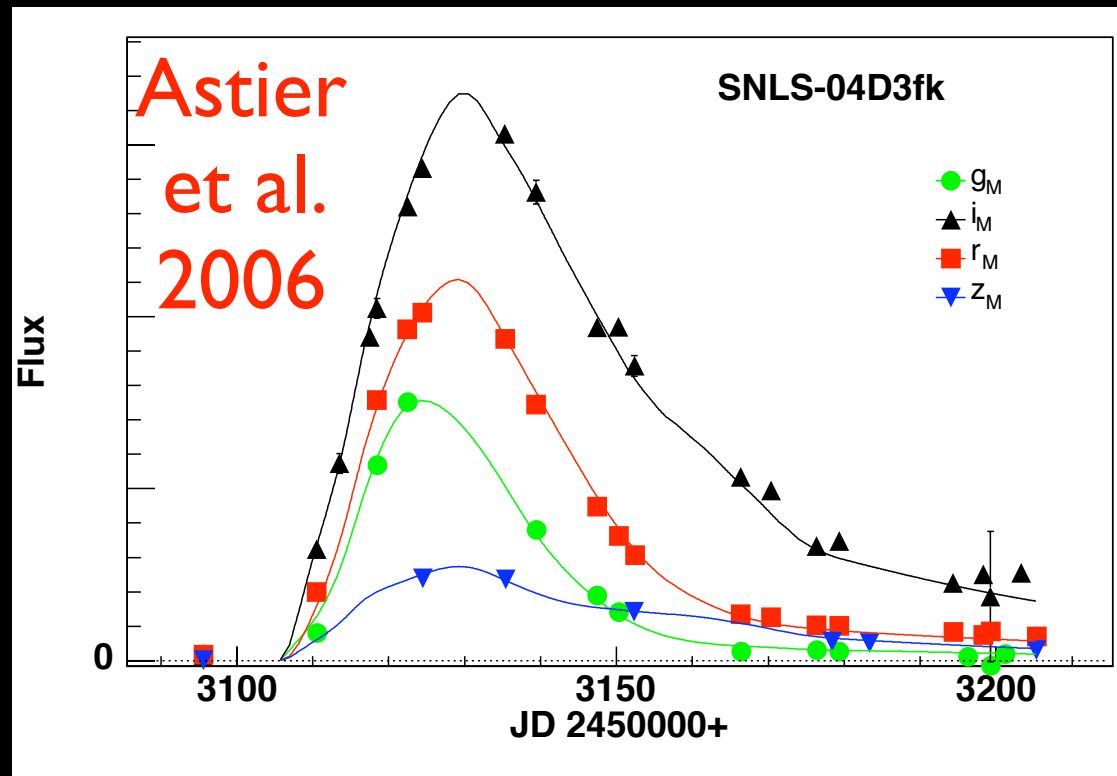
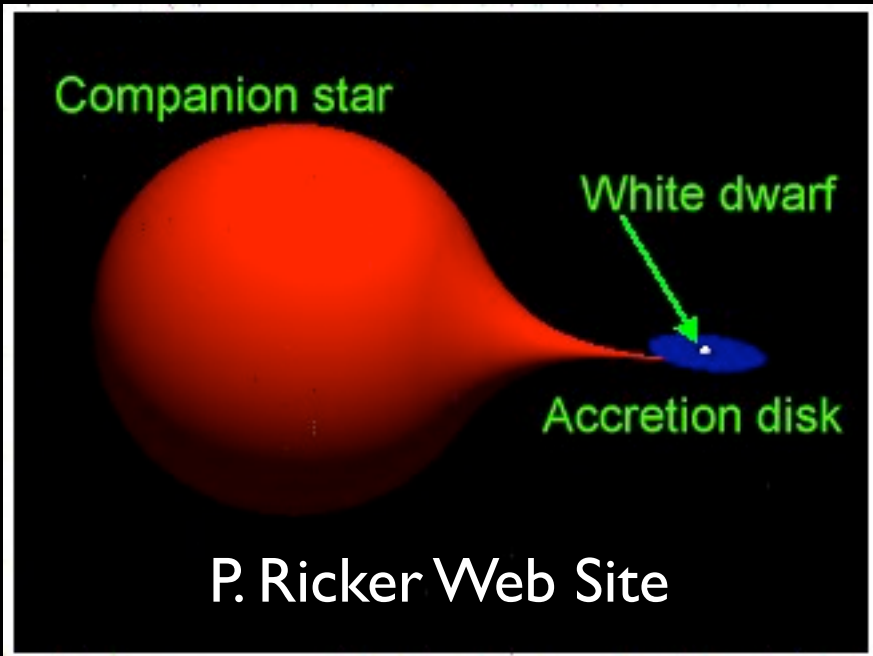


The Effect of Correlated Peculiar Velocities on Cosmology from Type Ia Supernovae

Christopher Gordon

Based on work done with Kate Land and Anze Slosar



- White dwarf explodes when Chandrasekhar limit of 1.44 solar masses is reached.
- Scatter reduced by taking into account correlations between rate of decline, colour, and intrinsic luminosity, e.g. MLCS2K2, SALT.

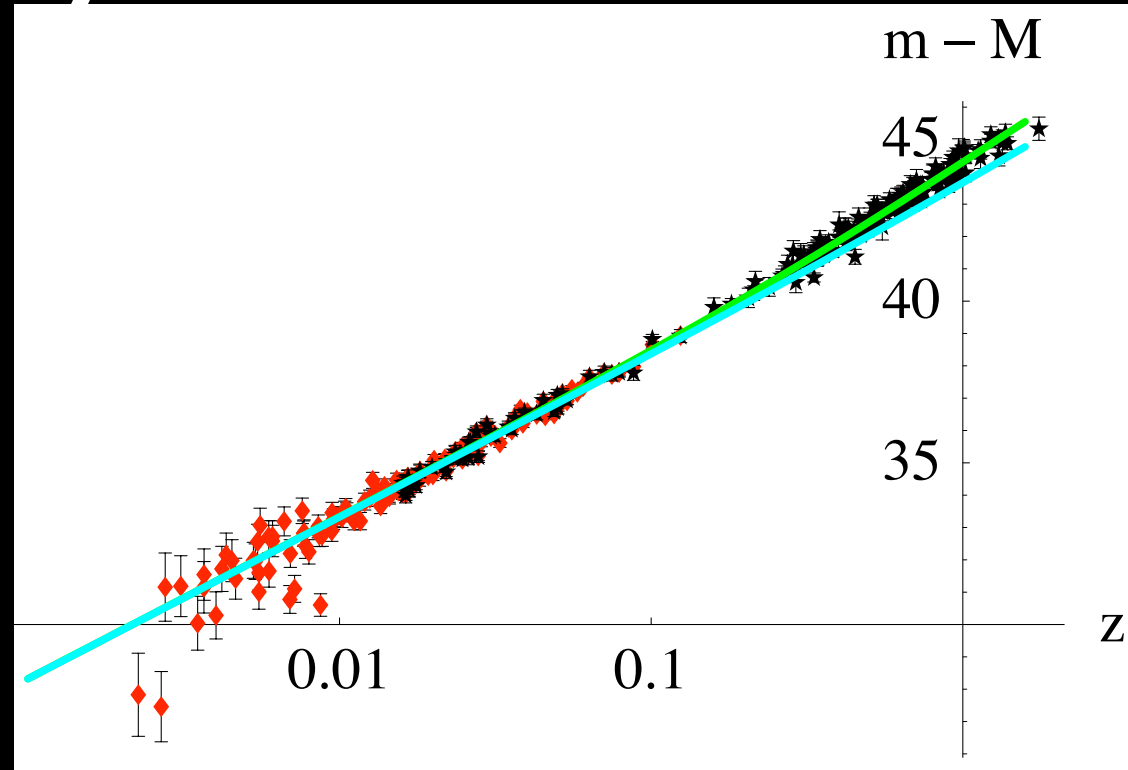
Luminosity Distance

$$\mathcal{F} = \frac{\mathcal{L}}{4\pi d_L^2}$$

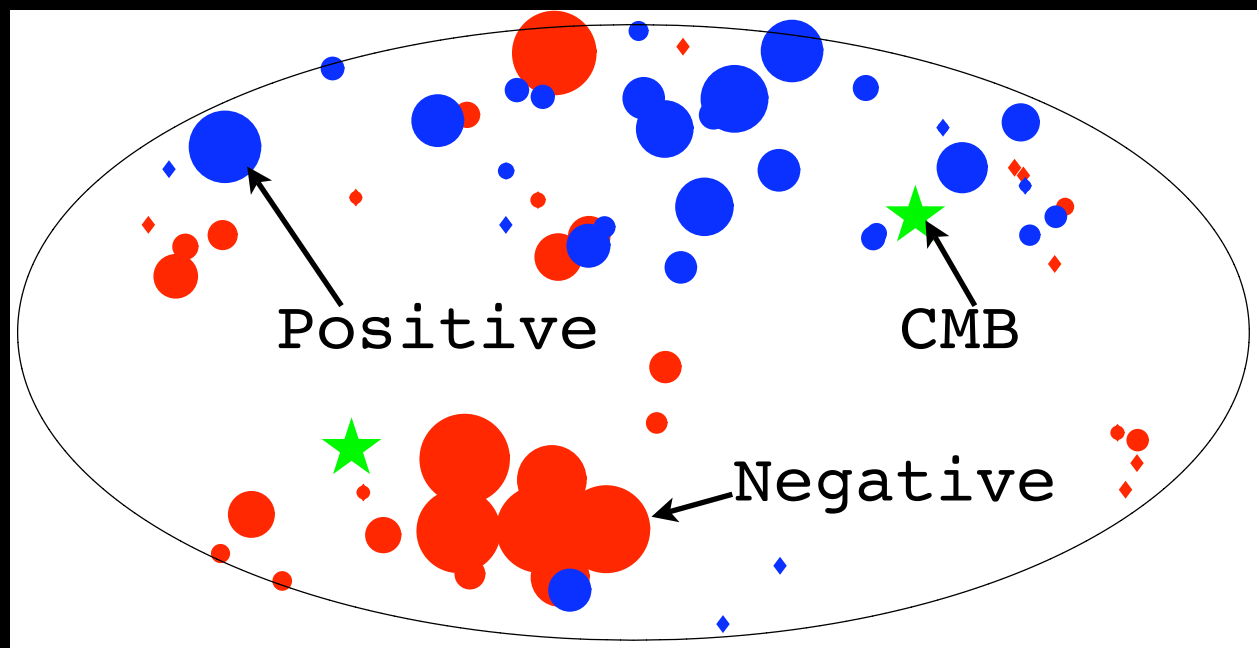
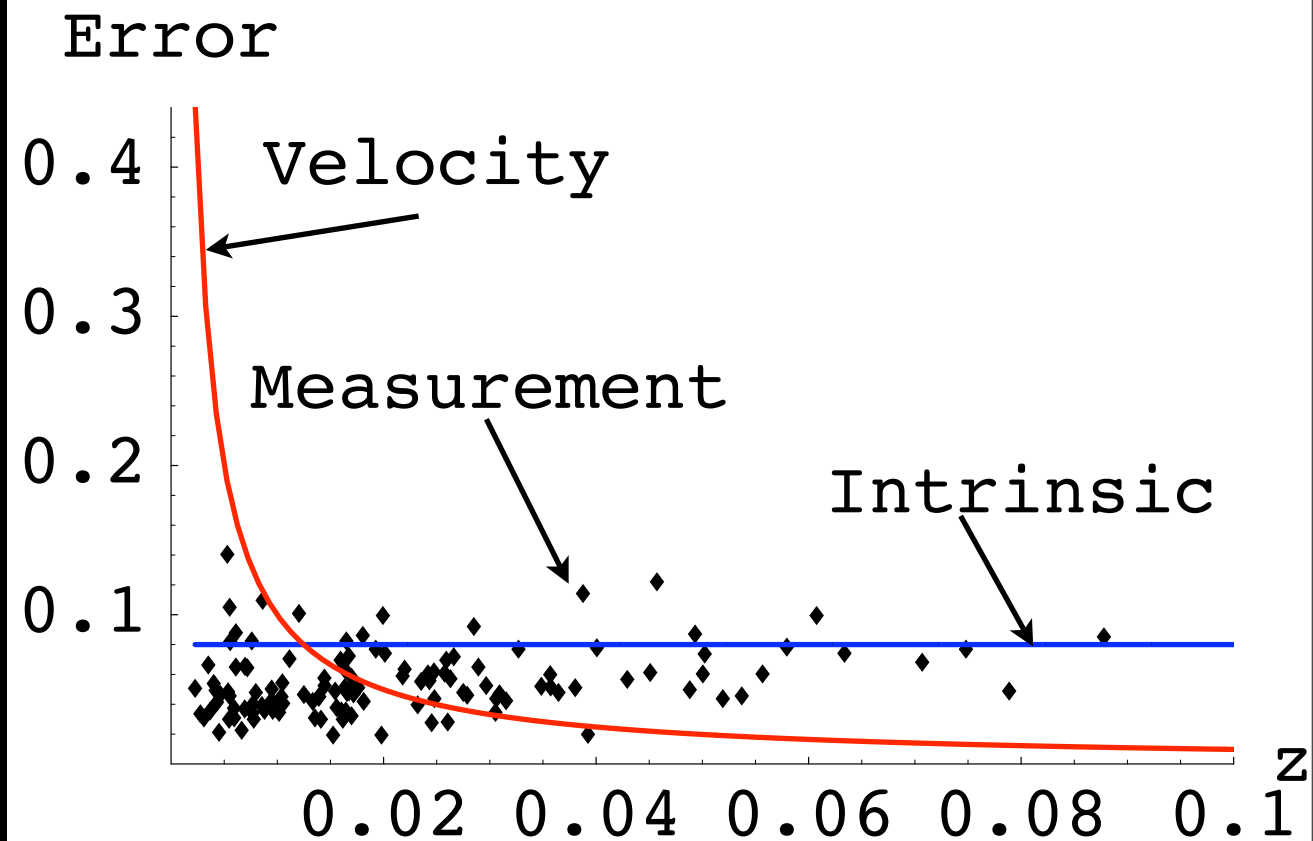
$$m = 5 \log d_L + 25 + M$$

$$d_L(z) = (1+z) \int_0^z \frac{dz'}{H(z')}$$

$$H^2 = \frac{8\pi G}{3} \rho_0 \left(\Omega_m (1+z)^3 + \Omega_d (1+z)^{3(1+w)} \right)$$



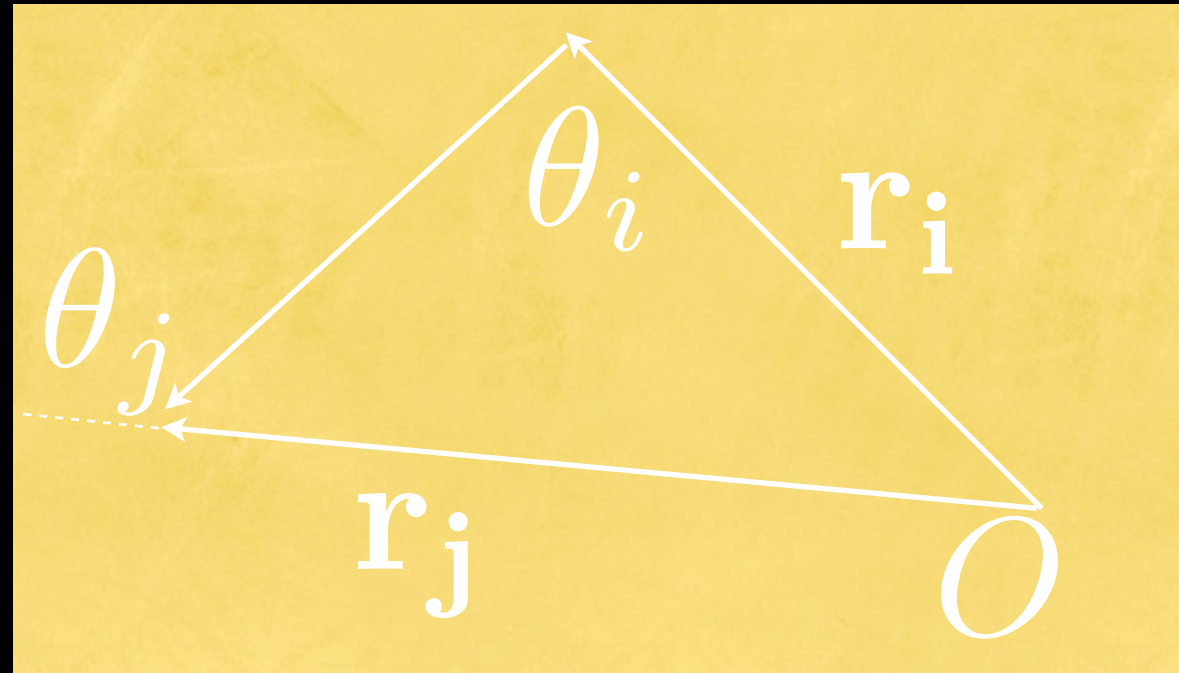
$$\frac{\delta d_L}{d_L} \approx \frac{\hat{r} \cdot (v_{\text{obs}} - v)}{z} + \delta m + \delta n$$



Peculiar Velocities

$$\delta_{\mathbf{k}}(z) = D(z)\tilde{\delta}_{\mathbf{k}}$$

$$\mathbf{v}_{\mathbf{k}} = -iD'\frac{\tilde{\delta}_{\mathbf{k}}}{k^2}\mathbf{k}$$



$$\xi = \sin \theta_i \sin \theta_j \xi_{\perp} + \cos \theta_i \cos \theta_j \xi_{\parallel}$$

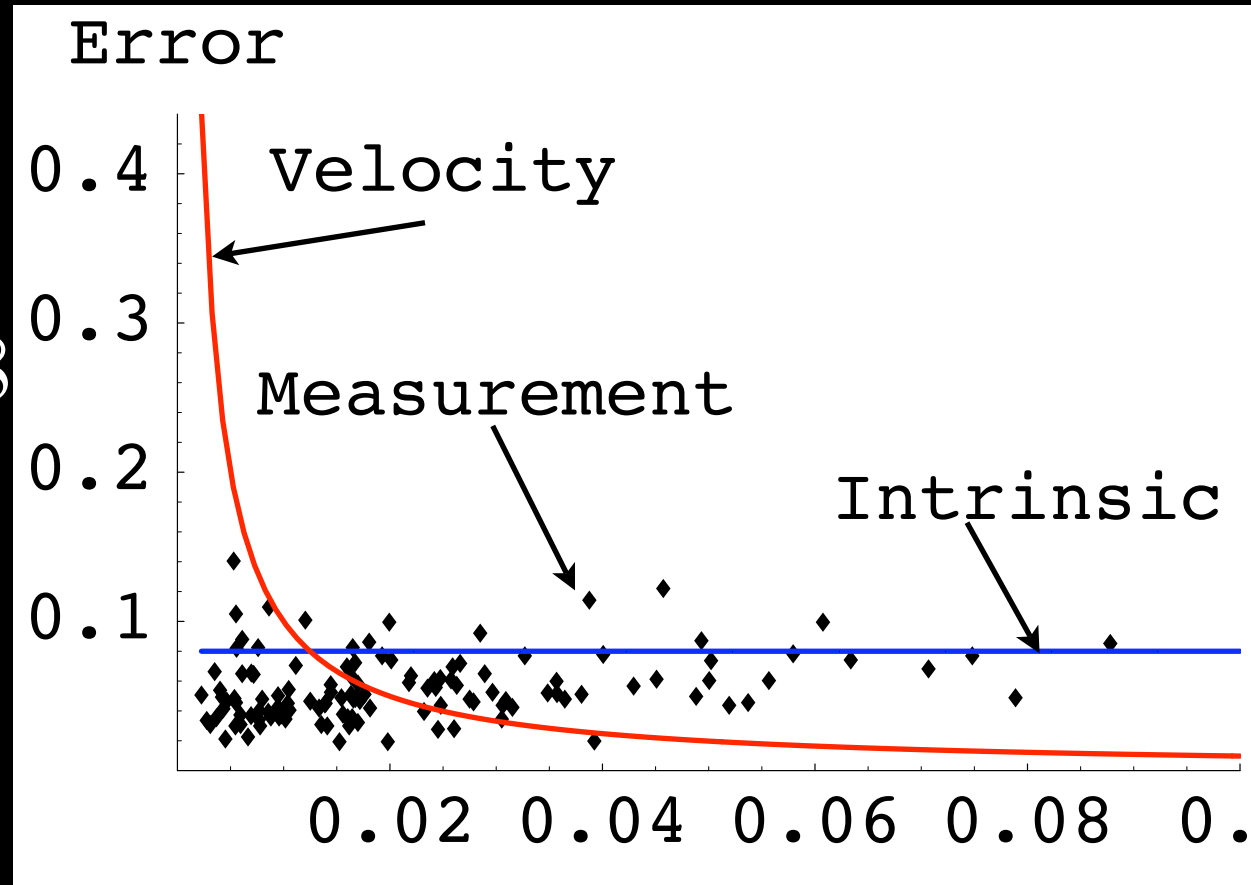
$$\xi_{\parallel, \perp} = D'(z_i) D'(z_j) \int_0^{\infty} \frac{dk}{2\pi^2} P(k) K_{\parallel, \perp}(kr)$$

$$K_{\parallel}(x) = j_0(x) - \frac{2j_1(x)}{x}, \quad K_{\perp}(x) = j_1(x)/x$$

- $z > 0.016$ (192 SNe), Davis et al. 2007.

- SN+WMAP:

$$w = -0.94 \pm 0.08$$

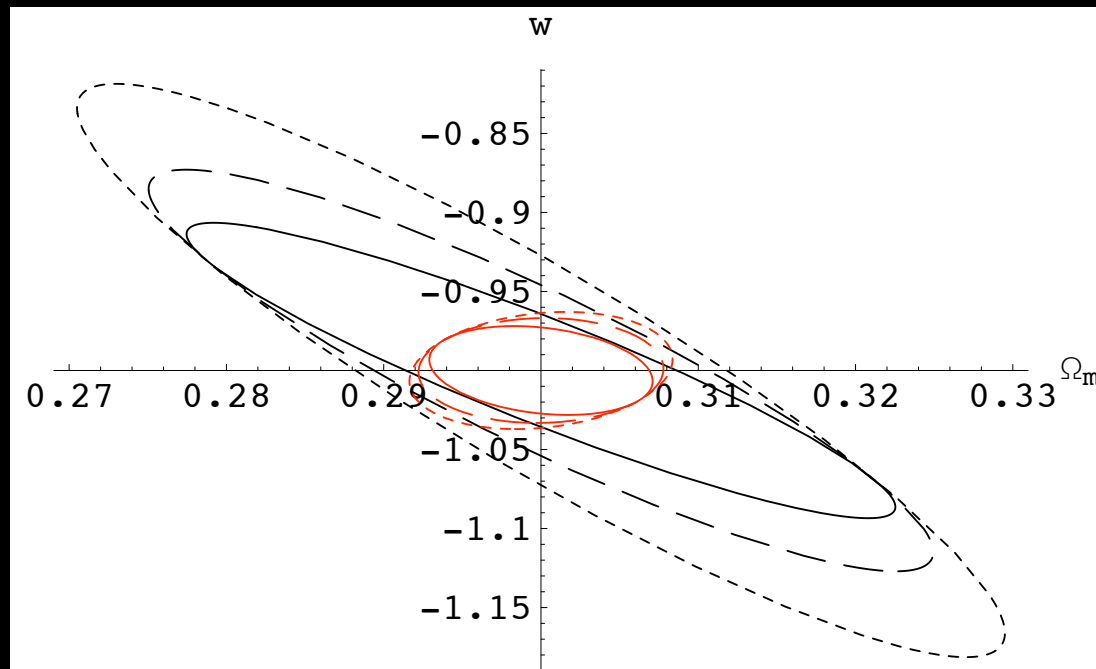


- SN+WMAP, no correlated PV:

$$w = -0.96 \pm 0.09$$

Forecasts

- Fisher Matrix:
$$F_{\alpha\beta} \equiv - \left\langle \frac{\partial^2 \mathcal{L}}{\partial p_\alpha \partial p_\beta} \right\rangle$$
$$F_{\alpha\beta} = \mathbf{d}_{,\alpha} \mathbf{C}^{-1} \mathbf{d}^T_{,\beta} + \frac{1}{2} \text{Tr} (\mathbf{C}^{-1} \mathbf{C}_{,\alpha} \mathbf{C}^{-1} \mathbf{C}_{,\beta})$$
- SNe factory: 300 SNe, $0.03 < z < 0.08$
- SNAP: 2000 SNe, $0.2 < z < 1.7$



Conclusions

- Current observations can ignore correlated peculiar velocities if $z > 0.016$.
- SN factory will need to account for correlated peculiar velocities even if $z > 0.03$.