



Reweighting

- NNPDF Users Guide
- Reweighting
- Unweighting
- NNPDF2.2

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Monte Carlo PDFs (eg NNPDF)

Giele & Kosower 1998

Forte & Latorre 2002

- Choose a very flexible functional form for each PDF:
(eg a neural network: ~ 250 params)
 - Generate data replicas ($\sim 100-1000$) using exp uncertainties
 - Find a **good** fit to each data replica by optimising χ^2
(best fit useless – fitting statistical noise:
instead use genetic algorithm + cross-validation)
 - Treat resulting PDF replicas as statistical ensemble:
each equally probable (importance sampling)
- So simple averages give central values, uncertainties etc.

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Advantages:

No theoretical bias due to parametrization

Statistically meaningful uncertainties: no need for tolerance

Technical stability: improved data/theory, same parametrization

NNPDF progress

- 2002: Structure Functions
- 2005: More Structure Functions
- 2007: Nonsinglet DIS partons
- 2008: First NLO DIS: NNPDF1.0
- 2009: Strange PDFs: NNPDF1.2
- 2010: First NLO DIS+DY+J: NNPDF2.0
- 2010: **Rewighting** (W-ev asymmetry) arXiv:1012.0836
- 2011: Global NLO + HQ: NNPDF2.1 NLO arXiv:1101.1300
- 2011: NNPDF2.1 LO and NNLO arXiv:1107.2652
- 2011: First fit with LHC data: **NNPDF2.2** arXiv:1108.1758

Major software development project

Users Guide

NNPDFs

- Download a set of NNPDFs (eg NNPDF2.1) from LHAPDF
- Each set of contains an ensemble of N ‘replicas’ (N=100,1000)
- Each replica f_k , $k=1 \dots N$ is a set of PDFs:
 $\{g, u, \bar{u}, d, \bar{d}, \dots\}$ on a grid in x and Q^2 – just as usual
- Each replica f_k is **equally** probable as a candidate PDF.

For any observable $O[f]$ depending on PDFs f :

$$\langle O[f] \rangle = \frac{1}{N} \sum_{k=1}^N O[f_k]$$

“Master formula”: all results are obtained using this

There are no “eigenvector sets” in NNPDF

Example 1: the PDFs

- Central values:

$$f_0 = \langle f \rangle = \frac{1}{N} \sum_{k=1}^N f_k$$

Note: f_0 is also given on LHAPDF as “set zero”

- Variances:

$$\text{Var}[f] = \langle (f - \langle f \rangle)^2 \rangle = \frac{1}{N} \sum_{k=1}^N (f_k - f_0)^2$$

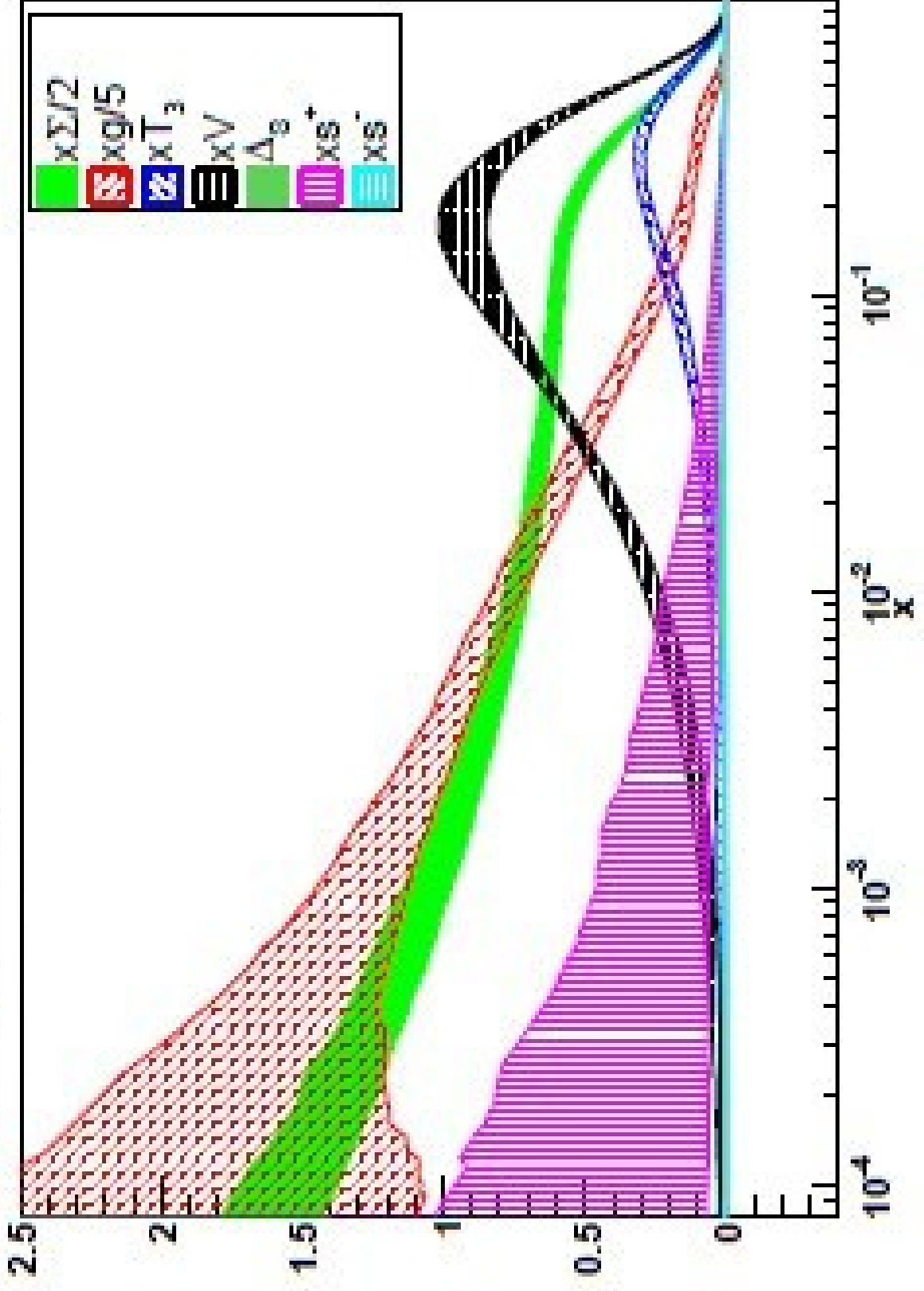
- Correlations: e.g.

$$\text{Corr}[f, f'] = \langle (f - \langle f \rangle)(f' - \langle f' \rangle) \rangle = \frac{1}{N} \sum_{k=1}^N (f_k - f_0)(f'_k - f'_0)$$

- Confidence levels (e.g. interval with 68% replicas inside)
- etc, etc

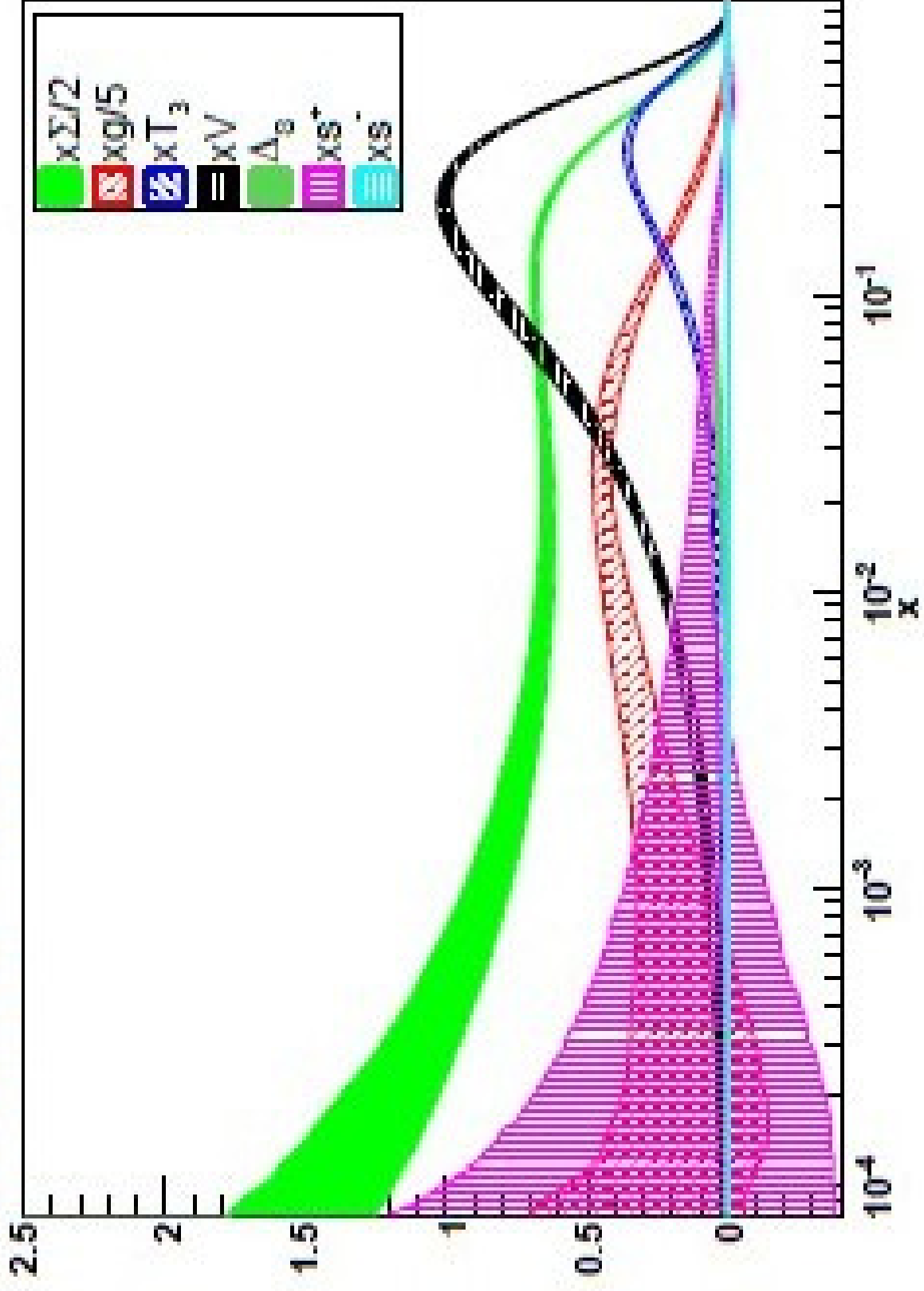
NNPDF2.1 LO, $Q^2 = 2 \text{ GeV}^2$

68% C.L.



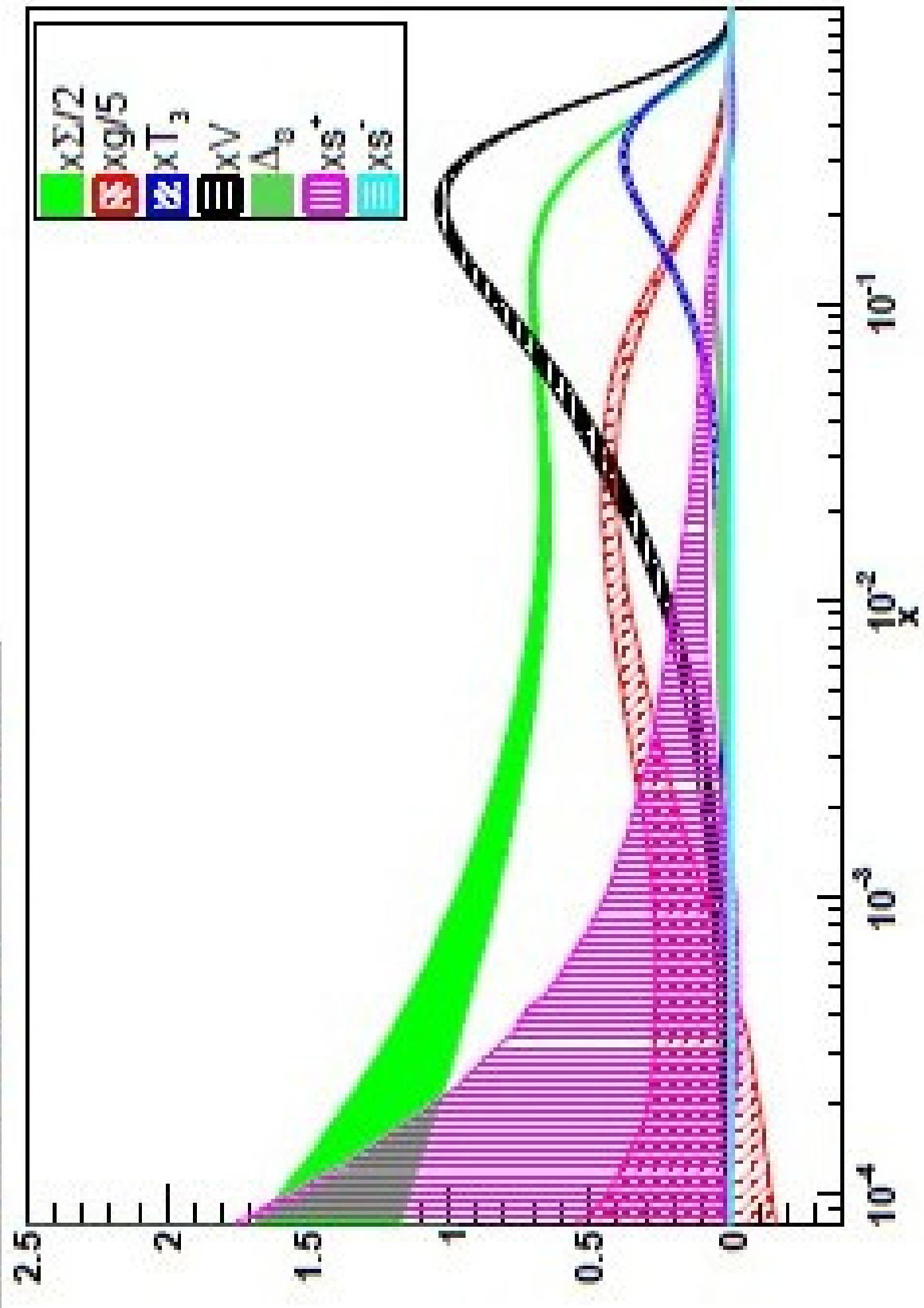
NNPDF2.1 NLO, $Q^2 = 2 \text{ GeV}^2$

68% C.L.



NNPDF2.1 NNLO, $Q^2 = 2 \text{ GeV}^2$

68% C.L.



Example 2: DIS xsecs

DIS xsecs $\sigma[f]$ depend **linearly** on the PDFs

- Central values:

$$E[\sigma] = \langle \sigma[f] \rangle = \frac{1}{N} \sum_{k=1}^N \sigma[f_k] = \sigma[f_0]$$

- Variances:

$$\begin{aligned} \text{Var}[\sigma] &= \left\langle (\sigma[f] - \langle \sigma[f] \rangle)^2 \right\rangle \\ &= \frac{1}{N} \sum_{k=1}^N (\sigma[f_k] - \sigma[f_0])^2 \end{aligned}$$

- etc, etc

$\sigma[f]$ can be anything you like: str fn, red xsec, jet xsec,

Example 3: Hadronic xsecs

Hadronic xsecs $\sigma[f,f]$ depend **quadratically** on the PDFs

- Central values:

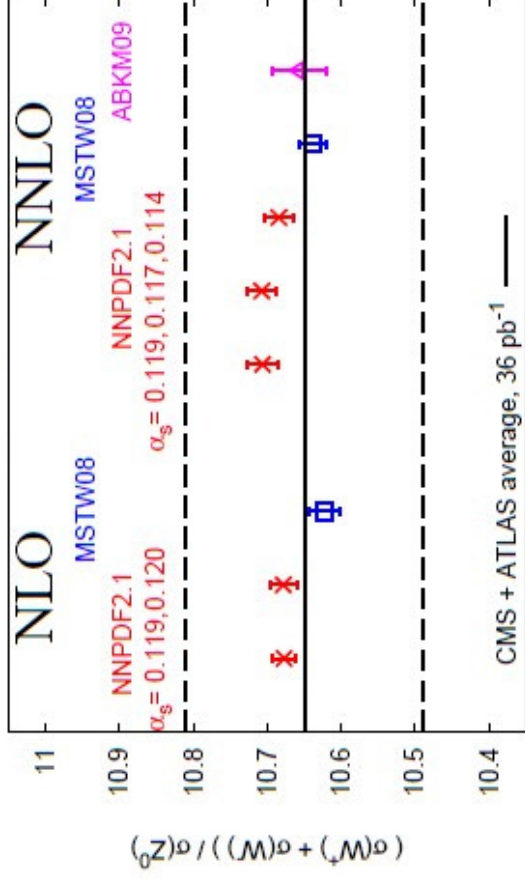
$$E[\sigma_h] = \langle \sigma_h[f, f] \rangle = \frac{1}{N} \sum_{k=1}^N \sigma_h[f_k, f_k] \approx \sigma_h[f_0, f_0]$$

- Variances:

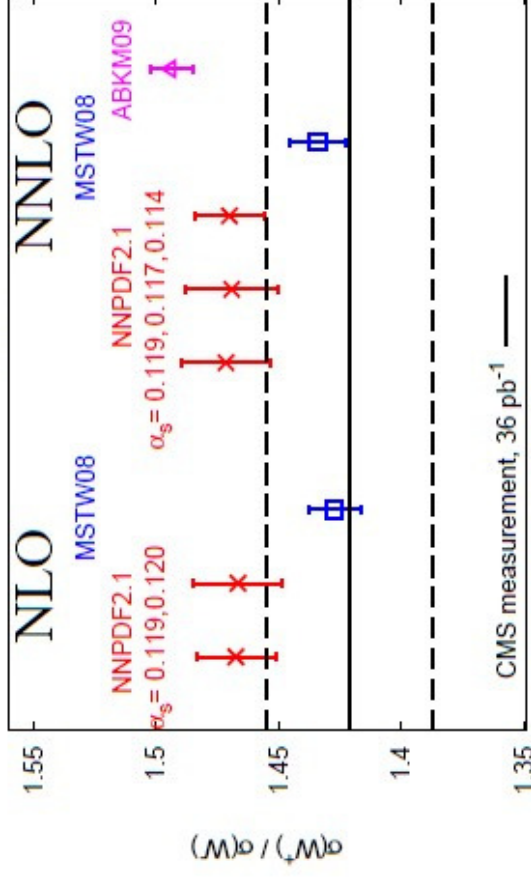
$$\begin{aligned} \text{Var}[\sigma_h] &= \left\langle (\sigma_h[f, f] - \langle \sigma_h[f, f] \rangle)^2 \right\rangle \\ &= \frac{1}{N} \sum_{k=1}^N \left(\sigma_h[f_k, f_k] - \frac{1}{N} \sum_{k=1}^N \sigma_h[f_k, f_k] \right)^2 \\ &\approx \frac{1}{N} \sum_{k=1}^N (\sigma_h[f_k, f_0] + \sigma_h[f_0, f_k] - 2\sigma_h[f_0, f_0])^2 \quad \text{if Gaussian} \end{aligned}$$

The approximate expressions can be evaluated more quickly (smaller N)

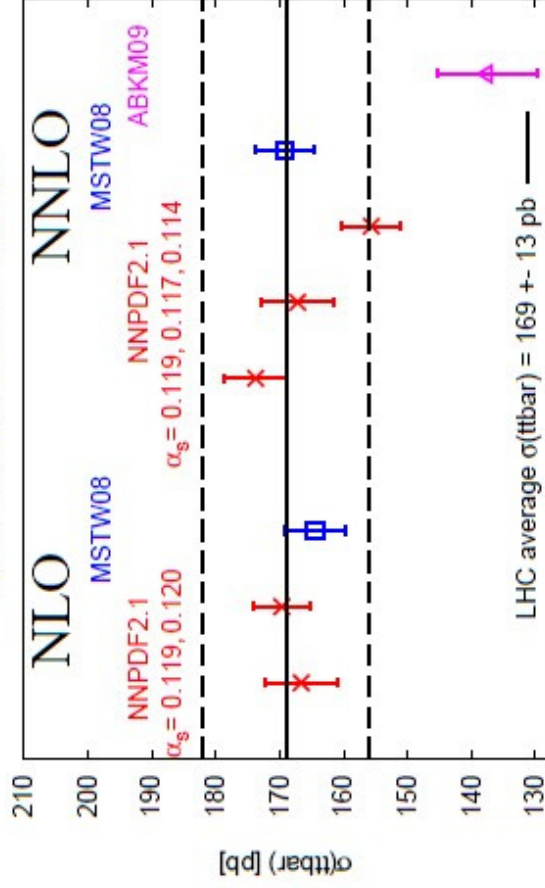
LHC 7 TeV, VRAP



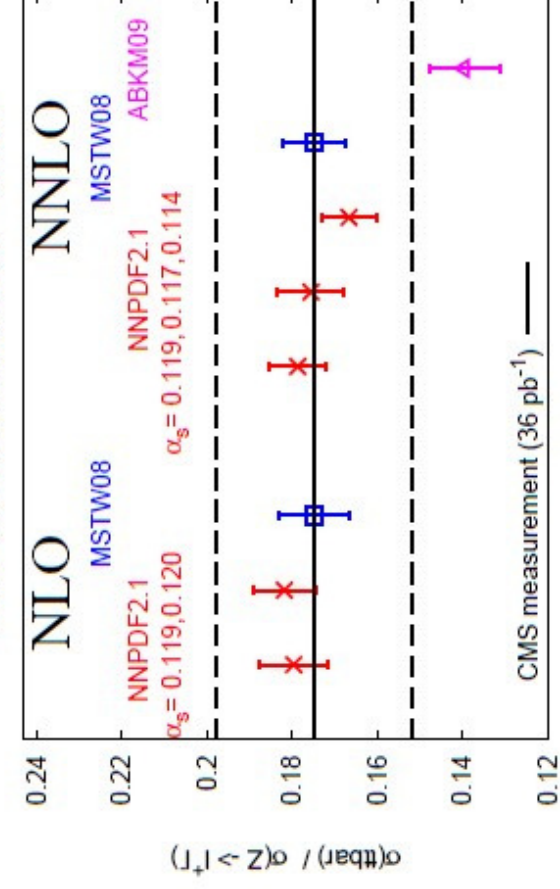
LHC 7 TeV, VRAP



LHC 7 TeV, HATHOR, mt = 172 GeV



LHC 7 TeV, HATHOR + VRAP, mt = 172 GeV



Three FAQs

Q: how many replicas N do I need?

A: depends on required accuracy: fluctuations fall as $1/\sqrt{N}$
typically use f_0 for central values, $\sim 100 f_k$ for variances etc

Q: which replicas should I use?

A: any random selection! – all replicas are equally probable

Q: for hadronic xsecs, should I use the exact or approx formulae?

A: error from using approx is $O(\text{Var}/E^2)$

(so for typical 10% uncertainty, error is $O(1\%)$)

and variance formula neglects non-Gaussian errors

But: when you use MSTW or CTEQ, you do this all the time...

Reweighting

Reweighting

All replicas are equally probable (importance sampling):

$$\langle \mathcal{O}[f] \rangle = \frac{1}{N} \sum_{k=1}^N \mathcal{O}[f_k]$$

Now add a **new** dataset $\{y_i, i = 1, \dots, n\}$

Q. What effect does this have on the PDFs?

A. The replicas are no longer equally probable: instead

$$\langle \mathcal{O}[f] \rangle_{\text{new}} = \frac{1}{N} \sum_{k=1}^N w_k \mathcal{O}[f_k]$$

w_k are the ‘weights’: probability of replica f_k given new data:

No need to refit!

Calculating the weights

w_k are the probabilities of replica k given n new data:

$$w_k \propto (\chi_k^2)^{\frac{1}{2}(n-1)} e^{-\frac{1}{2}\chi_k^2} \quad \frac{1}{N} \sum_{k=1}^N w_k = 1$$

$$\chi_k^2 = \sum_{i,j=1}^n (y_i - y_i[f_k]) \sigma_{ij}^{-1} (y_j - y_j[f_k]).$$

So... if you can plot new data y_i and compare with theory $y_i[f_k]$, then you can compute w_k and thus effect of data on PDFs

DIY PDFs

Loss of efficiency: replicas no longer have equal probability

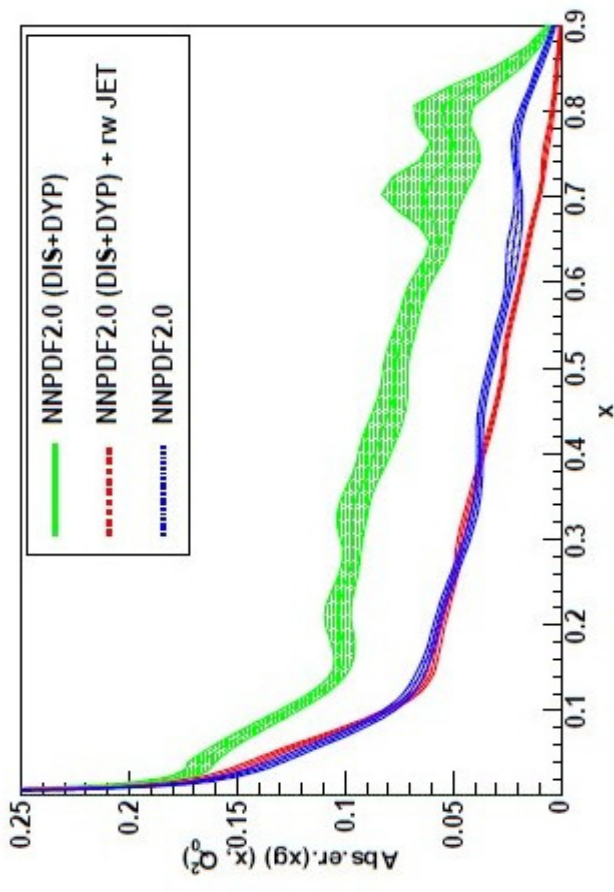
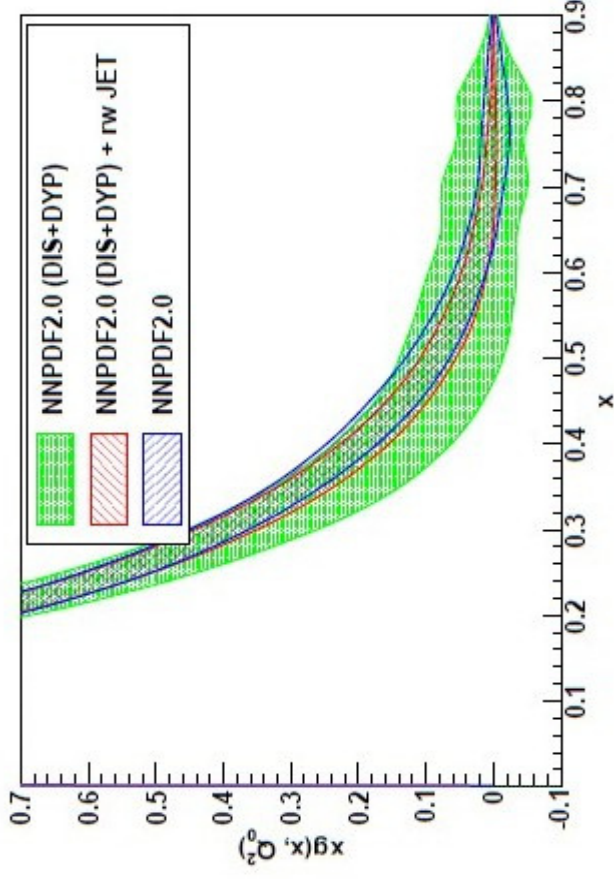
$$N_{\text{eff}} \equiv \exp \left[- \sum_{k=1}^N p_k \ln p_k \right] \quad p_k \equiv w_k / N \quad \text{cf Shannon entropy}$$

N_{eff}/N : gives measure of **impact** of new data

Does reweighting work?

Example:

- 1) take fit of DIS+DY data only
- 2) add (CDF+D0) inclusive jet data by reweighting
- 3) compare to result of fit using all the data DIS+DY+jet



Impact of jet data: with $N=10000$, have $N_{\text{eff}}=332$ left

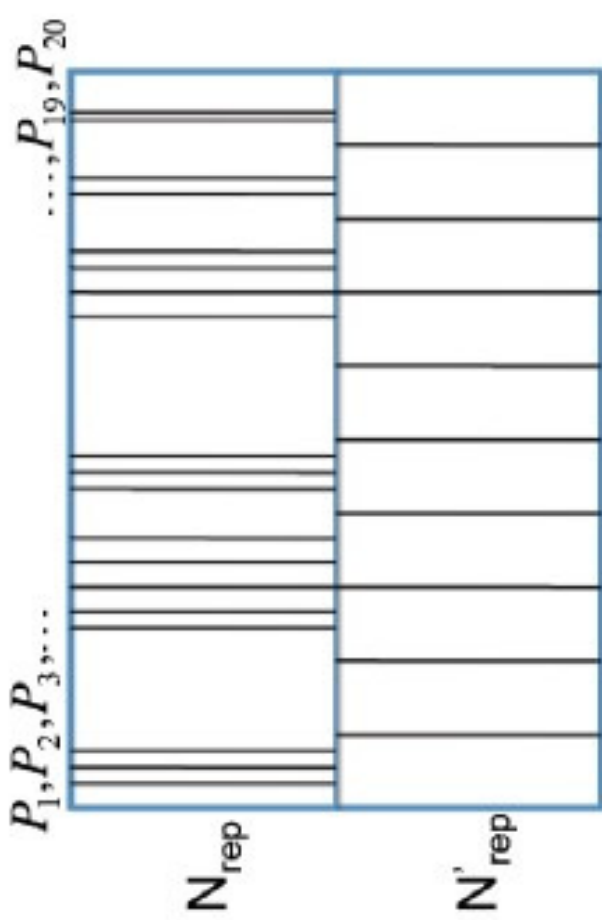
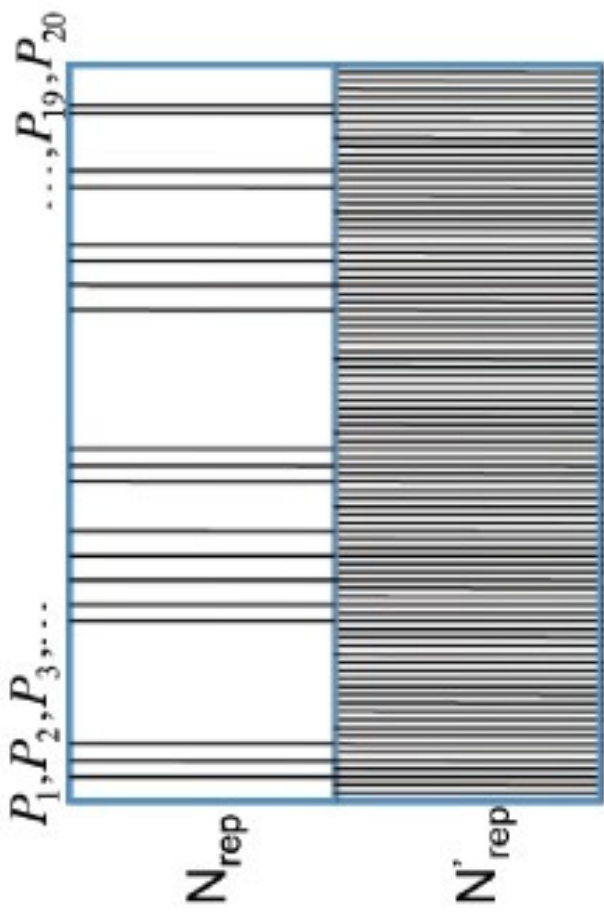
Note that if N_{eff} had been **too** small, would need to refit (or start with more replicas)

Unweighting

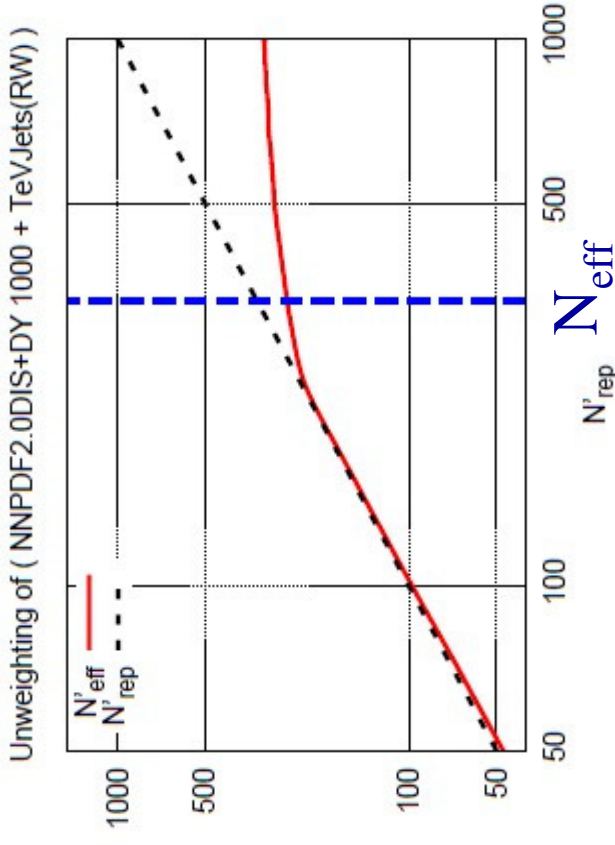
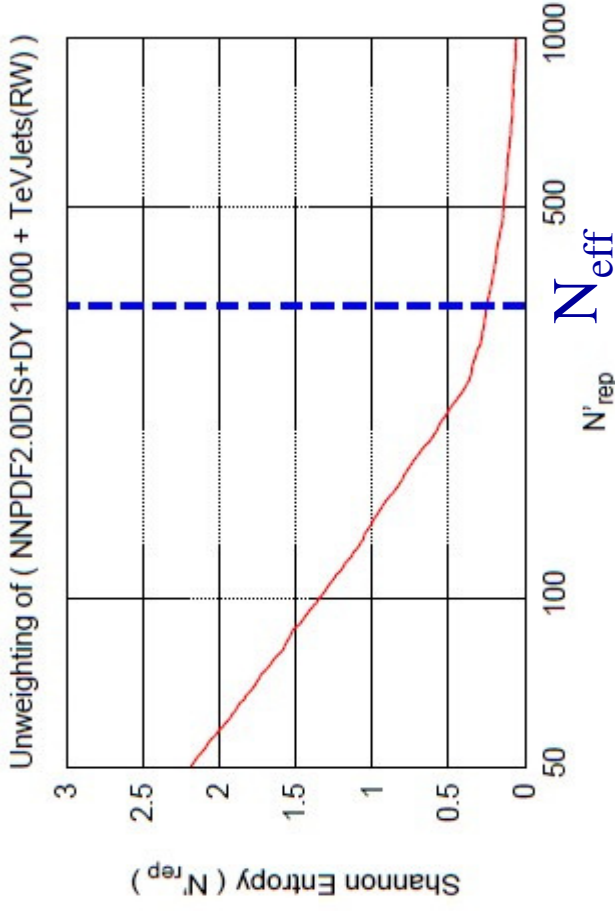
- Choose replicas according to weights, treated as probabilities
- Get subset of equally probable replicas equivalent to weighted set
- Treat just the same as a fitted set!
- In practice:

$$N'_{\text{rep}} \lesssim N_{\text{eff}} < N_{\text{rep}}$$

(so need N_{rep} large: typically $N_{\text{rep}} \sim 1000$, $N'_{\text{rep}} \sim 100$)



Does unweighting work?



$$H_R(N'_{\text{rep}}) = \sum_{k=1}^{N_{\text{rep}}} p'_k \ln \frac{p'_k}{p_k}$$

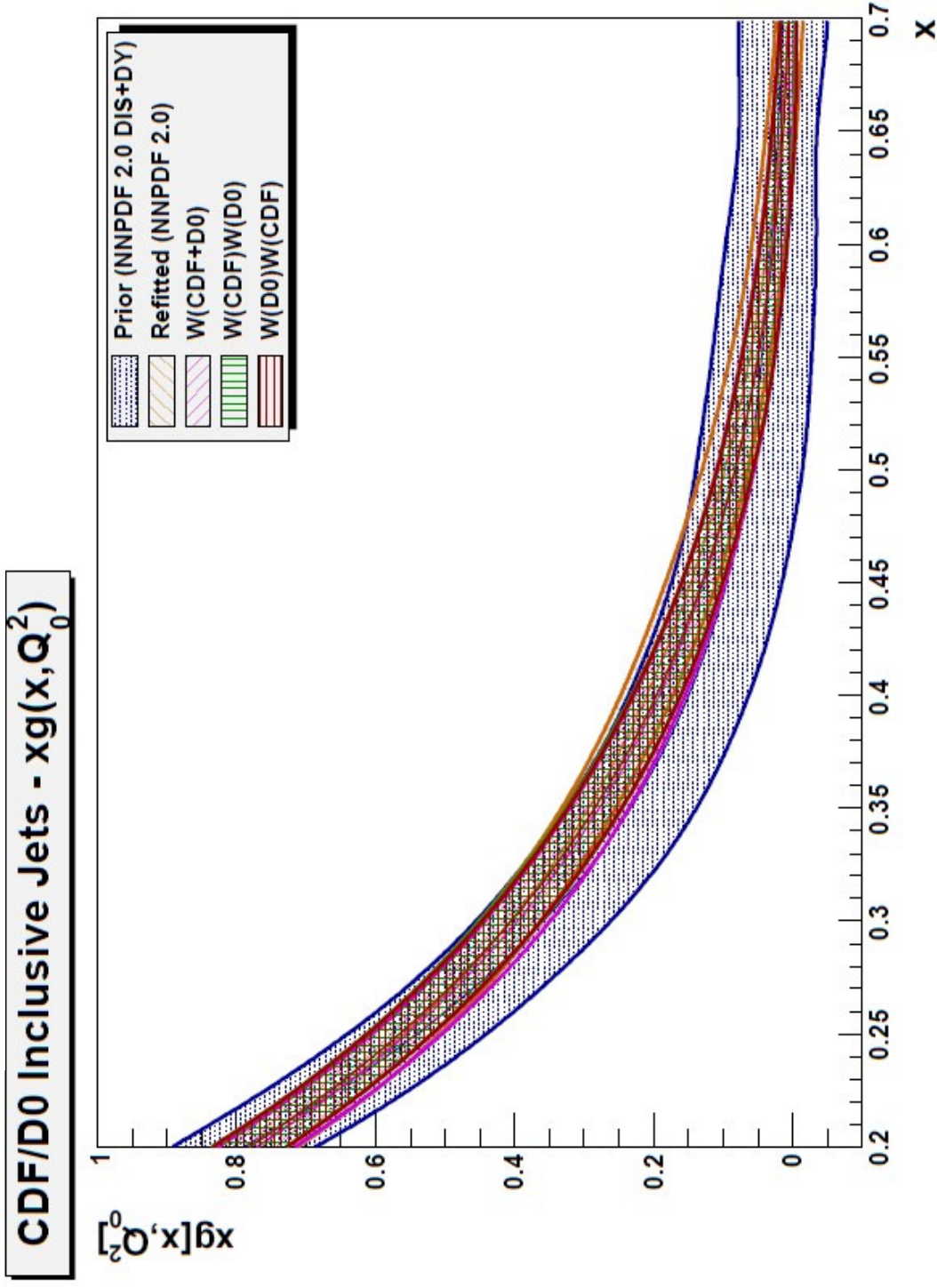
Relative entropy of unweighted
and weighted

$$N'_{\text{eff}} \equiv \exp \left[- \sum_{k=1}^{N_{\text{rep}}} p'_k \ln p'_k \right]$$

Number of effective replicas in
unweighted set

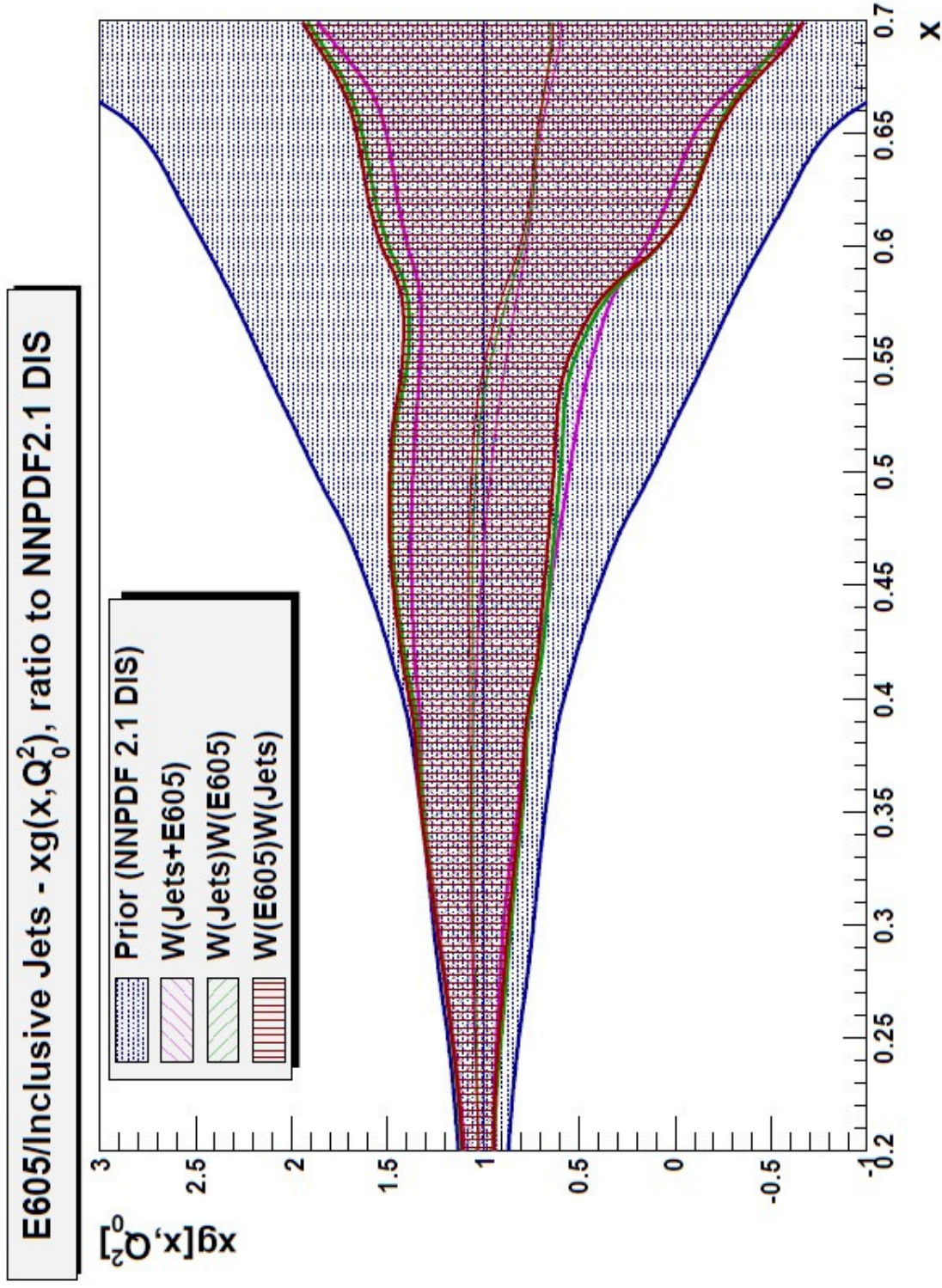
Consistency: two new sets, or one?

$$\hat{U} \hat{R}_{12} = \hat{U} \hat{R}_2 \hat{U} \hat{R}_1 = \hat{U} \hat{R}_1 \hat{U} \hat{R}_2.$$



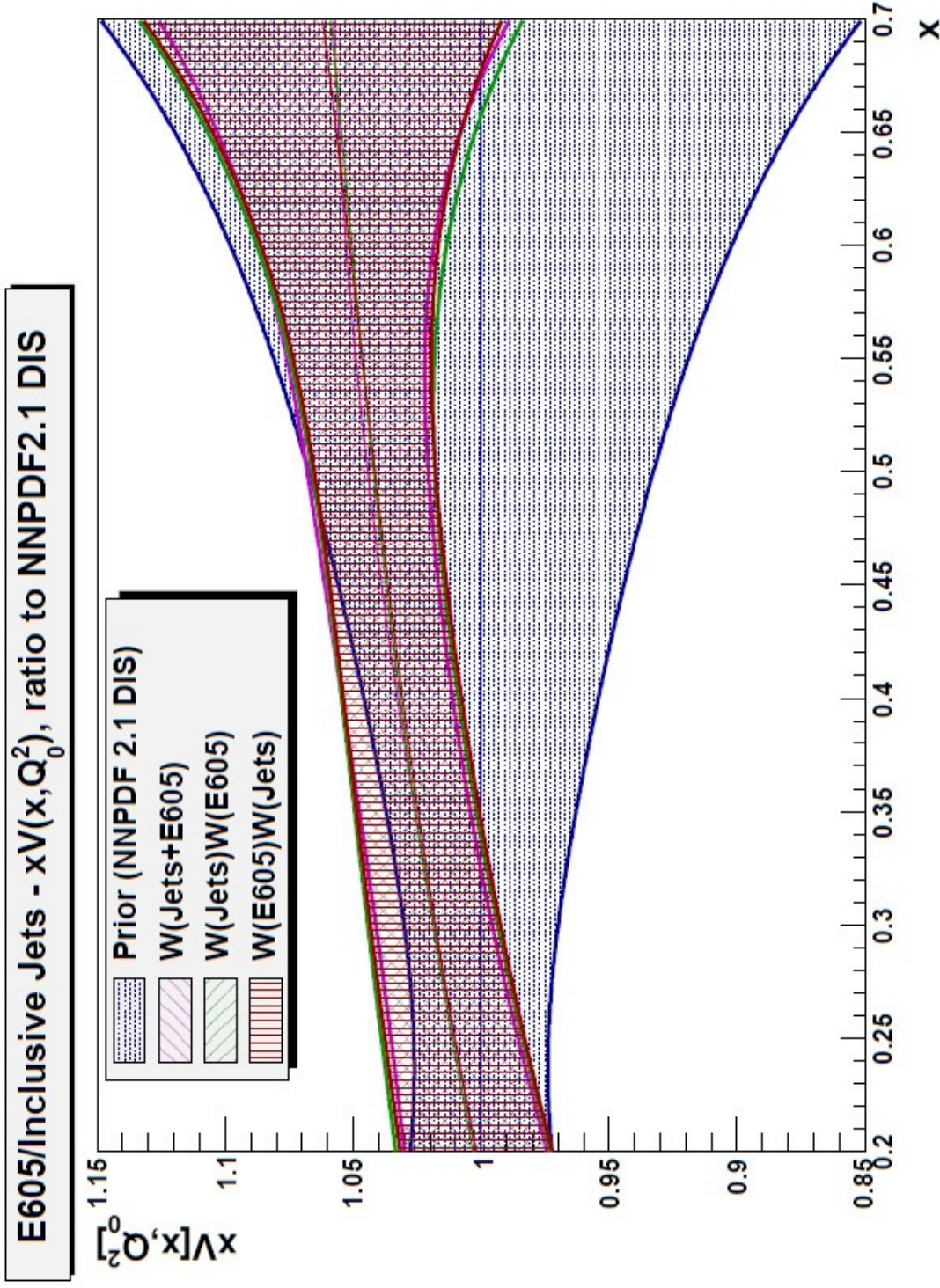
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Consistency: two new sets, or one?

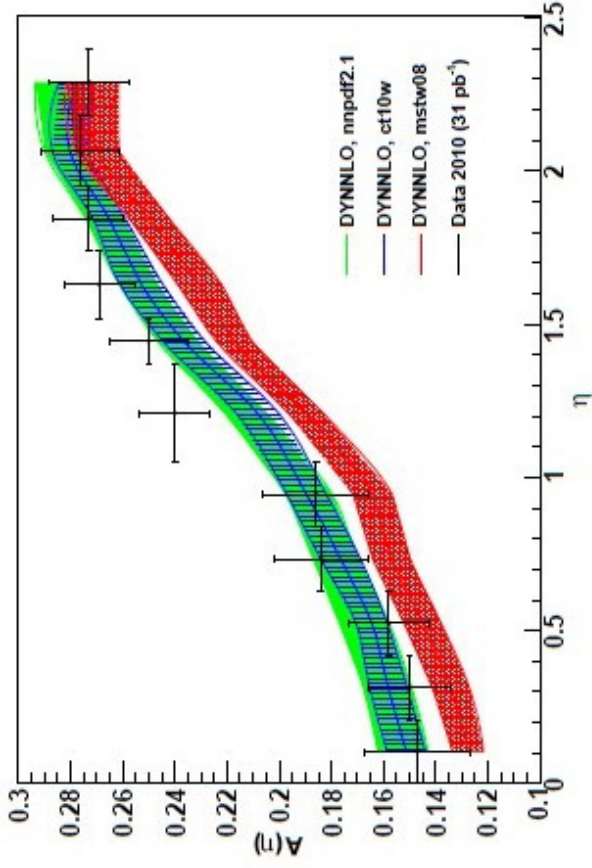
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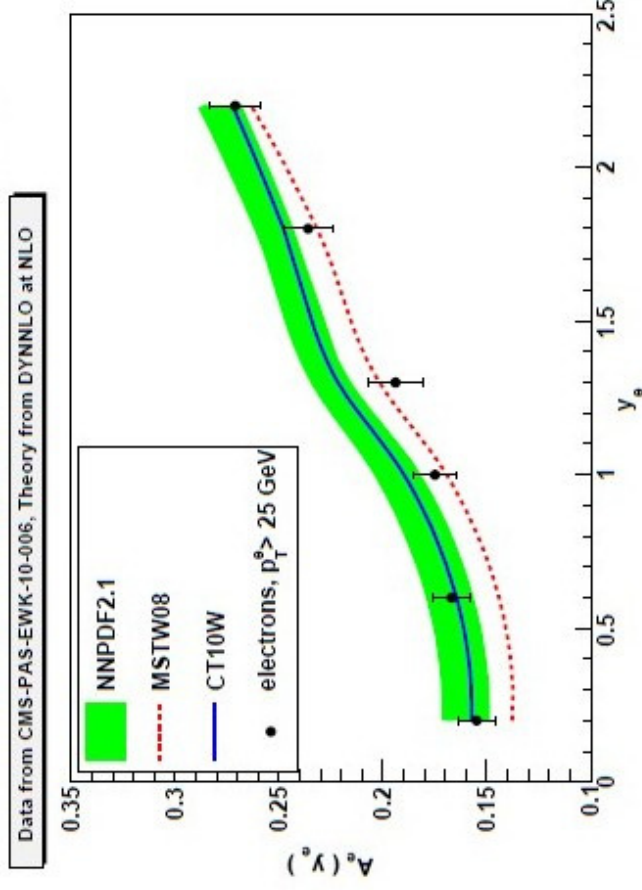
Mar 2011

W-lepton asymmetry

ATLAS



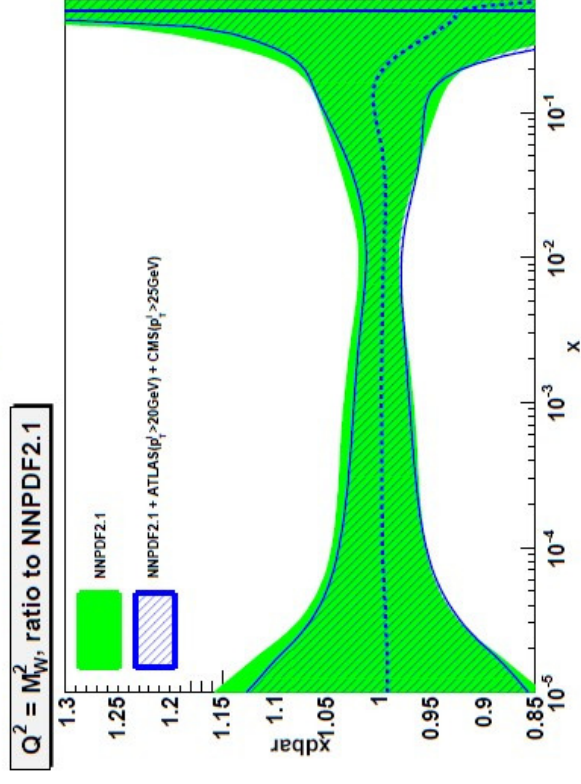
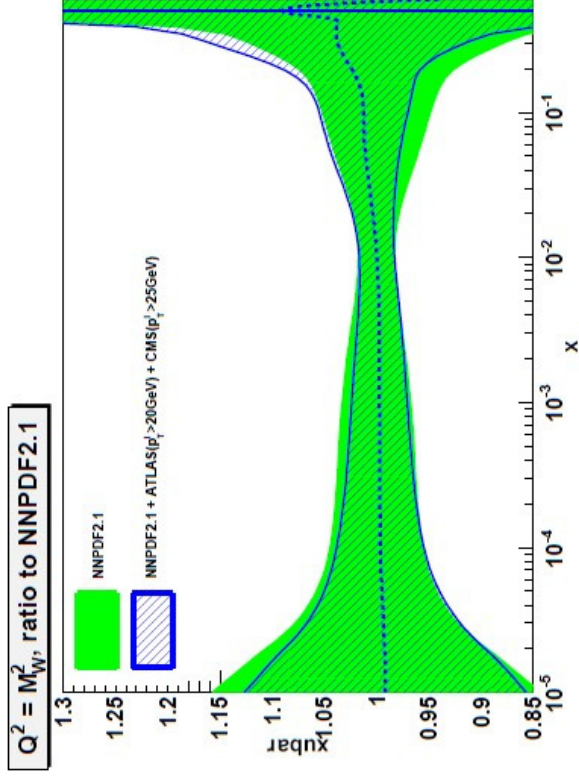
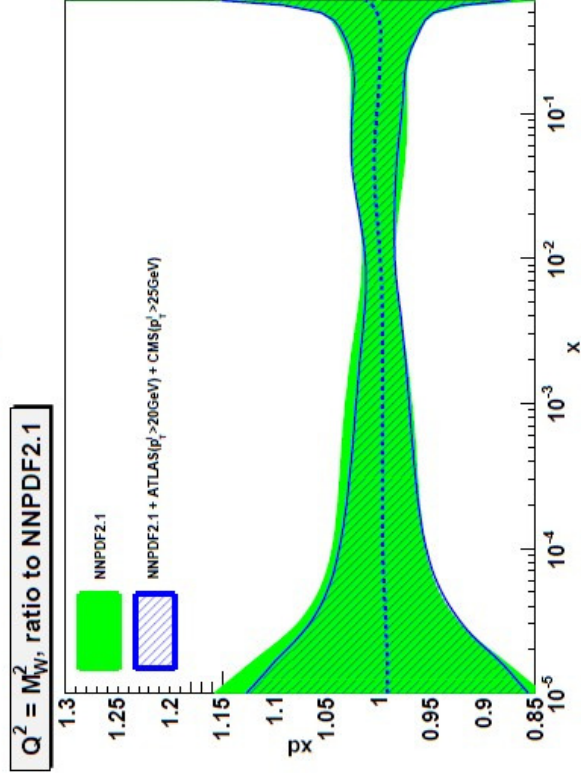
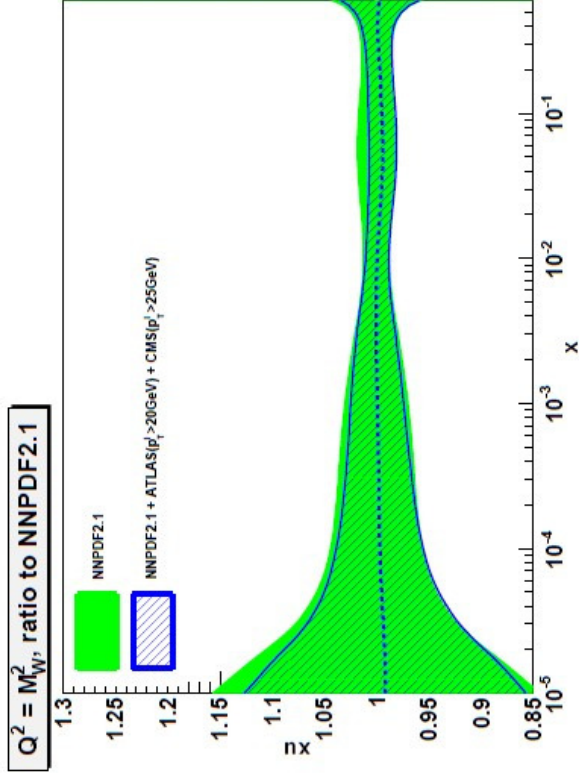
CMS



	N_{dat}	NNPDF2.1	CT10	MSTW08
ATLAS(31pb ⁻¹)	11	0.76	0.77	3.32
CMS(36pb ⁻¹) electron $p_T > 25$ GeV	6	1.83	1.19	1.70
CMS(36pb ⁻¹) muon $p_T > 25$ GeV	6	1.24	0.73	0.77

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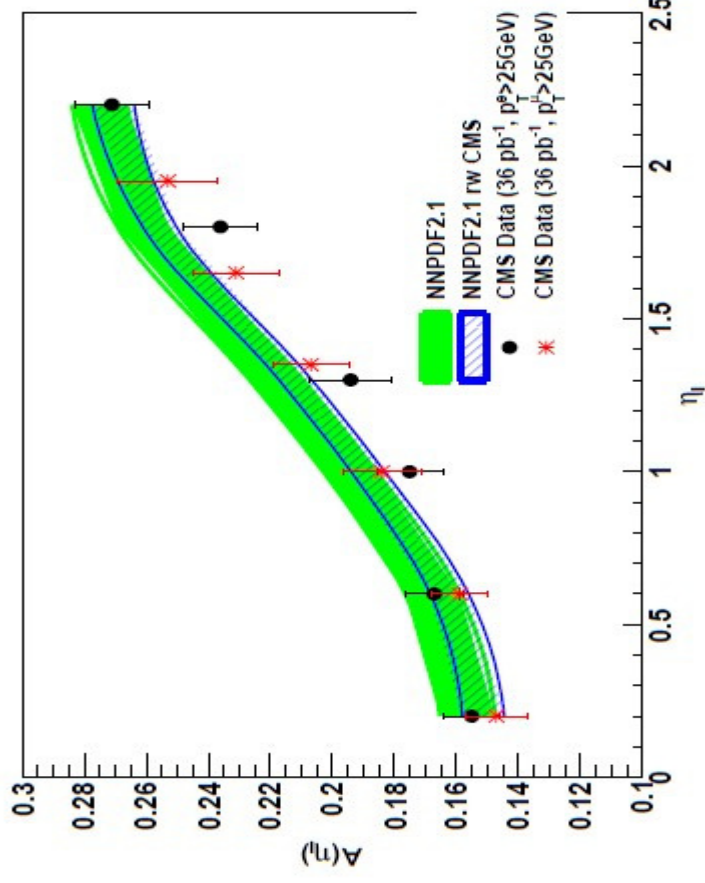
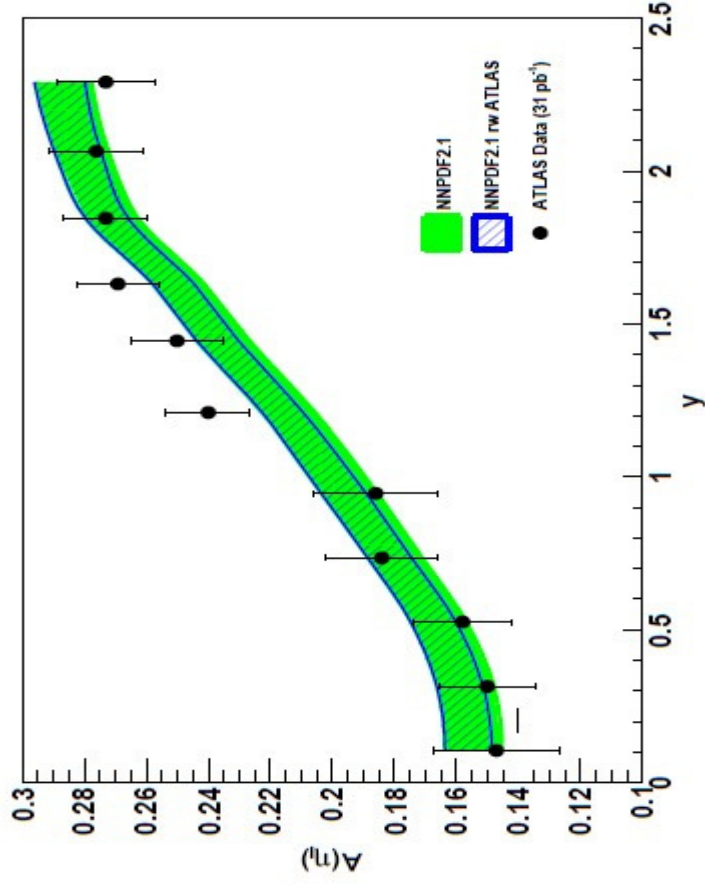
W-lepton asymmetry



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W-lepton asymmetry

After Reweighting



	N_{dat}	NNPDF2.1	+ LHC Wasy
ATLAS(31pb ⁻¹)	11	0.76	0.97
CMS(36pb ⁻¹) electron $p_T > 25$ GeV	6	1.83	1.23
CMS(36pb ⁻¹) muon $p_T > 25$ GeV	6	1.24	0.63

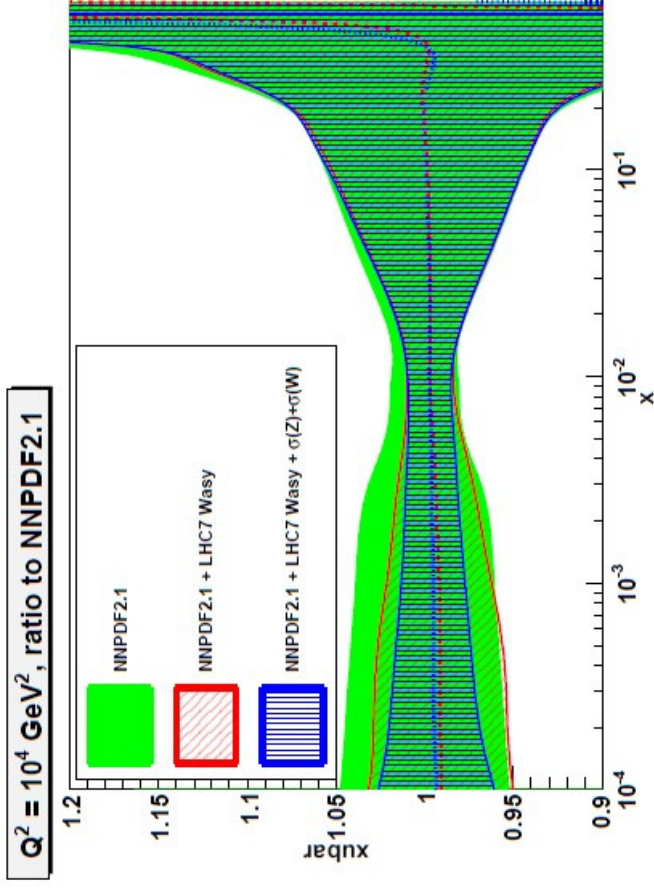
W-lepton asymmetry: future prospects

ATLAS LHC7

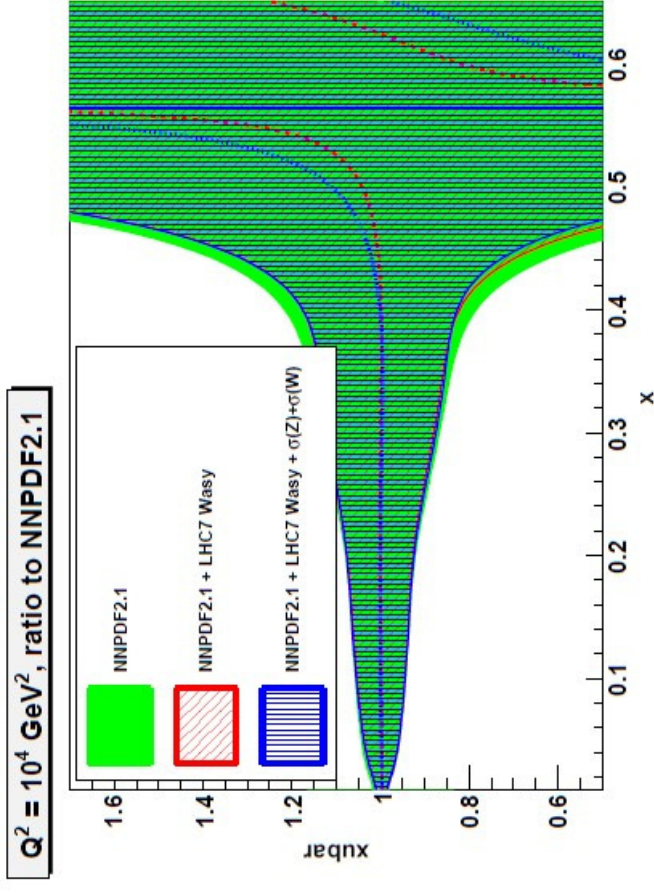
W lepton asymmetry measured to $\sim 5\%$ (kinematics courtesy A. Glazov)

W & Z total xsecs measured to 2%

ANTIUP SMALL x



ANTIUP LARGE x



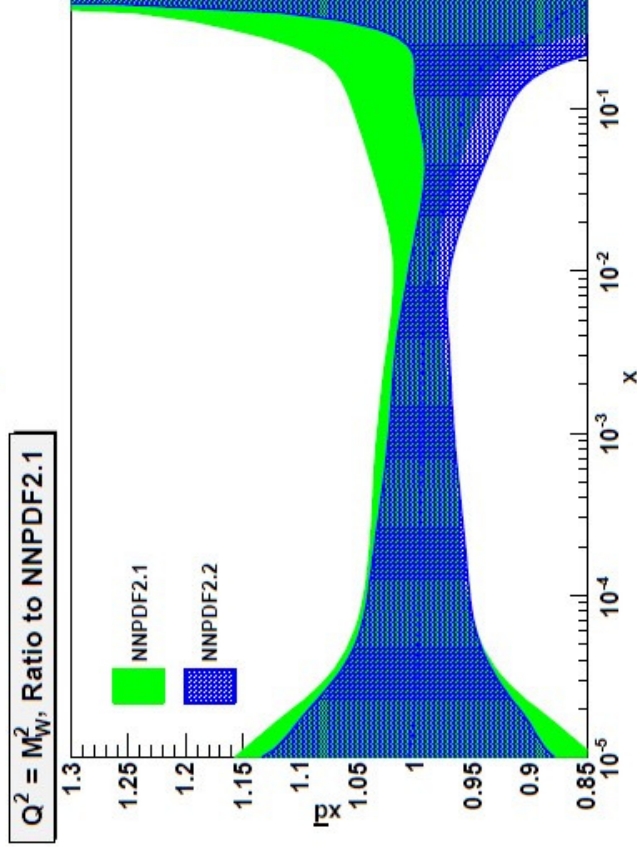
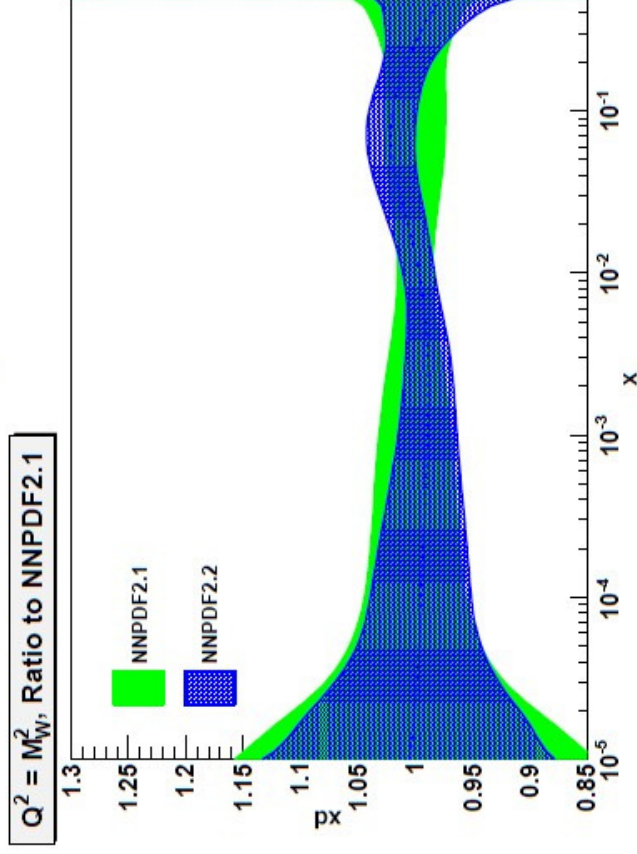
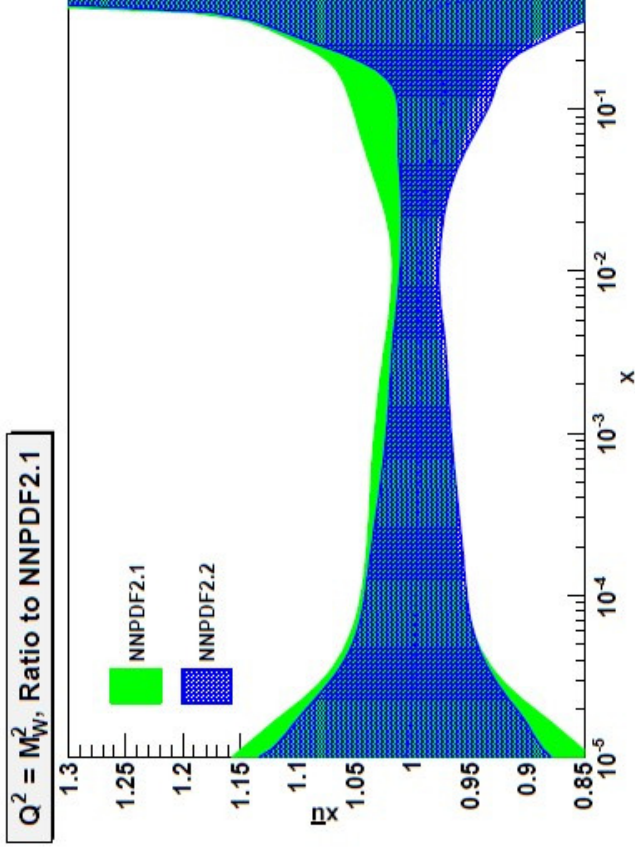
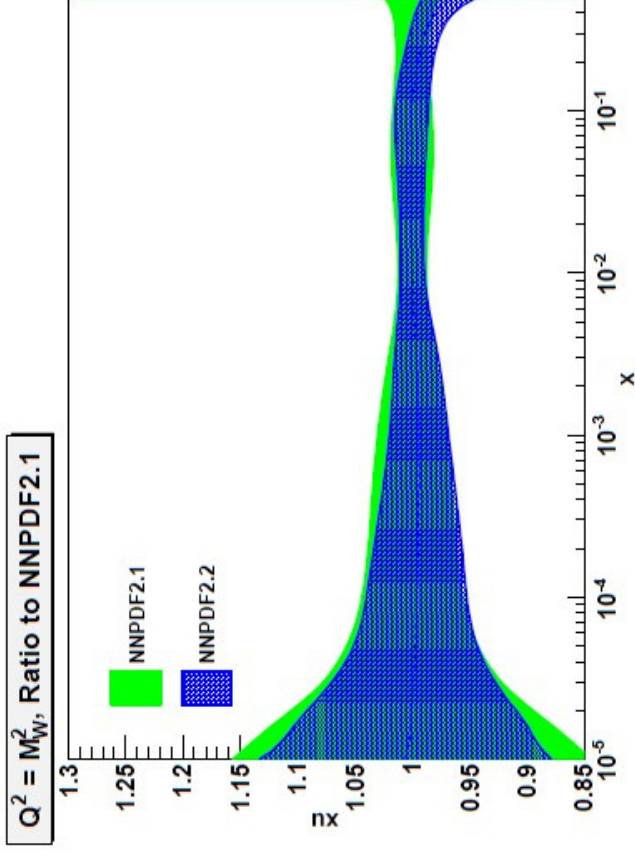
Improvement seen for all flavours and antiflavours

NNPDF2.2

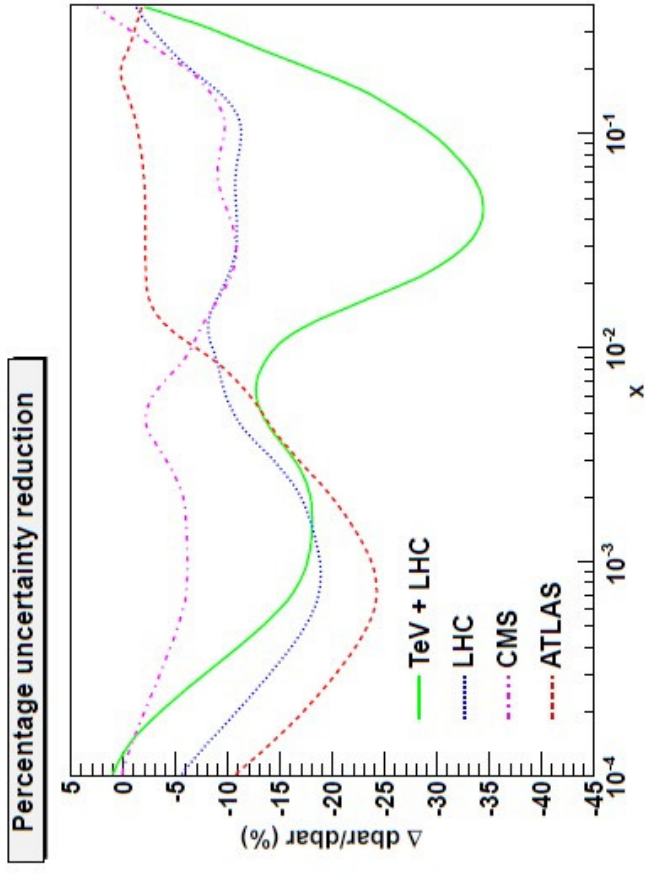
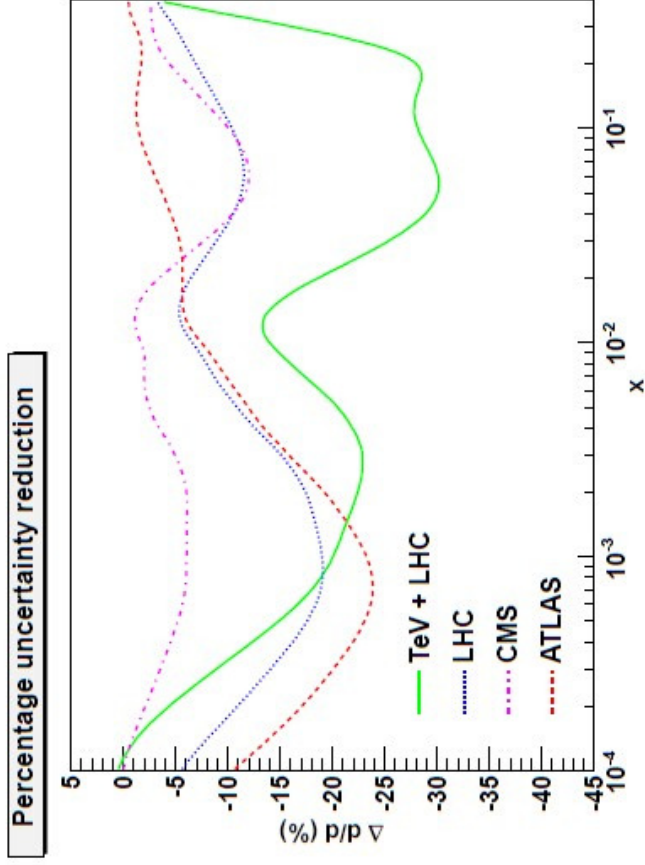
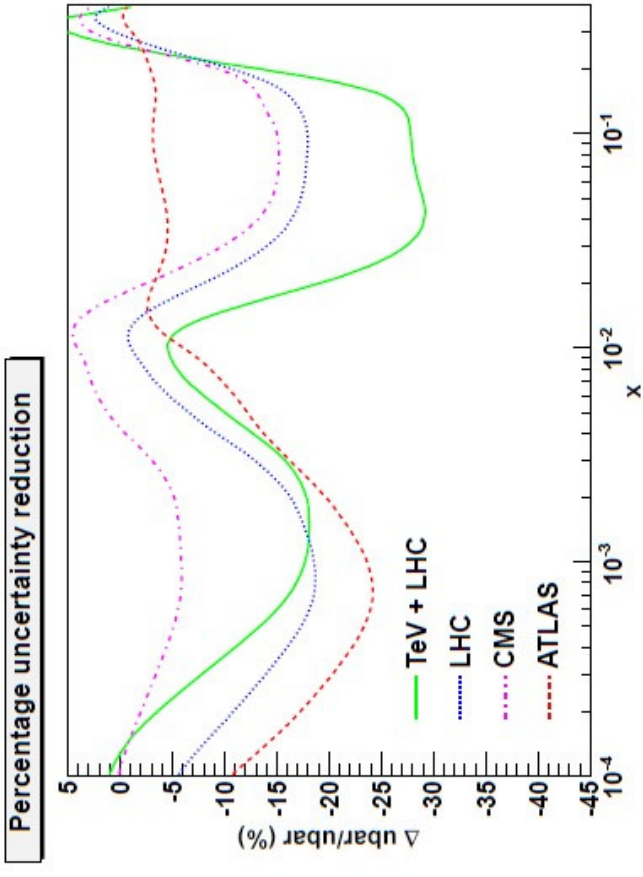
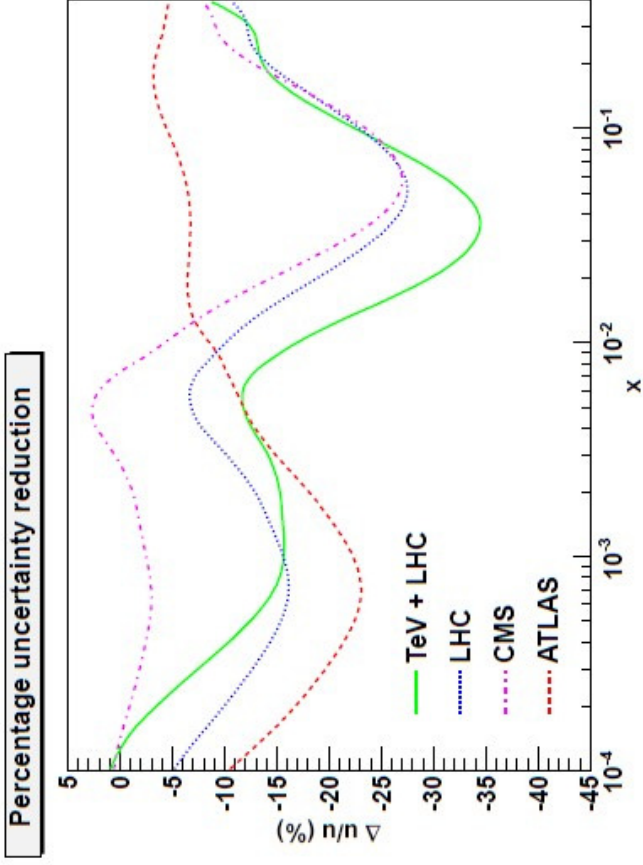
NNPDF2.1 + (Atlas+CMS+D0 W lepton asy)

Experiment	N_{dat}	NNPDF2.1	NNPDF2.1 LHC	NNPDF2.2
NMC-pd	132	0.97	0.95	0.97
NMC	221	1.73	1.72	1.72
SLAC	74	1.33	1.26	1.28
BCDMS	581	1.24	1.23	1.23
HERAI-AV	592	1.07	1.07	1.07
CHORUS	862	1.15	1.15	1.15
FLH108	8	1.37	1.37	1.37
NTVDMN	79	0.79	0.74	0.70
ZEUS-H2	127	1.29	1.28	1.28
ZEUSF2C	50	0.78	0.79	0.78
H1F2C	38	1.51	1.52	1.51
DYE605	119	0.84	0.84	0.86
DYE886	199	1.25	1.23	1.27
CDFWASY	13	1.85	1.81	1.81
CDF'ZRAP	29	1.66	1.61	1.70
D0ZRAP	28	0.60	0.60	0.58
CDFR2KT	76	0.98	0.98	0.96
D0R2CON	110	0.84	0.84	0.83
ATLASmuASY	11	[0.77]	0.97	1.07
CMSseASY	6	[1.83]	1.23	1.08
CMSmuASY	6	[1.24]	0.63	0.56
D0eASY	12	[4.39]	[3.46]	1.38
D0muASY	10	[1.48]	[1.17]	0.35
Total		1.165	1.158	1.157

NNPDF2.2: ups and downs



Percentage uncertainty reduction





Summary & Outlook

- NNPDF 2.1 LO, NLO, NNLO
- See for yourself: <http://projects.hepforge.org/lhapdf>
- Reweighting:
 - You can update NNPDFs yourself: new tool
- NNPDF2.2:
 - First Global PDF fit with LHC data
- For the future:
 - Lots and lots of new LHC data!