Neutrino Theory

1105-1130 "Neutrino Theory" 20' + 5' discussion

- Should include general broad overview for non-experts
- Topic to be addressed include:
- o Neutrino physics and BSM at the TeV and GUT scale
- o Connection with leptogenesis
- o Relation to LFV processes such as $\mu \rightarrow e$ conversion and $\mu \rightarrow e\gamma$
- o lepton flavour physics, 013, CP violation, neutrino-less double beta decay o What are the important measurements and what are the phenomenological issues

(set the scene for the following experimental talks on reactor/LBL oscillation experiments/neutrinoless double beta decay/neutrino factory). ^VT o target for 20 years

Executive Summary (now since I will run out of time)

- Neutrino physics has surprised us all
- We have witnessed a revolution in the past decade
- Lepton Flavor is not conserved
- Neutrinos have tiny masses, not very hierarchical
- Neutrinos mix a lot
- At least 7 new parameters for SM
- Quite unlike quark mass and mixing
- Of all fermions, neutrinos are least understood
- Play an essential role in the Universe $n_v \sim n_v >> n_e$

Tiny neutrino mass indicates New Physics BSM

Impact on particle physics and cosmology

- Origin of tiny neutrino mass
 Extra dimensions, See-saw mechanism, SUSY
- Unification of matter, forces and flavour GUTs, Family Symmetry, Strings,...
- Did neutrinos play a role in our existence?
 Leptogenesis
- Did neutrinos play a role in forming galaxies?
 Hot/Warm Dark matter component → limits
- Did neutrinos play a role in birth of the universe?
 Sneutrino inflation
- Can neutrinos shed light on dark energy? $\Lambda \sim m_{v^4}$

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Three neutrino mass and mixing $\left|\begin{array}{c} \left(\begin{array}{c} V_{e} \\ e^{-} \end{array}\right) & \left(\begin{array}{c} V_{\mu} \\ \mu^{-} \end{array}\right) & \left(\begin{array}{c} V_{\tau} \\ \tau^{-} \end{array}\right) & \left(\begin{array}{c} \text{Standard Model} \\ \text{states} \end{array}\right)\right|$ Neutrino mass states m_3 Pontecorvo $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \checkmark$ Maki m_{γ} Nakagawa m_1 Sakata $U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\beta_{1}/2} & 0 \\ 0 & e^{i\beta_{2}/2} \\ 0 & 0 \end{pmatrix}$ $s_{ij} = \sin \theta_{ij}$ Reactor Solar Majorana Atmospheric $c_{ii} = \cos \theta_{ii}$ 3 masses + 3 angles + 1(3) phase(s)Oscillation phase δ = 7(9) new parameters for SM Majorana phases β_1, β_2

Neutrino mass squared splittings and angles



Global Fit to Atmospheric and Solar Data



There is a 2σ hint for θ_{13} being non-zero



The 2009 estimate includes the MINOS results which show a 1.5σ excess of events in the electron appearance channel

Tri-bimaximal (TB) mixing

Harrison, Perkins, Scott

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

TB angles $\theta_{12} = 35^{\circ}$, $\theta_{23} = 45^{\circ}$, $\theta_{13} = 0^{\circ}$.

c.f. data

$$\theta_{12} = 34.5^{\circ} \pm 1.4^{\circ}, \ \theta_{23} = 43.1^{\circ} \pm 4^{\circ}, \ \theta_{13} = 8^{\circ} \pm 2^{\circ}$$

Current data is consistent with TB mixing (ignoring the 2σ hint for θ_{13})

Useful to parametrize the PMNS mixing matrix in terms of deviations from TBM

$$s_{13} = \frac{r}{\sqrt{2}}, \quad s_{12} = \frac{1}{\sqrt{3}}(1+s), \quad s_{23} = \frac{1}{\sqrt{2}}(1+a)$$

$$SFK_{07}$$

$$0.14 < r < 0.24, \quad -0.05 < s < 0.02, \quad -0.04 < a < 0.10$$

r = reactor s = solar a = atmospheric

e.g. **r**≠ **0**, **s=0**, **a=0** gives Tri-bimaximal-reactor (TBR) mixing $U_{TBR} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}}re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1+re^{i\delta}) & \frac{1}{\sqrt{3}}(1-\frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}}(1-re^{i\delta}) & -\frac{1}{\sqrt{3}}(1+\frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}} \end{pmatrix}$ SFK '09

TBR not as simple as TB but is required if $\theta_{13} \neq 0$

Oscillation formulae in terms of r,s,a, δ

$$\begin{split} \hline P_{\alpha\beta} &= P(\nu_{\alpha} \to \nu_{\beta}) & \Delta_{ij} = 1.27 \Delta m_{ij}^2 L/E \\ \text{Reactor} \Big\{ P_{ee} &= 1 - 2r^2 \sin^2 \Delta_{31} - \frac{8}{9} \Delta_{21}^2 & \text{See talks by E.Falk, D.Wark} \\ \text{L} \begin{bmatrix} P_{\mu e} &= r^2 \sin^2 \Delta_{31} + \frac{4}{9} \Delta_{21}^2 + \frac{4}{3} r \Delta_{21} \sin \Delta_{31} \cos(\Delta_{31} + \delta) \\ P_{\mu \mu} &= 1 - (1 - 4a^2) \sin^2 \Delta_{31} - \frac{2}{9} (1 + 3 \cos 2\Delta_{31}) \Delta_{21}^2 \\ &+ \frac{2}{3} (1 - s - r \cos \delta) \Delta_{21} \sin 2\Delta_{31}. \end{split}$$

For formulae including matter effects see SFK arXiv:0710.0530

Steve King, Birmingham

14/07/2009

mass

mixing

C P

violatior

Implications of mass/mixing measurements





Dirac mass

 $0\nu\beta\beta$ experiments will decide See talk by Y. Ramachers

Cosmology and $0\nu\beta\beta$ experiments are complementary







Leptogenesis

- •Right-handed neutrinos are produced in early universe and decay out of equilibrium giving net lepton numbers L_e , L_u , L_τ
- •CP violation from complex Yukawa couplings
- •Out of equilibrium Boltzmann eqs lead to L_{e} , L_{μ} , L_{τ} partial washouts
- -Surviving L_e, L_{μ}, L_{τ} are processed into B via B-L conserving sphalerons





Clearly no link in general since 3x3 Yukawa matrix involves $9 \rightarrow 6$ phases

However in particular models the number of Yukawa phases may be reduced e.g. sequential dominance decouples a RH neutrino \rightarrow 3 phases, and a texture zero removes another \rightarrow 2 phases, then can show $|\phi_{COSMO}| = |\phi_{\beta\beta0\nu}|$ SFK

Other interesting examples in flavour dependent LG: Antusch, SFK, Riotto; Pascoli, Petcov, Riotto





Lepton Flavour Violation in mSUGRA



New Physics at the GUT scale t S GUT Family —

Tri-bimaximal mixing suggests a family symmetry (c.f. quark **Colour**)

GUT x Family symmetry $\rightarrow \theta_{12} \approx 35.3^{\circ} + \theta_{13} \cos \delta$

