Replica analysis of Entanglement Properties and Conditions for Islands

Arvind Shekar

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Entanglement Entropy

2 Motivation: Why study EE?

Main motivation: Black hole information, islands

- Conditions for island
- 5 How to calculate scaling of EE on static spacetimes
- 6 Results



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von Neumann or Entanglement Entropy: Density matrix $ho = \sum_i p_i |\psi_i\rangle \langle \psi_i |$

$$S = -\sum_{i} p_i \log(p_i) = -\operatorname{Tr}(\rho \log \rho)$$

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EE quantifies entanglement: A unique feature of QM not present in CM are non-local correlations i.e., $|\psi\rangle_{AB}$ in \mathcal{H}_{AB} cannot always be factorised as $|\psi\rangle_A \otimes |\psi\rangle_B$ in $\mathcal{H}_A \otimes \mathcal{H}_B$ (also \mathcal{H}_{AB} does not always factorise).

$$|\psi\rangle = \sqrt{p_k} \sum_k |\psi_k\rangle_A \otimes |\psi_k\rangle_B \xrightarrow{\rho_A = \sum_k p_k |\psi_k\rangle_A \langle\psi_k|} \rho_B = \sum_k p_k |\psi_k\rangle_B \langle\psi_k|$$

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Entangled pure state in \mathcal{H}_{AB} imply mixed states in \mathcal{H}_A and \mathcal{H}_B

 $S_{AB} = 0$, $S_A = S_B$ and is non zero. Hence captures both non local entanglement and local correlations (if mixed in \mathcal{H}_{AB}). Also notice $S(U\rho U^{-1}) = S(\rho)$.

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Quantifies correlation: $S(AB) \leq S(A) + S(B)$: equal if $\rho_{AB} = \rho_A \otimes \rho_B$. Difference I(A:B) quantifies correlation. Alternative to $\langle \mathcal{O}_A \mathcal{O}_B \rangle - \langle \mathcal{O}_A \rangle \langle \mathcal{O}_B \rangle$. EE and correlators (*Moitra, Sensarma PRB 108 (2023)*). $S(AB) < S(A) \implies$ entanglement.

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Condensed matter: Understanding many body quantum systems. Thermalization, Many Body Localization (*A.Pal, and D. A. Huse, PRL (2010)*), Measurement Induced phase transitions (*B Skinner, J. Puhmann, A.Nahum, PRX (2019)*)...

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AdS/CFT - Ryu Takayanagi:

$$S_A = \min_{R \sim A} \frac{Area(R)}{4G}$$

Ryu, Takayanagi PRL 96 (2006)



Main motivation: Black hole information and page curve

$$S_{gen} = rac{ ext{Area of horizon}}{4G} + S_{rad}$$

Start with black hole in pure state. As it evaporates it loses area $\frac{dM}{dT} \sim \sigma AT^4 \sim \frac{1}{M^2}$. After a time $S_{rad} > S_{Area}$ and there are not enough dof in BH to purify radiation. Non unitary!

Unitary evolution follows page curve

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Quantum extremal surface and islands

$$S = \min \operatorname{ext} (S_{QFT} + S_{area})$$

(Wall, Engelhardt (2014))





Island

Important in black hole evaporation. We need to minimize generalised entropy for candidate subregions. The presence of a QES not at asymptotic infinity implies that the complementary subregion has a region disconnected from the asymptotic boundary - **Island**. (*Penington, Shenker, Stanford, Yang; Almheri, Mahajan, Maldacena, Zhao...*)

Correlation between island and region between QES and asymptotic boundary brings down EE at later times and preserves unitarity.

This is a statement about how the EE of the black hole/ radiation (provided we started with a pure black hole) must behave to preserve unitarity. What does this translate to for the QFT spectrum, correlation functions of the radiation?



$$ds_d^2 = -g_{tt}(r)dt^2 + g_{rr}(r)dr^2 + r^2 d\Omega_{d-2}^2$$

subregion: $r_1 \leq r \leq r_2 \ \forall \ \Omega$, t constant.





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 S_{QFT} has information about the QFT spectrum and S_{area} is purely geometric. Hence this imposes constraints on the QFT spectrum in terms of background metric parameters.

Replica trick

Gap in literature: Calculating S_{QFT} and its r-scaling is hard! Known results in

- 1+1 dim CFTs (Holzhey, Larsen, Wilczek (1994); Calabrese, Cardy (2004))
- 2+1 dim (Casini, Huerta)
- Universal features, results restricted to CFTs in flat, vacuum state
- Non-linear differential equation for RT surface even in 2+1 dim.

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We need EE on static background

$$egin{aligned} S_{QFT} &= \lim_{q o 1} S_q = \lim_{q o 1} rac{1}{1-q} \log(\mathit{Tr}_A
ho_A^q) \ &= \lim_{q o 1} rac{1}{1-q} \left(\mathcal{W}_q - q \mathcal{W}_1
ight) \end{aligned}$$



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Scaling of EE on static background

We need scaling of EE on static background: Restrict to CFTs



$$\delta(\Delta r) \frac{\delta}{\delta(\Delta r)} S_{QFT} = -\Delta_{\Lambda} S_{QFT}$$

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$$\delta(\Delta r) \frac{\delta}{\delta(\Delta r)} S_{QFT} = -\Delta_{\Lambda} S_{QFT}$$

$$= -\int d^d x \left(\delta_{\Lambda} g_{\mu\nu} \frac{\delta}{\delta g_{\mu\nu}} + \sigma(x) \beta^i(\lambda) \frac{\delta}{\delta \lambda^i(x)} \right) \left(\lim_{q \to 1} \frac{1}{(1-q)} (W_q - q W_1) \right)$$

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Replica metric $g^{(q)}_{\mu u}$

Consider spatial entangling region $x = x_0$ on a $\tau = \tau_0$ slice.



Around the entangling boundary, replica boundary conditions imply a conical singularity. Metric remains unaffected elsewhere. Therefore it is enough to consider the metric expansion around the entangling boundary.

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We use coordinates
$$x - x_0 = \rho \cos(\psi)$$
, $\tau - \tau_0 = \rho \sin(\psi)$; $\rho \in [0, \infty)$, $\psi \in [0, 2\pi q)$

$$ds_{M_q}^2 \text{ around boundary} = g_{\mu\nu}^q dx^\mu dx^\nu |_{\text{around }\Sigma}$$

= $U(\rho, \mathbf{a})d\rho^2 + \rho^2 d\psi^2 + (\gamma_{ij} + 2\rho^\rho c^{1-\rho} \cos(\psi)K_{1ij} + 2\rho^\rho c^{1-\rho} \sin(\psi)K_{2ij})dy^i dy^j$
+ $A_i \varepsilon_{ac} x^a dx^c dy^i + \mathcal{O}\left(x^2\right)$

 $U|_{
ho=0} o q^2$, $U|_{
ho>>a} o 1$

Replica Stress Tensor $\langle T^{\mu\nu} \rangle_q$ and (q-1) expansion

The EE is equal to the difference $W_q - qW_1$ at O(q - 1). We consider a systematic O(q - 1) expansion which reveals the trivial terms in EE scaling. This picks up contributions only due to the localised conical singularity(ies).

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$$egin{aligned} g^{(q)}_{\mu
u} &= g_{\mu
u} + (q-1) \Big(g^{[1]}_{\mu
u} &= (\partial_q g^{(q)}_{\mu
u})|_{q=1} \Big) + \mathcal{O}\left((q-1)^2
ight), \ \langle T^{\mu
u}
angle_q &= \langle T^{\mu
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angle + (q-1) \, \langle T^{\mu
u}
angle^{[1]} + \mathcal{O}\left((q-1)^2
ight) \end{aligned}$$

Replica stress tensor

 $\nabla_{\mu}^{(q)} \langle T_{\mu\nu} \rangle_{q} = 0$ $\langle T_{\mu\nu}^{\mu} \rangle_{q} = \mathcal{A}[g_{\mu\nu}^{(q)}]$ The (d+1) equations completely determine $\langle T_{\mu\nu} \rangle^{[1]}$ in terms of $\langle T_{\mu\nu} \rangle$ for states with same symmetry as the static background.

EE scaling: General d, subregion, state

$$\{-\alpha(r_2-r_1)\}\frac{\partial S_{QFT}}{\partial r_1} = \lim_{q \to 1} \left(\frac{1}{2} \int_{M_q} d^d x \sqrt{g} \left(\delta_{\Delta r} g_{\mu\nu}\right) \langle T^{\mu\nu} \rangle^{[1]}\right)$$

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Vacuum and thermal state (odd dim), planar boundary in flat space

$$\langle T^{\mu}_{\ \nu} \rangle^{vacuum}_{q, a=0} = rac{C_d(q)}{
ho^d} \operatorname{diag}((1-d), 1, 1, 1, ..., 1)$$

$$\Delta r \frac{\partial S_{d \text{ dim}}^{\text{finite, thermal}}}{\partial \Delta r} = \pi \text{Vol}(M_{d-2}) \left\{ \frac{C_d^{[1]} 2^{d-2}}{(\Delta r)^{d-2}} + \frac{p}{4} (\Delta r)^2 \right\}$$

We can now use this to get explicit conditions for islands.

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Since scaling of S_{QFT} with region depends on $\langle T_{\mu\nu} \rangle$ for QFTs on static backgrounds, therefore demanding island imposes constraints on $\langle T_{\mu\nu} \rangle$.

Examples: QES in flat!

Thermal state: Odd dimension

$$C_d^{[1]}\left(rac{2^{d-2}}{(\Delta x)^{d-2}}
ight)+rac{p}{4}(\Delta x)^2=0\implies x_2=x_1+2\left(rac{-C_d^{[1]}}{p}
ight)^{1/d}$$

Vacuum: Spherical boundary - 4 dim

$$\frac{4a}{\Delta r} + 8\pi r_2 = 0 \implies r_2 = r_1 \pm \frac{1}{2} \left(r_1^2 - \frac{2a}{\pi} \right)^{1/2}$$

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• Use scaling result to explicitly calculate EE scaling for more complicated states and subregions (say with extrinsic curvature) on flat as well as more general static spacetimes.

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- Extend to stationary spacetimes. Lack of spherical symmetry implies more complicated subregions would reveal more dynamics. Defining equations of the replica stress tensor will be related to static case by corresponding generating transformations.

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- EE in Carrollian and celestial CFTs. Replica trick and scaling can be used on such backgrounds too. Defining equations for replica stress tensor must respect the symmetries of the degenerate metric and direction i.e., the Carroll structure. Does this have a holographic proposal?

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- EE in Carrollian and celestial CFTs. Replica trick and scaling can be used on such backgrounds too. Defining equations for replica stress tensor must respect the symmetries of the degenerate metric and direction i.e., the Carroll structure. Does this have a holographic proposal?
- Explicit conditions on QFT spectrum for islands by making use of the scaling result. Interpret these conditions in the holographic dual to black holes.
- Use our result to test and refine the Quantum focussing conjecture.

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Thank you for your attention!

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