

Indirect searches of dark matter



Dark matter in the
Universe

Dark Matter Candidates

See Dark Matter Theory Overview
by Malcolm Fairbairn

See Axion Dark Matter Overview
by Martin Bauer



Mass of dark matter unknown!



Indirect searches of dark matter

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 - ❑ Neutrinos
 - ❑ Charged cosmic rays

Indirect searches of dark matter

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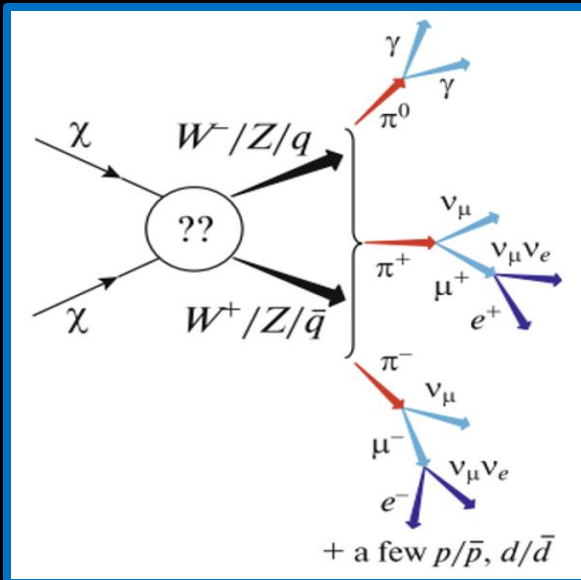
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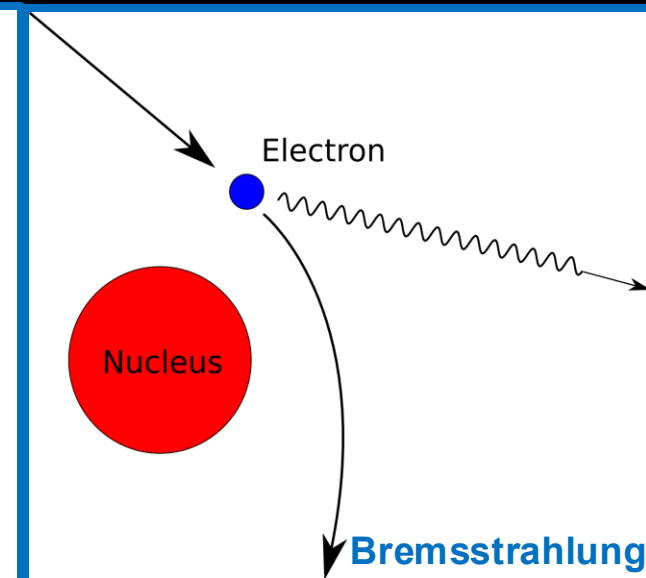
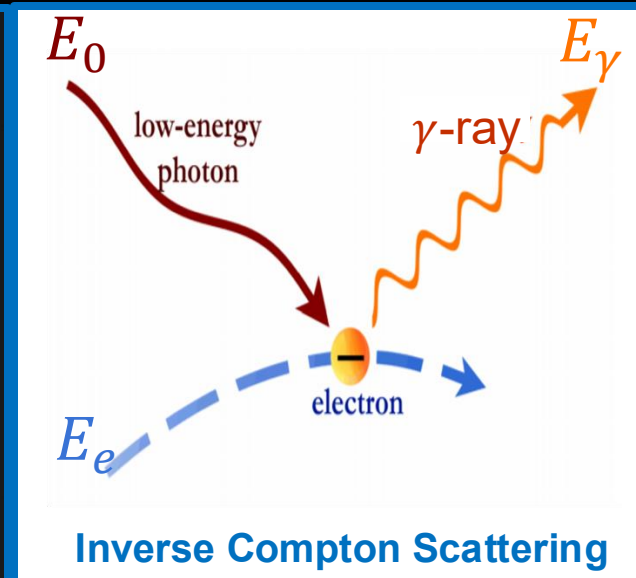
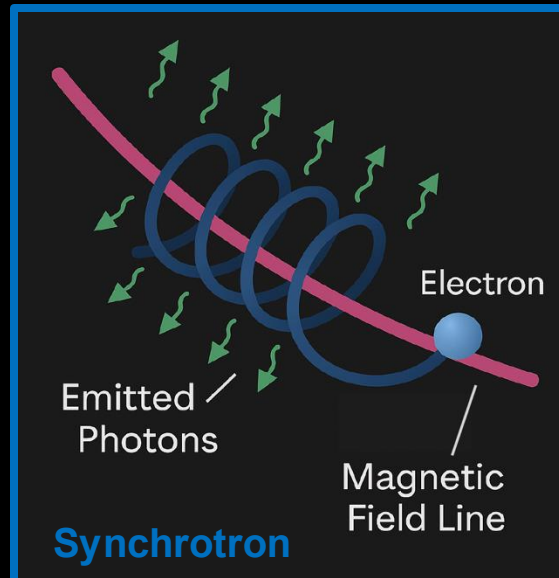
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Primary products



Secondary emission



HEROES vs VILLAINS



Can we discover dark matter with indirect detection?

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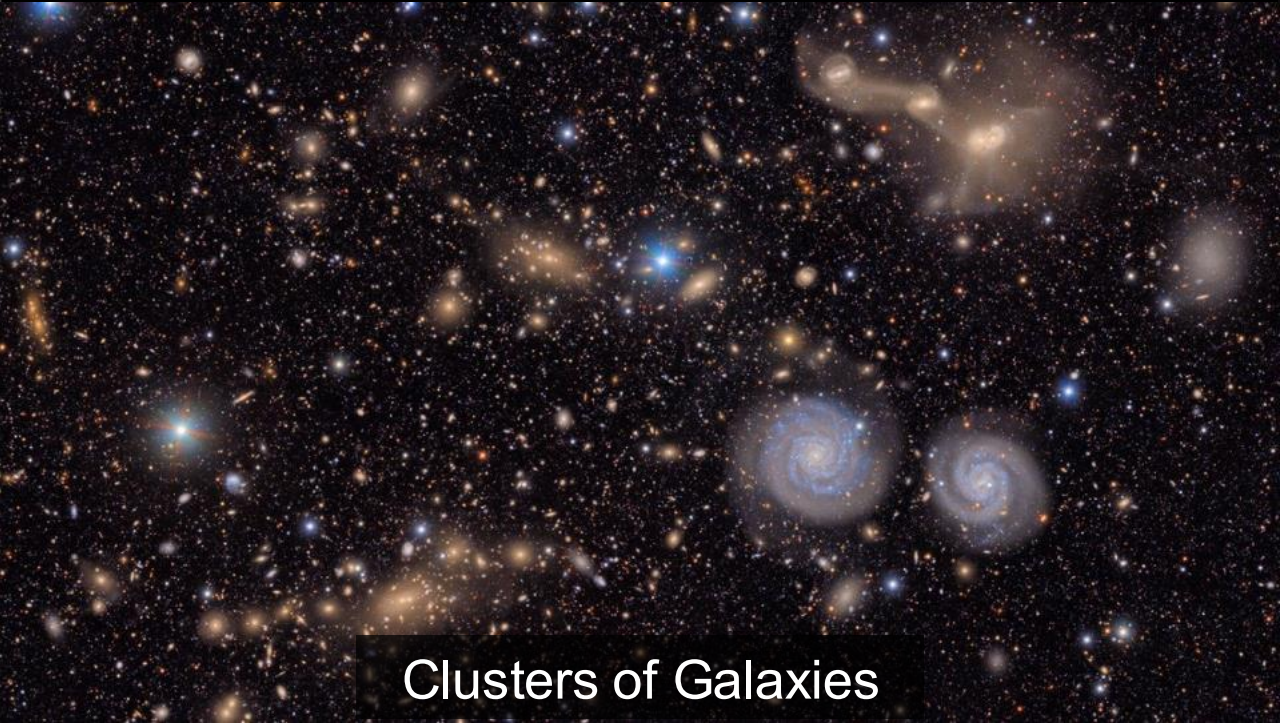
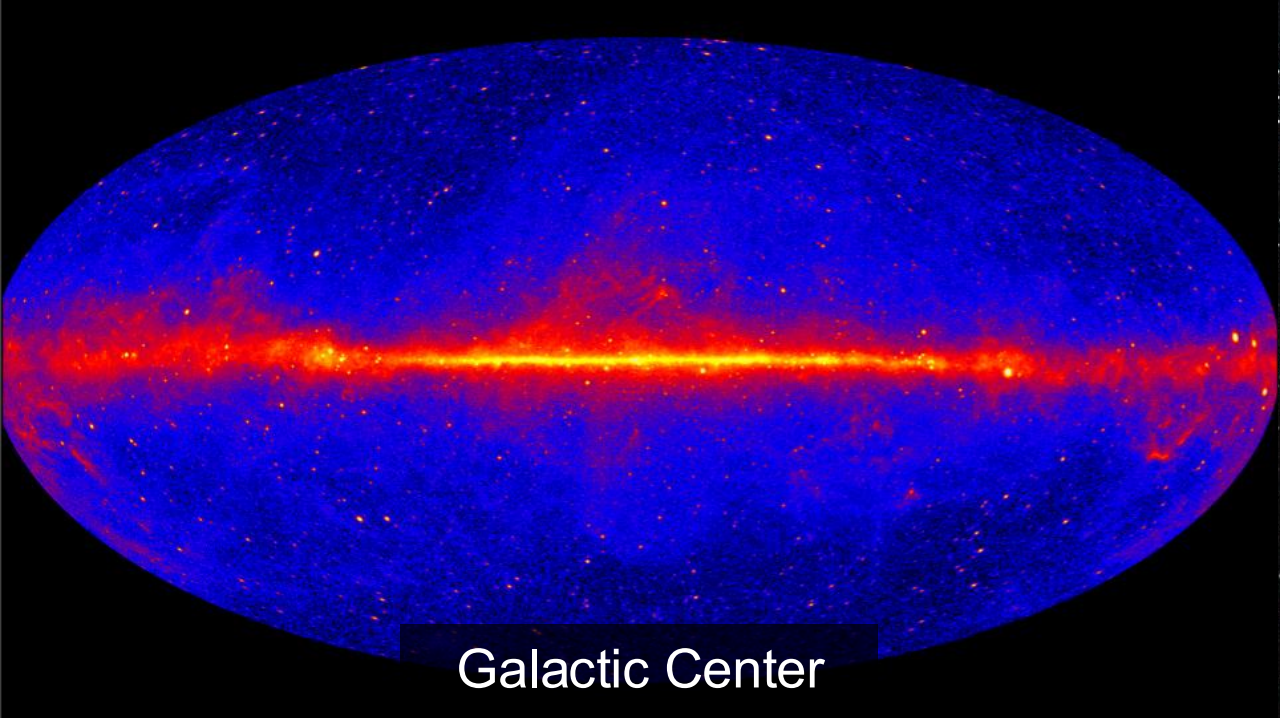
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Can we discover dark matter with indirect detection?

- ① Unique signatures of dark matter (model dependent)
- ② Modelling the astrophysical background (hard and scary)
- ③ Promising targets/astrophysical environments (creative and fun!)

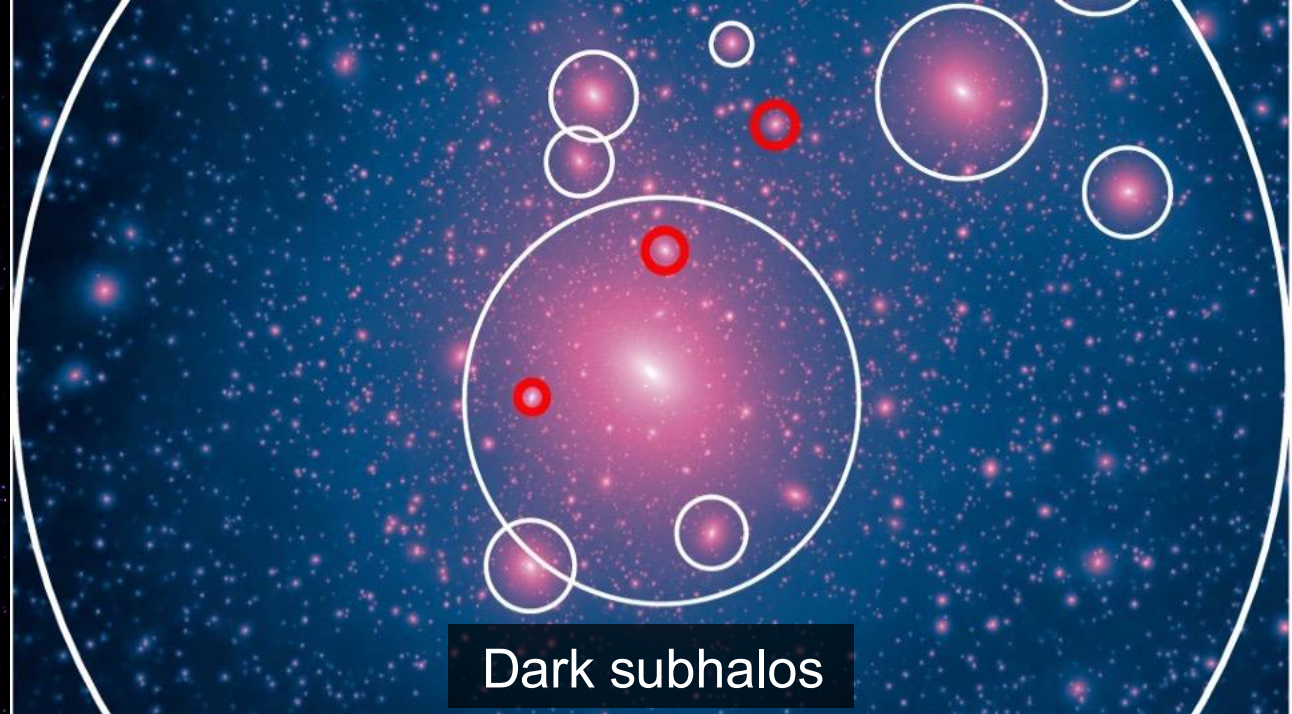


Where should we look?

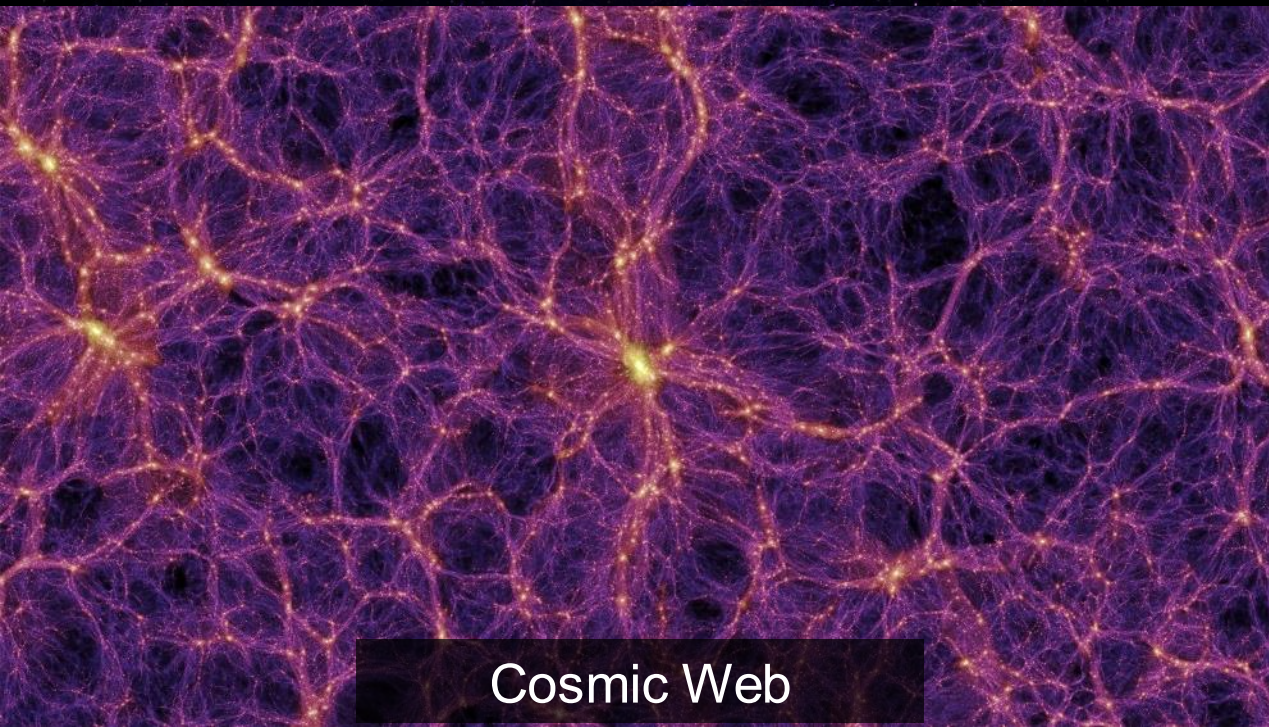




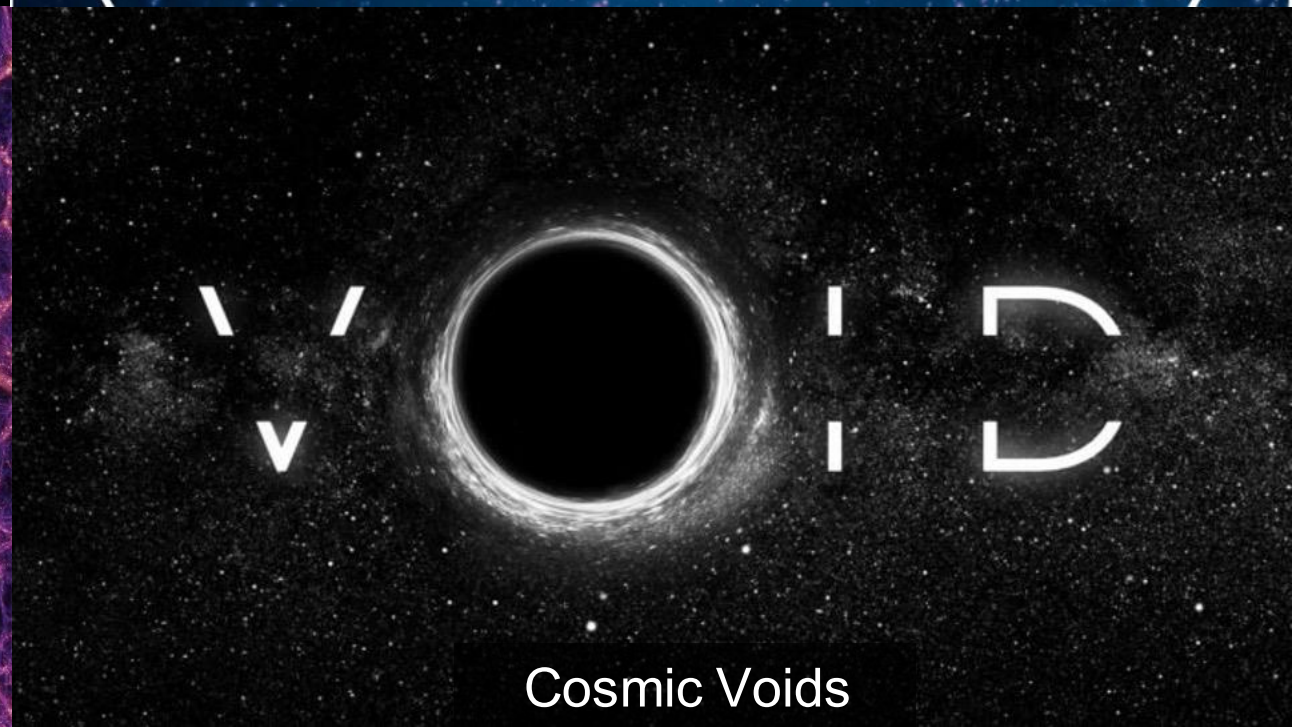
Blank Sky



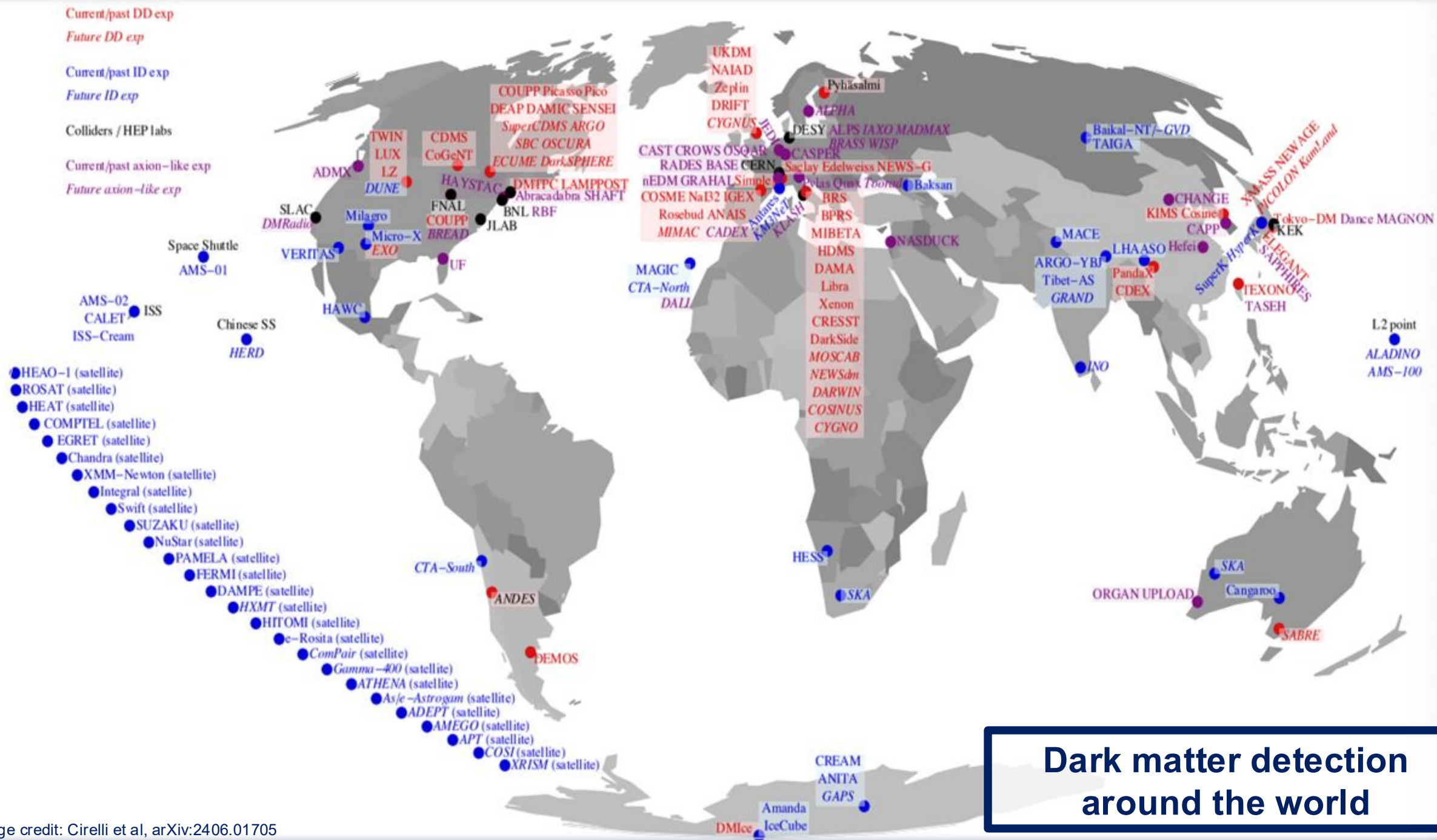
Dark subhalos



Cosmic Web



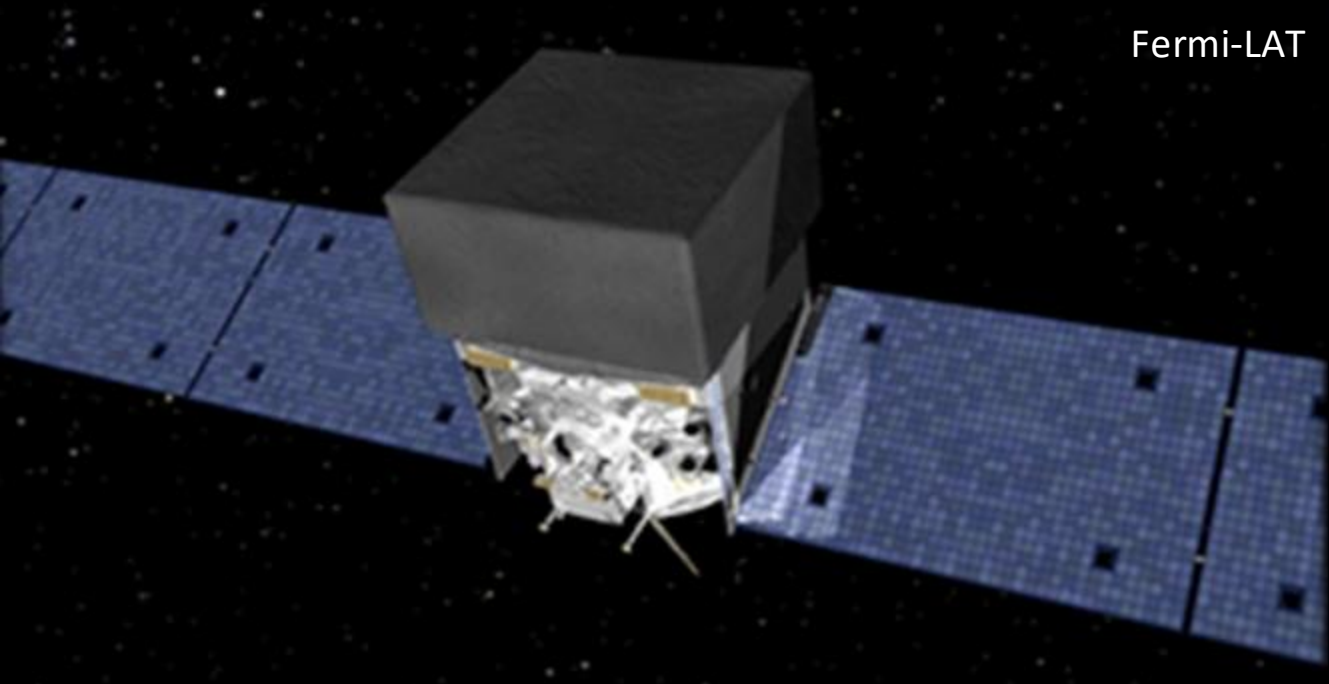
Cosmic Voids



**Dark matter detection
around the world**

Outline

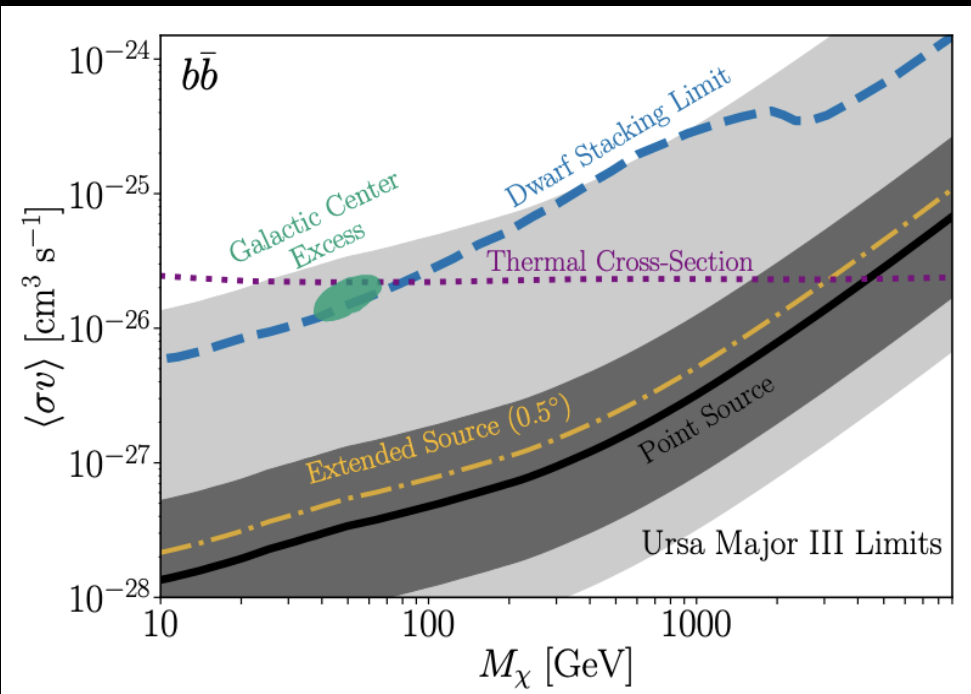
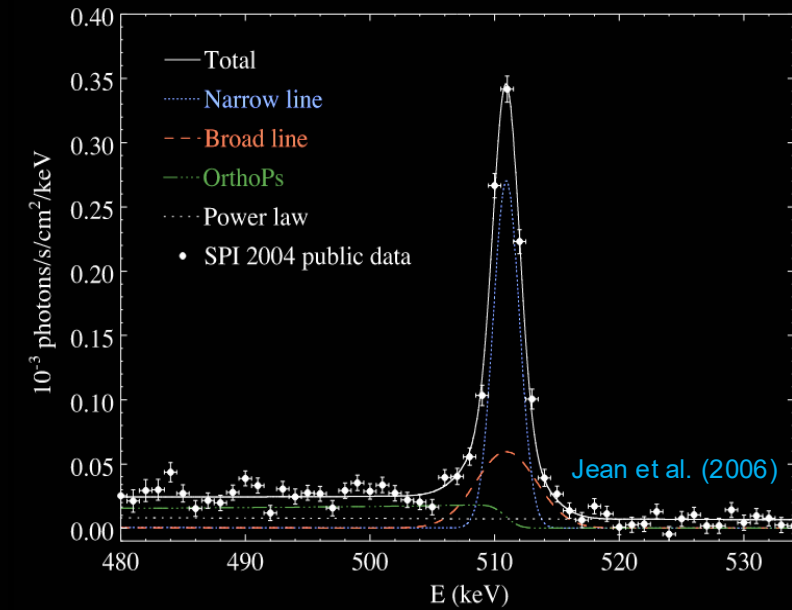
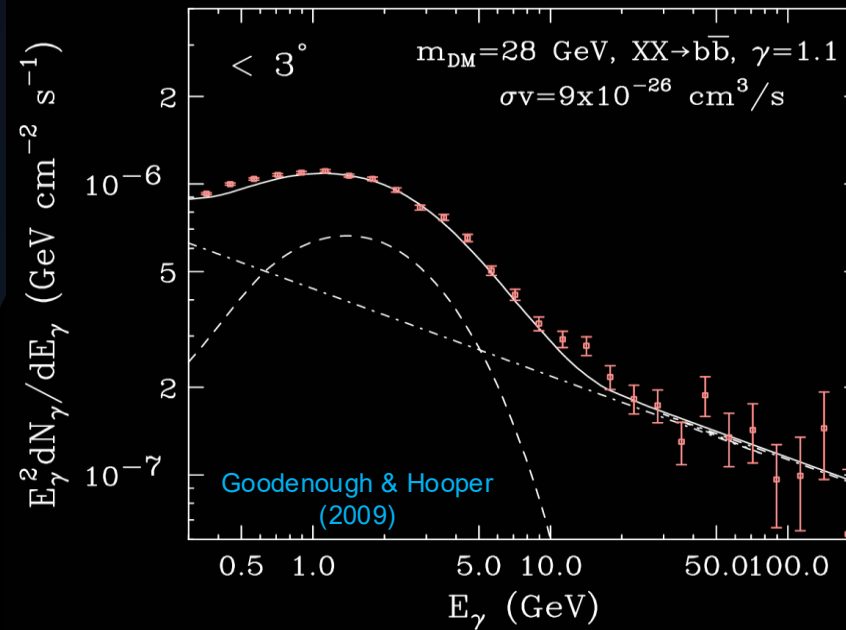
- ❑ Current status of indirect searches
- ❑ Longstanding and emerging anomalies
- ❑ Promising prospects for future searches



Dark matter searches with gamma-rays



Gamma-ray Anomalies



No Observational Evidence for Dark Matter Nor a Large Metallicity Spread
in the Extreme Milky Way Satellite Ursa Major III / UNIONS 1

WILLIAM CERNY,¹ DAISY BISSONETTE,² ALEXANDER P. JI,^{2,3,4} MARLA GEHA,¹ ANIRUDH CHITI,^{2,3} SIMON E.T. SMITH,⁵
JOSHUA D. SIMON,⁶ ANDREW B. PACE,⁷ EVAN N. KIRBY,⁸ KIM A. VENN,⁵ TING S. LI,⁹ AND ALICE M. LUNA^{2,3}

Sub-GeV Dark Matter

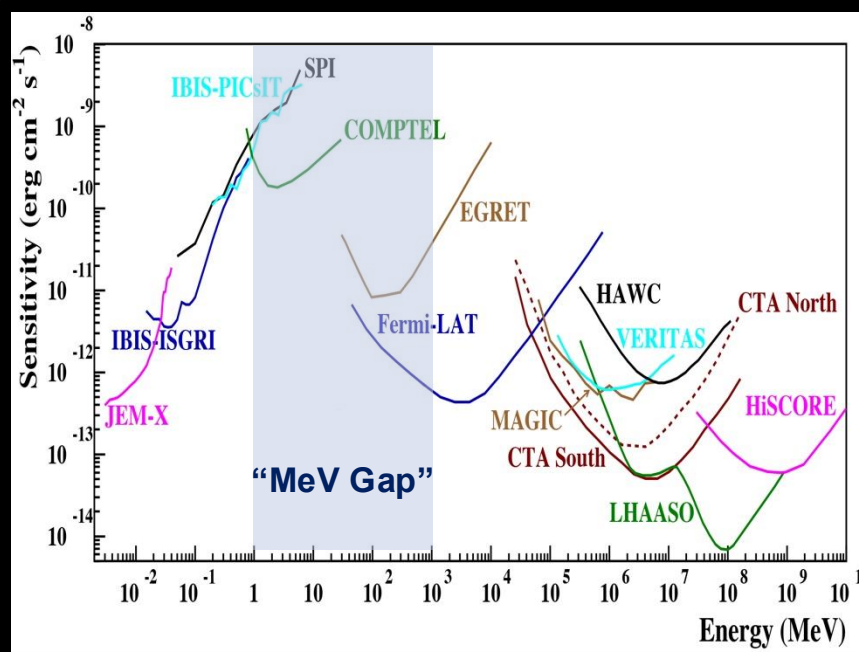
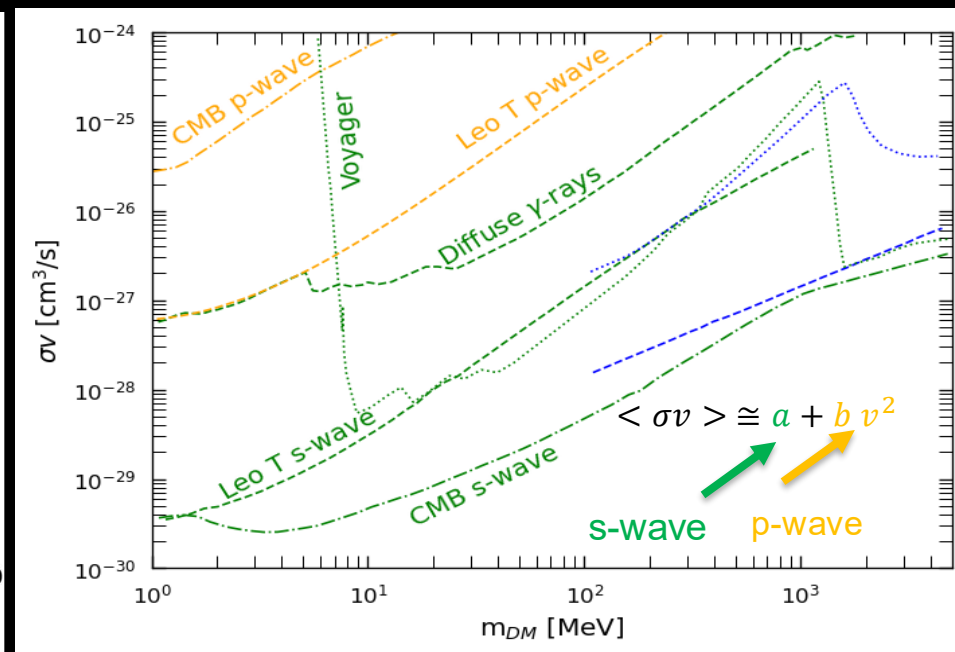


Figure adapted from Tatischeff+ arxiv:1805.06435



Slatyer, PRD 93 (2016) 023527 Wadekar et al, PRD106 (2022) 7, 075007
Essig+, JHEP 11 (2013) 193 Boudaud et al., PRL. 119 (2017) 021103

See Light Dark Matter Searches
by Ashlea Kemp on Tuesday

Sub-GeV Dark Matter

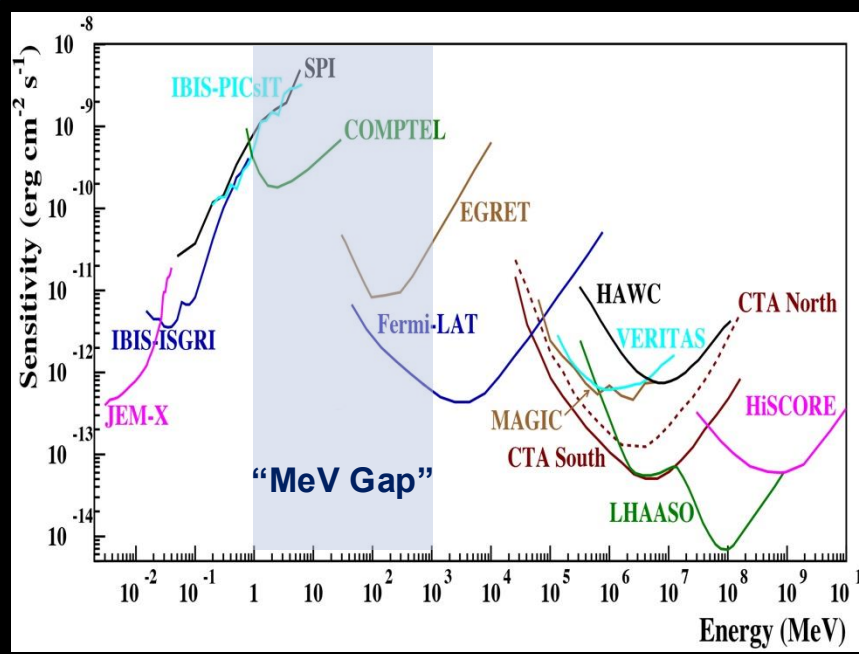
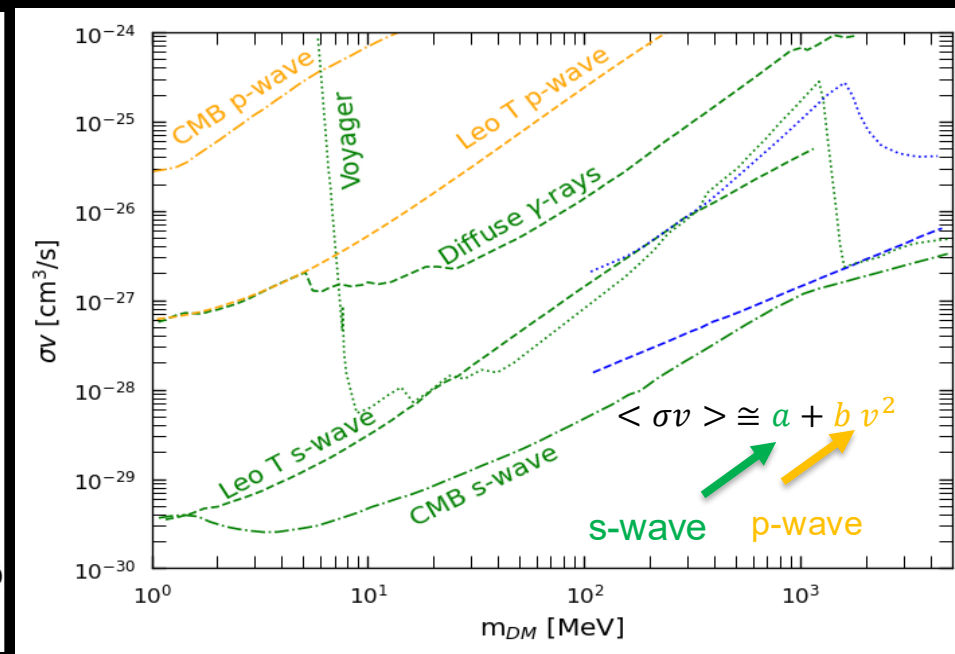


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Slatyer, PRD 93 (2016) 023527 Wadekar et al, PRD106 (2022) 7, 075007
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- ① Can we have “unique signatures” of sub-GeV DM?
- ② Can we place strong constraints beyond CMB?

Sub-GeV Dark Matter

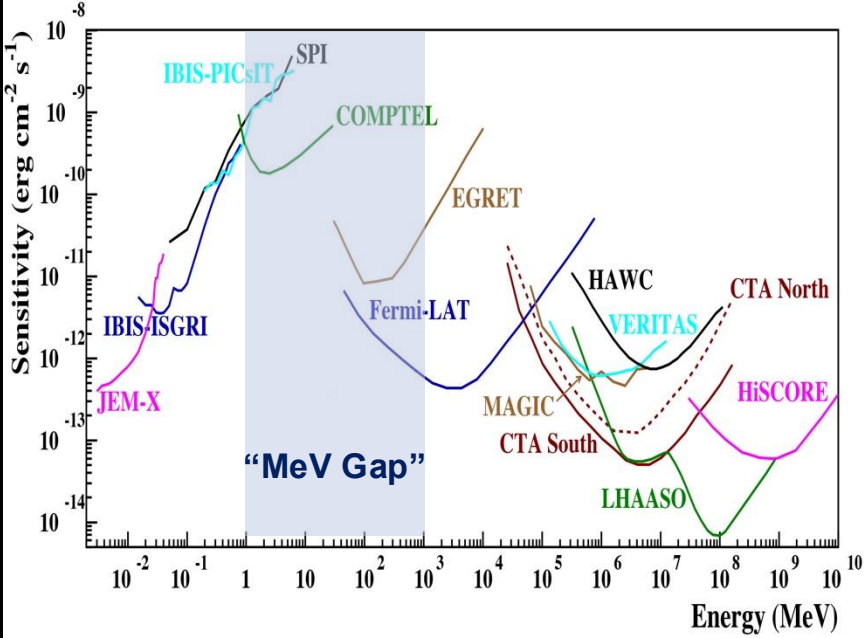
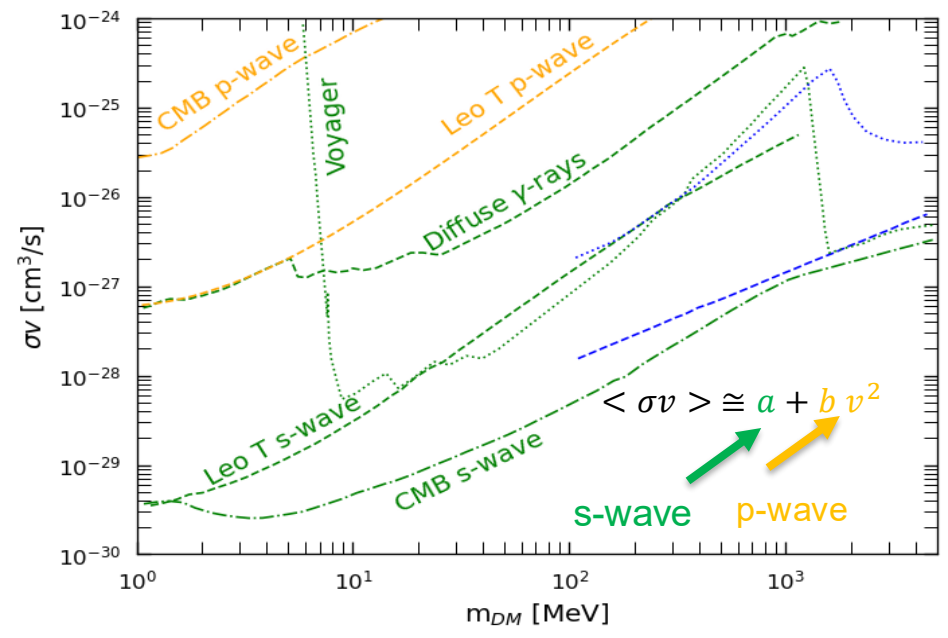
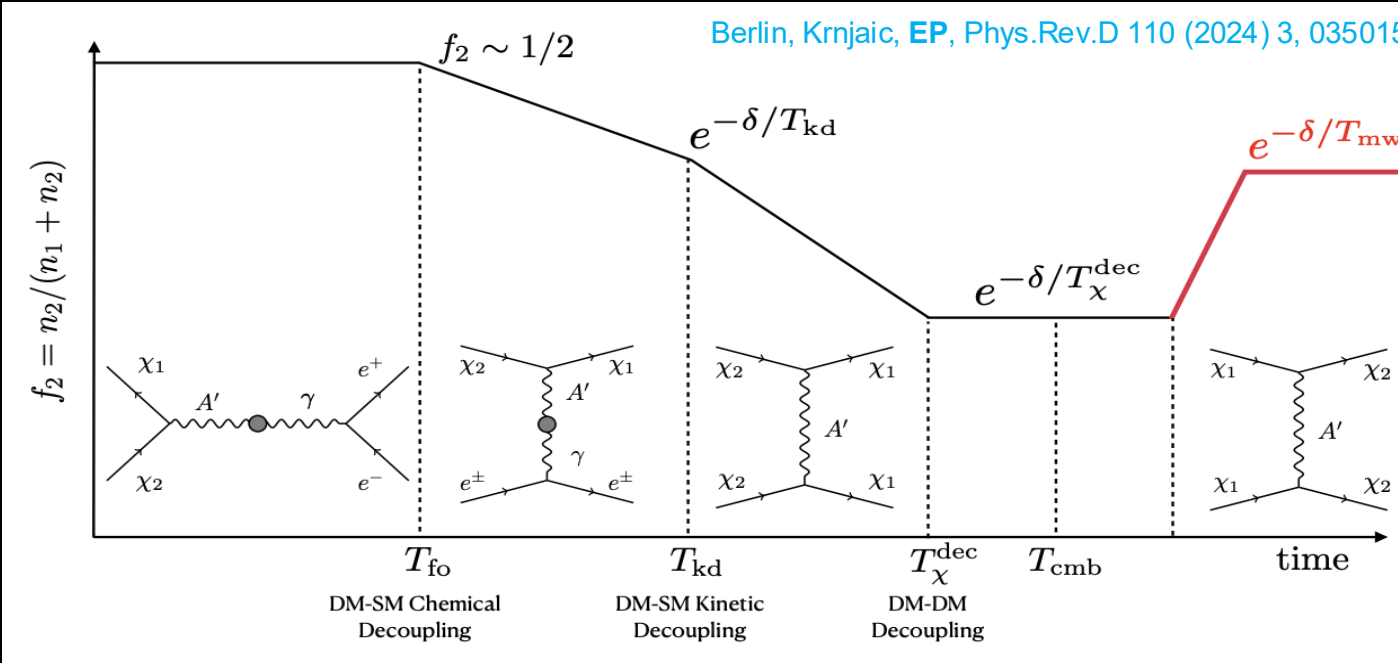


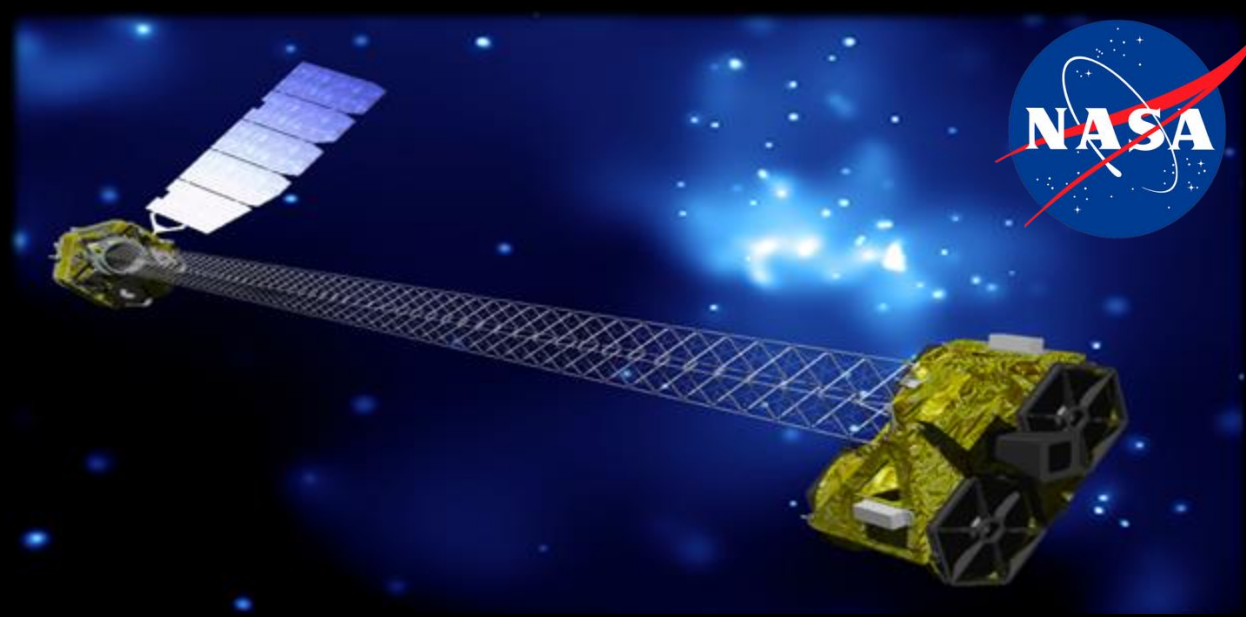
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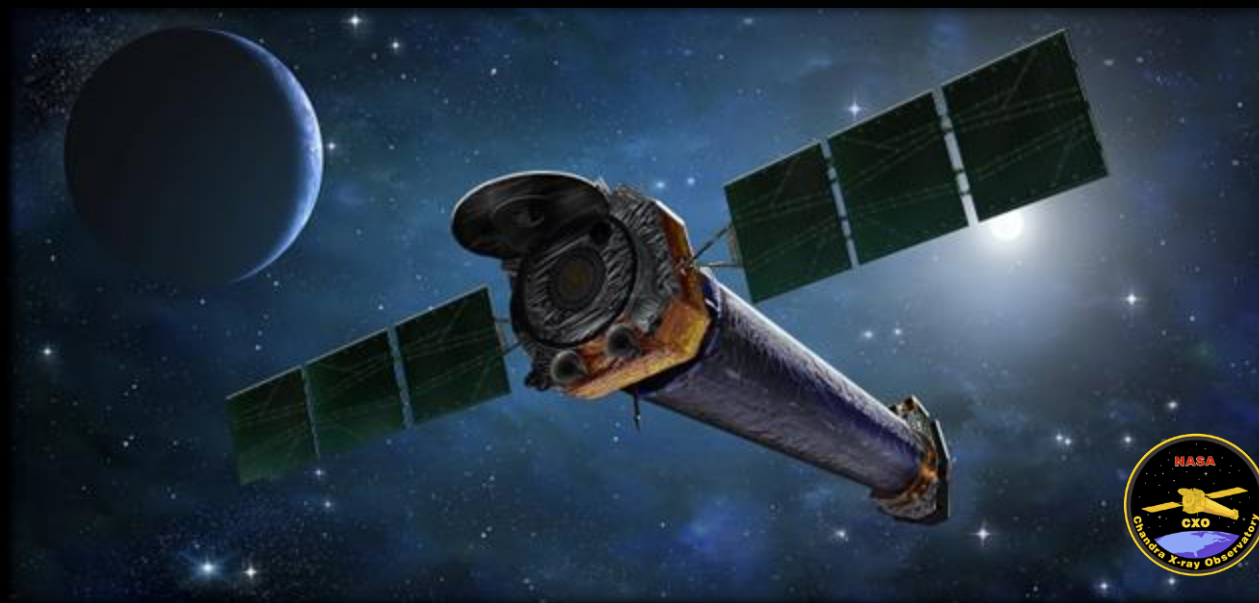
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Essig+, JHEP 11 (2013) 193 Boudaud et al., PRL. 119 (2017) 021103



Berlin, Krnjaic, EP, Phys.Rev.D 110 (2024) 3, 035015



Dark matter searches with X-rays



Sub-GeV Dark Matter

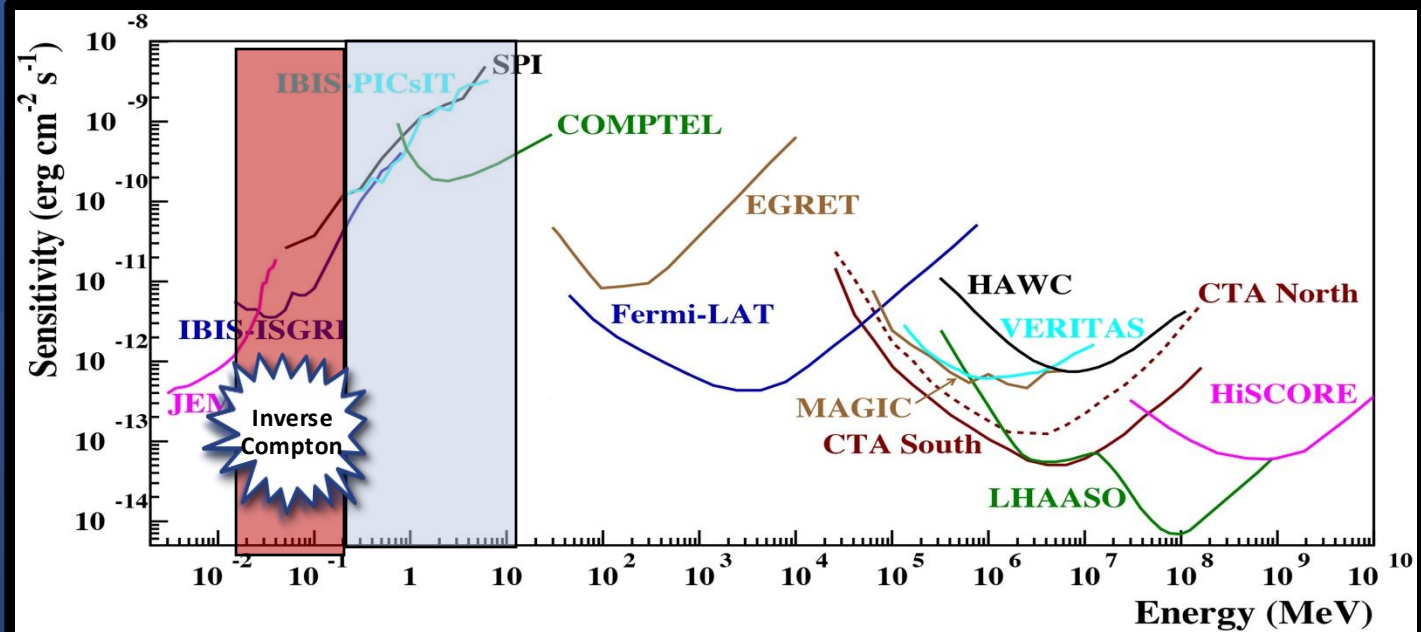
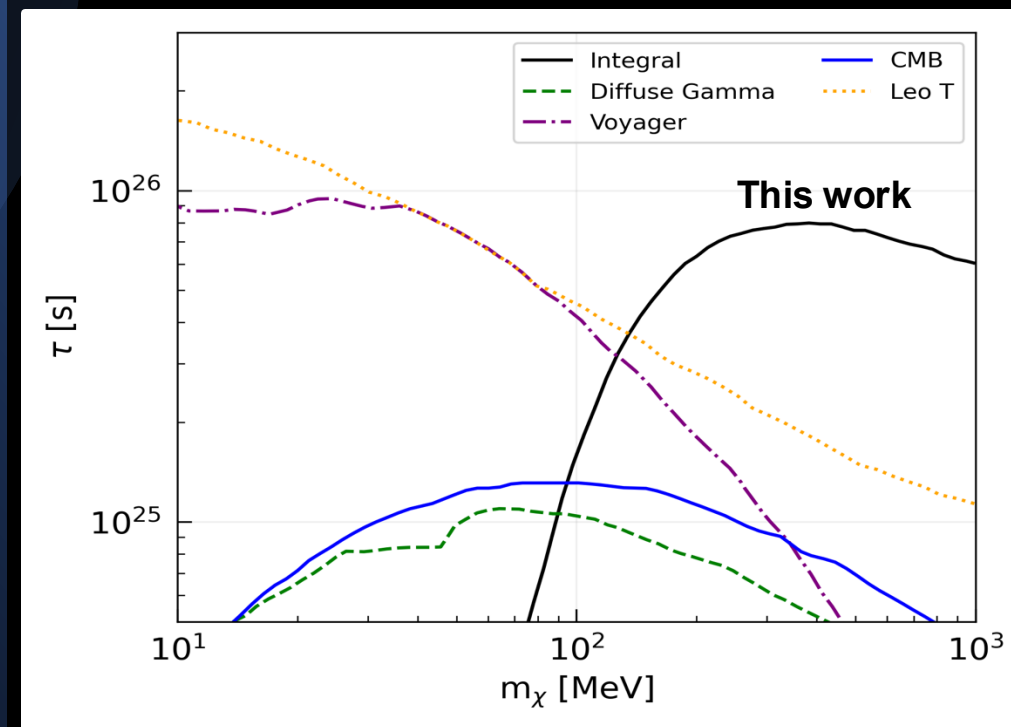
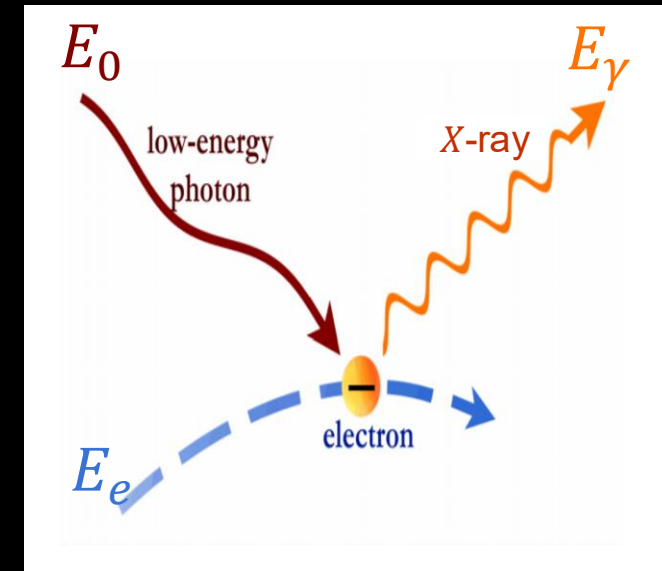


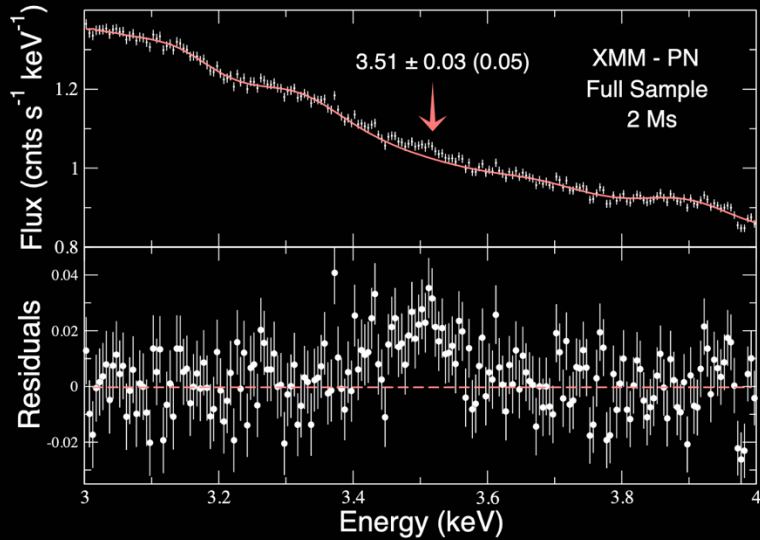
Figure adapted from Tatischeff+ arxiv:1805.06435



$$\chi(\chi) \rightarrow (\dots) \rightarrow e^+ e^-$$



3.5 keV line?



In 2014 two independent groups reported an unidentified line at 3.5 keV with XMM-Newton and Chandra in clusters and M31

Boyarsky et al, PRL 113, 251301

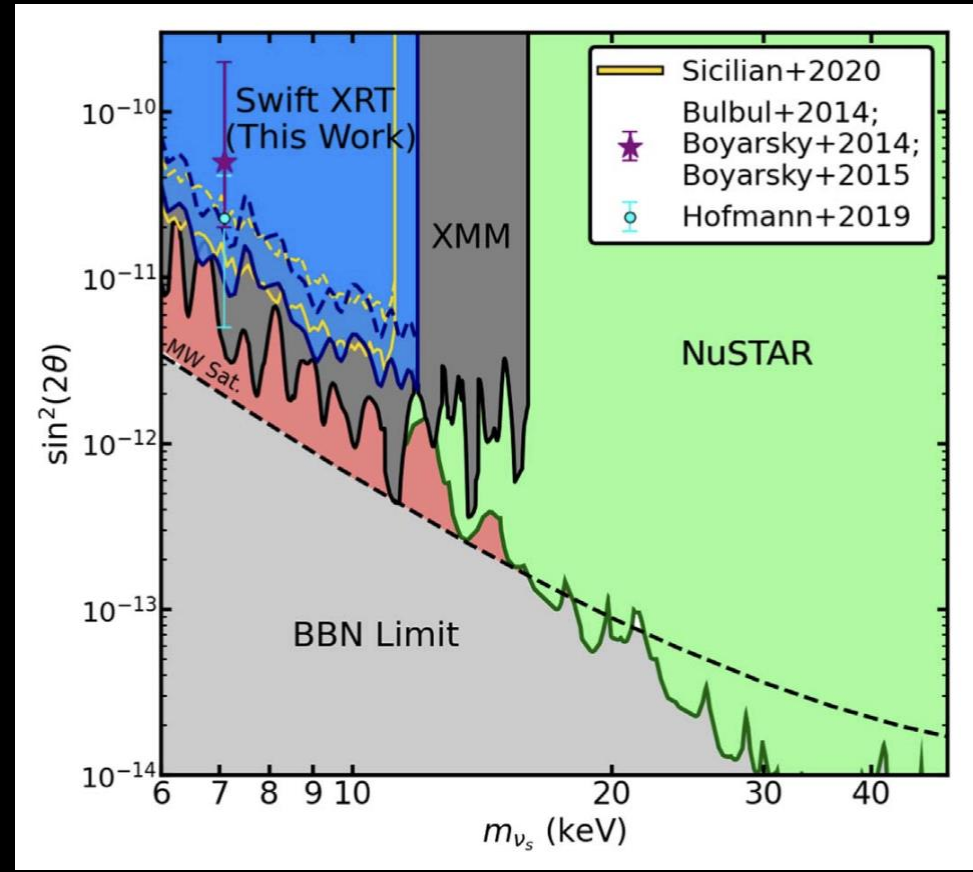
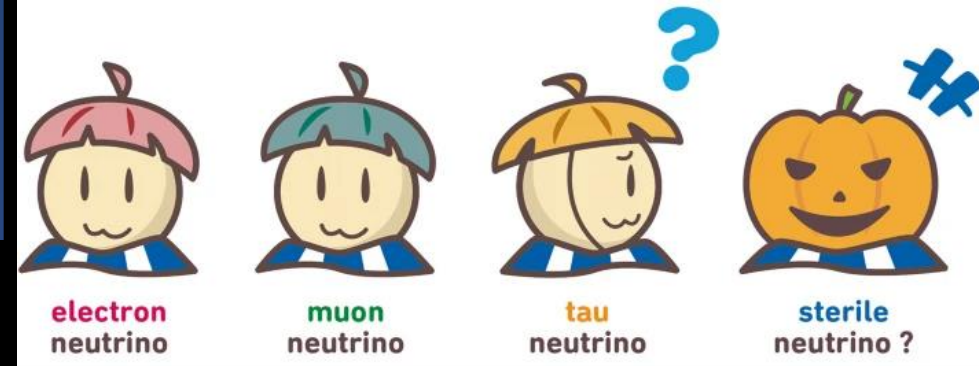
Bulbul et al, APJ 789, 13

$$\nu_s \longrightarrow \nu + \gamma \quad \text{with } m_{\nu_s} = 7 \text{ keV}$$

☐ **Atomic lines** (e.g. potassium, sulfur etc) [Jeltema et al, MNRAS 450 \(2015\) 2](#)

☐ The existence of the line is highly **disputed** [Dessert et al, APJ 964 \(2024\) 2,185 1, 2](#)

☐ **Excluded** (except Shi-Fuller mechanism with large lepton asymmetry) [Vogel et al, arXiv:2509.05631](#)
[Akita et al, arXiv:2507.20659](#)



[Sicilian et al, APJ 941 \(2022\) 1, 2](#)



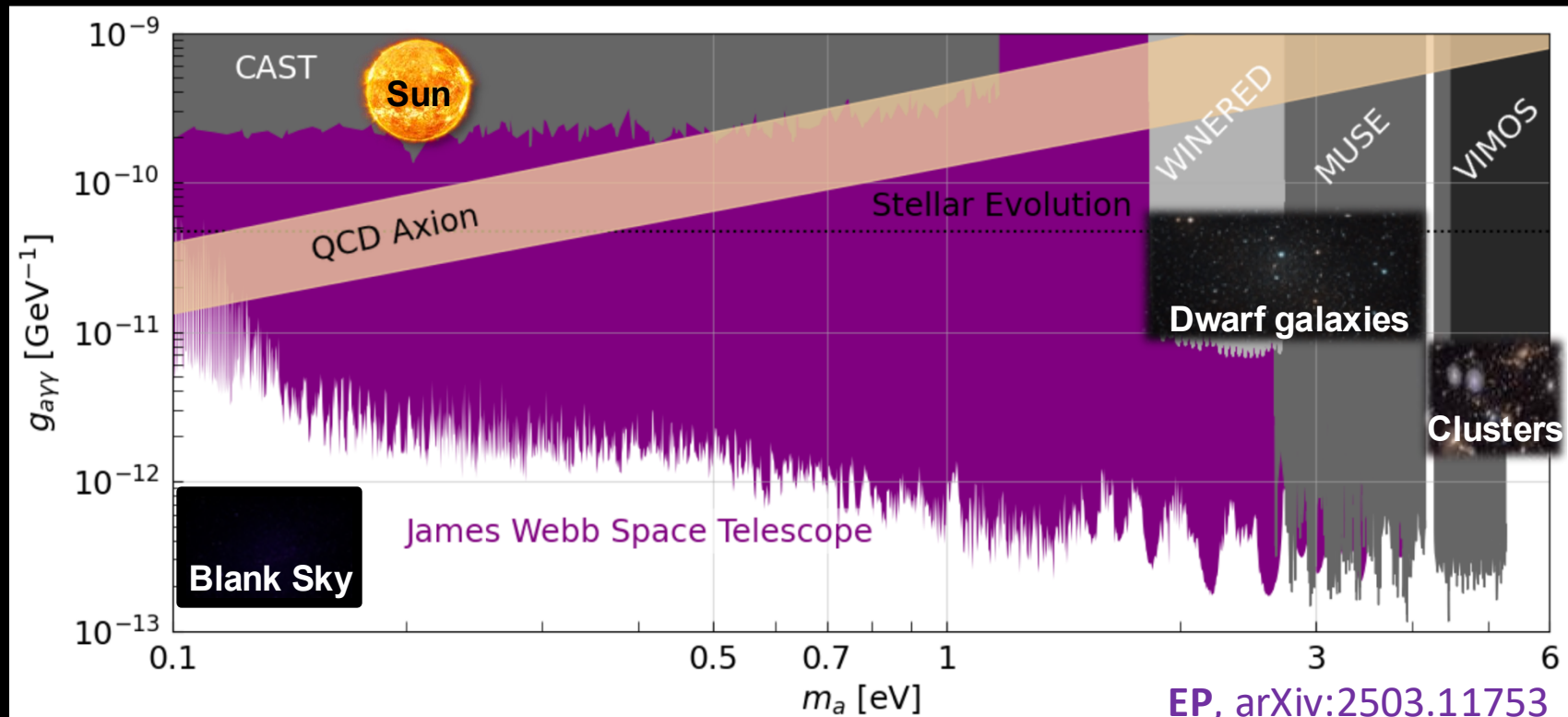
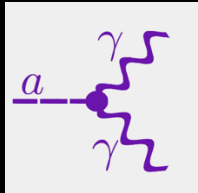
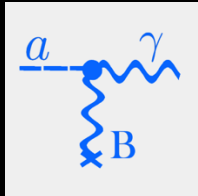
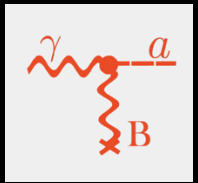
Axion searches with infrared and optical light



Axions at the eV scale

See Axion Dark Matter Overview
by Martin Bauer on Tuesday

- 1 Primakoff production in stars: $\gamma \rightarrow a$
- 2 Conversion $a \leftrightarrow \gamma$ in laboratory and astrophysical B-fields
- 3 Axion decay $a \rightarrow \gamma\gamma$



EP, arXiv:2503.11753

MWA



FAST



Radio searches in the cosmic web

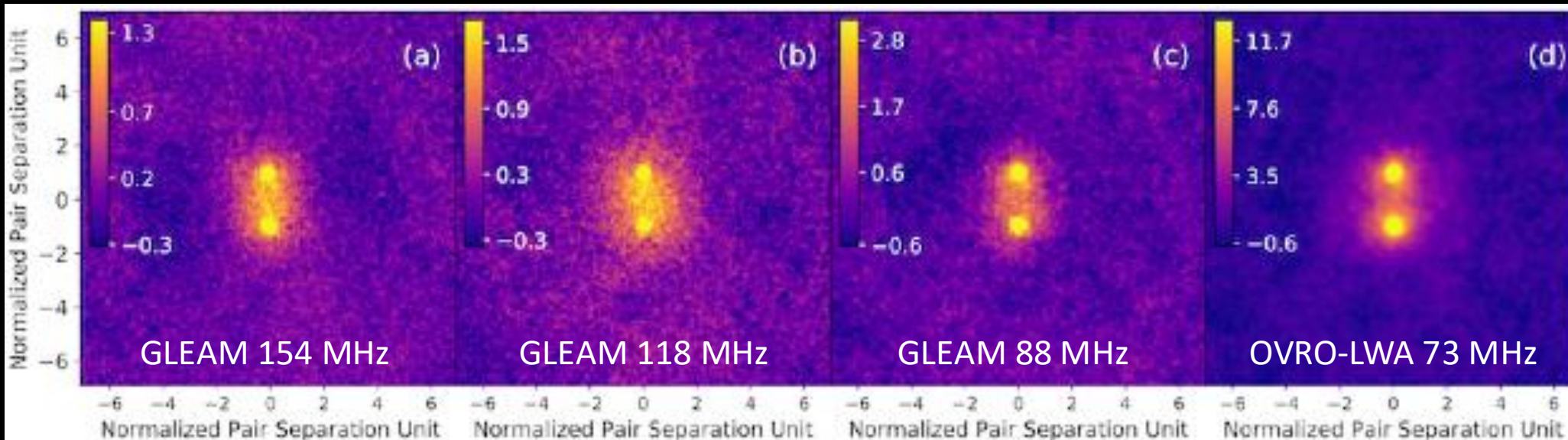
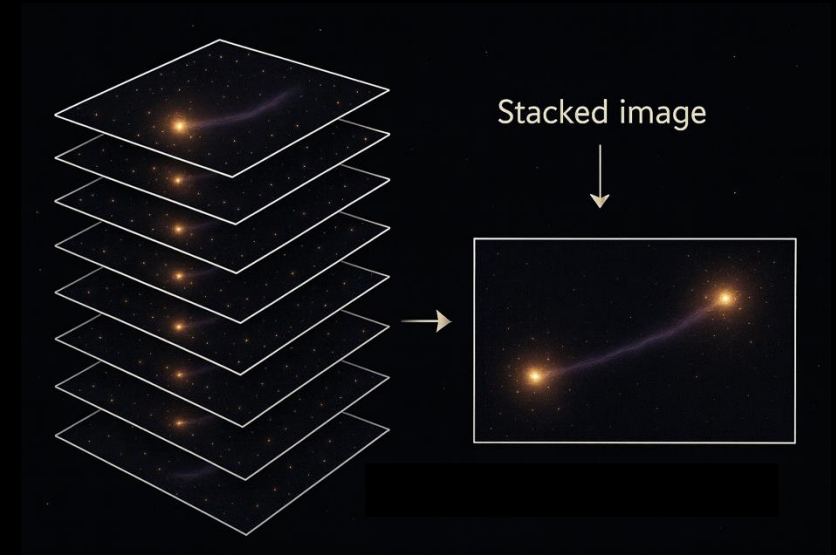
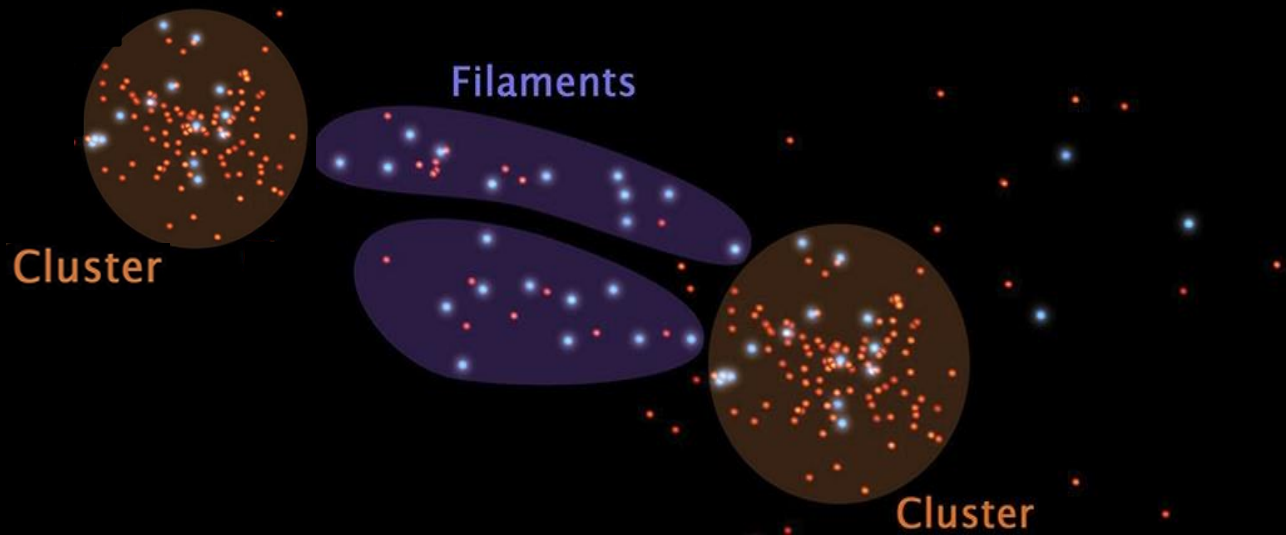
LWA



LOFAR



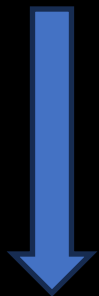
Dark matter searches with the cosmic web



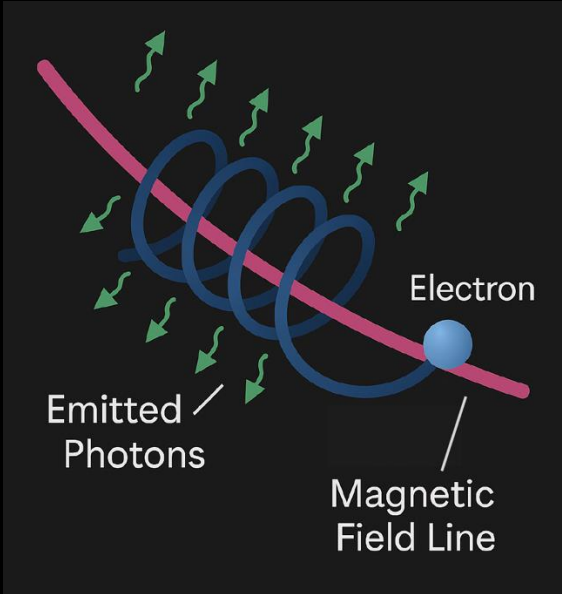
Vernstrom et al
(including EP), MNRAS
505 (2021) 3, 4178-4196

Dark matter searches with the cosmic web

$$\chi \rightarrow e^+ e^-$$



Synchrotron radiation

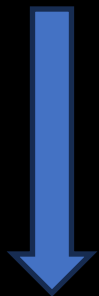


$m_{\text{DM}} = 30 \text{ GeV}$
 $\tau = 2 \times 10^{27} \text{ s}$
 $B = 30 \text{ nG}$
Mean DM density profile

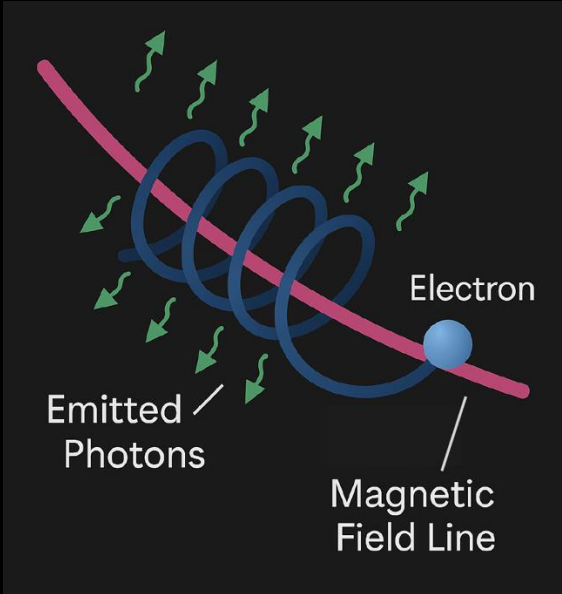
ν [MHz]	T_b^{obs} [K]	T_b^{DM} [K]
154	0.10 ± 0.04	0.10
118	0.22 ± 0.06	0.22
88	0.44 ± 0.09	0.51
73	1.1 ± 0.2	0.85

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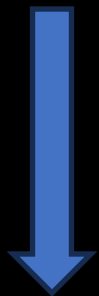
Why haven't we observed this signal elsewhere?

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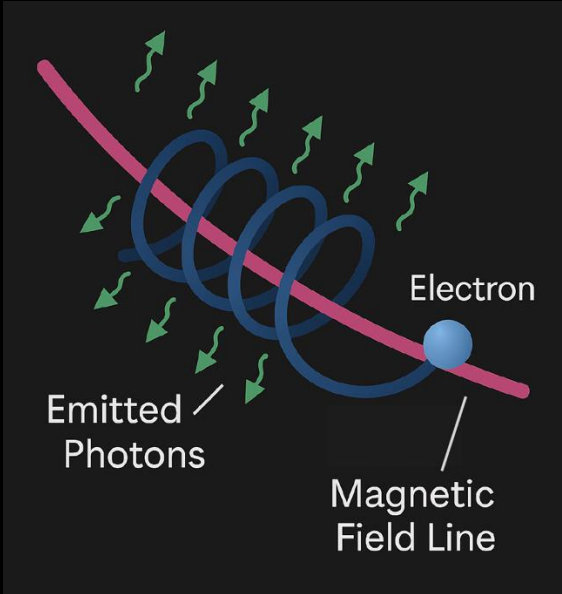
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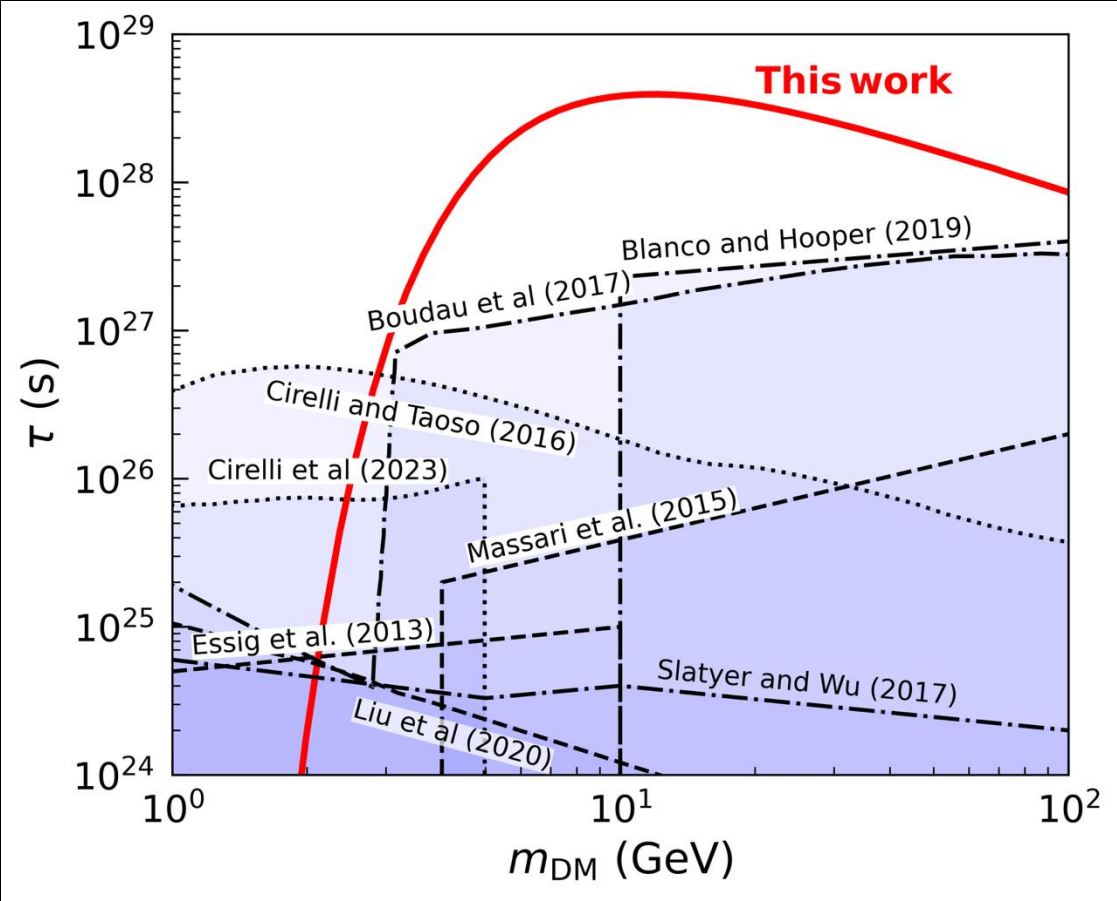
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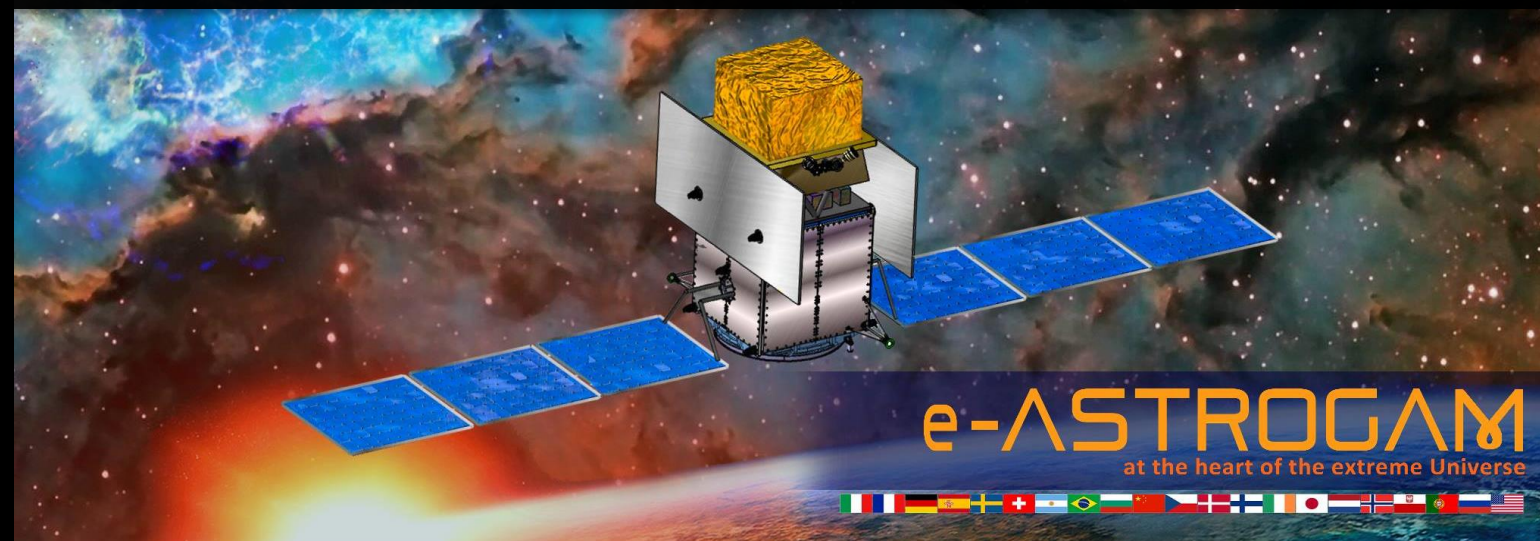




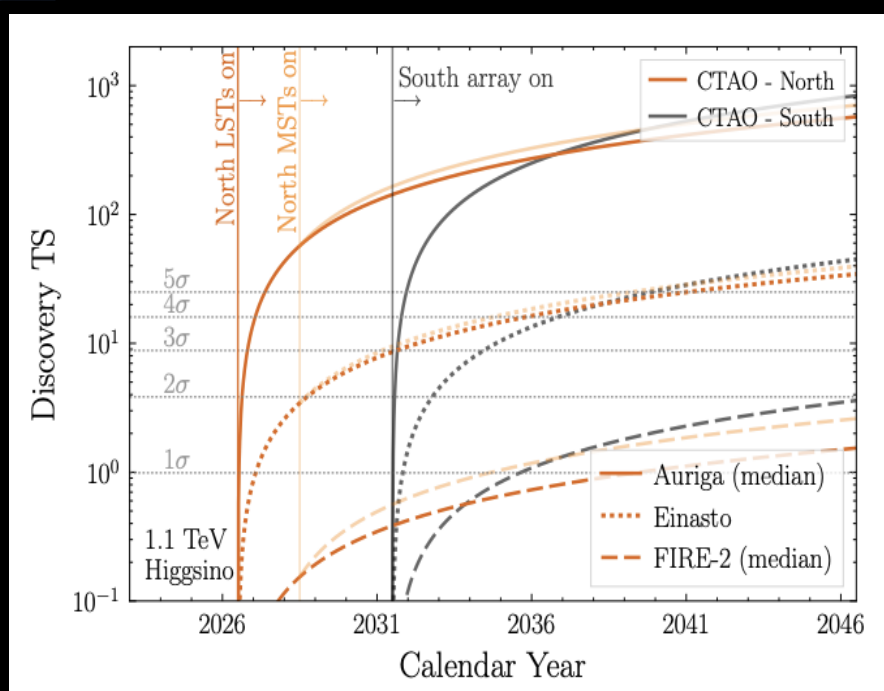
An eye toward the future



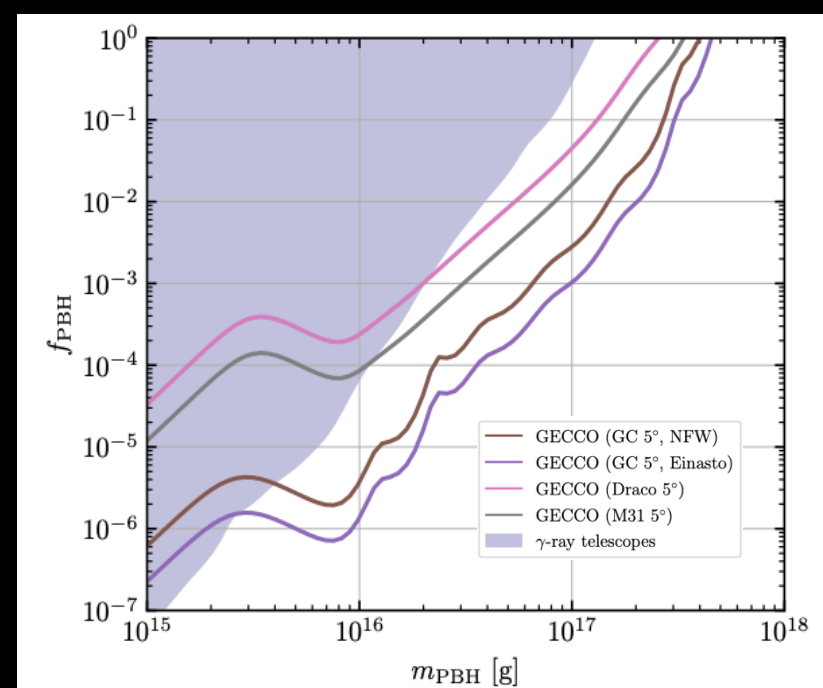
Future gamma-ray searches



Future γ -ray searches



Abe et al, arXiv:2506.08084

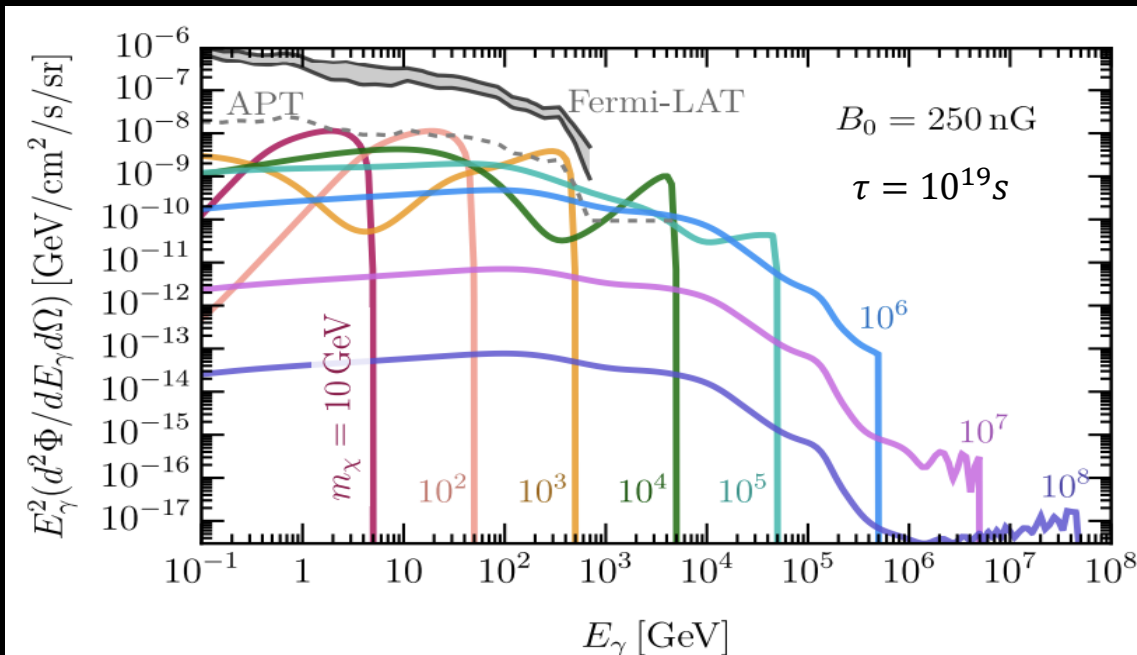


Coogan et al, PRD 107(2), 023022 (2023)



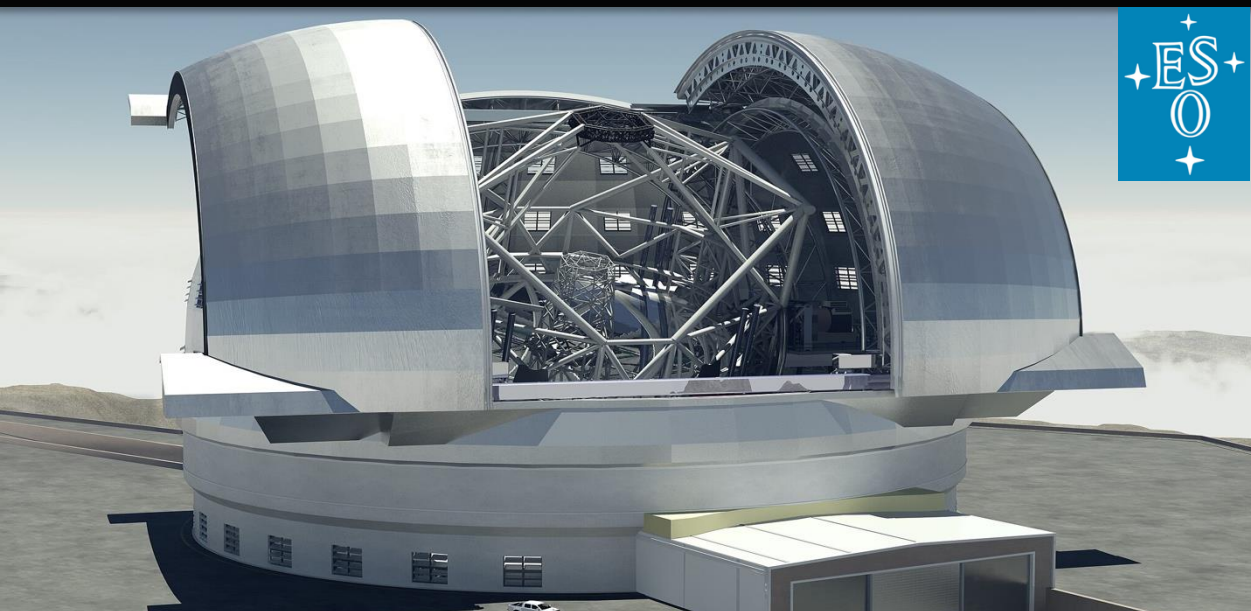
Graviton-Photon Conversion in cosmological filaments produce an irreducible gamma-ray signal

Dunsky, Krnjaic, EP, 2503.19019





Future infrared and optical searches



Exoplanets & Stars

Planets & Exoplanets



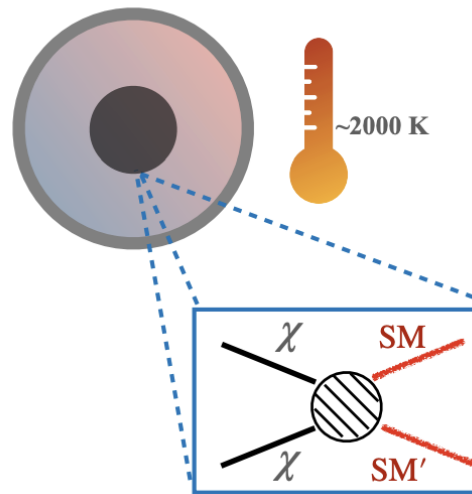
Up to $\sim 13x$
Jupiter's mass

Brown Dwarfs

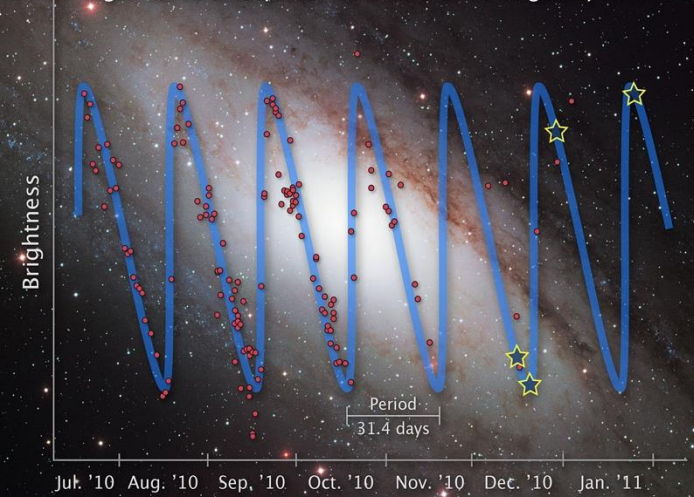


$\sim 13x$ to $80x$
Jupiter's mass

annihilation heating

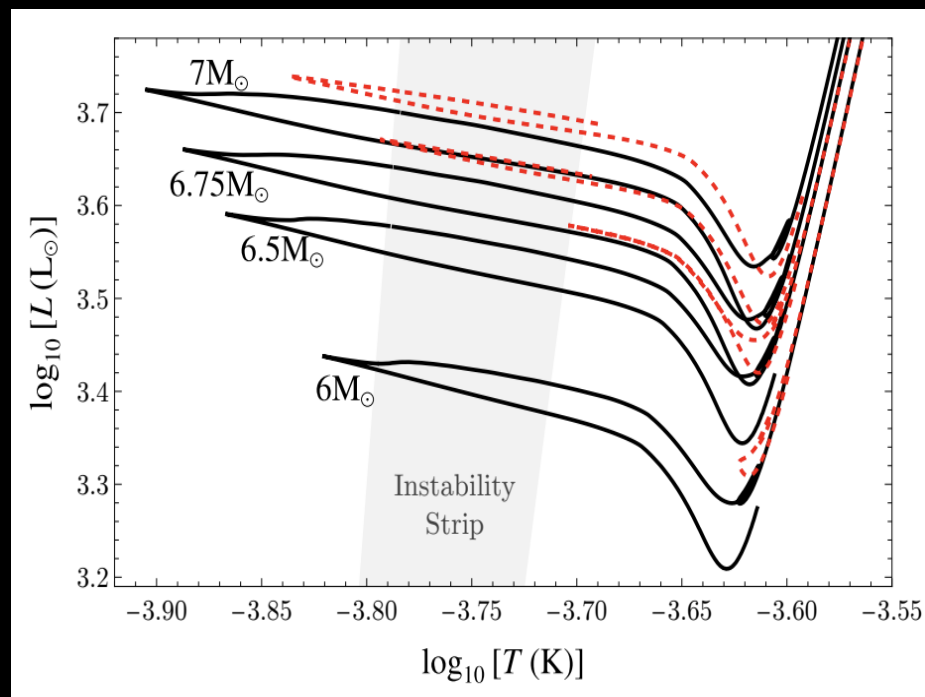
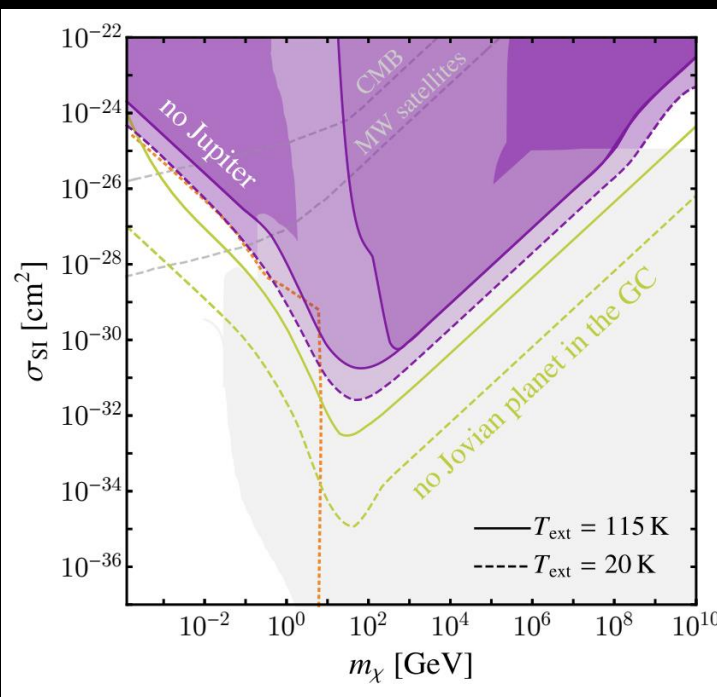
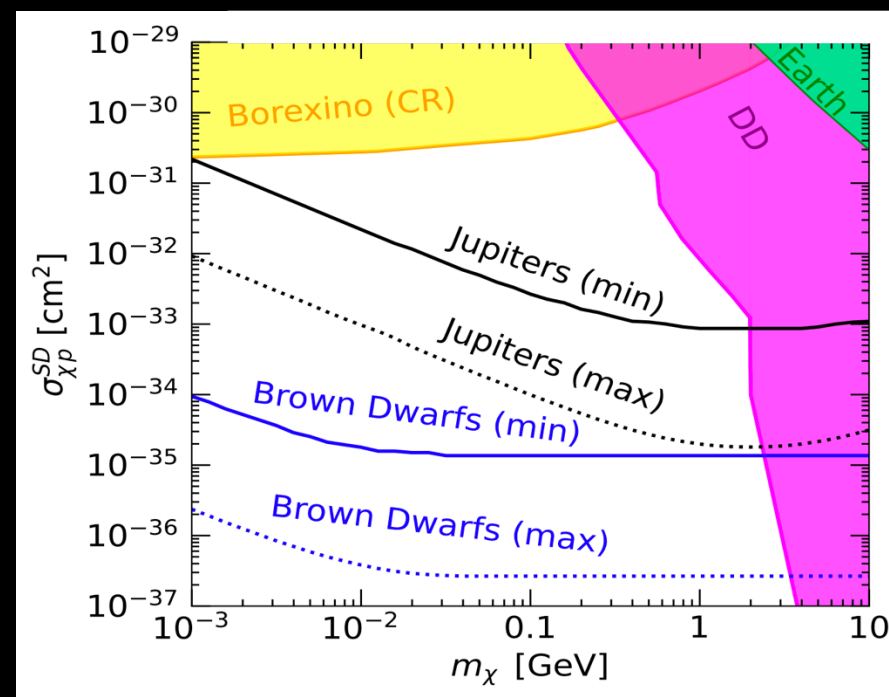


Light curve of Cepheid variable star V1 in galaxy M31



See also white dwarf: McCullough & Fairbairn, PRD 81 (2010) 083520

See also Hubble Tension's talk by Di Valentino



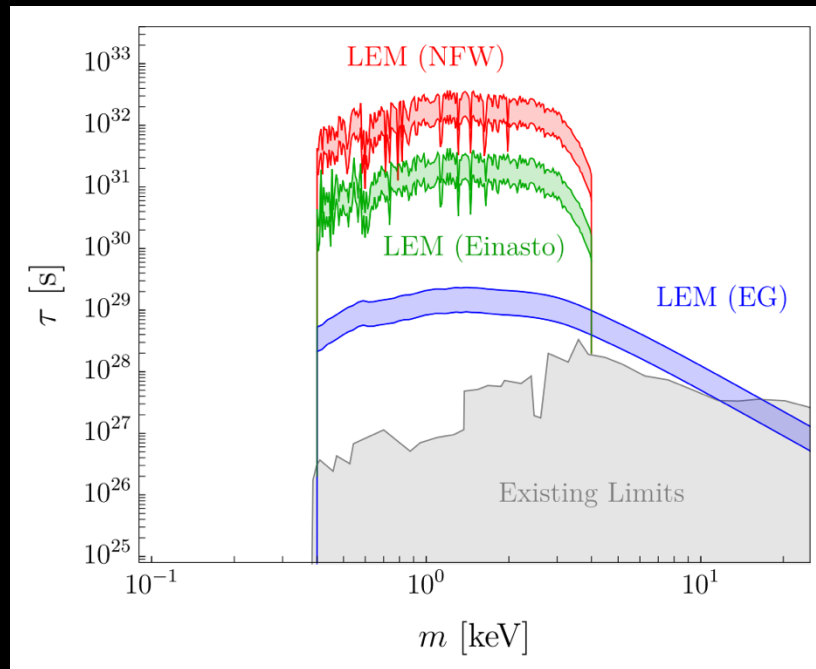
Leane & Smirnov, PRL 126 (2021) 16, 161101

Croon & Smirnov, JCAP 11 (2024), 046

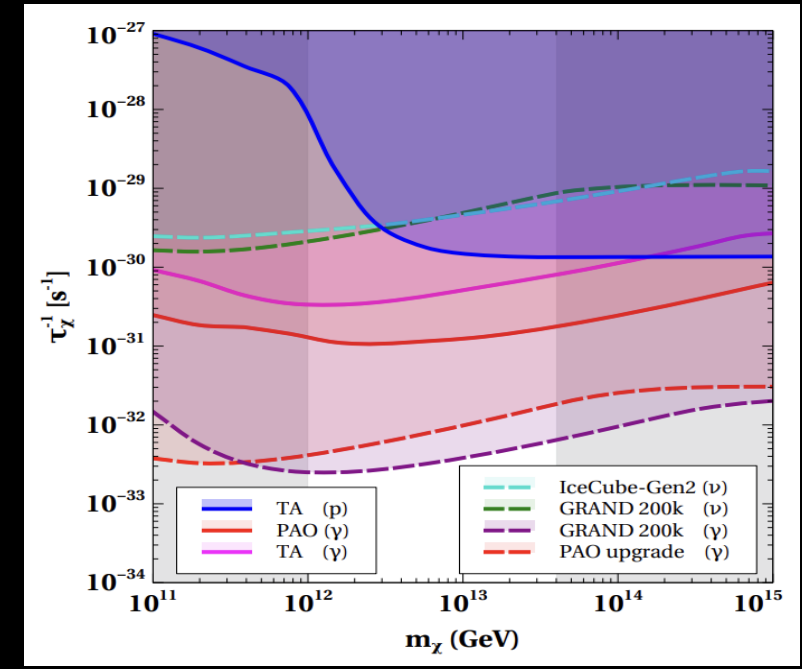
Croon et al, arXiv:2510.02641



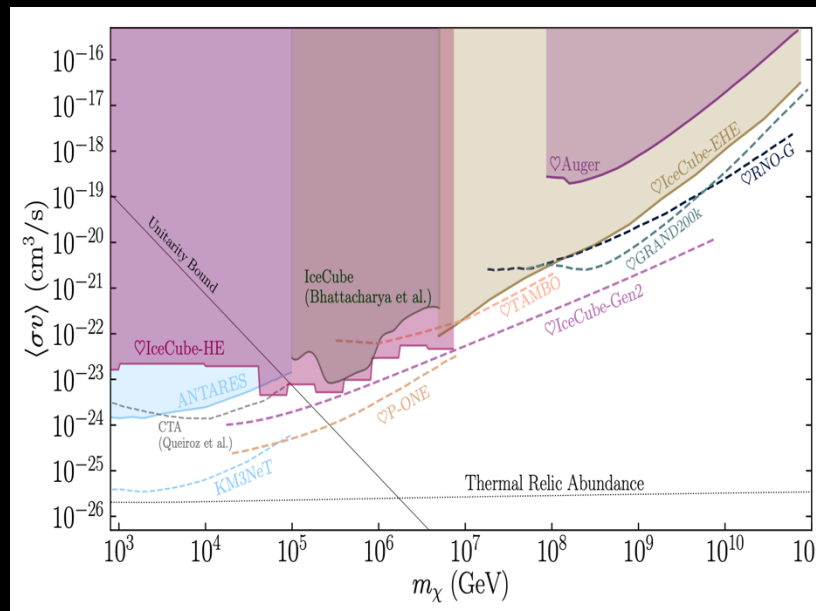
X-rays, neutrinos and cosmic rays



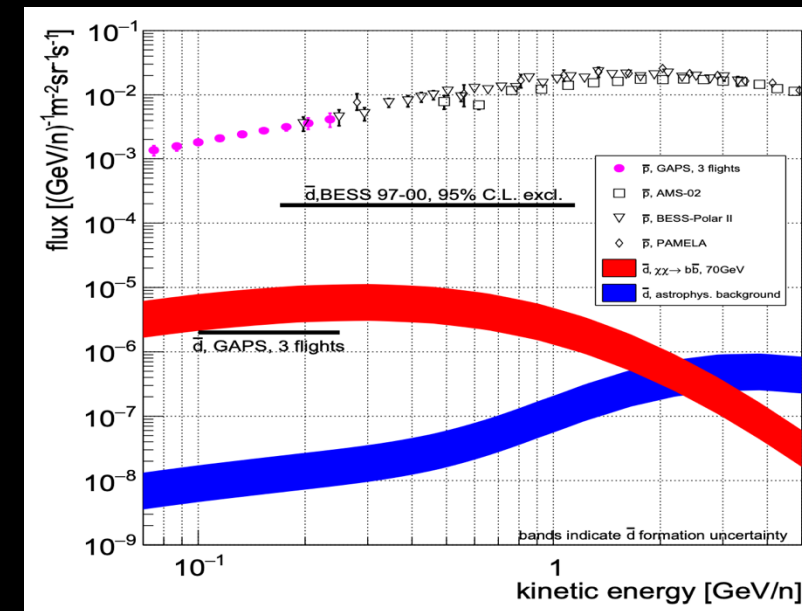
Krnjaic & EP, PRD 109 (2024) 9, 095027



Sarmah et al, PRD 111 (2025) 8, 083048

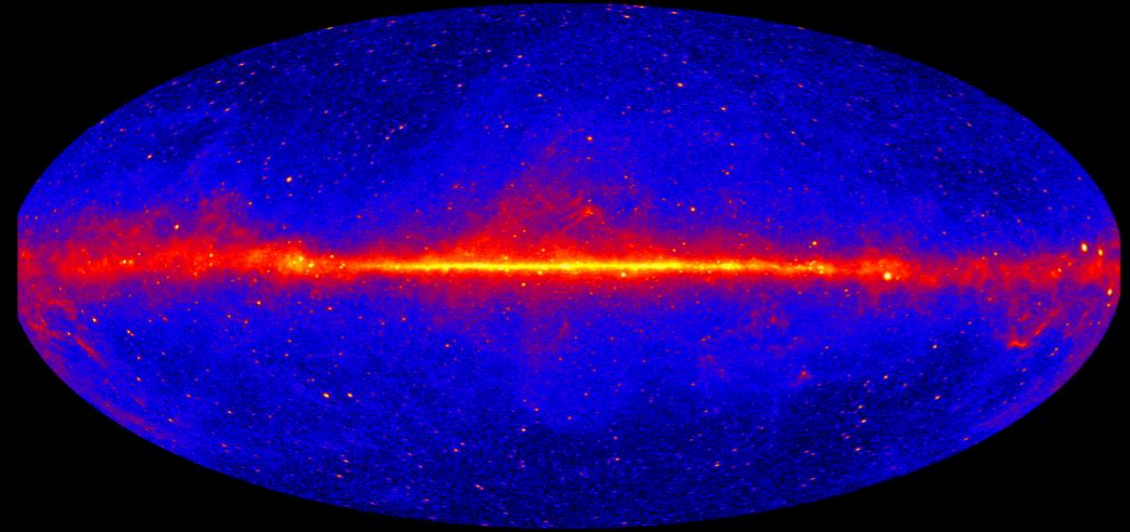
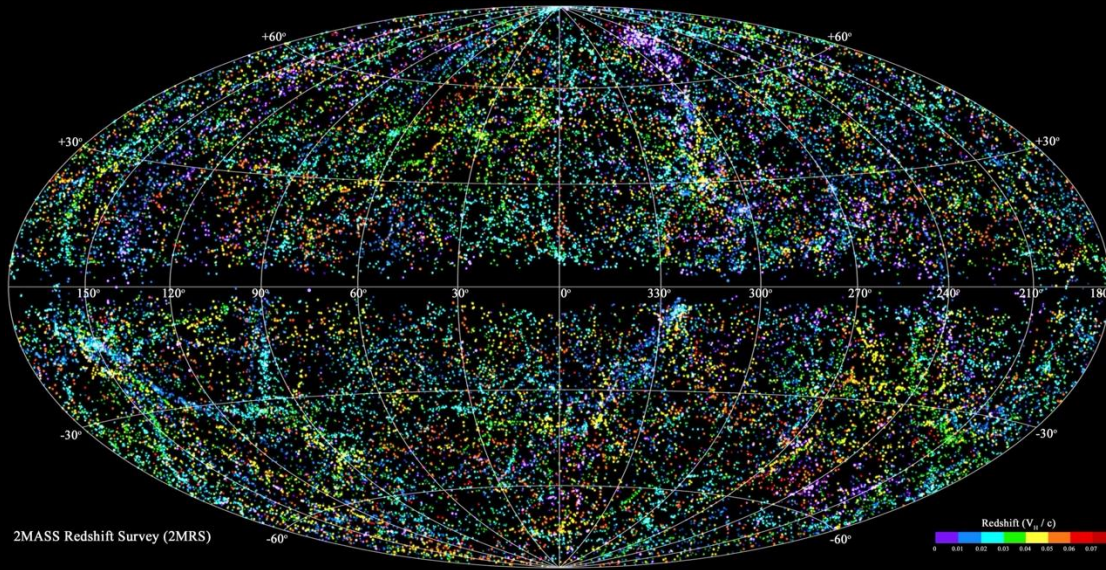


Argüelles et al, Rev. Mod. Phys. 93 (2021) 3, 035007



GAPS Collab, arXiv:2310.10181

Observational Synergies



MORE MATTER, I SENSE!



Gravitational tracers

Galaxy catalogues
Clusters catalogues
Neutral hydrogen
Weak lensing



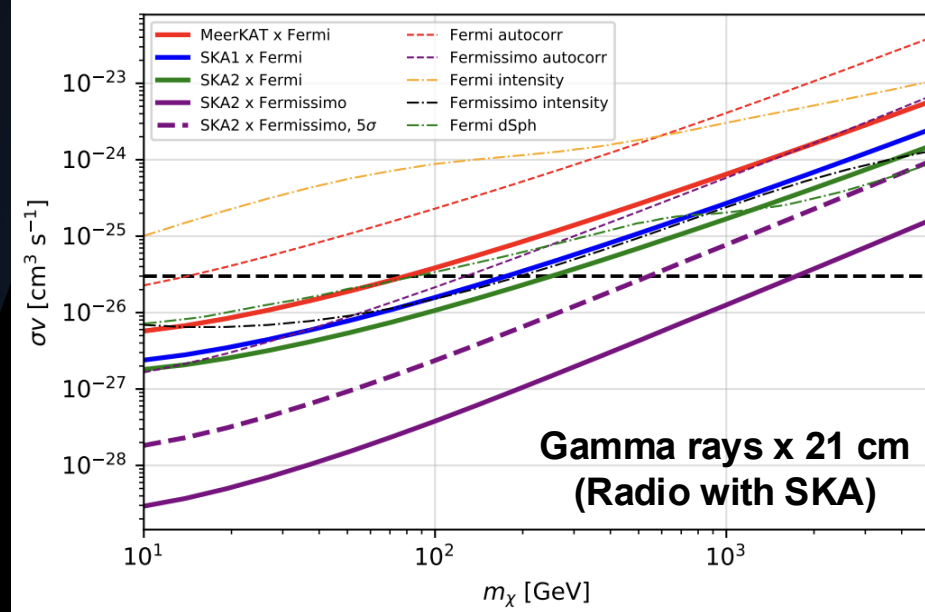
DM signals

- γ rays
- X rays
- IR emission
- Radio waves

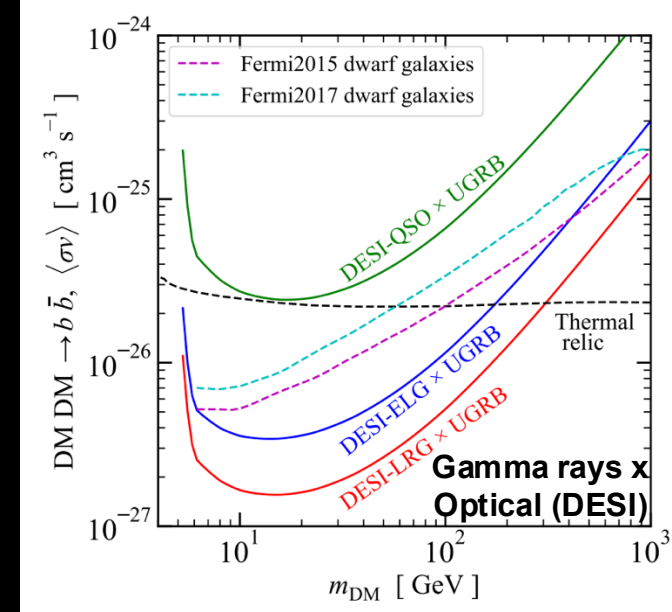


I'M A PARTICLE, BAZINGA!

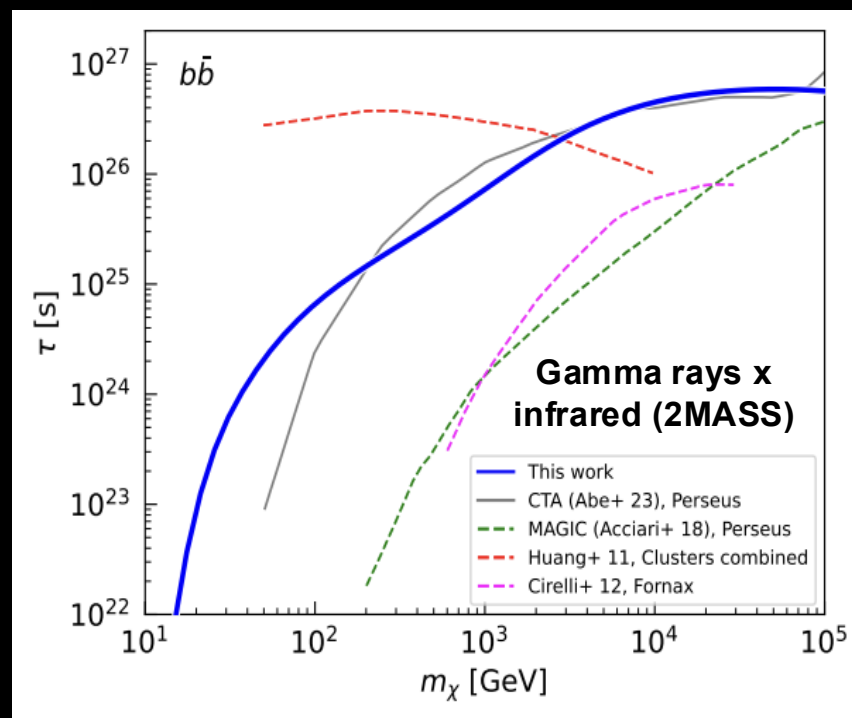
Observational Synergies



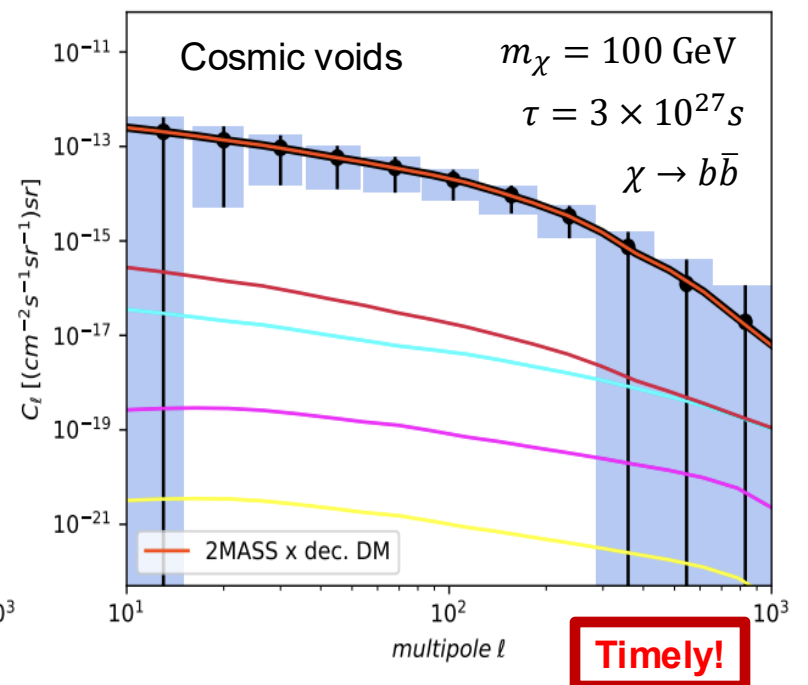
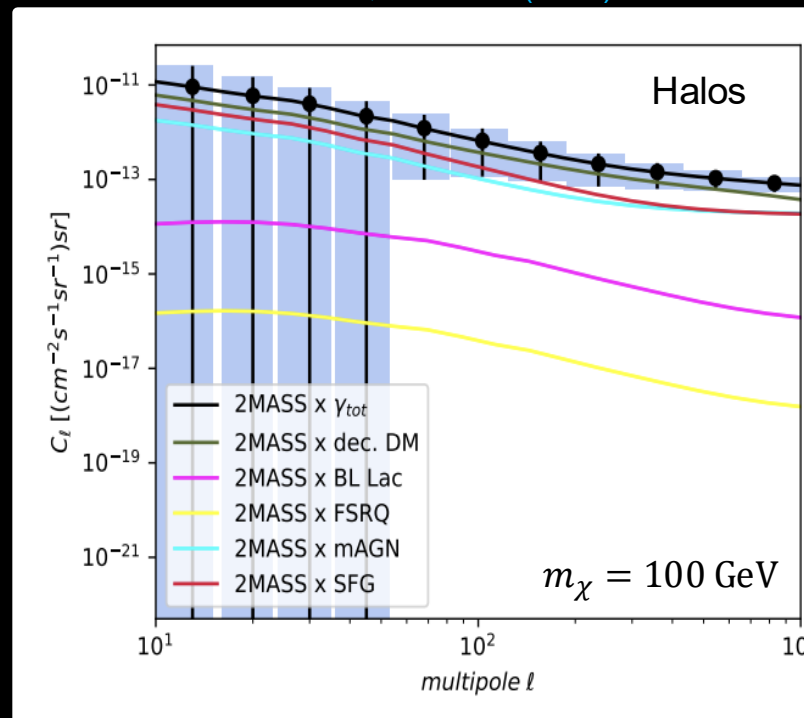
EP et al, JCAP 07 (2020) 044



Zhou, Bernal, EP et al, arXiv:2410.00375



EP et al, arXiv:2505.20383



Timely!

Arcari, EP, Fornengo, JCAP 11 (2022) 011

Summary & Conclusions

- 1 Indirect searches provide a powerful avenue to probe dark matter signals
- 2 Vibrant field with new ideas and detection strategies emerging at the intersection of particle physics, cosmology and astrophysics
- 3 Future prospects are exciting with upcoming telescopes set to deliver unprecedented sensitivity!



Summary & Conclusions

*Thank you for
your attention!*

- ① Indirect searches provide a powerful avenue to probe dark matter signals
- ② Vibrant field with new ideas and detection strategies emerging at the intersection of particle physics, cosmology and astrophysics
- ③ Future prospects are exciting with upcoming telescopes set to deliver unprecedented sensitivity!

