

Low-Scale Leptogenesis in the ν SMEFT

Sascha Weber

JGU Mainz

In collaboration with
Kaori Fuyuto (LANL) and Julia Harz (JGU)
[In preparation]

Another paper

Asymgenesis

Martin A. Mojahed^{1,2,*} and Sascha Weber^{2,†}

¹*Physics Department T70, Technical University of Munich, 85748 Garching, Germany*

²*PRISMA⁺ Cluster of Excellence & Mainz Institute for Theoretical Physics, FB 08 - Physics, Mathematics and Computer Science, Johannes Gutenberg-Universität Mainz, Staudingerweg 9, 550099 Mainz, Germany*

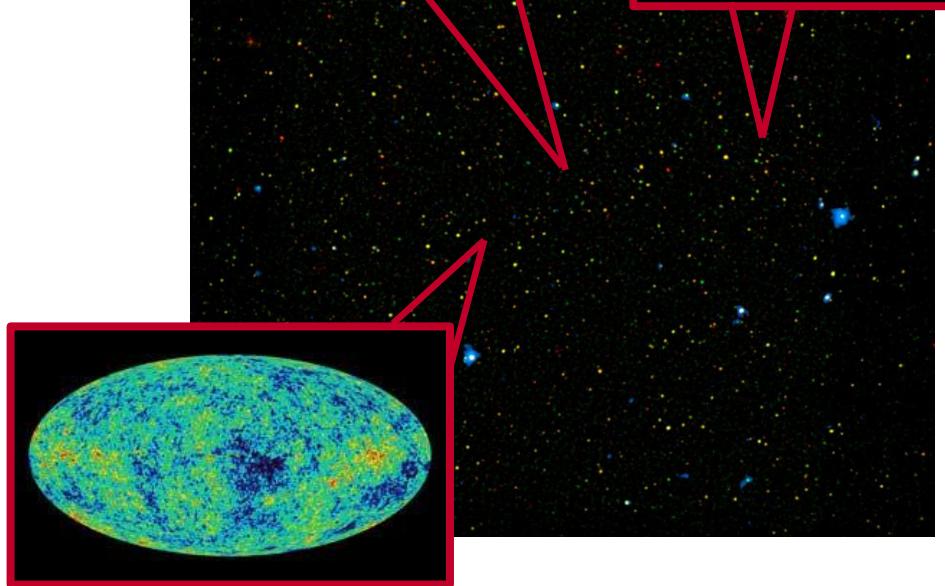
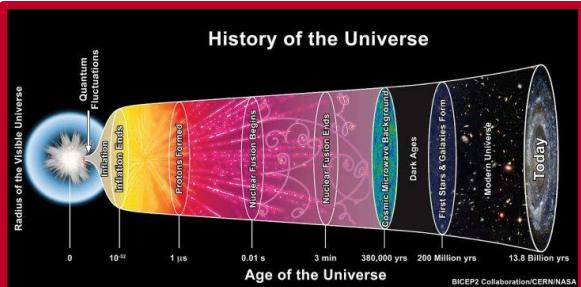
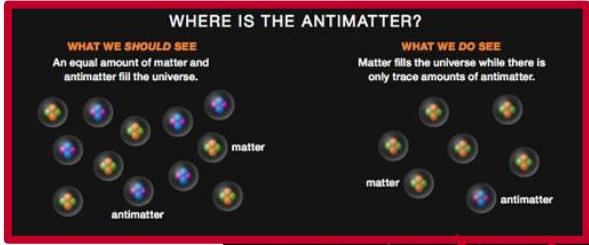
(Dated: July 16, 2025)

We present a framework based on the standard type-I seesaw model that relates the baryon asymmetry of the universe to the dark matter (DM) density. The framework, which we name *Asymgenesis*, relies on the presence of primordial charge asymmetries seeded either in the dark sector or in the visible sector. A higher-dimensional portal operator reshuffles this initial asymmetry into both sectors, eventually resulting in a nonzero $B - L$ asymmetry and an asymmetric DM component. Compared to conventional asymmetric-dark-matter (ADM) schemes, our framework imposes far milder requirements on the portal interaction. In particular, the portal interaction need not violate $B - L$, and the temperature scales of efficient $B - L$ violation and efficient charge-transfer interaction mediated by the portal operator can be separated. We develop the formalism in detail and argue that the flexibility of our framework enlarges the model-building landscape for ADM.

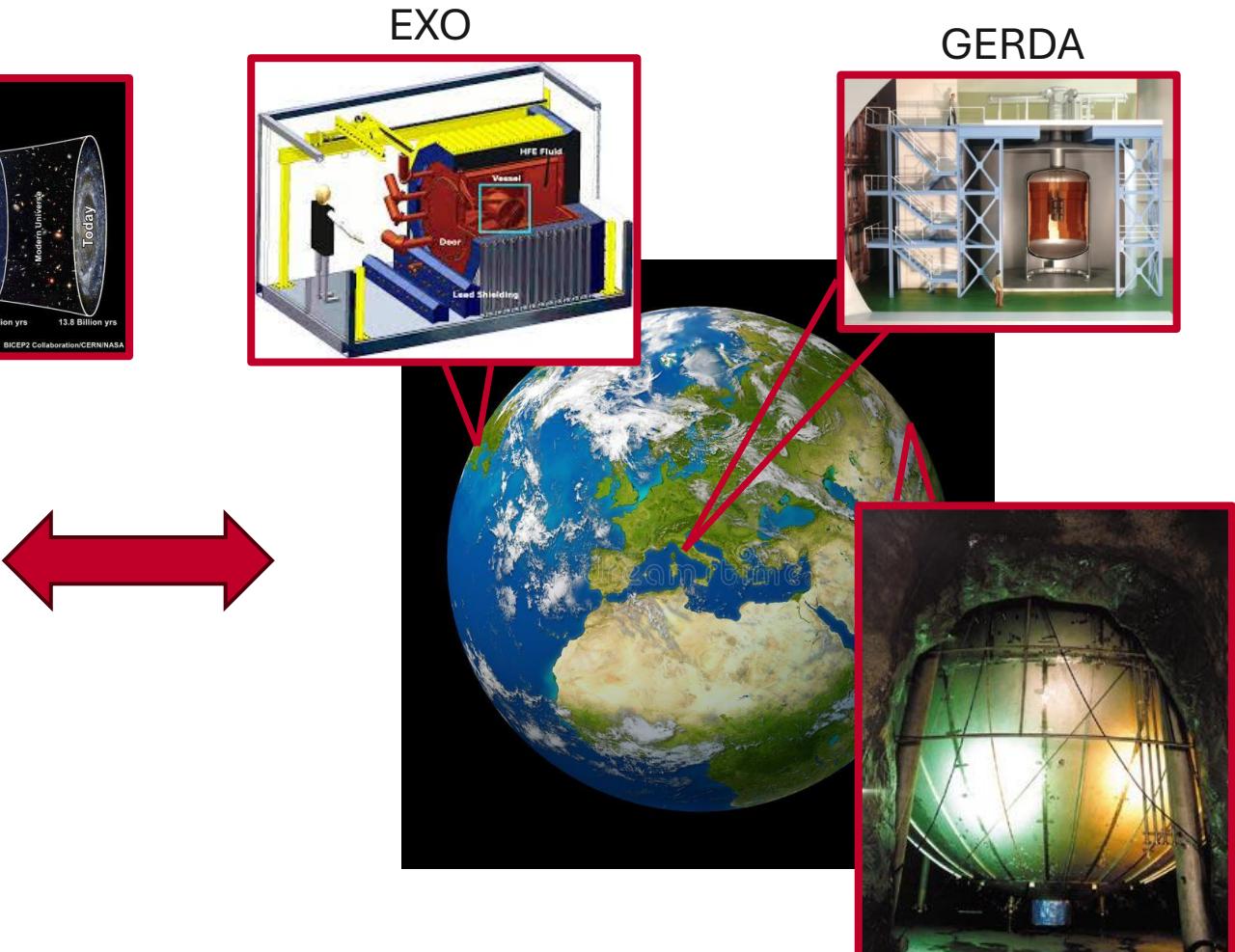
[M.Mojahed and S.W. arXiv:2507.10655]

Motivation

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[<https://www.mpi-hd.mpg.de/gerda/>]
[<https://www-project.slac.stanford.edu/exo/about.html>]
[<https://cerncourier.com/a/kamland-experiment-discovers-that-reactor-antineutrinos-disappear/>]



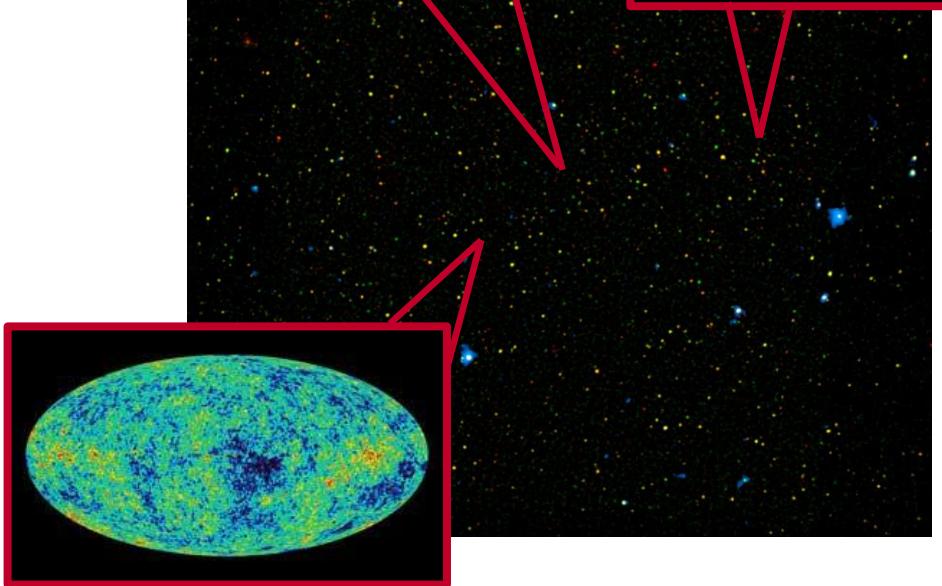
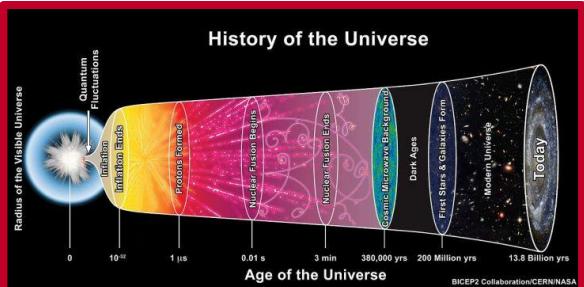
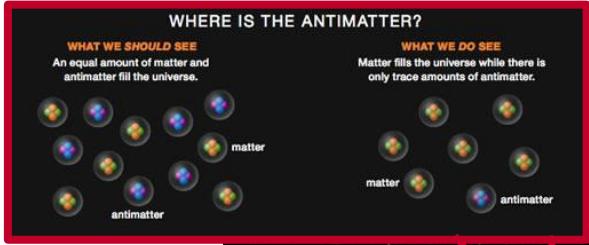
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[<http://www.spaceandmotion.com/cosmic-microwave-background-radiation.htm>]
[<https://www.astroblogs.nl/2013/03/23/wordt-het-universum-geregeerd-door-antineutrinos/baryon-asymmetry/>]
[https://de.m.wikipedia.org/wiki/Datei:The_History_of_the_Universe.jpg]



KamLAND-Zen

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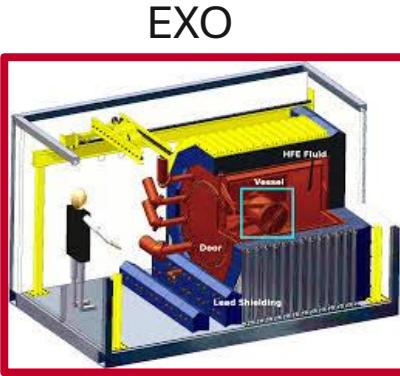
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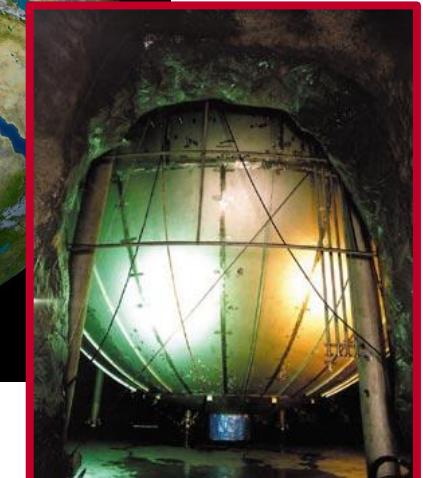
How robust?



GERDA



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Low-Scale Leptogenesis

Baryogenesis via neutrino oscillations

E. Kh. Akhmedov^(a,b) V. A. Rubakov^(c,a,d) and A. Yu. Smirnov^(a,c)

The ν MSM, Dark Matter and Baryon Asymmetry of the Universe

Takehiko Asaka* and Mikhail Shaposhnikov†

Kinetic Equations for Baryogenesis via Sterile Neutrino Oscillation

Takehiko Asaka^{1,2}, Shintaro Eijima^{2,3} and Hiroyuki Ishida^{2,3}

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Laurent Canetti^a, Marco Drewes^{b,c}, Mikhail Shaposhnikov^a

Uniting low-scale leptogeneses

Juraj Klarić,¹ Mikhail Shaposhnikov,¹ and Inar Timiryasov¹

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Low-scale leptogenesis with three heavy neutrinos

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How robust?

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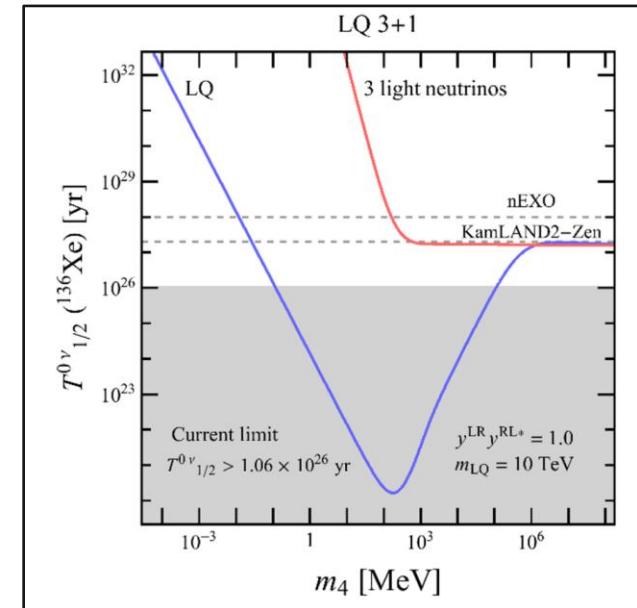
REVISED: May 6, 2020

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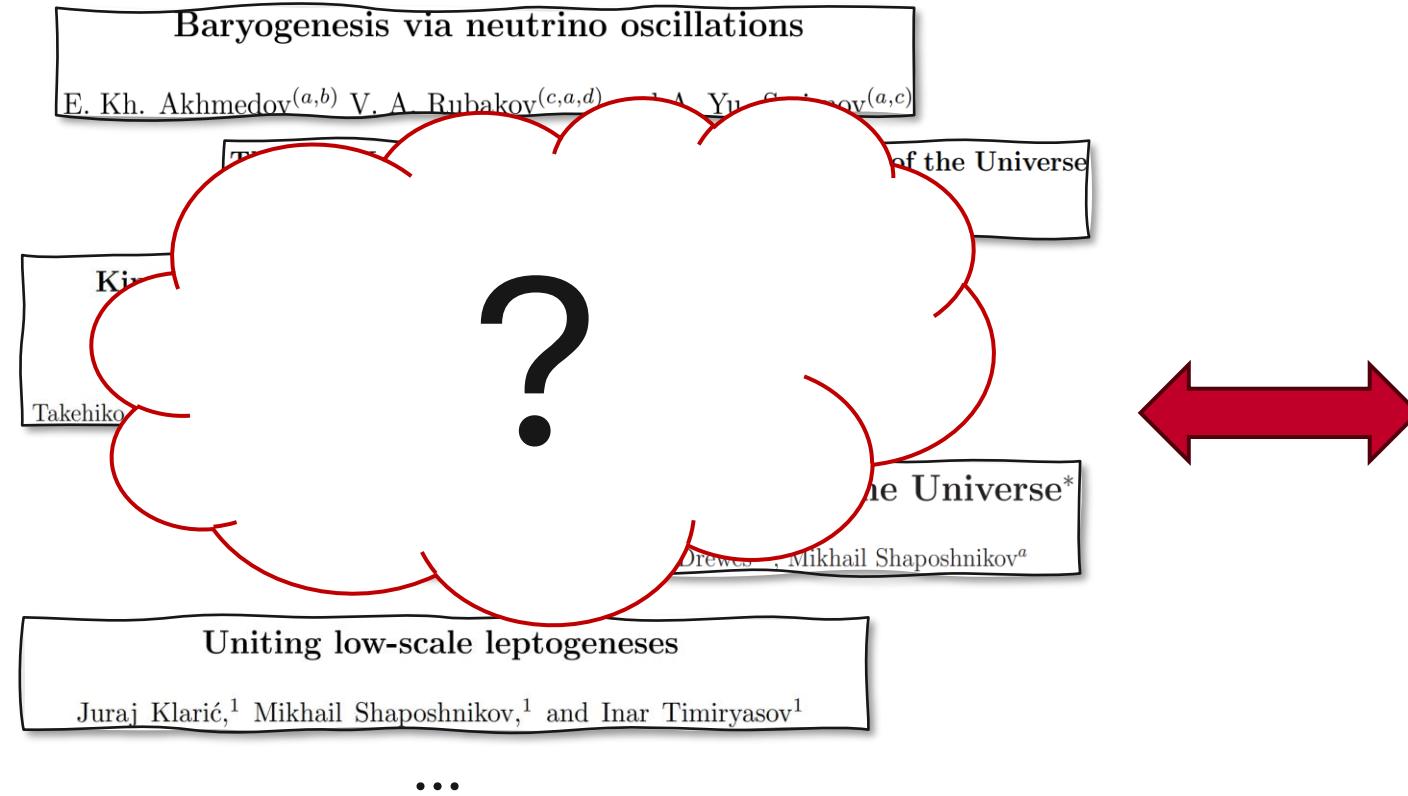
Sterile neutrinos and neutrinoless double beta decay in effective field theory

W. Dekens,^a J. de Vries,^{b,c} K. Fuyuto,^{b,d} E. Mereghetti^d and G. Zhou^b

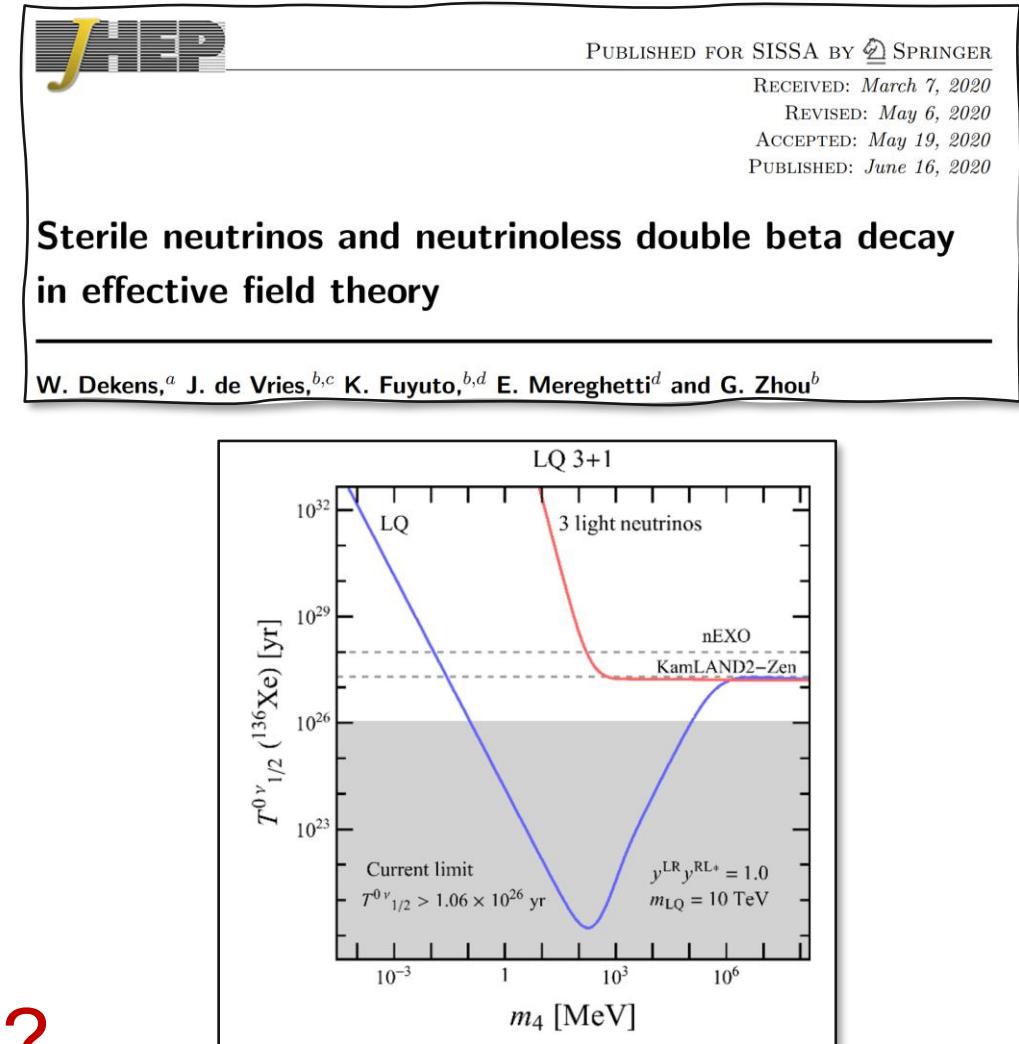


Low-Scale Leptogenesis

[Dekens et. al. JHEP 2020]

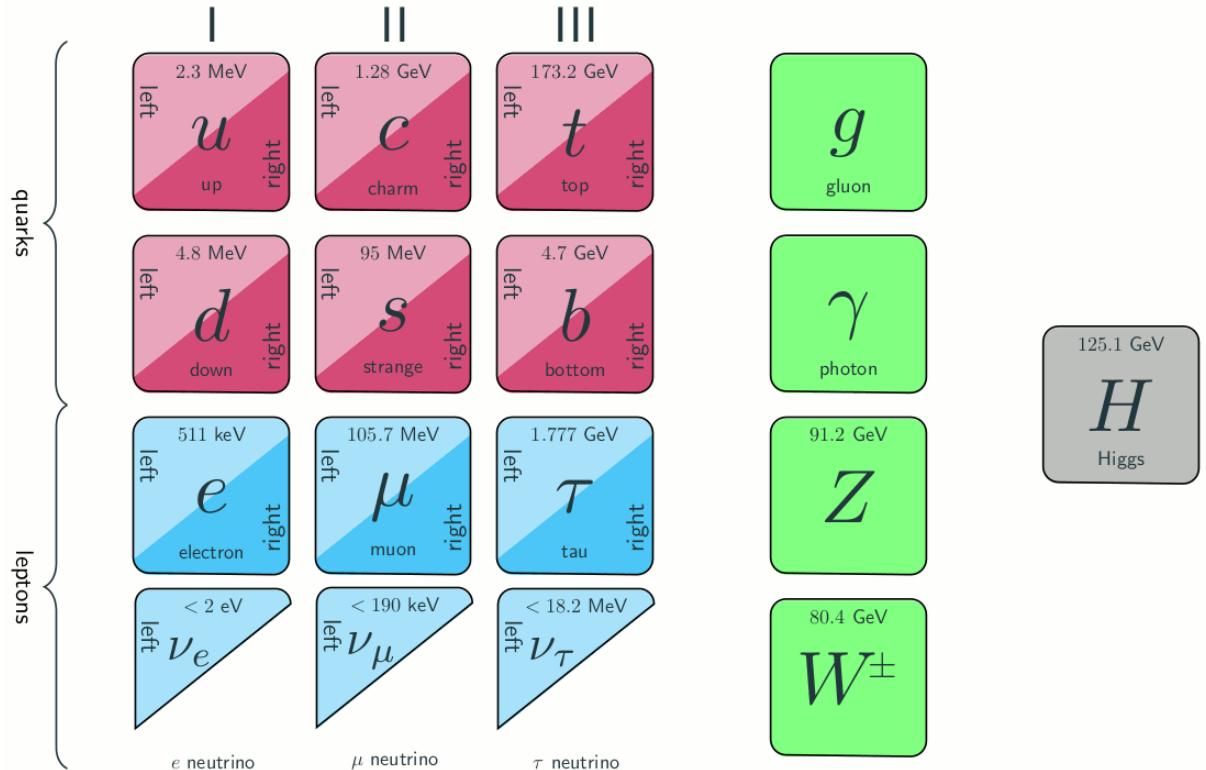


How robust?



Standard Case

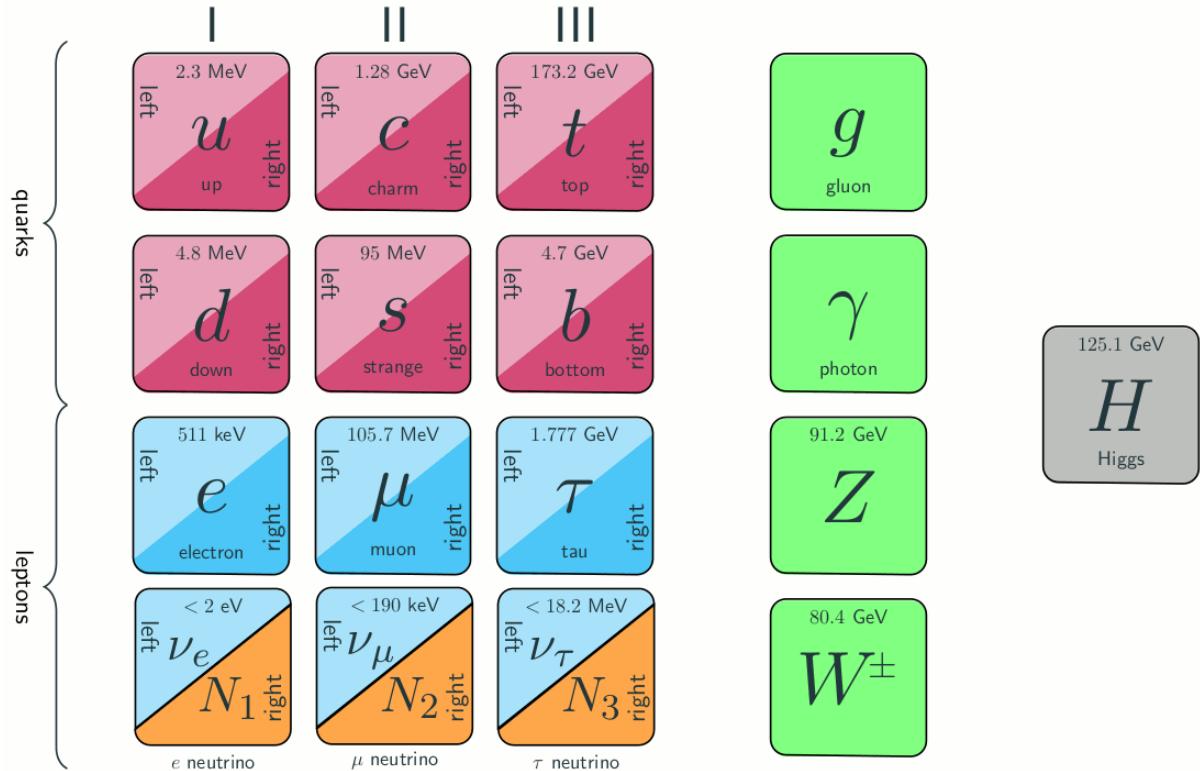
Underlying model



$$\mathcal{L} = \mathcal{L}_{\text{SM}}$$

[<https://ep-news.web.cern.ch/uniting-leptogeneses>]

Underlying model

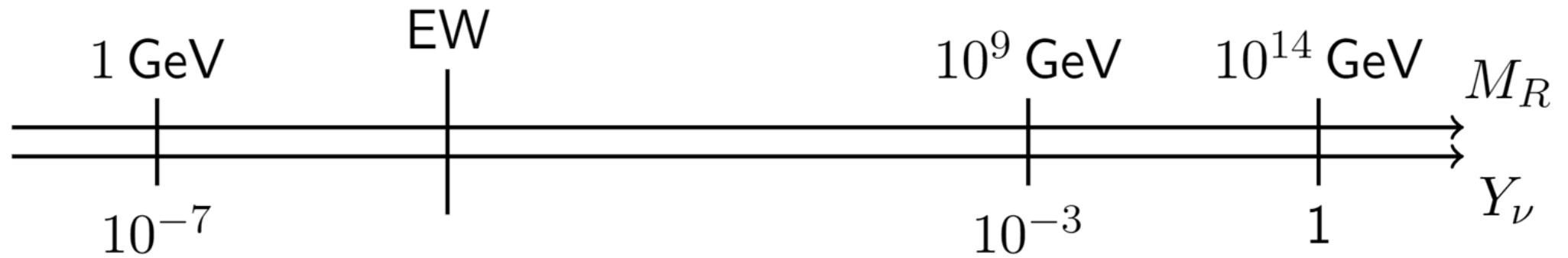


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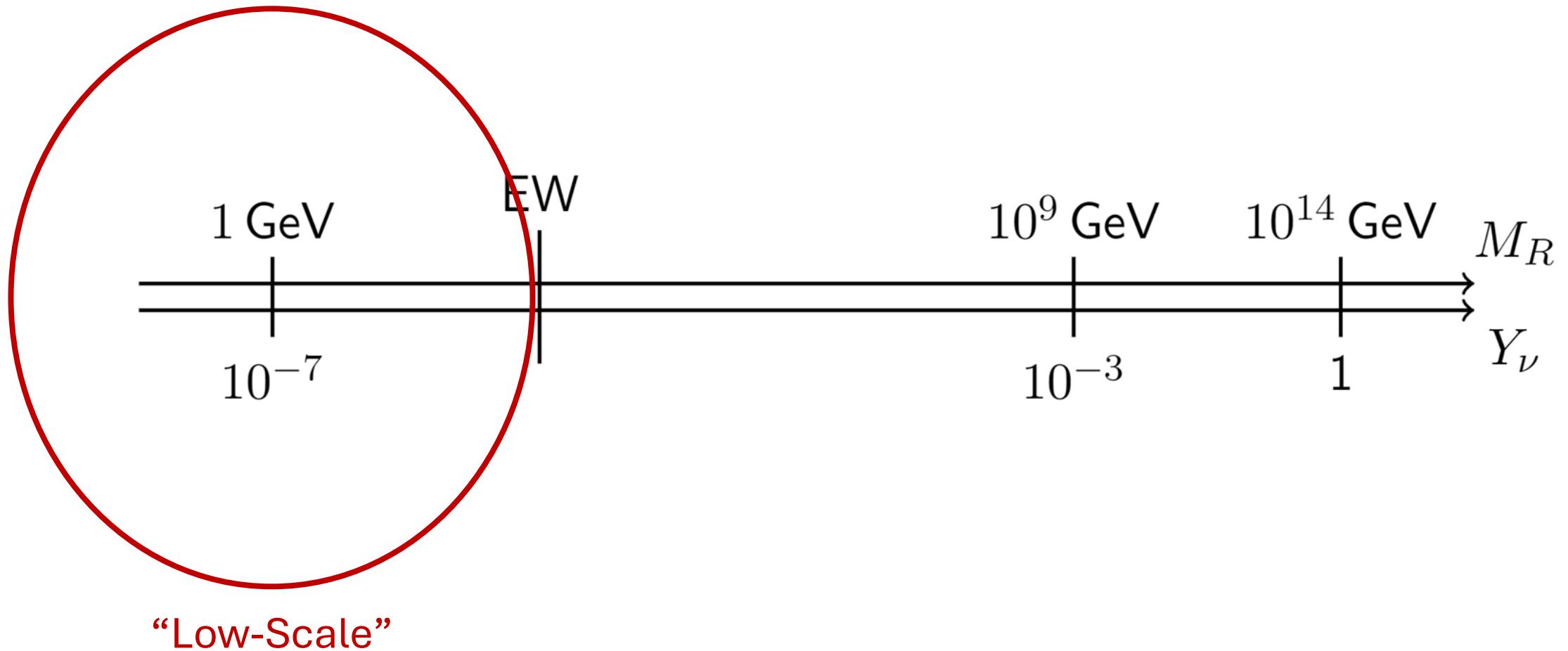
$$\mathcal{L} = \mathcal{L}_{\text{SM}}$$

$$+ \mathcal{L}_N \left\{ \begin{array}{l} + \bar{N}(i\partial)N \\ - Y_{i\alpha} \bar{N}_i H L_\alpha + \text{h.c.} \\ - \bar{N}_i^c M_i N_i + \text{h.c.} \end{array} \right.$$

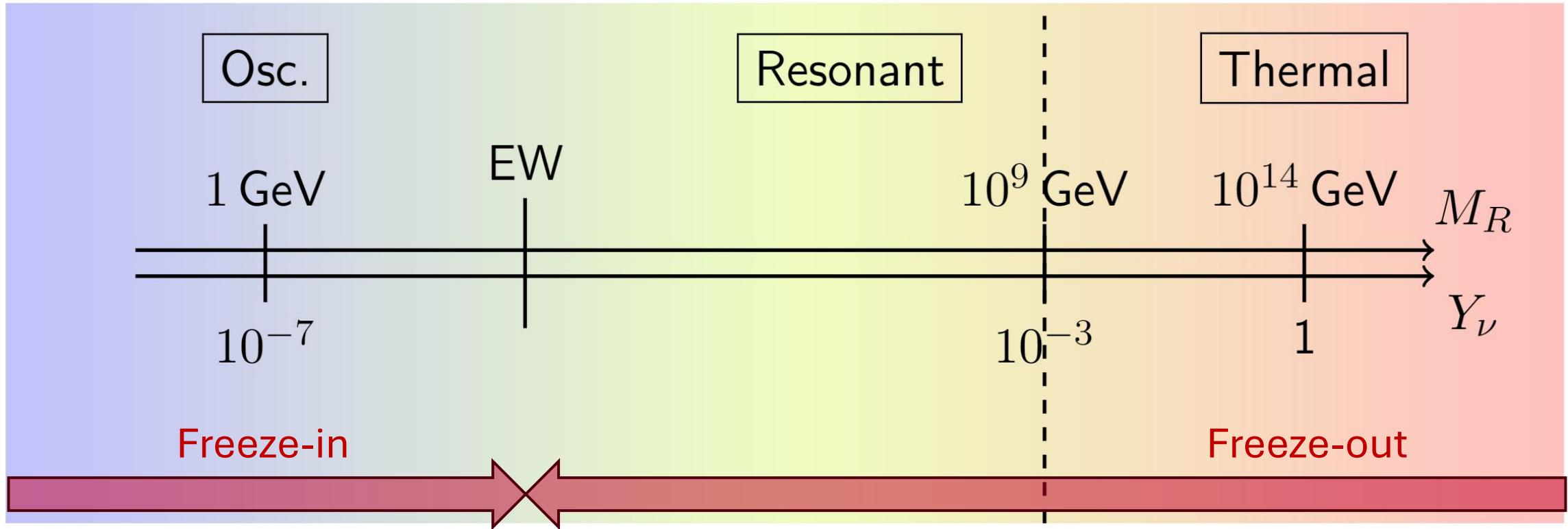
Mass Range



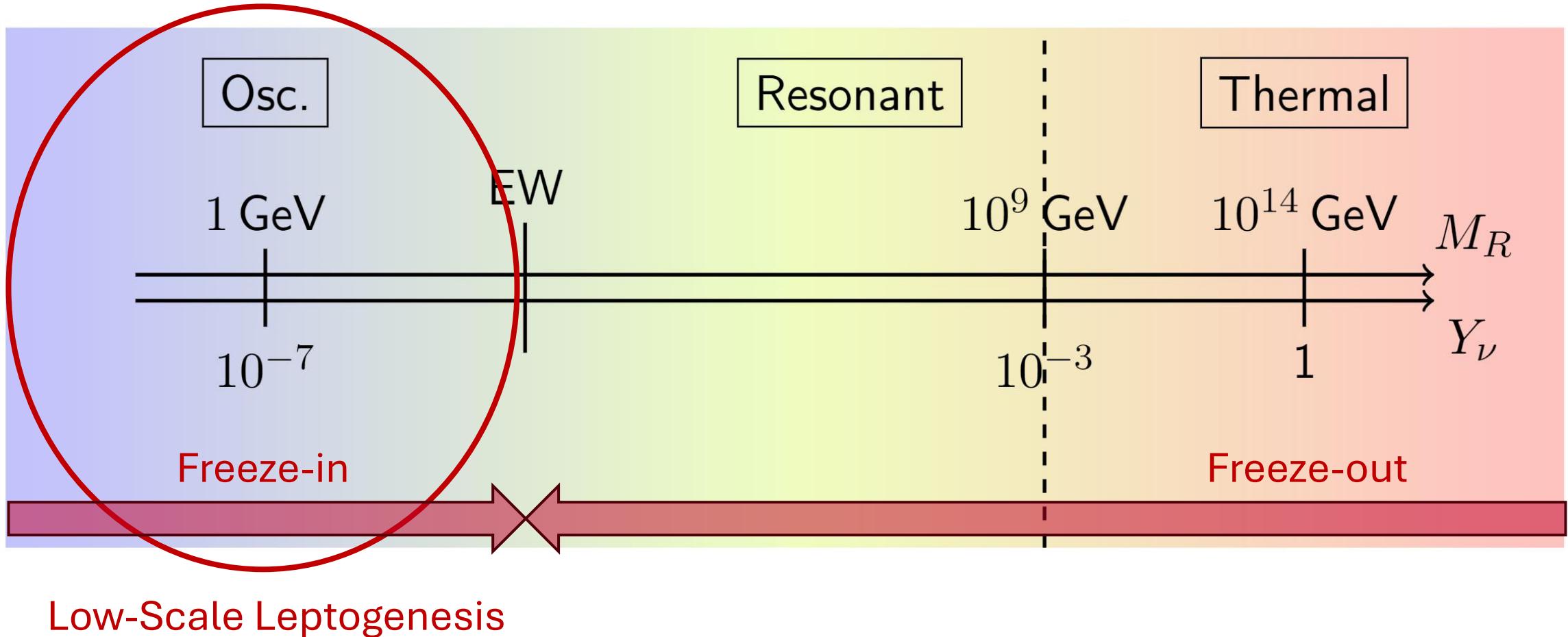
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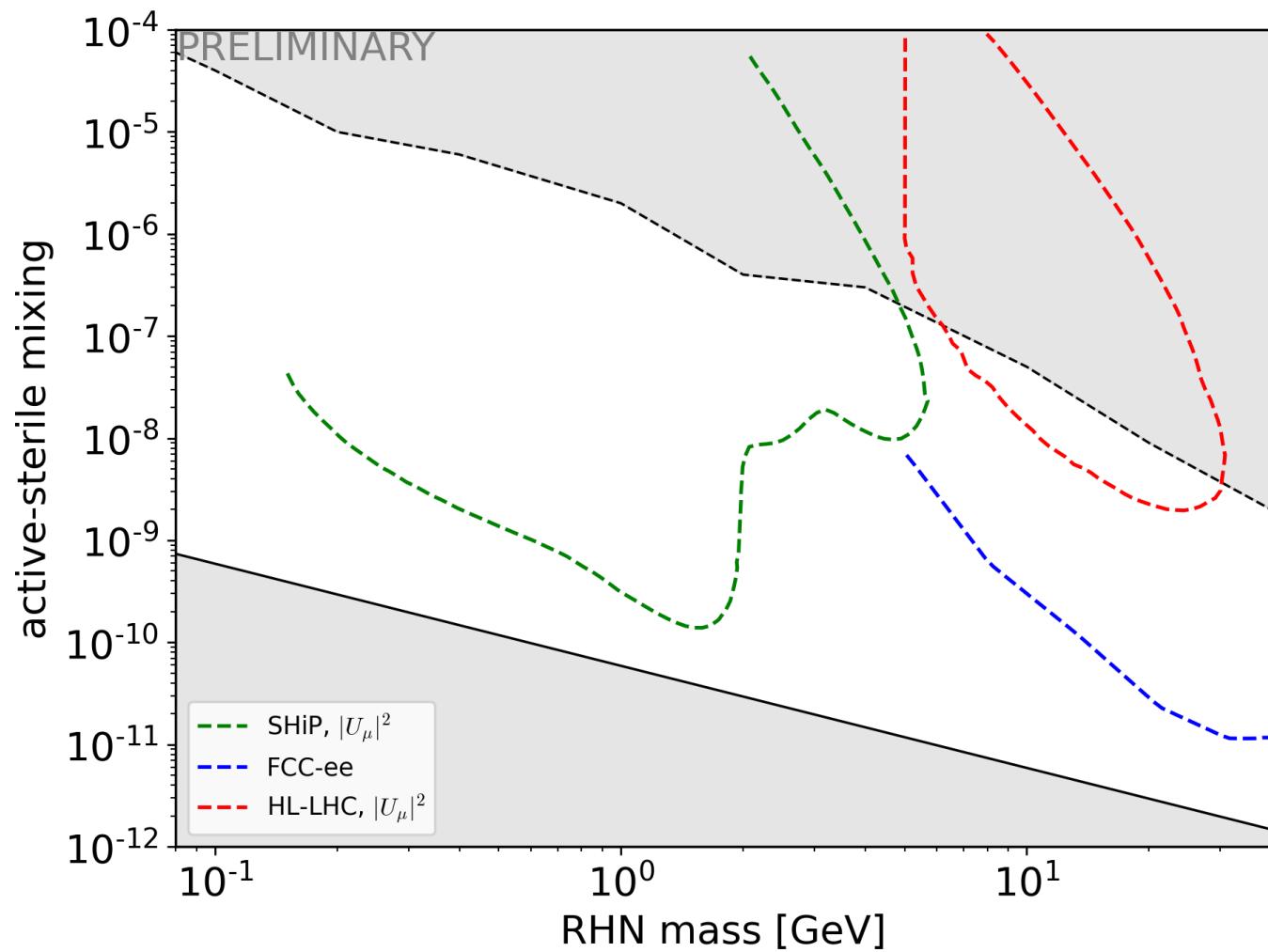
Leptogenesis regimes



Leptogenesis regimes

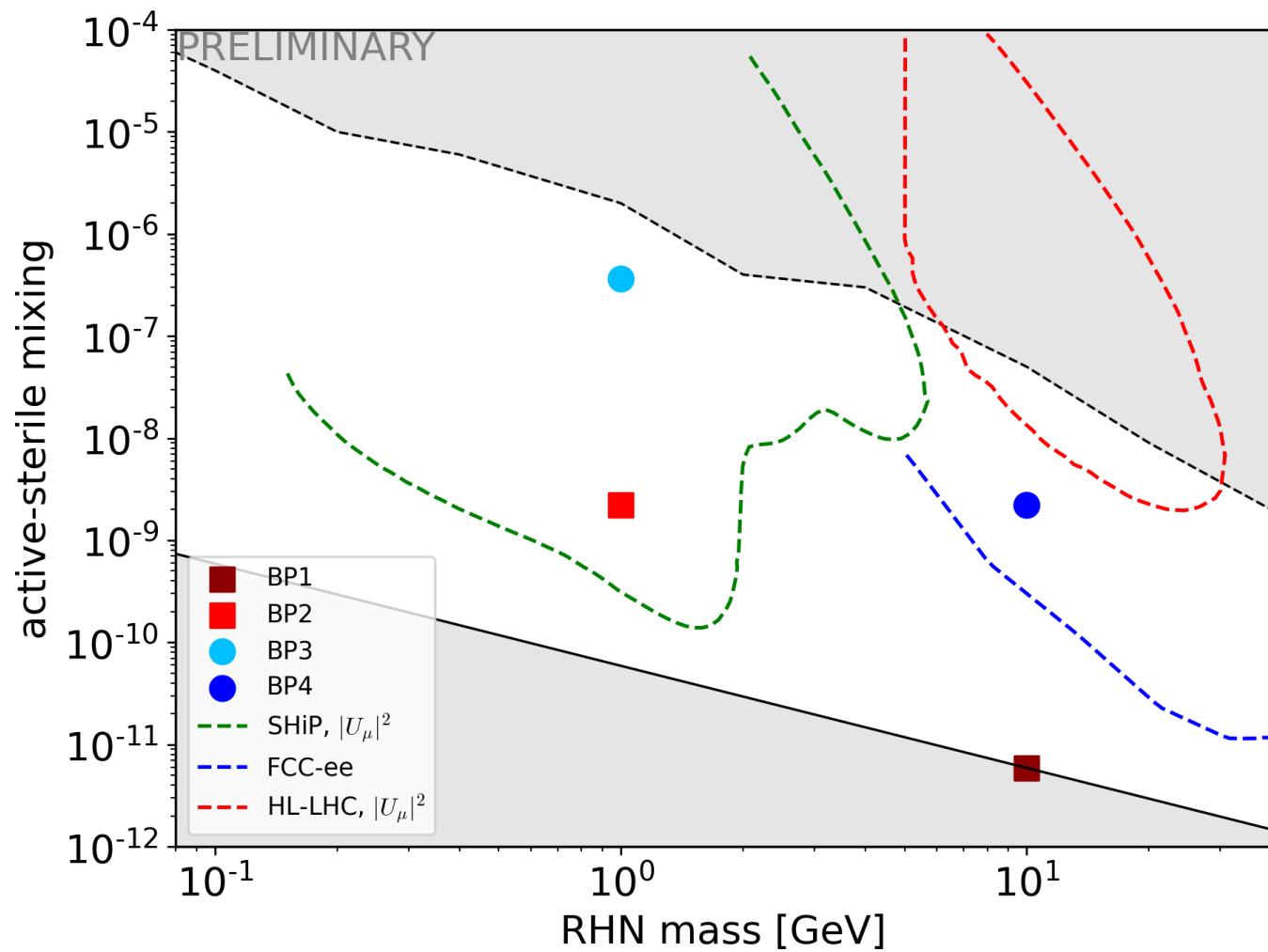


Available parameter space



[Fuyuto, Harz, SW in preparation]

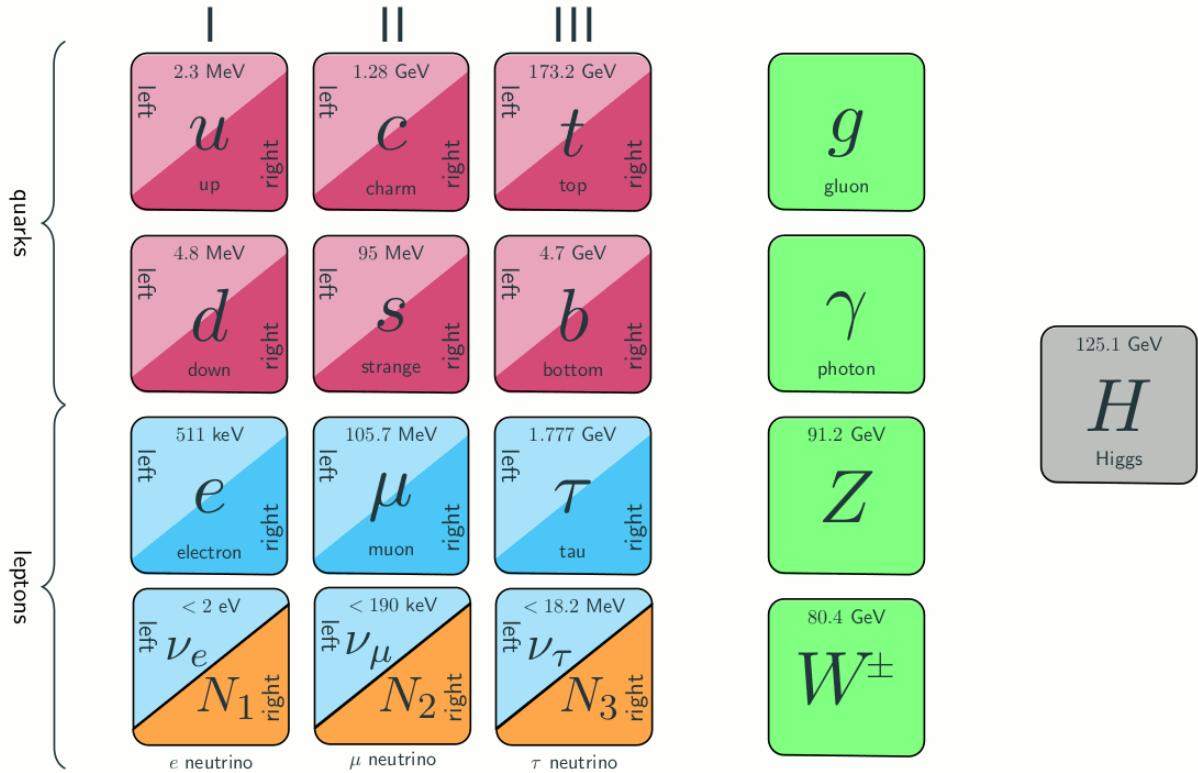
Available parameter space



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Non-Standard Case

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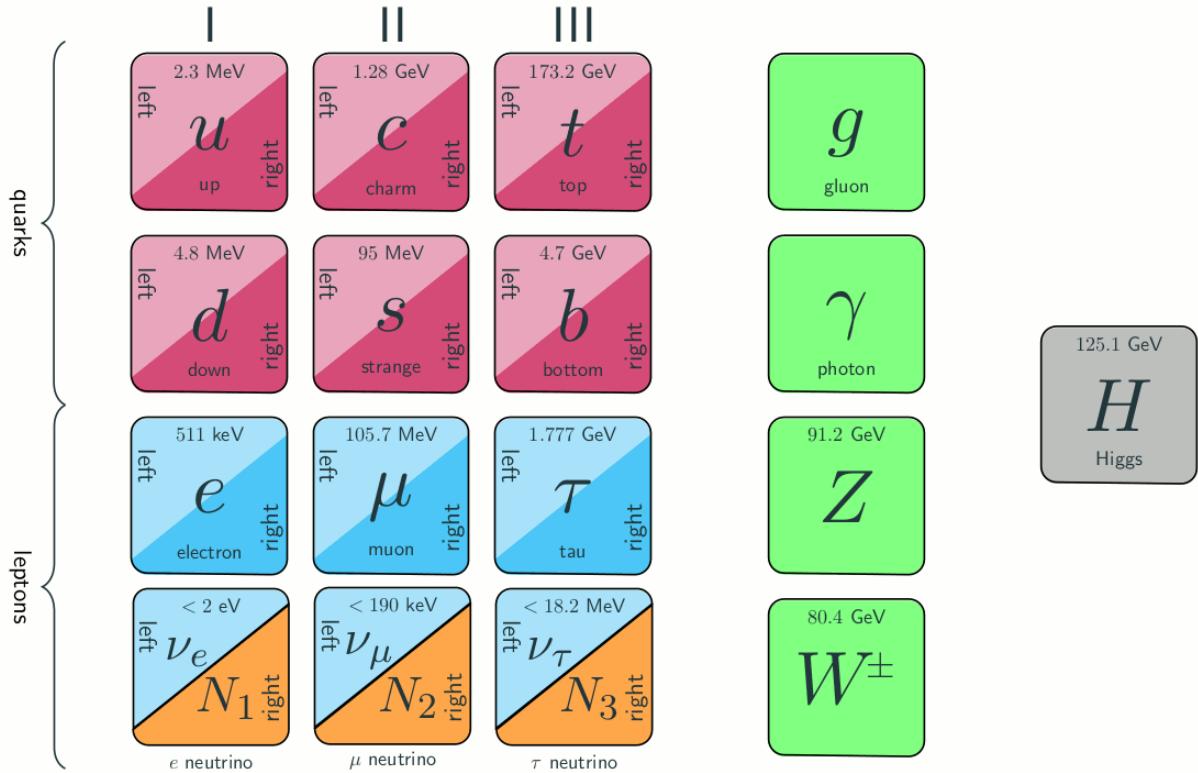


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$$\mathcal{L} = \mathcal{L}_{\text{SM}}$$

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Non-Standard Case



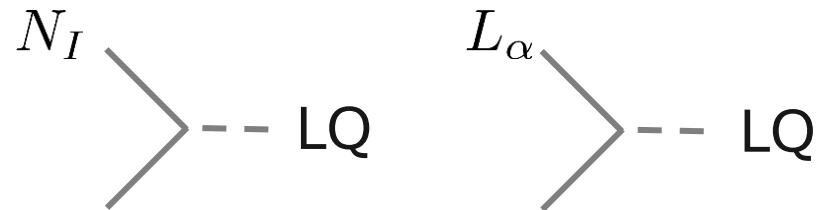
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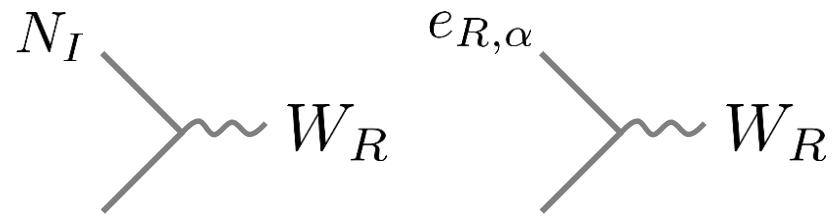
$$+ \mathcal{L}_N \left\{ \begin{array}{l} + \bar{N}(i\partial)N \\ - Y_{i\alpha} \bar{N}_i H L_\alpha + \text{h.c.} \\ - \bar{N}_i^c M_i N_i + \text{h.c.} \end{array} \right.$$

+ more?

Non-Standard Case



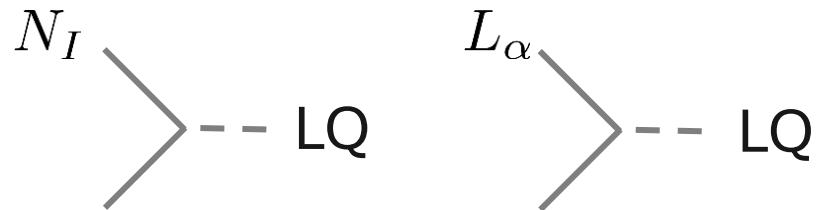
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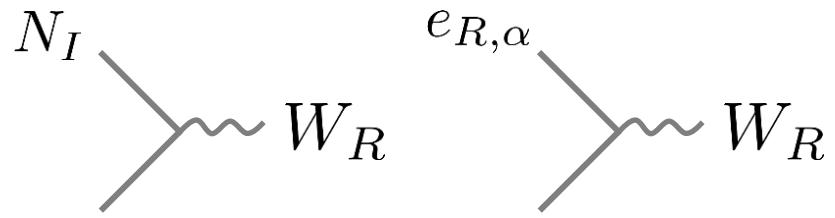
or

Any new particle coupling
to RHNs and/or leptons

Non-Standard Case

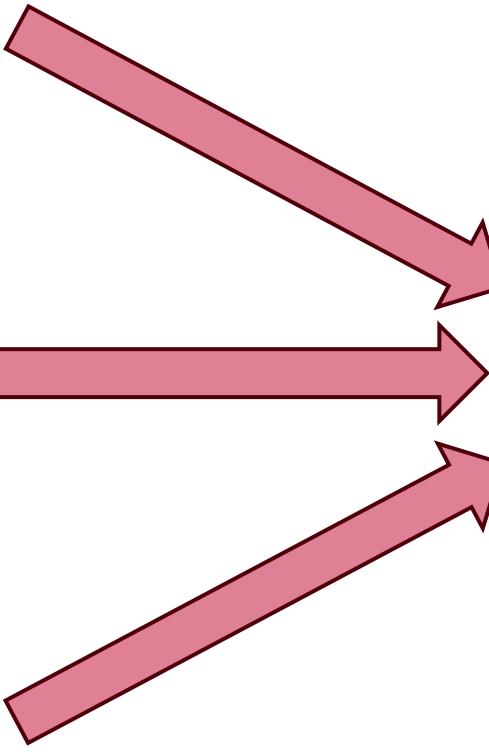


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or

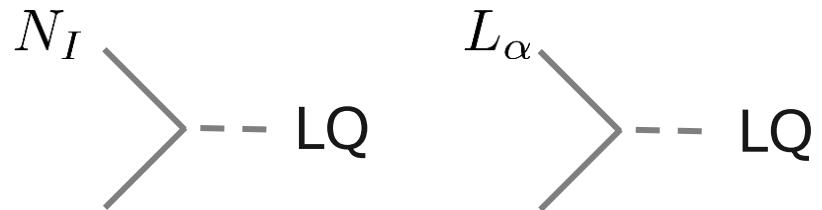
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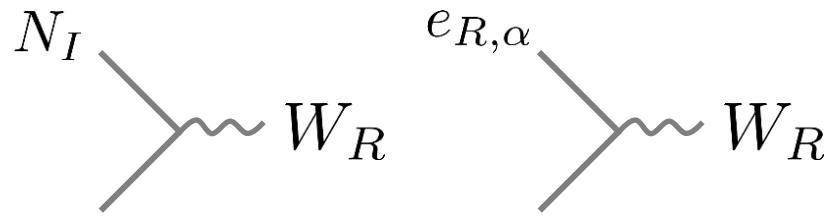
$$N_I \times \text{---} \sim \frac{1}{\Lambda^2}$$

Effective operator
description

Non-Standard Case

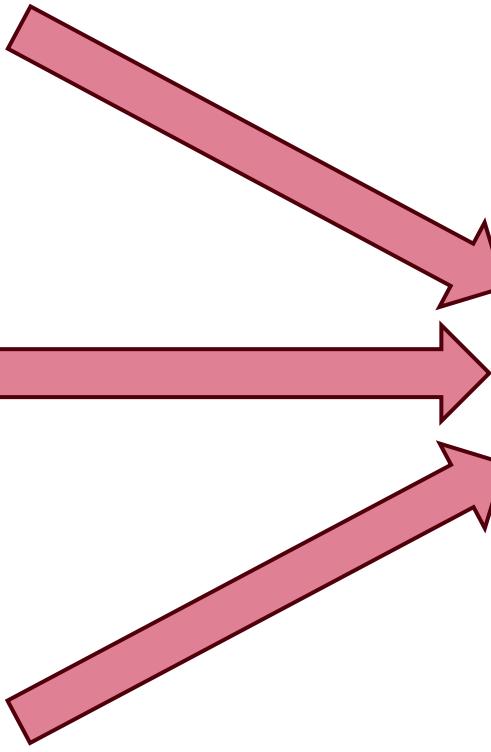


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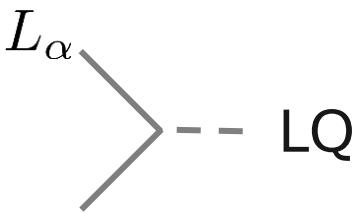
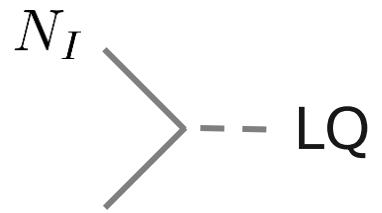
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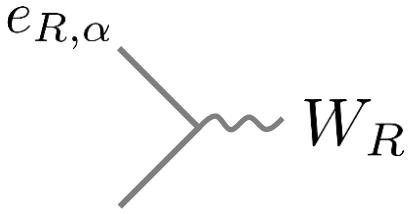
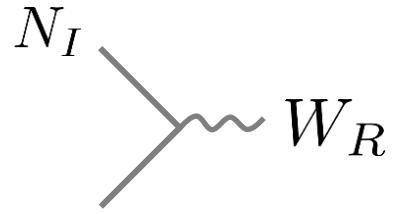
$$N_I \xrightarrow{\quad} \sim \frac{1}{\Lambda^2}$$

Effective operator
description
 ν SMEFT

Non-Standard Case

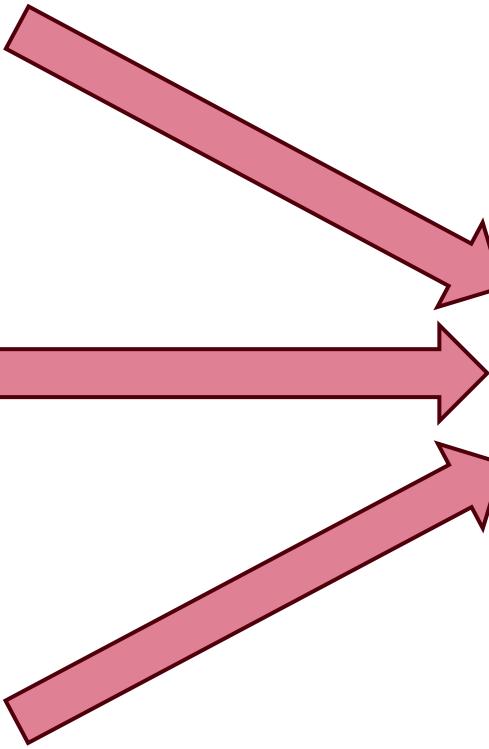


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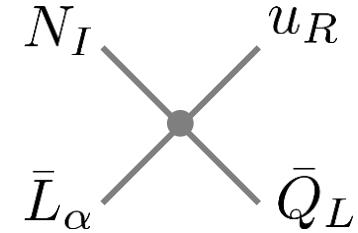


or

Any new particle coupling to RHNs and/or leptons



Example operator:



Feynman diagram of an effective operator: N_I (solid line) and \bar{L}_α (solid line) couple to u_R (solid line) and \bar{Q}_L (solid line) via a central vertex, with a scale factor $\sim \frac{1}{\Lambda^2}$.

Effective operator
description
 ν SMEFT

Lepton Number Violation

- Assignment of LN: $\mathcal{L} \supset -Y_{i\alpha} \overline{N_i} H L_\alpha - \overline{N_i^c} M_i N_i + \text{h.c.}$

$$L(L_\alpha) = 1$$

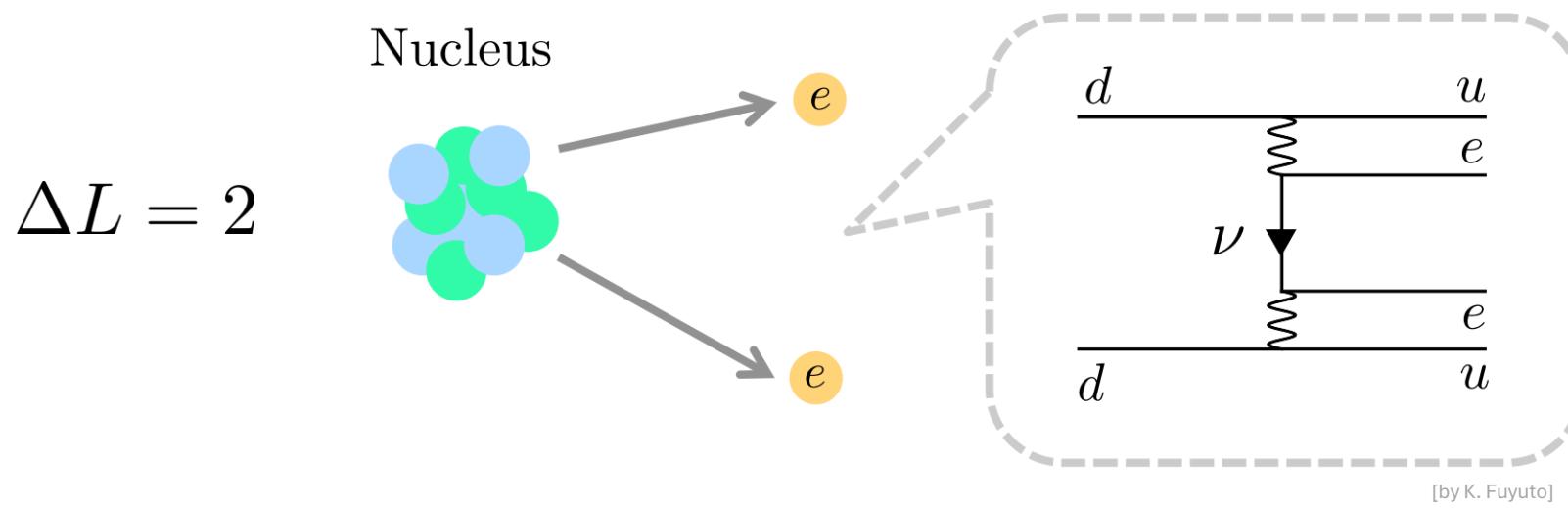
$$L(H) = 0$$

$$L(N_i) = 1$$

Lepton Number Violation

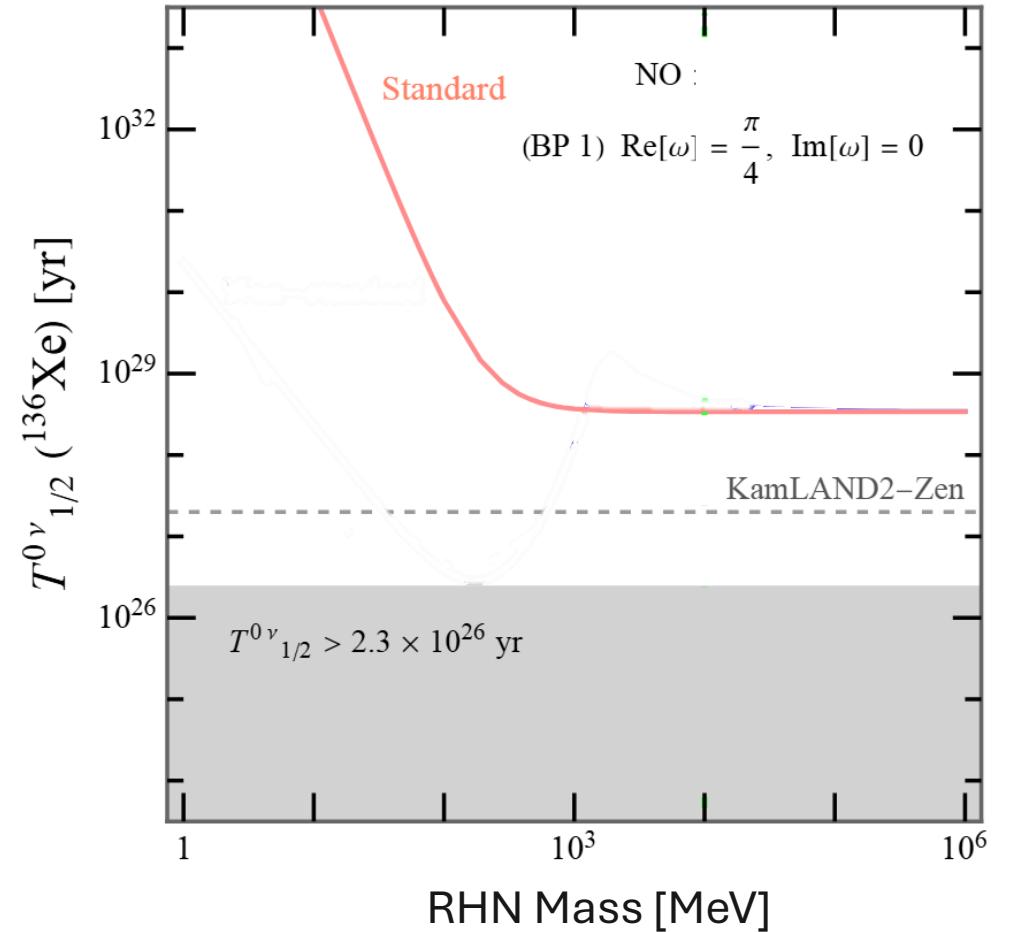
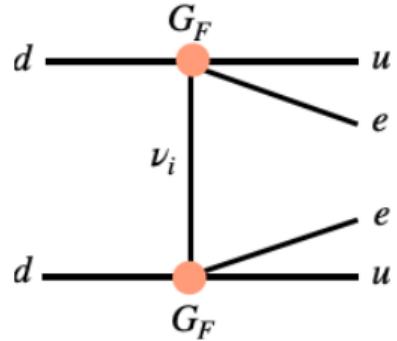
- Assignment of LN: $\mathcal{L} \supset -Y_{i\alpha}\overline{N_i}HL_\alpha - \overline{N_i^c}M_iN_i + \text{h.c.}$
LNC LNV
- “Most” promising observable: $0\nu\beta\beta$ decay

$$\begin{aligned} L(L_\alpha) &= 1 \\ L(H) &= 0 \\ L(N_i) &= 1 \end{aligned}$$



$0\nu\beta\beta$ decay

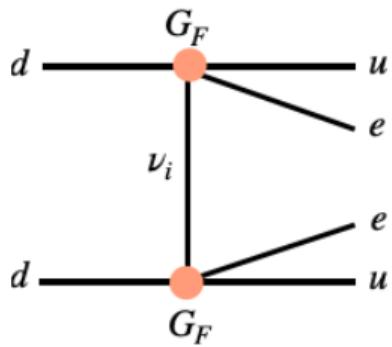
- 4-fermion interaction at low scales



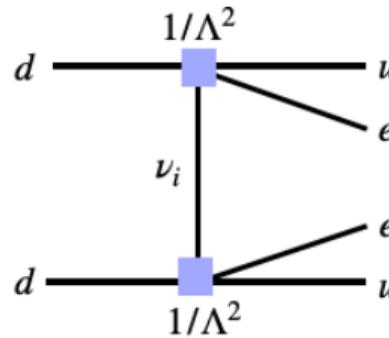
[Fuyuto, Harz, SW in preparation]

$0\nu\beta\beta$ decay

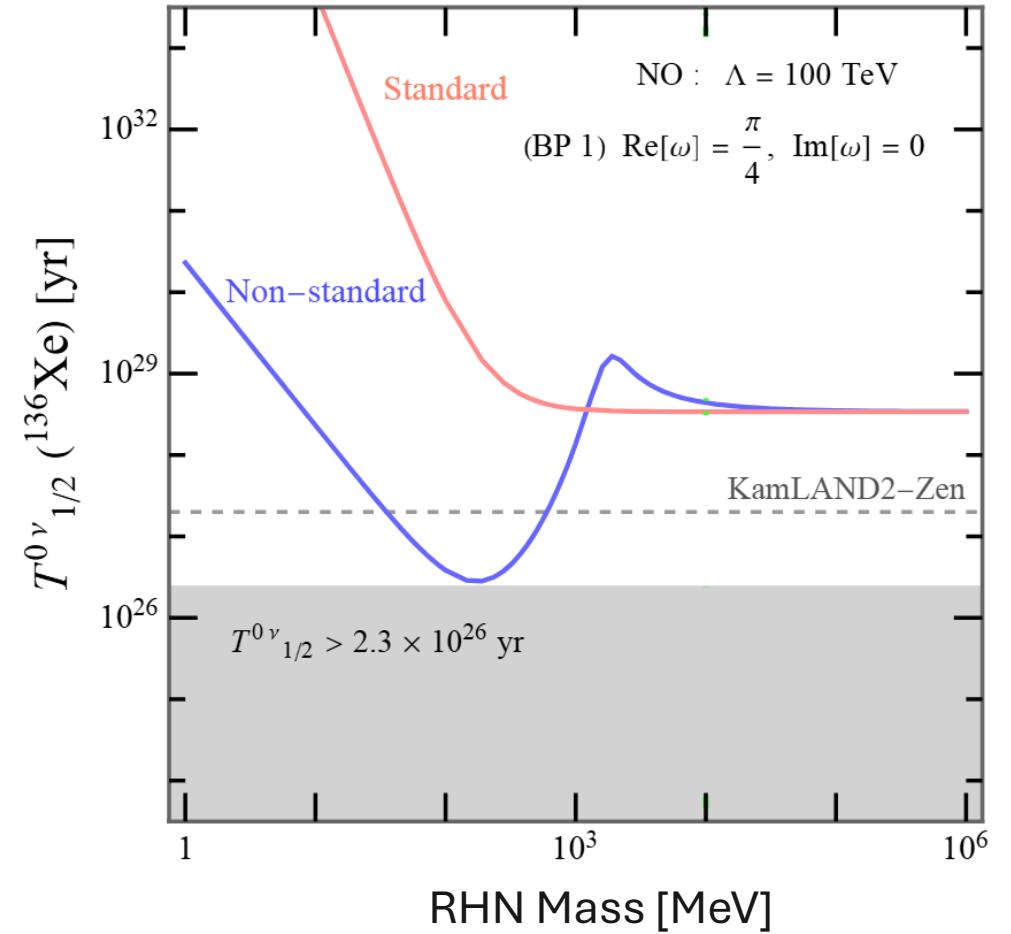
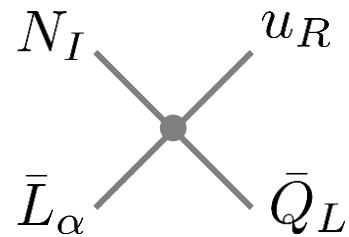
- See also [Dekens et. al. JHEP 2020]



vs.



LNC operator:

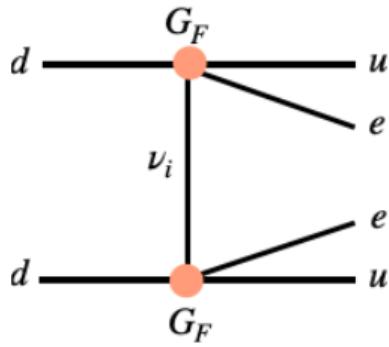


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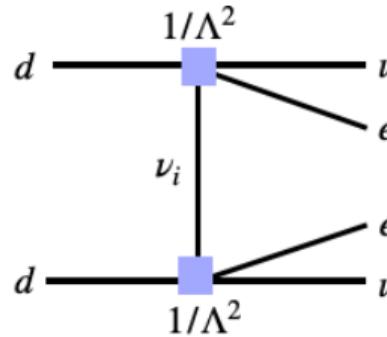
$0\nu\beta\beta$ decay

Order of magnitude effect!

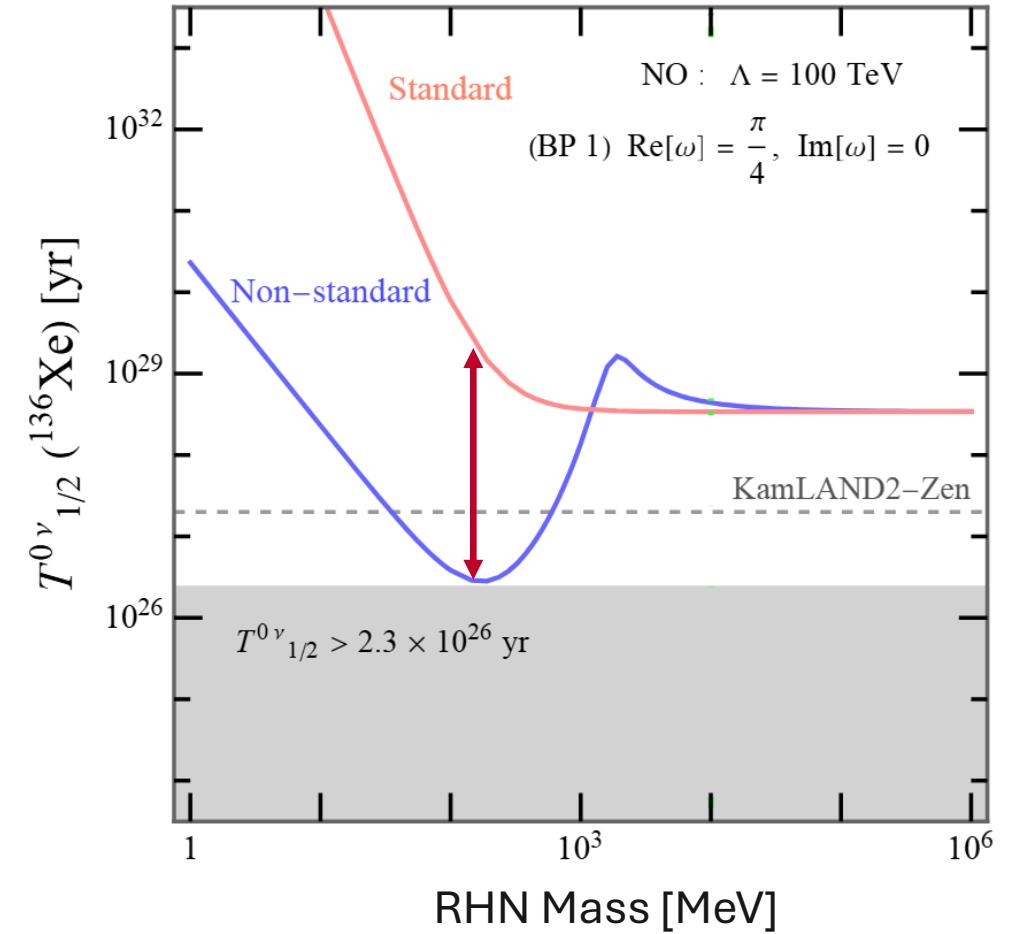
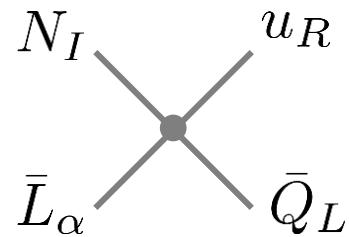
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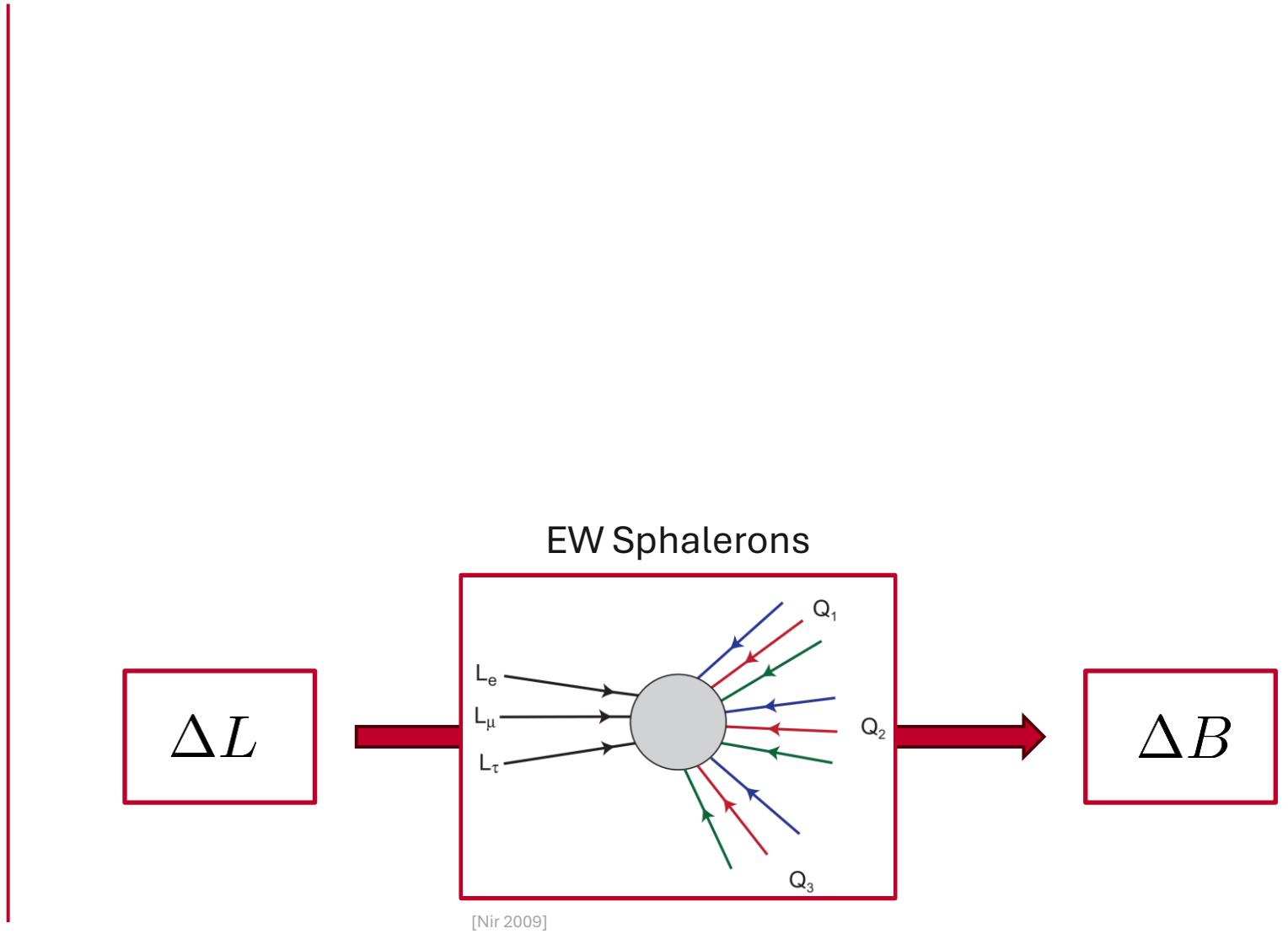


LNC operator:



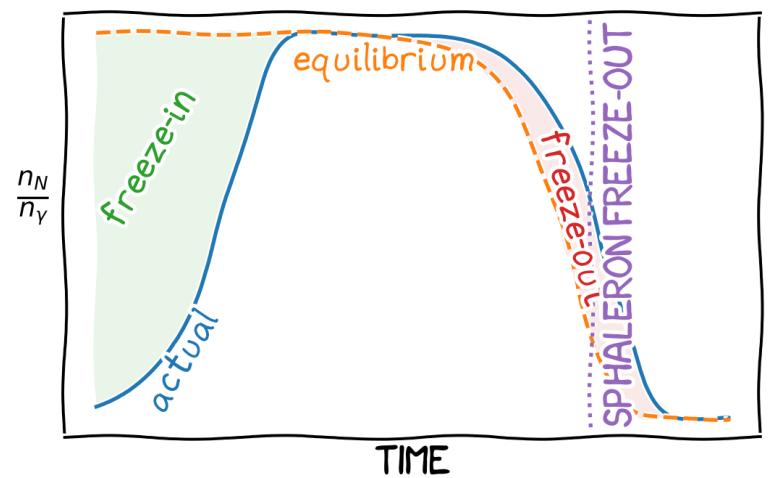
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Baryon Asymmetry

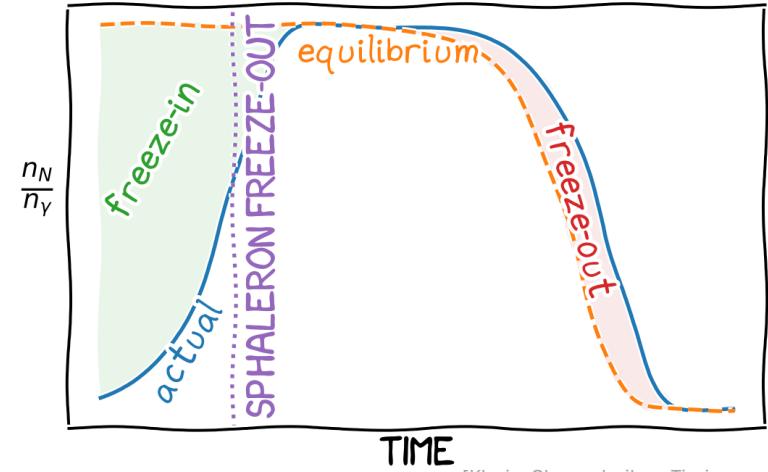


Baryon Asymmetry

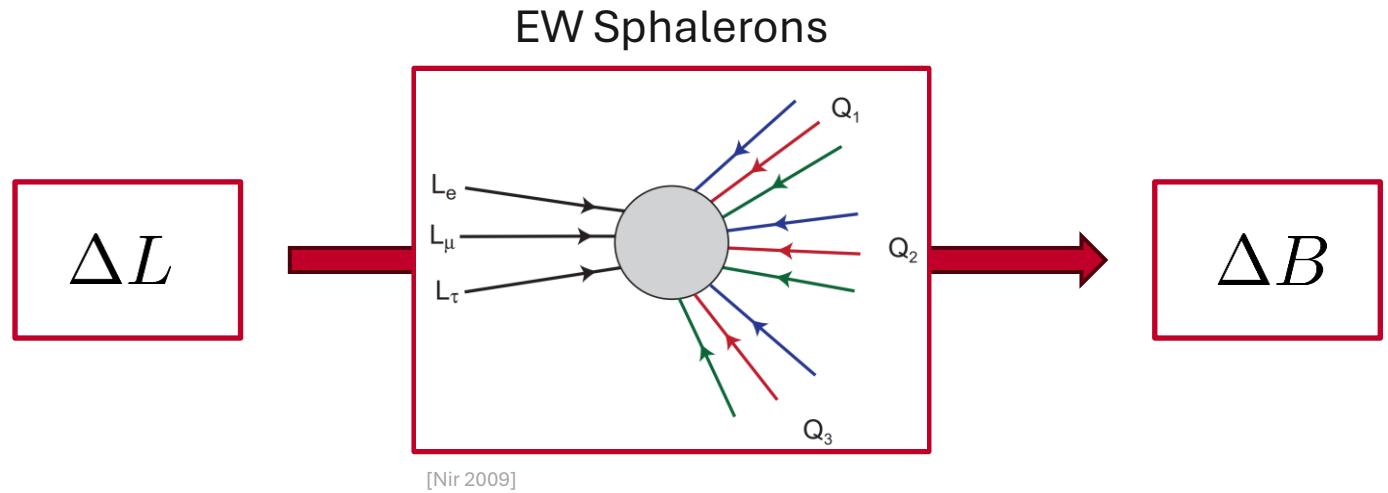
Thermal/Resonant



Via Oscillations

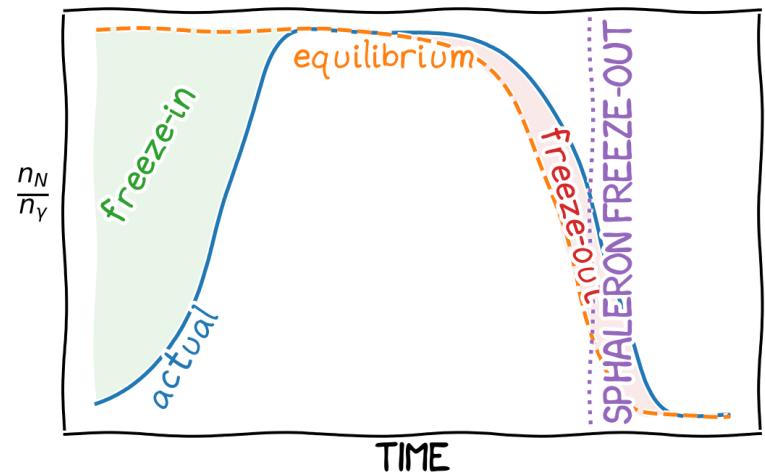


[Klaric, Shaposhnikov, Timiryasov 2103.16545]

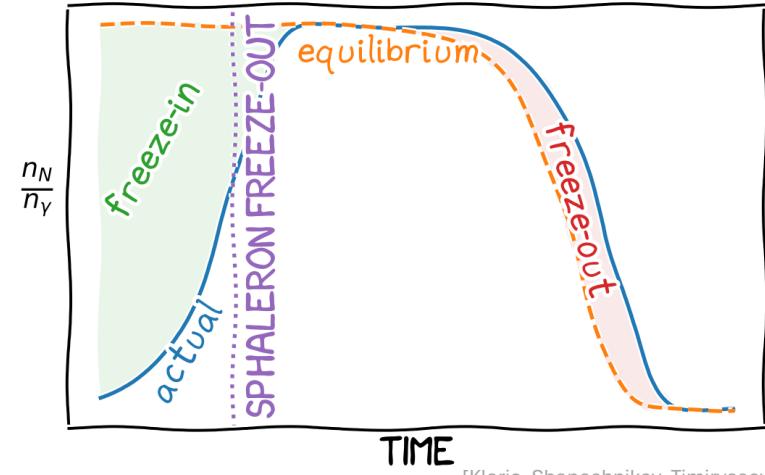


Baryon Asymmetry

Thermal/Resonant

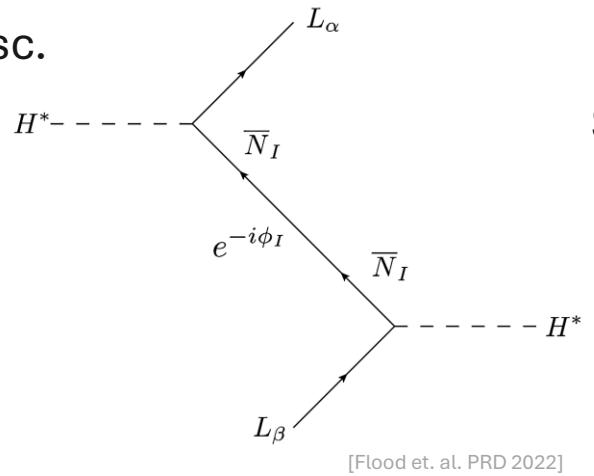


Via Oscillations



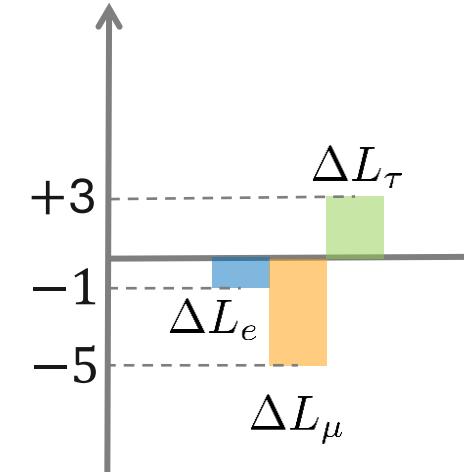
[Klaric, Shaposhnikov, Timiryasov 2103.16545]

Osc.

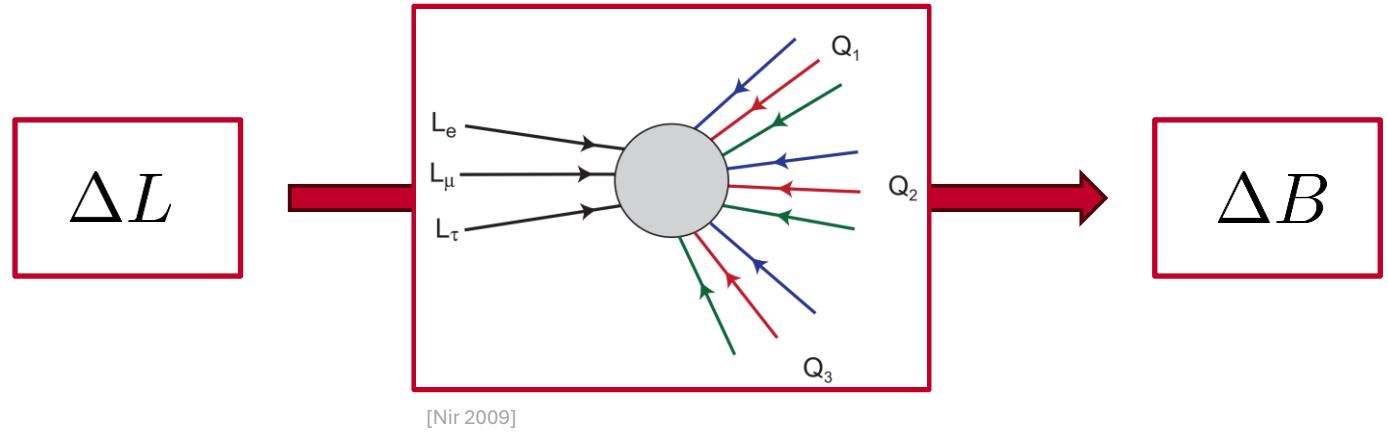


Coherent oscillations

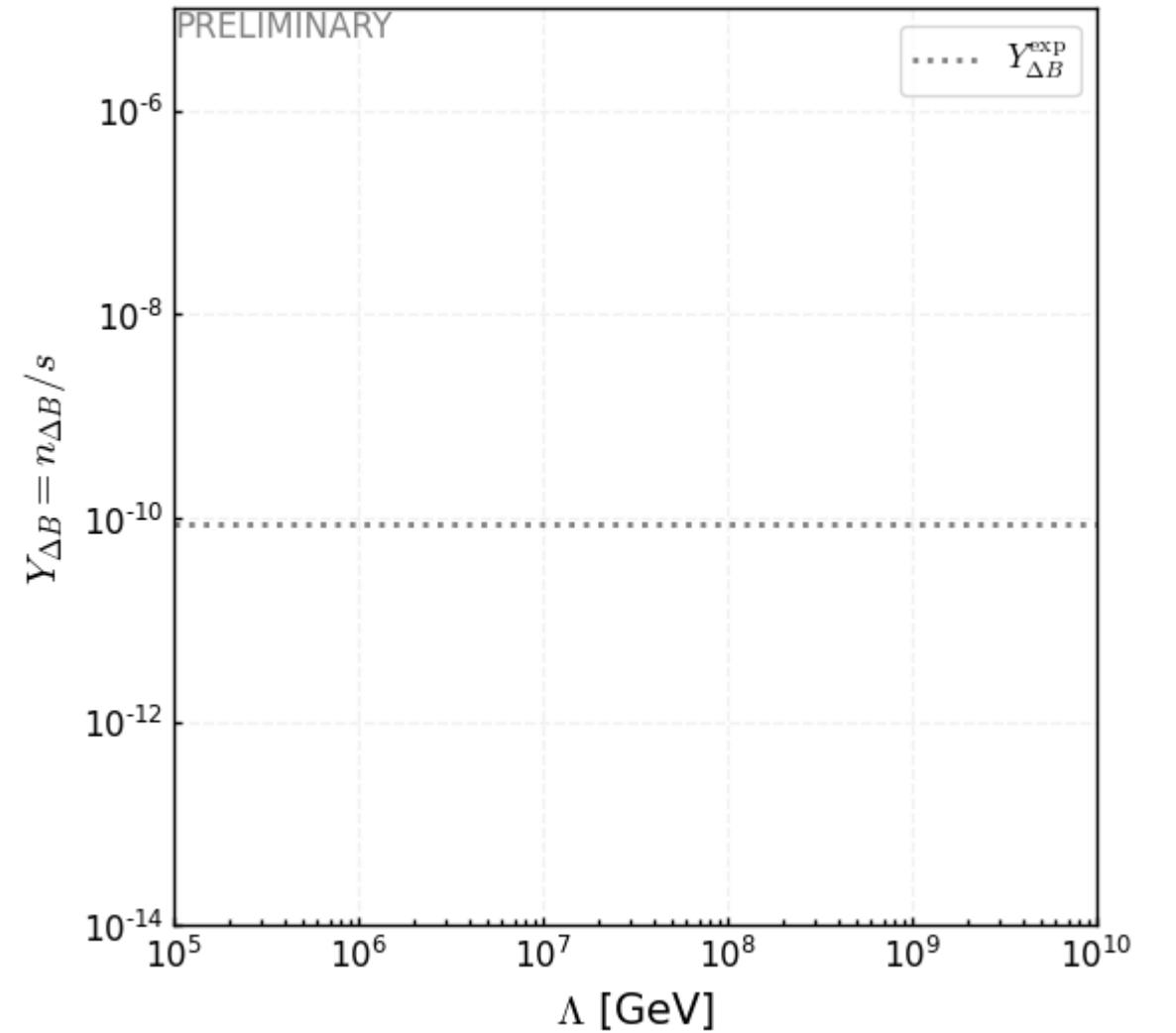
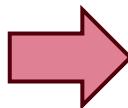
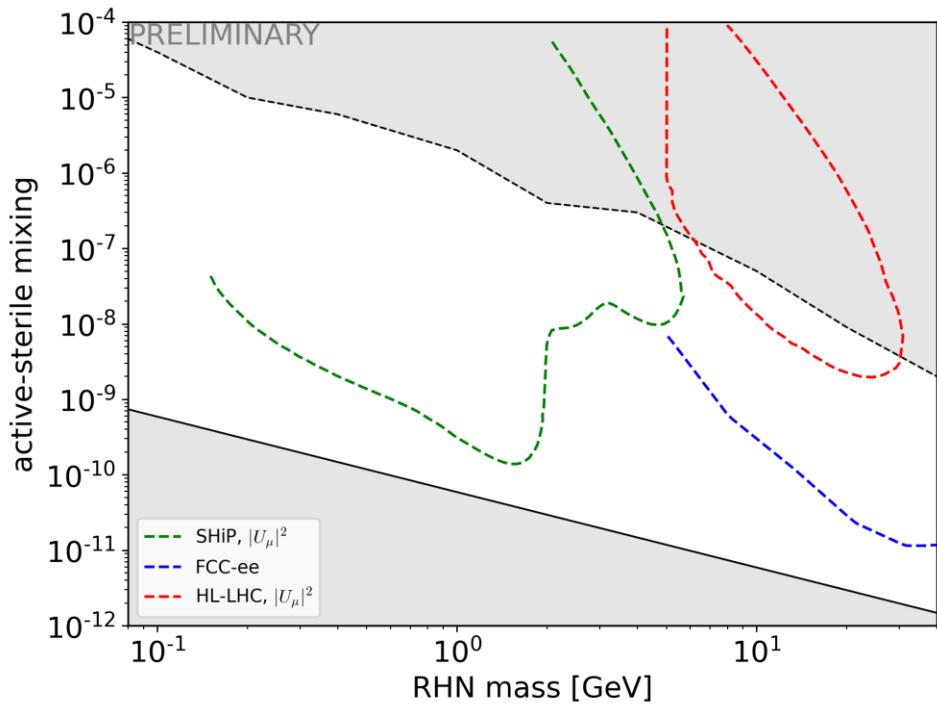
Scatterings



EW Sphalerons

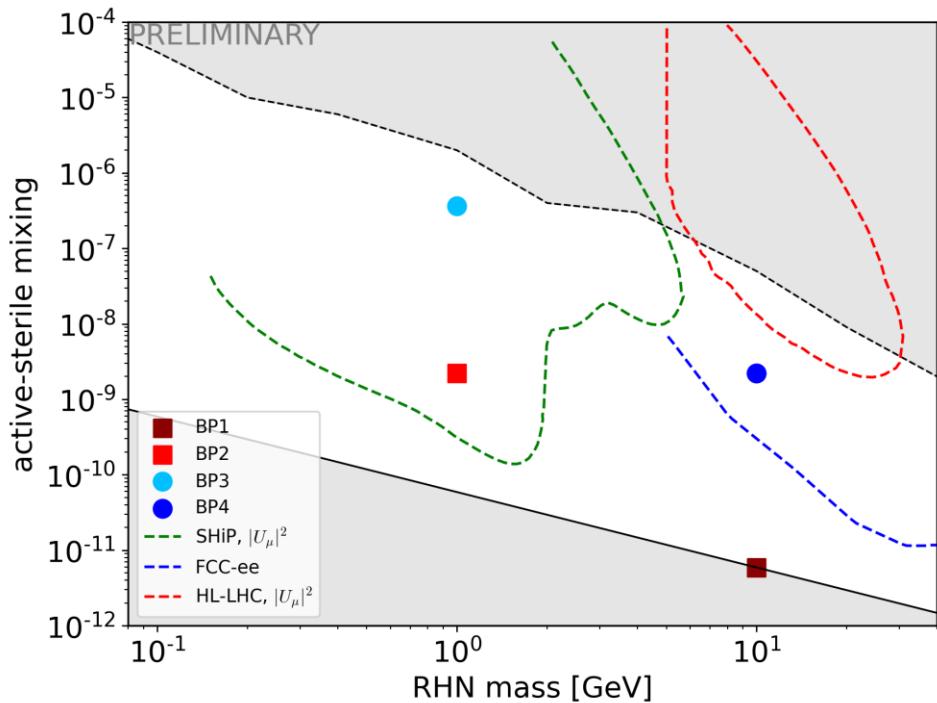


Low-Scale Leptogenesis in the vSMEFT

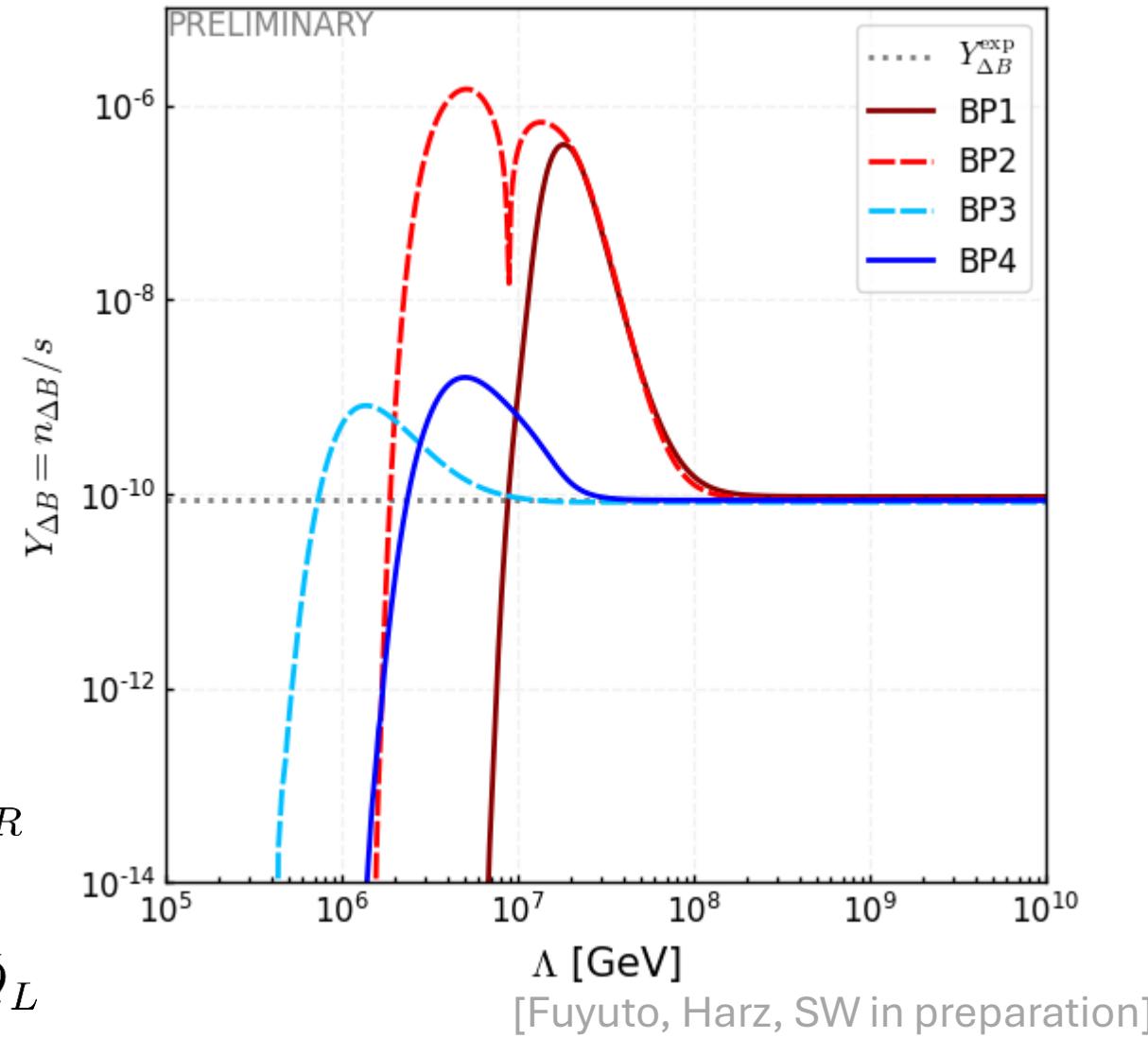
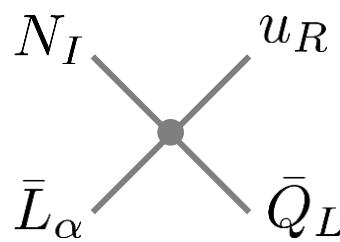


[Fuyuto, Harz, SW in preparation]

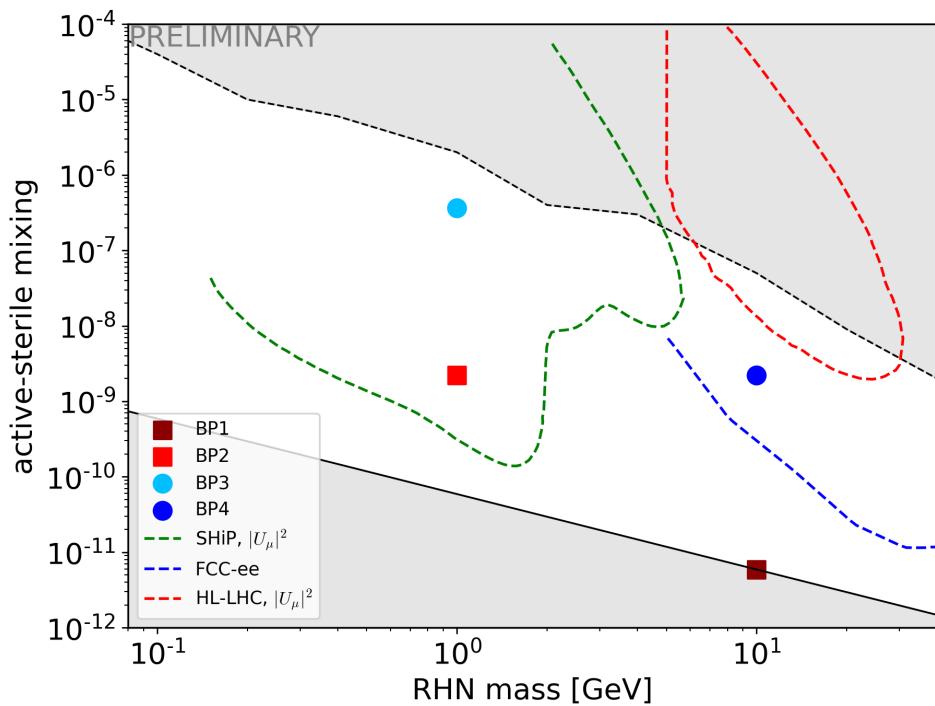
Low-Scale Leptogenesis in the vSMEFT



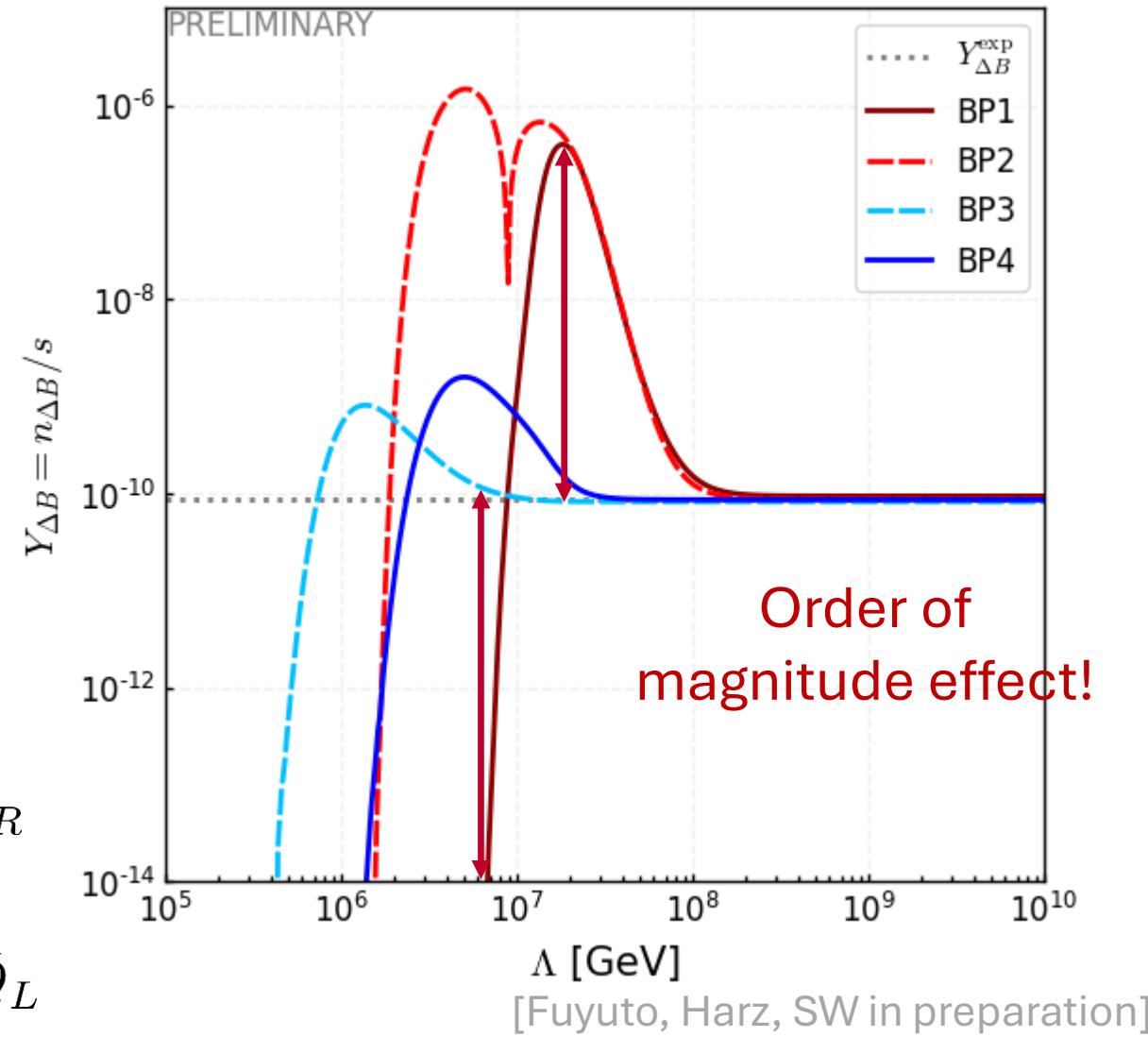
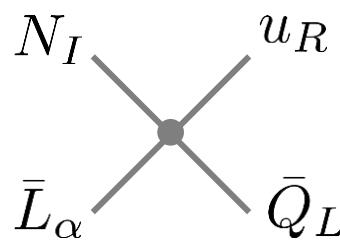
LNC operator:



Low-Scale Leptogenesis in the vSMEFT

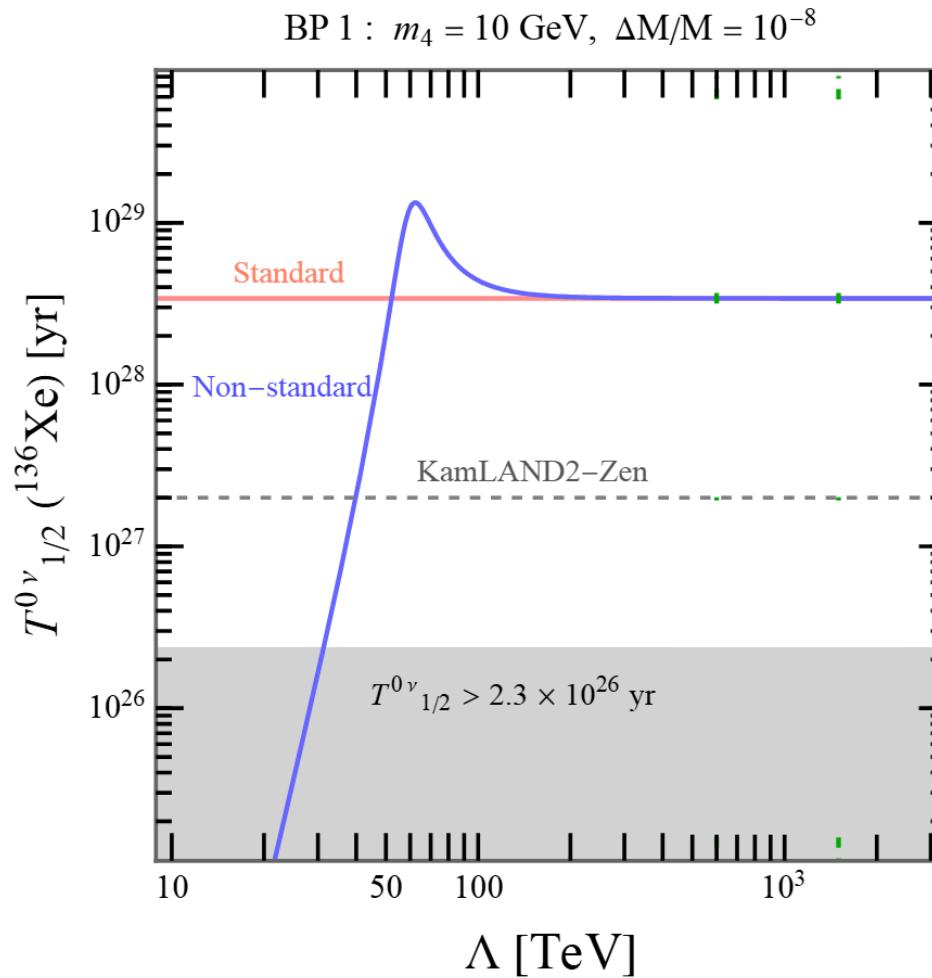


LNC operator:



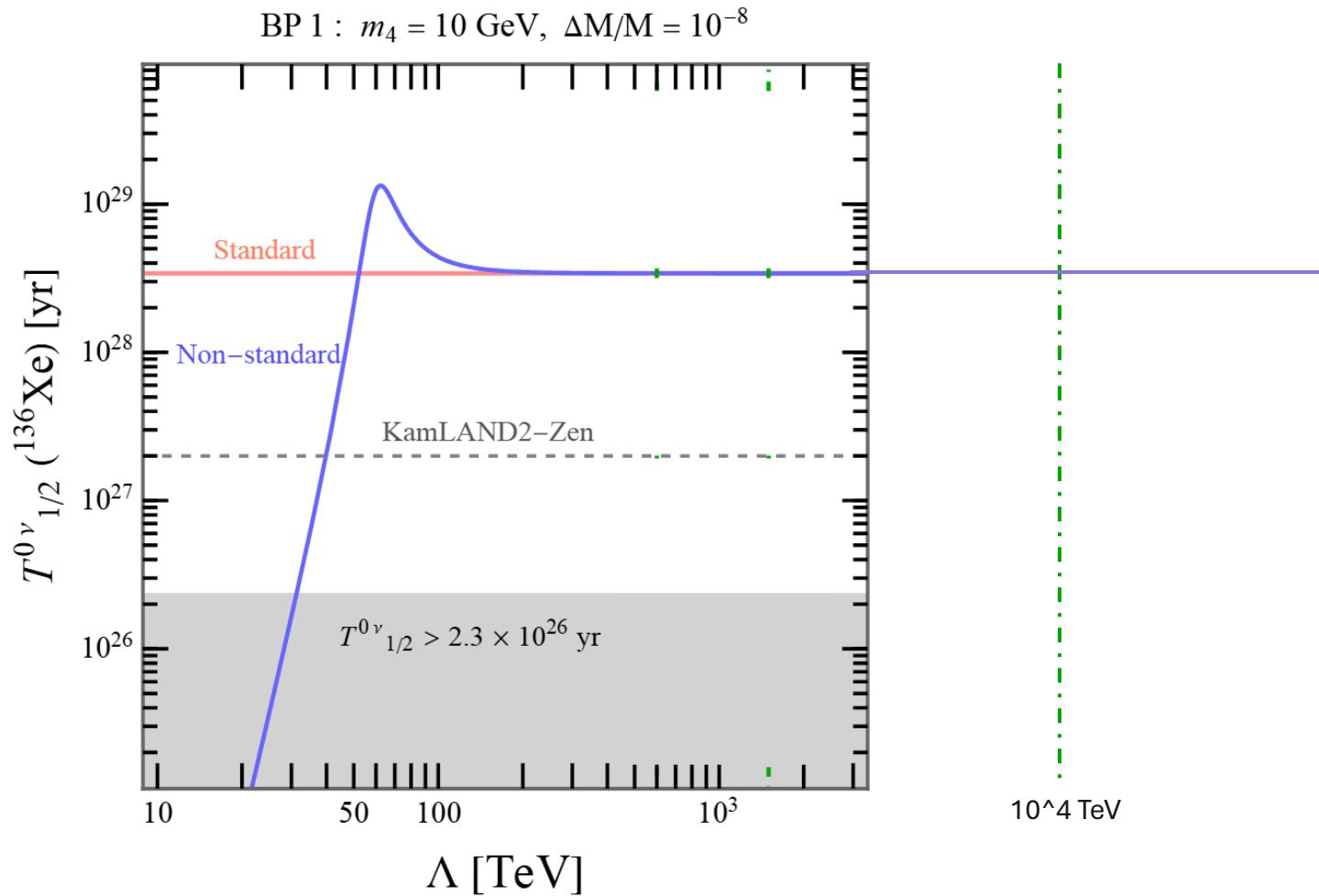
Combine Leptogenesis & $0\nu\beta\beta$ decay

Falsifying Low-Scale Leptogenesis?



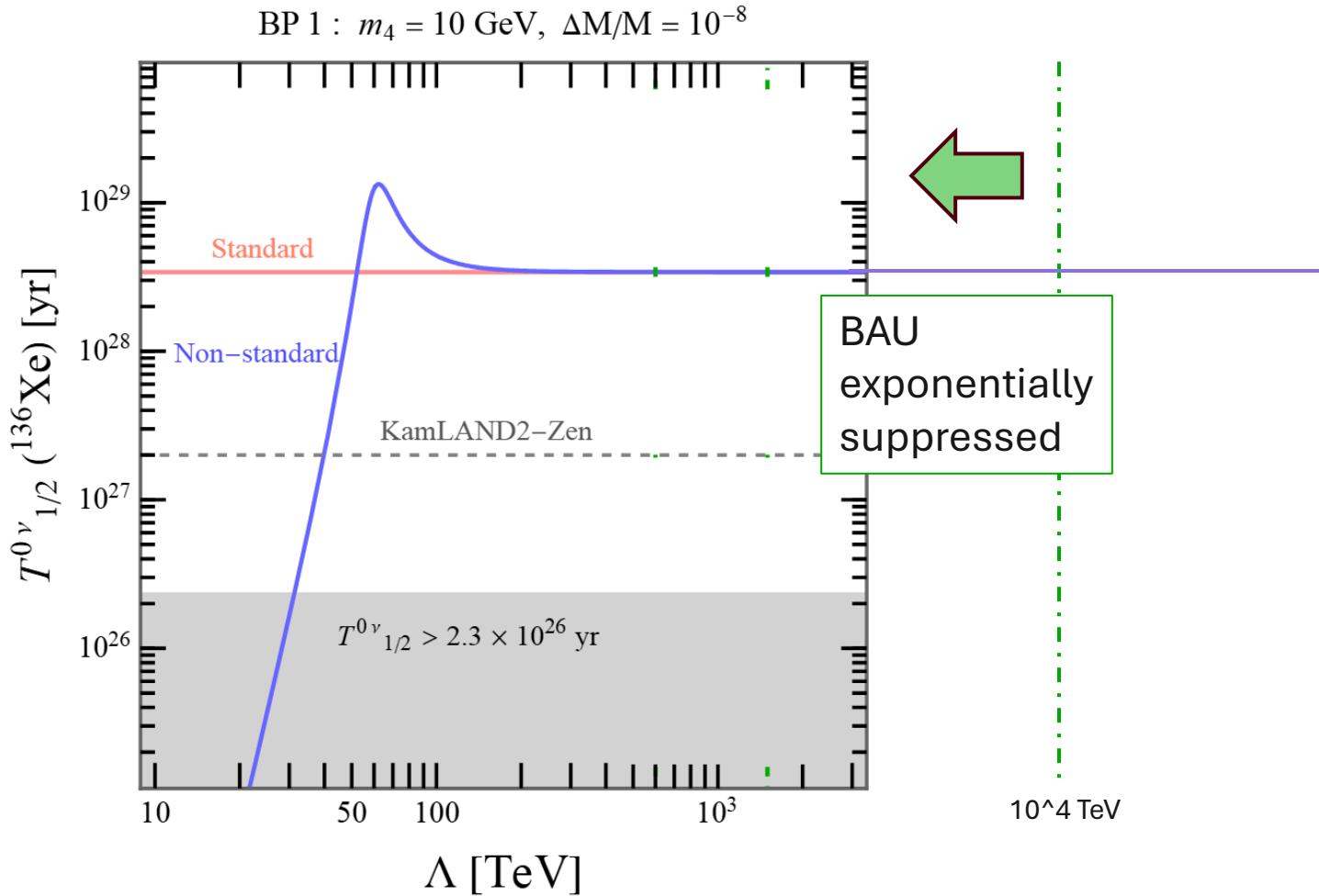
[Fuyuto, Harz, SW in preparation]

Falsifying Low-Scale Leptogenesis?



[Fuyuto, Harz, SW in preparation]

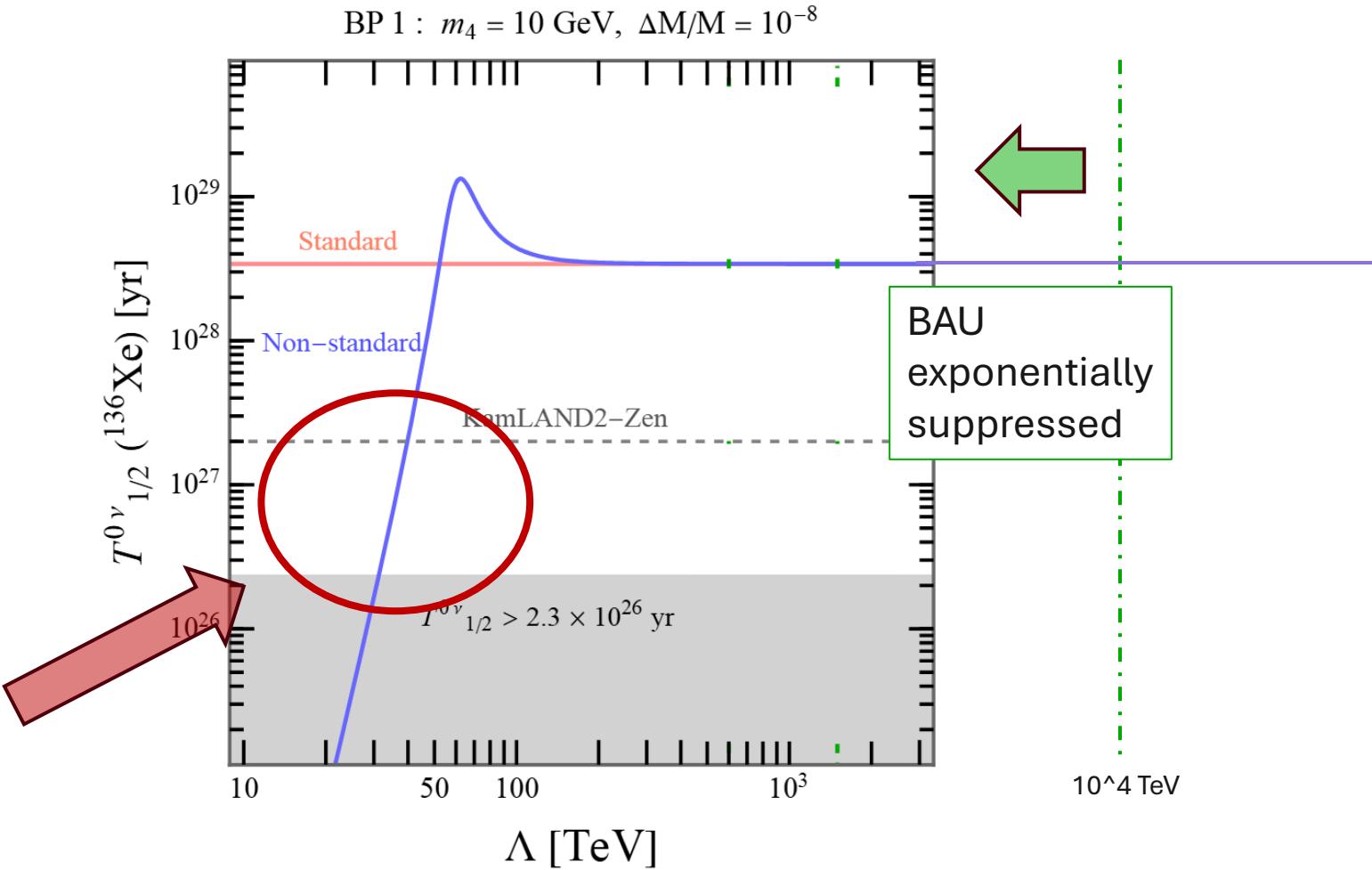
Falsifying Low-Scale Leptogenesis?



[Fuyuto, Harz, SW in preparation]

Falsifying Low-Scale Leptogenesis?

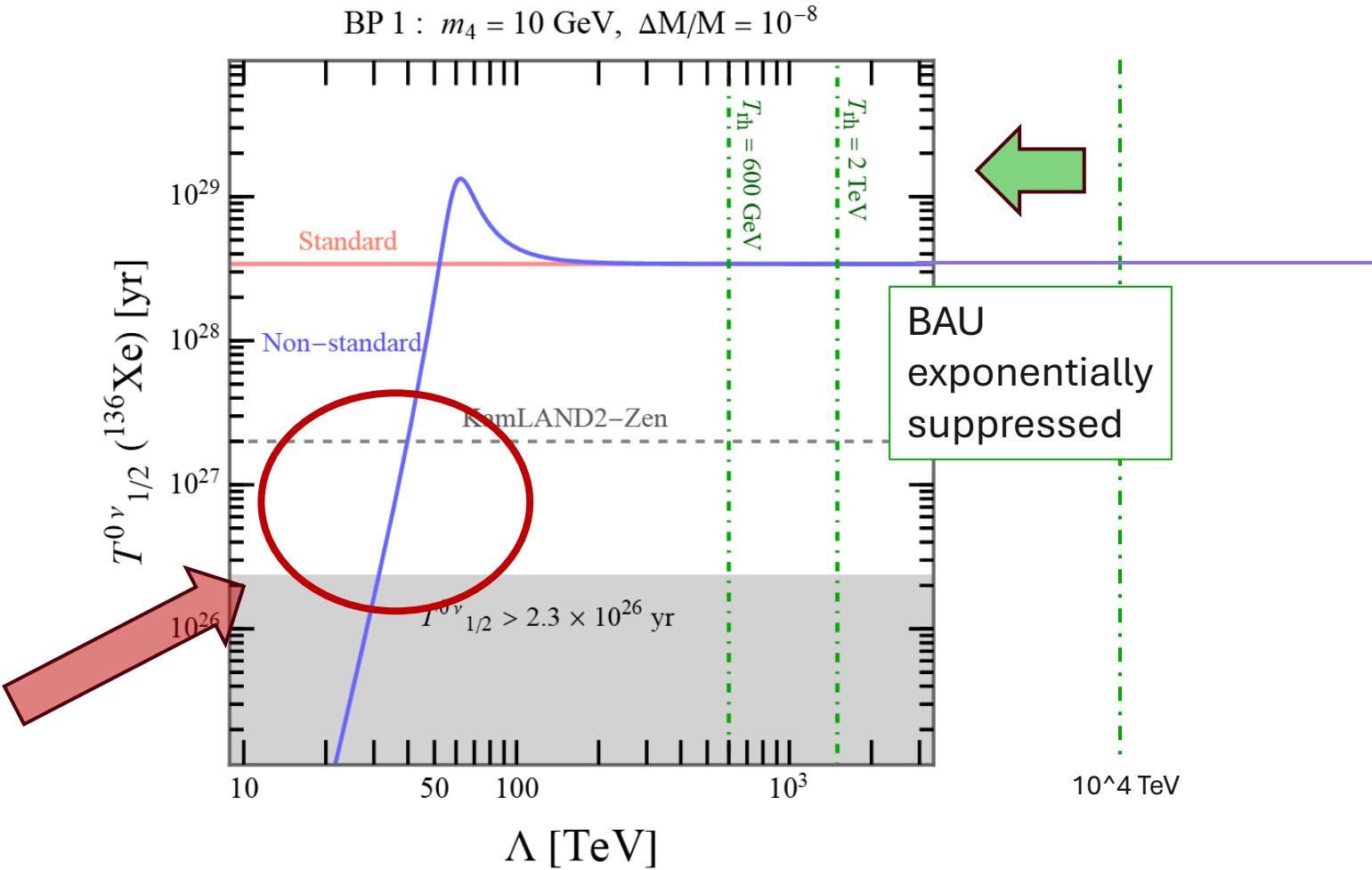
Possible
future
detection



[Fuyuto, Harz, SW in preparation]

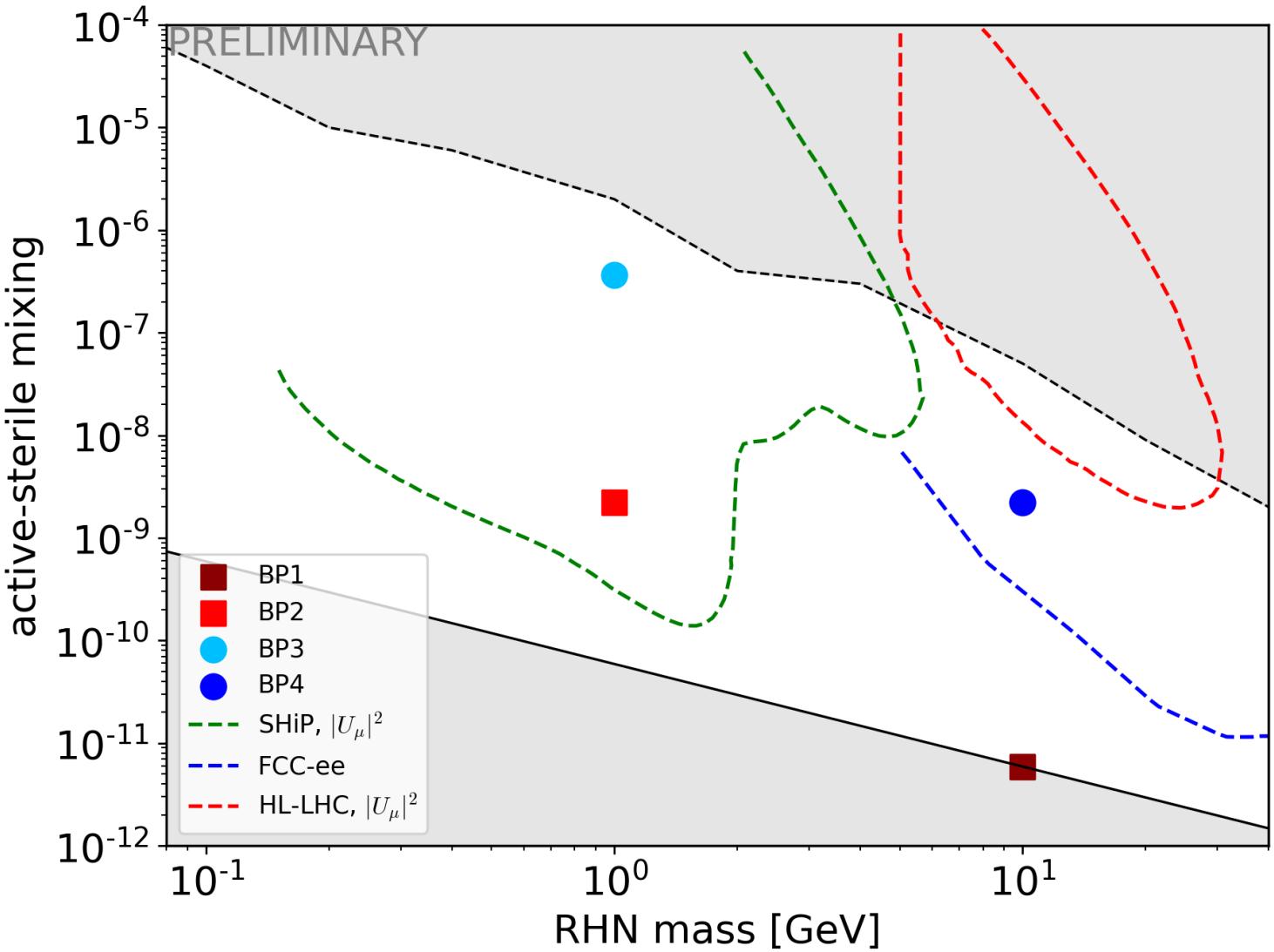
Falsifying Low-Scale Leptogenesis?

Possible
future
detection

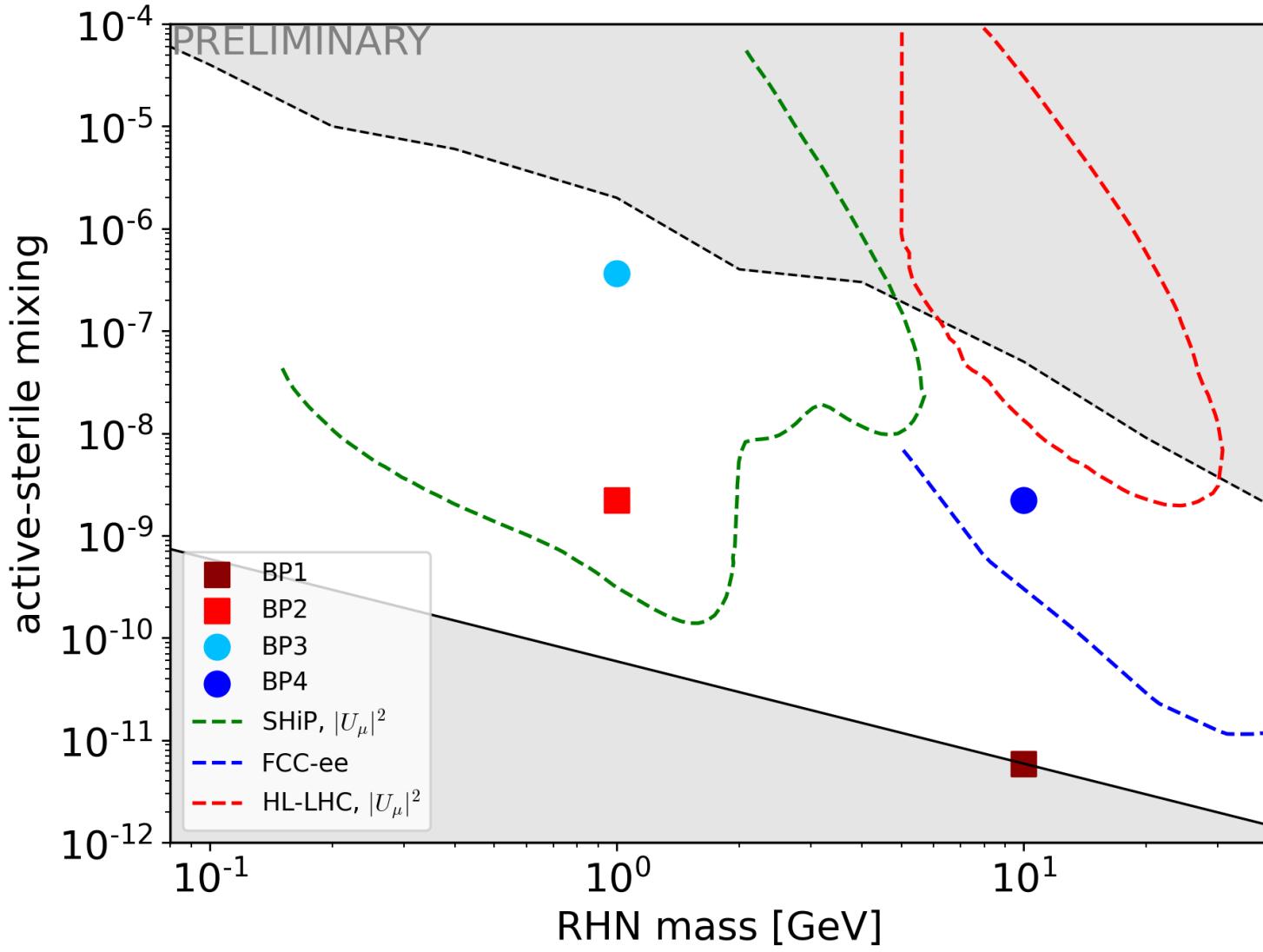


[Fuyuto, Harz, SW in preparation]

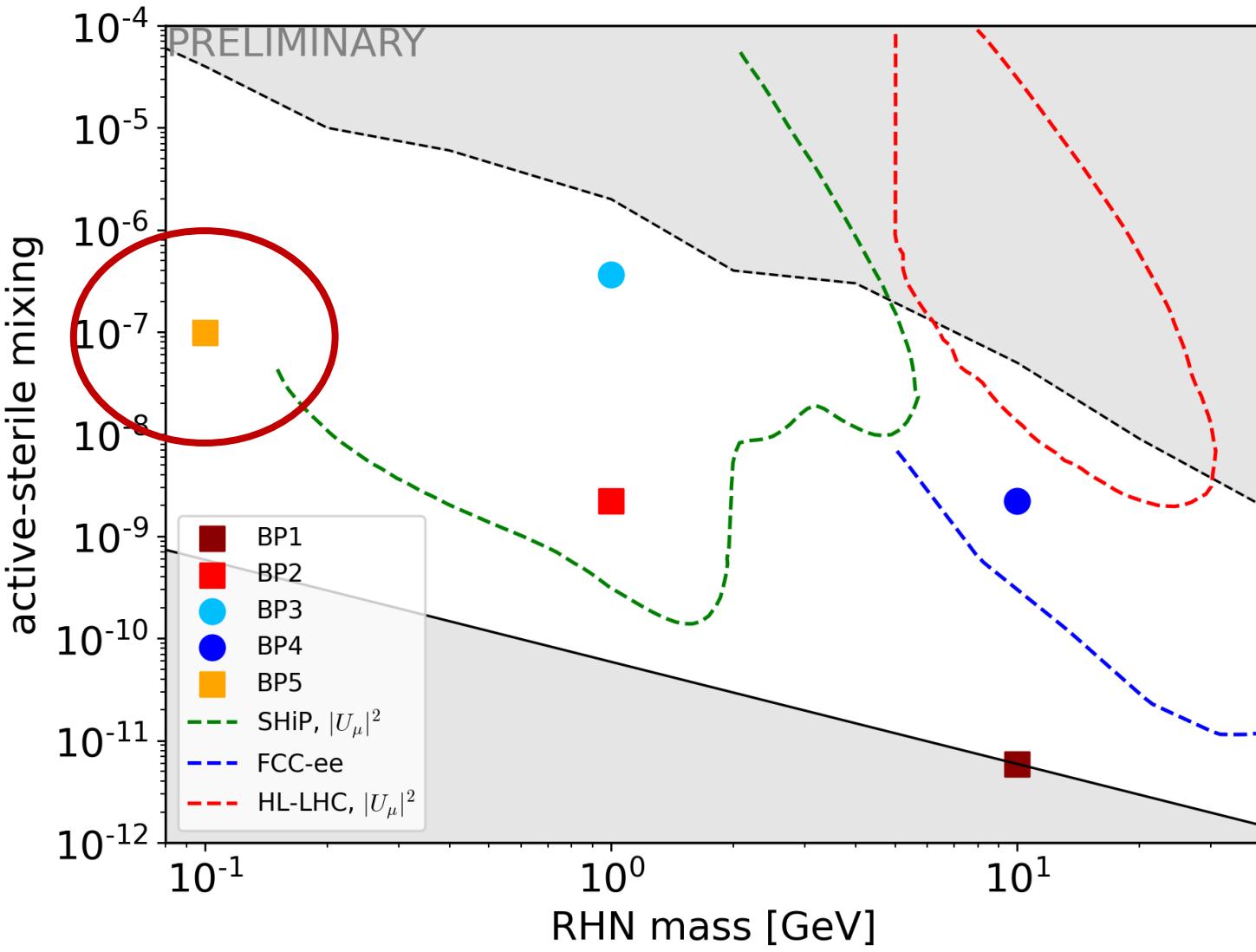
Always True?



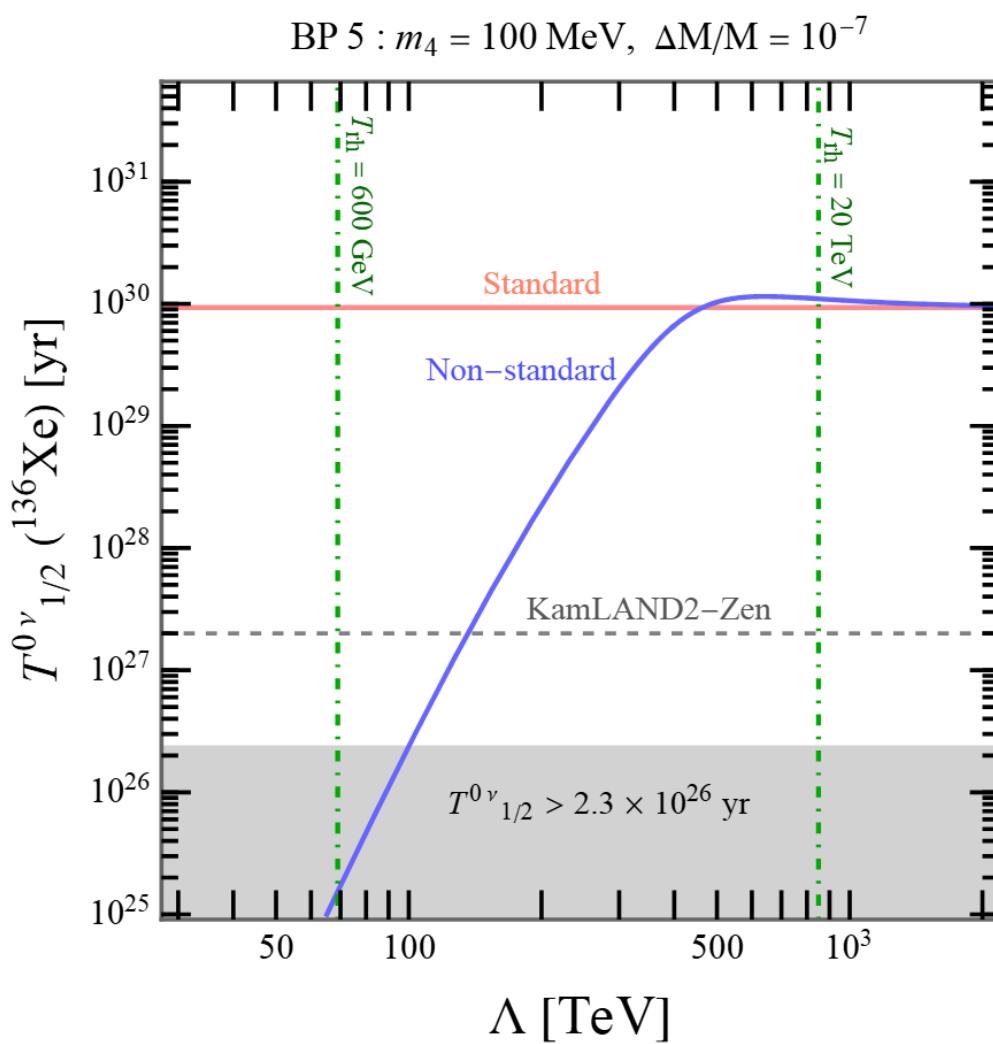
Always True?



Always True?



Falsifying Low-Scale Leptogenesis? – No

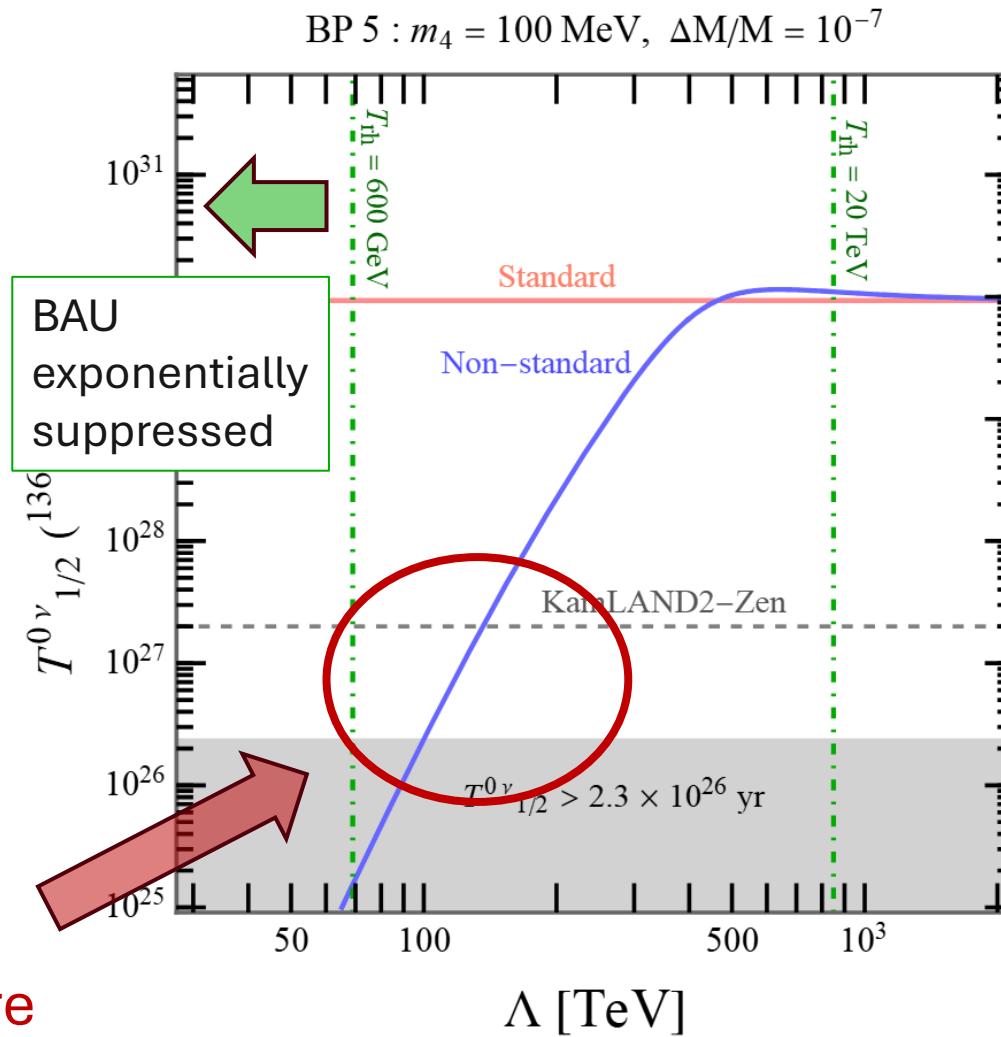


But still exciting!

[Fuyuto, Harz, SW in preparation]

Falsifying Low-Scale Leptogenesis? – No

Need low
reheating
temperature

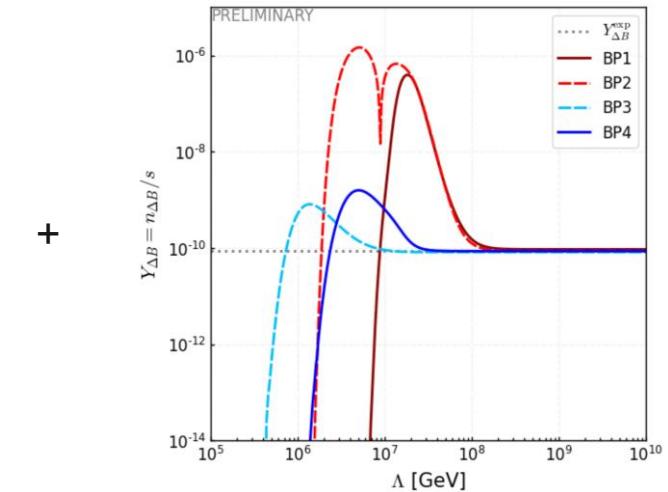
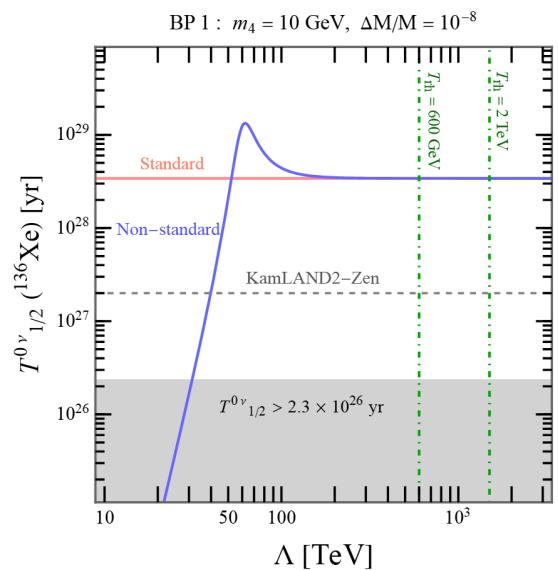
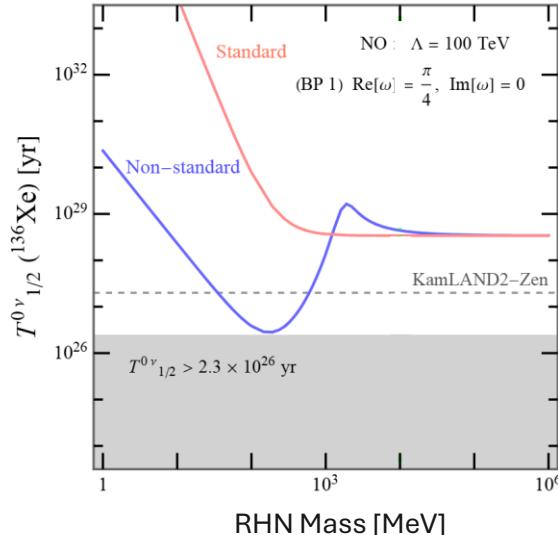
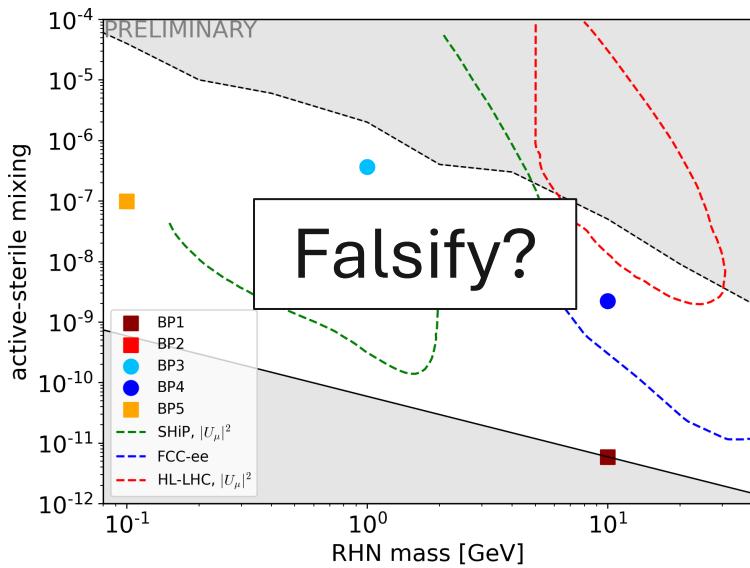
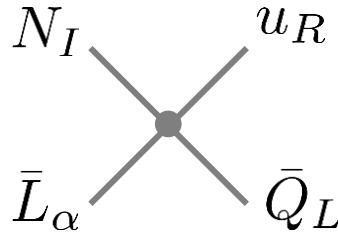


But still
exciting!

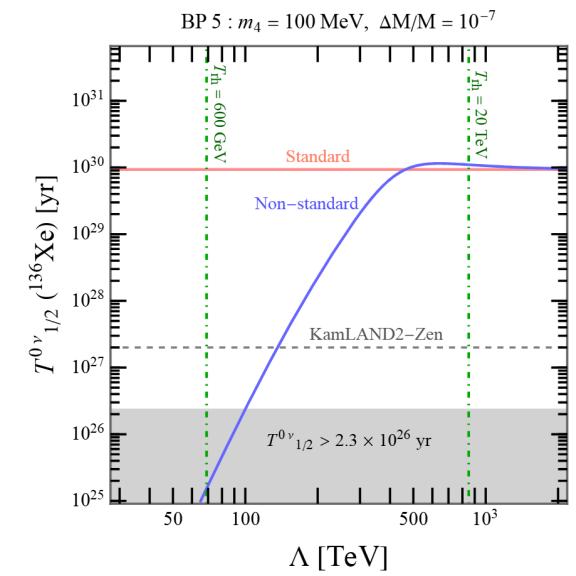
[Fuyuto, Harz, SW in preparation]

Conclusion

Example operator:



BUT





Thank You

Backup

Key predictions

1

Neutrino masses – Seesaw mechanism

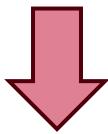
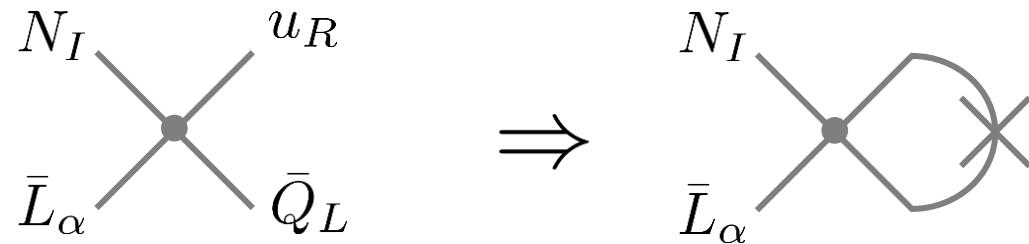
2

Lepton number violation – $0\nu\beta\beta$ decay

3

Baryon Asymmetry of the Universe - Leptogenesis

1) Contribution to neutrino mass?



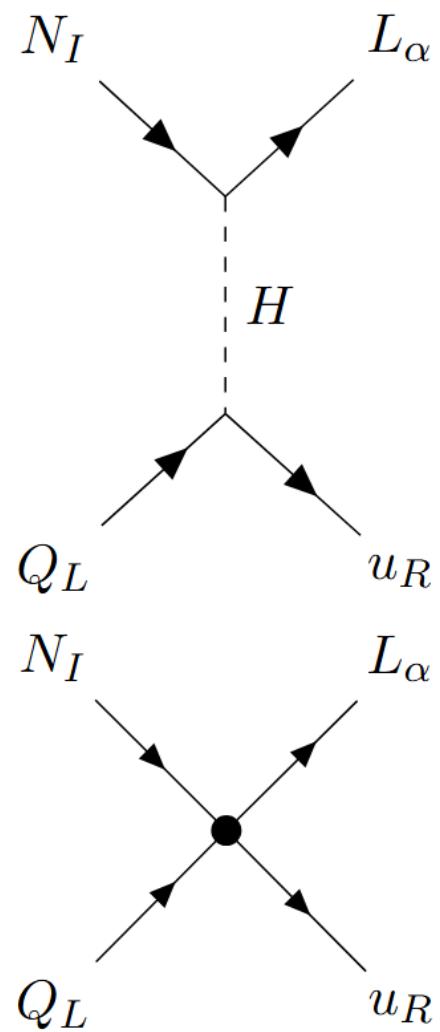
$$\Delta F_{\alpha I} \sim 1.8 \times 10^{-12} \left(\frac{1 \times 10^5 \text{ GeV}}{\Lambda} \right)^2 \left(\frac{[Y_u]_{aa}}{1.3 \times 10^{-5}} \right) [G_{\alpha I}]_{aa}$$

Quantum Kinetic Equations

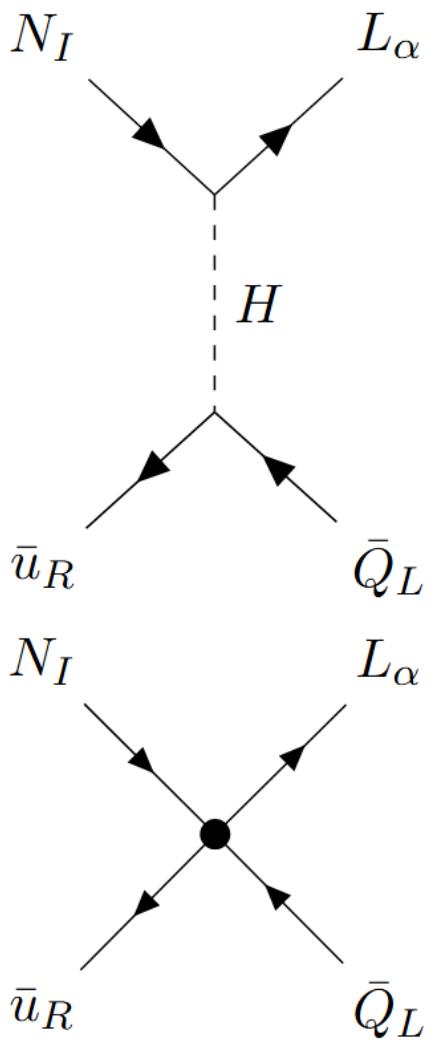
$$zH \frac{dR_N}{dz} = -i [\langle H_{\text{eff}} \rangle, R_N] - \frac{1}{2} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(0),XY} \right\rangle \{ X^\dagger Y, R_N - 1 \} \\ + \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(1),XY} \right\rangle X^\dagger \mu Y - \frac{1}{2} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(2),XY} \right\rangle \{ X^\dagger \mu Y, R_N \} \quad (4.5)$$

$$zH \frac{dR_{\bar{N}}}{dz} = -i [\langle H_{\text{eff}} \rangle, R_{\bar{N}}] - \frac{1}{2} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(0),XY} \right\rangle \{ X^T Y^*, R_{\bar{N}} - 1 \} \\ - \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(1),XY} \right\rangle X^T \mu Y^* + \frac{1}{2} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(2),XY} \right\rangle \{ X^T \mu Y^*, R_{\bar{N}} \} \quad (4.6)$$

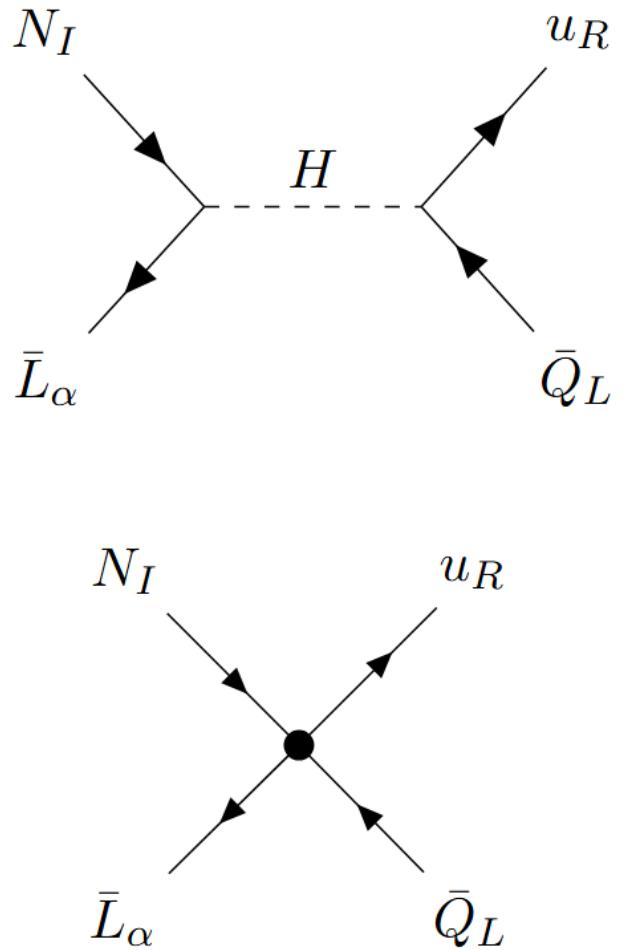
$$zH \frac{d\mu_\alpha}{dz} = \frac{1}{2N_D} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(0),XY} \right\rangle \left[Y R_N X^\dagger - Y^* R_{\bar{N}} X^T \right]_{\alpha\alpha} \\ - \frac{1}{N_D} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(1),XY} \right\rangle \left[Y X^\dagger \right]_{\alpha\alpha} \mu_\alpha \\ + \frac{1}{2N_D} \sum_{X,Y \in \{F,G\}} \left\langle \gamma_N^{(2),XY} \right\rangle \left[Y R_N X^\dagger + Y^* R_{\bar{N}} X^T \right]_{\alpha\alpha} \mu_\alpha , \quad (4.7)$$



(A)

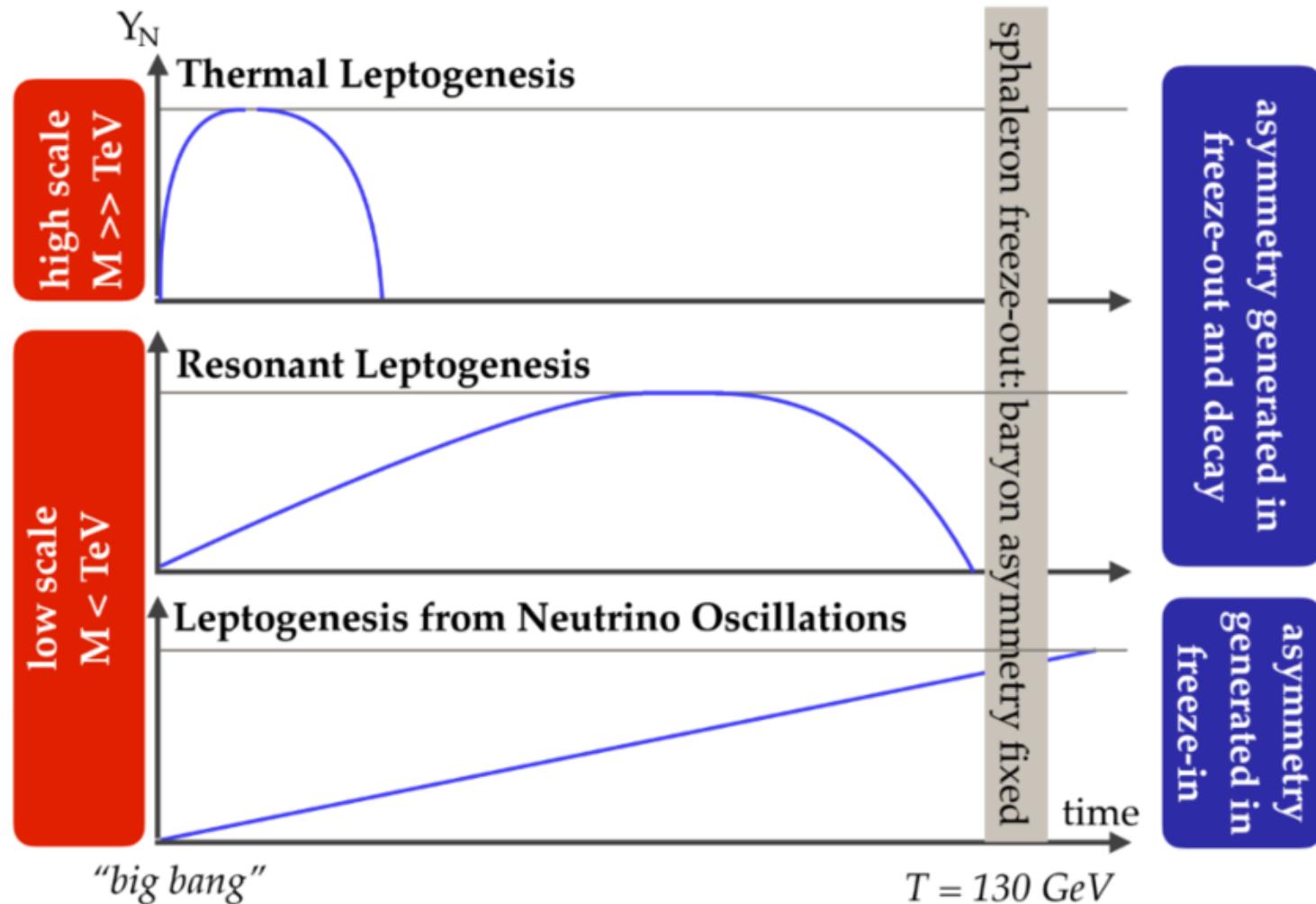


(B)

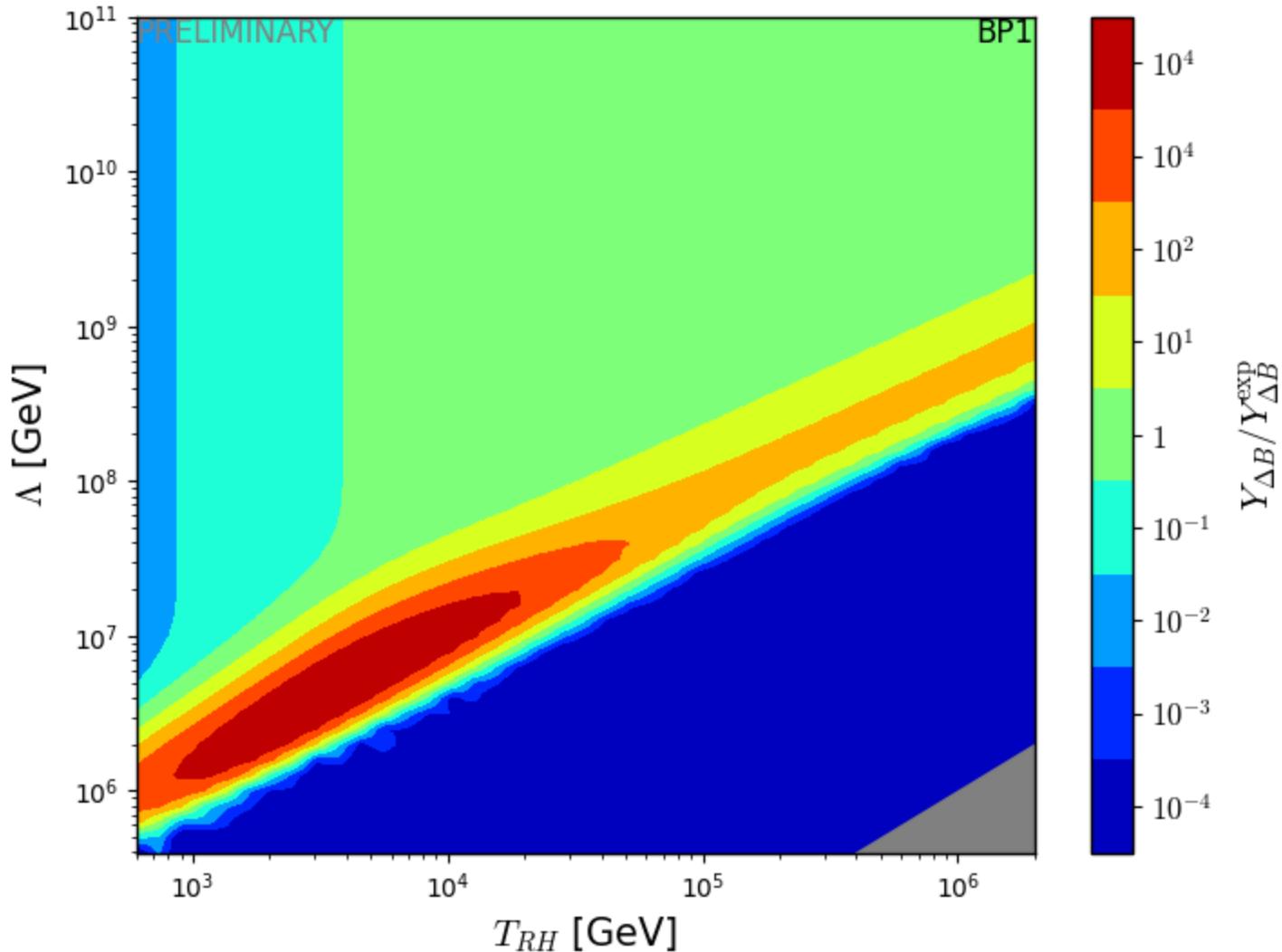


(C)

Types of Leptogenesis

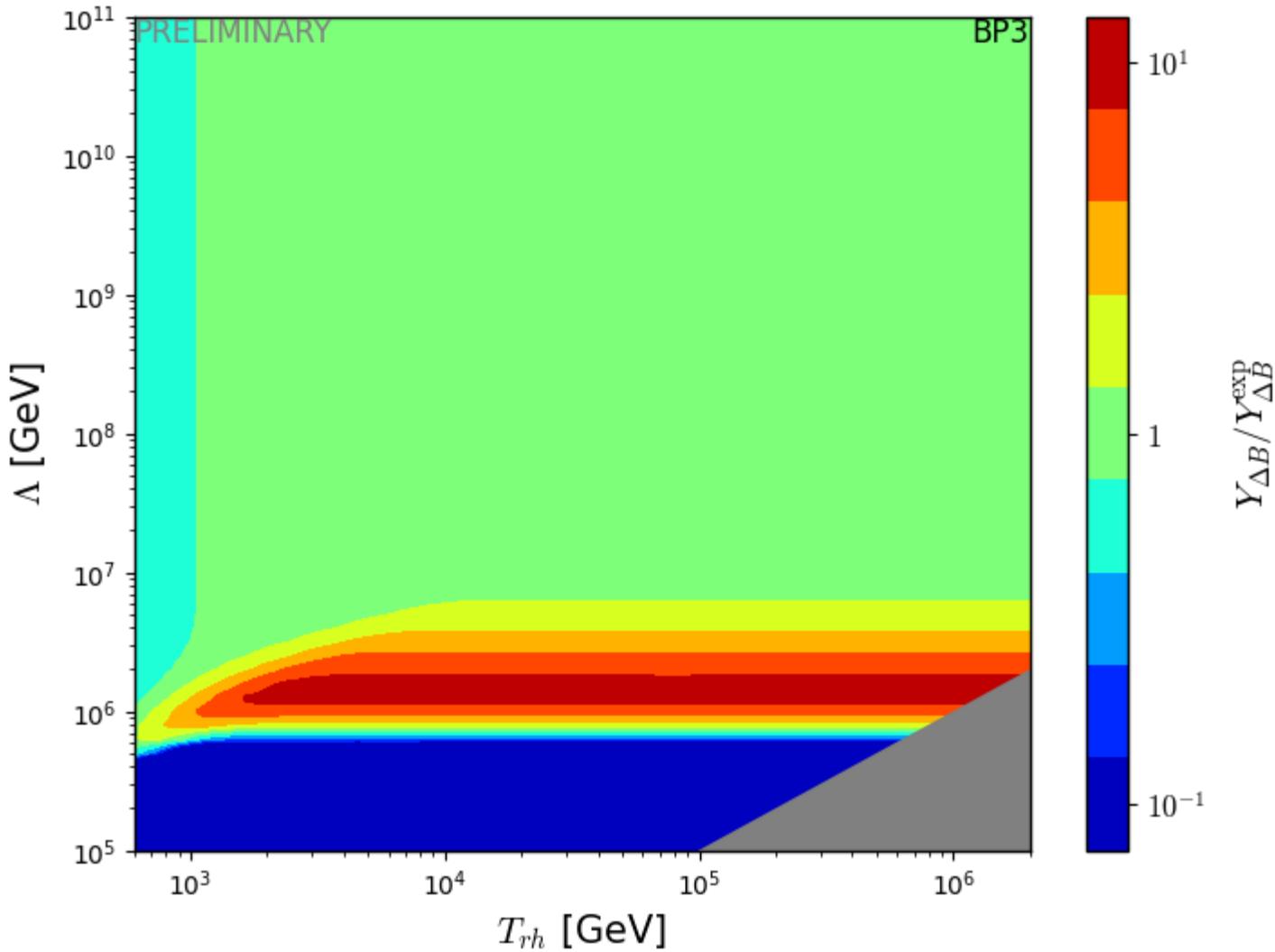


Dependence on reheating temperature



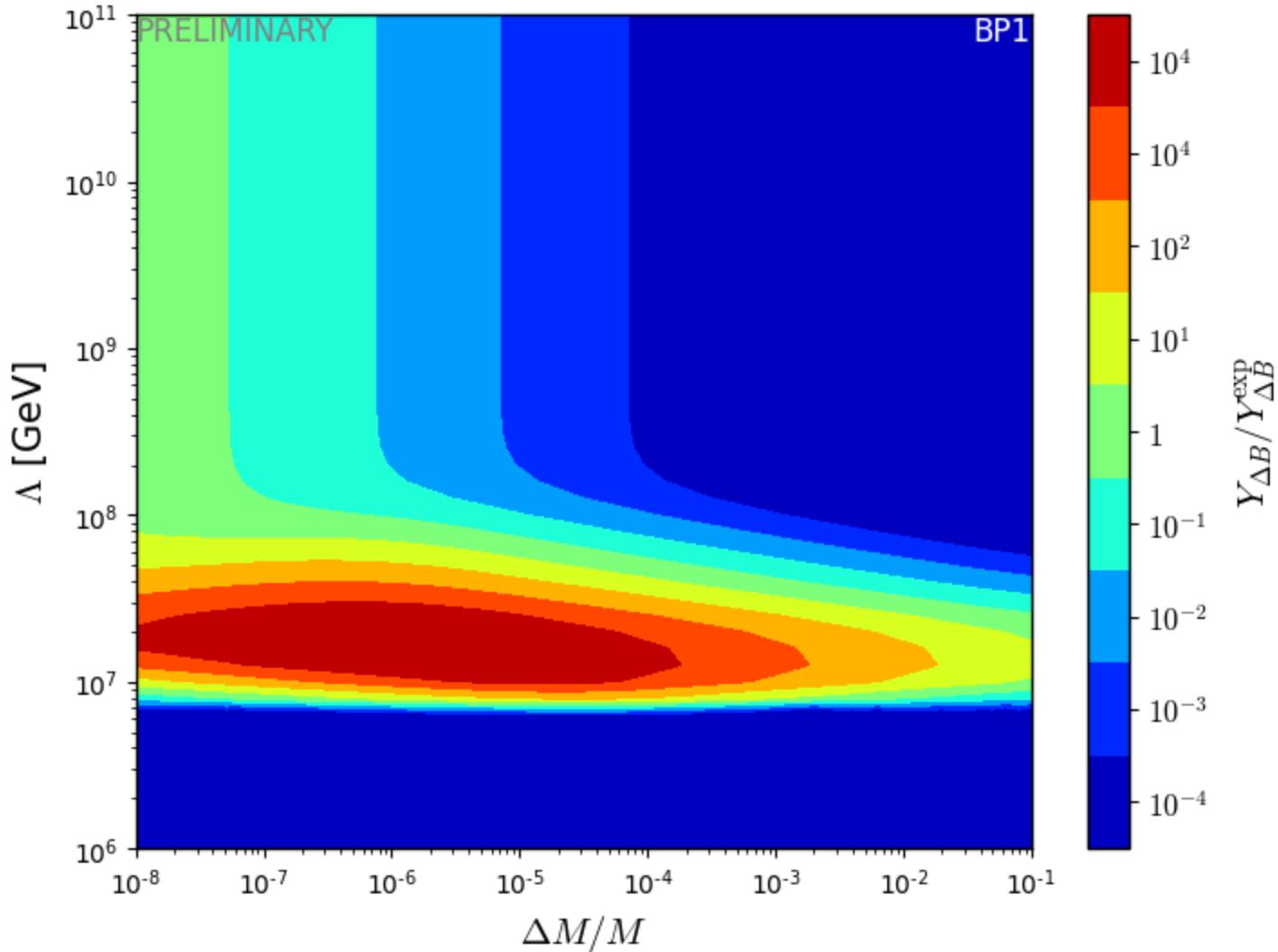
[Fuyuto, Harz, SW in preparation]

Dependence on reheating temperature



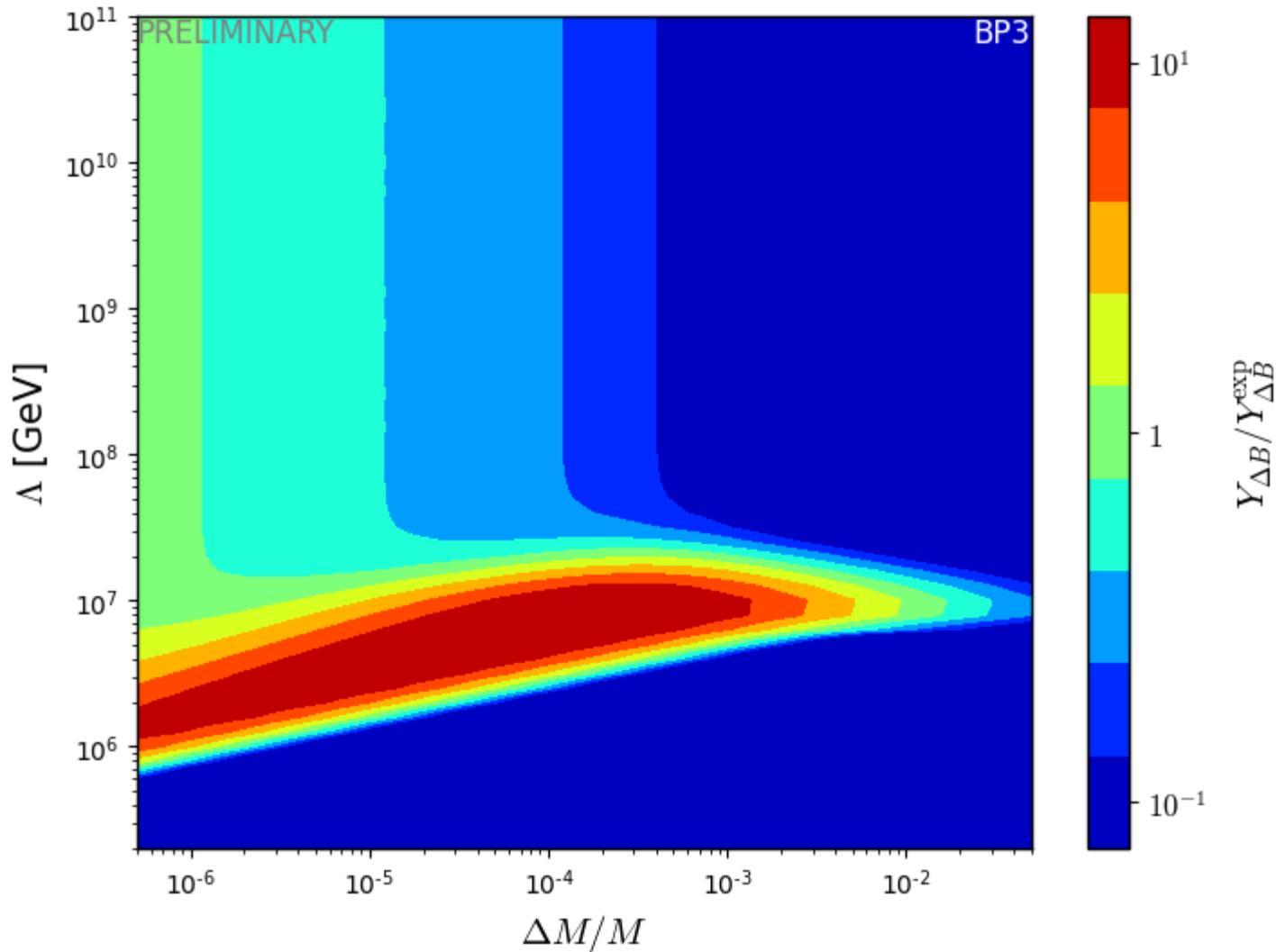
[Fuyuto, Harz, SW in preparation]

Lifting mass degeneracy

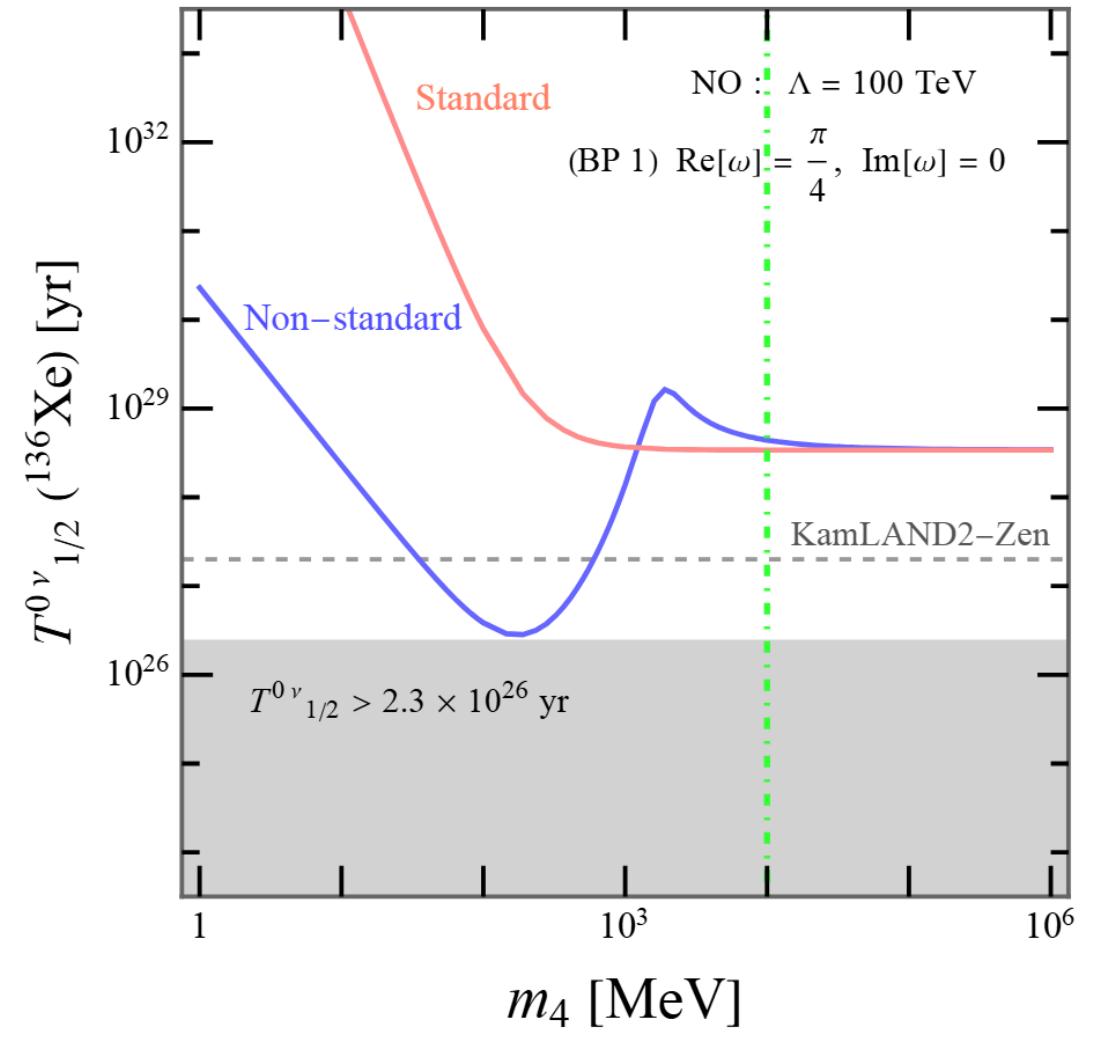
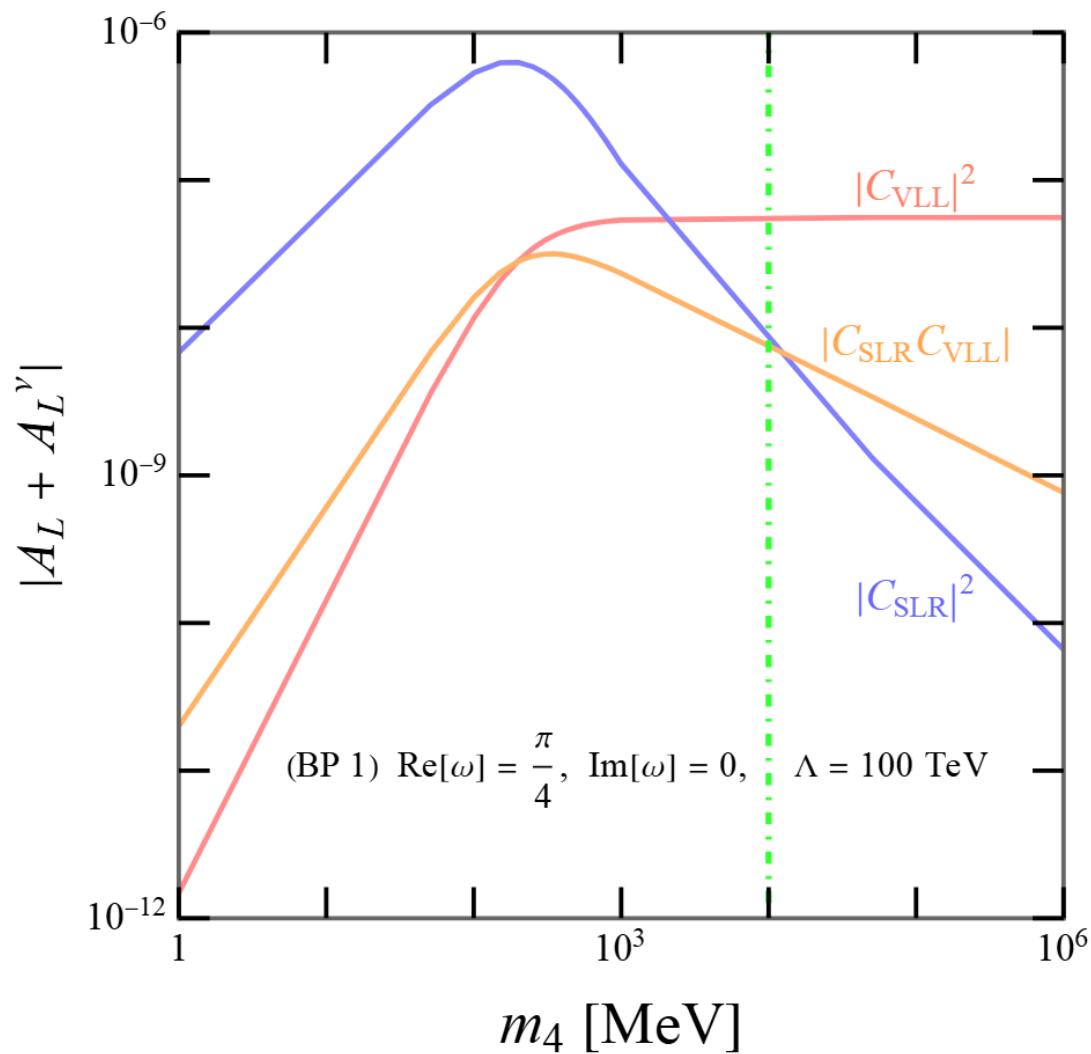


[Fuyuto, Harz, SW in preparation]

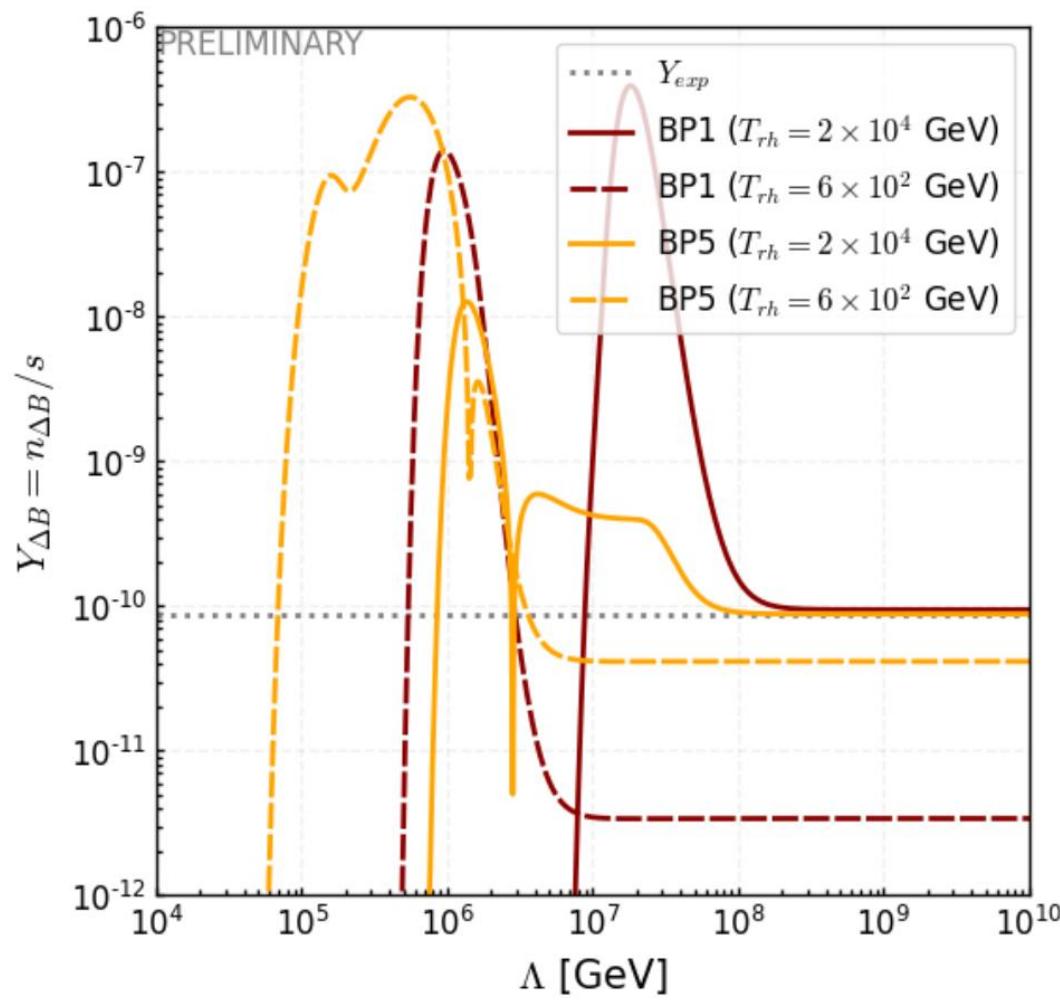
Lifting mass degeneracy



[Fuyuto, Harz, SW in preparation]



[Fuyuto, Harz, SW in preparation]



[Fuyuto, Harz, SW in preparation]