

Something

Do we live on the End of the World?

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Based on [2411.05912] w/ Tony Padilla and Paul Saffin

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"You only realise the importance of someone when they are gone."

1. What is Nothing?

This talk

2. How to get a dS_4 universe from Nothing

[Witten '81]











Witten's Bubble of Nothing

[Witten '81]

"End of the World"









Nothing as a Cobordism







Nothing as a Cobordism







Nothing as a Cobordism

 \bigotimes



Cobordism Conjecture

Non-vanishing cobordism groups $\Omega_k \neq [\emptyset]$ give rise to (D - k - 1)-form global symmetries, with cobordism invariants as preserved global charges



[McNamara, Vafa '19]

All cobordism classes in Quantum Gravity must vanish

 $\Omega_{k}^{QG} = [\emptyset]$







What QG is trying to tell us

1. Nothing is always an allowed configuration

->> we should properly understand what 'nothing' is!

2. ETW branes better be in the spectrum of extended objects

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Compute the gravitational instanton following standard Coleman-de Luccia

$$S_E = -\frac{1}{2\kappa^2} \int_{\text{AdS}_5} d^5 x \sqrt{g} \left(R + \frac{12}{\ell^2} \right) - \frac{1}{\kappa^2} \int_{\text{ETW}} d^4 \xi \sqrt{h} \left(K - \kappa^2 \sigma \right)$$

Make an O(5)-symmetric ansatz and solve Israel junction condition

$$ds^{2} = dr^{2} + \ell^{2} \sinh^{2}\left(\frac{r}{\ell}\right) d\Omega_{4}^{2}$$

$$\checkmark \qquad \sigma = \frac{3}{\kappa^2 \ell} \coth\left(\frac{r_0}{\ell}\right)$$

Attach this to the Lorentzian solution at "t = 0"

$$ds^{2} = dr^{2} + \left(\frac{\sinh(r/\ell)}{\sinh(r_{0}/\ell)}\right)$$

Hubble constant on the bubble is

$$H_0 \sim \ell^{-1} e^{-r_0/\ell} \ll \ell^-$$

 $\int_{0}^{1} \int_{0}^{1} \left[-dt^{2} + H_{0}^{-2} \cosh(H_{0}t) d\Omega_{3}^{2} \right]$



AdS_5 t = 0

ETW brane hosts a dS universe!



Positive tension $\sigma_{\rm ETW} > 0$

Searching for dS

Landscape approach

No-go theorems

Swampland constraints

20+ years of attempts

Domain wall approach

Apparently straightforward

Abundance

QG ----> ETW branes!







Nothing = AdS_5 with $\ell \rightarrow 0$

[Brown, Dahlen '11]

Claim:

 $(\Lambda \rightarrow -\infty)$







Exactly the same as before, but now subject to a gluing condition

$$\mathscr{C}_{-} \sinh\left(\frac{r_{-}}{\mathscr{C}_{-}}\right)$$

and tension

$$\sigma = \frac{3}{\kappa^2} \left[\frac{1}{\ell_-} \coth\left(\frac{r_-}{\ell_-}\right) - \frac{1}{\ell_+} \cot\left(\frac{r_-}{\ell_+}\right) \right]$$

that is negative for up-tunneling.





AdS_5^-



General AdS Bubbles

Oh no! Tension diverges! $\sigma = \sigma_{\rm ETW} - \frac{1}{\kappa^2 \ell_{\perp}} + \mathcal{O}(\ell_{+})$

 \mathscr{C}_+ : AdS length scale ($\rightarrow 0$) $\sigma_{\rm ETW} > 0$: Tension of the ETW brane from before



Infinite tension stems from infinite volume of AdS

Dual CFT calls for UV renormalization



Boundary counterterms



Holographic Renormalization



[c.f. Skenderis '02, Papadimitriou '16] [de Haro, Solodukhin, Skenderis '01]

Introduce holographic counterterms on the conformal boundary



Holographic Renormalization

To better understand the theory in the limit $r \to \infty$ we rewrite in Fefferman-Graham coordinates



$\hat{g}_{ij}(\varrho,\xi) = \hat{g}_{ij}^{(0)}(\xi) + \varrho \hat{g}_{ij}^{(2)}(\xi) + \varrho \hat{g$

$$g^2 + \frac{\ell^2}{\varrho} \hat{g}_{ij} \mathrm{d}\xi^i \mathrm{d}\xi^j$$

$$\varrho^2 \left(\hat{g}_{ij}^{(4)}(\xi) + \log(\varrho) \hat{h}_{ij}^{(4)}(\xi) \right) + \mathcal{O}(\varrho^3)$$

Holographic Renormalization

1. Regularise at some $\varrho = \varepsilon \ll 1$

$$S_{\text{reg},\varepsilon} = \int_{\varrho=\varepsilon} d^4 \xi \sqrt{-\hat{g}^{(0)}} \left(\varepsilon\right)^2$$

2. Introduce counterterms:

3. Renormalise:

$$S_{\text{ren}} = \lim_{\varepsilon \to 0} S_{\text{ren}},$$

 $-2\hat{\mathscr{L}}^{(0)} + \varepsilon^{-1}\hat{\mathscr{L}}^{(2)} - \log(\varepsilon)\hat{\mathscr{L}}^{(4)} + \mathcal{O}(\varepsilon)$

 $S_{\text{ct},\varepsilon} = -\operatorname{div}\left(S_{\text{reg},\varepsilon}\right)$

 $\varepsilon = \lim_{\varepsilon \to 0} S_{\operatorname{reg},\varepsilon} + S_{\operatorname{ct},\varepsilon}$





Our claim

















Symbols proof

$S_{\text{bubble}} = S_{\text{in}}[-] + S_{\text{bare}}[\text{brane}] + S_{\text{out}}[+]$

$S_{\text{bubble}} = S_{\text{in}}[-] + (S_{\text{bare}}[\text{brane}] + S_{\text{bare}}]$

$S_{out}[+] = S_{ren}[+] - S_{reg,\varrho_+}[+]$ = $S_{ren}[+] - S_{ren,\varrho_+}[+] + S_{ct,\varrho_+}[+]$

$$S_{ct,Q_{+}}[+]) + S_{ren}[+] - S_{ren,Q_{+}}[+]$$

Symbols proof

One can readily compute the holographic counterterms



$$R + \frac{\ell_{+}^{2} \log \varrho_{+}}{8} \left(R_{ij} R^{ij} - \frac{1}{3} R^{2} \right)$$

$$\sigma = \sigma_{\rm ETW} - \frac{3}{\kappa^2 \ell_+} + \mathcal{O}(\ell_+)$$







Our claim



Summary Two fun take-home lessons

Absence of spacetime is a spacetime

of infinite negative curvature. It is

able to decay.

everywhere singular metric.

holographic in nature.

dual "CFT" with $c \rightarrow 0$

dS on ETW branes is straightforward works in d dimensions and to all orders in perturbations

->> embedding in Type IIB using topological terms Some ongoing work:

→ early universe (cosmologists?)

a "dual AdS Distance Conjecture"?



Thank you