Elements of Discovery Process Lecture I



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Elements of Discovery Process



Theoretical Insight







Gravitation electromagnetism

weak force strong/nuclear force

Forces due to exchange of particles:



Effect of force ~ intrinsic strength ("muscle power") + mass of carrier (~ range)

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Gravitation electromagnetism weak force strong/nuclear force

Gravity	Electromagnetic	Weak	Strong
Graviton (not observed)	Photon	W⁺, W⁻, Z	Gluon
10 -41	1	0.8	25





Gravitation electromagnetism weak force strong/nuclear force



Add another particle, the Higgs gives mass to other particles

- A few fundamental particles
- A few forces
 mediated by bosons
- Higgs to give mass.



The Standard Model has been incredibly successful in explaining all data... ...but there are problems too.

The Standard Problems: Higgs?

From precision Electroweak measurements Mass of SM Higgs boson: 87⁺³⁵₋₂₆ GeV but we know 114 < m_H < 157 GeV

Where is the Higgs?



Standard Problems: antimatter?

What happened to all the antimatter?



The imbalance is a trillion times bigger than the model predicts.

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The Standard Problems: Dark Matter?

23 % Dark Matter

- Inferred from gravitational effects
- Rotational speed of galaxies
- Orbital velocities of galaxies in clusters
- Gravitational lensing

- 73 % Dark Energy
 - Accelerated expansion of universe

Only 4 % made out of known matter.....





More Problems!

"Cracks" have started to appear in the Standard Model...

Many problems identified over time

- No explanation of
 - masses,
 - coupling constants
- Why three families?
- Gravity not included
- The "hierarchy" problem, fine tuning.

...and yet it explains the data

The Standard Model isn't so much wrong as it is *incomplete*

Is SUSY the "next SM"?

- Supersymmetry (SUSY)
 - SM particles get partners (-spin ½)
- Unifications of forces possible
 - SUSY changes running of coupling
- Dark matter candidate exists:
 - The lightest neutral partner of the gauge bosons
- No (or little) fine-tuning required
 - cancellation of loop corrections





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What else is there beside SUSY?

Extra spatial dimensions

- Addresses hierarchy problem
- make gravity strong at TeV scale
- Extra gauge groups: Z', W'
 - Occur naturally in GUT scale theories

Leptoquarks:

 Would combine naturally the quark and lepton sector

New/excited fermions

- More generations? Compositeness?
- something not thought of yet



Redman's Theorem

Sorting out the structure of the universe can be frustrating...

... just ask the astronomers

"Any competent theoretician can fit any given theory to any set of facts."



Quoted in M. Longair's *High Energy Astrophysics*, sect 2.5.1: "The Psychology of Astronomers and Astrophysicists": Prof. R. Redman, 1905-1975

First we need data...

S.Worm

Accelerator Advances

Accelerator Advances

At the LHC we can create and study new interactions



Accelerator Advances

- Each advance is a revolution...
 - but sadly only once or twice per generation
- To help understand our excitement about the LHC
 - Previous energy record-holder (Tevatron) started in 1983 27 years
 - LEP at CERN stopped in 2000 10 years ago



Role of Colliders

- Colliders key tool for discovering particles we know today
 - Anti-proton (LBNL, 1955)
 - Quarks (SLAC 1969)
 - W- and Z-boson (CERN, 1983)
 - Top-quark (FNAL, 1994)
 - Image: plus many more

Basic principle follows from E=mc²

- If collider energy ≥ mass of particle
 - particle can be produced

New Energy Regimes

- Each advance in territory makes new discovery possible
- Many historic examples...
- Hunting for "bumps" in the mass spectra



S.Worm

More energy, more particles



Some Fun Facts about the LHC



http://www.sustain.ucla.edu/media/images/facts_general.jpg







Proton on Proton Collisions at 14 TeV*



7 TeV

7 TeV

- 1 TeV ~ kinetic energy (KE) of a mosquito
- 10¹¹ protons in a bunch = KE of a London bus
- LHC beam stores 700 MegaJoules



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The emptiest (largest) vacuum in the solar system





Ten times more atmosphere on the Moon than inside LHC beam pipes

Colder than outer space

LHC operates at 1.9 K 38000 tones cool mass 120 tones of Helium



The largest refrigerator ever

Sources of ultra high energetic particles



27 highest energy cosmic rays detected by Auger 2004 - 2007

E > 6x10¹⁹ eV=60x10⁶ TeV

LHC is safe!

J. Ellis, http://indico.cern.ch/conferenceDisplay.py?confld=39099

- LHC@14 TeV=cosmic ray@10¹⁷eV
- ~ 3.10²² cosmic rays >10¹⁷ e\
 have struck Earth
- Equivalent to 10⁵ LHC programmes
- Area of Sun 10⁴ larger
- 10¹¹ stars in galaxy
- 10¹¹ galaxies in Universe
- Nature has performed 10³¹ LHC programmes
- Nature carries out 3.10¹³ LHC programmes per second

arXiv:0806.3414v2 [hep-ph]



ultra-high-energy cosmic rays up to 10²⁰ eV

Luminosity

- Single most important quantity
 - Drives ability to observe new rare processes

$$L = \frac{f * n_{\text{bunch}} * N_{p}^{2}}{4\pi * \sigma_{x} * \sigma_{y}}$$

- revolving frequency f = 11245.5/s
- n_{bunch} = 2808
- $N_p = 1.15 \times 10^{11}$ Protons/Bunch
- Area of beams: $4\pi\sigma_x\sigma_y\sim40 \ \mu m$
- Rate of physics processes per unit time ~ L

$$N_{Obs} = \int Ldt * \epsilon * \sigma_{process}$$
 Cross section; given by nature; predicted by theory finited by experimentalists Maximize $N_{obs} \rightarrow max \epsilon and L$

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2010 Proton Proton Collisions

- First 7 TeV collision:
 - 30.03.2010
- End of proton run:

04.11.2010

$$\int Ldt = 45 pb^{-1}$$



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Detectors and Computing



The Detector Development Loop Today



Interplay of Detectors and Discoveries





Georges Charpak (1924 – 2010)

Invented Multi-wire proportional chambers (1968) which rapidly replaced bubble chambers

Nobel Prize in Physics in 1992
Searches and (old fashioned) Detectors:



New/better detector → new physics found → Nobel prize (simple, isn't it?)

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Modern Detectors











Detector Mass in Perspecitive



CMS is 30% heavier than the Eiffel tower

Durham Cathedral and ATLAS



Silicon Detectors







- Silicon strip and pixel detectors
 - Pixels used for first time at hadron colliders
 - Huge!
 - area of CMS silicon ~200 m²

gdem Issever Like a football field!

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Muon Systems and Calorimeters







Particle Collisions - "Events"

The raw data recorded is later reconstructed, filtered, and then analysed.



Finding the interesting Events

- A lot more "uninteresting" than "interesting" processes
 - at design luminosity (L=10³⁴ cm⁻²s⁻¹)
 - Any event: 10⁹ / second
 - W boson: 150 / second
 - Top quark: 8 / second
 - Higgs (150 GeV): 0.2 / second

The Trigger Systems



What is it for?







A.Barr

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Analysis of the Data



S. Worm



We do not know what is out there for us... A large variety of possible signals. We have to be ready for that

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Early discoveries? E.g. Di-lepton Resonance



How to make a Discovery?

This is a tricky business!

Lot's of complicated statistical tools needed at some level

- But in a nutshell:
 - Show that we have a signal that is inconsistent with being background
 - Number of observed data events: N_{Data}
 - Number of estimated background events: N_{Bg}
 - Number of observed data events to be inconsistent with background fluctuation:

Background fluctuates statistically: $\sqrt{N_{Bq}}$

- Significance: S/ $\sqrt{B}=(N_{Data}-N_{Bg})/\sqrt{N_{Bg}}$
 - Require typically 5σ
 - Increases with increasing luminosity: S/ $\sqrt{B} \sim \sqrt{L}$
 - All a lot more complex with systematic uncertainties...

B.Heinemann

The Discovery of the Oops-Leon

Physical Review Letters 36: 1236–123

New subatomic particle "discovered" at Fermilab in



╢

"less than one chance in fifty" that this is due to random coincidence BUT 1977 data showed that is was such a coincidence

"Five-sigma Rule"

Commonly-accepted standard (acid test)

- \mathbb{N}_{obs} > 5 σ above the expected level of the background
 - 99.9999% of events fall within 5σ
 - Iess than one in a million chance
- Today :Oops-Leon "discovery" would have not been published.

Five-sigma rule is far from golden

5σ "discoveries" can vanish overnight

"The statistical analysis is based upon the assumption that you know everything and that everything is behaving as it should. But after everything you think of, there can be things you don't think of. A fivesigma discovery is only five sigma if you properly account for systematics."

Val Fitch, 1980 Nobel Prize for discovering charge-parity violation in K mesons.



Kinematic Constraints and Variables

Longitudinal momentum and energy, p_z and E

Visible p_z is not conserved

Polar angle θ

• θ is not Lorentz invariant

$$y = \frac{1}{2} ln \frac{E + p_z}{E - p_z}$$

$$y = \eta = -\ln(\tan\frac{\theta}{2})$$



Kinematic Constraints and Variables

Transverse momentum, pT, very useful!
Colliding partons p_T≈0
parton
parton
z

- Vector sum p_T conserved: ∑_i p_{Ti}≈0
 - If non zero something escaped detection
 - Missing transverse momentum: ∑_i p_{Ti}
- Scalar sum p_T: ∑_i |p_{Ti}
 - Measure of "Umpf" in the final state

2010 (Exotics) Searches at the LHC

Bump Hunting

- Dijet, dilepton, diphoton, dijet+dilepton final states
- Excited quarks, Z', W', RS Graviton, Leptoquarks....
- Search for deviations in the tails
 - Digamma + Met final states (UED, SUSY)
 - Multi-object final states (Strong Gravity, Black Holes)

Search for odder things

Long lived particles

Overview of Dijet Resonance Search

Select inclusive dijet events, and plot dijet mass

$$m_{jj} = \sqrt{(E_{j1} + E_{j2})^2 - (\vec{p}_{j1} + \vec{p}_{j2})^2}$$

Background: fit a smooth function to DATA*

$$f(x) = p_1 (1 - x)^{p_2} x^{p_3 + p_4 lnx}$$
$$x = m_{jj} / \sqrt{s}$$

- Search for discrepancies between data and background.
- If no discrepancies found, set limits *CDF, Phys.Rev.D79:112002,2009

Fitting MC



Event Selection

- High Data Quality
- p_T^{j1}>150 GeV
- p_T^{j2}>30 GeV
- M_{ii}>350 GeV

- **Reject poorly** measured jets
- | η | < 2.5 and | $\Delta\eta$ |<1.3



ATLAS Preliminary

 $|\eta_{j1} - \eta_{j2}|_{max}$

Data and Background



Model independent search for discrepancy

- BumpHunter *
- TailHunter*
- Likelihood, χ², KS, Jeffreys Divergence

*<u>http://arxiv.org/abs/1101.0390</u>, Phys.Rev.D79:011101,2009 and Phys.Rev.D78:012002,2008

BumpHunter and TailHunter

BumpHunter:

- scan the spectrum for a local excess
- surrounded by agreeing sidebands.

TailHunter:

Similar. Check all high-mass tails for an excess.

Nothing assumed about mass, or width of the signal.

BumpHunter Demo



BumpHunter Demo



Georgios Choudalakis

BumpHunter Demo



Georgios Choudalakis

Results

- 6 tests, independently, indicated that the data are consistent with the background-only hypothesis.
- Data were so well-fitted that the p-values were ~ 99%.
 - \rightarrow Perfectly consistent.
 - Multiple tests agreed that this was just a coincidence.

NOTHING FOUND

Back to setting limits...

Limits on excited Quarks



	Observed Limit	Expected Limit
MRST2007	1.53 TeV	1.51 TeV
CTEQ6LI	1.45 TeV	1.43 TeV

Observed ~ **Expected**

	/1					
			PDF Set	Stat.	Stat. only	Stat.
ATLAS-CONF-2010-093		MC09 [3]	MRST2007 [4]	1.53	1.64	1.51
	2010-003	MC09'	CTEQ6L1 [5]	1.45	1.56	1.43
	2010-095	Perugia0 [6]	CTEQ5L [7]	1.49	1.60	1.46
	Cigdem i	2267CI				/1

Why set limits and publish?

Mapping of our knowledge



Quantifying search with help of bench mark models
 Comparison with other models via bench mark models
Credits

Some slides taken from

- Alan Barr
- Beate Heinemann
- Georgios Choudalakis
- Chris Llewellyn Smith
- Tom Lecompte
- Phil Burrows
- Tara Shears
- John Ellis
- David Britton
- Steven Worm
- Theory friends: Glenn Starkman, Dejan Stojkovic and Dechang Dai
- Wikipedia and CERN public web pages