

Moduli Stabilisation in the Presence of Matter Fields

Lisa Hall

University of Sheffield, UK

work in progress with
Carsten van de Bruck
and Ki-Young Choi

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- Uplifting of AdS vacua

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 - Kachru,Kallosh,Linde,Trivedi (2003):

$$\delta V = \frac{k}{\text{Re}(T)^2}$$

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 - Kacrh, Kallosh, Linde, Trivedi (2003):
- D-term

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 - Burgess, Kallosh, Quevedo (2003)
$$\delta V = \frac{1}{2g^2} D^2$$
 - Achúcarro, deCarlos, Casas, Doplicher (2006)

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- Matter Superpotentials: (F-term uplifting)
 - Lebedev,Nilles,Ratz (2006)

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- Stabilisation from additional matter field
 - (with background fluid)

$$\delta V = \frac{k}{\text{Re}(T)^2}$$

$$\delta V = \frac{1}{2g^2} D^2$$

KKLT (Kachru, Kalosh, Linde, Trivedi) Model

Kähler potential:

$$K = -3 \ln (T + \bar{T})$$

integrate out
complex-structure
moduli

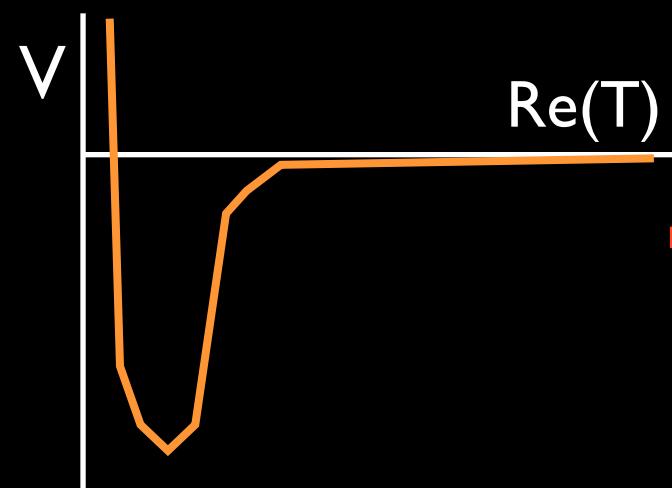
Superpotential:

$$W = W_0 + A e^{-aT}$$

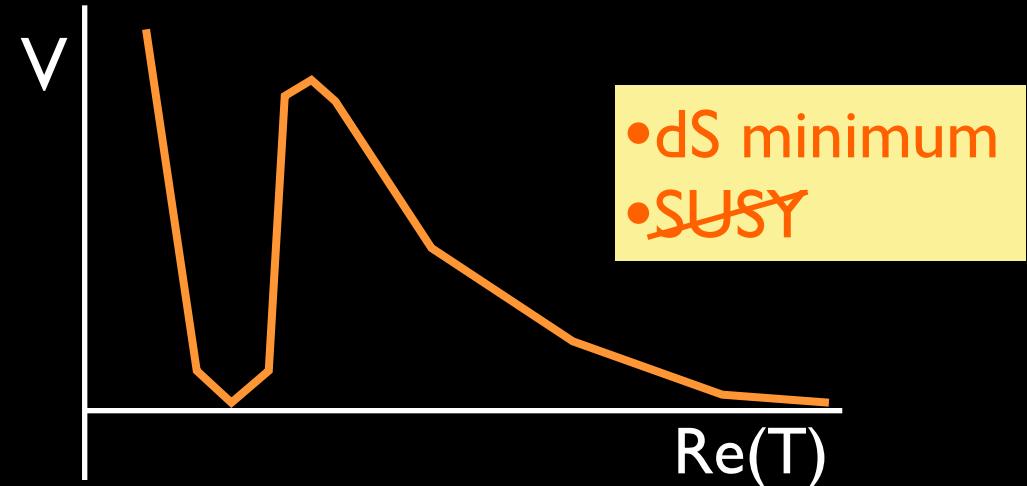
non-perturb.
corrections
e.g. Witten (1996)

$$V = e^K \left(K^{i\bar{j}} D_i W \overline{D_j W} - \underline{3W\bar{W}} \right)$$

AdS minimum



Add $\overline{\text{D3}}$ -brane: $\delta V = \frac{k}{\text{Re}(T)^2}$



KKLT and the Polonyi Model

$$K = -3 \ln(T + \bar{T}) + |C|^2$$

$$V_0 \sim \epsilon \mu^4 \quad c \sim \mu^2 \left(2 - \sqrt{3} - \frac{\sqrt{3}}{6} \epsilon \right)$$

$$C \sim \sqrt{3} - 1 - \frac{\sqrt{3} - 3}{6} \epsilon$$

Lebedev,Nilles,Ratz (2006)

Lebedev,Löwen,Mambrini,Nilles,Ratz (2006)

$$\mathcal{W}(C) = c + \mu^2 C$$

$$\mathcal{W}(T) = W_0 + A e^{-\alpha T}$$

**modulus heavy
compared to
Polonyi field**

$$\mu^2 \sim W_0$$

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$$C = C_r + iC_i$$

$$T = T_r + iT_i$$

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$$\phi = \sqrt{\frac{3}{2}} \ln T_r$$

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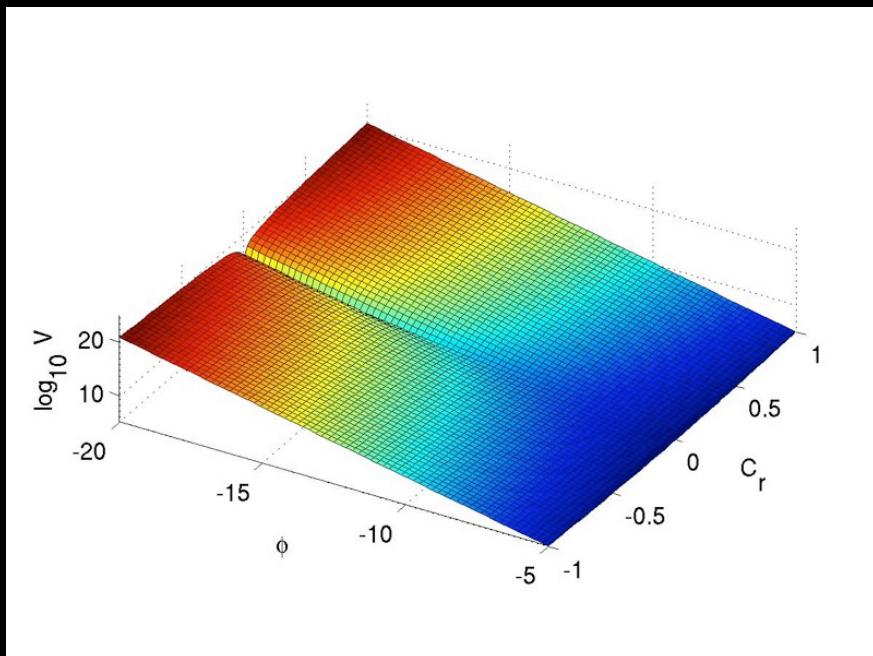
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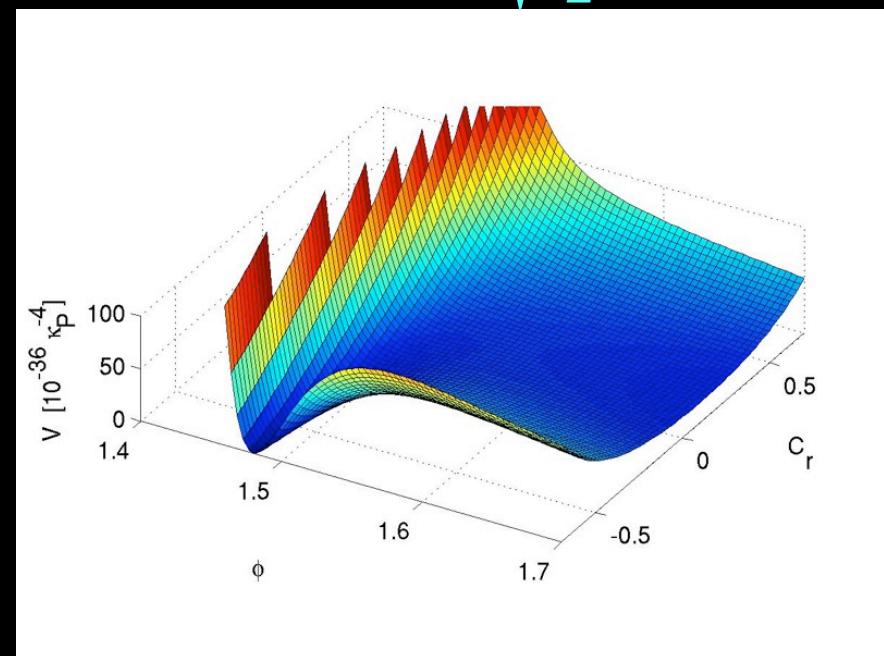
modulus heavy
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$$\mu^2 \sim W_0$$

$$\phi = \sqrt{\frac{3}{2}} \ln T_r$$



$$\begin{aligned} A &= 1 \\ \alpha &= 12 \\ \mu^2 &= 10^{-8} \end{aligned}$$



Lebedev,Nilles,Ratz (2006)

Lebedev,Löwen,Mambrini,Nilles,Ratz (2006)

Single Field Dynamics

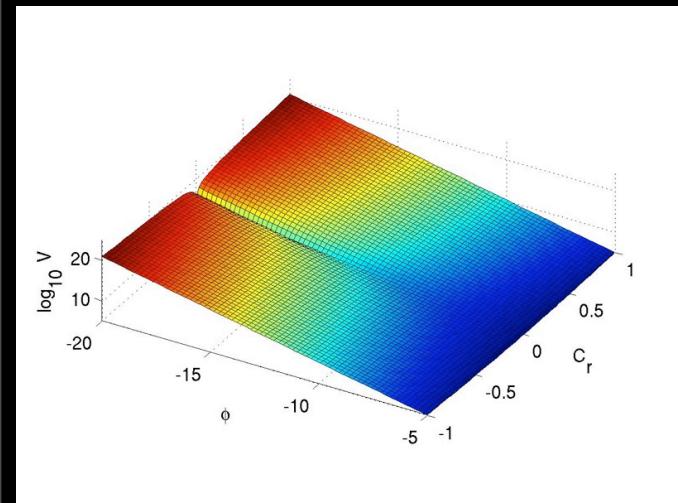
Kaloper,Olive (1993)

Barreiro, de Carlos, Nunes (1998)

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$$C_i = T_i = 0$$

$$C_r(N=0) = 0$$



Single Field Dynamics

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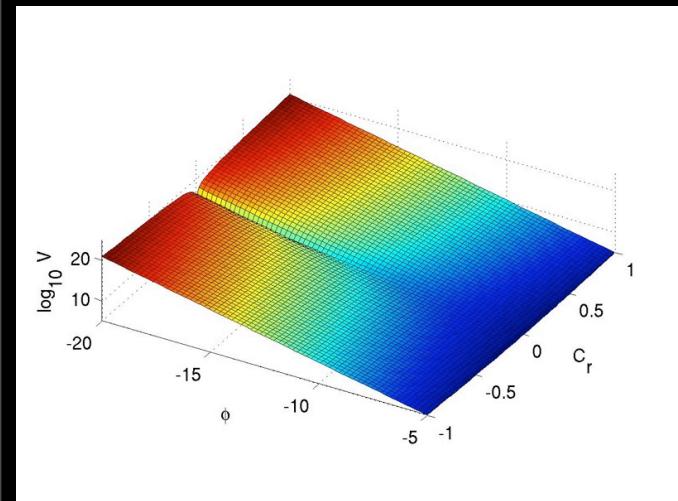
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$$\phi(N=0) = -15$$



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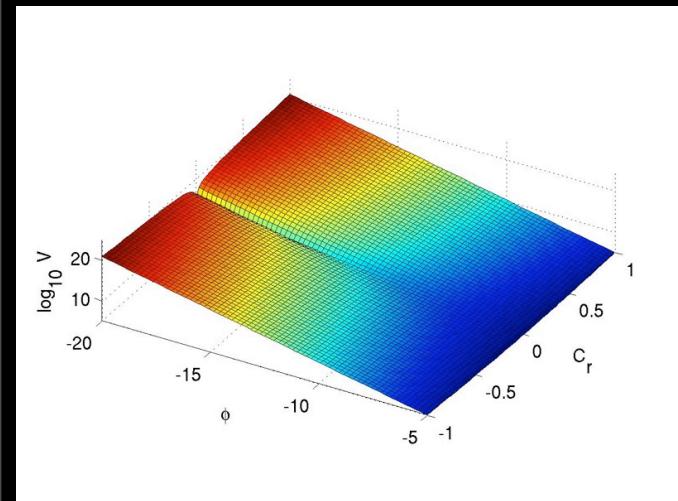
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$$\Omega_b(N=0) = 0.93$$

matter
background field



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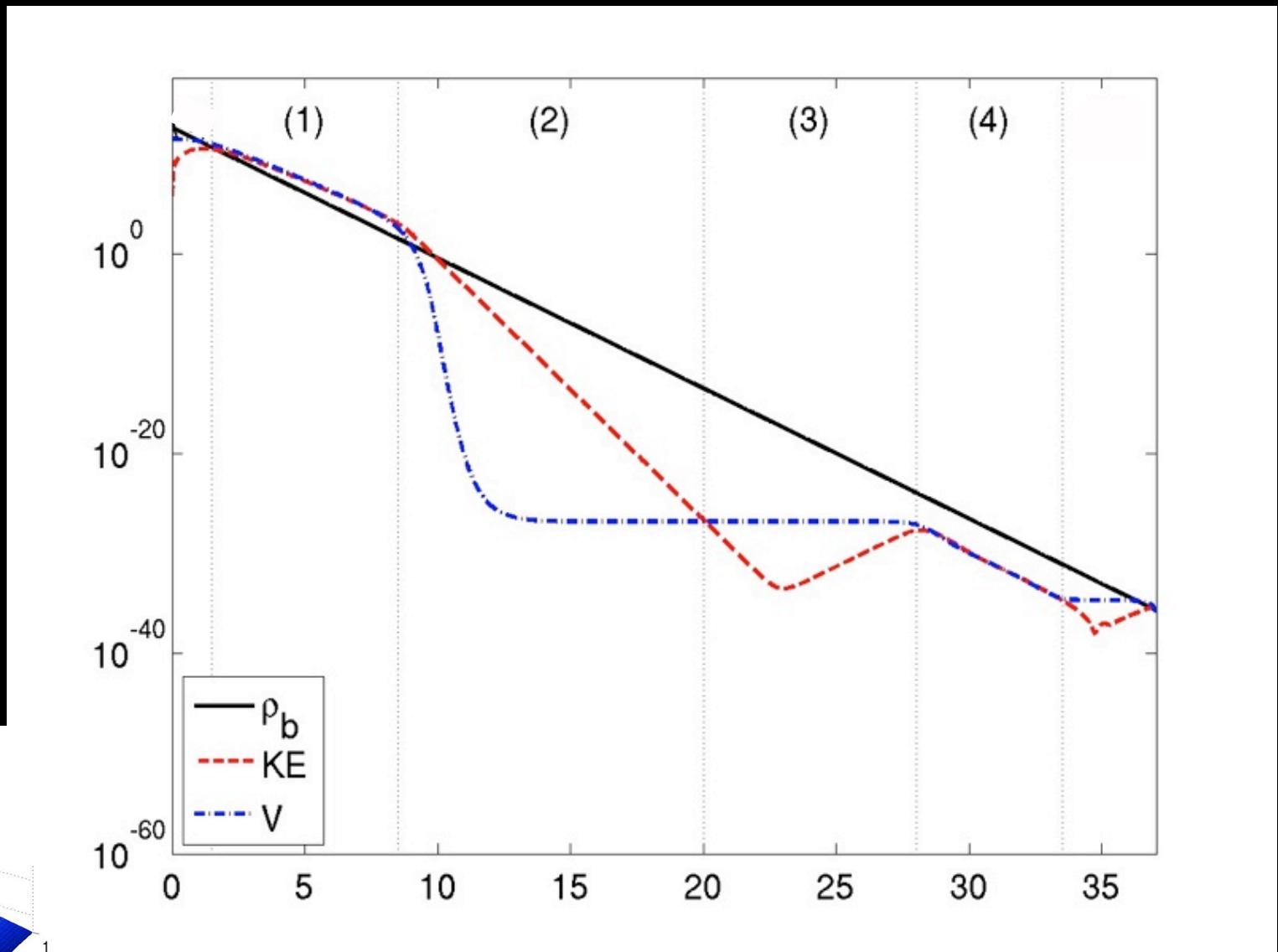
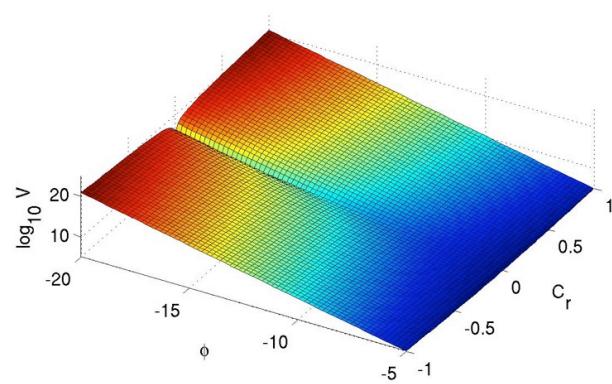
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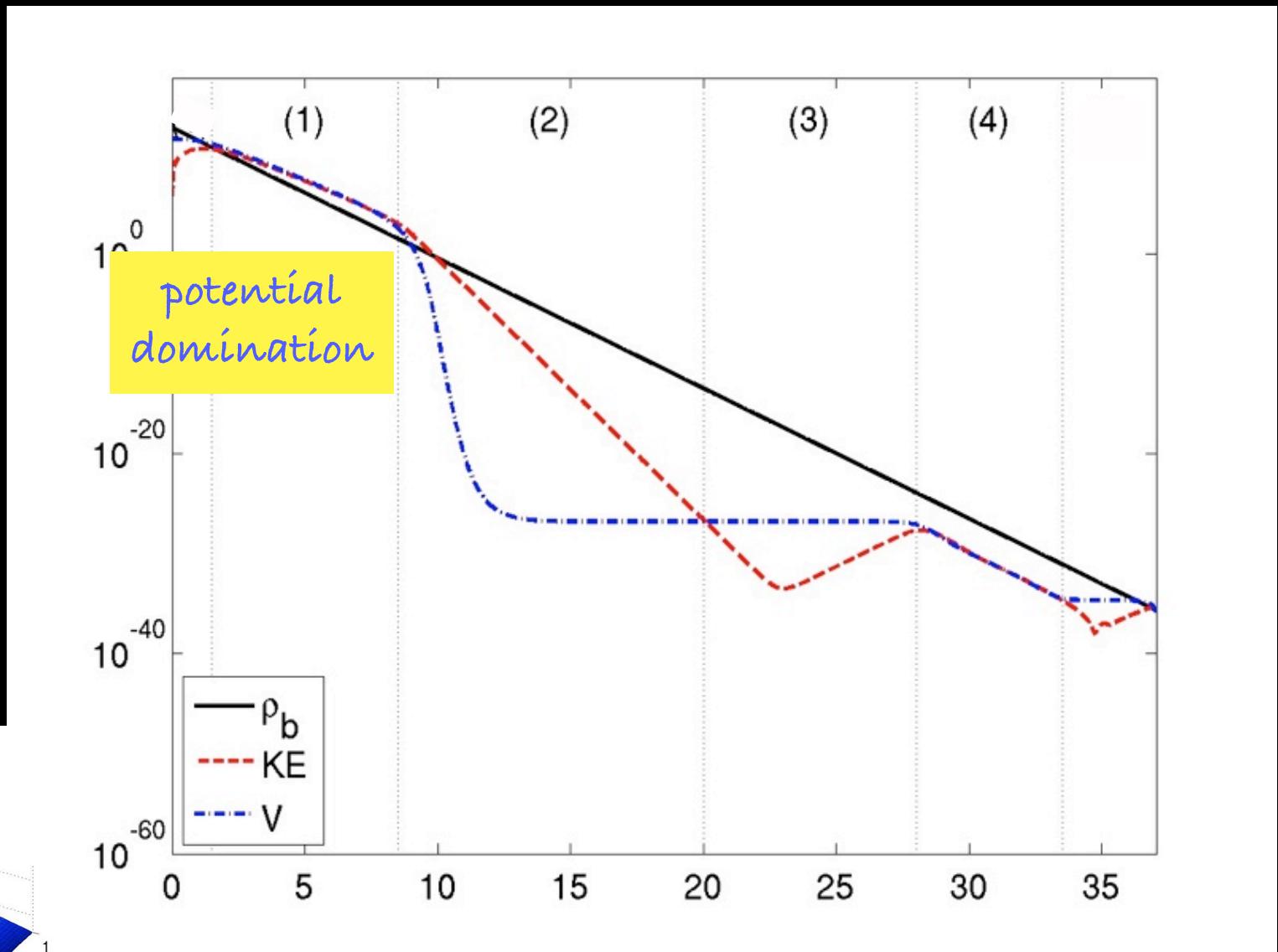
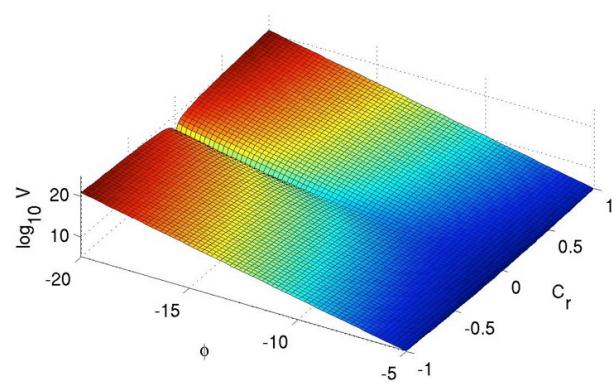
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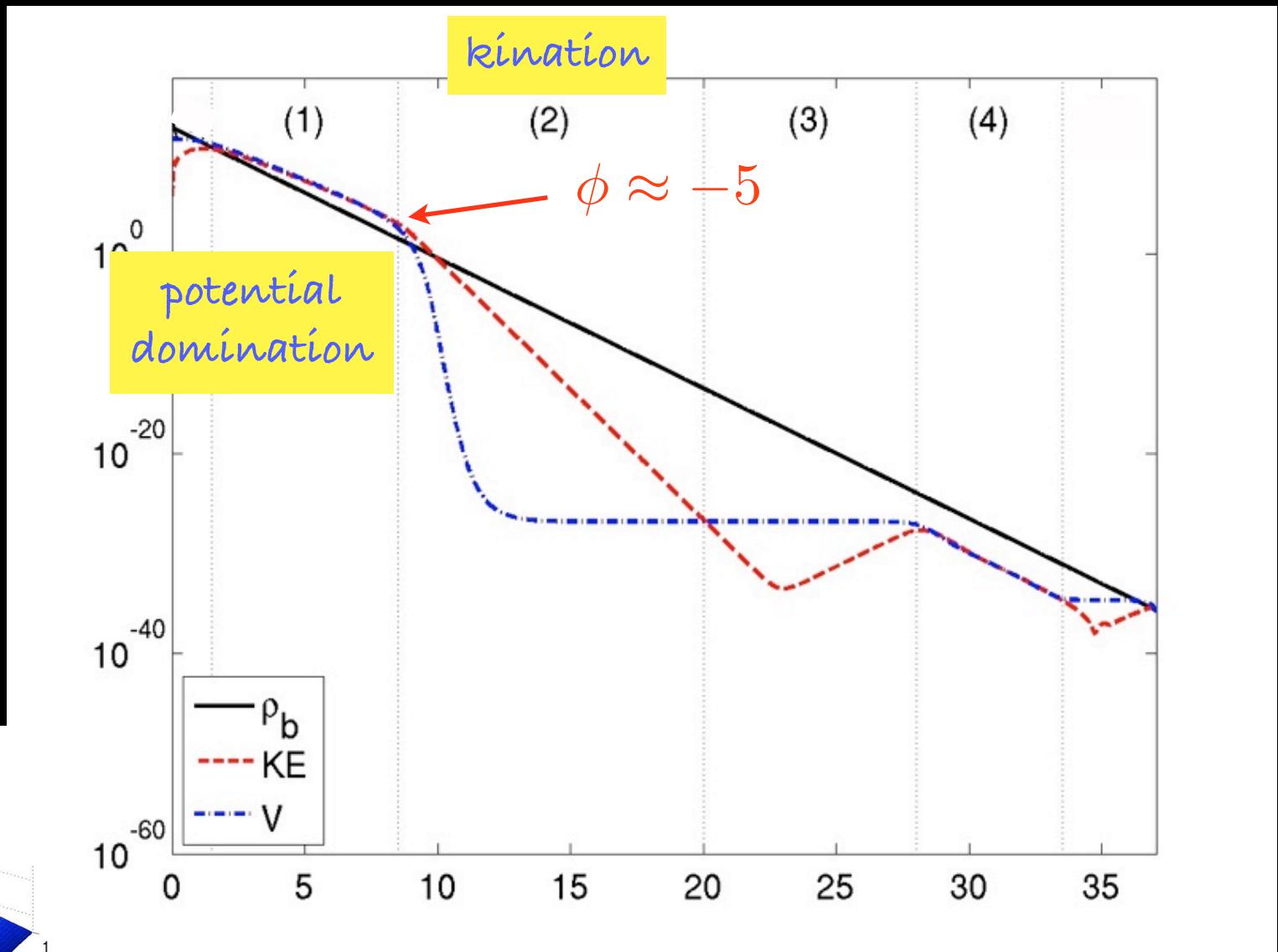
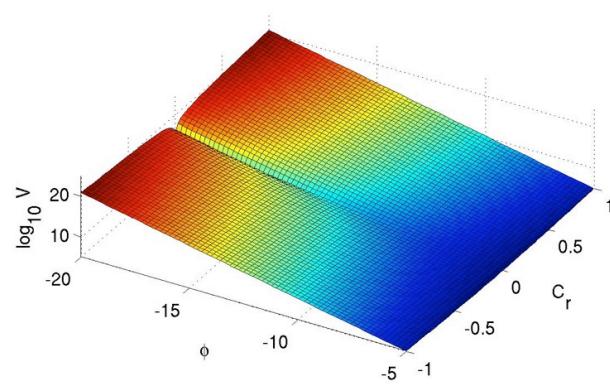
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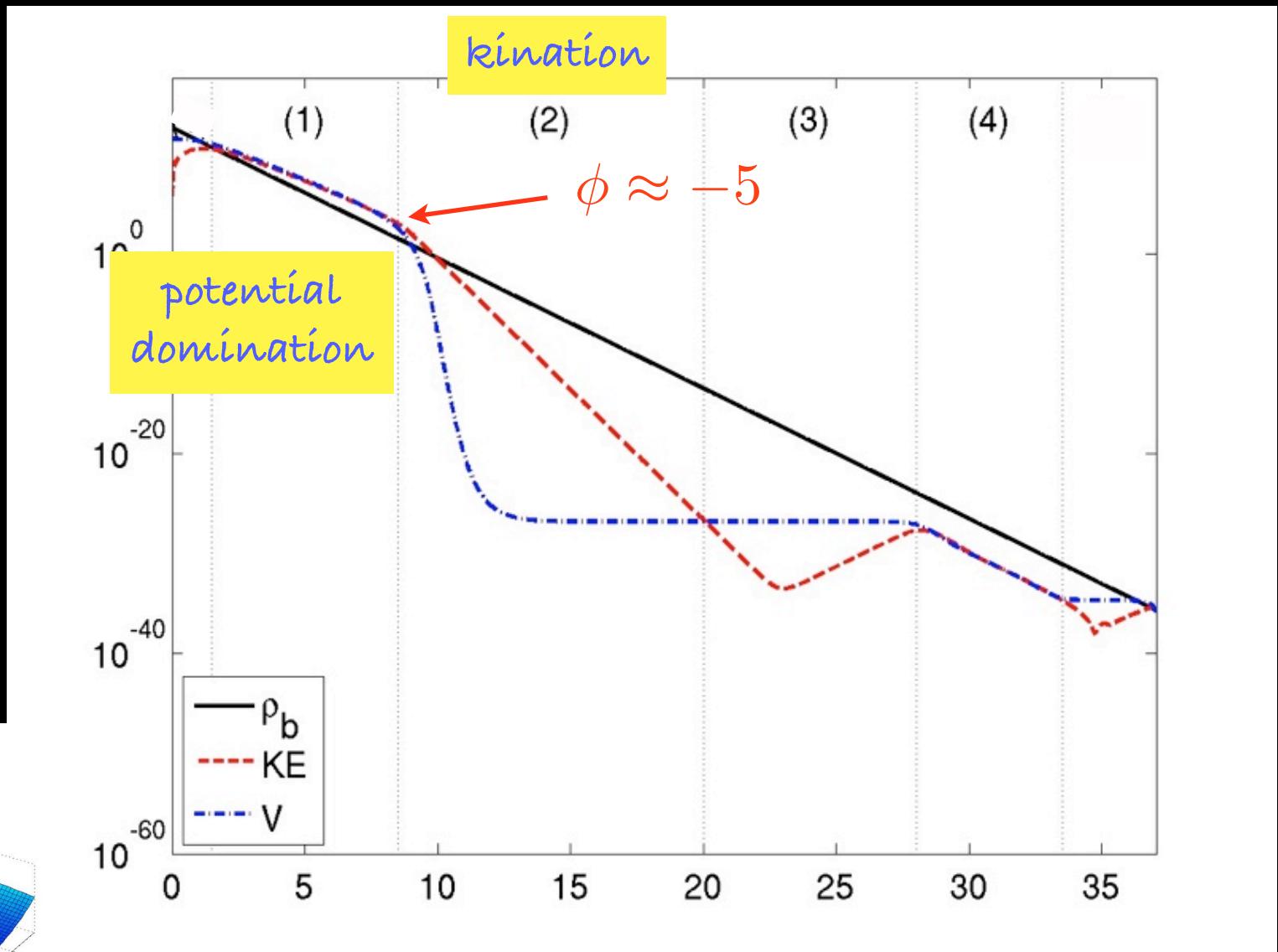
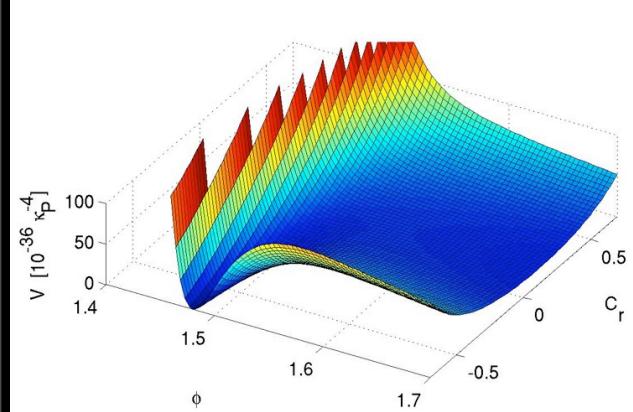
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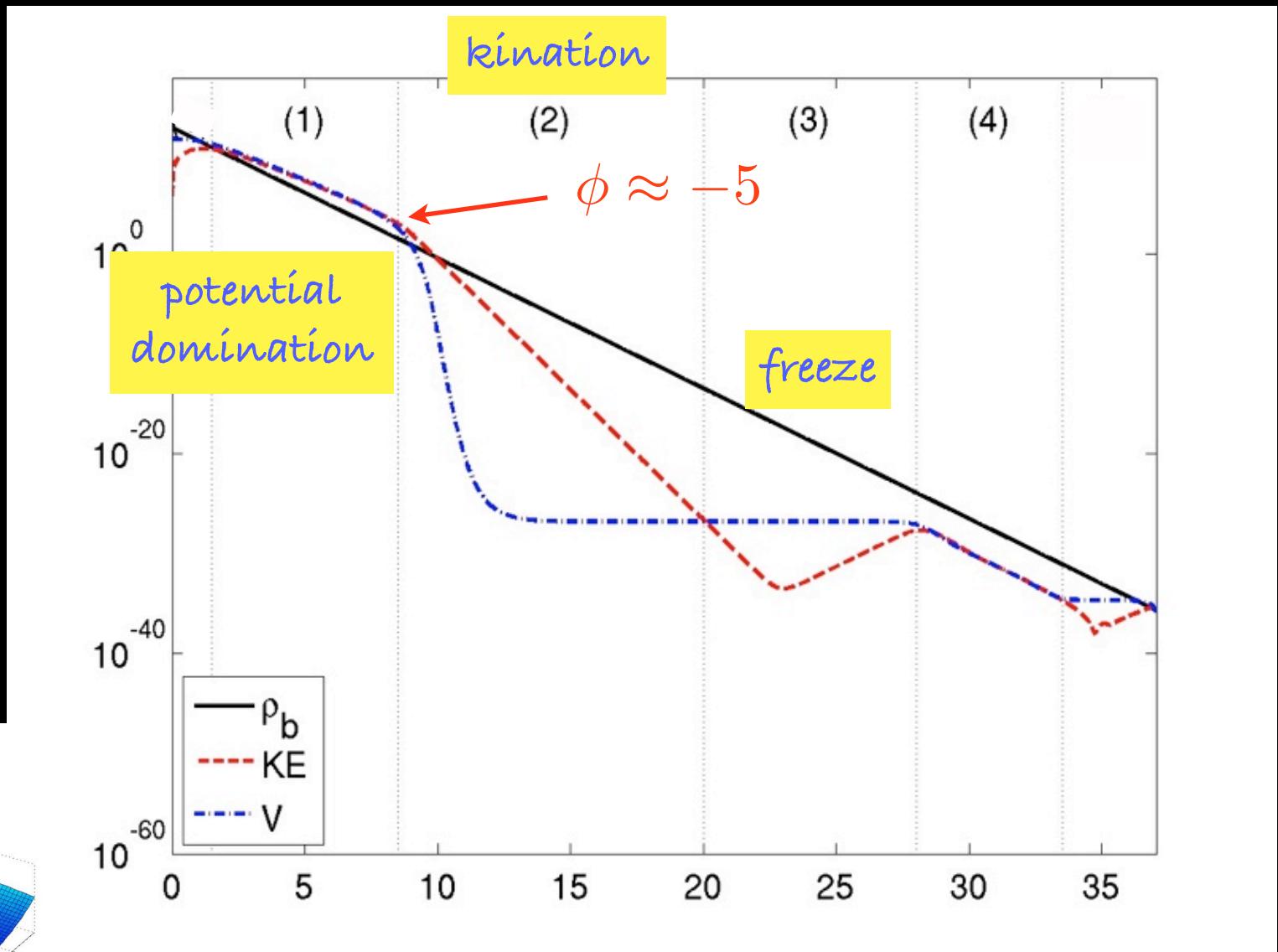
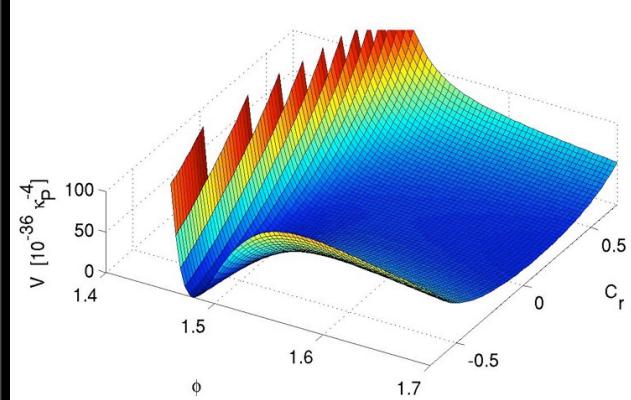
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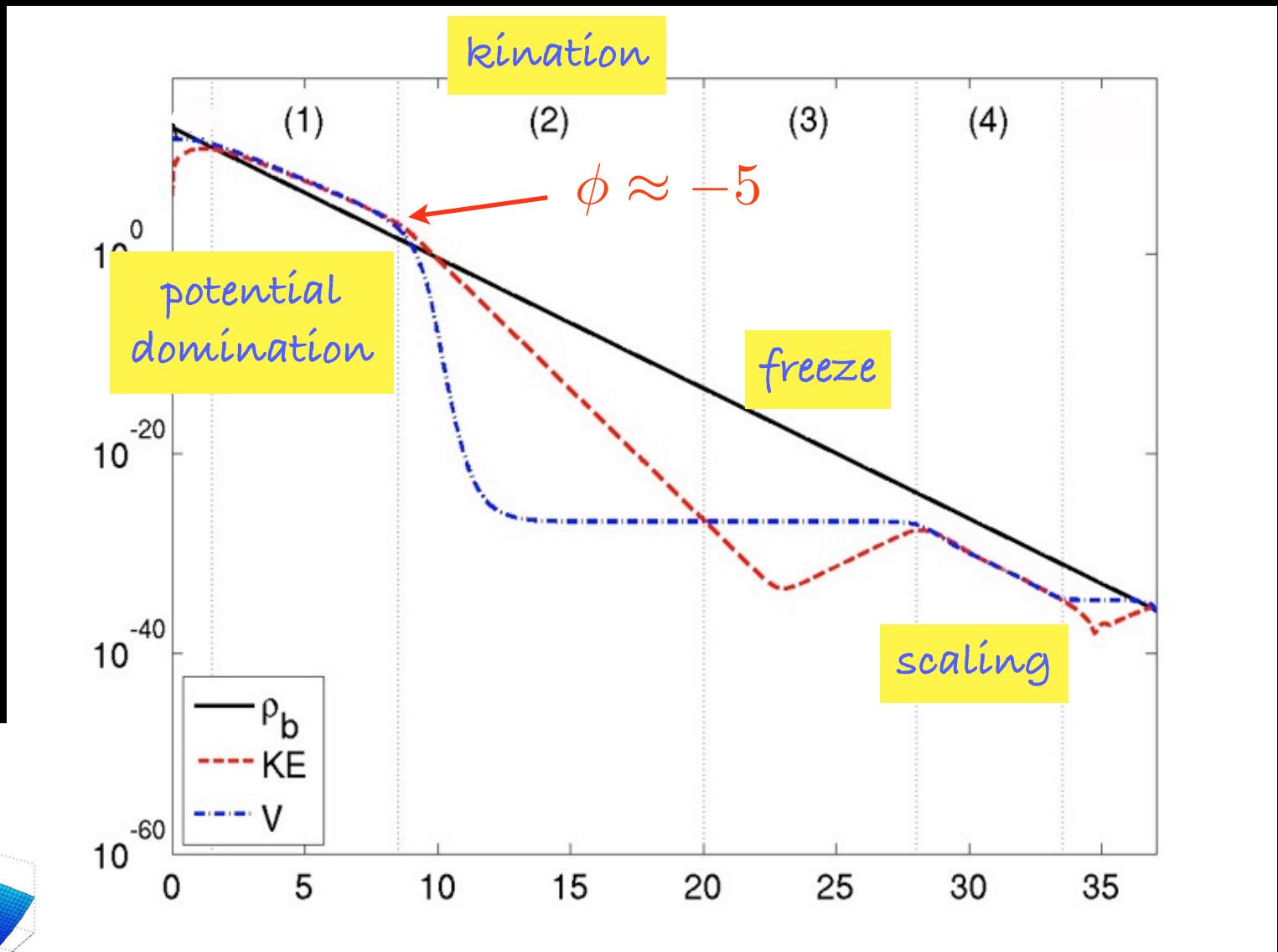
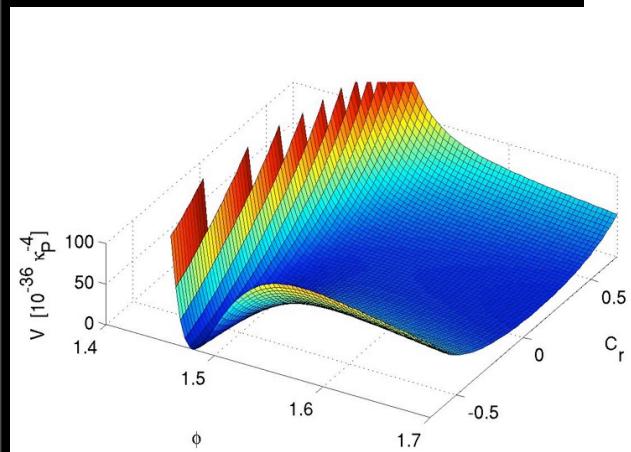
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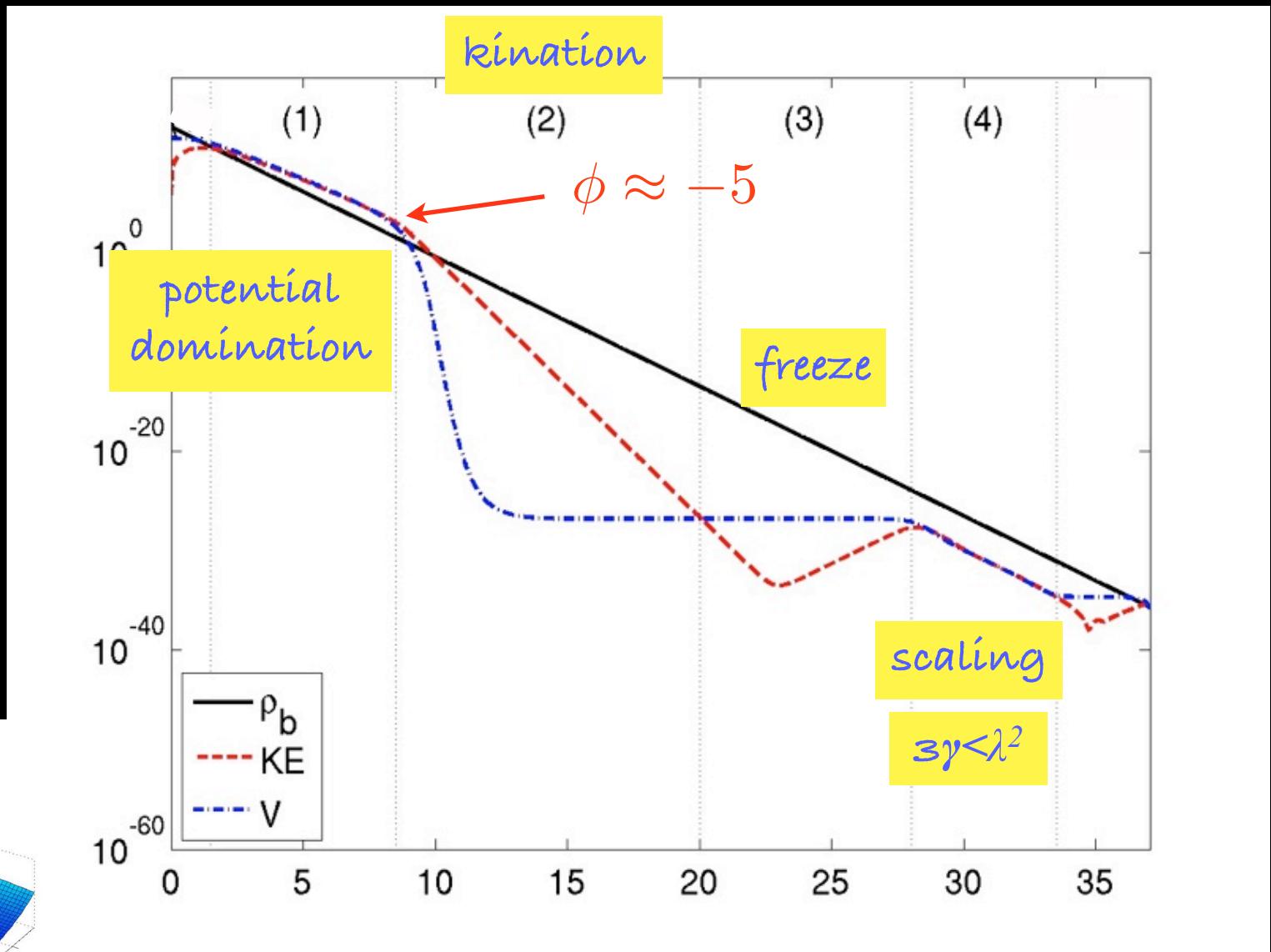
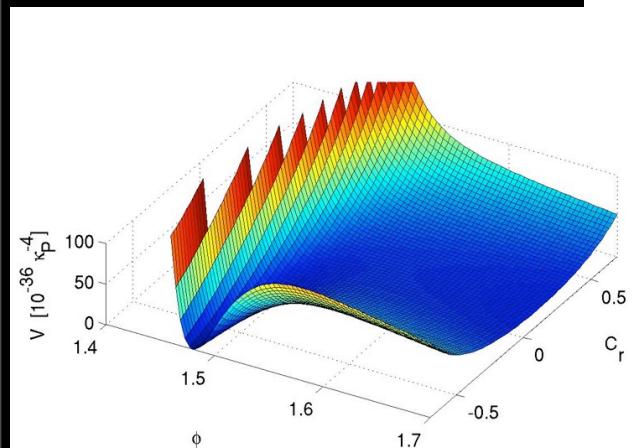
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$$\lambda \equiv -\frac{1}{V} \frac{\partial V}{\partial \phi}$$

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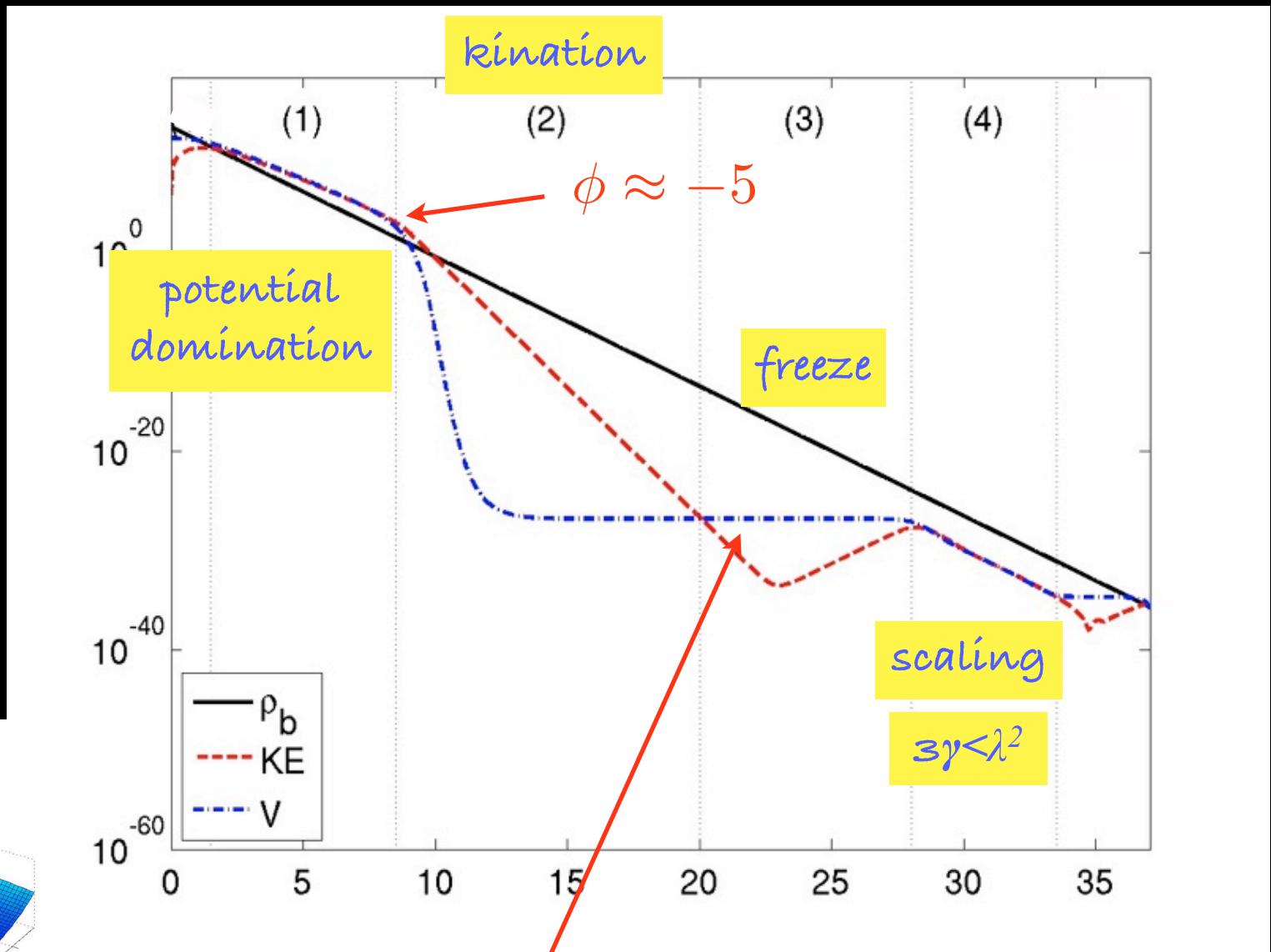
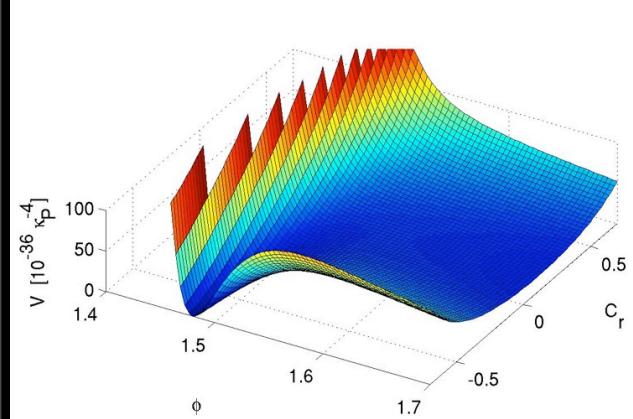
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stabilisation depends on freeze point

$$\lambda \equiv -\frac{1}{V} \frac{\partial V}{\partial \phi}$$

$$C_i = T_i = 0$$

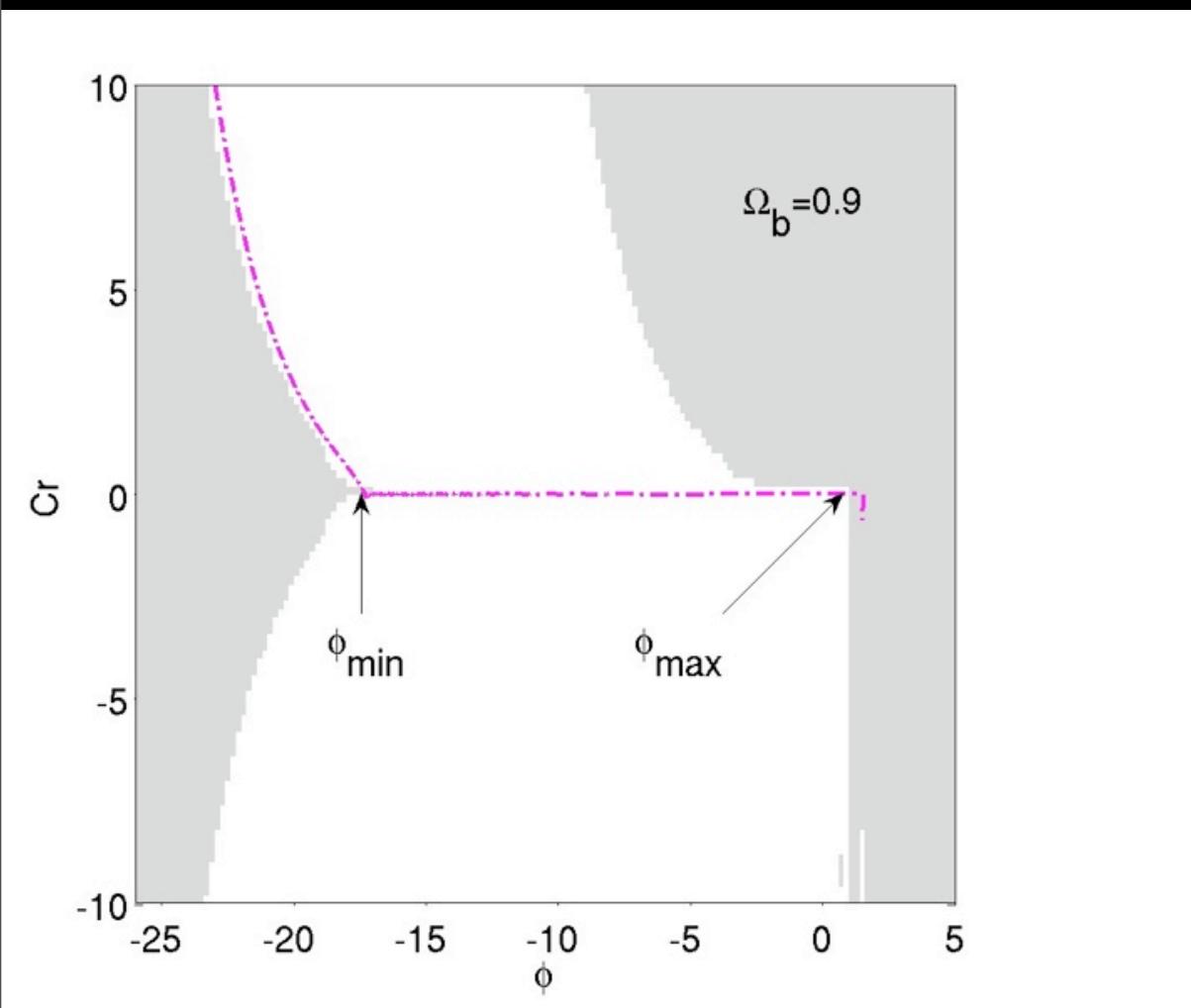
Two REAL Fields

$$\Omega_{\mathrm{b}} = 0.9$$

$C_i = T_i = 0$

Two REAL Fields

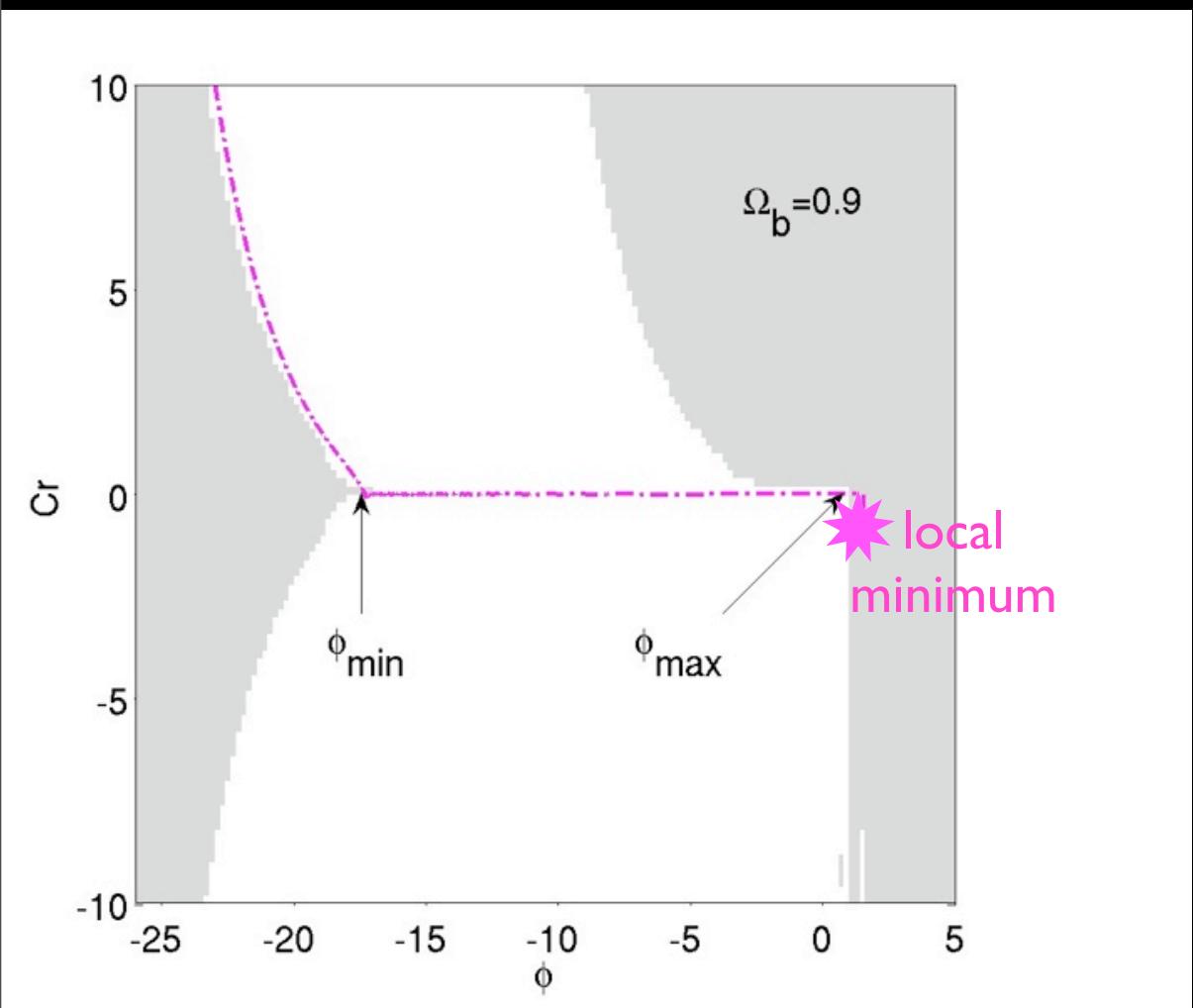
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$C_i = T_i = 0$

Two REAL Fields

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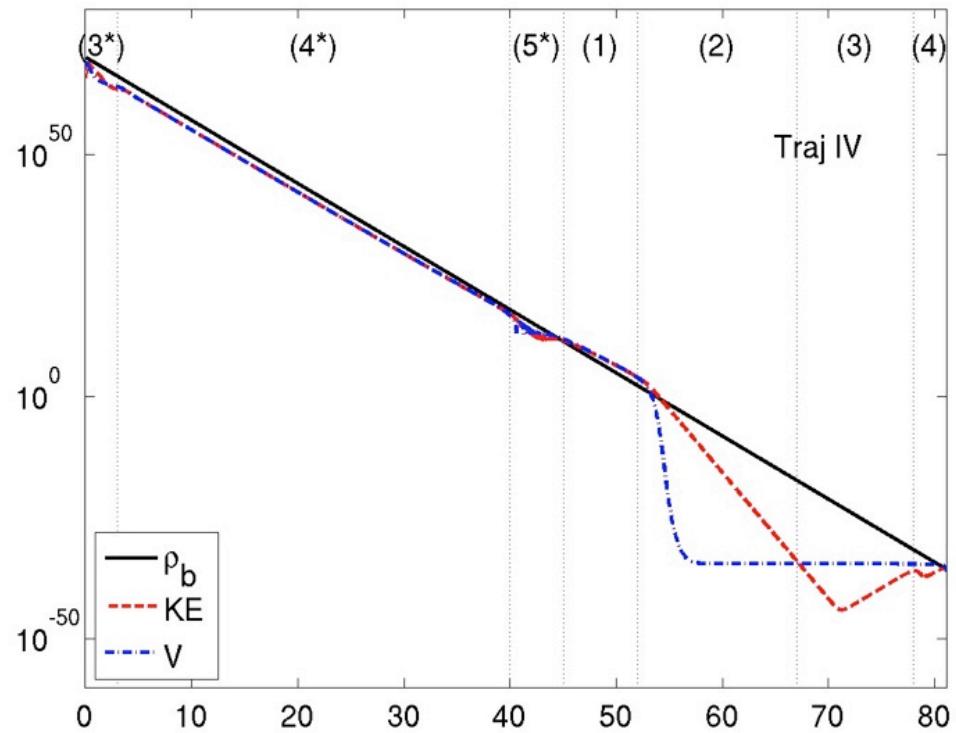
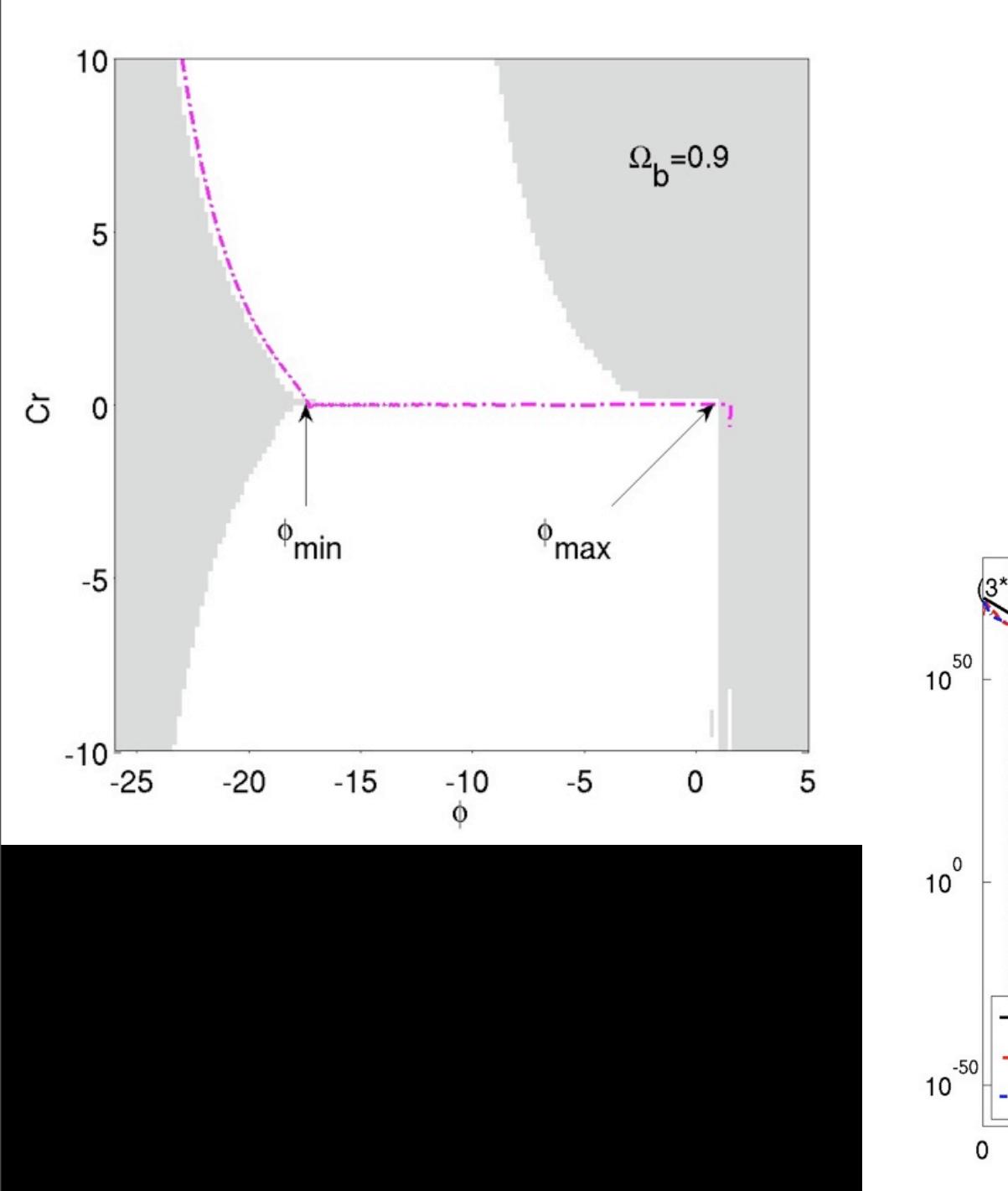


$$C_i = T_i = 0$$

Two REAL Fields

$$\Omega_b = 0.9$$

1. Potential dom.
2. Kination
3. Freeze out
4. Scaling
5. Oscillation

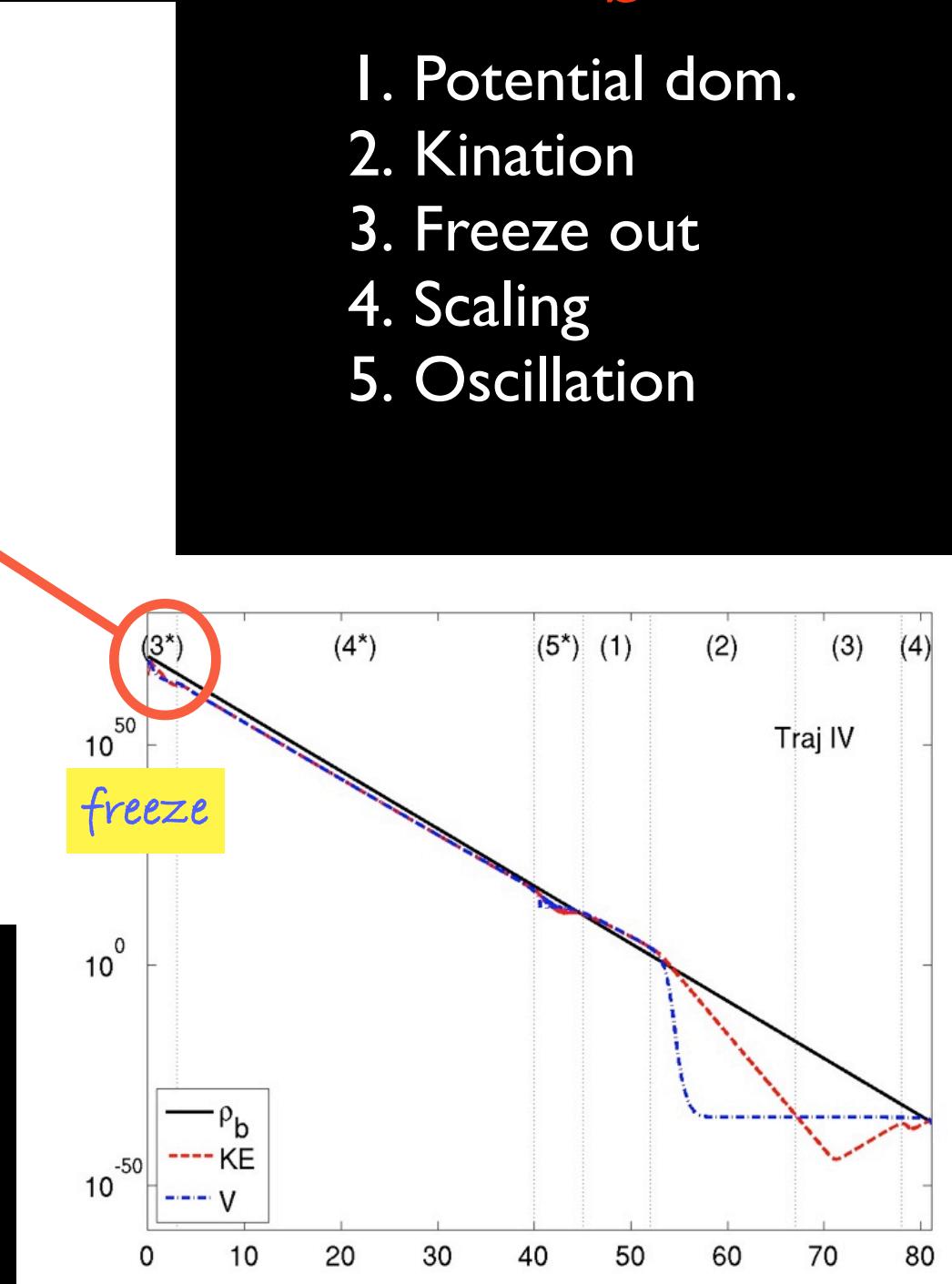
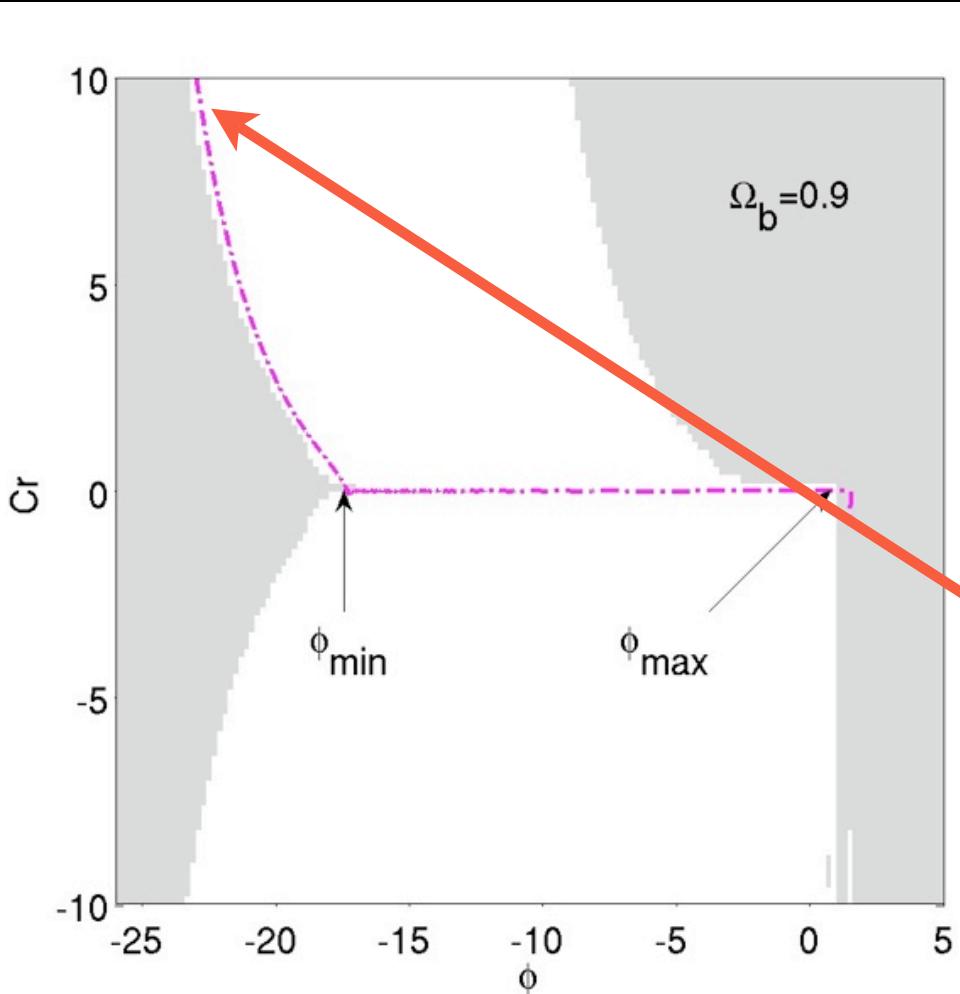


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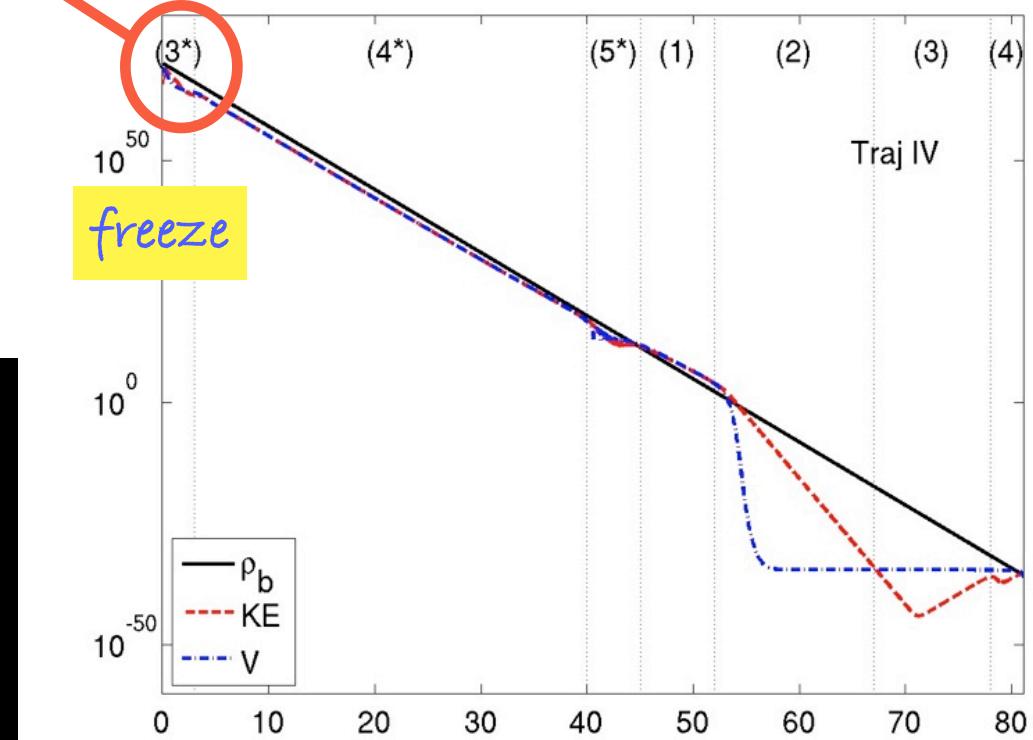
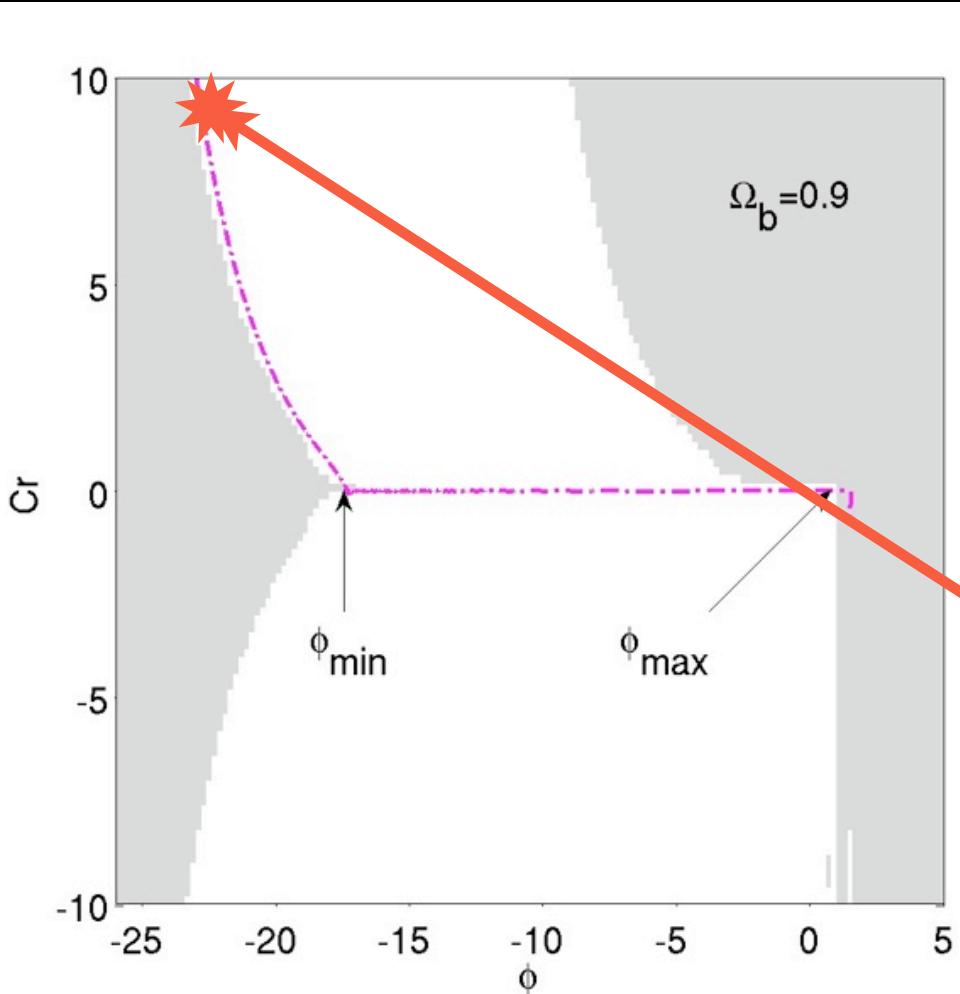


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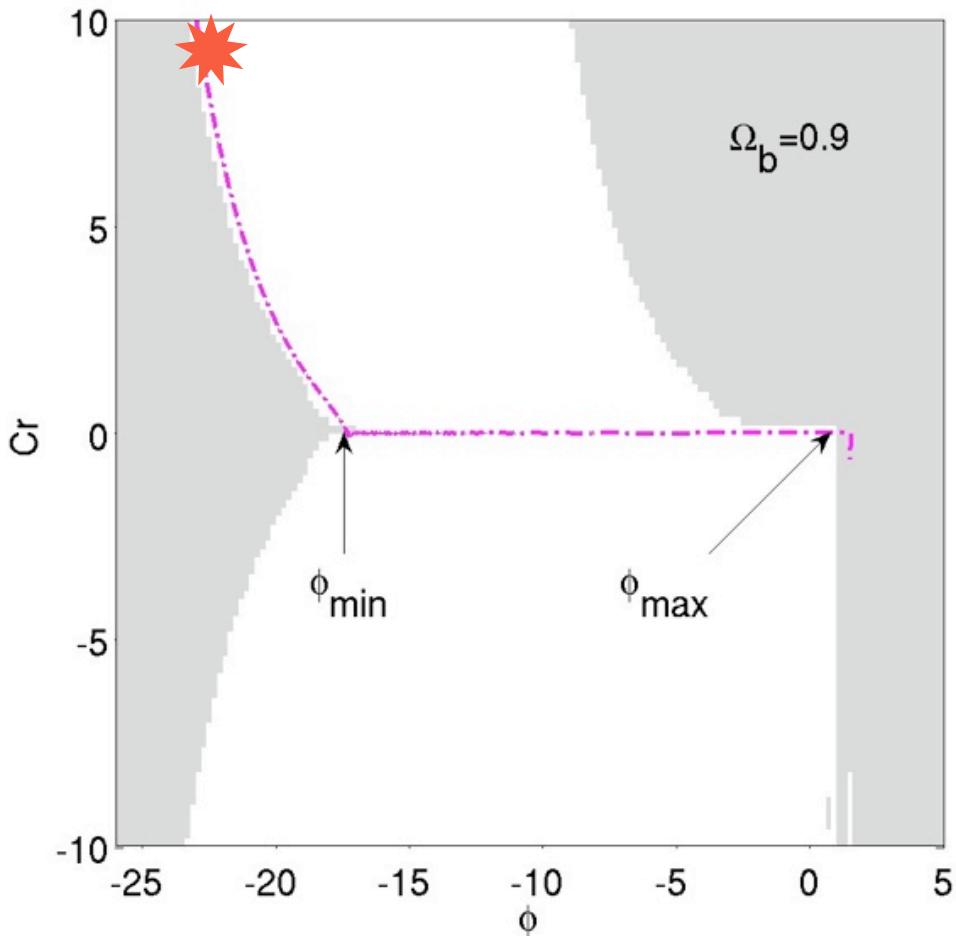
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Two REAL Fields

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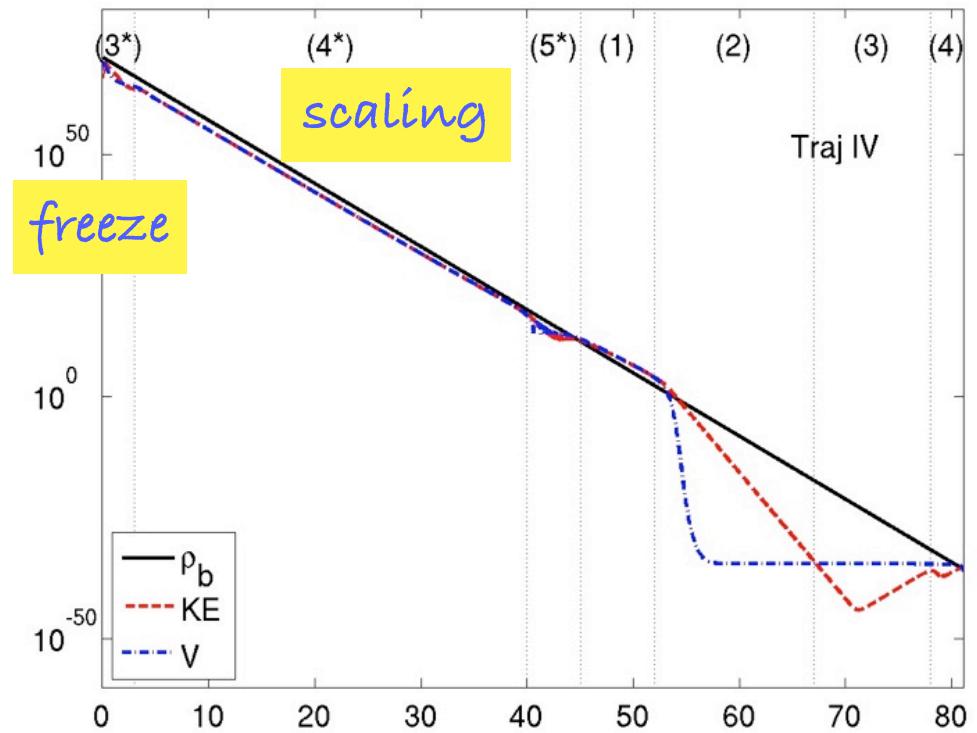
scaling condition:

$$3\gamma < \lambda^2 + \delta^2/2$$

$$\lambda \equiv -\frac{1}{V} \frac{\partial V}{\partial \phi}$$

$$\delta \equiv -\frac{1}{V} \frac{\partial V}{\partial C_r}$$

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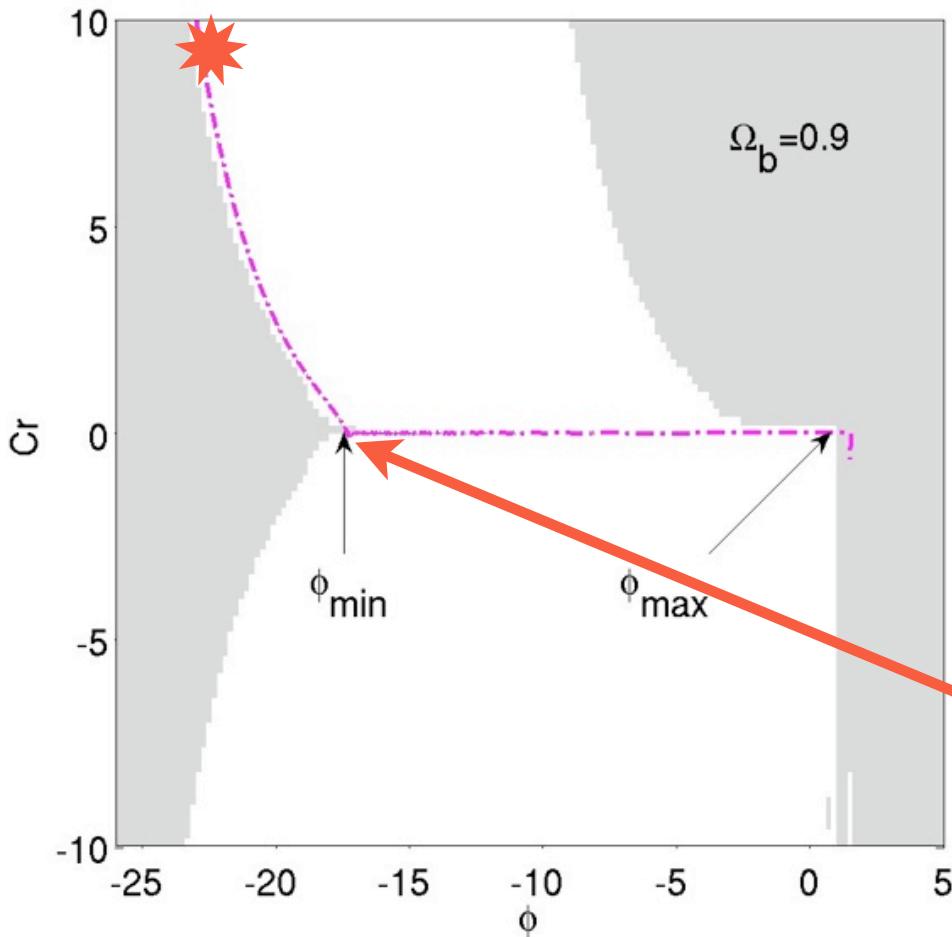


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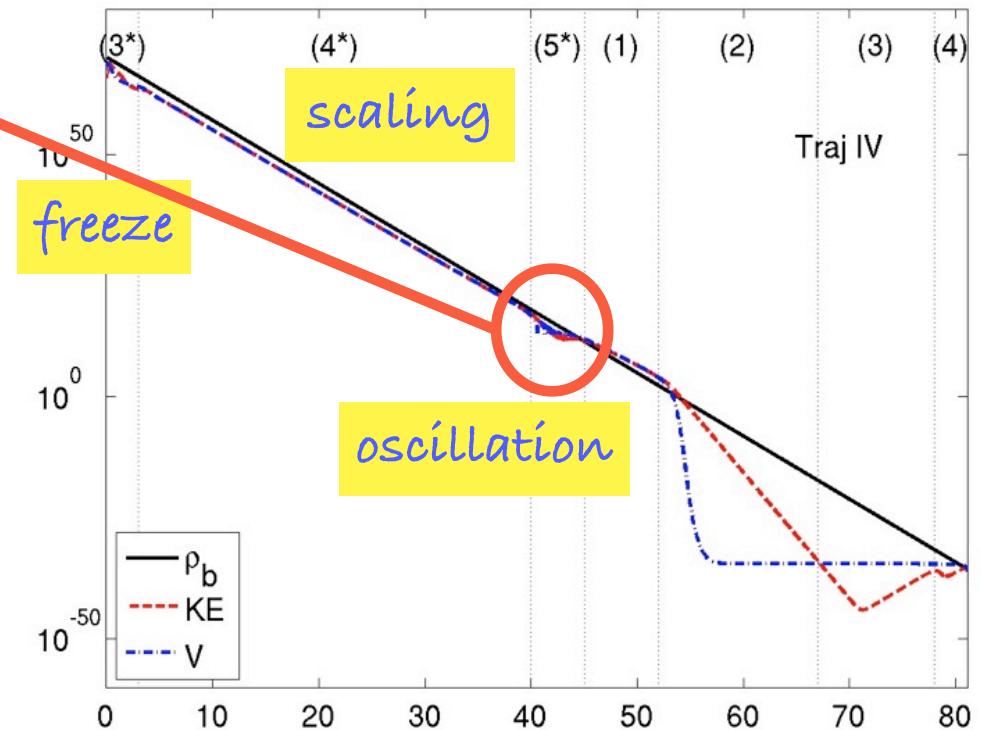


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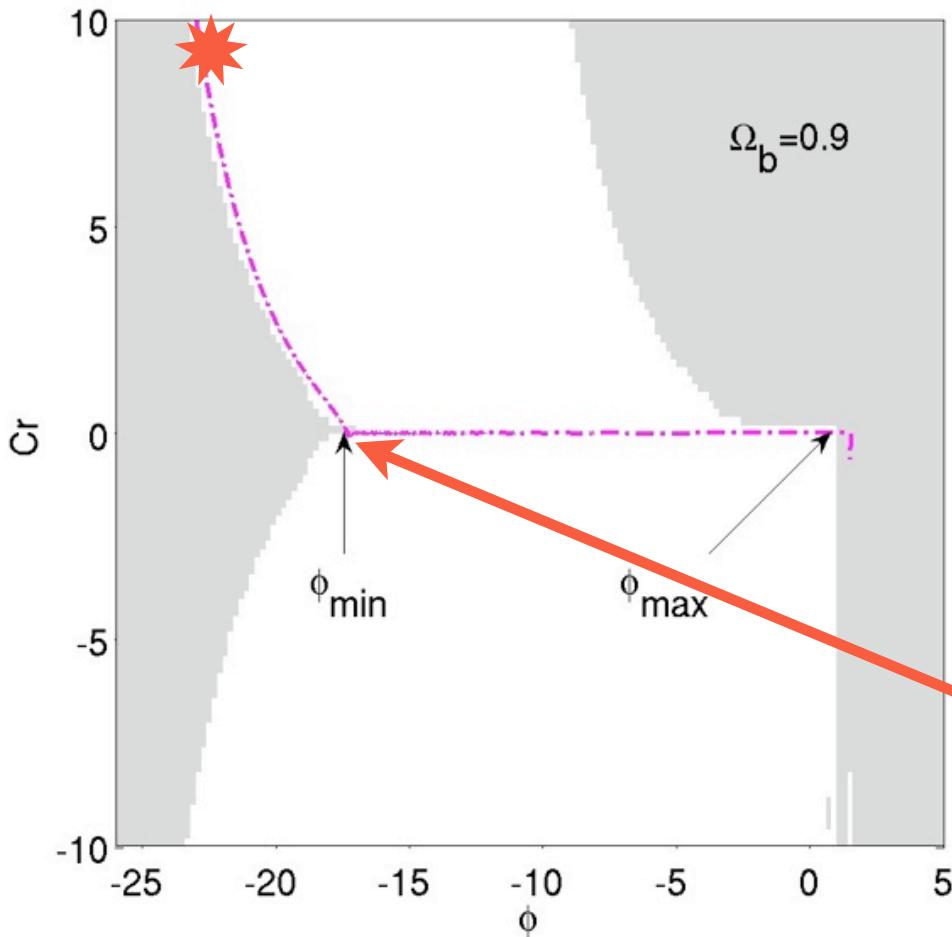
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Two REAL Fields

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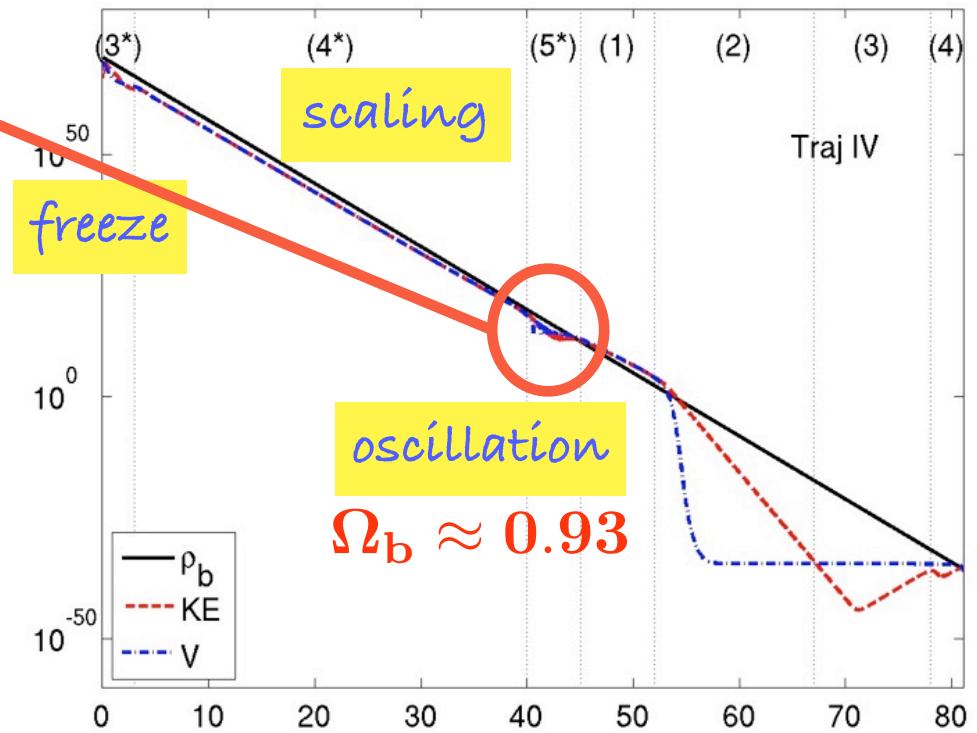
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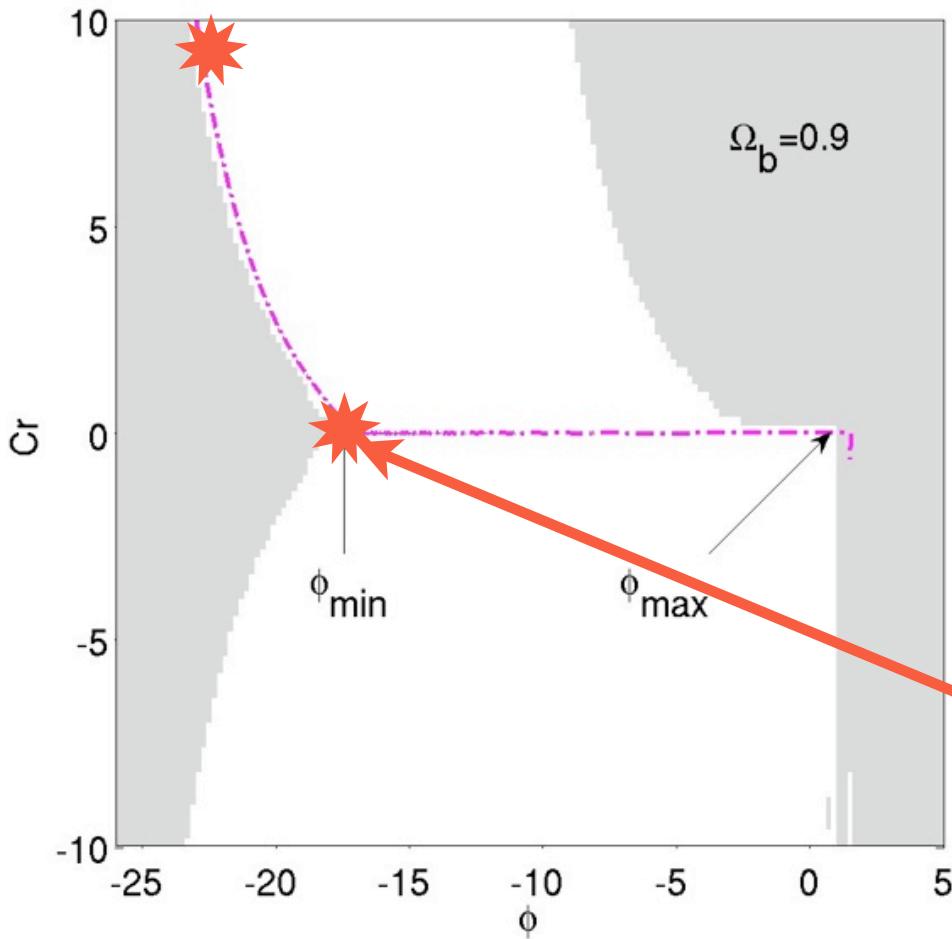


$$C_i = T_i = 0$$

Two REAL Fields

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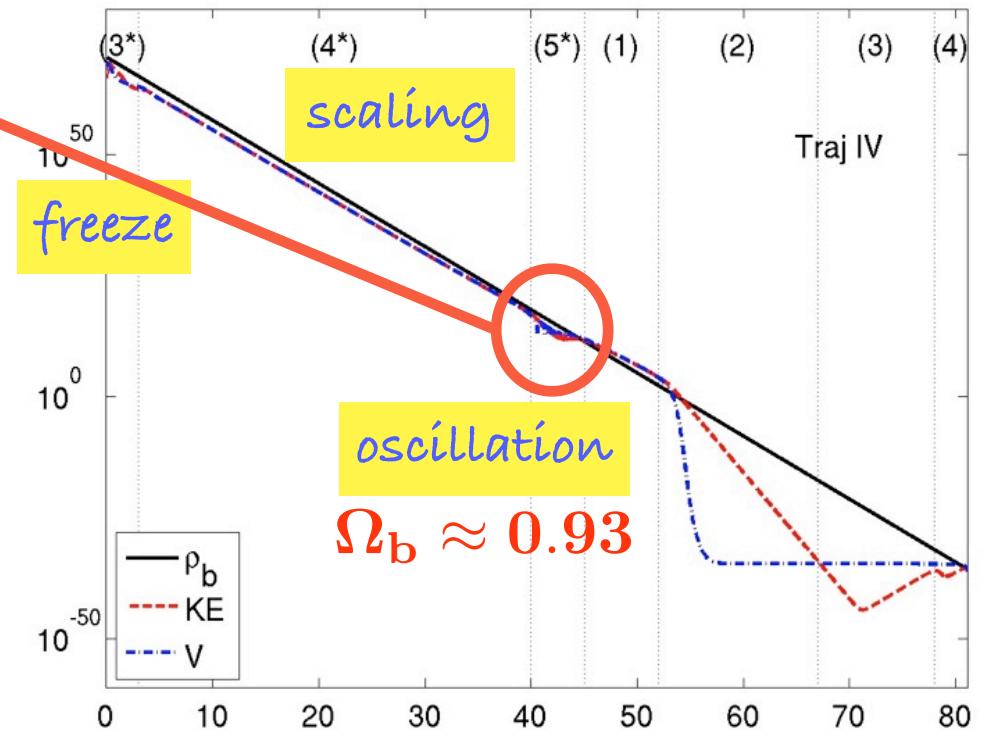


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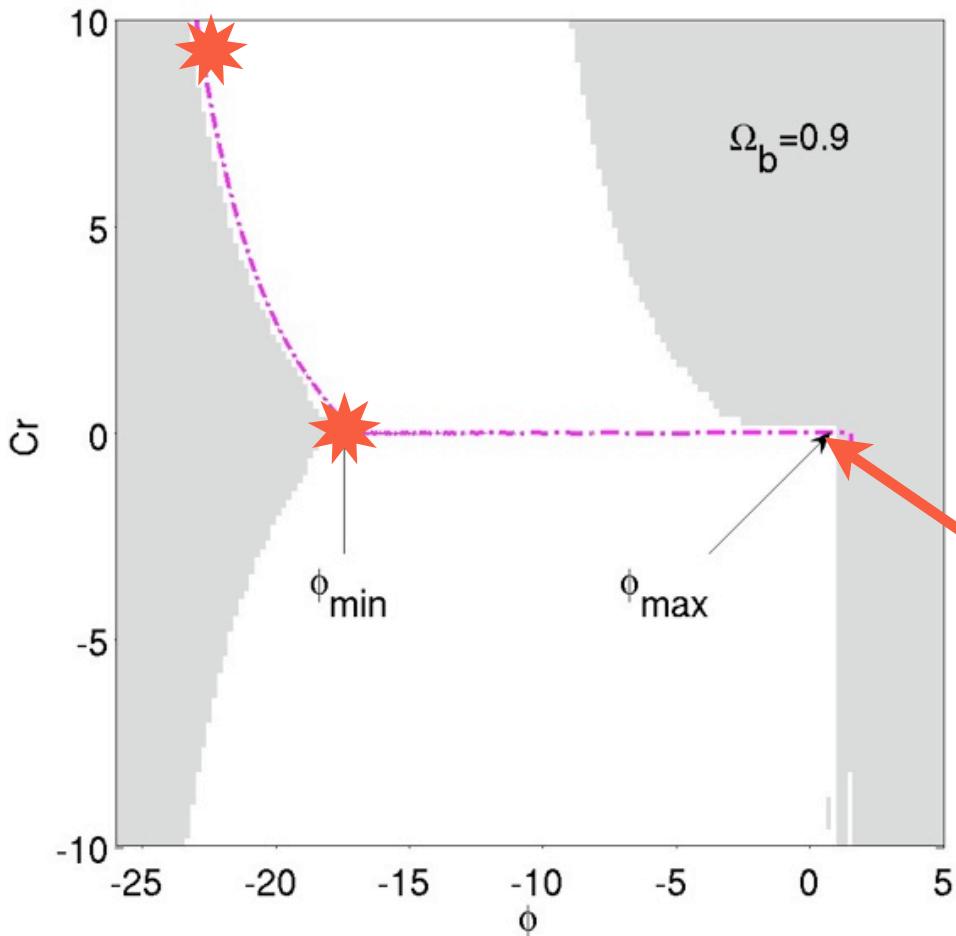


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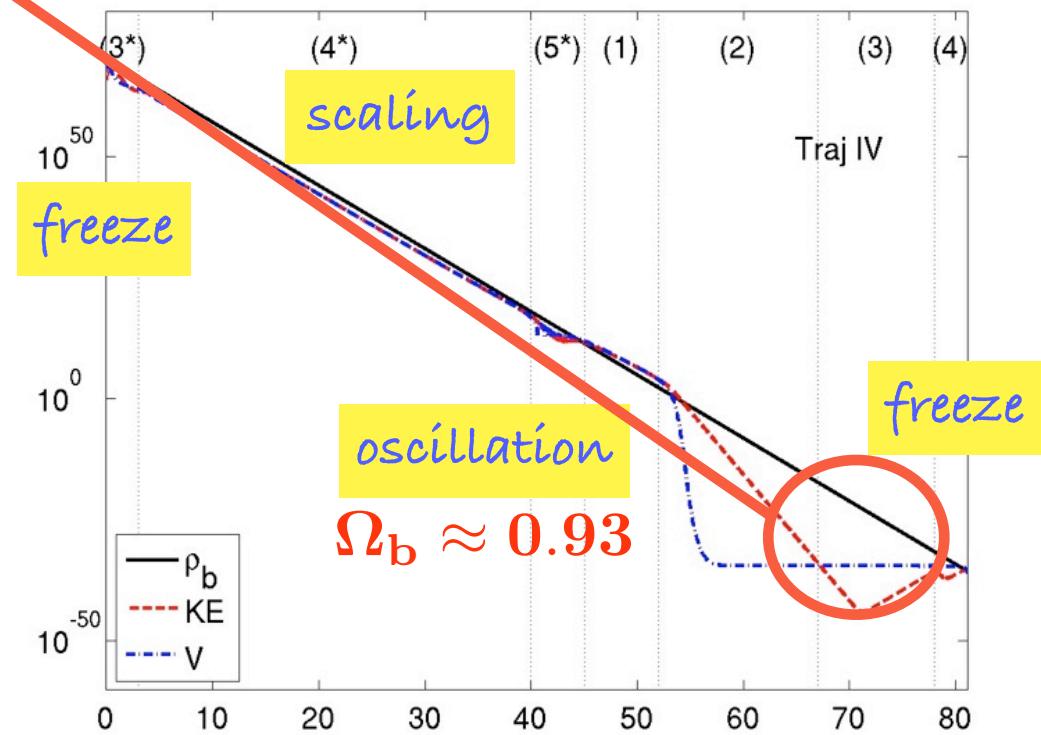


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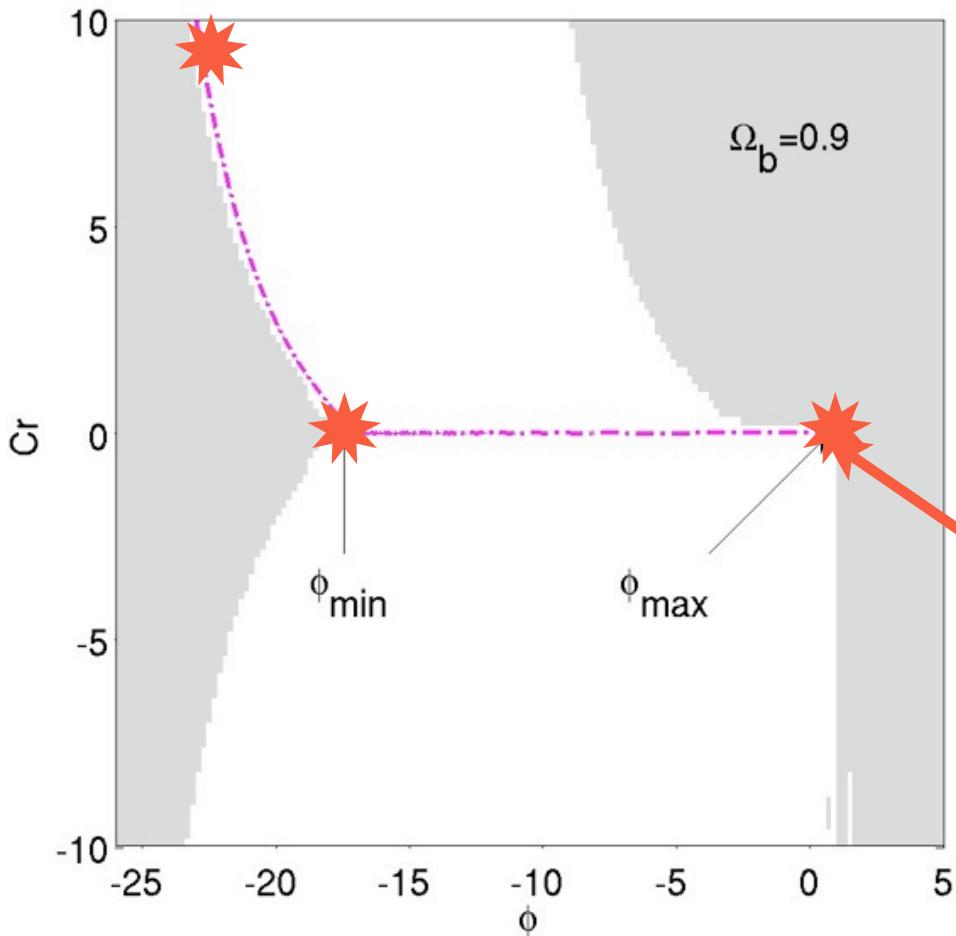


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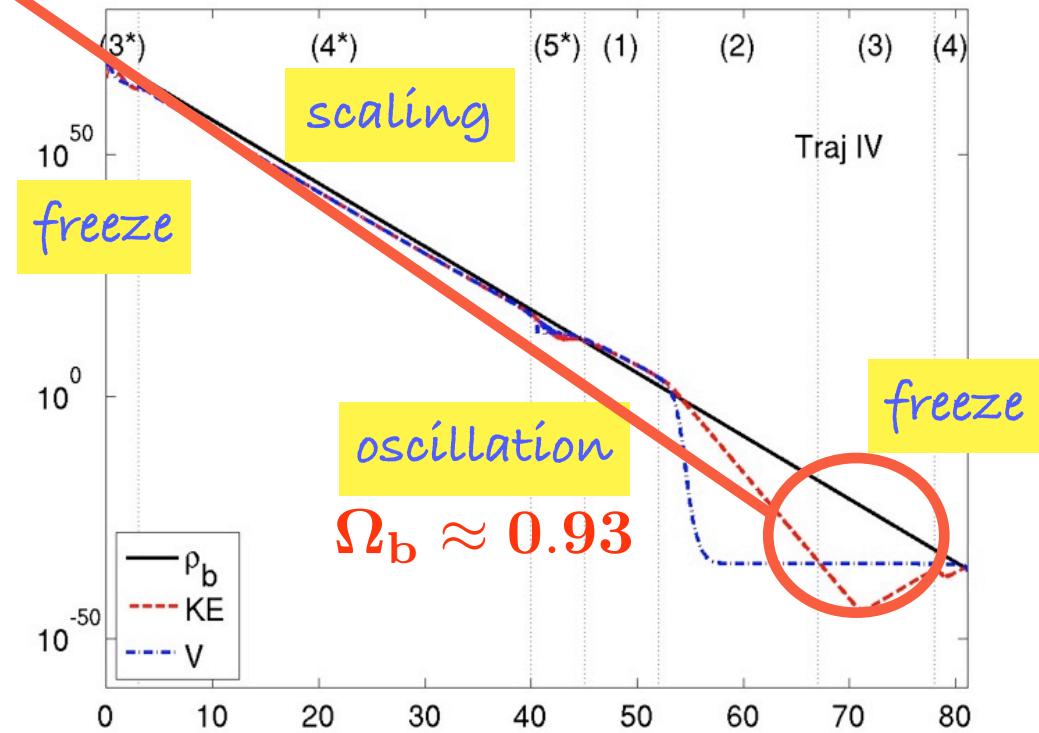


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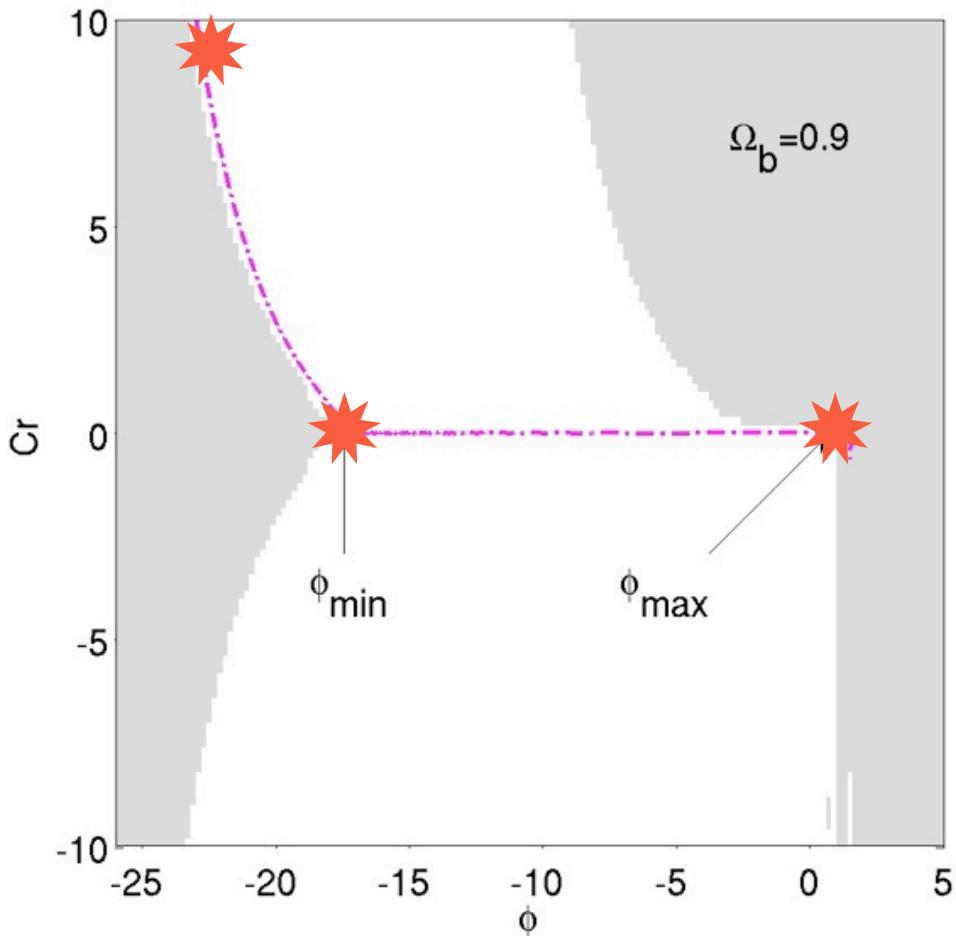
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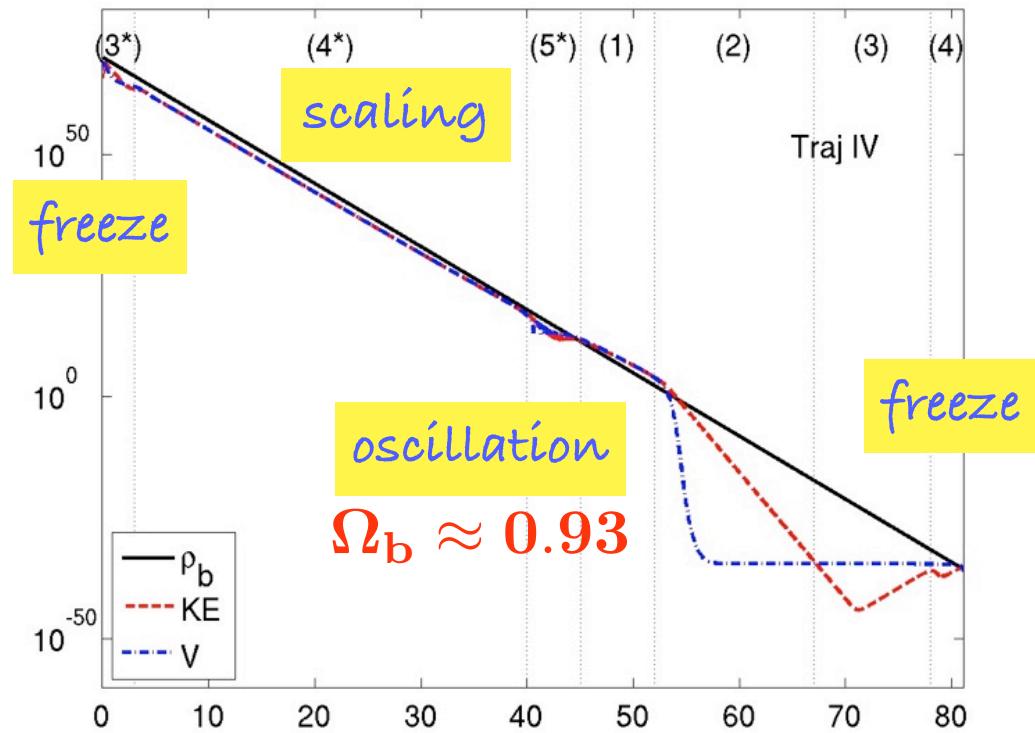
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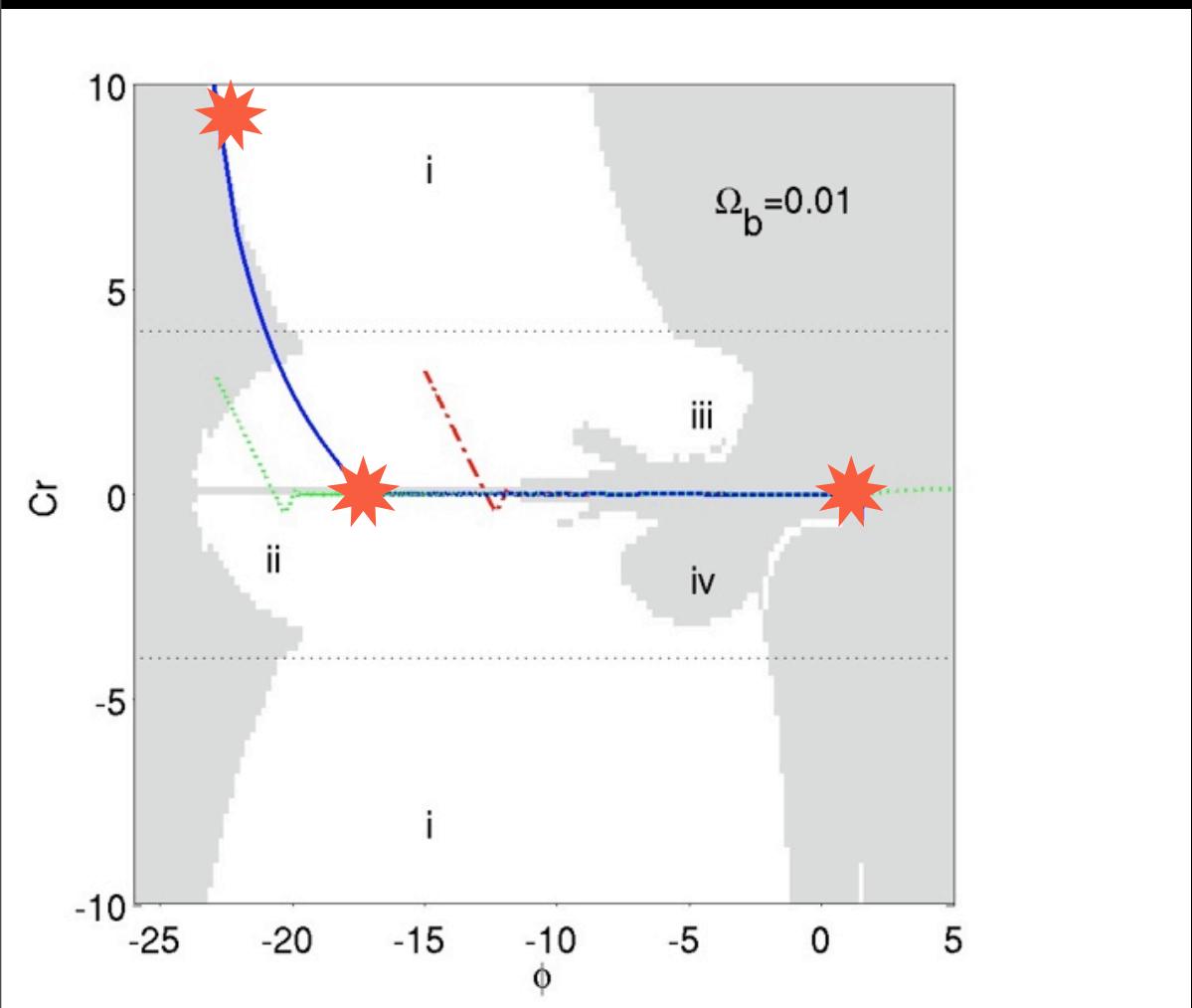


$$C_i = T_i = 0$$

Two REAL Fields

$$\Omega_b = 0.01$$

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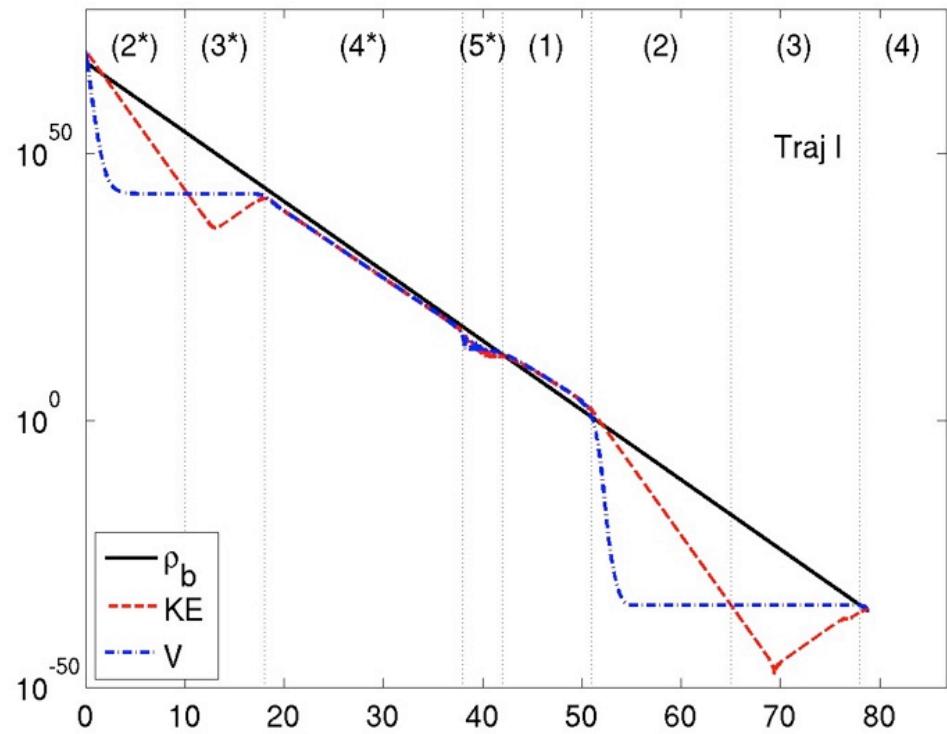
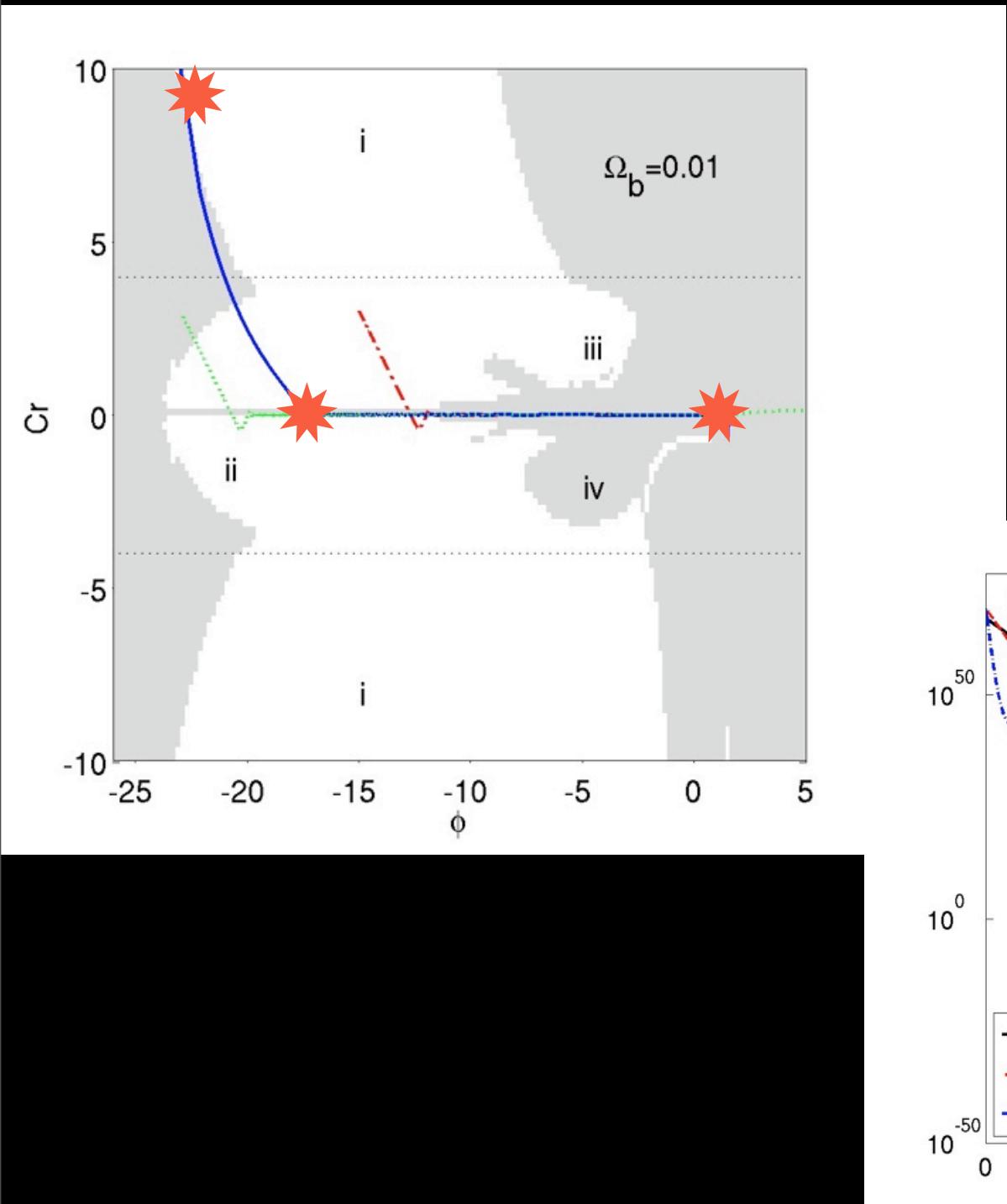


$$C_i = T_i = 0$$

Two REAL Fields

$$\Omega_b = 0.01$$

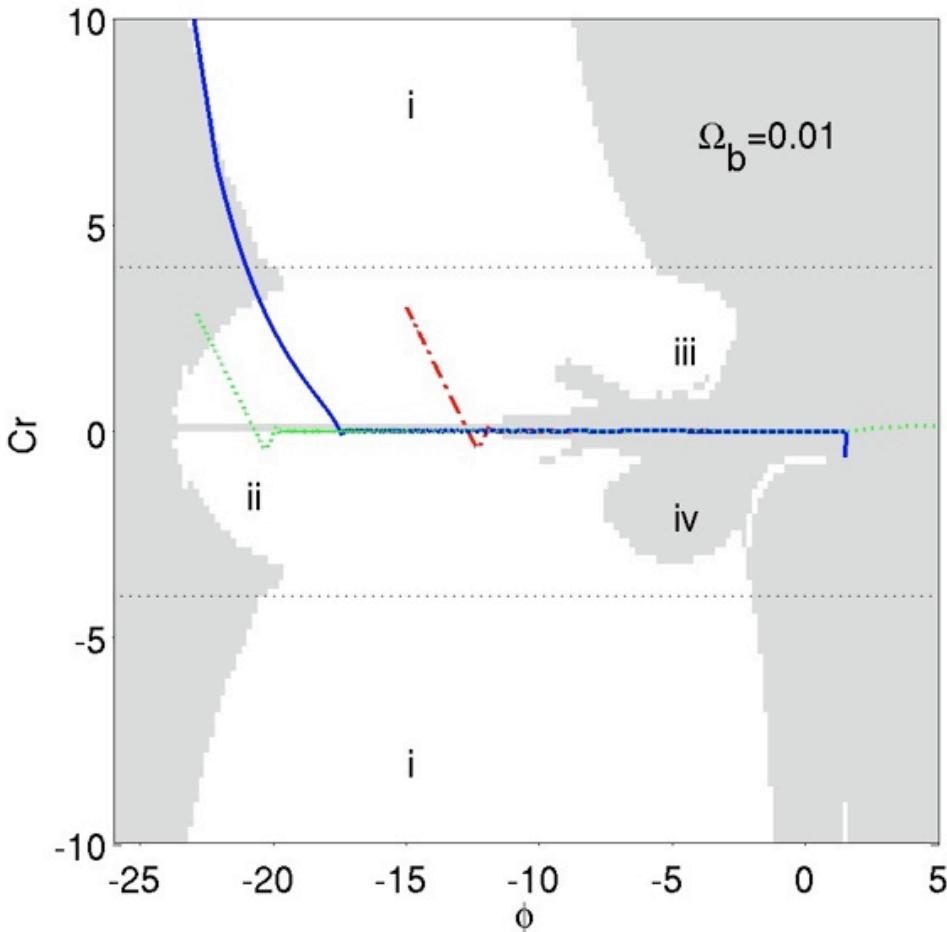
- 1. Potential dom.
- 2. Kination
- 3. Freeze out
- 4. Scaling
- 5. Oscillation



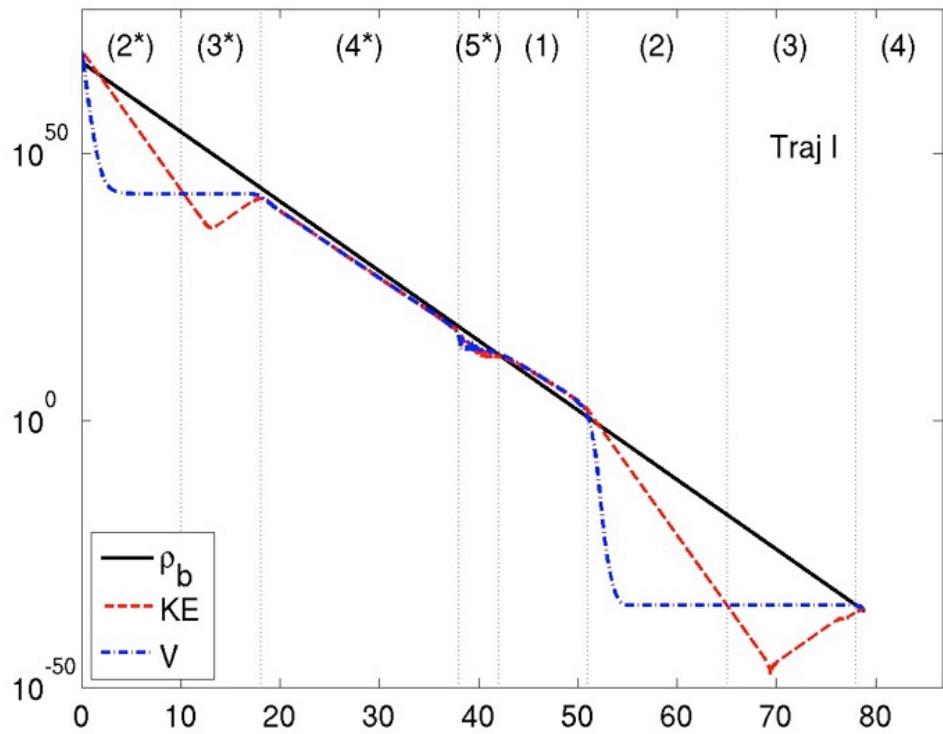
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Two REAL Fields

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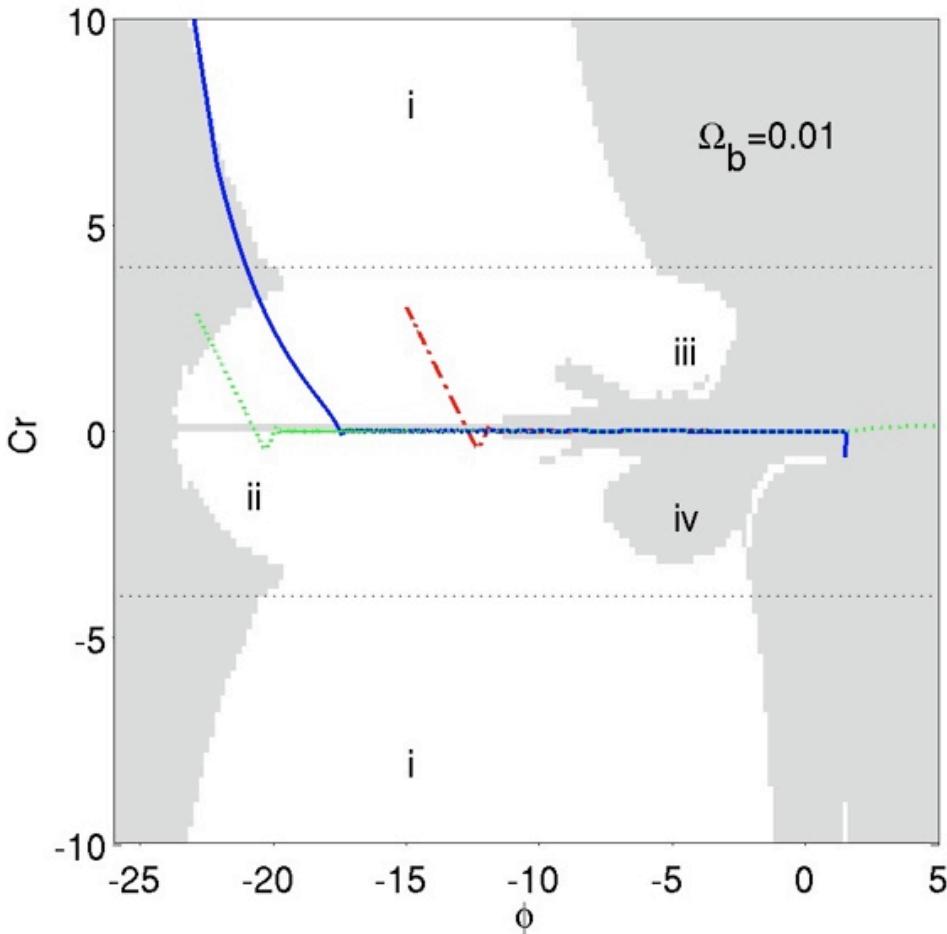
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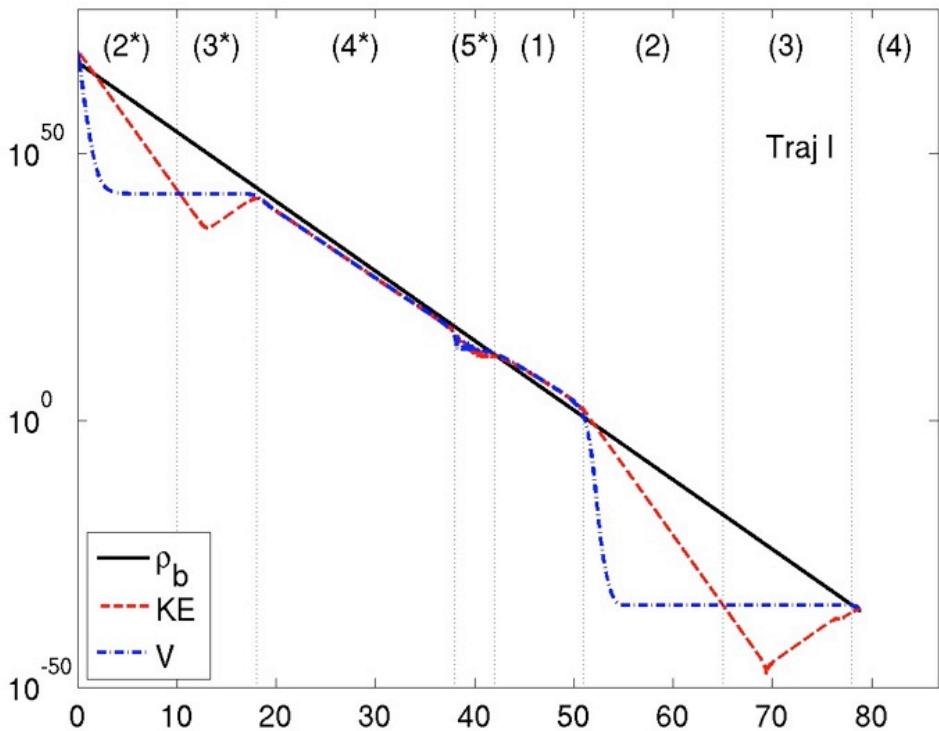
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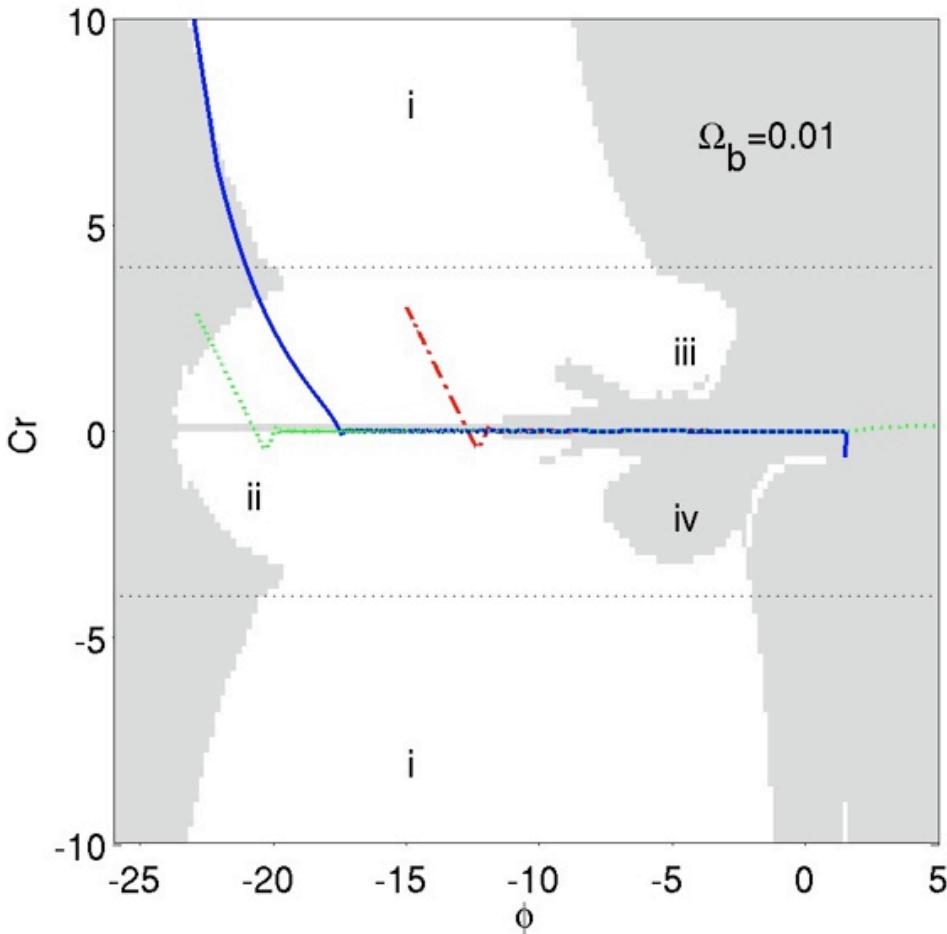
At end of scaling (oscillation):

$\Omega_b \approx 0.93$

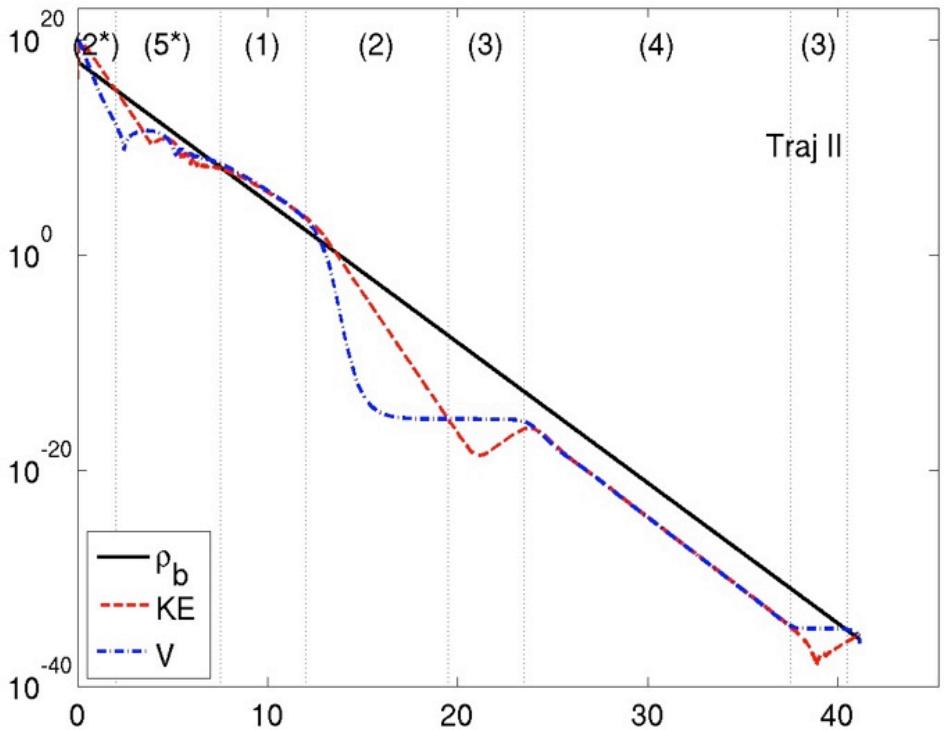
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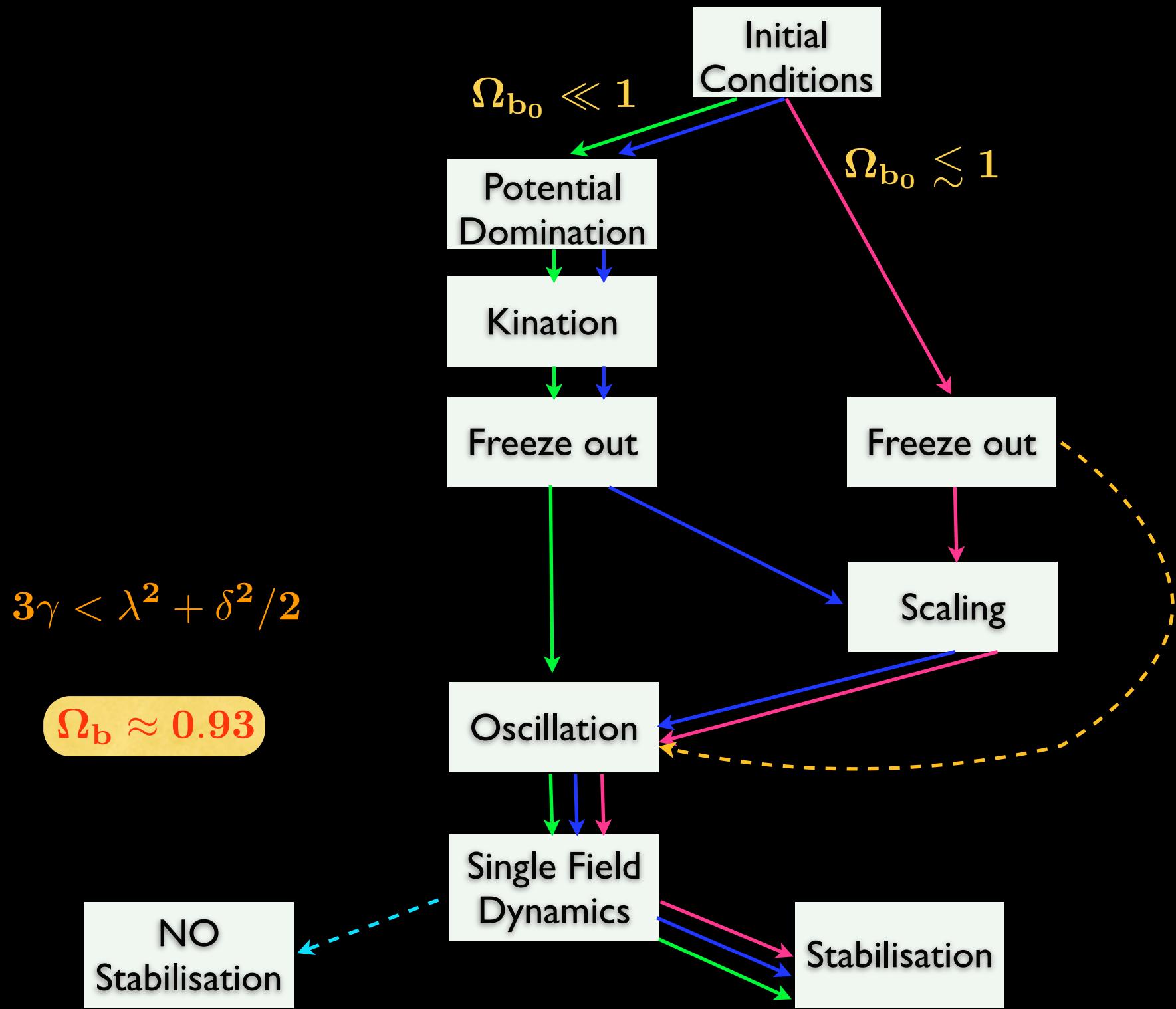


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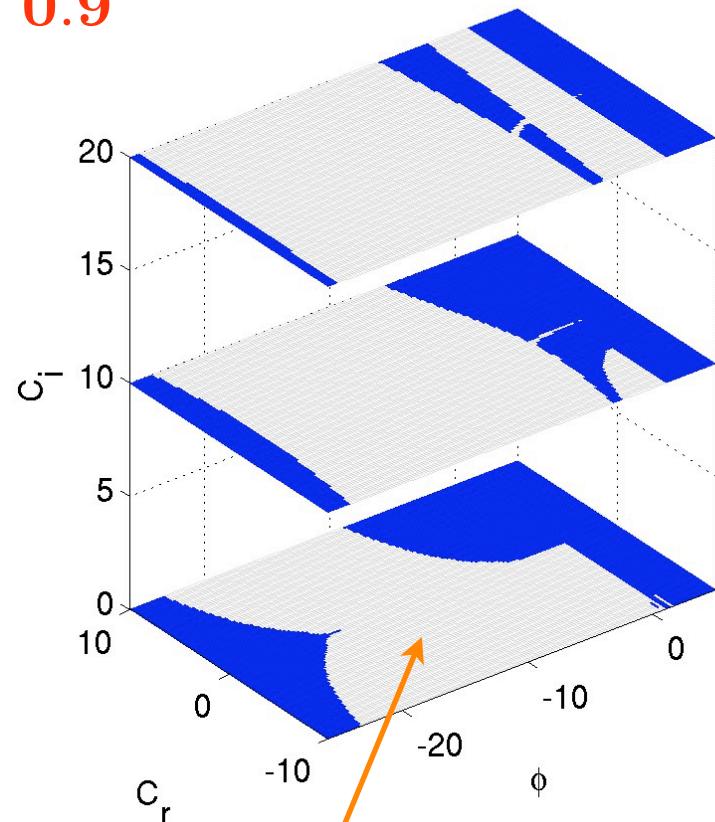


At end of scaling (oscillation):

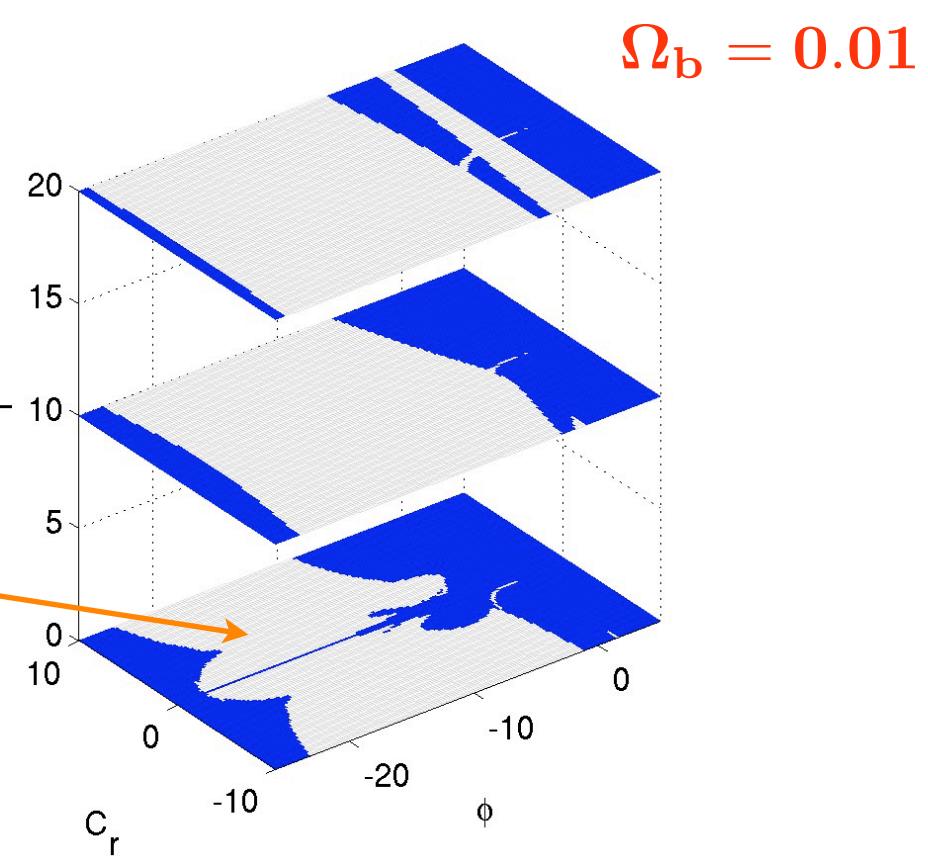
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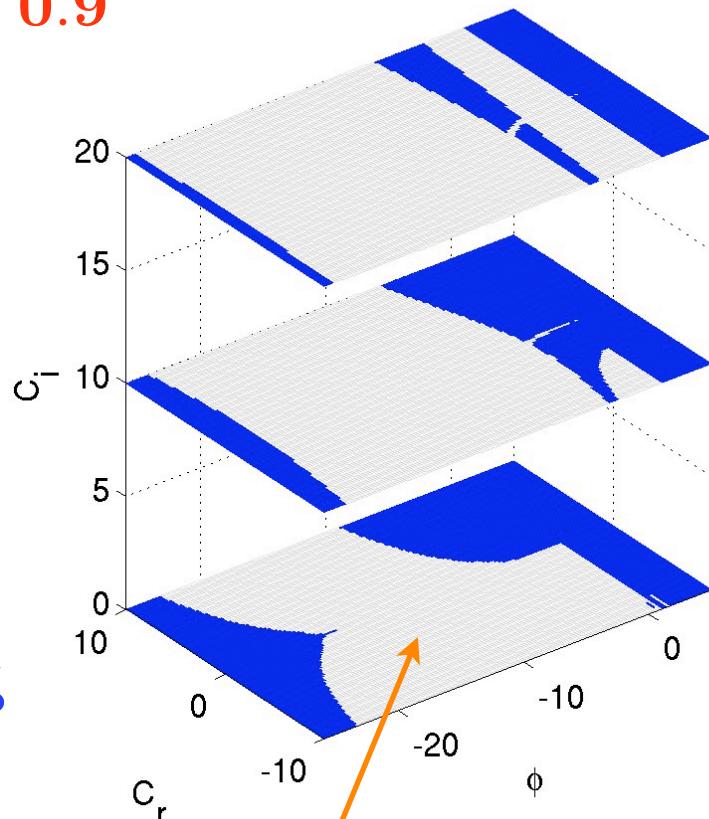
Real fields



$\Omega_b = 0.01$

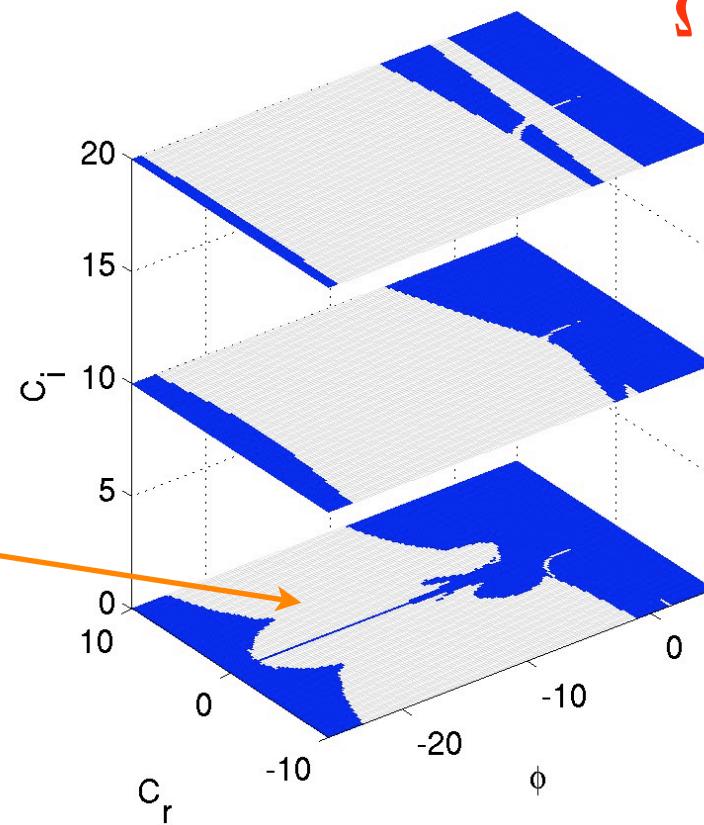
$\Omega_b = 0.9$

59.0%



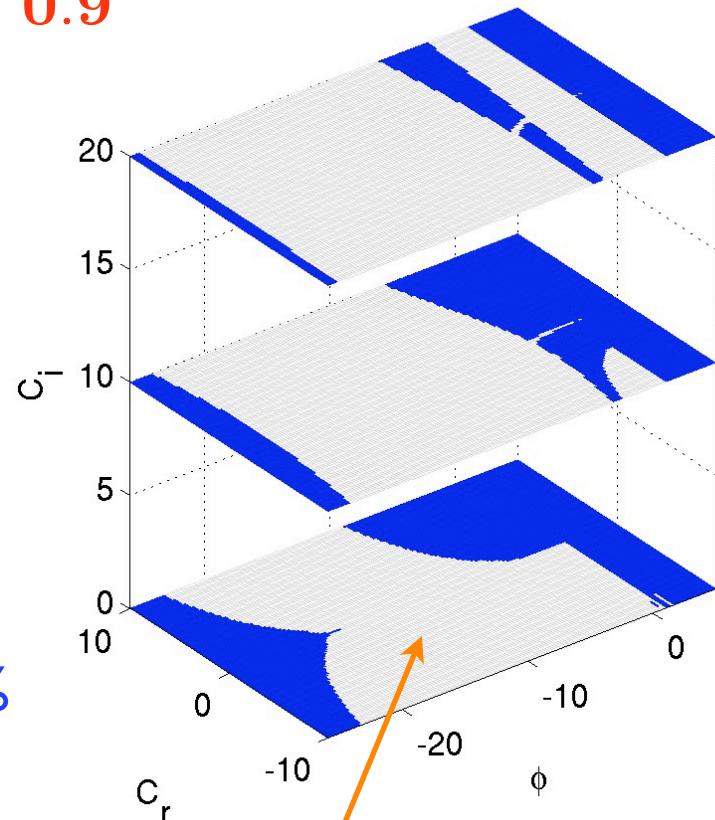
Real fields

$\Omega_b = 0.01$



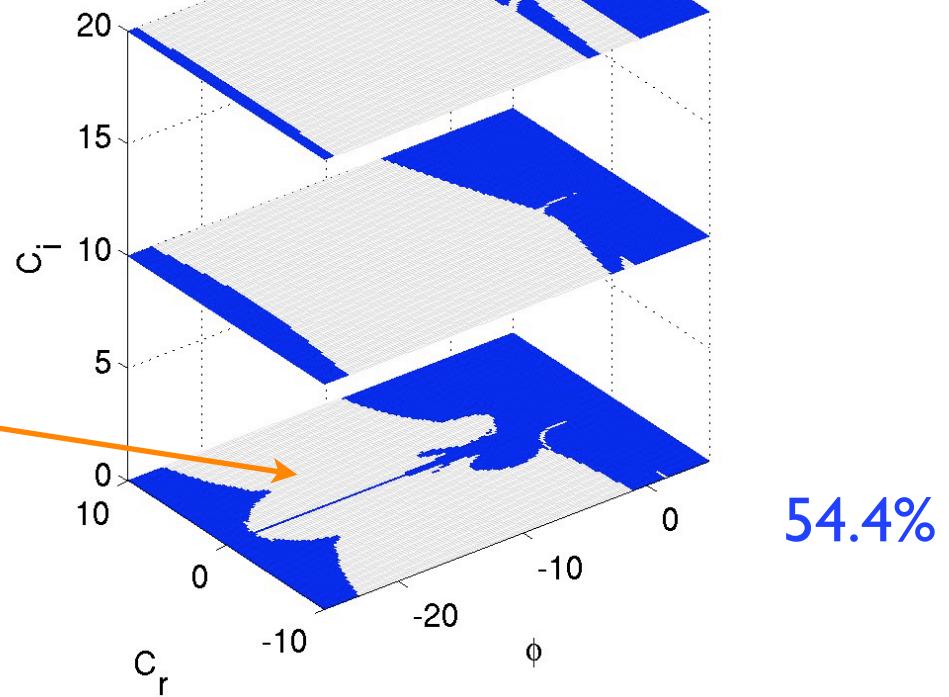
$\Omega_b = 0.9$

59.0%



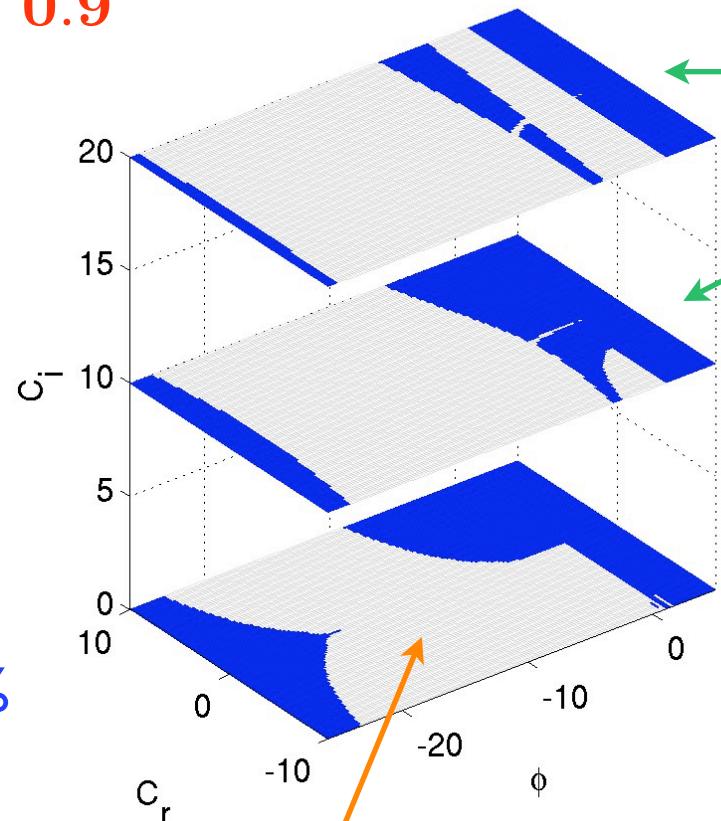
Real fields

$\Omega_b = 0.01$



$\Omega_b = 0.9$

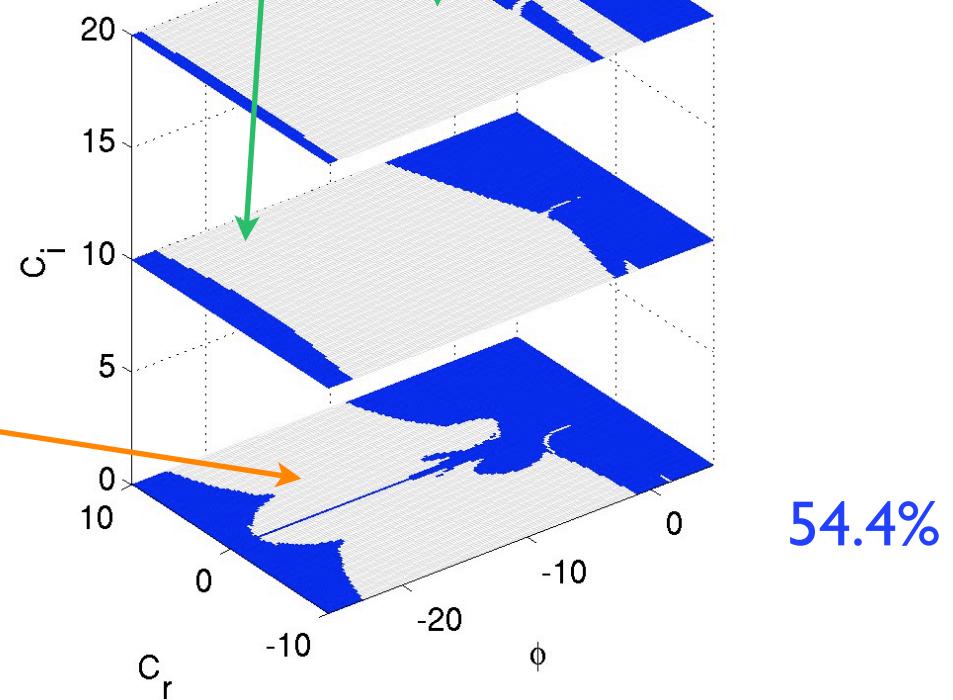
59.0%



Complex fields

Real fields

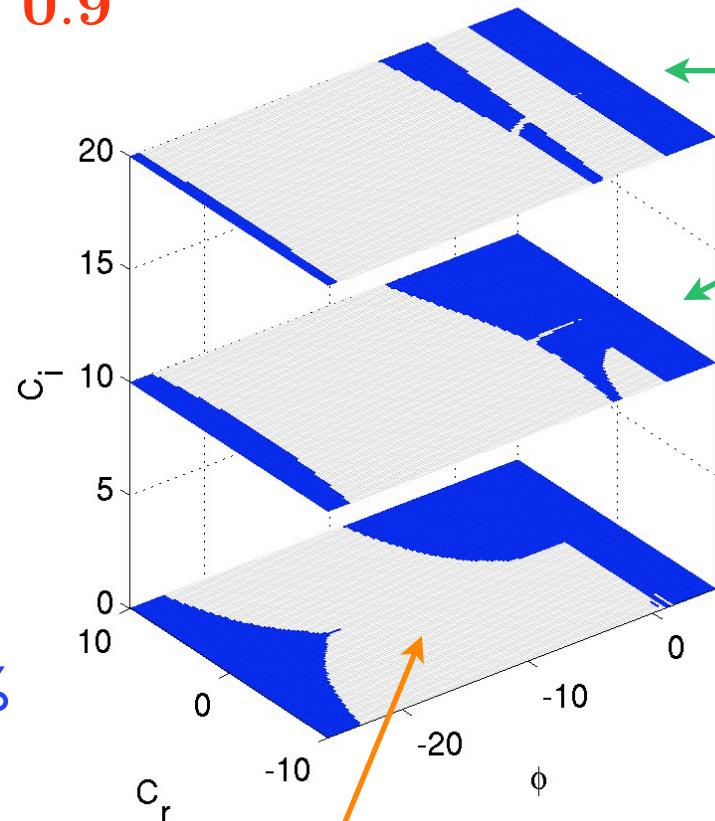
$\Omega_b = 0.01$



54.4%

$\Omega_b = 0.9$

59.0%

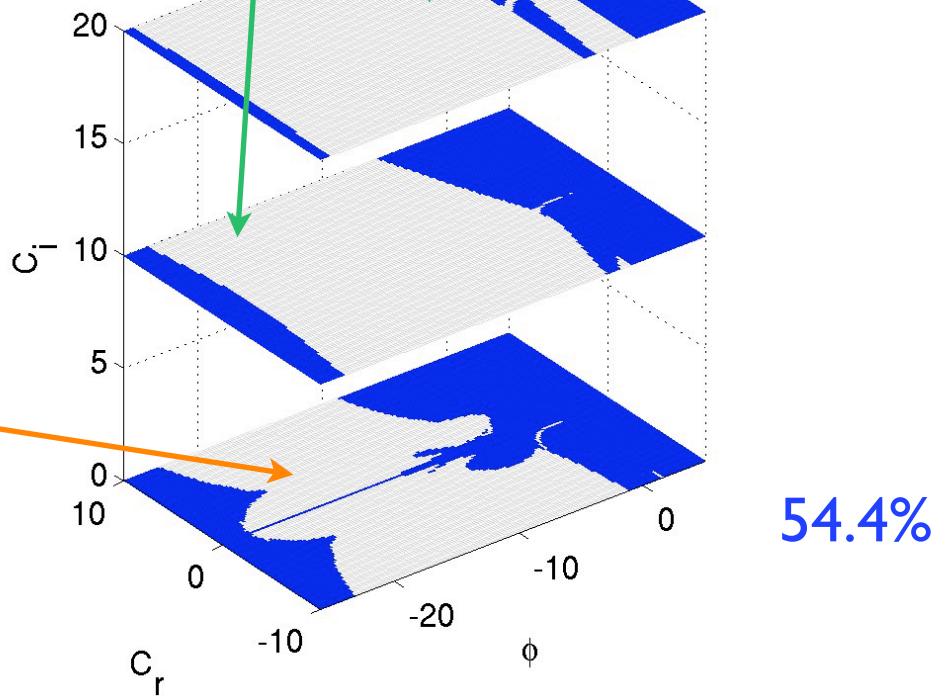


Real fields

Complex fields

potential changes
with dynamics

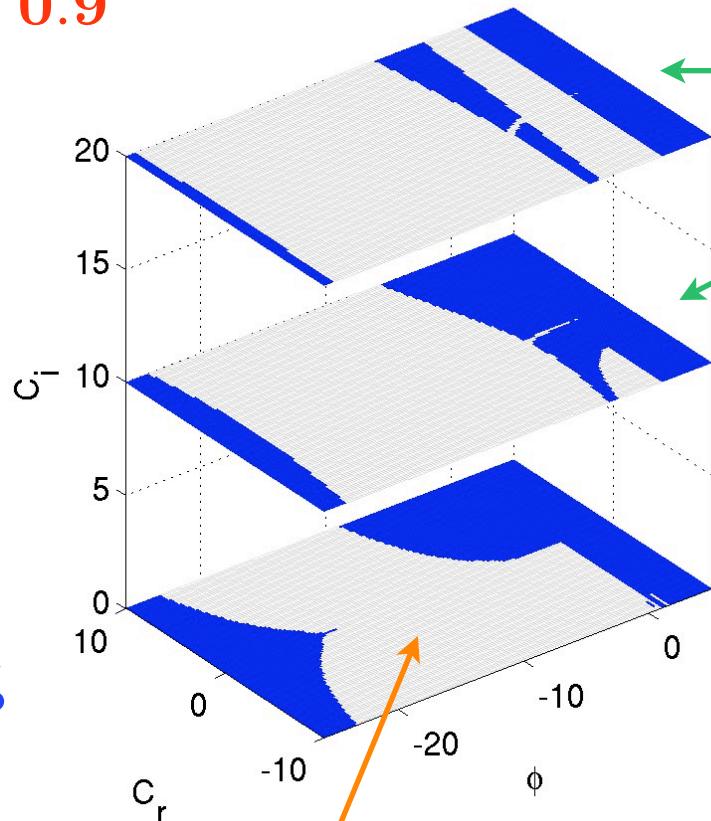
$\Omega_b = 0.01$



54.4%

$\Omega_b = 0.9$

59.0%



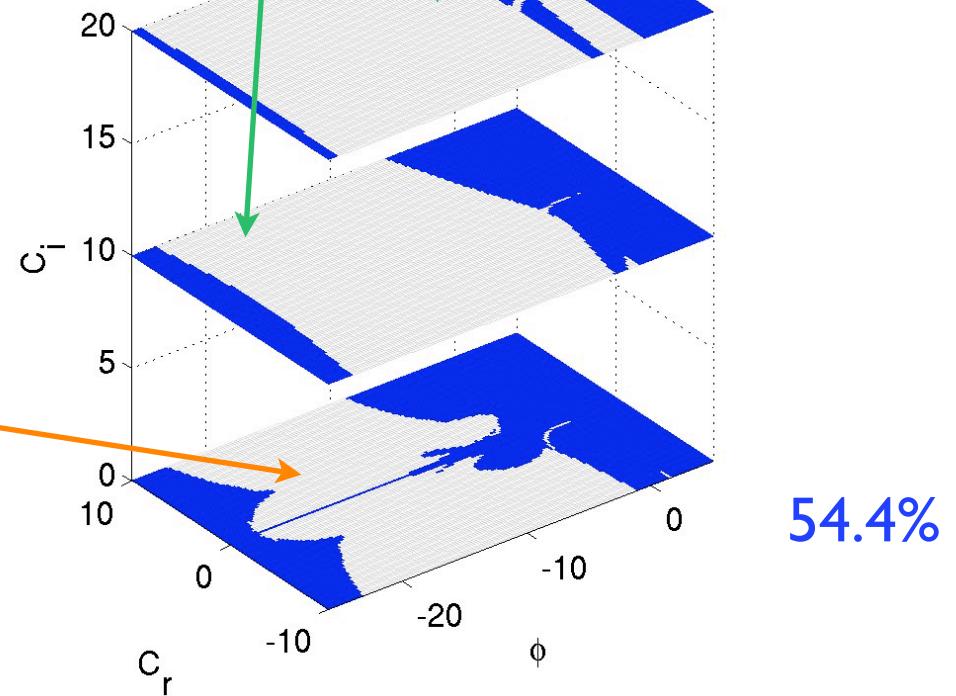
Real fields

Complex fields

potential changes
with dynamics

aids stabilisation
further !!!

$\Omega_b = 0.01$



54.4%

$\Omega_b = 0.9$

76.8%

66.8%

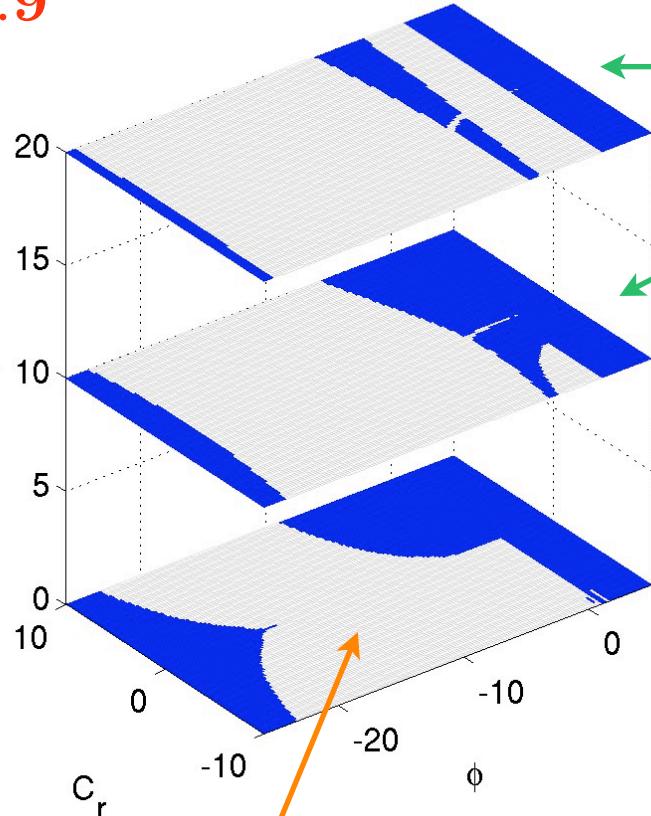
59.0%

C_r

-10 0 -20

ϕ

Real fields



Complex fields

potential changes
with dynamics

aids stabilisation
further !!!

$\Omega_b = 0.01$

71.2%

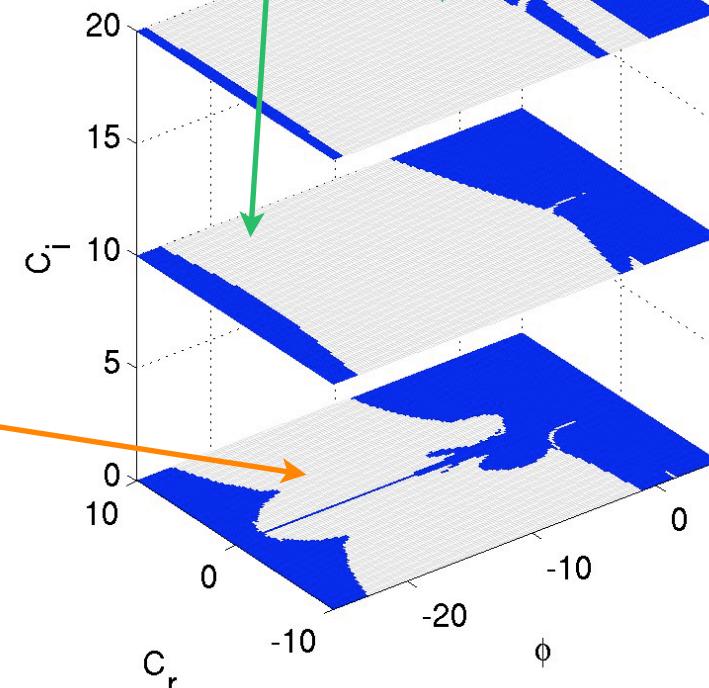
66.1%

54.4%

C_r

-10 0 -20

ϕ



Summary

KKLT + Polonyi Model:

$$K = -3 \ln(T + \bar{T}) + |C|^2$$

$$\mathcal{W}(C) = c + \mu^2 C$$

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...with background fluid

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Background: Matter/Radiation ($\gamma = 1, 4/3$)

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Carsten van de Bruck,
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