

Dark matter: then and now

Christopher McCabe

IPPP 2011 - 2014

Then: circa 2011 - 2014



July 2012



July 2012



Status of Standard Model Higgs searches in ATLAS

Using the full datasets recorded in 2011 at $\sqrt{s}=7$ TeV and 2012 at $\sqrt{s}=8$ TeV: up to 10.7 fb^{-1}

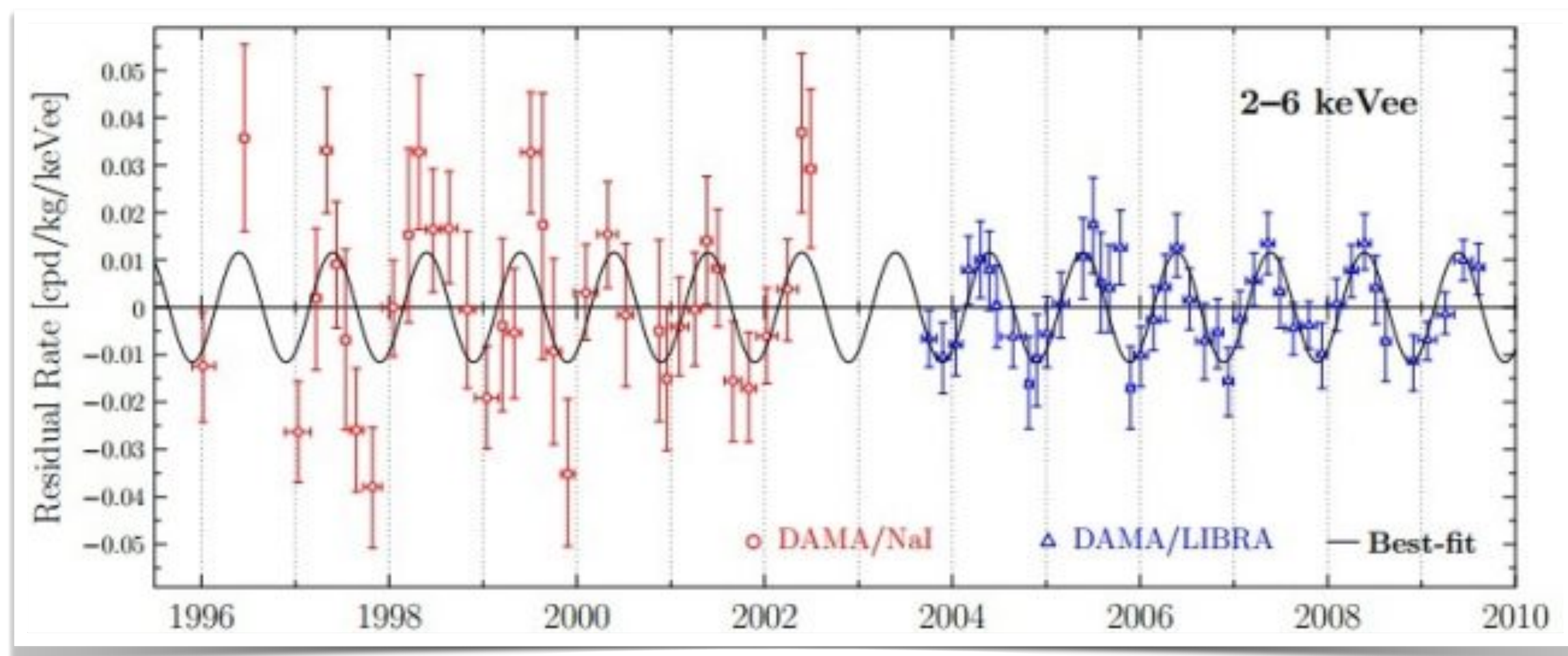
Fabiola Gianotti (CERN), representing the ATLAS Collaboration



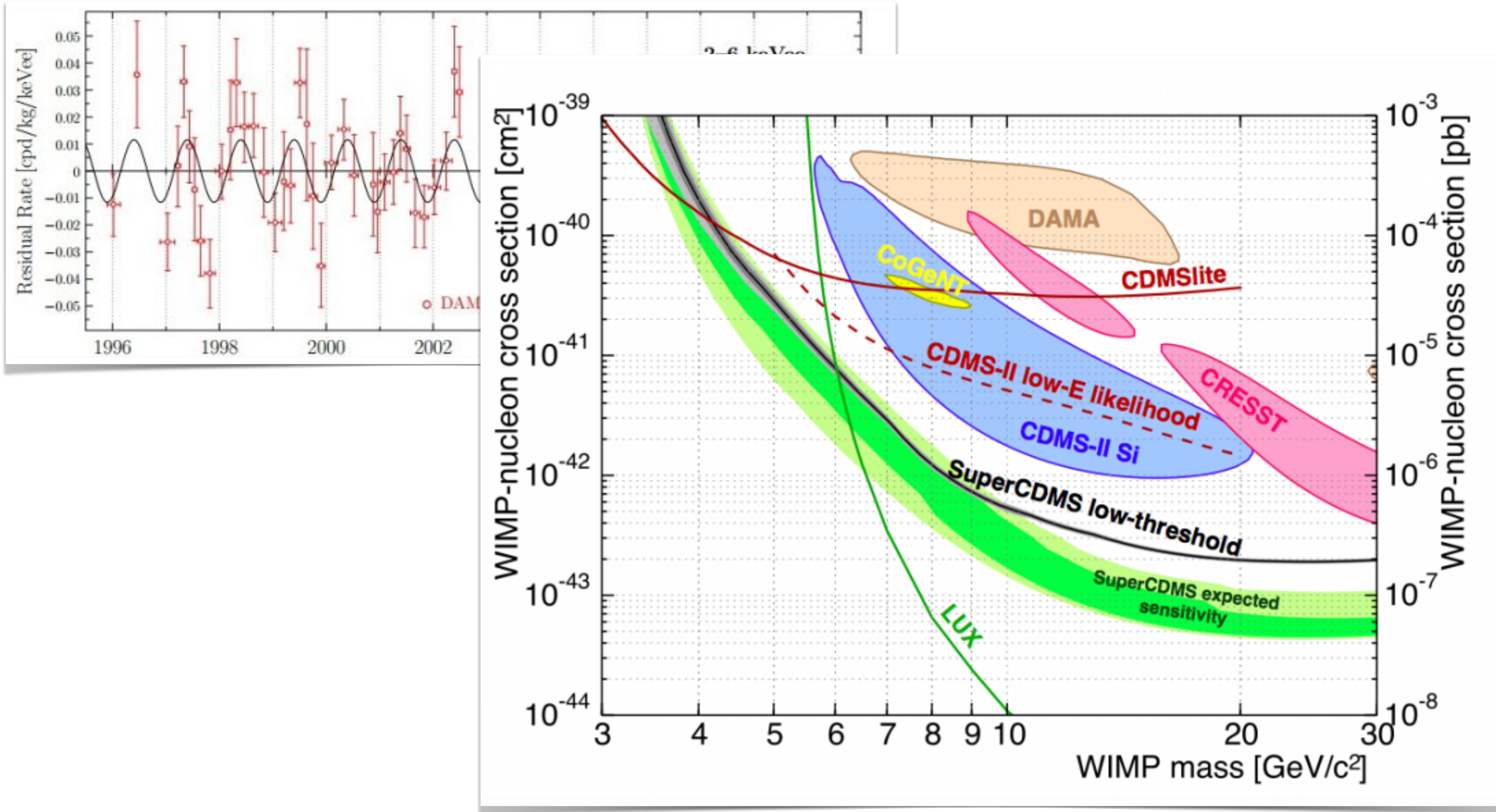


Chasing dark
matter signals

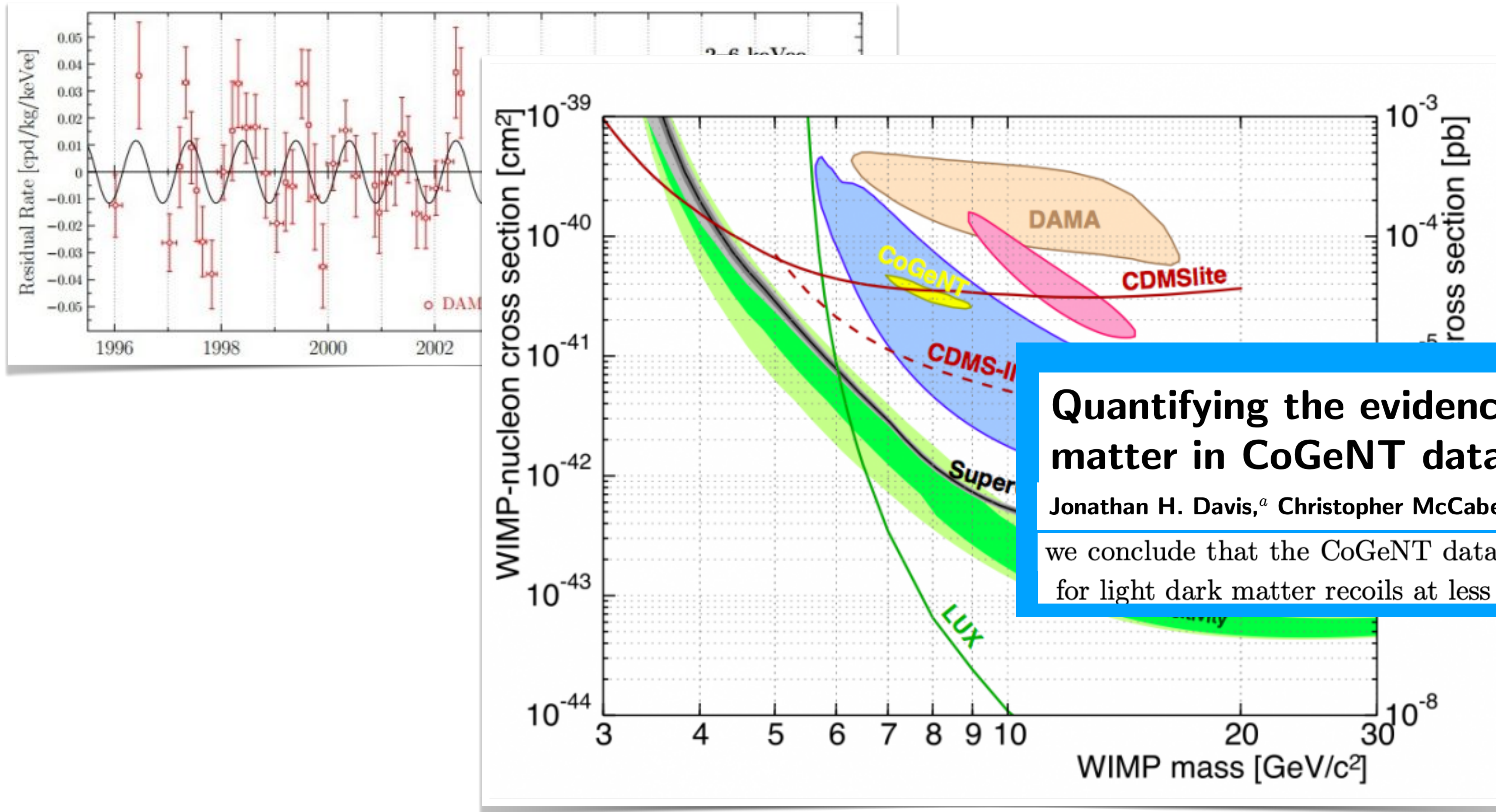
Direct detection anomalies



Direct detection anomalies



Direct detection anomalies



Quantifying the evidence for dark matter in CoGeNT data

Jonathan H. Davis,^a Christopher McCabe,^a and Céline Boehm^{a,b}

we conclude that the CoGeNT data show a preference for light dark matter recoils at less than 1σ .

arXiv: 1405.0495
JCAP

Indirect detection anomalies



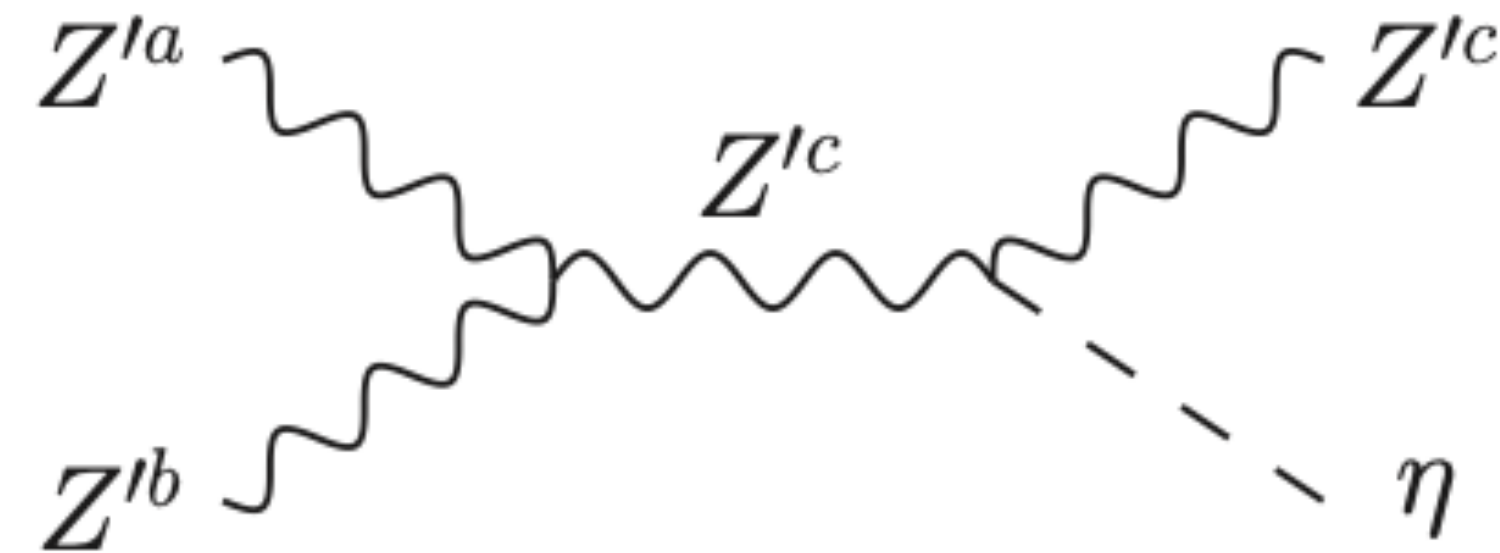
Indirect detection anomalies



A weighty interpretation of the Galactic Centre excess

Phys. ReV. D title: Interpretation of the Galactic Center excess of gamma rays with heavier dark matter particles

Céline Boehm,^{1,2} Matthew J. Dolan,³ and Christopher McCabe¹



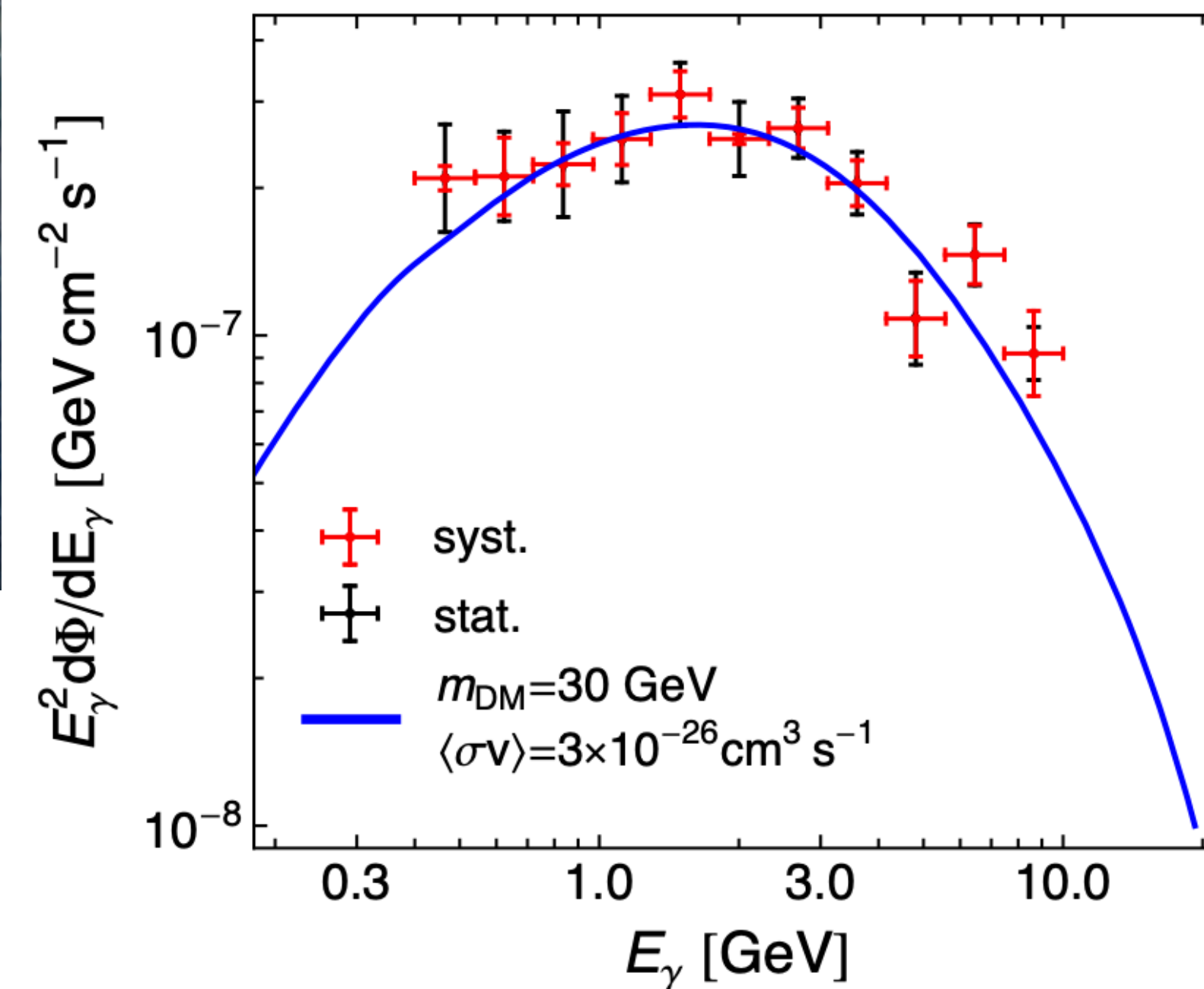
arXiv: 1404.4977, PRD

Indirect detection anomalies

Extended gamma-ray emission from Coy Dark Matter

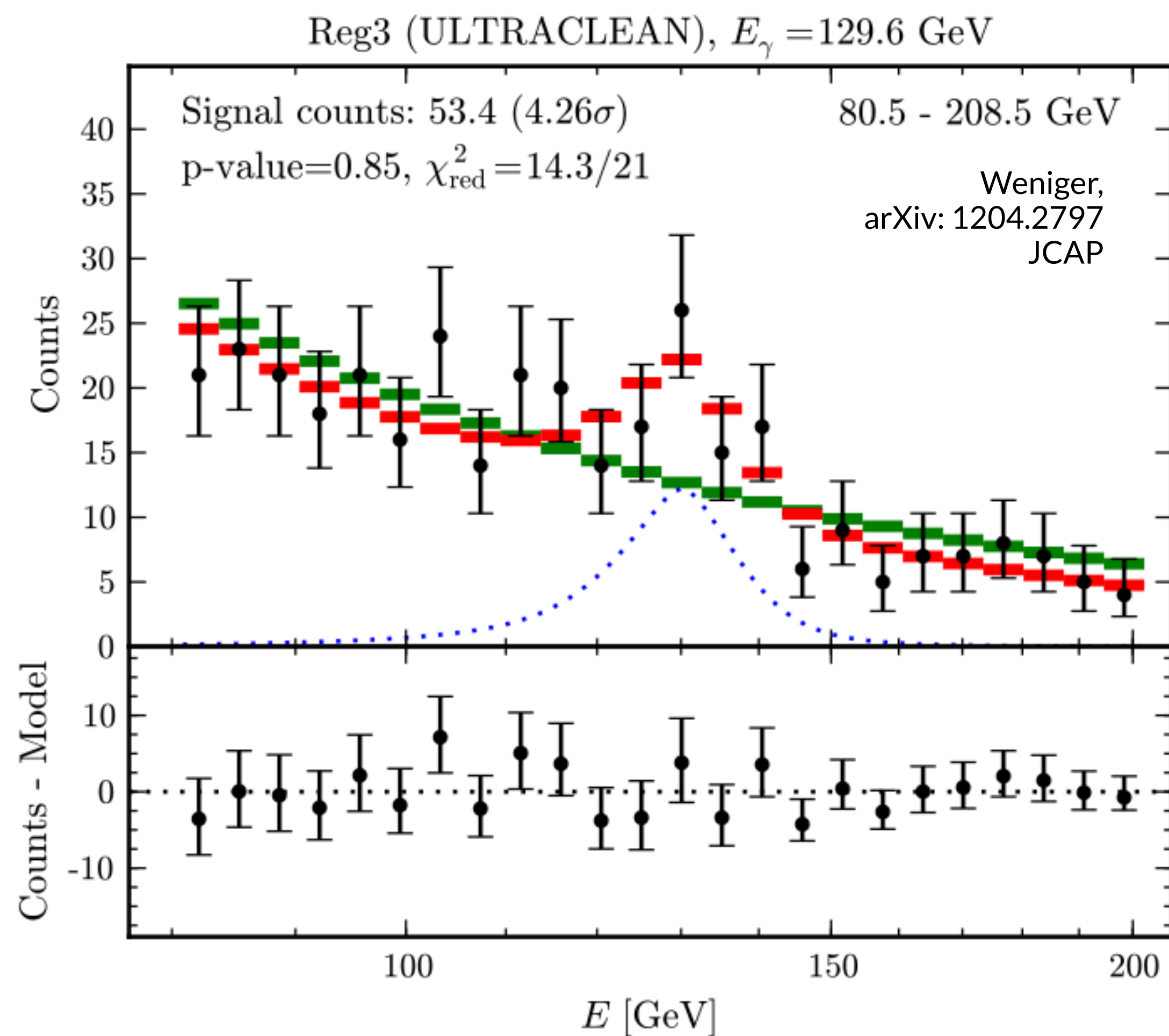
Céline Boehm,^{1,2} Matthew J. Dolan,³ Christopher McCabe,¹ Michael Spannowsky,¹ and Chris J. Wallace¹

Galactic Centre Excess (of gamma-rays)

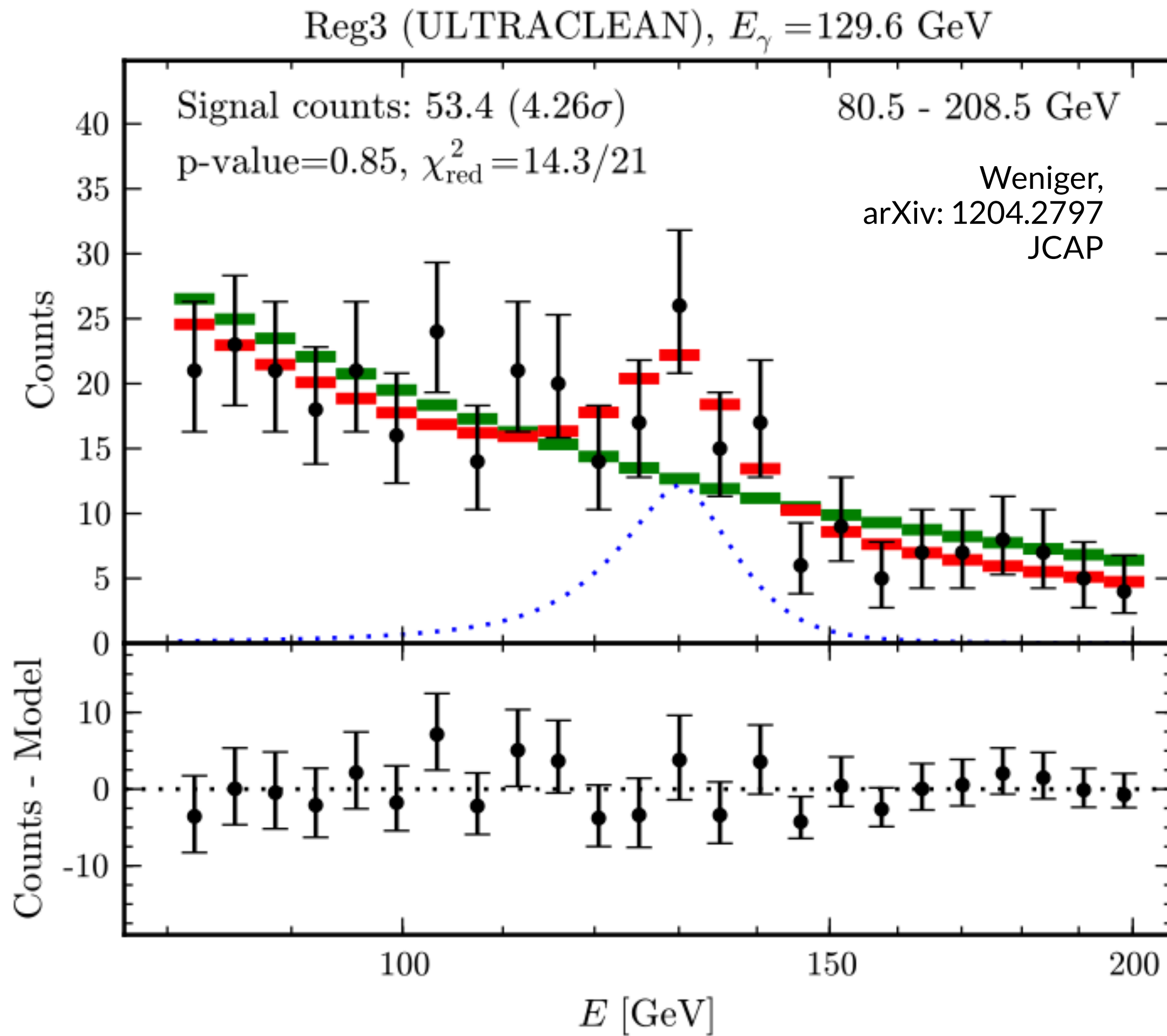


arXiv: 1401.6458, PRD

Indirect detection anomalies

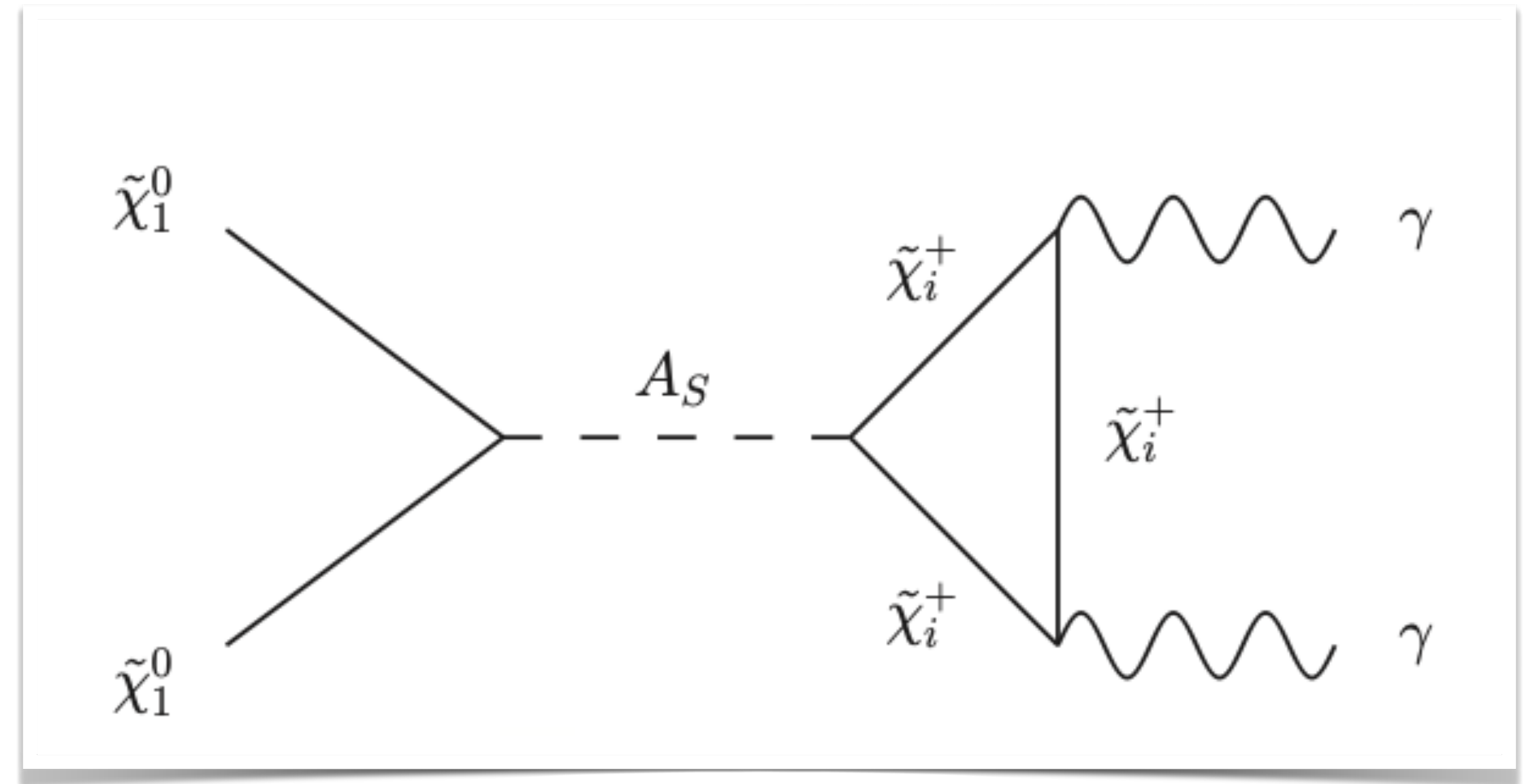


Indirect detection anomalies



Neutralino dark matter and the Fermi gamma-ray lines

Guillaume Chalons,^{1,*} Matthew J. Dolan,^{2,†} and Christopher McCabe^{2,‡}



arXiv: 1211.5154, JCAP

Now





**Chasing dark
matter signals:
old and new frontiers**

Cosmic Frontier's recommendation

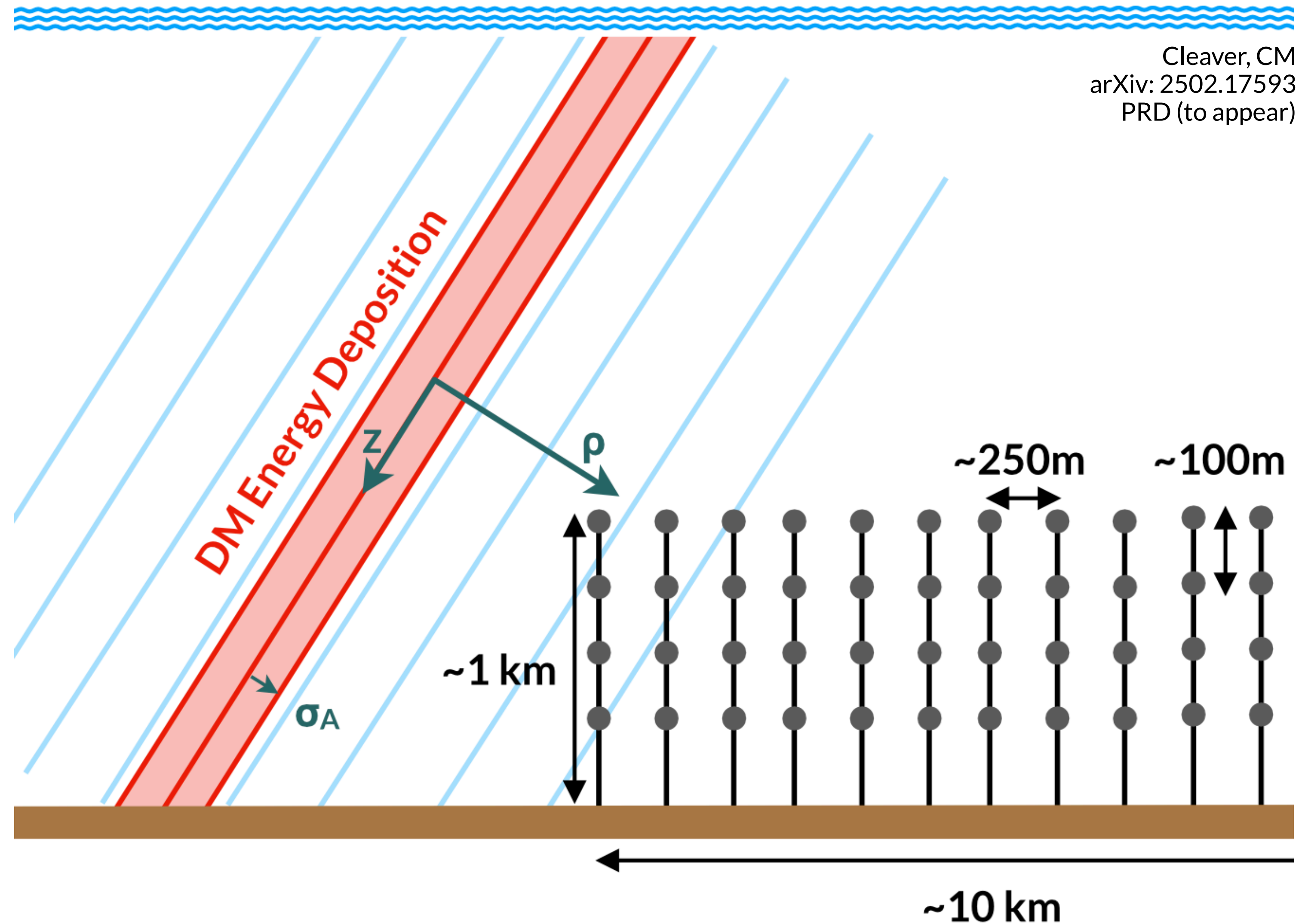
Delve deep (cover high priority targets e.g., WIMPs)

Search wide (explore as much DM parameter space as possible)

Cooley et al,
arXiv:2209.07426

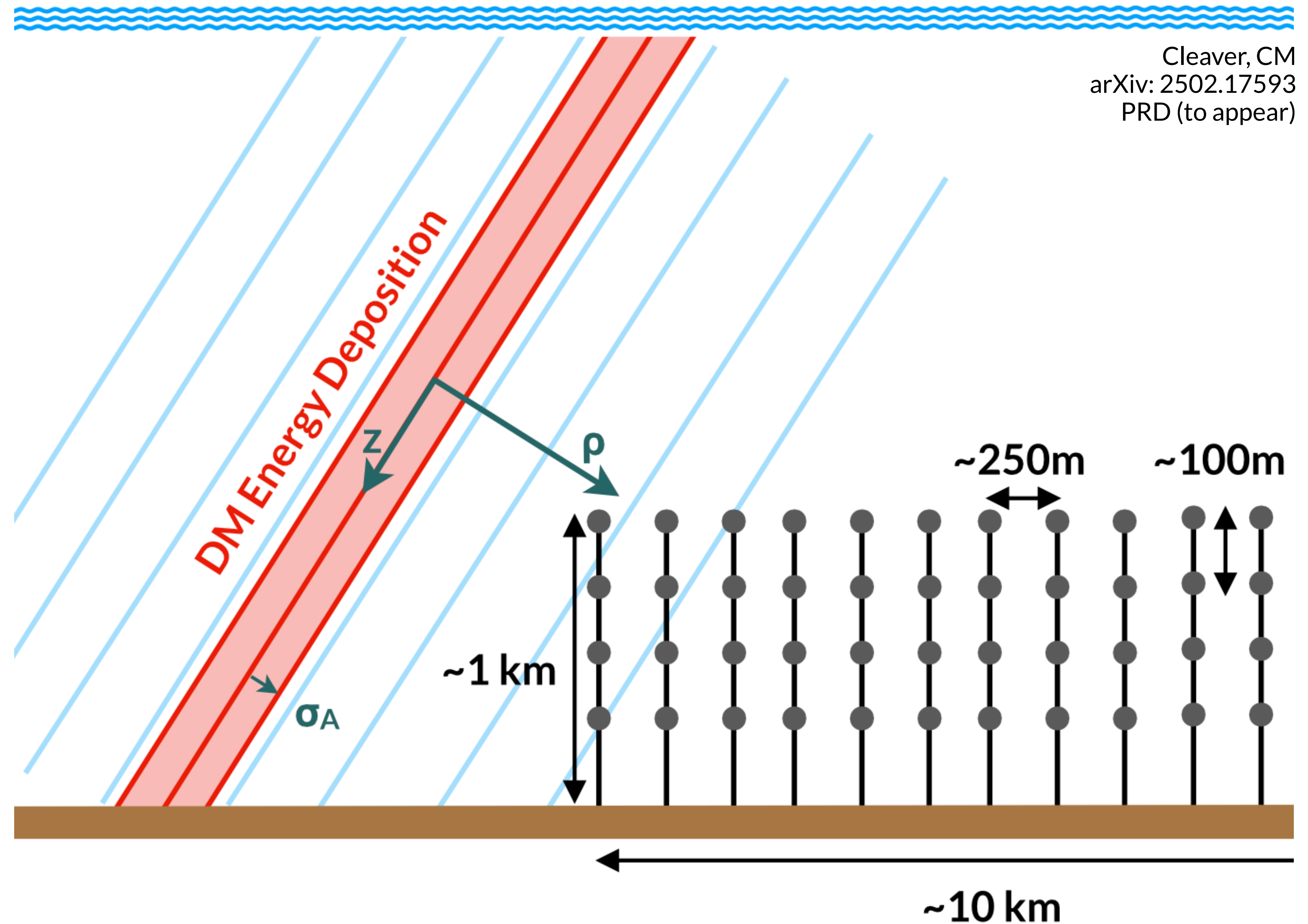


Dark matter: Macroscopic masses and cross-sections



Dark matter: Macroscopic masses and cross-sections

Lesson from IPPP:
Quantifying backgrounds
is an important as
quantifying the signal



AMBIENT NOISE IN THE SEA



DISTRIBUTION STATEMENT A
Approved for Public Release
Distribution Unlimited

NAVAL SEA SYSTEMS COMMAND

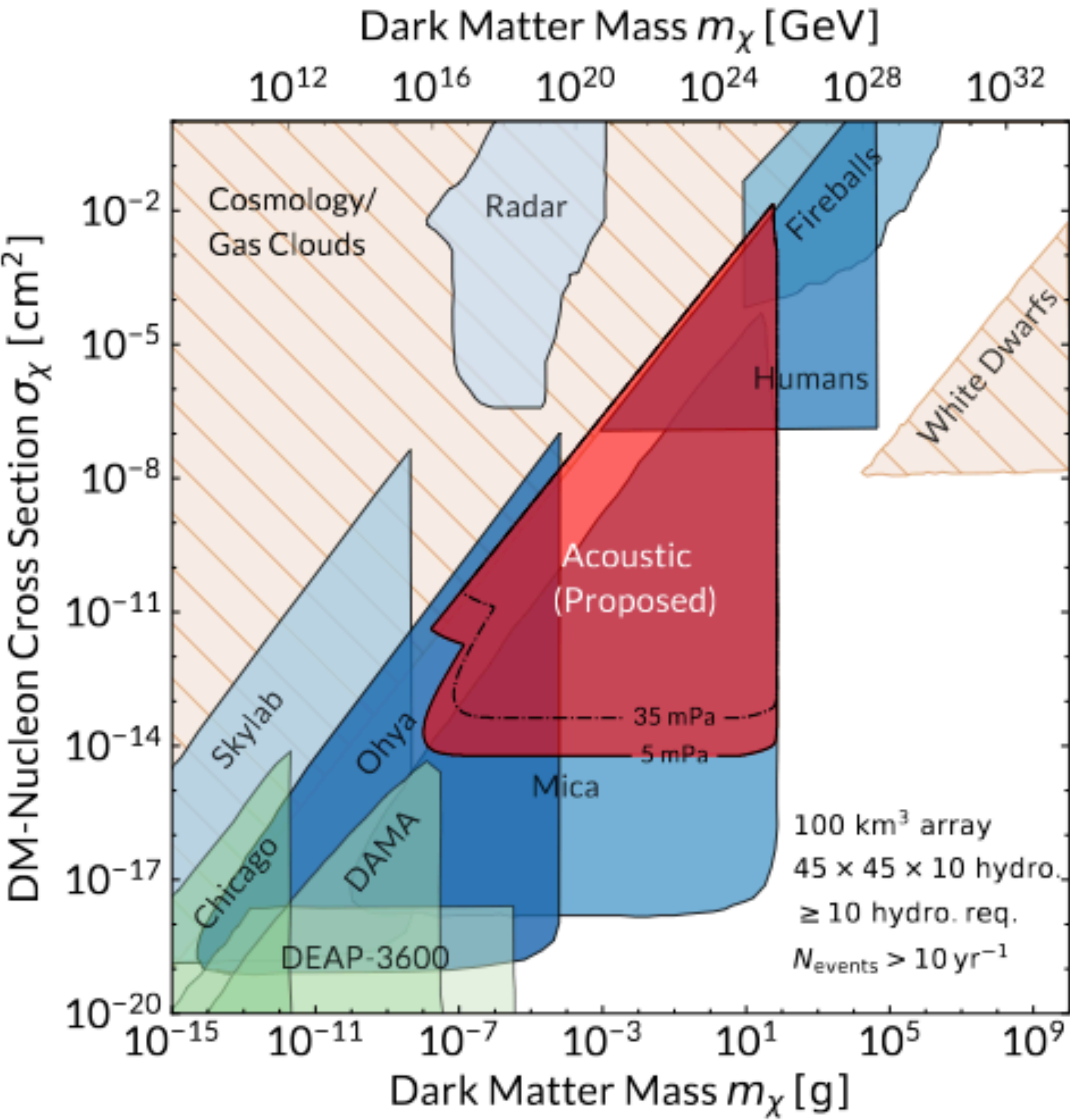
UNCLASSIFIED UNLIMITED
QinetiQ Proprietary



SEA 7 Technical report:
Underwater ambient noise
(Non-technical summary)

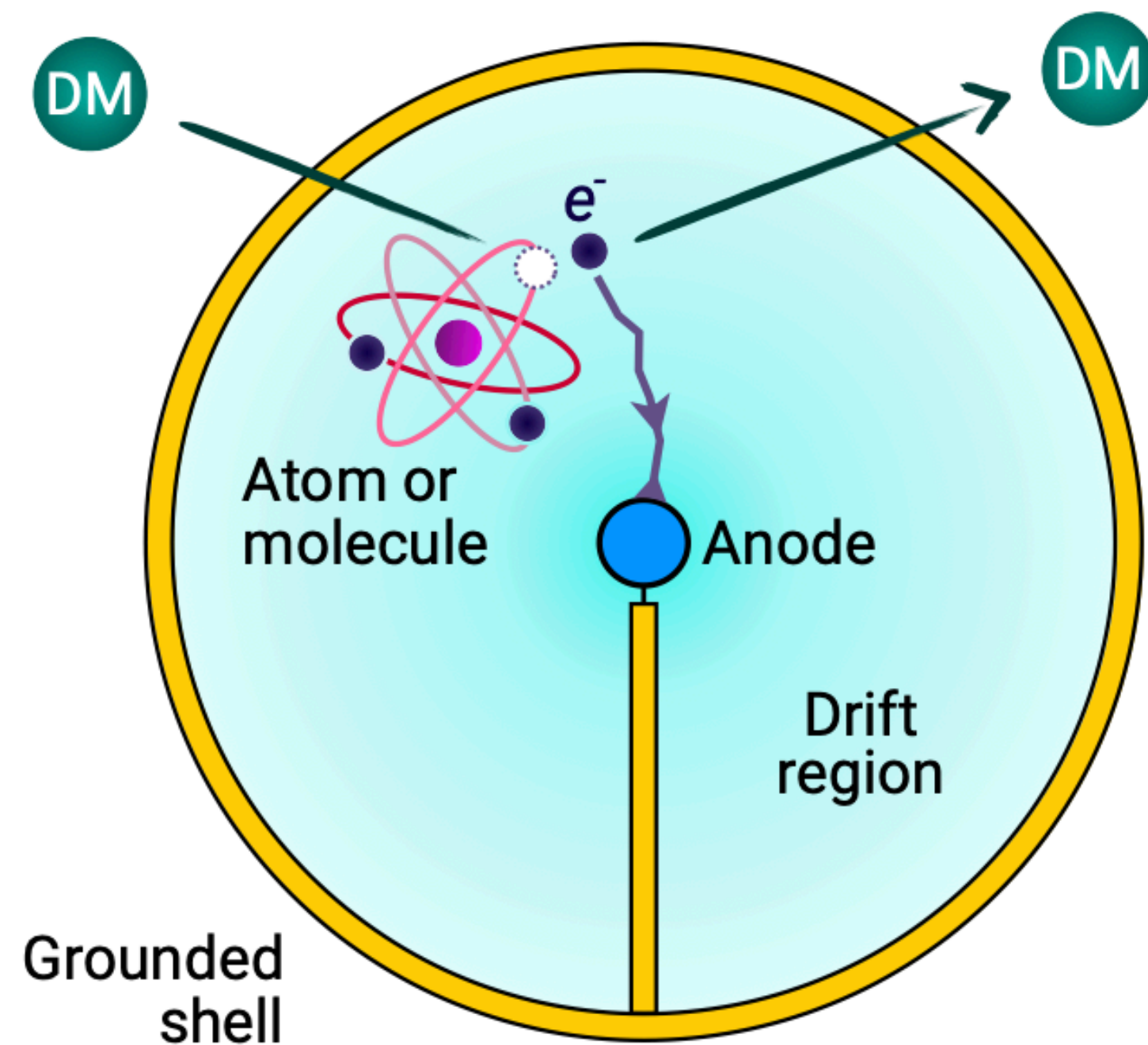
Eavesdropping on whales

*The acoustic detection of ultra high energy neutrinos in the
Mediterranean Sea*

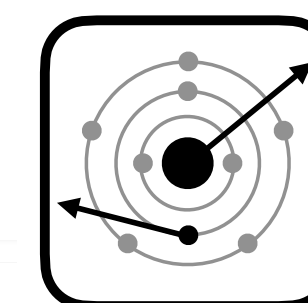


Cleaver, CM
arXiv: 2502.17593
PRD (to appear)

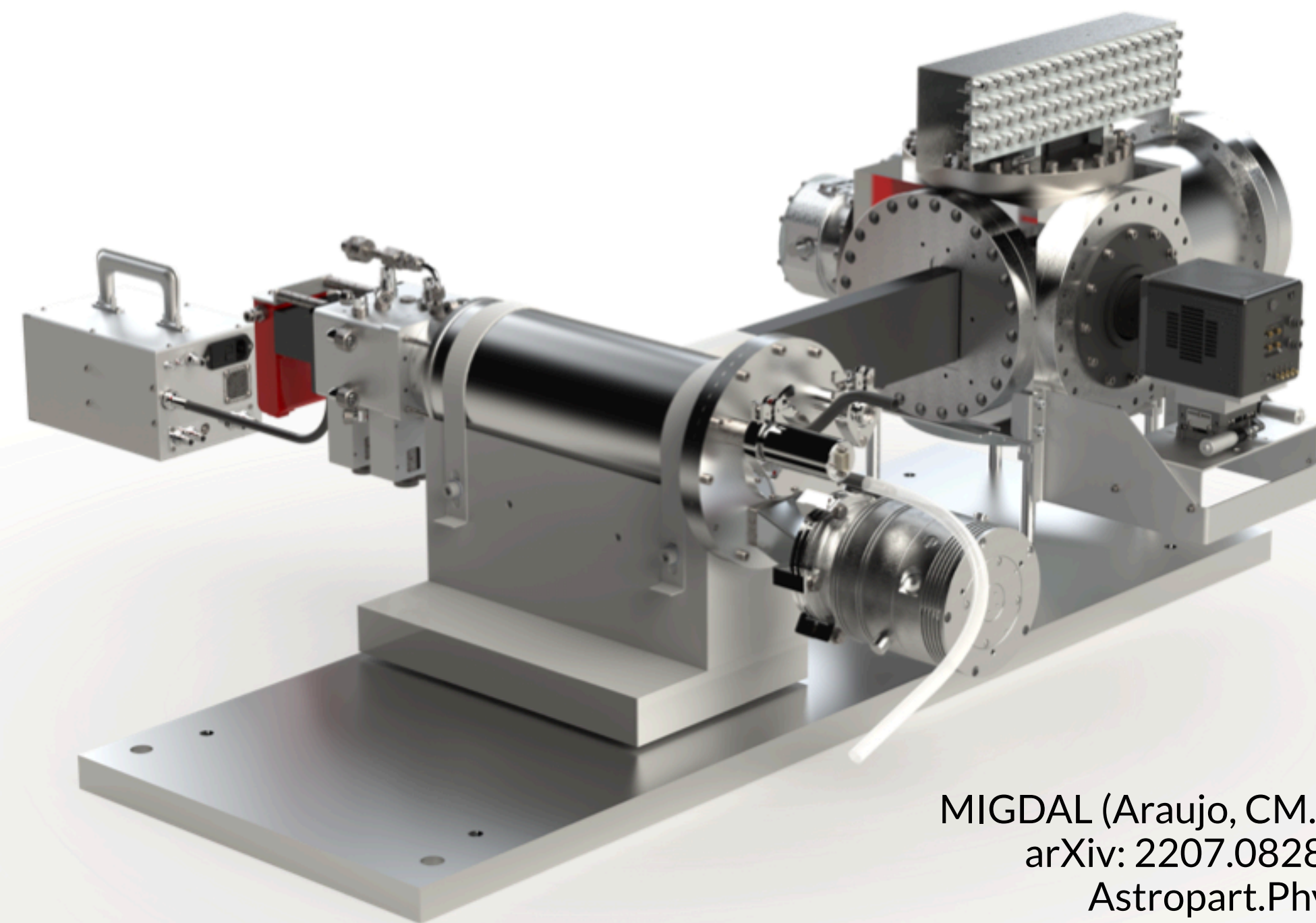
Dark matter: MeV - GeV scale



DarkSPHERE (Balogh, Hamaide, CM...)
arXiv: 2301.05183, PRD

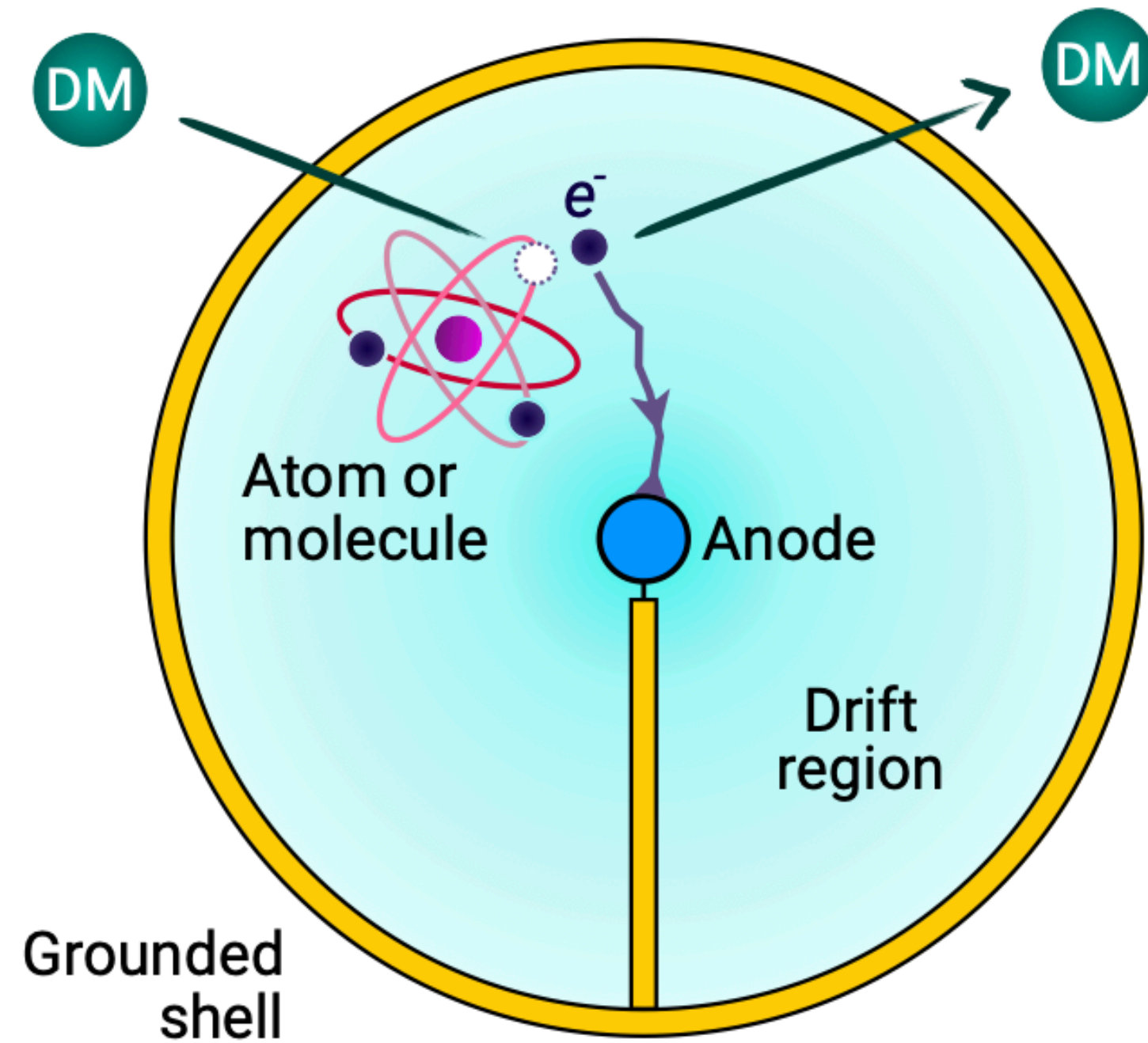


MIGDAL
Migdal In Galactic Dark mAtter expLoration



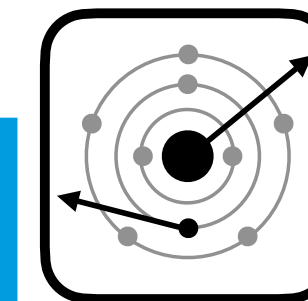
MIGDAL (Araujo, CM...)
arXiv: 2207.08284
Astropart.Phys

Dark matter: MeV - GeV scale

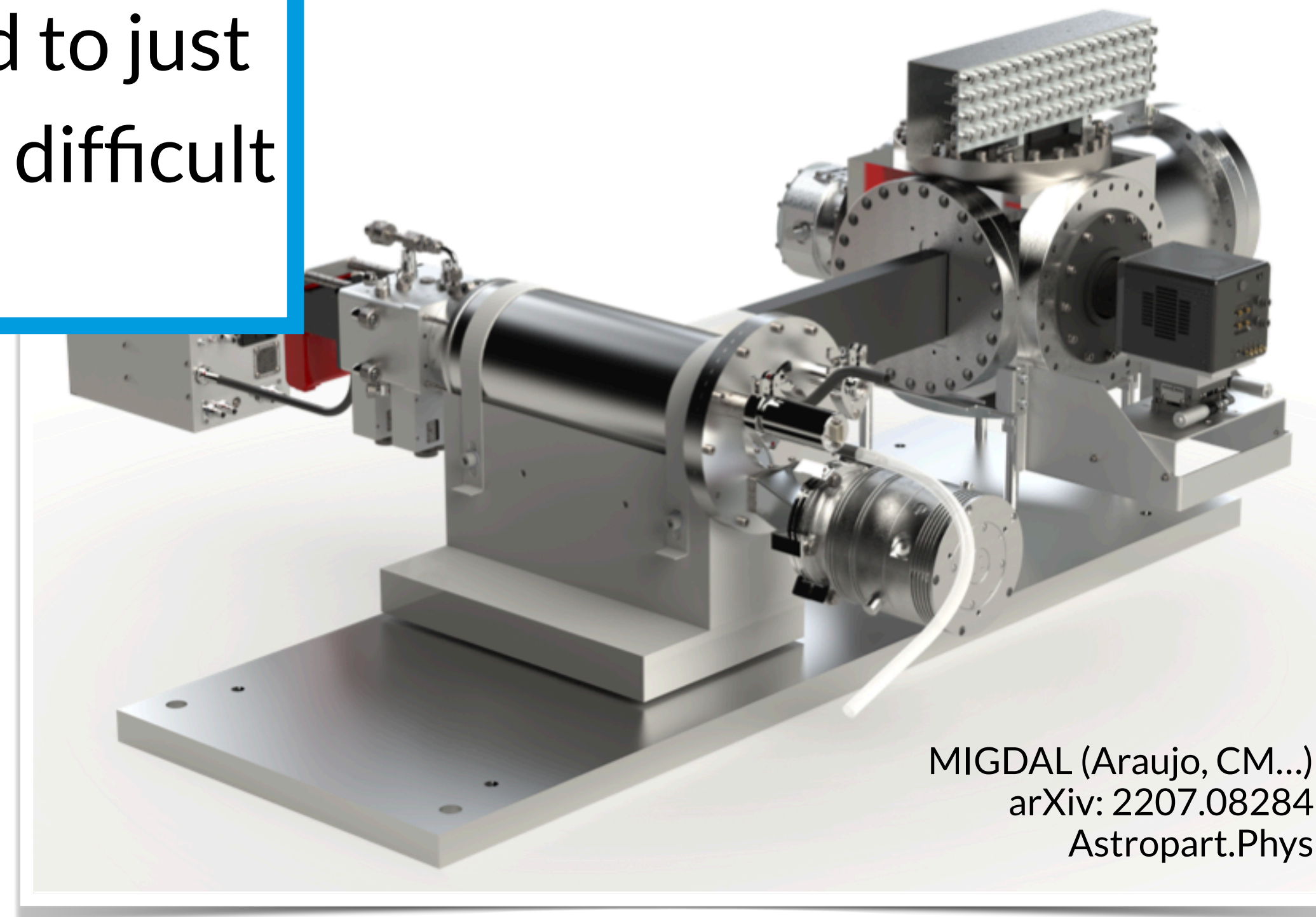


DarkSPHERE (Balogh, CM...)
arXiv: 2301.05183, PRD

Lesson from IPPP:
Sometimes you need to just
get on and tackle the difficult
calculations

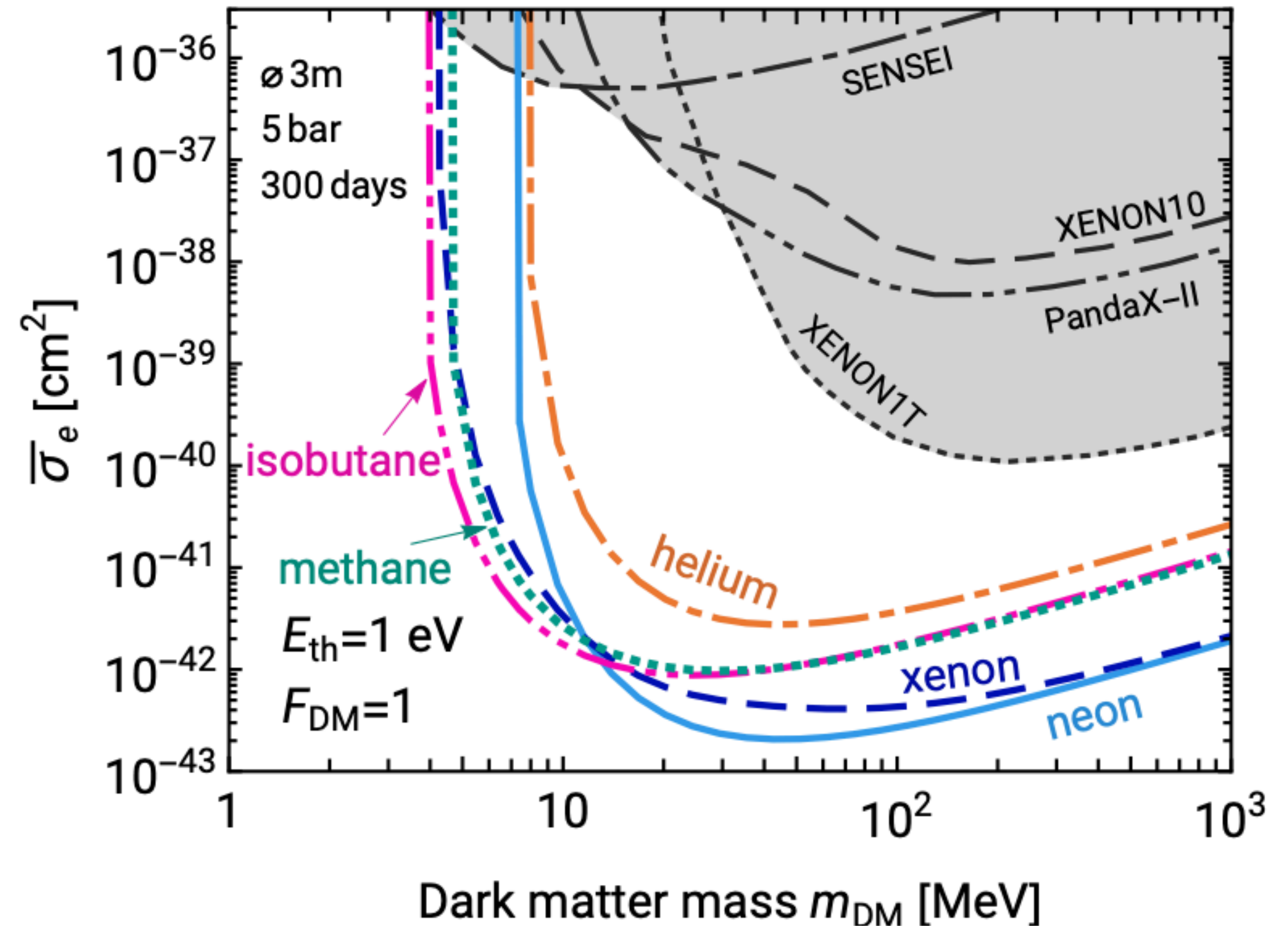
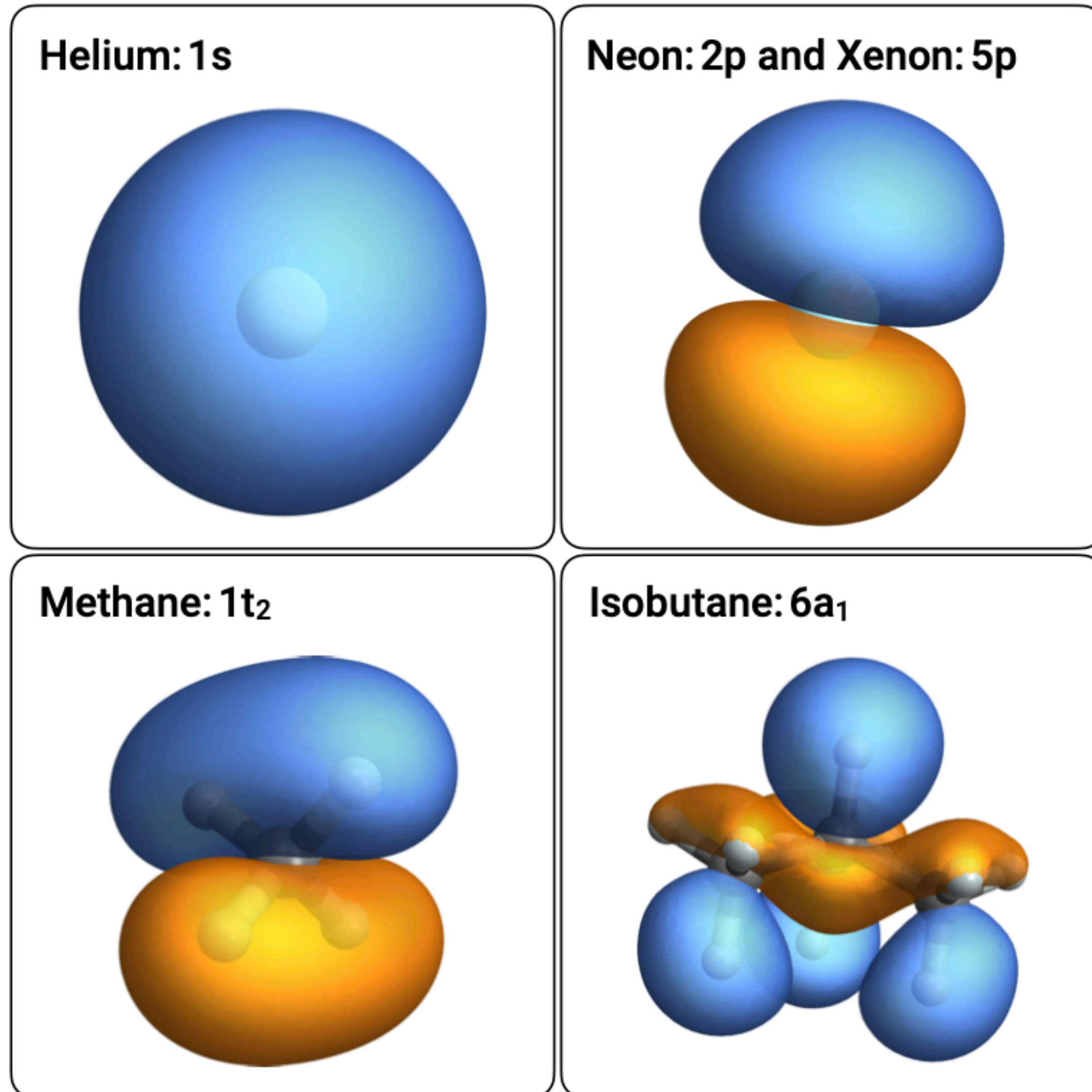


MIGDAL
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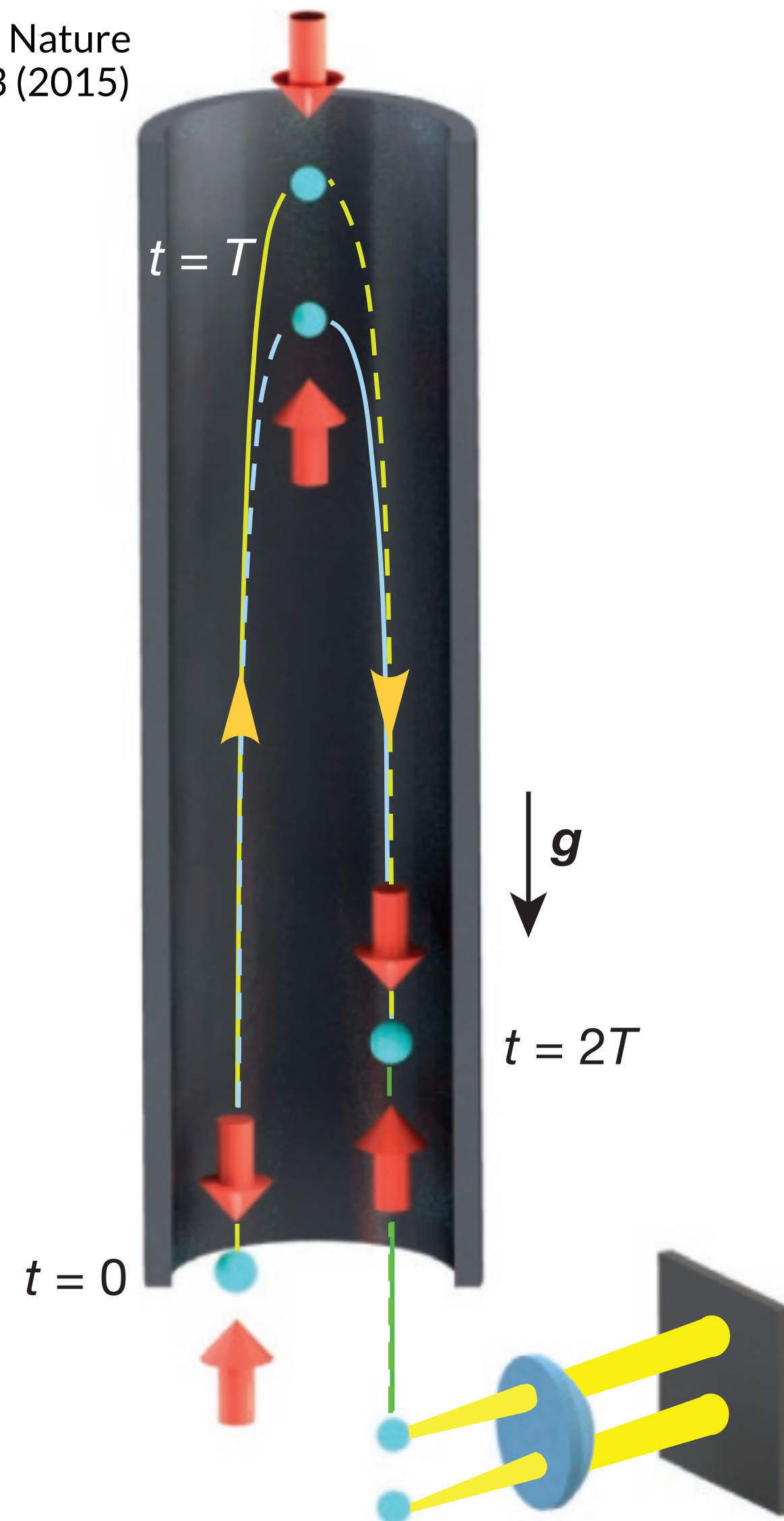
MIGDAL (Araujo, CM...)
arXiv: 2207.08284
Astropart.Phys

Dark matter: MeV - GeV scale



New frontiers with atom interferometry: neV scale

Kovachy et al, Nature
528, 530–533 (2015)

The logo for AION, featuring the word "AION" in a bold, black, sans-serif font. The letter "O" is stylized with two blue dots above and below it, and a series of concentric blue circles to its right, resembling a ripple effect or a signal.

AION (Badurina, CM, et al)
arXiv: 1911.11755, JCAP
arXiv: 2504.09158
arXiv: 2508.03491

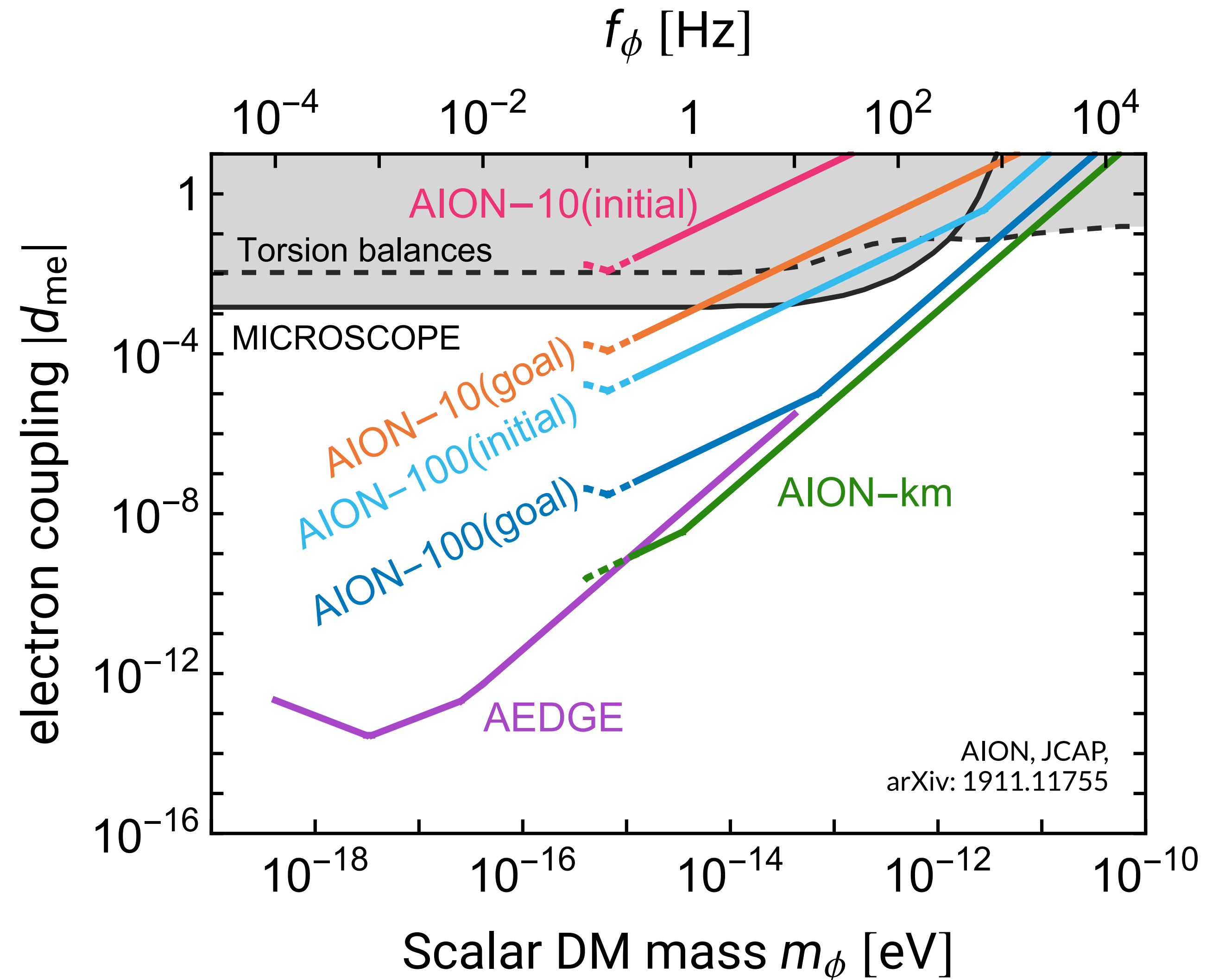
Initial puzzle: what should we do with curves below 1 Hz?

Simply cut-off the lines?

Almost certainly too conservative

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is an important as
quantifying the signal

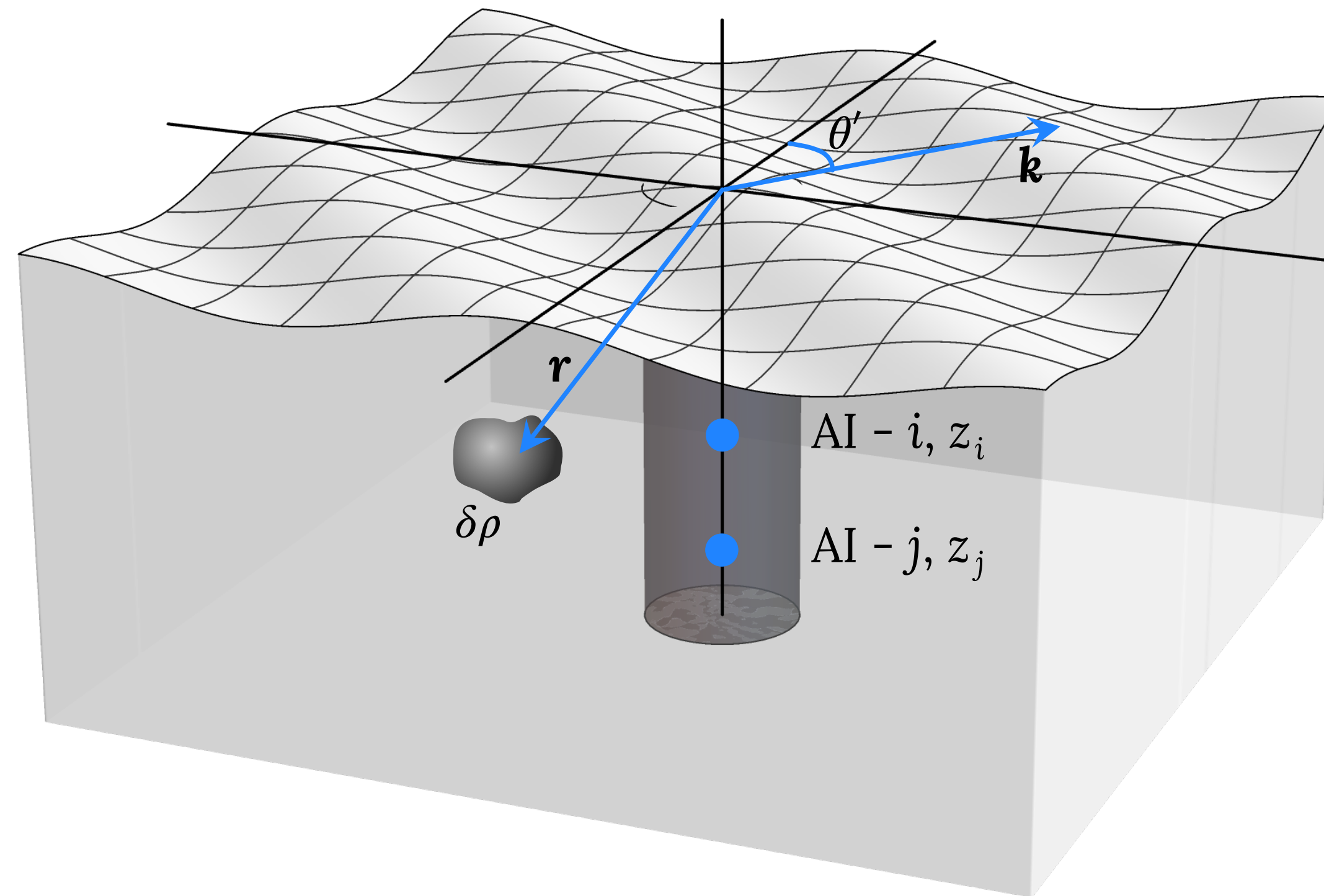
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Factors that could limit sub-Hz regime

Seismic effects

surface waves, body waves, ...



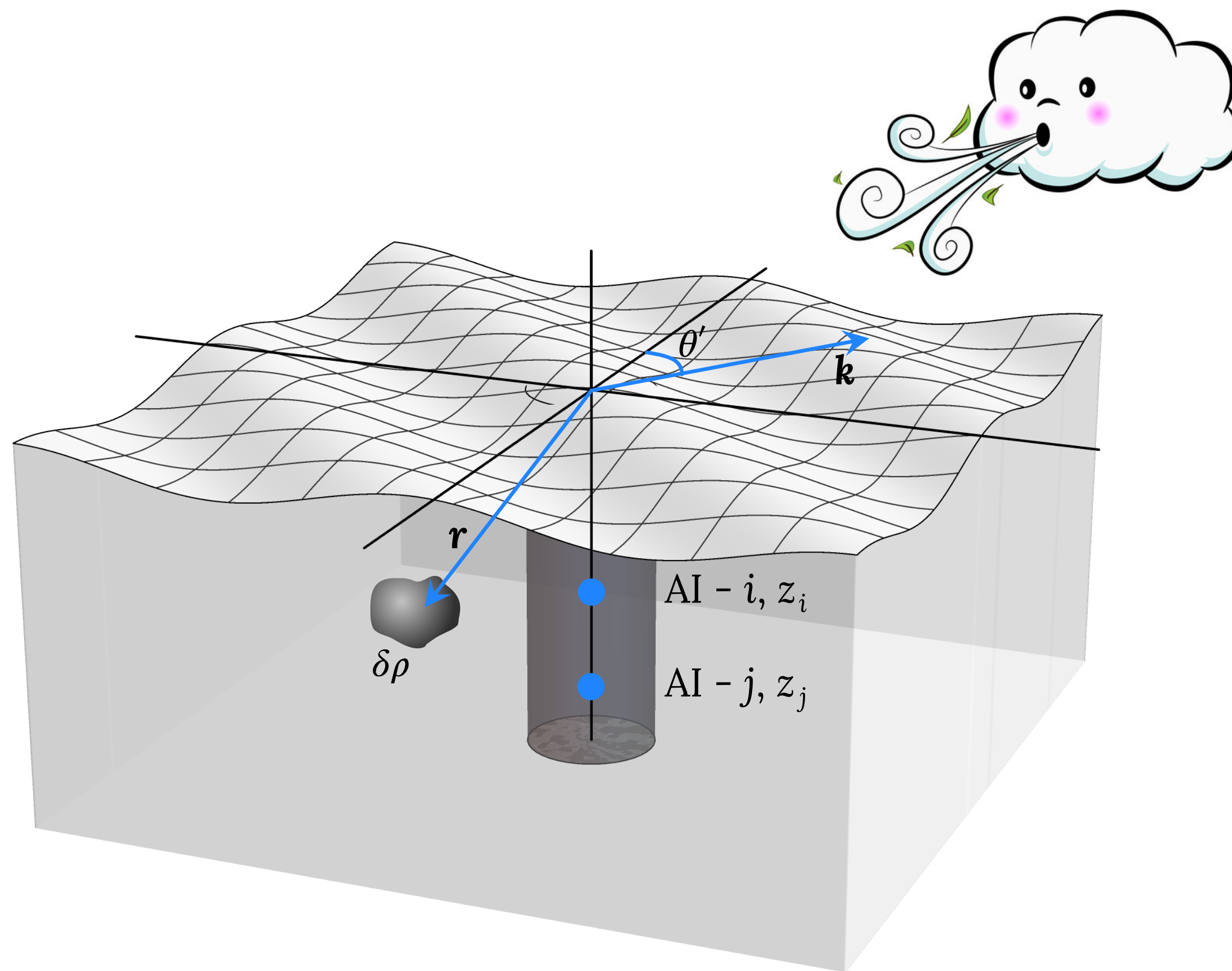
Factors that could limit sub-Hz regime

Seismic effects

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Atmospheric effects

temperature, pressure, ...



Factors that could limit sub-Hz regime

Seismic effects

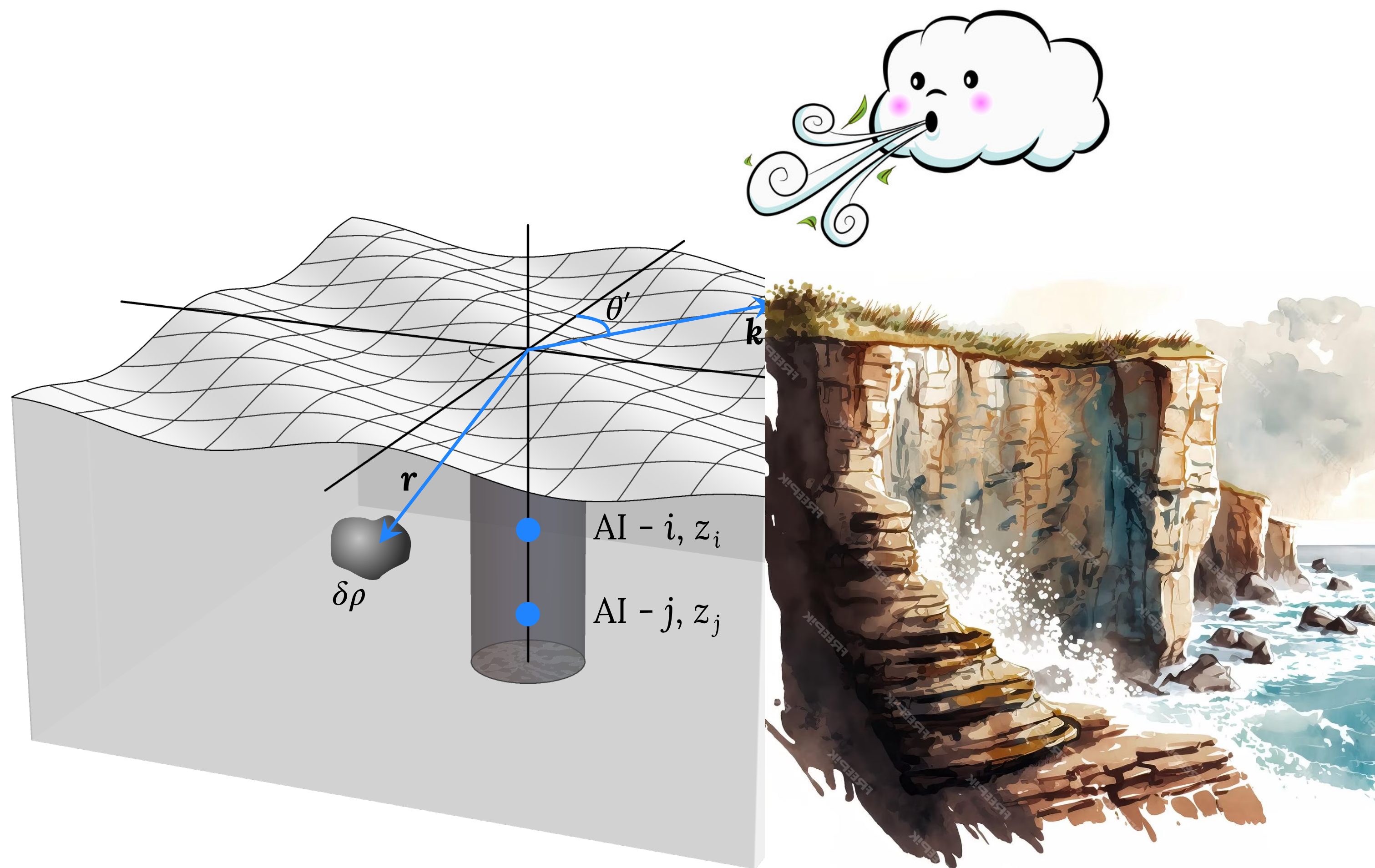
surface waves, body waves, ...

Atmospheric effects

temperature, pressure, ...

Ocean/sea effects

e.g. if hosted at Boubly



Factors that could limit sub-Hz regime

Seismic effects

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Atmospheric effects

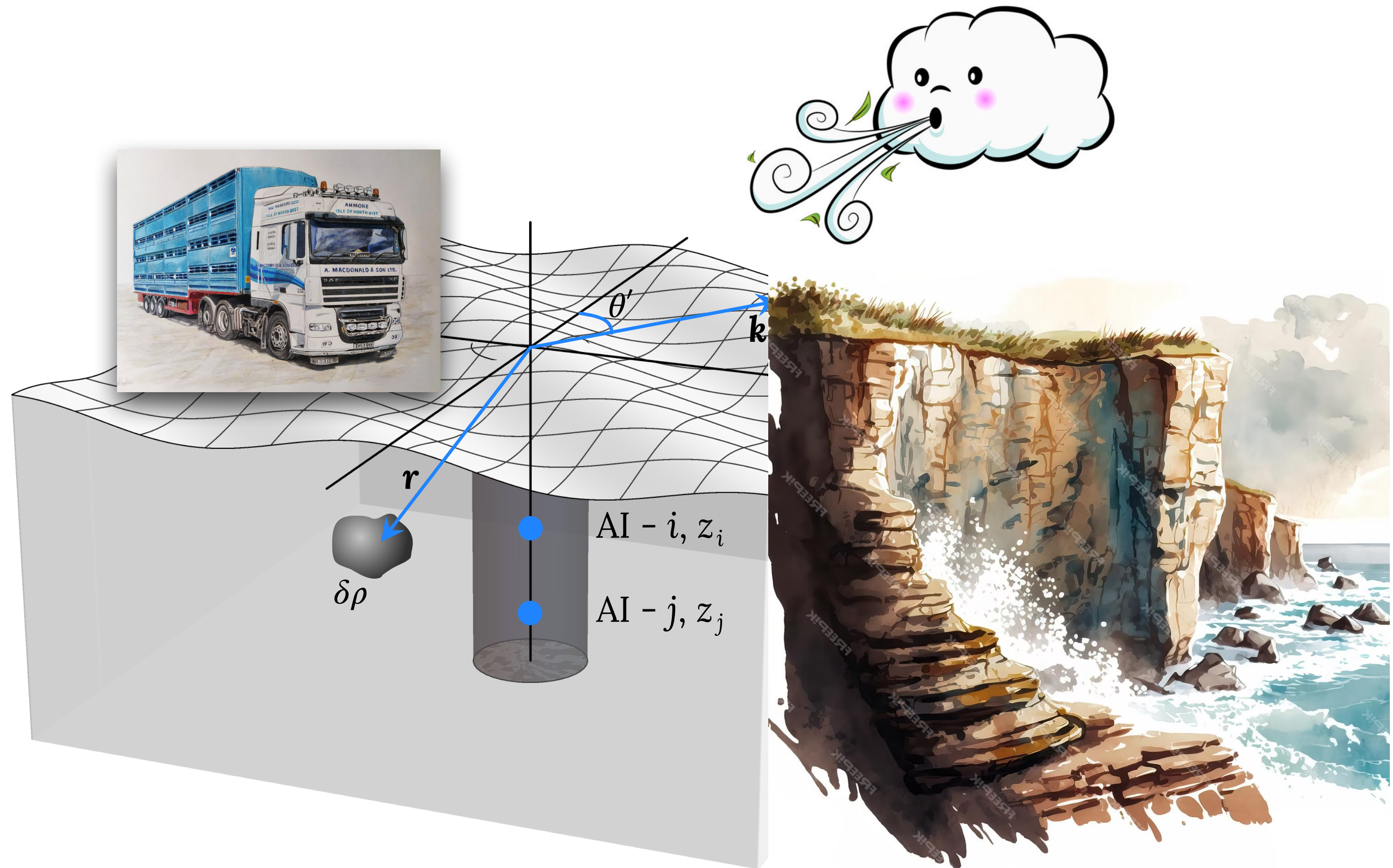
temperature, pressure, ...

Ocean/sea effects

e.g. if hosted at Boubly

Local moving objects

vehicles, etc



Factors that could limit sub-Hz regime

Seismic effects ✓

surface waves, body waves, ...

Badurina, CM, et al, arXiv:2211.01854 (PRD)

Atmospheric effects ✓

temperature, pressure, ...

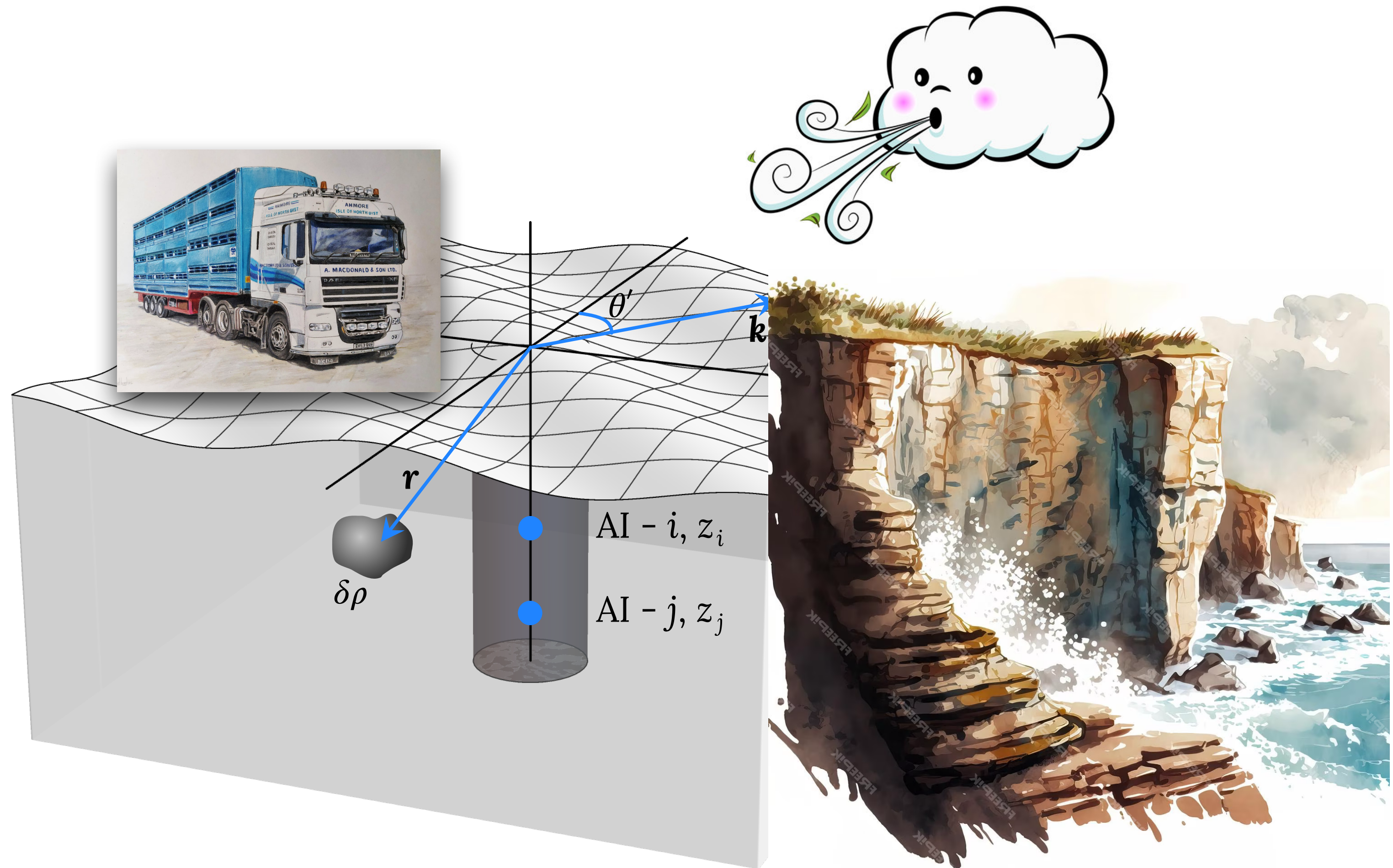
Carlton, CM, et al, arXiv:2412.05379 (PRD)

Ocean/sea effects

e.g. if hosted at Boubly

Local moving objects ✓

Carlton, CM, arXiv:2308.10731 (PRD)



Happy Birthday IPPP: in a
short time, you had a huge
impact on how I approach
dark matter research

still collaborating with
contacts made at that time

still following the lessons
learnt from observing
others here

