



The University of Manchester

Towards a non-Hermitian cosmology

Many thanks to: Jean Alexandre, Richard Battye, Carl M. Bender, Maxim N. Chernodub, Edmund J. Copeland, Steven Cotterill, Madeleine Dale, John Ellis, Robert Mason, Esra Sablevice, and Dries Seynaeve

Peter Millington (UKRI Future Leaders Fellow, Particle Theory Group, University of Manchester) peter.millington@manchester.ac.uk PASCOS 2025, Durham

Monday 21 July 2025







The state of fundamental physics at 17:40 on Monday 21 July 2025 ...





The state of fundamental physics at 17:40 on Monday 21 July 2025 ...

Actually, it's time for the reception.



But first this talk: What if Hermiticity (of the Hamiltonian) turns out to be a really good thing to get rid of?

Forget Hermiticity, get pseudo-Hermiticity: antilinear symmetry, not Hermiticity.

Some pseudo-Hermitian truths and untruths ...

Truth 1: All Hermitian matrices have real eigenvalues, but not all matrices with real eigenvalues are Hermitian.

Truth 2: Hermiticity is a mathematical statement, anti-linear symmetry is a physical statement.

An untruth: Only real eigenvalues are interesting. (This talk.)



What is a pseudo-Hermitian theory?

A Hamiltonian is **pseudo-Hermitian**, if there exists a **Hermitian operator** $\hat{\eta}$ s.t.

 \hat{H}^{\dagger}

Expectation values of \hat{H} are **real** with respect to the inner product $\langle \cdot | \cdot \rangle$

and the **time evolution** is η -pseudo-unitary.

PT-(parity-time-reversal-)symmetric theories are a subset of pseudo-Hermitian theories.

$$=\hat{\eta}\hat{H}\hat{\eta}^{-1}$$

$$_{\eta} := \langle \cdot | \hat{\eta} \cdot \rangle$$

Bender and Boettcher, Phys. Rev. Lett. 80 (1998) 5243, physics/9712001; Mostafazadeh, J. Math. Phys. 43 (2002) 205, math-ph/0107001; Bender, Brody and Jones, Phys. Rev. Lett. 89 (2002) 270401, guant-ph/0208076



Example: non-Hermitian mixing of 2 complex scalar fields

$$\mathscr{L} = -\partial \tilde{\phi}_a^* \cdot \partial \phi_a - m_a^2 \tilde{\phi}_a^* \phi_a - \mu^2 (\tilde{\phi}_1^* \phi_2 - \tilde{\phi}_2^* \phi_1), \qquad a = 1, 2$$

PT-symmetric under:

$$\phi_a \to P_{ab} \tilde{\phi}_b^*, \qquad P = \pm \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} = P^{-1}$$

Pseudo-Hermitian mass matrix: $M^2 = \begin{pmatrix} m_1^2 \\ -\mu^2 \end{pmatrix}$

Eigenspectrum:
$$m_{\pm}^2 = \frac{m_1^2 + m_2^2}{2} \pm \left[\left(\frac{m_1^2 - m_2^2}{2} \right)^2 - \mu^4 \right]^{1/2} \in \mathbb{R} \text{ if } \zeta = \frac{2|\mu^2|}{|m_1^2 - m_2^2|} \le 1$$

Exceptional point at $\zeta = 1$, and **complex-conjugate pairs** of eigenvalues for $\zeta > 1$ (**PT-broken regime**).

Alexandre, PM and Seynaeve, Phys. Rev. D 96 (2017) 065027, 1707.01057; Alexandre, Ellis and PM, Phys. Rev. D 102 (2020) 12, 125030, 2006.06656; Sablevice and PM, Phys. Rev. D 109 (2024) 6, 2307.16805

$$\begin{pmatrix} \mu^2 \\ 2 \\ m_2^2 \end{pmatrix} \neq (M^2)^{\dagger}, \qquad PM^2P = (M^2)^{\dagger}$$





Example: non-Hermitian mixing of 2 complex scalar fields

$$\mathscr{L} = -\partial \tilde{\phi}_a^* \cdot \partial \phi_a - m_a^2 \tilde{\phi}_a^* \phi_a - \mu^2 (\tilde{\phi}_1^* \phi_2 - \tilde{\phi}_2^* \phi_1), \qquad a = 1, 2$$

PT-symmetric under:

$$\phi_a \to P_{ab} \tilde{\phi}_b^*, \qquad P =$$

Pseudo-Hermitian mass matrix:

Eigenspectrum:
$$m_{\pm}^2 = \frac{m_1^2 + m_2^2}{2} \pm \left[\left(\frac{m_1^2 - m_2^2}{2} \right)^2 - \mu^4 \right]^{1/2} \in \mathbb{R} \text{ if } \zeta = \frac{2}{|m|}$$

Exceptional point at $\zeta = 1$, and **complex-conjugate pairs** of eigenvalues for

Alexandre, PM and Seynaeve, Phys. Rev. D 96 (2017) 065027, 1707.01057; Alexandre, Ellis and PM Sable

$$\pm \begin{pmatrix} 1 & 0\\ 0 & -1 \end{pmatrix} = P^{-1}$$

$$P_{ab}\tilde{\phi}_{b}^{*}, \quad P = \pm \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} = P^{-1}$$

$$M^{2} = \begin{pmatrix} m_{1}^{2} & \mu^{2} \\ -\mu^{2} & m_{2}^{2} \end{pmatrix} \neq (M^{2})^{\dagger}, \quad PM^{2}P = (M^{2})^{\dagger} \quad \text{isomorphotonom}$$

$$\frac{m_{2}^{2}}{2} \pm \left[\left(\frac{m_{1}^{2} - m_{2}^{2}}{2} \right)^{2} - \mu^{4} \right]^{1/2} \in \mathbb{R} \text{ if } \zeta = \frac{2}{|m|} \quad \text{isomorphotonom}$$
where regime isomorphoton is the set of the

Jken regime).

<u>(v) 12, 125030, 2006.06656;</u> <u>. D 109 (2024) 6, 2307.16805</u>

Ì





Back to representations of the Poincaré group

Consider the **Heisenberg equation** for a scalar field operator $\hat{\phi}(t, \vec{x})$:

$$\left[\hat{\phi}(t,\vec{x}),\hat{H}\right] = i\partial_t\hat{\phi}(t,\vec{x}) \Rightarrow$$

Thus, if $\hat{P}^0 = \hat{H}$ is not Hermitian, so too are other **generators of the Poincaré group**, and $\hat{\phi}^{\dagger}(t,\vec{x})\hat{\phi}(t,\vec{x})$ does **not** transform properly under the **Poincaré group**.

Which is the operator $\hat{\phi}^{\dagger}(x)$ transforming in the representation "dual" to $\hat{\phi}(x)$?

Chernodub and PM, Phys. Rev. D 105 (2022) 7, 076020, 2110.05289; Sablevice and PM, Phys. Rev. D 109 (2024) 6, 2307.16805

$$\left[\hat{\phi}^{\dagger}(t,\vec{x}),\hat{H}^{\dagger}\left(\neq\hat{H}\right)\right]=i\partial_{t}\hat{\phi}^{\dagger}(t,\vec{x})$$

- $\hat{\phi}(t,\vec{x})$ evolves with the Hamiltonian $\hat{H} \Rightarrow \hat{\phi}^{\dagger}(t,\vec{x})$ evolves with the Hamiltonian \hat{H}^{\dagger}



The dual field operator is

$$\hat{\tilde{\phi}}^{\dagger}(x) =$$
the same map η that det
pseudo-Hermiticity of the



Sablevice and PM, Phys. Rev. D 109 (2024) 6, 2307.16805



An observation about field theory on dynamical spacetimes: Eigenspectra are, in general, not real.

So this talk: Can we exploit the PT-broken regime for cosmology?

Example: a non-Hermitian theory with symmetry breaking

$$-\mathscr{L} = -g^{\rho\sigma}\partial_{\rho}\tilde{\phi}_{a}^{\dagger}\partial_{\sigma}\phi_{a} + m_{1}^{2}\tilde{\phi}_{1}^{\dagger}\phi_{1} - m_{2}^{2}\tilde{\phi}_{2}^{\dagger}\phi_{2} - \mu^{2}\left(\tilde{\phi}_{1}^{\dagger}\phi_{2} - \epsilon\tilde{\phi}_{2}^{\dagger}\phi_{1}\right) - \frac{\lambda}{4}\left(\tilde{\phi}_{1}^{\dagger}\phi_{1}\right)^{2}$$

Non-Hermitian for $\epsilon = +1$; **Hermitian** for $\epsilon = -1$. (Note the overall sign of the Lagrangian.)

The U(1)-breaking minima of the potential lie at

$$\phi_1 \to v_1 = \sqrt{(2/\lambda)(m_1^2 - \epsilon \mu^4/m_2^2)} \qquad \tilde{\phi}_1^\dagger \to v_1 \qquad \phi_2 \to v_2 = \epsilon \mu^2 v_1/m_2^2 \qquad \tilde{\phi}_2^\dagger \to -\epsilon v_2$$

Expanding around the minima in Minkowski spacetime, the eigenmasses are

$$M_G^2 = 0 \qquad M_0^2 = m_2^2 - \frac{\epsilon\mu^2}{m_2^2} \qquad M_{\pm}^2 = \frac{m_2^2}{2} \left[1 + 2\frac{m_1^2}{m_2^2} + 3\frac{\epsilon\mu^4}{m_2^4} \pm \sqrt{\left(1 - 2\frac{m_1^2}{m_2^2}\right)^2 + 2\left(1 - 6\frac{m_1^2}{m_2^2}\right)\frac{\epsilon\mu^4}{m_2^4} + 9\frac{\epsilon^2\mu^8}{m_2^8}} \right]$$

There is an **exceptional point** at $\mu^2/m^2 = 1$, where $M_0^2 \to M_G^2$.

Alexandre, Ellis, PM and Seynaeve, Phys. Rev. D 98 (2018) 045001, 1805.06380; Mannheim, Phys. Rev. D 99 (2019) 4, 045006, 1808.00437; see also related works by Fring and Taira



Now on FLRW ...

In conformal time τ , the de-Sitter phase FLRW line element $ds^2 = 1/(H\tau)^2(-d\tau^2 + d\vec{x}^2)$ is consistent with anti-linear symmetry under $\tau \mapsto -\tau$ and $\vec{x} \mapsto -\vec{x}$.

Defining $\varphi_i = \phi_i / v_i$ and $\varphi_i^{\dagger} = \tilde{\phi}_i^{\dagger} / \tilde{v}_i$ (with $\tilde{v}_1 = v_1$ and $\tilde{v}_2 = -\epsilon v_2$), and $\bar{H} = H/m_2$, $\bar{m} = m_1/m_2$, $\bar{\mu} = \mu^2/m_2^2$ and $\overline{t} = m_2 t$, the equations of motion and Friedmann constraint take simpler forms:

$$-\ddot{\varphi}_1 - 3\bar{H}\dot{\varphi}_1 + \bar{m}^2\varphi_1 - \epsilon\bar{\mu}^2\varphi_2 - \left(\bar{m}^2 - \epsilon\bar{\mu}^2\right)\varphi_1^{\dagger}\varphi_1^2 = 0$$

$$-\ddot{\varphi}_2 - 3\bar{H}q$$

$$\bar{H}^{2} = \frac{8\pi G v_{1}^{2}}{3} \left[\epsilon \bar{\mu}^{2} \dot{\varphi}_{2}^{\dagger} \dot{\varphi}_{2} - \dot{\varphi}_{1}^{\dagger} \dot{\varphi}_{1} + \bar{m}^{2} |\varphi_{1}|^{2} + \epsilon \bar{\mu}^{2} |\varphi_{2}|^{2} - \epsilon \bar{\mu}^{2} \left(\varphi_{1}^{\dagger} \varphi_{2} + \varphi_{2}^{\dagger} \varphi_{1} \right) - \frac{1}{2} \left(\bar{m}^{2} - \epsilon \bar{\mu}^{2} \right) |\varphi_{1}|^{4} \right]$$

 $\dot{\phi}_{2} - \phi_{2} + \phi_{1} = 0$



At late times ...

In terms of **radial** and **azimuthal field components** ($\varphi_a = R_a e^{i\theta_a}$), and assuming: R_1 , R_2 and \bar{H} constant, and $\ddot{\theta}_a \rightarrow 0$ at late times, the asymptotic behaviour is governed by

$$R_1\dot{\theta}_1^2 + \bar{m}^2 R_1 - \epsilon \bar{\mu}^2 R_2 \cos \delta\theta - \left(\bar{m}^2 - \epsilon \bar{\mu}^2\right) R_1^3 = 0 \qquad R_2\dot{\theta}_2^2 - R_2 + R_1 \cos \delta\theta = 0$$
$$-3\bar{H}R_1\dot{\theta}_1 + \epsilon \bar{\mu}^2 R_2 \sin \delta\theta = 0 \qquad -3\bar{H}R_2\dot{\theta}_2 + R_1 \sin \delta\theta = 0$$

and the solution for the **relative phase** $\delta\theta = \theta_1 - \theta_2$ is

$$\delta\theta = 2\operatorname{arccot}\left\{A \exp\left[\left(\frac{R_1}{R_2} - \epsilon \bar{\mu}^2 \frac{R_2}{R_1}\right) \frac{\bar{t}}{3\bar{H}}\right]\right\} = \begin{cases} 0, & \epsilon \mu^4 / m_2^4 \frac{R_2^2}{R_1^2} < 1 \text{ (PT unbroken at vevs)} \\ \operatorname{const.}, & \epsilon \mu^4 / m_2^4 \frac{R_2^2}{R_1^2} \ge 1 \text{ (PT broken at vevs)} \end{cases}$$



Self-sustaining accelerated expansion

This late-time attractive behaviour is realised numerically:







Growth instability vs. Hubble damping

In the **PT-broken regime** (with complex eigenspectrum), the eigenfrequency of one of the fluctuations about the minimum of the potential has a **positive imaginary part**:

$$\bar{\omega}_{3(4)} = -\frac{3i\bar{H}}{2} + (-)\frac{i}{2}\sqrt{9\bar{H}^2 + 4\bar{\mu}^2 - 4}$$

Late-time equation of motion for the phases is

$$\dot{\theta} = \text{const}$$
.

 \Rightarrow effectively subject to a linear potential.

cf. M. C. Escher's "Ascending and Descending"



This model (and similar ones) also support defects.

Vortex solutions found by Begun, Chernodub and Molochkov (Phys. Rev. D 104 (2021) 5, 056024, 2105.07453).

See also Fring and Taira, Phys. Lett. B 807 (2020) 135583, 2006.02718; J. Phys. A 53 (2020) 45, 455701, 2007.15425; Correa, Fring and Taira, Nucl. Phys. B 971 (2021) 115516, 2102.05781; Nucl. Phys. B 979 (2022) 115783, 2110.06825; JHEP 10 (2022) 109, 2208.03199

Revisiting a similar model, with quartic self-couplings for both fields, the usual **vortex Ansatz** $\phi_a = f_a(r)e^{in\theta}$ leads to the equations of motion

$$\frac{\mathrm{d}^2 f_1}{\mathrm{d}r^2} + \frac{1}{r} \frac{\mathrm{d}f_1}{\mathrm{d}r} - \frac{n^2}{r^2} f_1 + m_1^2 f_1 - \mu^2 f_2 - 2\lambda_1 f_1^3$$

$$\frac{\mathrm{d}^2 f_2}{\mathrm{d}r^2} + \frac{1}{r} \frac{\mathrm{d}f_2}{\mathrm{d}r} - \frac{n^2}{r^2} f_2 - m_2^2 f_2 + \mu^2 f_1 - 2\lambda_2 f_2^3$$



Battye, Cotterill and PM in prep 2025







We find that these solutions are metastable. (Preliminary.)

Despite being in the **PT-symmetric regime**, we find evidence of an **unstable fluctuation**, which may be due to the loss of spatial homogeneity via an effect identified by Chernodub and PM. See Chernodub and PM, Phys. Rev. D 105 (2022) 7, 076020, 2110.05289; Phys. Rev. D 109 (2024) 10, 105006, 2401.06097.

This means that the string networks (right) may be metastable, with intriguing possibilities for cosmology.





Battye, Cotterill and PM in prep 2025

Closing remarks

- We can relax Hermiticity and still realise physically viable theories.
- The **PT-broken regime** can be viable in **dynamical spacetimes**.
- Example: self-sustaining cosmological expansion from PT-breaking.
- Example: metastable networks of topological defects.
- Phenomenologically relevant pseudo-Hermitian extensions of the Standard Models of particle physics and cosmology ...

Backup slides

Some intuition from thinking about QFTs as (X)QMs

Proximity to the exceptional points is determined by the parameters of the theory.

- In QM: We are in (1) the regime of exact antilinear symmetry (real eigenvalues), (2) at an exceptional point, or (3) in the broken regime (complex conjugate eigenvalues).
- Modes do not need to be in the same regime of the antilinear symmetry! In QFT:

To see this, we just need to make the parameters of the theory spacetime dependent ...



Local similarity transformations

$$\mathscr{L} = \bar{\psi} \left(i\gamma \cdot \partial - m - m_5 \gamma^5 \right) \psi$$

by the similarity transformation

$$\psi \to e^{\omega_5 \gamma^5} \psi, \qquad \bar{\psi} \to \bar{\psi} e^{\omega_5 \gamma^5}.$$

one and introduce a "similarity gauge field" C_{μ} with $C_{\mu} \rightarrow C_{\mu} - \partial_{\mu}\omega_5$:

$$\mathscr{L} = \bar{\psi}(i\gamma \cdot \partial + i\gamma \cdot C - m - m_5\gamma^5)\psi$$

Chernodub and PM, Phys. Rev. D 105 (2022) 076020, 2110.05289; Phys. Rev. D 109 (2024) 105006, 2401.06097; similar vector fields found in non-Hermitian holographic theories, Morales-Tejera and Landsteiner, SciPost Phys. 14 (2023) 030, 2203.02524

$M^2 = m^2 - m_5^2$ with eigenspectrum

Bender, Jones and Rivers, Phys. Lett. B 625 (2005) 333, hep-th/0508105

The theory belongs to a one-parameter family of similar theories $m + m_5 \gamma^5 = M e^{2\gamma^5 \theta}$ related

If we make $\omega_5 \equiv \omega_5(x)$ local, we must promote the global similarity transformation to a local





Two-flavour fermion mixing

$$\mathscr{L} = \bar{\psi}_a \left(i\gamma \cdot \partial - m_a \right) \psi_a - m_5 \left(\bar{\psi}_1 \gamma^5 \psi_2 + \bar{\psi}_2 \gamma^5 \psi_1 \right) \quad \equiv \quad \mathscr{L} = \bar{\Psi} \left(i\gamma \cdot \partial - M \right) \Psi$$

$$M = \begin{pmatrix} m_1 & m_5 \\ -m_5 & m_2 \end{pmatrix} \quad \text{with eigenspectrum} \quad M_{\pm} = \frac{m_1 + m_2}{2} \pm \left[\left(\frac{m_1 - m_2}{2} \right)^2 - m_5^2 \right]^{1/2}$$

Now make the mass parameters local s.t. $M_{\pm} = M_0 \pm m_0$ remain constant:

$$\hat{M}(x) = M_0 \mathbb{I} + m_0 [\sigma_3 c$$

Taking a time-local $\theta(x) = \theta_0 + C_0 x^0$ (time-like similarity gauge field), the **dispersion relation** is

$$\omega_{\pm,\mathbf{p}} = \mathbf{p}^2 + M_0^2 + m_0^2 - C_0^2 \pm 2\sqrt{M_0^2(m_0^2 - C_0^2) - C_0^2 \mathbf{p}^2}$$

 $\cosh(2\theta(x)) + i\sigma_2 \sinh(2\theta(x))]$

Chernodub and PM, Phys. Rev. D 109 (2024) 105006, 2401.06097

Inhomogeneous two-flavour fermion mixing

Momentum-dependent exceptional point: $p = p_c = M_0 \sqrt{\frac{m_0^2}{C_0^2} - 1}$

Superluminal propagation: $p > p_{SL}$, $v_p = \frac{\partial \omega_{\pm,p}}{\partial p} > 1$

Stopped propagation: $p = p_{stop} = \frac{1}{|C_0|} \sqrt{M_0^2 (m_0^2 - C_0^2) - C_0^4}$

Negative group velocities: $p_{stop} , <math>v_{+,p} < 0$

Anomalous dispersion:
$$\frac{\partial n_{-}}{\partial p} < 0$$



Chernodub and PM, Phys. Rev. D 109 (2024) 105006, 2401.06097



Phase diagram



Chernodub and PM, Phys. Rev. D 109 (2024) 105006, 2401.06097

Pseudo-reality

We can (in fact have to) generalise what we mean by a real field:

$$\hat{\phi}^{\dagger}(x) = \hat{\eta}^{-1} \hat{\phi}^{\dagger}(x_{\eta}) \hat{\eta} \pi = \hat{\phi}(x) \in \mathbb{R}_{\eta}$$

- complex field space, such that we still have **one degree of freedom**.
- Similarity transformations of pseudo-Hermitian theories map this hypersurface.
- This is the self-consistent resolution of the so-called **Hermiticity puzzle**.

Two degenerate constraints (real and imaginary parts) define a hypersurface in the

See Mannheim, Phys. Rev. D 99 (2019) 045006, 1808.00437

• Import: We can treat continuous symmetries in pseudo-Hermitian QFTs consistently.

Chernodub, PM and Sablevice, to appear in Phys. Rev. D (2025), 2501.09111



