 1 - 540 \sum_{n=1}^{\infty} \frac{n^5 q^n}{1 - q^n} = 1 - 504q - 16632q^2 - 122976q^3 - 532728q^4 + \mathcal{O}(q^5)$$

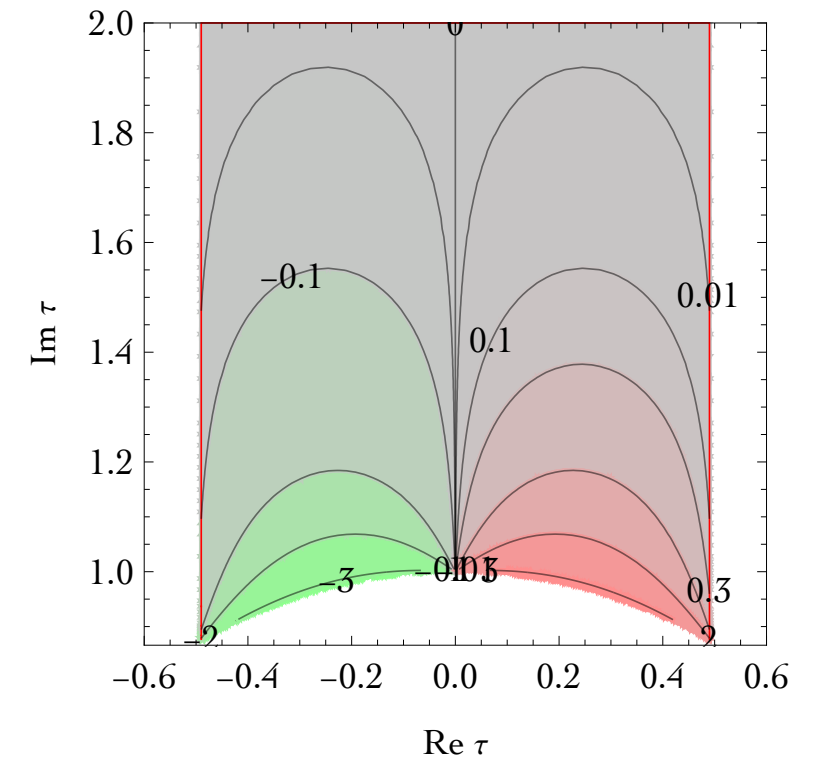
$|(\operatorname{Im} \tau)^2 E_4(\tau)|$



$|(\operatorname{Im} \tau)^3 E_6(\tau)|$



$\arg E_4^3/E_6^2$



# Modular invariance and SUGRA

---

$\mathcal{N} = 1$  SUGRA action depends on

$$G = \frac{K}{M_{\text{Pl}}^2} + \log \left| \frac{W}{M_{\text{Pl}}^3} \right|^2$$

For  $G$  to be invariant, both  $K$  and  $W$  have to transform

$$K \rightarrow K + M_{\text{Pl}}^2 (F + F^\dagger) \quad \text{and} \quad W \rightarrow e^{-F} W$$

In the case of modular transformations

$$F = \frac{h^2}{M_{\text{Pl}}^2} \log(c\tau + d)$$

$$W \rightarrow (c\tau + d)^{-k_W} W \quad \text{with} \quad k_W = \frac{h^2}{M_{\text{Pl}}^2} > 0$$

The superpotential is a **modular function**, having singularities at some values of  $\tau$

$$k_W \rightarrow 0 \quad \text{rigid SUSY limit}$$

# Modular invariance and SUGRA

$$W = Y_{ij}^u(\tau) U_i^c Q_j H_u + Y_{ij}^d(\tau) D_i^c Q_j H_d$$

$$Y_{ij}^q(\tau) \rightarrow (c\tau + d)^{k_{ij}^q} Y_{ij}^q(\tau) \quad \text{with} \quad k_{ij}^{u(d)} = k_{U_i^c(D_i^c)} + k_{Q_j} + k_{H_{u(d)}} - k_W$$

Furthermore, the Kähler transformation must be accompanied by a U(1) rotation

$$\psi \rightarrow e^{\frac{F - F^\dagger}{4}} \psi \quad \lambda \rightarrow e^{-\frac{F - F^\dagger}{4}} \lambda \quad \text{how gaugino enters the game}$$

$$\psi_{\text{can}} \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{\frac{k_W}{4} - \frac{k_\Phi}{2}} \psi_{\text{can}} \quad \lambda \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{-\frac{k_W}{4}} \lambda$$

Modular-QCD anomaly modifies as

$$A = \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} - 2k_W \right) + Ck_W$$

$C = 3$  is quadratic Casimir of **8** of  $\text{SU}(3)_C$

# Gluino mass

---

$$\bar{\theta} = \theta_{\text{QCD}} + \arg \det M_q + C \arg M_3$$

Assume  $k_{\text{det}} = 0$  and the quark contribution to  $A$  vanishes. Then

$$\bar{\theta} = \theta_{\text{QCD}} + C \arg M_3$$

Gluino mass requires SUSY breaking

$$M_3 = \frac{g_3^2}{2} e^{K/2M_{\text{Pl}}^2} K^{i\bar{j}} D_{\bar{j}} W^\dagger f_i$$

Assuming  $D_\tau W = 0$  and no additional phases from SUSY breaking

$$\arg M_3 = - \arg W$$

$$W = \dots + \frac{c_0 M_{\text{Pl}}^3}{\eta(\tau)^{2k_W}} \quad \text{and} \quad f = \dots + \frac{C k_W}{4\pi^2} \log \eta(\tau)$$

$$\bar{\theta} = -8\pi^2 \text{Im} f - C \arg W = 0$$

# More on modular invariance in SUGRA

---

$$\det M_q \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{\frac{k_{\text{det}}}{2}} \det M_q \quad k_{\text{det}} = \sum_{i=1}^3 \left( 2k_{Q_i} + k_{U_i^c} + k_{D_i^c} - 2k_W \right) + 3 \left( k_{H_u} + k_{H_d} \right)$$

$$M_3 \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{\frac{k_W}{2}} M_3 \quad (\text{gluino mass arises only if SUSY is broken})$$

# Finite modular groups

Infinite normal subgroups of  $SL(2, \mathbb{Z})$ ,  $N = 2, 3, 4, \dots$

$$\Gamma(N) = \left\{ \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in SL(2, \mathbb{Z}), \quad \begin{pmatrix} a & b \\ c & d \end{pmatrix} \equiv \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \pmod{N} \right\}$$

Principal congruence subgroups of the modular group

$$\bar{\Gamma}(2) \equiv \Gamma(2)/\{I, -I\} \quad \bar{\Gamma}(N) \equiv \Gamma(N), \quad N > 2$$

Finite modular groups

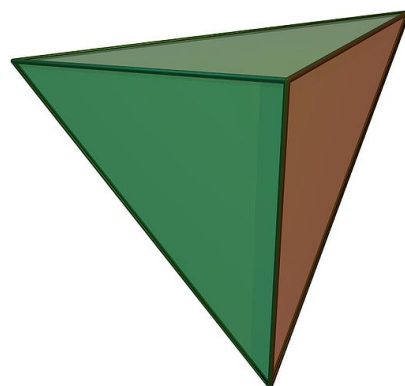
$$\Gamma_N \equiv \bar{\Gamma}/\bar{\Gamma}(N)$$

$$\Gamma_N = \langle S, T \mid S^2 = (ST)^3 = T^N = I \rangle, \quad N = 2, 3, 4, 5$$

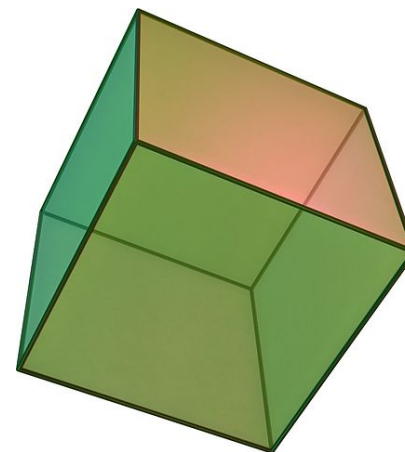
$$\Gamma_2 \cong S_3$$



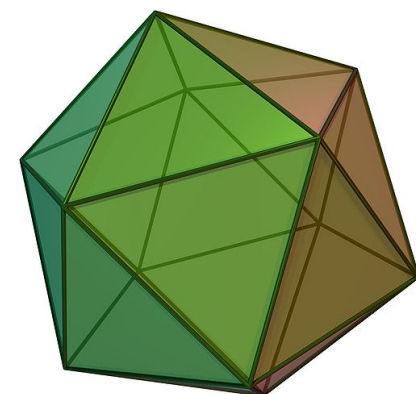
$$\Gamma_3 \cong A_4$$



$$\Gamma_4 \cong S_4$$



$$\Gamma_5 \cong A_5$$



# Theories based on finite modular groups

$\mathcal{N} = 1$  rigid SUSY matter action

$$\mathcal{S} = \int d^4x d^2\theta d^2\bar{\theta} K(\tau, \bar{\tau}, \psi, \bar{\psi}) + \int d^4x d^2\theta W(\tau, \psi) + \int d^4x d^2\bar{\theta} \bar{W}(\bar{\tau}, \bar{\psi})$$

Ferrara, Lust, Shapere, Theisen, PLB **225** (1989) 363

Ferrara, Lust, Theisen, PLB **233** (1989) 147

$$\begin{cases} \tau \rightarrow \frac{a\tau + b}{c\tau + d} \\ \psi_i \rightarrow (c\tau + d)^{-k_i} \rho_i(\tilde{\gamma}) \psi_i \end{cases} \Rightarrow \begin{cases} W(\tau, \psi) \rightarrow W(\tau, \psi) \\ K(\tau, \bar{\tau}, \psi, \bar{\psi}) \rightarrow K(\tau, \bar{\tau}, \psi, \bar{\psi}) + f_K(\tau, \psi) + \bar{f}_K(\bar{\tau}, \bar{\psi}) \end{cases}$$

Feruglio, 1706.08749

unitary representation of  $\Gamma_N$

$$W(\tau, \psi) = \sum_n \sum_{\{i_1, \dots, i_n\}} \sum_s g_{i_1 \dots i_n, s} \left( Y_{i_1 \dots i_n, s}(\tau) \psi_{i_1} \dots \psi_{i_n} \right)_{\mathbf{1}, s}$$

$$Y(\tau) \xrightarrow{\gamma} (c\tau + d)^{k_Y} \rho_Y(\tilde{\gamma}) Y(\tau)$$

$$k_Y = k_{i_1} + \dots + k_{i_n}$$

$$\rho_Y \otimes \rho_{i_1} \otimes \dots \otimes \rho_{i_n} \supset \mathbf{1}$$

Yukawa couplings are modular forms!

# Models with finite modular symmetries

$$\begin{cases} \Phi \rightarrow (c\tau + d)^{-k_\Phi} \rho_\Phi(\gamma) \Phi \\ Y \rightarrow (c\tau + d)^{k_Y} \rho_Y(\gamma) Y \end{cases}$$

Feruglio, 1706.08749

$\rho$  is a unitary representation of the finite modular group  $\Gamma_N = \Gamma/\Gamma(N)$

$$\Gamma_2 \cong S_3 \quad \Gamma_3 \cong A'_4 = T' \quad \Gamma_4 \cong S'_4 \quad \Gamma_5 \cong A'_5$$

Invariance of the superpotential  $W \sim Y(\tau) \Phi_I \Phi_J \Phi_K$  under  $\Gamma$  and  $\Gamma_N$  requires

$$\begin{cases} k_Y = k_I + k_J + k_K \\ \rho_Y \otimes \rho_I \otimes \rho_J \otimes \rho_K \supset \mathbf{1} \end{cases}$$

# Models with finite modular symmetries

$$\begin{cases} \Phi \rightarrow (c\tau + d)^{-k_\Phi} \rho_\Phi(\gamma) \Phi \\ Y \rightarrow (c\tau + d)^{k_Y} \rho_Y(\gamma) Y \end{cases}$$

Feruglio, 1706.08749

$\rho$  is a unitary representation of the finite modular group  $\Gamma_N = \Gamma/\Gamma(N)$

$$\Gamma_2 \cong S_3 \quad \Gamma_3 \cong A'_4 = T' \quad \Gamma_4 \cong S'_4 \quad \Gamma_5 \cong A'_5$$

Invariance of the superpotential  $W \sim Y(\tau) \Phi_I \Phi_J \Phi_K$  under  $\Gamma$  and  $\Gamma_N$  requires

$$\begin{cases} k_Y = k_I + k_J + k_K \\ \rho_Y \otimes \rho_I \otimes \rho_J \otimes \rho_K \supset \mathbf{1} \end{cases}$$

Penedo, Petcov, 2404.08032

Multi-dim irreps  $\rho \sim \mathbf{r}$ ,  $\mathbf{r} > 1$  for SM quarks do not allow to simultaneously

- realise the proposed mechanism for  $\bar{\theta} = 0$
- successfully describe quark masses and mixings

For the mechanism to work, SM quarks must furnish 1D irreps  $\rho \sim \mathbf{1}, \mathbf{1}', \mathbf{1}'', \dots$

# Models with finite modular symmetries

Minimal models (6 Lagrangian parameters per sector)

Penedo, Petcov, 2404.08032

$$Y^q = \begin{pmatrix} c_{11}^q & 0 & c_{13}^q F_{1*}^{(k')} + c_{13}'^q F_{1*}'^{(k')} \\ 0 & c_{22}^q & c_{23}^q F_{1*}^{(k)} \\ 0 & 0 & c_{33}^q \end{pmatrix}$$

(up to weak basis transformations)

Modular weights are large

String compactifications  
lead to smaller weights

(For models based on modular  $A_4$   
see also Petcov, Tanimoto, 2404.00858)

| Minimal models (I and II)                                |                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                   |
|----------------------------------------------------------|-------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| All $\Gamma'_N$ $(k, k') = (10, 12), (12, 14), (14, 16)$ |                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                   |
| $S_3$<br>only                                            | (10', 12), (10, 18'), (10', 18), (12, 12'),<br>(12, 14'), (12, 16'), (12', 16), (12, 20'),<br>(12', 20), (14', 16), (14, 18'), (14', 18),<br>(14, 22'), (14', 22), (16, 16'), (16', 18),<br>(16', 18'), (16, 20'), (16', 20), (18, 20'),<br>(18', 20'), (20, 20'), (20', 22), (20', 22')                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                          |
| $A'_4$<br>only                                           | (8', 12), (8', 18), (10', 12), (10, 16'),<br>(10', 16), (10, 20''), (10', 20), (12, 12'),<br>(12, 12''), (12, 14'), (12, 14''), (12, 16''),<br>(12', 16), (12'', 16'), (12, 18'), (12, 18''),<br>(12', 18), (12'', 18), (12, 22''), (12', 22),<br>(12'', 22'), (14, 16'), (14', 16), (14', 16'),<br>(14'', 16), (14'', 16'), (14', 18), (14, 20'),<br>(14, 20''), (14', 20), (14', 20''), (14'', 20),<br>(14'', 20'), (14, 24''), (14'', 24'), (16, 16''),<br>(16', 16''), (16, 18'), (16, 18''), (16', 18'),<br>(16', 18''), (16'', 18), (16'', 20'), (16, 22''),<br>(16', 22''), (16'', 22), (16'', 22'),<br>(16'', 26'), (18, 18'), (18, 18''), (18', 20),<br>(18', 20'), (18', 20''), (18'', 20), (18'', 20'),<br>(18'', 20''), (18, 22''), (18', 22), (18'', 22'),<br>(18', 24''), (18'', 24'), (20, 22''), (20', 22''),<br>(20'', 22''), (22, 22''), (22', 22''),<br>(22'', 24'), (22'', 24''), (22'', 26') |