Machine learning techniques for BSM phenomenology

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Work with Alistair Shilton and Martin White [arXiv:1106.4613]

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The problem

SUSY and other models have many parameters

Interpretation of experimental data fastest if unfolding of detector effects can be avoided

But detector simulation is *very* slow – 30 mins per event for ATLAS full sim!

How to get coverage of the model space for constraints?

In general *many* params. But almost all pheno done in restricted models such as 5 param CMSSM. ATLAS and CMS constrain in m_0 , $m_{1/2}$ only – definitely improveable!

Super-fast simulation for MC tuning

- We already deal with a similar problem in MC tuning!
- No full simulation, but again we have many parameters: MPI or hadronisation ⇒ O(5 – 10) params
- 1-10M events needed per energy/process to evaluate a param point. How to find the optimal point?
- Serial MCMC sampling intractable. PROFESSOR tuning system builds a polynomial *parameterisation* of every bin of every observable as a function of generator params
- Numerical optimisation is then trivial. It works!





Parameterising an ATLAS SUSY analysis

Idea: use PROFESSOR to parameterise detector level cut-pass yields as a function of BSM model params.

Example: ATLAS' 2010 data 0-lepton analysis, signal region D [arXiv:1102.5290]:

- ► \geq 3 jets with p_T > 40 GeV and p_T^{i1} > 120 GeV,
- ► $E_T^{\text{miss}} > 100 \text{ GeV},$ $\Delta \phi(\text{jet}, p_T^{\text{miss}})_{\text{min}} > 0.4,$
- $m_{\rm eff} > 1000 \, {\rm GeV}, E_T^{\rm miss}/m_{\rm eff} > 0.25$



ATLAS constrains $m_0 \& m_{1/2}$ for tan $\beta = 3$, $A_0 = 0$, $\mu > 0$

A first attempt...

1000 HERWIG+AcerDet points from 4D CMSSM with $\mu > 0$: parameterise #events passing Region D cuts. A 1-bin observable!



Unphysical negative values on boundary!

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Showing $\log_{10}(\%$ deviation): accuracy typically within 10%

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Definite artefacts in MC / ipol correlation. Not so good...

Improving the parameterisation

What went wrong? Problems particularly at the boundaries: fitted polynomials are not trustworthy near the edges of the space.

BUT – we don't need an algebraically constructed parameterisation since there is only one observable: fitting speed is not a serious issue. Thresholds and discontinuities in the param space.

Idea #2: use a parameterisation-free interpolating function:

- ▶ FBM Bayesian neural net
- SVMHeavy support vector machine

Training time for one bin \sim 12 hrs!



Sampling details

Sample from the 4D CMSSM with $\mu > 0$ again:

Parameter	Minimum	Maximum
$m_{1/2}$ [GeV]	50	1200
m_0 [GeV]	50	1200
A_0 [GeV]	-1000	1000
$\tan\beta$	2	60

20000 points each of 50k HERWIG+AcerDet events $\Rightarrow \sigma^{(D)}$

BNN/SVM will train to optimally fit the max *number* of training points. Implicit measure on the sampling space. Flat sampling undersamples high- σ = low $m_{1/2}$ region. Sample $m_{0,1/2} \sim e^{-x/\lambda}$ with λ = 200 GeV.



Validation

Predicted vs. true $\sigma^{\rm (D)}$ and the Poisson likelihood L used to make the exclusion contour

BNN with 5k points; independent training and test sets



Excellent performance on both measures

Validation

Predicted vs. true $\sigma^{\rm (D)}$ and the Poisson likelihood L used to make the exclusion contour

SVM with 5k points; independent training and test sets



Slightly better at high $\sigma^{(D)}$ but overall slightly worse *L*





































Exclusion contours from coarse grid scanning



BNN/SVM 5k



Compatible 95% exclusion results in $m_0-m_{1/2}$ plane for "ATLAS-like" fixed tan $\beta = 3$, $A_0 = 0$.

Summary and outlook

- BSM model exclusion becomes exp. weaker as # params increases
- Particularly when detector (full or fast) simulation needed
- Parameterisation has been a big success in MC tuning: idea has been applied to an example BSM parameter space and analysis
- PROFESSOR polynomial ipols and SVD are crippled by parameterisation bias. For small numbers of observables, neural net or SVM approaches are much better
- Outlook:
 - extend to more complex signatures and models
 - use multiple observables in likelihood, handle systematics
 - use newer/better MC generators + NLO cross-sections
 - better detector simulation: AtlFast-2 in ATLAS, ??? outside

Thank you!

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