

Deriving Weyl Double Copies with Sources

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- ▶ Introduction: Gravity and Gauge Theories
- ▶ The Classical Double Copy
- ▶ Finding a Classical Double Copy for Einstein Maxwell Gravity
- ▶ Outlook
- \triangleright An open problem in theoretical physics is the relationship between theories of gravity and theories of particle physics.
- \triangleright The last 50 years have seen various attempts to resolve this problem: String Theory, Quantum Loop Gravity, etc ...
- ▶ Though none of these attempts have completely solidified the relationship between gravity and particle physics, we have learned a great deal from these investigations.

The standard model of particle physics is described by gauge theories; quantum field theories which incorporate local symmetries defined at every point in spacetime.

The most relevant of these theories for us is the theory of quarks, gluons and the strong nuclear force: **Quantum Chromodynamics**.

We can also remove the quarks from QCD to receive a non-abelian gauge theory known as Yang-Mills theory.

- \blacktriangleright Currently, the best theory we have to describe gravity is general relativity.
- ▶ Recent experimental results on black holes and gravitational waves has increased interest in the subject.
- \triangleright Gravity is the only fundamental force that we cannot comfortably combine with quantum mechanics.
- ▶ Gravity is an example of a **non-renormalisable** theory.

Quantum Field Theory and Interactions

Scattering amplitudes in quantum field theory are quantities related to the probability for an interaction (also known as a scattering process) between particles to happen.

- ▶ Number of (external legs) points \rightarrow Number of arrows going in **and** out.
- ▶ Number of Loops \rightarrow Number of "self interactions".

$$
\mathcal{A}_m^{(L)} = i^{L-1} \underbrace{\widetilde{g}^{m-2+2L}}_{\text{Coupling Constant}} \sum_i \int \prod_{l=1}^L \frac{d^D \ell_l}{(2\pi)^D} \frac{1}{S_i} \frac{c_i n_i}{\prod_{i_j} d_{i_j}}, \qquad (1)
$$

$$
\mathcal{A}_{m}^{(L)}=i^{L-1}g^{m-2+2L}\sum_{i}\prod_{\text{Integral over loop momenta }l_{i}}\frac{1}{\int_{l=1}^{L}\frac{d^{D}\ell_{l}}{(2\pi)^{D}}}\frac{1}{S_{i}}\frac{c_{i}n_{i}}{\prod_{i_{j}}d_{i_{j}}},
$$

$$
\mathcal{A}_{m}^{(L)} = i^{L-1} g^{m-2+2L} \sum_{i} \int \prod_{l=1}^{L} \frac{d^{D} \ell_{l}}{(2\pi)^{D}} \frac{1}{S_{i}} \frac{c_{i} n_{i}}{\prod_{i_{j}} d_{i_{j}}},
$$

$$
\mathcal{A}_m^{(L)} = i^{L-1} g^{m-2+2L} \sum_i \int \prod_{l=1}^L \frac{d^D \ell_l}{(2\pi)^D} \frac{1}{S_i} \frac{\overbrace{C_i}^{\text{Color Factors}} n_i}{\prod_{i_j} d_{i_j}},
$$

$$
\mathcal{A}_{m}^{(L)}=i^{L-1}g^{m-2+2L}\sum_{i}\int\prod_{l=1}^{L}\frac{d^{D}\ell_{l}}{(2\pi)^{D}}\frac{1}{S_{i}}\frac{c_{i}}{\prod_{i_{j}}d_{i_{j}}},
$$

$$
\mathcal{A}_{m}^{(L)} = i^{L-1} g^{m-2+2L} \sum_{i} \int \prod_{l=1}^{L} \frac{d^{D} \ell_{l}}{(2\pi)^{D}} \frac{1}{S_{i}} \underbrace{\prod_{j} d_{i_{j}} d_{j_{j}}}_{\text{Internal Lines Propagator}},
$$

It turns out the kinematic numerators (n_i) can be made to obey (mirroring the colour factors c_i):

$$
c_i + c_j + c_k = 0
$$

\n
$$
n_i + n_j + n_k = 0
$$
\n(2)

This has become known as BCJ Duality (0805.3993, 1004.0476). Importantly, equation [\(3\)](#page-13-0) implies the existance of structures known as **kinematic algebras**!

This allows us to write down relevant scattering amplitudes in quantum gravity.

We can promote gravity to a quantum field theory (ignoring issues with renormalizability), to write down a scattering amplitude with L loops and m points:

$$
\mathcal{M}_m^{(L)} = i^{L-1} \left(\frac{\kappa}{2}\right)^{m-2+2L} \sum_i \int \prod_{l=1}^L \frac{d^D \ell_l}{(2\pi)^D} \frac{1}{S_i} \frac{n_i \tilde{n}_i}{\prod_{i_j} d_{i_j}},\qquad(4)
$$

For:

$$
\kappa=\sqrt{32\pi G_N}
$$

We can promote gravity to a quantum field theory (ignoring issues with renormalizability), to write down a scattering amplitude with L loops and m points:

second set
\n
$$
\mathcal{M}_m^{(L)} = i^{L-1} \left(\frac{\kappa}{2}\right)^{m-2+2L} \sum_i \int \prod_{l=1}^L \frac{d^D \ell_l}{(2\pi)^D} \frac{1}{S_i} \frac{n_i}{\prod_{i_j} d_{i_j}},
$$

Comparing both expressions:

$$
\mathcal{A}_{m}^{(L)} = i^{L-1} g^{m-2+2L} \sum_{i} \int \prod_{l=1}^{L} \frac{d^{D} \ell_{l}}{(2\pi)^{D}} \frac{1}{S_{i}} \frac{c_{i} n_{i}}{\prod_{i_{j}} d_{i_{j}}},
$$

$$
\mathcal{M}_{m}^{(L)} = i^{L-1} \left(\frac{\kappa}{2}\right)^{m-2+2L} \sum_{i} \int \prod_{l=1}^{L} \frac{d^{D} \ell_{l}}{(2\pi)^{D}} \frac{1}{S_{i}} \frac{n_{i} \tilde{n}_{i}}{\prod_{i_{j}} d_{i_{j}}},
$$

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$$

$$
\mathcal{M}_{m}^{(L)} = i^{L-1} \left(\frac{\kappa}{2}\right)^{m-2+2L} \sum_{i} \int \prod_{l=1}^{L} \frac{d^{D} \ell_{l}}{(2\pi)^{D}} \frac{1}{S_{i}} \frac{n_{i} \tilde{n}_{i}}{\prod_{i_{j}} d_{i_{j}}},
$$

We can turn the gauge theory amplitude into the gravity amplitude via the following replacements:

$$
\mathcal{M}_m^{(L)} = \mathcal{A}_m^{(L)}\Big|_{\substack{c_i \to \tilde{n}_i \\ g \to \kappa/2}} = i^{L-1} \left(\frac{\kappa}{2}\right)^{m-2+2L} \sum_i \int \prod_{l=1}^L \frac{d^D \ell_l}{(2\pi)^D} \frac{1}{S_i} \frac{n_i \tilde{n}_i}{\prod_{i_j} d_{i_j}},\tag{5}
$$

We then make the choice $n_i \equiv \tilde{n}_i$.

This is known as the **Double Copy**.

- \blacktriangleright The double copy can be extended to produce dualities between **exact** classical solutions (classical yang-mills and general relativity).
- \blacktriangleright The first instance of the classical double copy is the so-called Kerr-Schild Double Copy (1410.0239, 1606.04724).
- \blacktriangleright This relates gauge fields in classical yang mills (A_μ^a) with exact linearized vacuum solutions to the Einstein equations $(h_{\mu\nu})$. Where:

$$
g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu} \tag{6}
$$

- ▶ The Kerr Schild Double Copy **only** applies to Kerr Schild Solutions in General Relativity.
- ▶ We can extend this relationship to work for broad class of GR exact solutions, in a framework known as the **Weyl double copy** (1810.08183).
- ▶ The **Weyl** double copy relates spinor quantities in General Relativity to their counterparts in Electromagnetism.
- ▶ This relationship has only been seen to work for Petrov type D and N vacuum Spacetimes.
- ▶ We can **convert** tensorial objects into spinors using the Infeld-van der Waerden symbols *σ µ* μ
ΑÀ
- ▶ Normally, field theorists tend to work in the language of tensors.
- ▶ However, another formalism exists in the language of **2 component** (or *Weyl*)-spinors λ_A and their complex conjugates $\tilde{\lambda}_A$.
- \blacktriangleright We can also define their duals λ^A and $\tilde{\lambda}^{\dot{A}}$ respectively.
- \triangleright We raise and lower spinorial indices using the 2-dimension Levi-Civita symbol (*ε_{AB}*, *ε^{AB}*, *ε_{AB}*, *ε^{AB}*).
- ▶ We can also build multi-indexed objects under is formalism (such as $F_{A\dot{A}R\dot{B}}$).

For our purposes, we define:

$$
\sigma_{A\dot{A}}^{\mu} = (I_{2\times 2}, \sigma_{z}, \sigma_{x}, i\sigma_{y})
$$
 (7)

Where $\sigma_x, \sigma_y, \sigma_z$ are the standard Pauli matrices and $I_{2\times 2}$ is the 2 dimensional identity matrix.

The advantage of using spinors is that we can see things that otherwise would be **hidden** in the tensor language.

Constructing The Weyl Double Copy

Converting the Electromagnetic Field Strength tensor F*µν* into spinors we receive:

$$
F_{\dot{A}\dot{A}B\dot{B}} = \sigma^{\mu}_{\dot{A}\dot{A}} \sigma^{\nu}_{\dot{B}\dot{B}} F_{\mu\nu}
$$

$$
= \phi_{AB} \varepsilon_{\dot{A}\dot{B}} + \tilde{\phi}_{\dot{A}\dot{B}} \varepsilon_{AB}
$$
(8)

Where ϕ_{AB} is the Maxwell Spinor.

We can obtain ϕ_{AB} directly by:

$$
\phi_{AB} = \frac{1}{2} F_{ABC}{}^C \tag{9}
$$

Constructing The Weyl Double Copy

For vacuum solutions $(R_{\mu\nu}=0)$, the Riemann tensor reduces to the **Weyl** tensor:

$$
R_{\mu\nu\rho\lambda} = W_{\mu\nu\rho\lambda} \tag{10}
$$

Converting W*µνρλ* into a spinor we receive:

$$
W_{A\dot{A}B\dot{B}C\dot{C}D\dot{D}} = \sigma_{A\dot{A}}^{\mu} \sigma_{B\dot{B}}^{\nu} \sigma_{C\dot{C}}^{\rho} \sigma_{D\dot{D}}^{\lambda} W_{\mu\nu\rho\lambda}
$$

= $\Psi_{ABCD} \varepsilon_{\dot{A}\dot{B}} \varepsilon_{\dot{C}\dot{D}} + \tilde{\Psi}_{\dot{A}\dot{B}\dot{C}\dot{D}} \varepsilon_{AB} \varepsilon_{CD}$ (11)

We can obtain Ψ_{ABCD} from $W_{A\dot{A}B\dot{B}C\dot{C}D\dot{D}}$ by:

$$
\Psi_{ABCD} = \frac{1}{4} W_{A\dot{A}B\dot{B}C\dot{C}D\dot{D}c}{}^{\dot{A}\dot{B}}{}_{\varepsilon}{}^{\dot{C}\dot{D}} \tag{12}
$$

Constructing The Weyl Double Copy

Thus for **certain** vacuum solutions in general relativity yields the result:

$$
\Psi_{ABCD} = \frac{\phi_{(AB}\phi_{CD)}}{S} \tag{13}
$$

Where S is some scalar which is a solution to the equation of motion of Biadjoint Scalar Theory.

This relationship is known as the **Weyl Double Copy**.

The Weyl Double Copy works for **Exact** "Linear" solutions to the Einstein equations.

- ▶ Einstein-Maxwell gravity is a theory which allows for the matter inside our spacetime to be electromagnetically charged.
- ▶ For **non-vacuum solutions**, the theory admits a Riemann Curvature Spinor $R_{AABBCCDD}$, given in the form:

$$
R_{\dot{A}\dot{B}\dot{B}\dot{C}\dot{C}\dot{D}\dot{D}} = \Psi_{ABCD}\epsilon_{\dot{A}\dot{B}}\epsilon_{\dot{C}\dot{D}} + \bar{\Psi}_{\dot{A}\dot{B}\dot{C}\dot{D}}\epsilon_{AB}\epsilon_{CD} + \Phi_{AB\dot{C}\dot{D}}\epsilon_{\dot{A}\dot{B}}\epsilon_{CD} + \bar{\Phi}_{\dot{A}\dot{B}\dot{C}D}\epsilon_{AB}\epsilon_{\dot{C}\dot{D}} + 2\Lambda\left(\epsilon_{AC}\epsilon_{BD}\epsilon_{\dot{A}\dot{C}}\epsilon_{\dot{B}\dot{D}} - \epsilon_{AD}\epsilon_{BC}\epsilon_{\dot{A}\dot{D}}\epsilon_{\dot{B}\dot{C}}\right)
$$
\n(14)

Weyl Double Copy for Einstein-Maxwell Gravity

▶ It was conjectured in (2210.16339, 2110.02293) that one could write down a Weyl Double Copy like formula for non-vacuum Einstein-Maxwell gravity:

$$
\Psi_{ABCD} = \sum_{n=1}^{m} \frac{1}{\phi^{(n)}} \Phi_{(AB}^{(n)} \Phi_{CD)}^{(n)},
$$
(15)

- ▶ The right-hand side contains a **"tower"** of electromagnetic spinors $\Phi_{AB}^{(n)}$ and scalar fields $\phi^{(n)}$.
- \blacktriangleright The $n = 1$ term corresponds to the traditional **vacuum Weyl double copy**.

Weyl Double Copy for Einstein-Maxwell Gravity

 \blacktriangleright **Each** $n > 1$ electromagnetic spinor corresponds to a field strength $\mathit{F_{\mu\nu}^{(n)}}$ satisfying a non-vacuum Maxwell equation

$$
\partial^{\mu}F^{(n)}_{\mu\nu} = j^{(n)}_{\mu}.\tag{16}
$$

▶ **Each** scalar field $n > 1$ $\phi^{(n)}$ satisfies a non-vacuum equation for some charge density $\rho_S^{(n)}$ S :

$$
\partial^2 \phi^{(n)} = \rho_S^{(n)} \tag{17}
$$

 \blacktriangleright How can we systematically prove these results?

The Double Copy: Momentum to Position space

- \triangleright The Double Copy for classical (*position*) solutions and for scattering amplitudes (momentum) are related to each other by special integral transforms. (2208.08548)
- \triangleright Using this relationship, we can systematically derive Weyl Double Copy formulas from scattering amplitudes in a **three-stage process**.
- ▶ In particular, the above discussion **must** be carried out in (2*,* 2) signature. (n.b It is always possible to analytically continue to any choice of signature after one has obtained the final classical fields!)

Systematically deriving the Weyl Double Copy

- (i) Express spinorial solutions of (linearised) classical equations for scalar, gauge and gravity theory as **inverse Fourier transforms** of **momentum-space solutions**.
- (ii) One may **transform** the momentum integral to certain spinor variables, and carry out some of the integrals to yield an intermediate twistor-space representation of each classical solution. This step reproduces the **Twistor Double Copy** (2012.02479, 2103.16441).
- (iii) The remaining integral can be carried out to yield position-space representations of each **classical** solution, with their principal spinors clearly identified. These solutions are then found to obey the Weyl double copy.

Stage 1: Inverse Fourier Transforms

- ▶ Another way to obtain these inverse Fourier transforms of momentum-space solutions is to consider the sources of these fields (2112.05111).
- ▶ For our purposes, we choose *spherically symmetric* and *static* sources/currents for each field type (scalar, vector, tensor) :

$$
\rho_S = \rho(x), \quad j^{\mu}(x) = \rho(x)u^{\mu}, \quad T^{\mu\nu}(x) = \rho(x)u^{\mu}u^{\nu}, \quad (18)
$$

 \triangleright Which in momentum space is given by:

$$
\tilde{\rho}_{S}(k) = \delta(u \cdot k) \mathcal{J}(\vec{k}), \quad \tilde{j}^{\mu}(k) = \delta(u \cdot k) \mathcal{J}(\vec{k}) u^{\mu}, \n\tilde{T}^{\mu\nu}(k) = \delta(u \cdot k) \mathcal{J}(\vec{k}) u^{\mu} u^{\nu}, \quad (19)
$$

Stage 1: Inverse Fourier Transforms

The **Scalar Field** pertaining to S:

$$
\phi(x) = \int \frac{d^4k}{(2\pi)^4} \delta(u \cdot k) \frac{\mathcal{J}(\vec{k})}{k^2} e^{-ik \cdot x}.
$$
 (20)

The **Electromagnetic Field Strength** Tensor:

$$
F^{\mu\nu}(x) = \int \frac{d^4k}{(2\pi)^4} k^{[\mu} u^{\nu]} \delta(u \cdot k) \frac{\mathcal{J}(\vec{k})}{k^2} e^{-ik \cdot x}.
$$
 (21)

The **Riemann Curvature Tensor**:

$$
R^{\mu\nu\rho\sigma}(x) = \kappa \int \frac{d^4k}{(2\pi)^4} \left[U^{\rho[\mu}k^{\nu]}k^{\sigma} - U^{\sigma[\mu}k^{\nu]}k^{\rho} \right] \frac{\mathcal{J}(\vec{k})}{k^2} e^{-ik \cdot x},\tag{22}
$$

Stage 1: Inverse Fourier Transform

 \blacktriangleright Where u^{μ} is its 4-velocity:

$$
u^{\mu} = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix} \tag{23}
$$

- $\rho(x)$ represents the appropriate **charge** density, whereas it will be an **energy** density in gravity.
- ▶ *k* is the 4-momentum $k^{\mu} = (k^0, \vec{k})$ in a stationary frame, with $k^0 = u \cdot k$.
- \triangleright Where in [\(22\)](#page-32-0) we define:

$$
U^{\nu\rho} = u^{\nu}u^{\rho} - \frac{1}{2}\eta^{\nu\rho}.
$$
 (24)

 \triangleright By inspection of [\(20\)](#page-32-1), [\(21\)](#page-32-2) and [\(22\)](#page-32-0), we see that our expressions for the electromagnetic and Riemann Curvature tensors can be written **in terms** of a scalar field $\phi^{(n)}(x)$:

$$
F_{\mu\nu}^{(n)} = 2u_{[\mu}\partial_{\nu]}\phi^{(n)}.
$$
 (25)

$$
R^{(n)}_{\mu\nu\rho\sigma}(x) = -2 \left[U_{\rho[\mu} \partial_{\nu]} \partial_{\sigma} - U_{\sigma[\mu} \partial_{\nu]} \partial_{\rho} \right] \phi^{(n)}(r). \tag{26}
$$

Stage 2: Integral Transforms and Integration

- ▶ Moving to **stage 2** of the process, we now wish to convert our integrals into the language of 2-spinors, and then integrate out some of the degrees of freedom.
- \blacktriangleright Let's start with the case of the scalar integral.
- \triangleright Our momentum 4-vector is given by:

$$
k^{\mu} = \omega \ell^{\mu} + \xi q^{\mu}, \qquad (27)
$$

where $\ell^2=q^2=0$; the parameters ω $(o$ n-shell) and ξ (off-shell) have dimensions of energy; with a fixed 4-vector:

$$
q^{\mu} = \frac{1}{2}(0, 0, -1, 1) \tag{28}
$$

Momentum Spinorial Translation

▶ Spinorial translating our 4-momentum we get $k_{\Delta\Delta}$:

$$
k_{A\dot{A}} = \omega \lambda_A \tilde{\lambda}_{\dot{A}} + \xi q_{A\dot{A}}, \qquad (29)
$$

 \blacktriangleright Where λ_A and $\tilde{\lambda}_{\dot{A}}$ are now defined in terms of a complex number z and its complex conjugate \bar{z} :

$$
\lambda_A = \left(\frac{1}{\sqrt{z}}, \sqrt{z}\right), \quad \tilde{\lambda}_{\dot{A}} = \left(\frac{1}{\sqrt{-\tilde{z}}}, -\sqrt{-\tilde{z}}\right). \tag{30}
$$

▶ In (2, 2)-signature, complex numbers *z* and their complex conjugate \bar{z} **decouple**. E.g:

$$
\tilde{z} \neq (z)^* \tag{31}
$$

Delta Function Transformation

The delta function $\delta(u \cdot k)$ spinorial translation depends on the order in which we perform the integrals over dz and $d\tilde{z}$. The delta function first translates to:

$$
\delta(u \cdot k) \to \delta\left(u_{A\dot{A}}\lambda^{A}\tilde{\lambda}^{\dot{A}}\right) = \frac{1}{m}\delta(U^{A\dot{A}}(\omega \lambda_{A}\tilde{\lambda}_{\dot{A}}))
$$

$$
= \frac{2}{\omega}\delta\left(\frac{1+z\tilde{z}}{\sqrt{-z\tilde{z}}}\right)
$$
(32)

If we are integrating over dz:

$$
\frac{2}{\omega}\delta\left(\frac{1+z\tilde{z}}{\sqrt{-z\tilde{z}}}\right) = \frac{2}{\omega\tilde{z}}\delta\left(z + \frac{1}{\tilde{z}}\right)
$$
(33)

If we are integrating over $d\tilde{z}$:

$$
\frac{2}{\omega} \delta \left(\frac{1+z\tilde{z}}{\sqrt{-z\tilde{z}}} \right) = \frac{2}{\omega z} \delta \left(\tilde{z} + \frac{1}{z} \right)
$$
(34)

Delta Functions

An important implication of integrating over dz or $d\tilde{z}$ is the replacement rules it generates for $\lambda_{\mathcal{A}}$ and $\tilde{\lambda}_{\dot{\mathcal{A}}}$ respectively:

Integration over dz**:**

$$
\lambda_A = U_A{}^A \tilde{\lambda}_A \tag{35}
$$

Integration over $d\tilde{z}$:

$$
\tilde{\lambda}_{\dot{A}} = -U^A_{\dot{A}} \lambda_A \tag{36}
$$

Where we define the mixed raised and lowered indexed spinorial translation of the 4-velocity u^{μ} :

$$
U_A^{\dot{A}} = \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}, \tag{37}
$$

$$
U_{\dot{A}}^A = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}, \tag{38}
$$

Stage 2: Integral Transforms and Integration

We now have a *change of variables fr*om d^4k to (ω,ξ,z,\tilde{z}) :

$$
d^4k = \frac{dzd\tilde{z}d\omega d\xi}{4(2\pi)^3},\tag{39}
$$

We can then find:

$$
\phi(x) = \int \frac{d\omega d\xi dz d\tilde{z}}{2(2\pi)^4} \frac{\omega \sqrt{-\tilde{z}}}{\tilde{z}\sqrt{z}} \delta\left(\frac{1+z\tilde{z}}{\sqrt{-z\tilde{z}}}\right) \times
$$

$$
e^{-i\xi q \cdot x} \exp\left[-\frac{i\omega}{2} \tilde{\lambda}_A \tilde{\lambda}_B x^{A\tilde{A}} u_A{}^{\tilde{B}}\right] \mathcal{J}(\vec{k}). \tag{40}
$$

It is now useful to consider an example of this integral in practice.

Reissner-Nordstrom black hole

- ▶ The **Reissner-Nordstrom black hole** is simply a Schwarzschild Black hole that has been allowed to be electrically charged.
- ▶ In particular, its Weyl Double Copy is given by:

$$
\Psi_{ABCD}
$$
\n
$$
= \Psi_{ABCD}^{(1)} + \Psi_{ABCD}^{(2)}
$$
\n
$$
= \frac{1}{\phi^{(1)}} \Phi_{(AB}^{(1)} \Phi_{CD}^{(1)} + \frac{1}{\phi^{(2)}} \Phi_{(AB}^{(2)} \Phi_{CD}^{(2)} \tag{41}
$$

▶ We will be focusing on deriving $\phi^{(2)}$, $\Phi_{AB}^{(2)}$ and $\Psi_{ABCD}^{(2)}$. From the literature, we know that:

$$
\phi^{(2)} \propto \frac{1}{r^2} \tag{42}
$$

Scalar Field Calculation

 \triangleright This implies our source current is given by:

$$
\mathcal{J}(\vec{k}) = 2\pi^2 |\vec{k}|.\tag{43}
$$

▶ Applying this to our integral for *ϕ* (2) and integrating over *ξ*, *ω* and z we find:

$$
\phi^{(2)}(x) = \frac{1}{\sqrt{32}\pi} \frac{1}{\sqrt{q \cdot x}} \int d\tilde{z} \frac{1}{\tilde{z}^{3/2}} \frac{1}{\left(\tilde{\lambda}_A \tilde{\lambda}_B x^{A\dot{A}} u_A \dot{B}\right)^{3/2}}.
$$
 (44)

 \blacktriangleright Where $q \cdot x$ is given by:

$$
q \cdot x = \frac{1}{2} q_{A\dot{A}} x^{A\dot{A}} \tag{45}
$$

It is now convenient to define the rescaled spinor

$$
\tilde{\chi}_{\dot{A}} = \sqrt{-\tilde{z}} \tilde{\lambda}_{\dot{A}} = \begin{pmatrix} 1 \\ \tilde{z} \end{pmatrix}, \tag{46}
$$

such that [\(44\)](#page-41-0) can be written more compactly as

$$
\phi^{(2)}(x) = \frac{1}{\sqrt{32}\pi} \frac{1}{\sqrt{q \cdot x}} \int d\tilde{z} \frac{1}{\left(\tilde{\chi}_{\tilde{A}} \tilde{\chi}_{\tilde{B}} x^{A\tilde{A}} u_{A} \tilde{B}\right)^{3/2}}.
$$
 (47)

One complication of [\(47\)](#page-42-0) is that its singularity structure involves branch cuts rather than simple poles, and it is not then immediately clear how to choose an appropriate contour.

Scalar Field Calculation

However, we can instead carry out the **inverse Fourier transform** of our original expression for $\phi^{(2)}$ in [\(20\)](#page-32-1) to obtain:

$$
\phi^{(2)}(x) = \int \frac{d^4k}{(2\pi)^4} \frac{2\pi^2 \delta(u \cdot k)|\vec{k}|}{k^2} e^{-ik \cdot x} \n= \frac{1}{(x^2 - (u \cdot x)^2)} = \frac{1}{r^2}.
$$
\n(48)

Comparison with [\(47\)](#page-42-0) then yields

$$
\int d\tilde{z} \frac{1}{\left(\tilde{\chi}_{\tilde{A}}\tilde{\chi}_{\tilde{B}} \times^{A\tilde{A}} u_{A} \tilde{B}\right)^{3/2}} = \frac{\sqrt{32}\pi \sqrt{q \cdot x}}{r^{2}}.
$$
 (49)

With denominator being give as:

$$
\tilde{\chi}_{\dot{A}}\tilde{\chi}_{\dot{B}}x^{A\dot{A}}u_{A}{}^{\dot{B}} = \tilde{z}^{2}(y+z) + 2x\tilde{z} + z - y.
$$
 (50)

Recurrance Relations in Practice

Recalling that $\phi^{(2)}$ can be used to find the electromagnetic and curvature fields respectively, we define a family of integrals:

$$
I_{m,n}=(-1)^{m+1}\int d\tilde{z}\frac{\tilde{z}^n}{[\tilde{z}^2(y+z)+2x\tilde{z}+z-y]^{3/2+m}},\qquad(51)
$$

where the scalar field in (49) is given by:

$$
\phi^{(2)}(x) = \frac{1}{\sqrt{32}\pi} \frac{1}{\sqrt{q \cdot x}} l_{0,0}.
$$
 (52)

with

$$
I_{0,0} = \frac{\sqrt{32}\pi\sqrt{q \cdot x}}{r^2} = \frac{4\pi\sqrt{y+z}}{x^2 + y^2 - z^2}.
$$
 (53)

By differentiating [\(51\)](#page-44-0) with respect to the spatial coordinates (x, y, z) , one can derive the recurrence relations

$$
\frac{\partial}{\partial x} I_{m,n} = -(3+2m)I_{m+1,n+1},
$$

$$
\left(\frac{\partial}{\partial y} + \frac{\partial}{\partial z}\right) I_{m,n} = -(3+2m)I_{m+1,n+2},
$$

$$
\left(\frac{\partial}{\partial y} - \frac{\partial}{\partial z}\right) I_{m,n} = (3+2m)I_{m+1,n},
$$
(54)

which we will use **repeatedly** in what follows.

Deriving Maxwell Fields

Spinorially translating the Fourier transform of the electromagnetic field strength tensor in [\(21\)](#page-32-2), then applying [\(9\)](#page-23-0) to obtain the Maxwell Spinor. Then, following the **same** process as done for the scalar field previously we find:

$$
\Phi_{AB}^{(2)} = -\frac{3}{4\pi} \frac{1}{(x^{01})^{1/2}} \int d\tilde{z} \frac{\tilde{\chi}_A \tilde{\chi}_B}{(\tilde{\chi}_A \tilde{\chi}_B x^{A\dot{A}} u_A \dot{B})^{5/2}} + \frac{1}{4\pi} \frac{1}{(x^{01})^{3/2}} \int d\tilde{z} \frac{u_A^A q_A \tilde{q}_B + u_B^B q_B \tilde{q}_A}{(\tilde{\chi}_A \tilde{\chi}_B x^{A\dot{A}} u_A \dot{B})^{3/2}}.
$$
(55)

Using our integral relations, we can write:

$$
\Phi_{\dot{A}\dot{B}}^{(2)} = \frac{1}{4\pi} \left[\frac{1}{(q \cdot x)^{3/2}} \begin{pmatrix} 0 & 0 \\ 0 & -l_{0,0} \end{pmatrix} - \frac{6}{\sqrt{q \cdot x}} \begin{pmatrix} l_{1,0} & l_{1,1} \\ l_{1,1} & l_{1,2} \end{pmatrix} \right],
$$
\n(56)

Deriving Maxwell Fields

$$
l_{1,0} = -\frac{32\pi (q \cdot x)^{3/2}}{3r^4},
$$

\n
$$
l_{1,1} = \frac{16\pi x (q \cdot x)}{3r^4},
$$

\n
$$
l_{1,2} = \frac{16\pi}{3} (y - z) \frac{\sqrt{q \cdot x}}{r^4} - \frac{1}{6} \frac{l_{00}}{q \cdot x}.
$$
\n(57)

Thus, the Maxwell spinor becomes

$$
\Phi_{AB}^{(2)} = -\frac{2}{r^4} \begin{pmatrix} -y - z & x \\ x & -y + z \end{pmatrix}, \tag{58}
$$

$$
\Phi_{AB}^{(2)} = 2 \frac{u_{(\dot{A}}{}^{B} r_{\dot{B})B}}{r^4},\tag{59}
$$

where

$$
r_{B\dot{B}} = (x - (u \cdot x)u) \cdot \sigma_{B\dot{B}} = \begin{pmatrix} -x & z - y \\ -z - y & x \end{pmatrix}.
$$
 (60)

Weyl Spinor Calculation

The Weyl Spinor can be obtained from the Riemann Curvature spinor via:

$$
\Psi_{ABCD} = \frac{1}{4} R_{(A \dot{X} B} \dot{X}_{C \dot{Y} D)} \dot{Y}
$$
 (61)

.

Then applying the same procedure as before to [\(22\)](#page-32-0):

$$
\Psi_{\dot{A}\dot{B}\dot{C}\dot{D}}^{(2)} = \frac{3\kappa}{2} \int d\tilde{z} \left[\frac{5\tilde{\chi}_{\dot{A}}\tilde{\chi}_{\dot{B}}\tilde{\chi}_{\dot{C}}\tilde{\chi}_{\dot{D}}}{(\tilde{\chi}_{\dot{A}}\tilde{\chi}_{\dot{B}}\kappa^{A\dot{A}}u_{A}\dot{B})^{7/2}(q \cdot x)^{1/2}} + \frac{2\tilde{\chi}_{(\dot{A}}\tilde{\chi}_{\dot{B}}(u \cdot q)\dot{C}\tilde{q}_{\dot{D})}}{(\tilde{\chi}_{\dot{A}}\tilde{\chi}_{\dot{B}}\kappa^{A\dot{A}}u_{A}\dot{B})^{5/2}(q \cdot x)^{3/2}} + \frac{(u \cdot q)_{(\dot{A}}\tilde{q}_{\dot{B}}(u \cdot q)\dot{C}\tilde{q}_{\dot{D})}}{(\tilde{\chi}_{\dot{A}}\tilde{\chi}_{\dot{B}}\kappa^{A\dot{A}}u_{A}\dot{B})^{3/2}(q \cdot x)^{5/2}} \right]
$$
(62)

Weyl Spinor Calculation

Using our recurrence relations, the Weyl spinor becomes:

$$
\Psi_{(2)}^{\dot{A}\dot{B}\dot{C}\dot{D}} = \frac{1}{6r^6} \left[u_C^{(\dot{A}} u_B^{(\dot{D})} r^{\dot{B}B} r^{\dot{C}C} - u_B^{(\dot{C}} u_A^{(\dot{D})} r^{\dot{A}A} r^{\dot{B}B} - u_C^{(\dot{B}} u_A^{(\dot{D})} r^{\dot{A}A} r^{\dot{C}C} \right. \\
\left. - \left(u_D^{\dot{C}} \left(u_A^{\dot{B}} r^{\dot{A}A} + u_B^{\dot{A}} r^{\dot{B}B} \right) + u_D^{\dot{B}} \left(u_A^{\dot{C}} r^{\dot{A}A} + u_C^{\dot{A}} r^{\dot{C}C} \right) \right. \\
\left. + u_D^{\dot{A}} \left(u_B^{\dot{C}} r^{\dot{B}B} + u_C^{\dot{B}} r^{\dot{C}C} \right) \right) r^{\dot{D}D} \right],
$$
\n(63)

The various terms combine in such a way as to yield the combination

$$
\Psi_{\dot{A}\dot{B}\dot{C}\dot{D}}^{(2)} = -2 \frac{\Phi_{(\dot{A}\dot{B}}^{(2)} \Phi_{\dot{C}\dot{D})}^{(2)}}{\phi^{(2)}},\tag{64}
$$

in terms of the electromagnetic spinor of [\(59\)](#page-47-0). Thus, a Weyl double copy does indeed hold for the **integrated Weyl spinor**, even in the presence of a non-trivial source!

Sourced Weyl Double Copy from Tensor Methods

The Weyl Double for Einstein Maxwell Gravity can also be expressed in the **language of tensors**:

$$
W^{\mu\nu\rho\sigma} = \sum_{n=1}^{m} \frac{2+n}{n\phi^{(n)}} \mathcal{P}^{\mu\nu\rho\sigma}_{\tau\lambda\eta\omega} F^{(n),\tau\lambda} F^{(n),\eta\omega}.
$$
 (65)

where $\mathcal{P}_{\tau\lambda\eta\omega}^{\mu\nu\rho\sigma}$ is some *projector* given by:

$$
\mathcal{P}^{\mu\nu\rho\sigma}_{\tau\lambda\eta\omega} = \delta^{\mu}_{\tau}\delta^{\nu}_{\lambda}\delta^{\rho}_{\eta}\delta^{\sigma}_{\omega} + \frac{1}{2}g_{\tau\eta}\delta^{\left[\mu}_{\lambda}g^{\nu\right]}\left[\rho_{\delta^{\sigma}_{\omega}}\right] + \frac{1}{6}g_{\tau\eta}g_{\lambda\omega}g^{\mu\left[\rho}g^{\sigma\right]\nu}, \quad (66)
$$

Using this language, we can prove that the Double Copy holds for **arbitrary** n:

$$
\phi^{(n)} = \frac{1}{r^n},\tag{67}
$$

- ▶ So far, we have proven that there exists a Weyl Double Copy for the case of non-vacuum solutions in Einstein-Maxwell gravity sourced by static spherically symmetric sources.
- ▶ It was also conjectured the Weyl Double Copy holds for **spinning** sources.
- ▶ Performing the **Janis Newman Shift** on our results for the Reissner-Nordstrom black hole, we obtain the Kerr-Newman Black Hole (a rotating black hole sourced by a charged disk).

- ▶ The Weyl Double Copy **relates** Petrov Vacuum Type D and N solutions in General relativity to phenomena in electromagnetism.
- ▶ We **prove** conjectures of a Weyl Double Copy for Einstein-Maxwell gravity for scalar fields in arbitrary of r^{-n} .
- ▶ Further work should probe whether we can extend our discussion of the Kerr-Newman black hole to arbitrary r^{−n}.
- ▶ It would be interesting to whether our results could have applications to **astrophysical settings**.