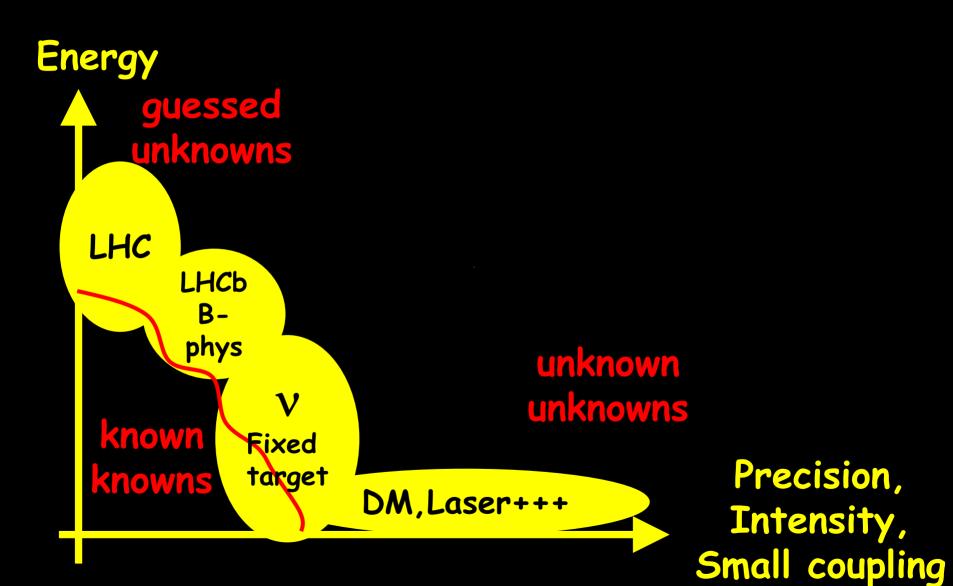


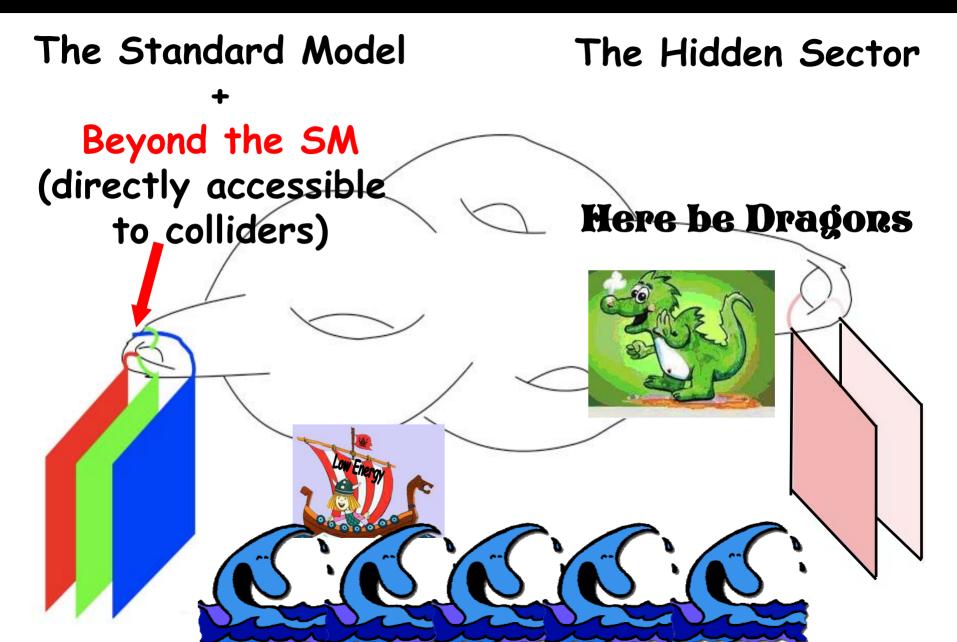
Exploring...





Where we want to go...





Hints for new Physics

Uglyness of old models

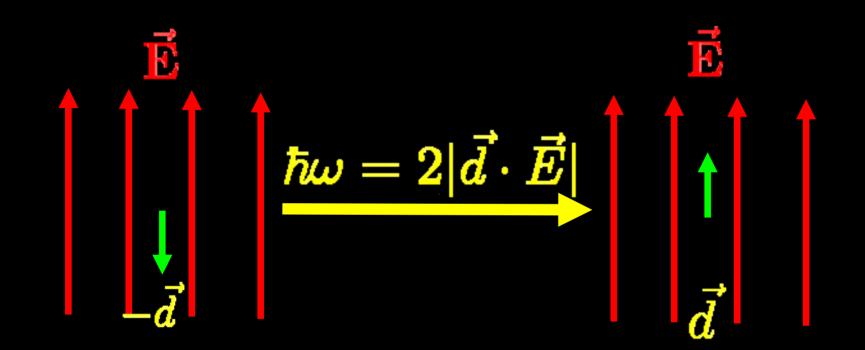


- The Standard Model has many free parameters: O(30)
- Naturalness problems. Finetuning.
 Examples: Higgs mass, θ-angle (strong CP-problem)

Neutron electric dipole moment



θ would cause neutron EDM Experiment



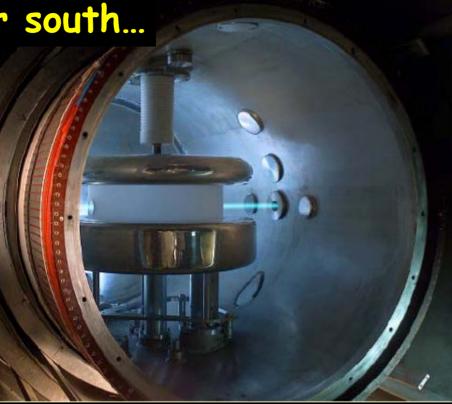


No neutron electric dipole moment



Somewhere far far south...





 $|\vec{d}| < 3 \, 10^{-26} e \, cm$ $= 3 \, 10^{-13} e \, fm$

No neutron electric dipole moment...





$\begin{aligned} |\vec{d}| &< 3\,10^{-26} e\,cm \\ &= 3\,10^{-13} e\,fm \,\, \mbox{\em set} \,\, \frac{\theta}{16\pi^2}\,e\,fm \end{aligned}$



Uglyness of old models



- The Standard Model has many free parameters: O(30)
- Naturalness problems. Finetuning.
 Examples: Higgs mass, θ-angle (strong CP-problem)
- Gravity separate, i.e. not unified.
- (Probably) Breaks down at a finite energy scale
 - Landau poles etc.

Unexplained Stuff

- University of Durham
- Dark Matter (25%)
 (astrophysical + cosmological observations)
- Dark Energy (70%) (astrophysical + cosmological observations)
- Mass Hierarchies (colliders, neutrino exp, etc)
- Small parameters (θ-angle, again) (neutron electric dipole measurements)





- $(g-2)_{\mu}$ deviations from SM prediction
- DAMA anomaly
- · CoGeNT etc.
- PAMELA observation
- WMAP observes extra "neutrinos"
- Proton radius in muonic hydrogen

Hints for new Physics Model Building Top-down Bottom-up (theory) (pheno)

Fix problem `here and now'

Go back to drawing board `Start from scratch'

The strong CP problem: Axions

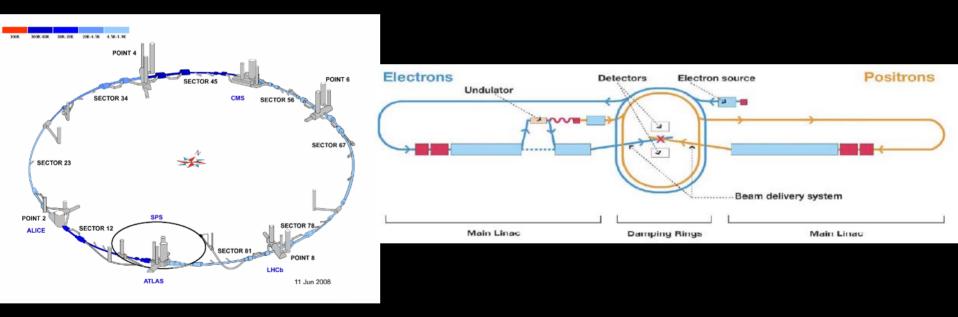


- Introduce new Peccei-Quinn symmetry to solve naturalness problem
- Predict as a consequence a new particle: The Axion (it's a Weakly Interacting Sub-eV Particle) Dark matter candidate Good motivation for axion/WISP experiments

Hints for new Physics Model Building Bottom-up Top-down (theory) (pheno)

Experiments

Exploring fundamental high energy physics... The direct approach: MORE POWER LHC, Tevatron + ILC, CLIC



- Detects most things within energy range
- E.g. may find SUSY particles, WIMPs etc.





- May miss very weakly interacting matter (Axions, WIMPs, WISPs...)
- Current maximal energy few TeV

• Man its DANGEROUS...

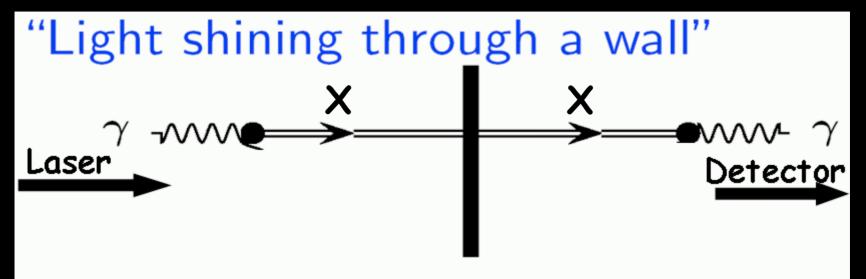
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Recycling... Complementary approaches

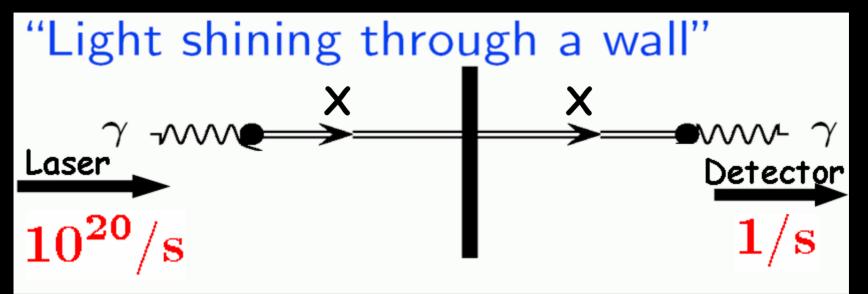
Light shining through walls





Light shining through walls





\cdot Test $P_{\gamma \to X \to \gamma} \lesssim 10^{-20}$

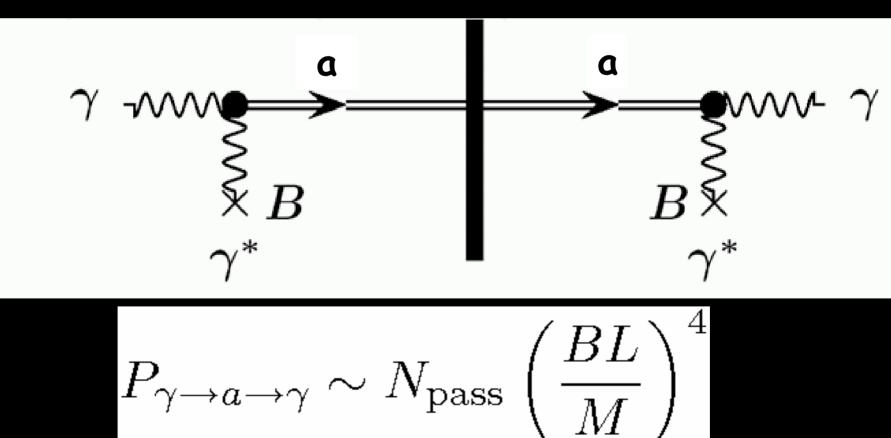
- Enormous precision!
- Study extremely weak couplings!

Photons coming through the wall!



- It could be Axion(-like particle)s!
- Coupling to two photons:

$$\frac{1}{M}a\tilde{F}F\sim\frac{1}{M}a\vec{\mathbf{E}}\cdot\vec{\mathbf{B}}$$



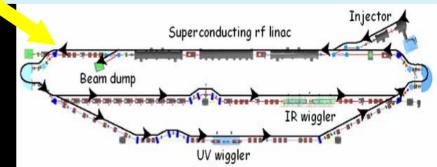
Light Shining Through Walls





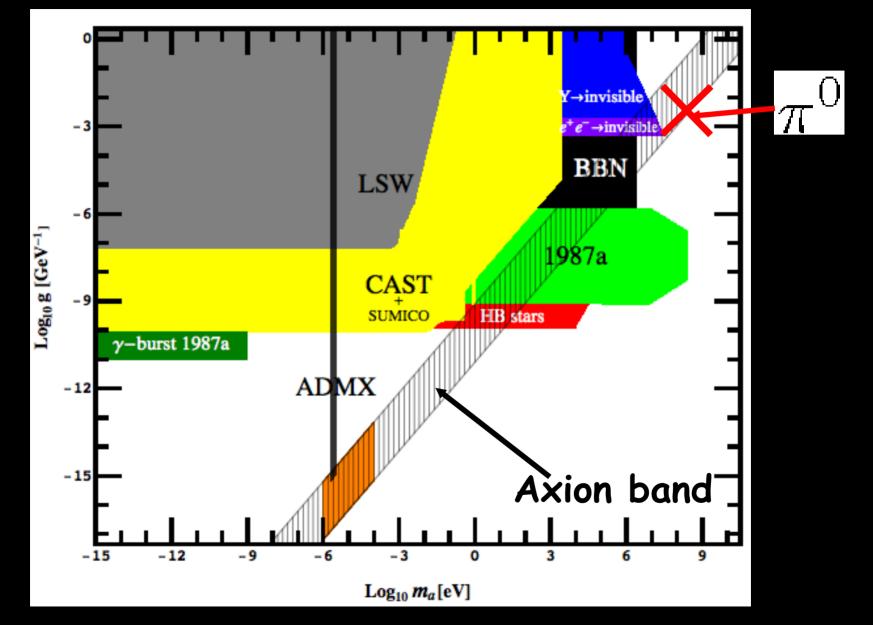
- ALPS
- BMV
- Gamme V 25 cm
- LIPPS
- OSQAR

	Calibration dio	
Laser Box	Tevatron magnet (6m) Plunger	PMT Box
aser n-		
	k	PMT
	(2m)	
/	Warm bore "wali"	
Monitor sensor		



Small coupling, small mass





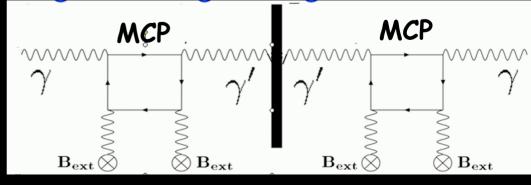
WISPS=Weakly interacting sub-eV particles



• Axions $\gamma \rightarrow \gamma \rightarrow \varphi$ $\chi = B \times B \times B \times Y^*$ γ^*

 Massive hidden photons (without B-field)
 =analog v-oscillations

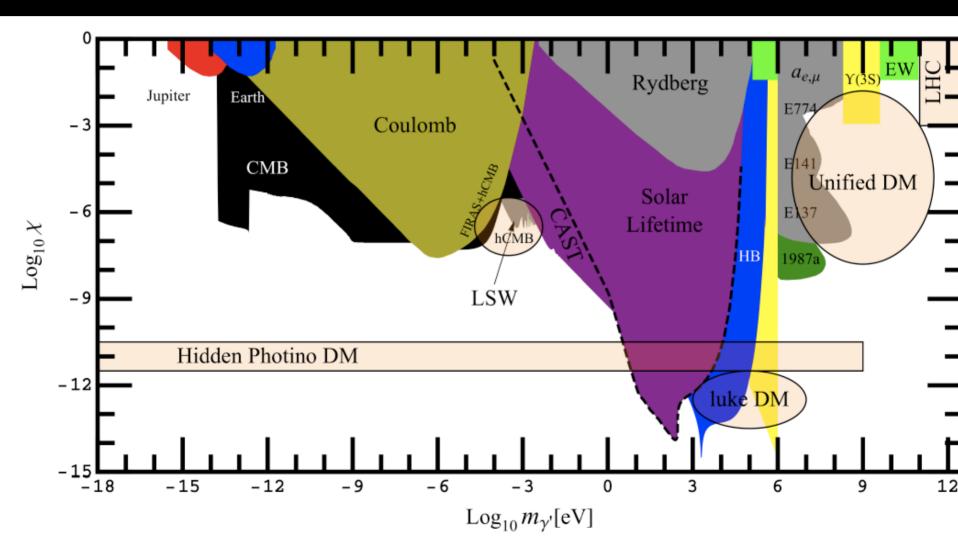
 Hidden photon + minicharged particle (MCP)



Hidden Photons



LSW already competitive + testing interesting area



Hints for new Physics Model Building

Hope for light particles?

YES, we can!

Hope for light particles?

YES, or you'll be shot!





Neutrino masses:

 $m_
u \sim {
m meV}$

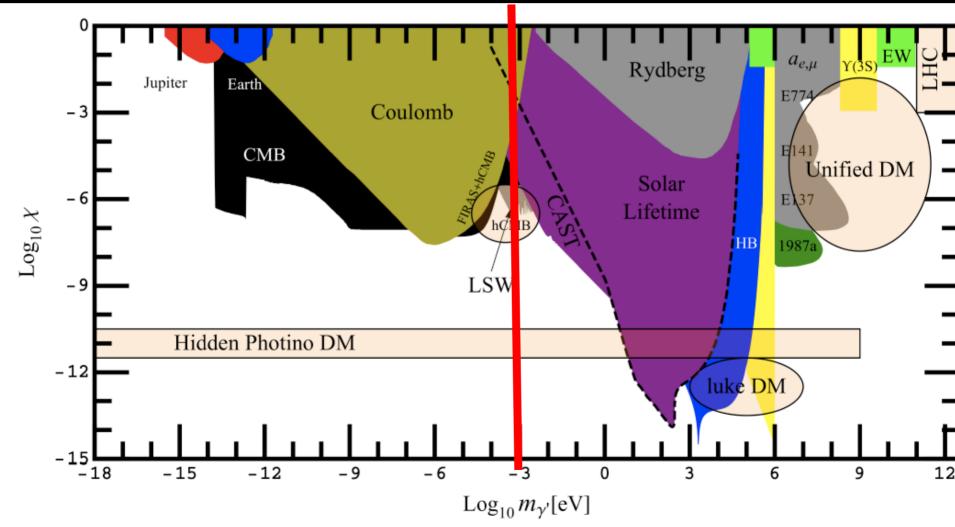
- Scale of dark energy: $\rho_{\Lambda} \sim (\mathrm{meV})^4$
- Energy density of the Universe:

 $ho_{
m today} \sim ({
m meV})^4$

Hidden Photons



LSW already competitive + testing interesting area Dark energy scale



High Scale Couping Small

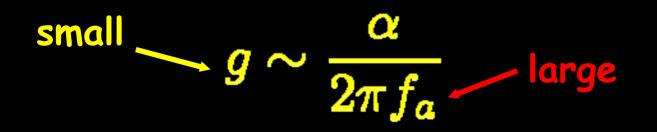




Effective higher dimensional coupling

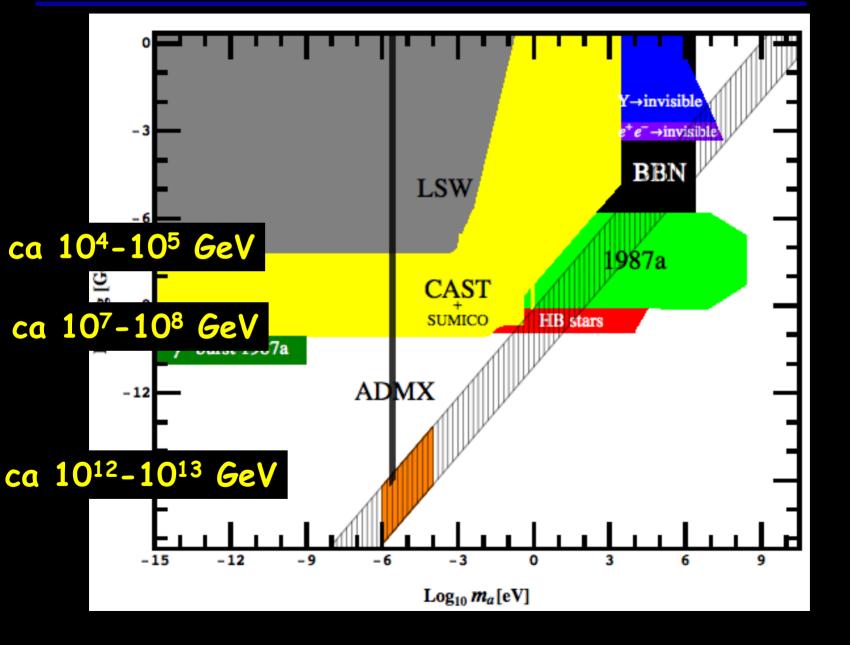
$\mathcal{L}_{Int} = -rac{1}{4}gaF^{\mu u}\widetilde{F}_{\mu u} = -ga\mathbf{E}\cdot\mathbf{B}$

• Small coupling for large axion scale:



Huge Scale >> LHC Energy!

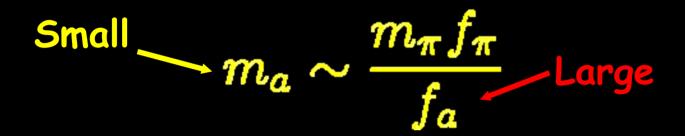




High Scale Small Mass



• The axion mass is small, too!





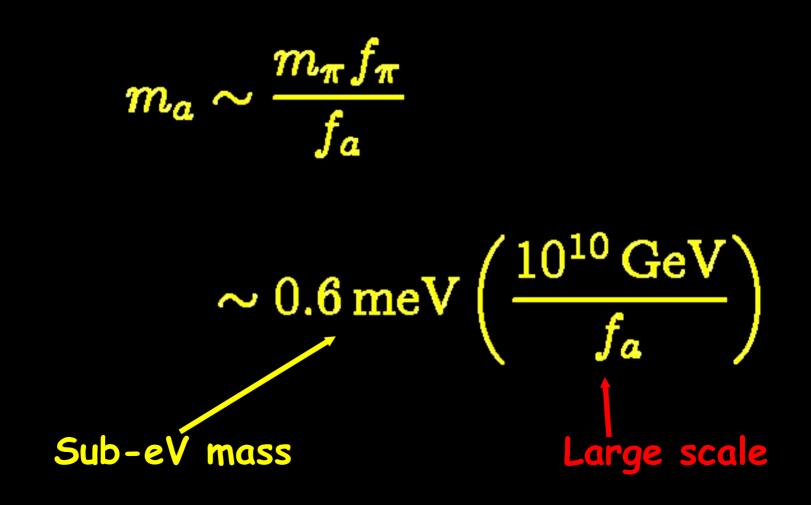
• The axion mass is small, too!



Pseudo-Goldstone Boson!

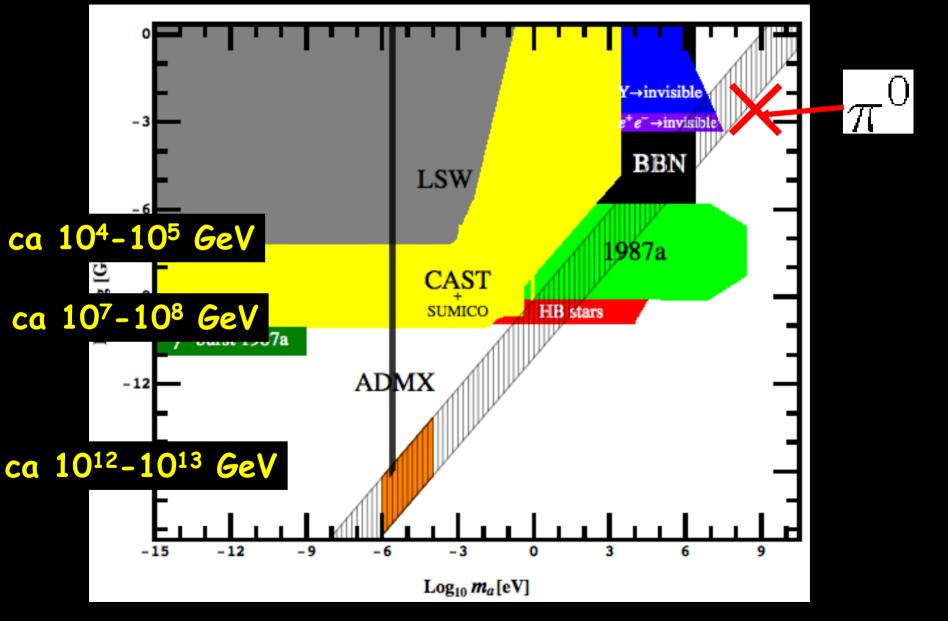


• The axion mass is small, too!



Large Scale but light!





Hints for new Physics Model Building Bottom-up Top-down (theory) (pheno)

Go back to drawing board `Start from scratch' WISPs from String Theory

String theory



- Attempt to unify SM with gravity
- New concept: strings instead of point particles

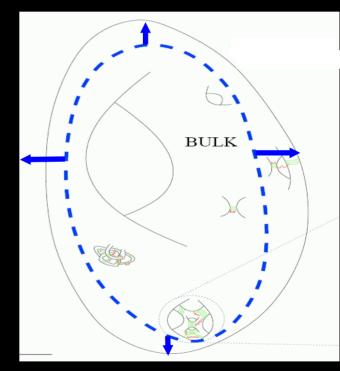
Axion(-like particles)

String theory: Moduli and Axions

String theory needs Extra Dimensions

Must compactify

 Shape and size deformations correspond to fields: Moduli (WISPs) and Axions Connected to the fundamental scale, here string scale

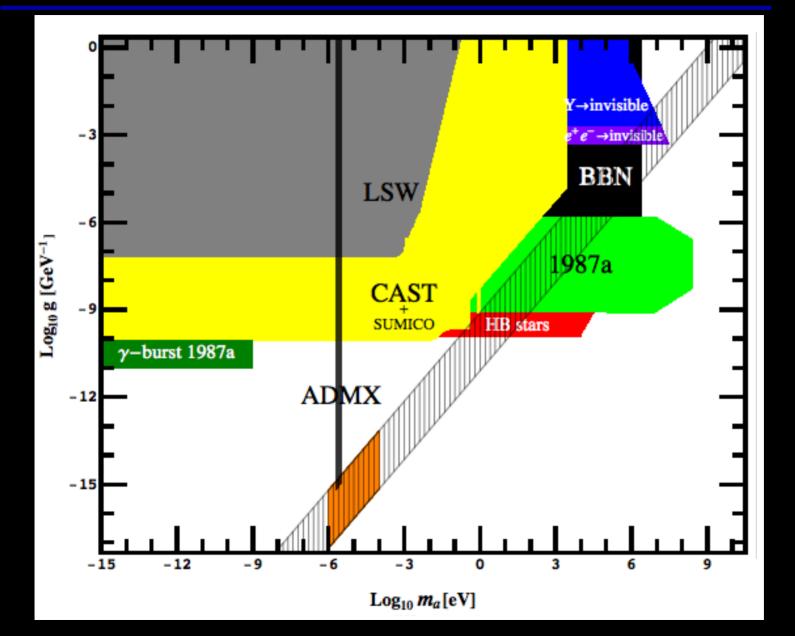






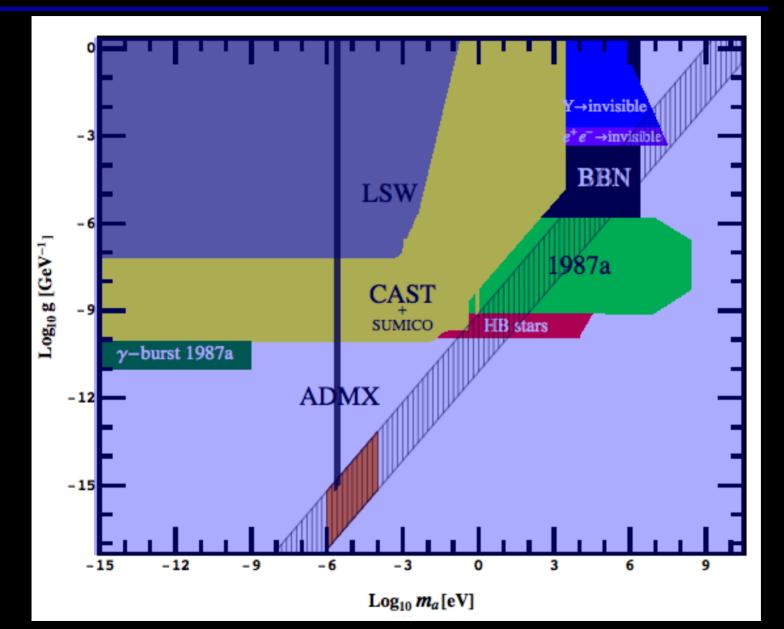
Axion (like particles): Where are we?





Axion (like particles): Where are we?

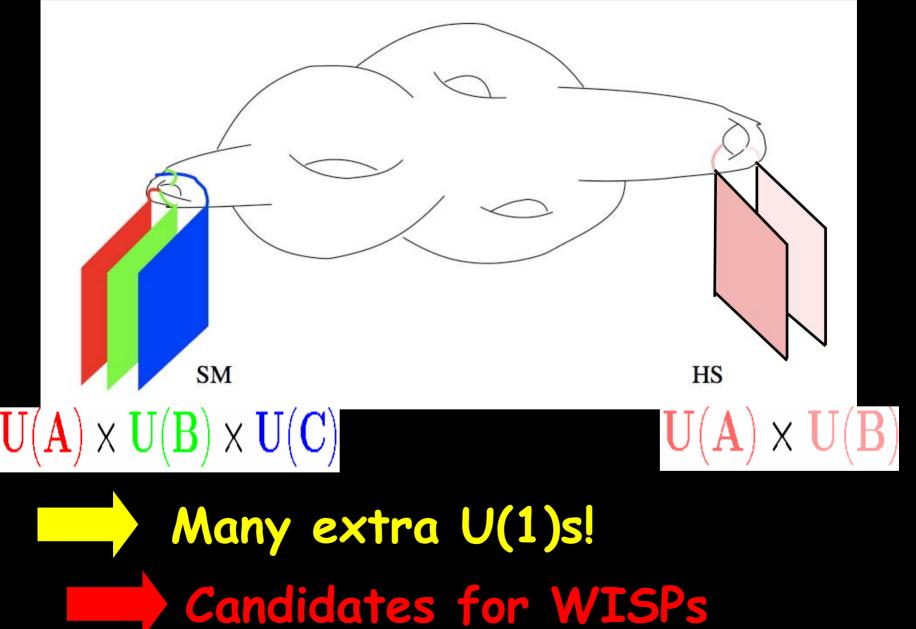




Hidden Photons

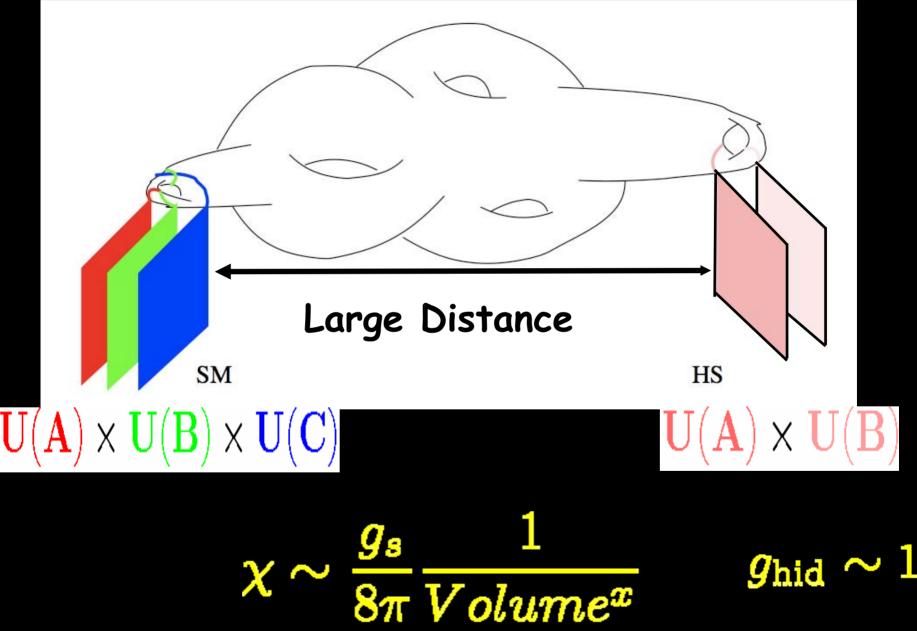
String theory likes extra gauge groups





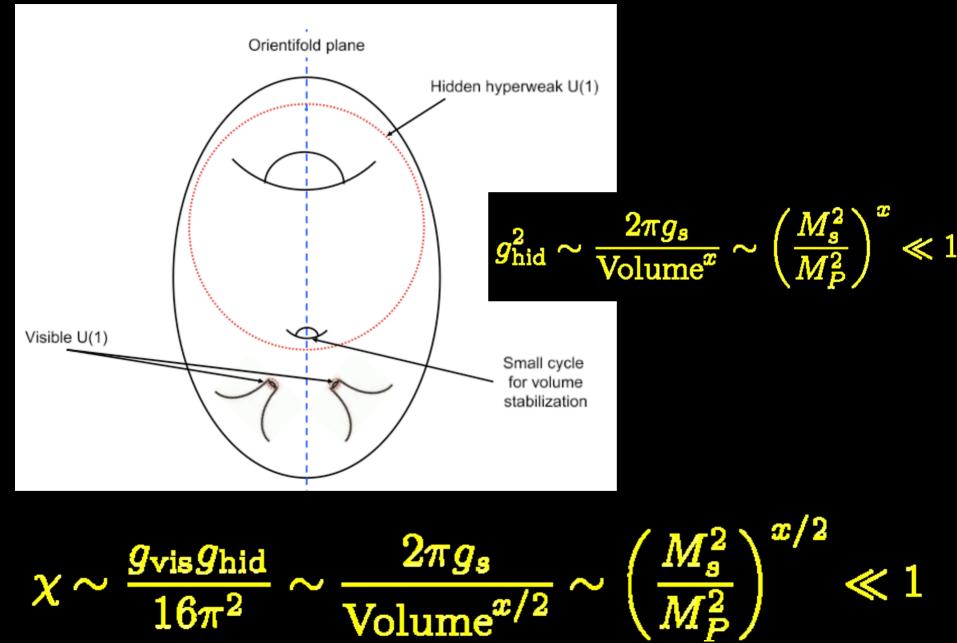
Hidden by distance





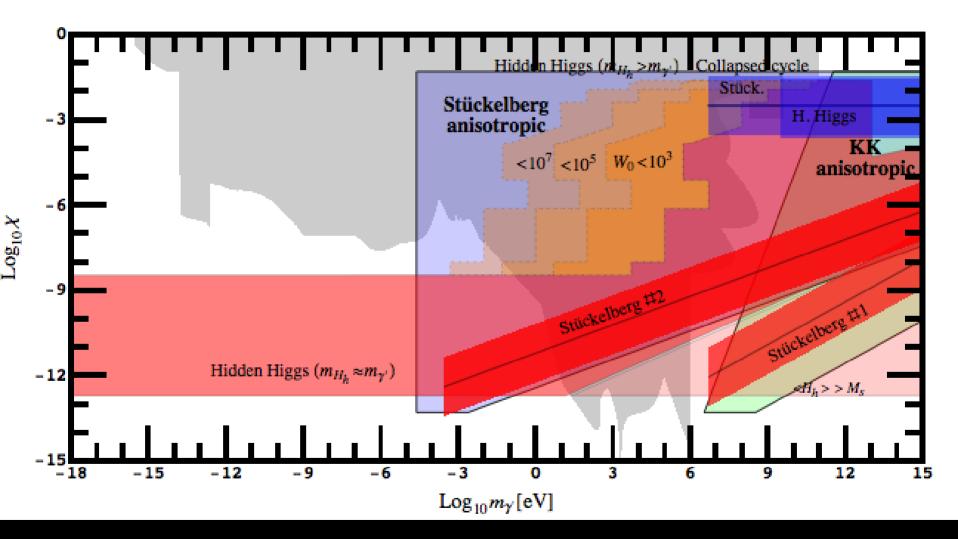
Hidden by weakness





Hidden Photons, all over the place

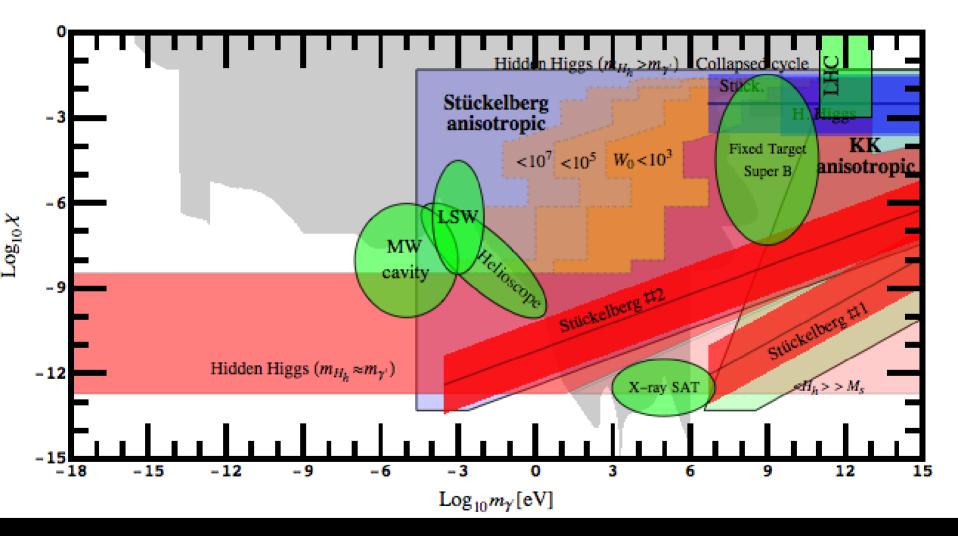




Hints for new Physics Model Building Bottom-up (pheno) Top-down (theory) New, cool Experiments

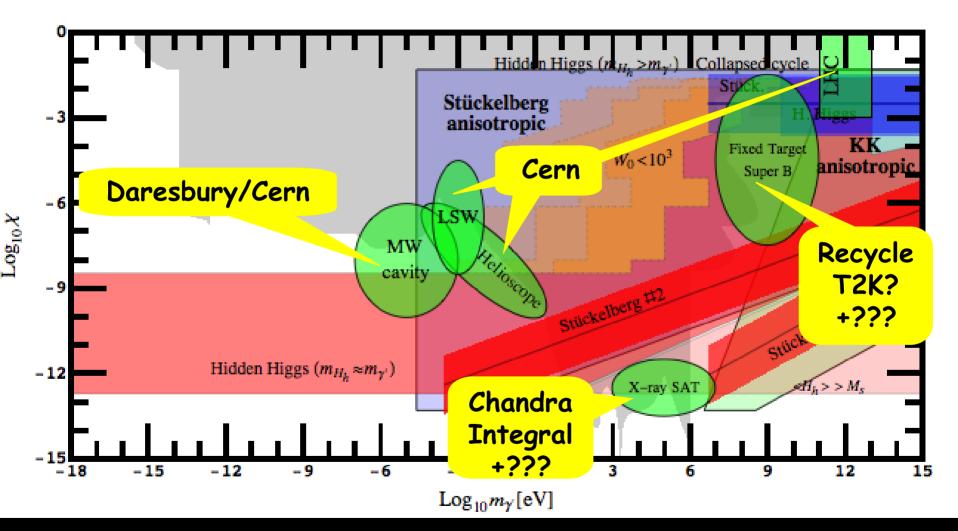
Hidden Photons: Back to Experiment





Hidden Photons: Back to Experiment



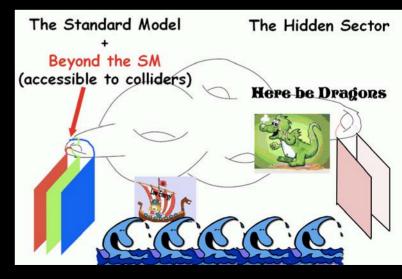


Conclusions

Conclusions



- Good Physics Case for NEW STUFF @ Low Energies
 explore `The Low Energy Frontier' High Precision
- Low energy experiments test energy scales much higher than accelerators
 - **Complementary!**
- May provide information on hidden sectors and thereby into the underlying fundamental theory



Discover the Hidden Islands

THE R. O'Monday's read of "Manufactor Distance" States of Long.