

# Other Opportunities in Neutrino Experiments



- Neutrino oscillations lead to rich phenomenology
- Global  $\nu$  programme has a wide range of future options
- UK physicists are working on many of these ideas
  - In some cases, leadership positions already
  - In others, will be in strong position to assume leadership roles when the time comes
- STFC strategy will group small experiments into broad science areas
  - this talk is a few details on one: **neutrinos**

# Outline

- Reminder of current picture
- Sterile neutrinos
  - Will focus on experiments that are important to the 3- $\nu$  paradigm or otherwise add value to global programme
- Reactor neutrinos ( $\theta_{12}$ )
- UHE neutrinos from astrophysical sources

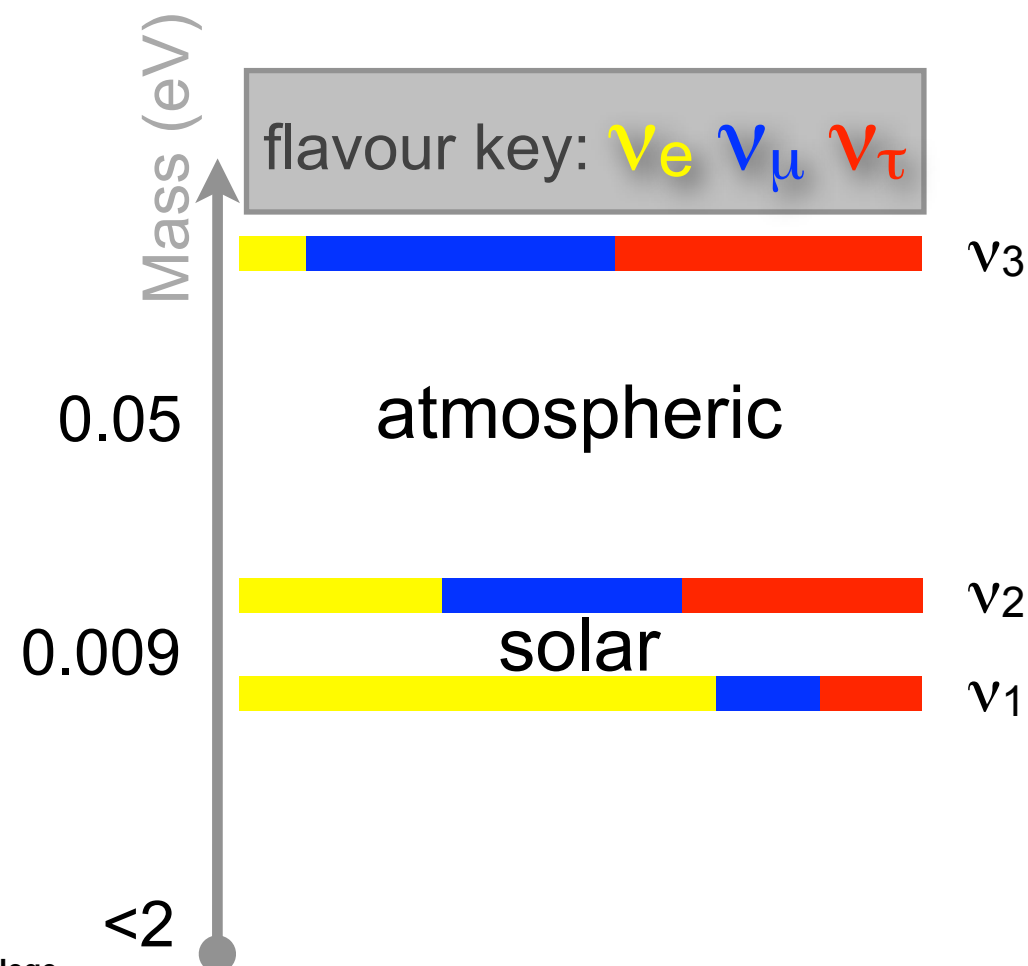


# Reminder: neutrino picture

$$\begin{array}{c} \text{flavour} \end{array} \quad \begin{array}{c} \text{atmospheric} \end{array} \quad \begin{array}{c} \text{accelerator} \end{array} \quad \begin{array}{c} \text{solar} \end{array} \quad \begin{array}{c} \text{Majorana} \end{array} \quad \begin{array}{c} \text{mass} \end{array}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\frac{\alpha_1}{2}} & 0 & 0 \\ 0 & e^{i\frac{\alpha_2}{2}} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

*Neutrino physics is making discoveries!  
Big results in past year - more discoveries on horizon.*



## OPEN QUESTIONS:

- Hierarchy?
- CP violation?
- Majorana or Dirac?
- Absolute mass scale?
- A nice picture with a clear path forward...

# Sterile Neutrinos?



# Sterile Neutrinos?

## Gallium Anomaly: $\nu_e$ Disappearance?

- SAGE and GALLEX gallium solar neutrino experiments used  $\text{MCI } ^{51}\text{Cr}$  and  $^{37}\text{Ar}$  sources to calibrate their detectors

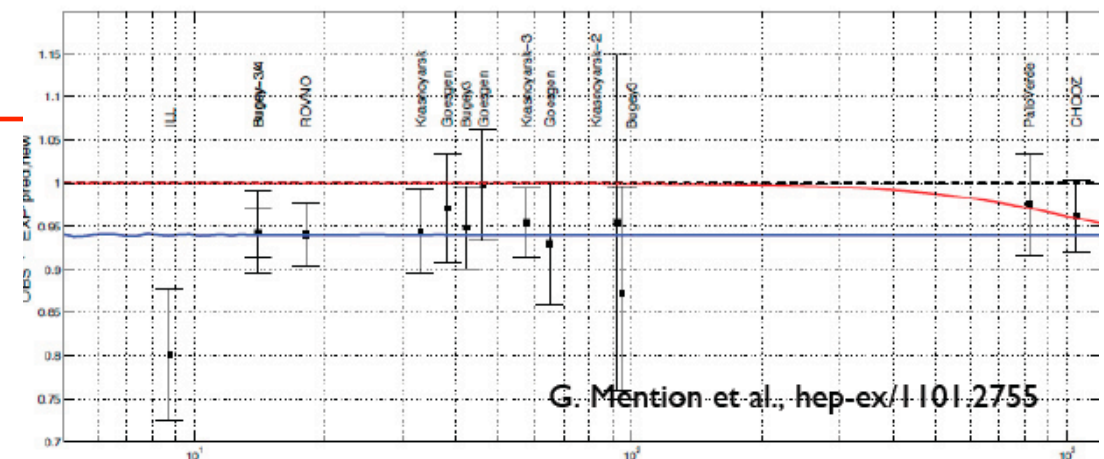
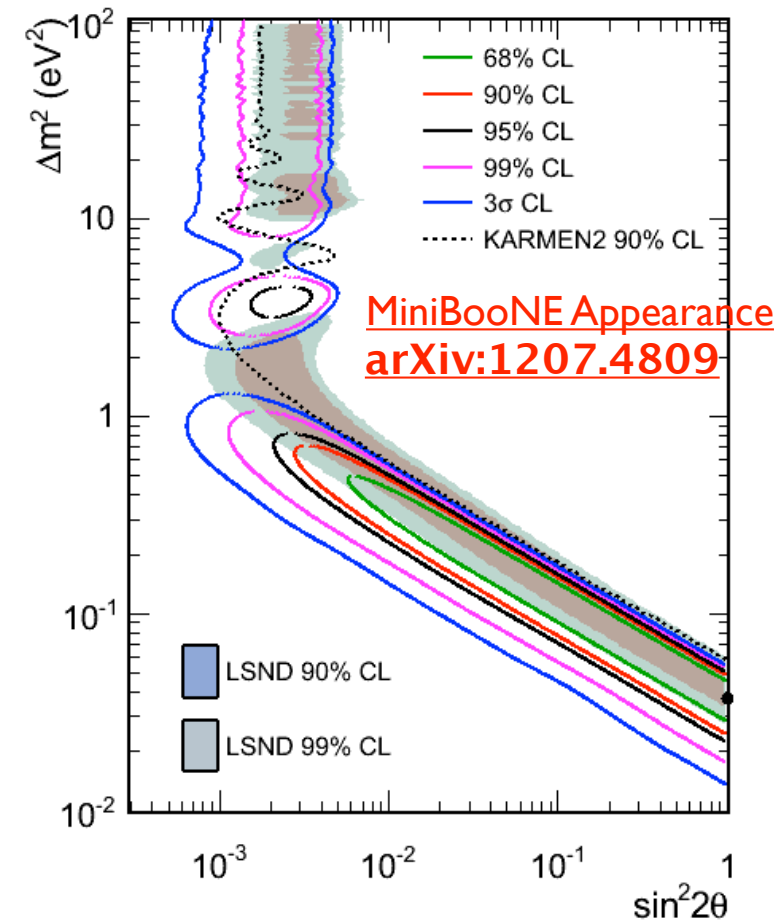
- A recent analysis claims a significant ( $3\sigma$ ) deficit (Giunti and Laveder, 1006.3244v3 [hep-ph])
  - Ratio (observation/prediction) =  $0.76 \pm 0.09$
  - An oscillation interpretation gives  $\sin^2 2\theta > 0.07, \Delta m^2 > 0.35 \text{eV}^2$

## Reactor Antineutrino Anomaly

Re-analysis of predicted reactor fluxes based on a new approach for the conversion of the measured electron spectra to anti-neutrino spectra.

- Reactor flux prediction increases by 3%.
- Re-analysis of reactor experiments show a deficit of electron anti-neutrinos compared to this prediction – at the  $2.14\sigma$  level
- Could be oscillations to sterile with  $\Delta m^2 \sim 1 \text{eV}^2$  and  $\sin^2 2\theta \sim 0.1$

Red: Oscillations assuming 3 neutrino mixing  
Blue: Using a 3+1 (sterile neutrino) model



N.B.: several 2-3  $\sigma$  results don't constitute compelling evidence...



# Active-sterile Neutrino Oscillation

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_{s1} \\ \dots \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} & \dots \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} & \dots \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} & \dots \\ U_{s11} & U_{s12} & U_{s13} & U_{s14} & \dots \\ \dots & \dots & \dots & \dots & \dots \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \\ \dots \end{pmatrix}$$

$\nu_\mu \rightarrow \nu_e$  appearance

$$P(\nu_\mu \rightarrow \nu_e) = 4|U_{e4}|^2|U_{\mu4}|^2 \sin^2 \left[ 1.27 \Delta m_{41}^2 \frac{L}{E} \right]$$

$\nu_e$  disappearance

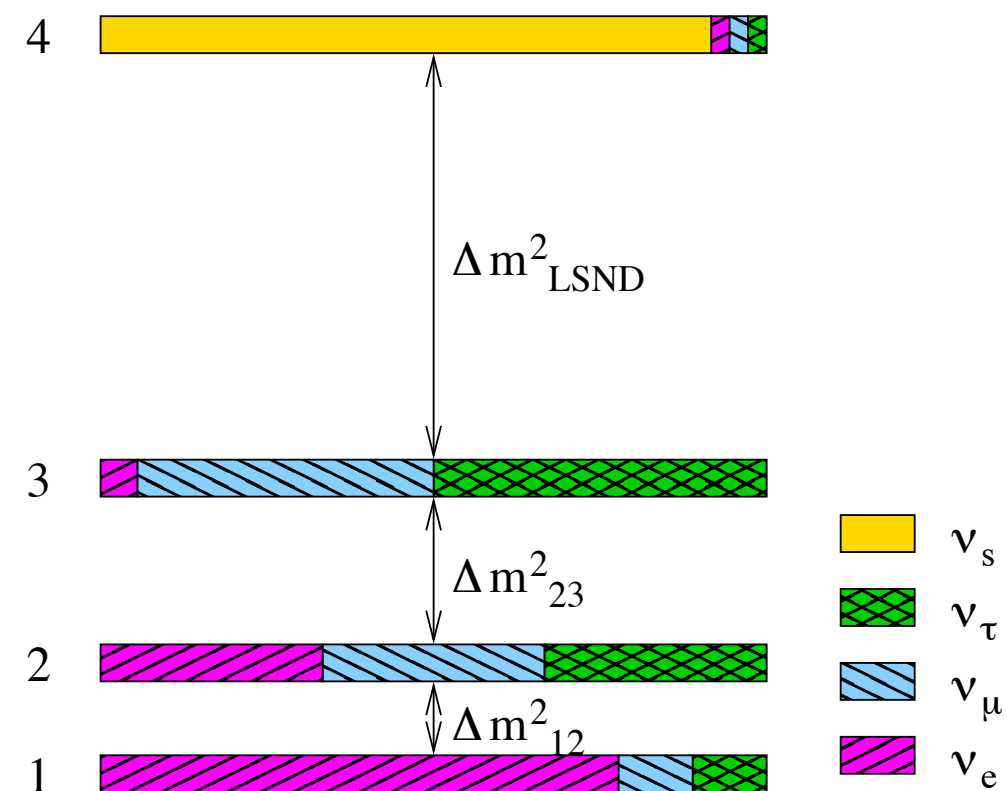
$$P(\nu_e \rightarrow \nu_x) = 1 - 4|U_{e4}|^2(1 - |U_{e4}|^2) \sin^2 \left[ 1.27 \Delta m_{41}^2 \frac{L}{E} \right]$$

$\nu_\mu$  disappearance

$$P(\nu_\mu \rightarrow \nu_x) = 1 - 4|U_{\mu4}|^2(1 - |U_{\mu4}|^2) \sin^2 \left[ 1.27 \Delta m_{41}^2 \frac{L}{E} \right]$$

Close relationships between appearance and disappearance channels

3+1 sterile neutrino scheme

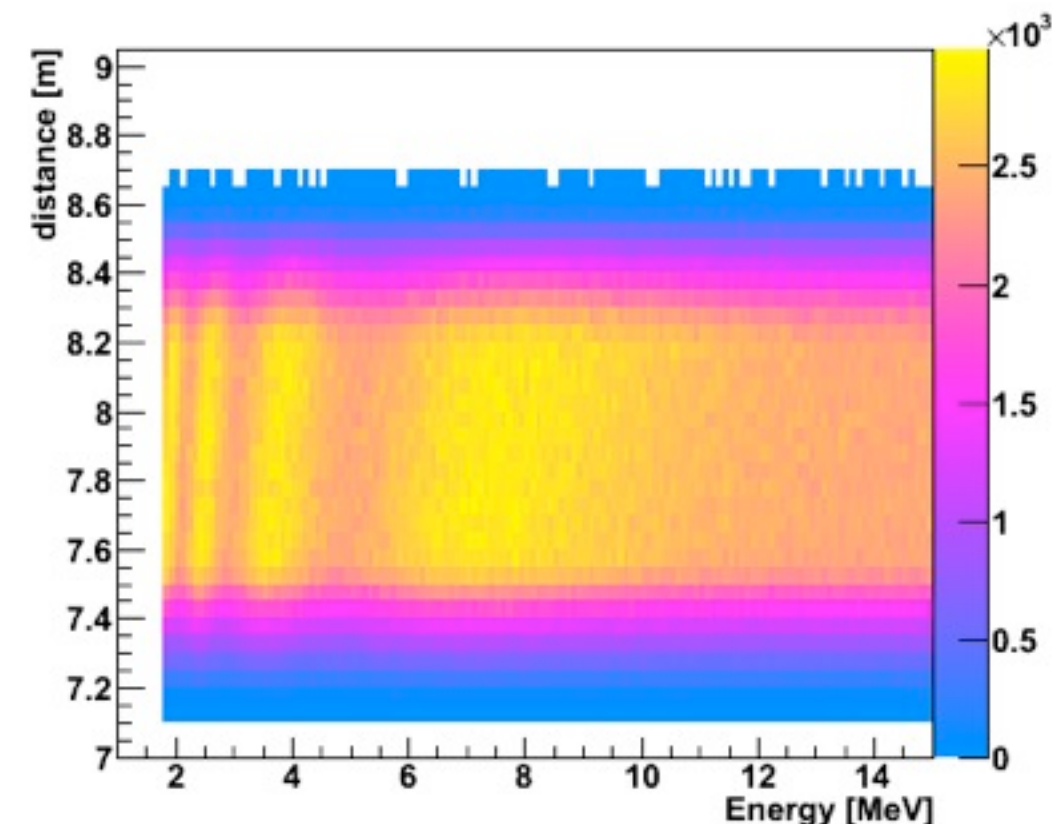
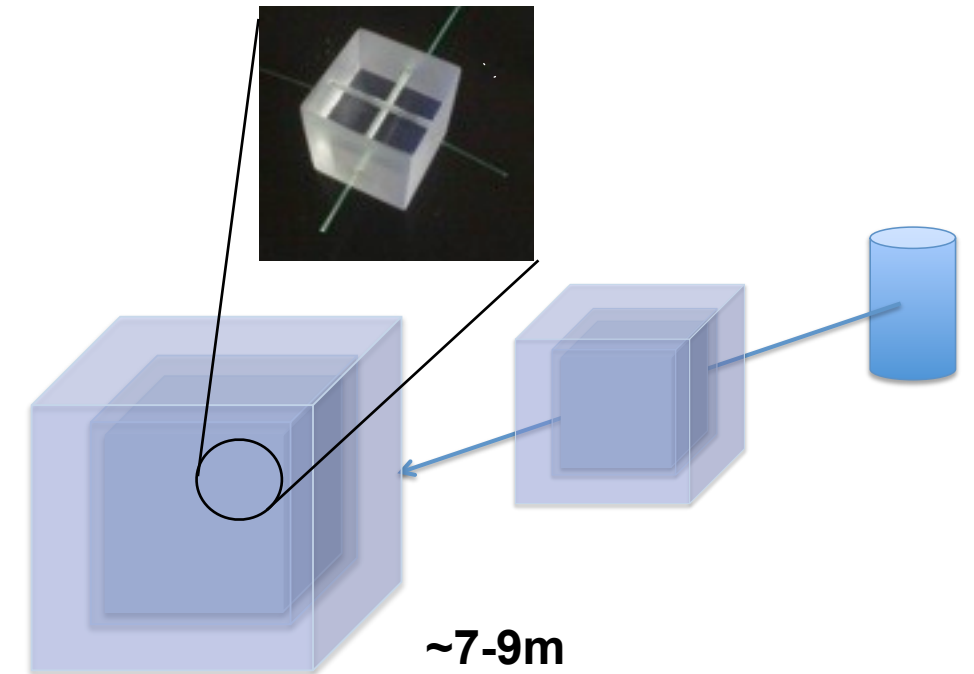


To test reactor  
flux and Ga  
anomalies

# TWIX

*solid segmented plastic scintillator detectors*

- Novel approach to detect antineutrinos at reactors
- composite scintillator cells with  $\text{Li}^6$
- compact system with minimal shielding (1.5m footprint for 1T Fiducial mass)
- very low sensitivity to gamma background
- can achieve better signal to background ratio than traditional liquid scintillator system
- Originally developed for reactor monitoring purposes



$$\Delta m^2 = 2.35, \sin^2 2\theta_{ee} = 0.165$$

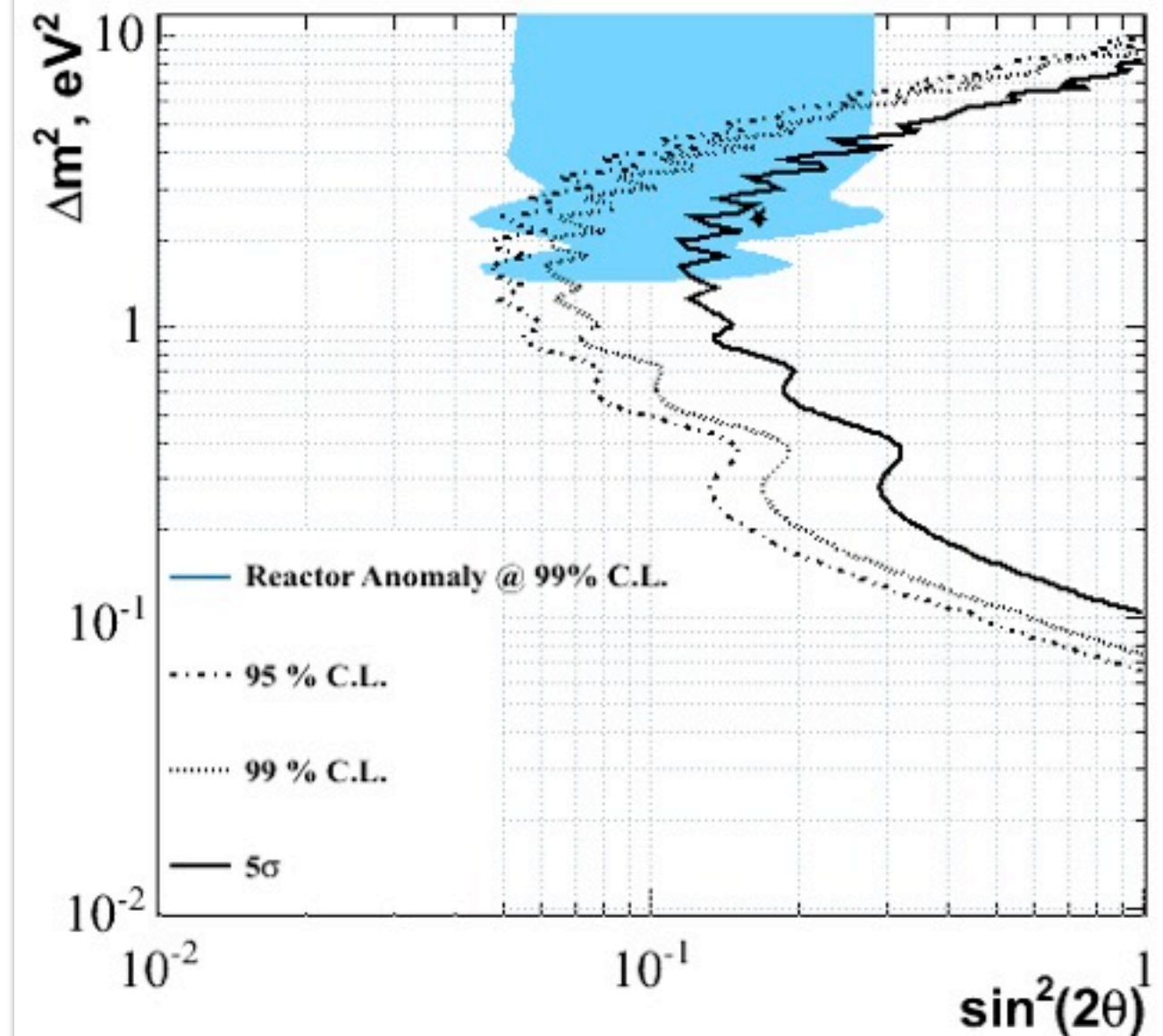


To test reactor  
flux and Ga  
anomalies

# TWIX

*solid segmented plastic scintillator detectors*

- Measurement at ILL (2 years)  
(~50k events)
- Baselines assumed: 7.5 m near  
and 9 m far (being optimised)
- (ILL 0.8m x 0.4m core can  
provide best resolution on SBL  
oscillations)
- shape analysis using two detector  
baseline
  - signal from ratio of spectra
  - 3D vertex reconstruction (< 10  
cm resolution)
  - $\sigma_E/E \sim 0.1$  MeV



# DAEδALUS

[arXiv:1006.0260 \[physics.ins-det\]](https://arxiv.org/abs/1006.0260)

osc max ( $\pi/2$ )  
at 40 MeV

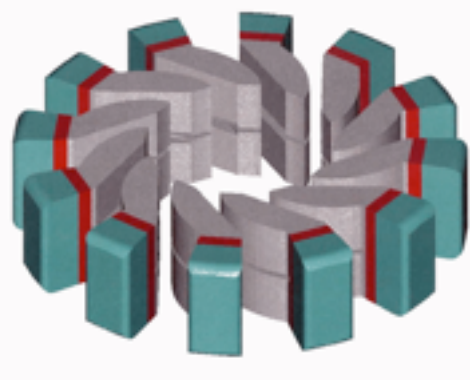
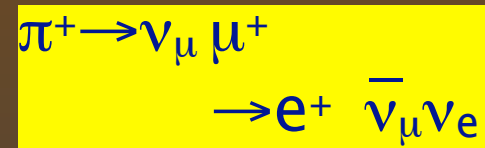
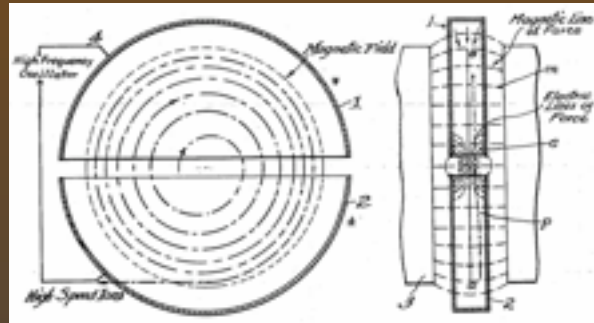
off max ( $\pi/4$ )  
at 40 MeV

Constrains  
flux

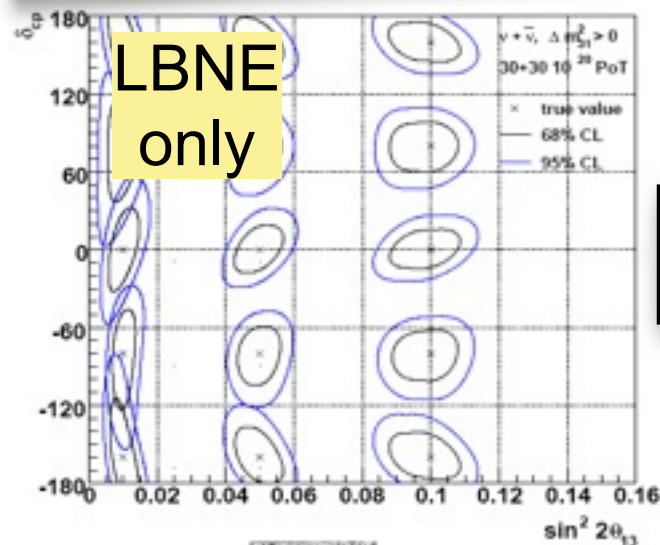
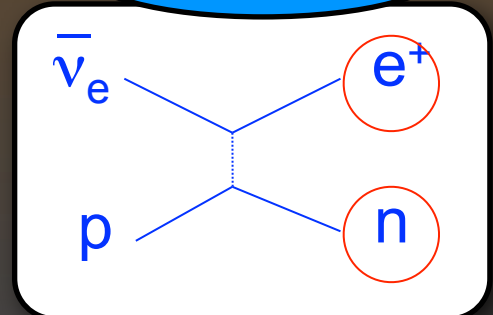
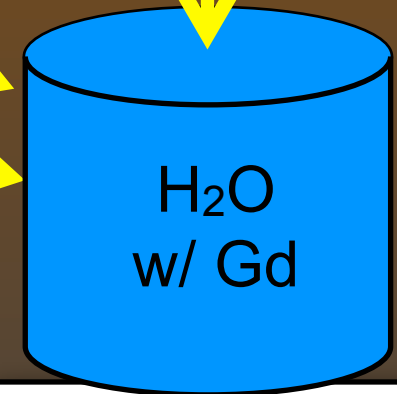
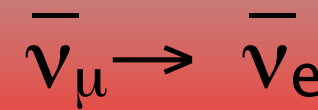
20km

8km

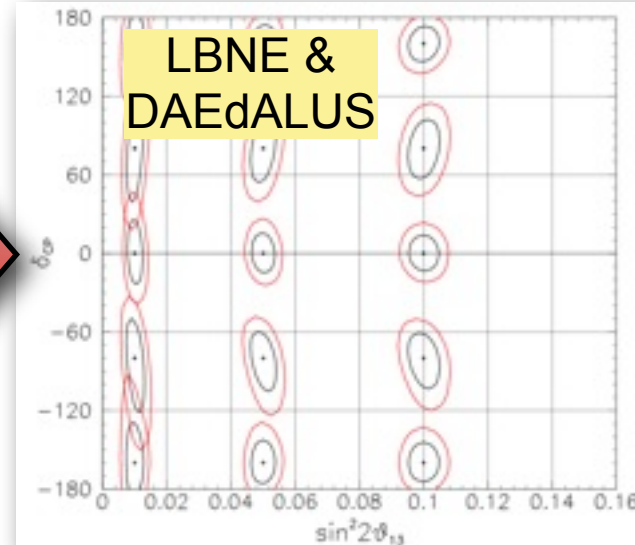
1.5km



High power cyclotrons  
create massive  $\bar{\nu}_\mu$  flux  
at multiple baselines



DAEdALUS



Physics studies done  
assuming H<sub>2</sub>O detector in  
LBNE, but same  
performance achievable  
with Hyper-K or LBNO

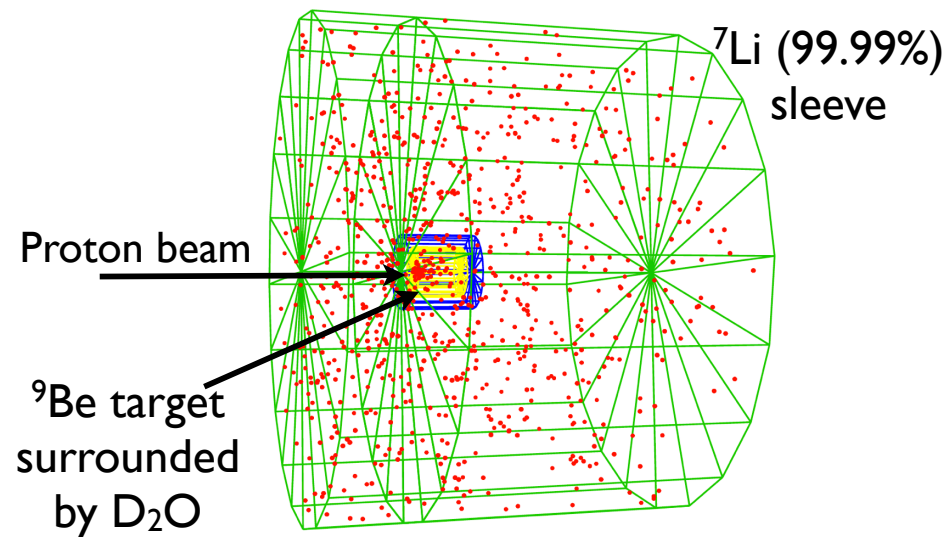
To test reactor  
flux and Ga  
anomalies

DAEδALUS

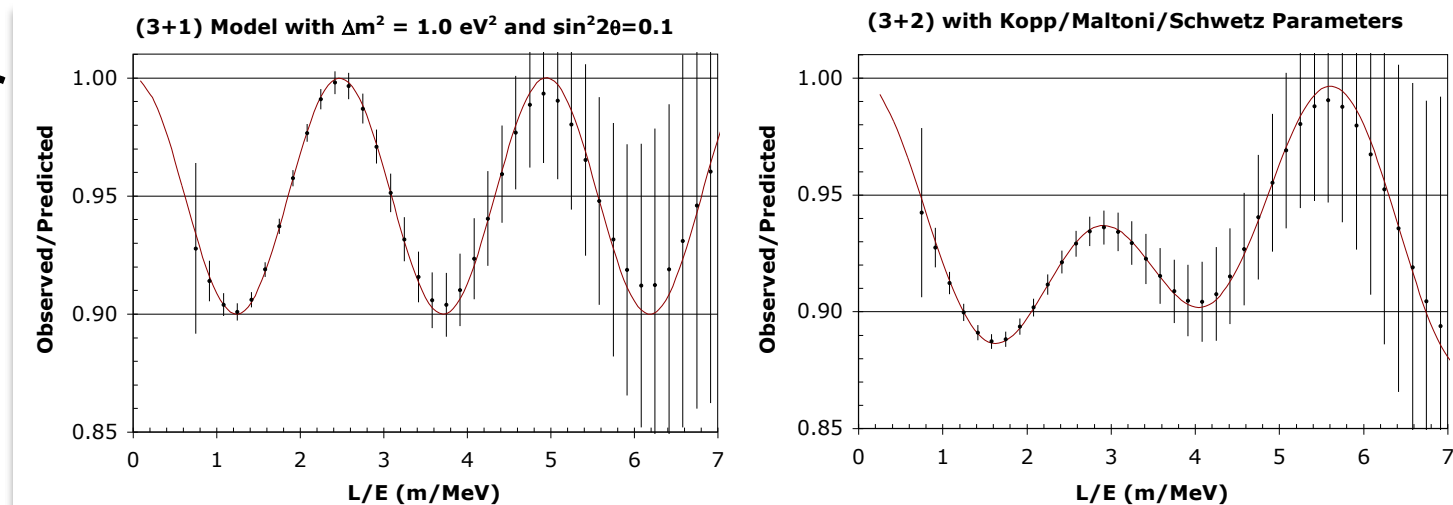
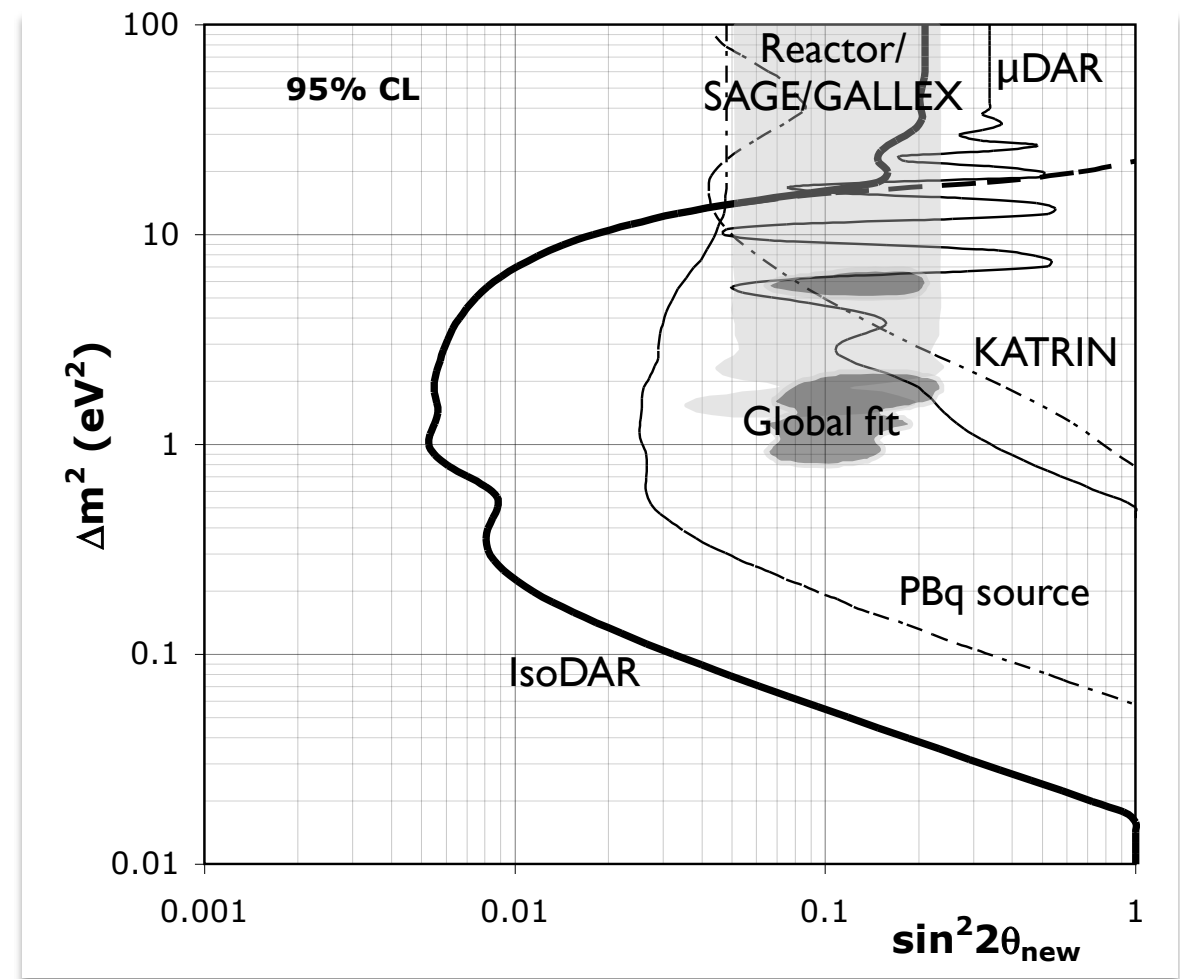
Adriana Bungau <A.Bungau@hud.ac.uk>

# Medium term: IsoDAR

[arXiv:1205.4419 \[hep-ex\]](https://arxiv.org/abs/1205.4419)



- High power cyclotrons create high  $\bar{\nu}_e$  flux
- $n + \text{Li}^7 \rightarrow \text{Li}^8$   
 $\hookrightarrow \bar{\nu}_e$  ,  $\langle E_\nu \rangle = 6.4 \text{ MeV}$
- Placed near a good  $\bar{\nu}_e$  detector (e.g. KamLAND) gives excellent sensitivity to sterile oscillation
- UK involved in accelerator and beam dump studies





To test LSND & MiniBooNE, Ga, and reactor anomalies

# NuSTORM



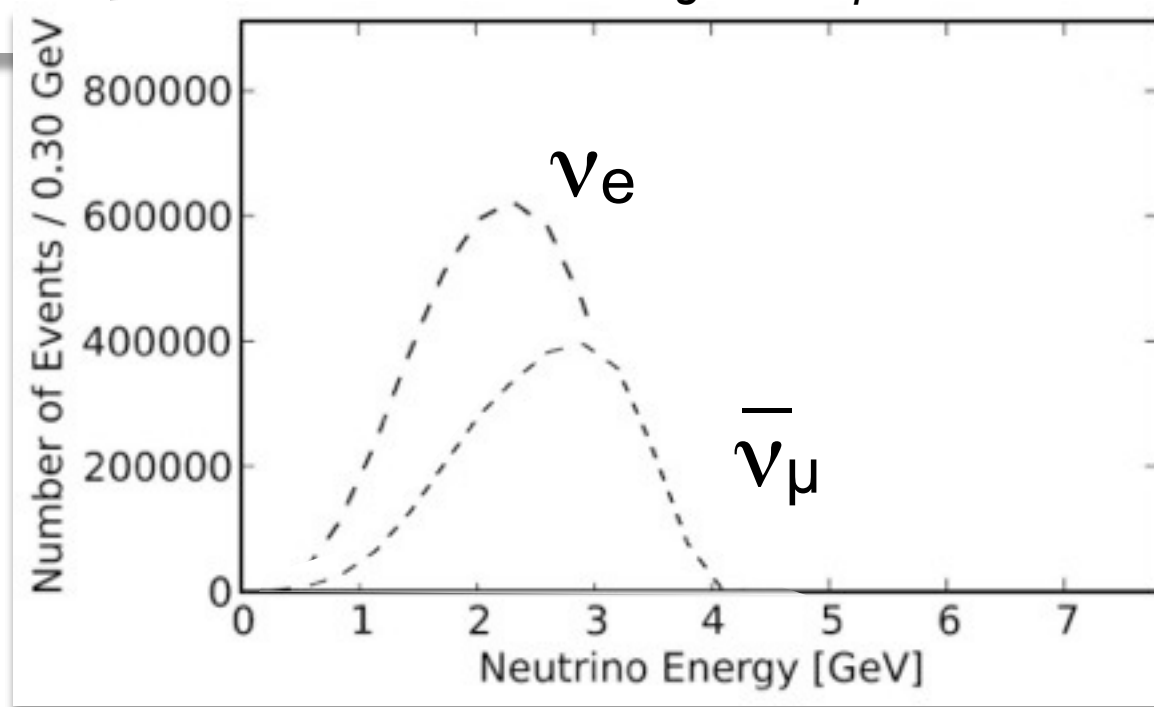
[arXiv:1206.0294 \[hep-ex\]](https://arxiv.org/abs/1206.0294)

Multiple sterile  $\nu$  channels  
Appearance Channel:

$$\nu_e \rightarrow \nu_\mu$$

Must reject the wrong sign  $\mu$  with **high** efficiency

Event rates/100T at Fe ND 50m from straight with  $\mu^+$  stored

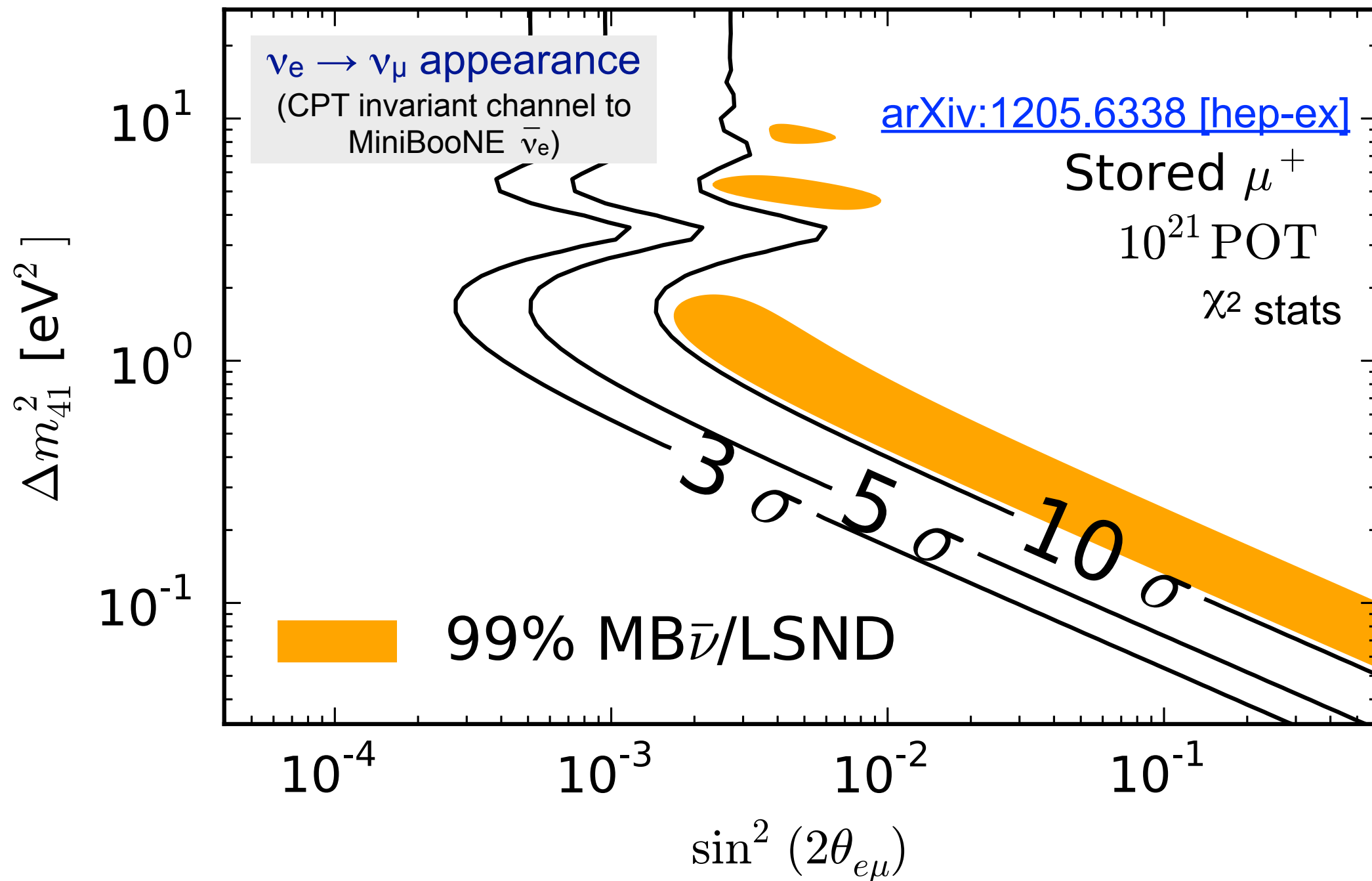


Received positive feedback from Fermilab PAC

[http://www.fnal.gov/directorate/program\\_planning/phys\\_adv\\_com/PAC%20Comments%20and%20Recommendations.pdf](http://www.fnal.gov/directorate/program_planning/phys_adv_com/PAC%20Comments%20and%20Recommendations.pdf)

[http://www.fnal.gov/directorate/program\\_planning/June2012Public/P-1028\\_LOI\\_Final.pdf](http://www.fnal.gov/directorate/program_planning/June2012Public/P-1028_LOI_Final.pdf)

# NuSTORM: oscillations



3+1  
Assumption

# NuSTORM physics programme

- As an experiment, NuSTORM can:
  - ✓ Perform direct tests of the LSND and MiniBooNE anomalies.
  - ✓ Perform direct tests of the Gallium and reactor anomalies.
  - ✓ Test the CP- and T-conjugated channels, constrain with disappearance.
  - ✓ Make precise and **unique** measurements of  $\nu_\mu$  and  $\nu_e$  cross-sections
- As a facility, NuSTORM:
  - ✓ Provides an accelerator technology test bed
  - ✓ Provides a powerful  $\nu$  detector test facility
- As a programme, NuSTORM:
  - ✓ Provides an important step on the path toward discovery in neutrinos and collider physics

Valuable physics  
input for  $\delta_{CP}$   
searches

Jaroslav's  
talk

Excellent  
synergy with  
superbeams!



# NuSTORM $\nu$ Cross-sections

- NuSTORM presents only way to measure  $\nu_e$ ,  $\bar{\nu}_\mu$  (&  $\bar{\nu}_e, \nu_\mu$ ) cross-sections in the same detector(s)
- Supports future long-baseline experiments!
- $E_\nu$  matched well to needs of these experiments

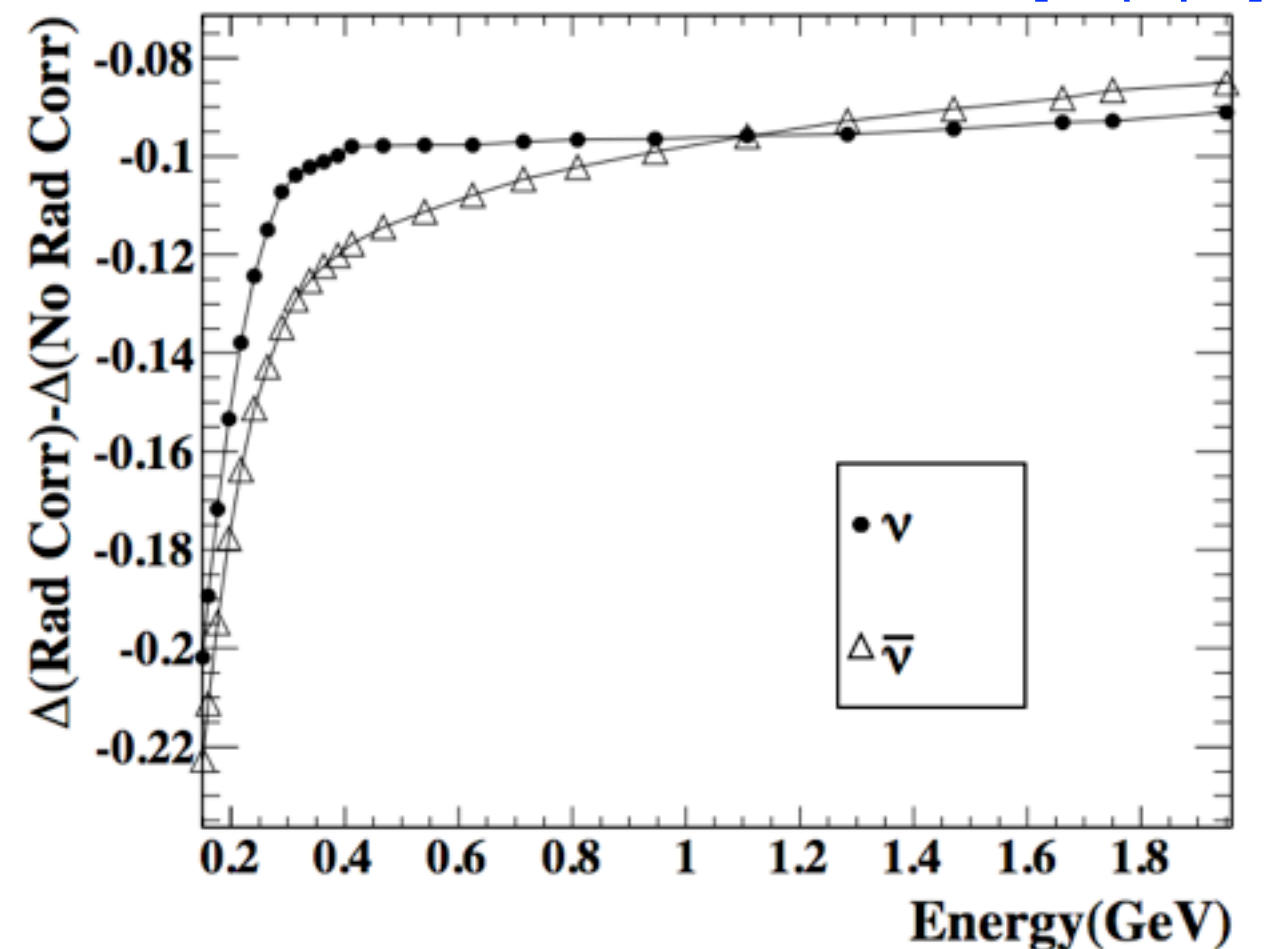
Recent calculations showing expectations for differences between

$\nu_e$  and  $\nu_\mu$  cross-sections

***We need data!***

NuSTORM members have submitted a statement to the PPAP and the CERN Strategy Committee

[arXiv:1206.6745 \[hep-ph\]](https://arxiv.org/abs/1206.6745)

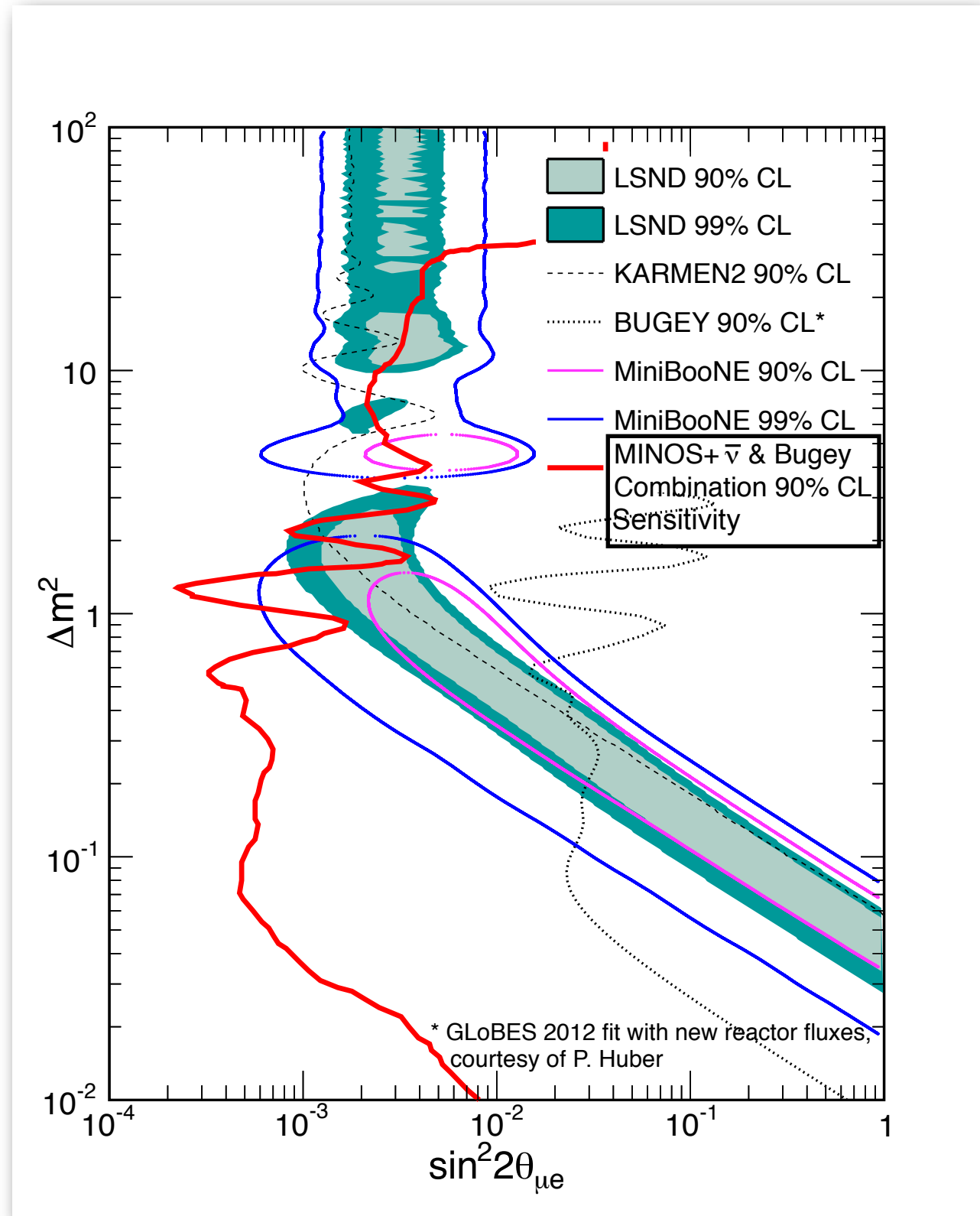


# MINOS+

- Extension of MINOS run in the medium-energy NuMI beam
- Runs concurrently with NOvA
- Sterile sensitivity:  

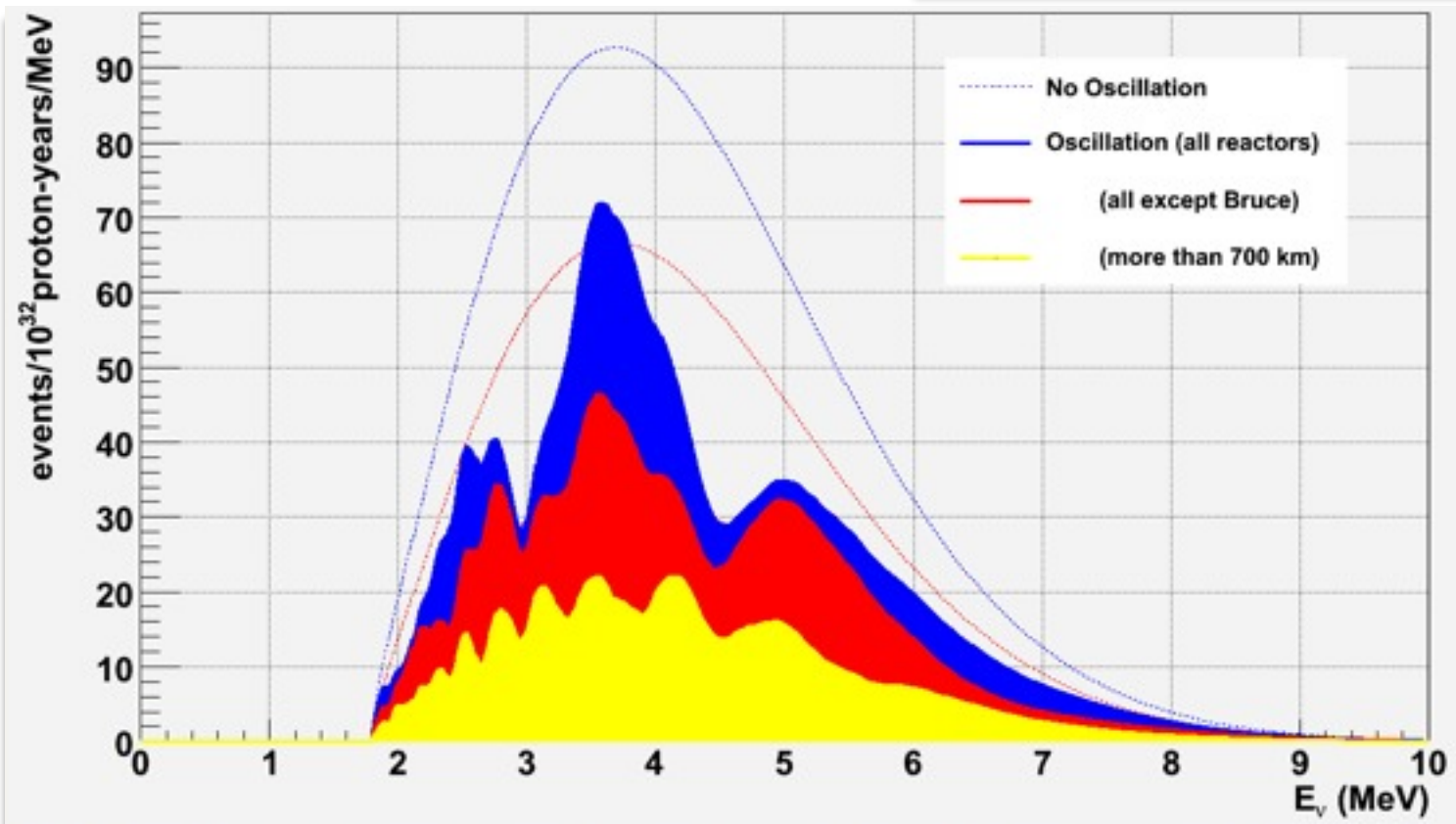
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) < 4P(\bar{\nu}_\mu \rightarrow \bar{\nu}_x) * P(\bar{\nu}_e \rightarrow \bar{\nu}_x)$$

Uses reanalysis of Bugey data and assumed sensitivity of MINOS+ with  $1.2 \times 10^{21}$  POT in nubar mode





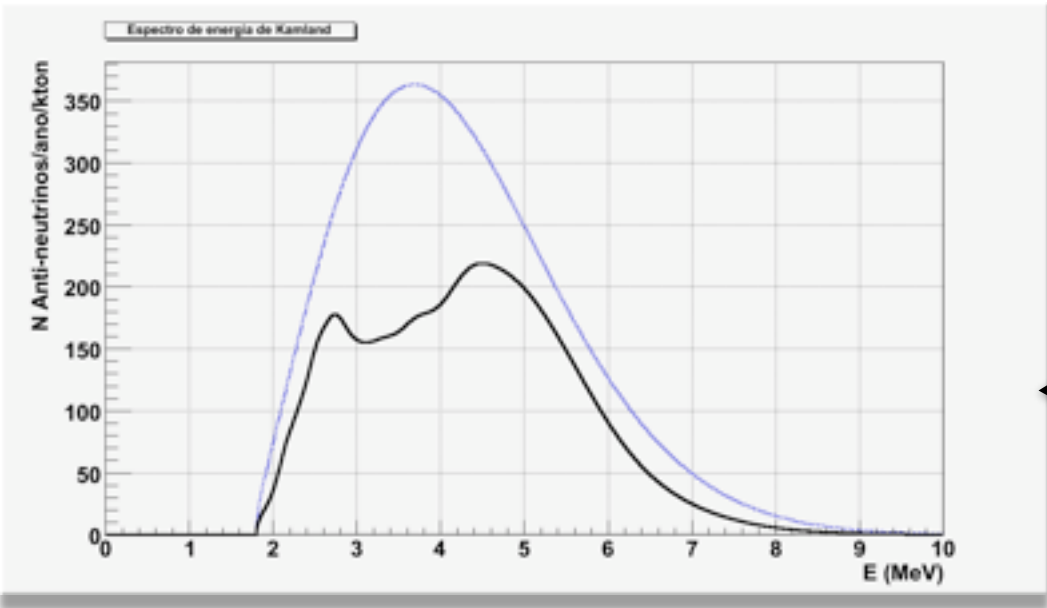
measurement of  $\Delta m^2_{12}$  with precision comparable to or better than KamLAND (despite lower statistics!)



Bruce reactor will contribute mainly to the central peak.



Reactor	d (km)	Th. Power (GW)
Bruce	281	10,32
Pickering	330	6,192
Darlington	340	10,572
R.E. Ginna	455	1,41
James A. Fitzpatrick	488	2,34
Nine Mile Point	488	5,07
Perry	530	3,1615
Enrico Fermi	559	3,255
Kewaunee	568	1,509
Davis-Besse	588	2,531
Point Beach	589	2,91
Palisades	617	2,34
Gentilly	648	1,914
Beaver Valley	657	4,929
Donald C. Cook	685	3,06



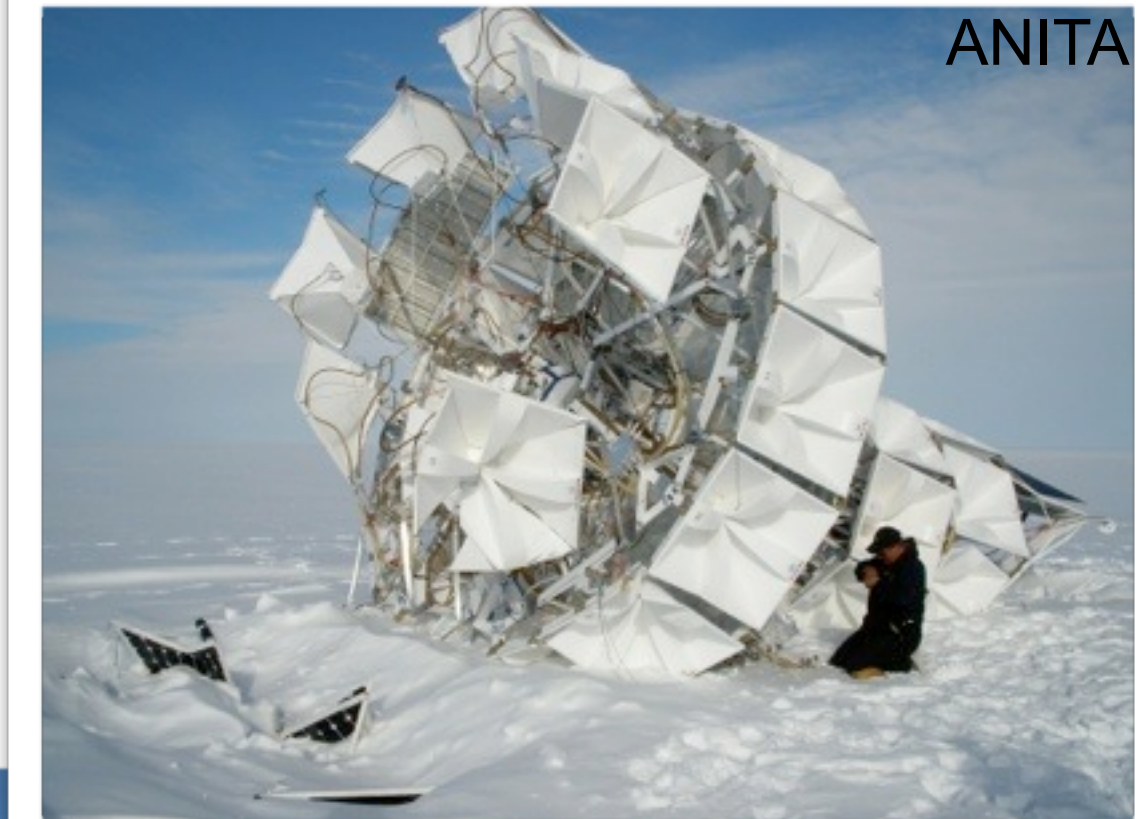
Much better defined peak structure compared to KamLAND



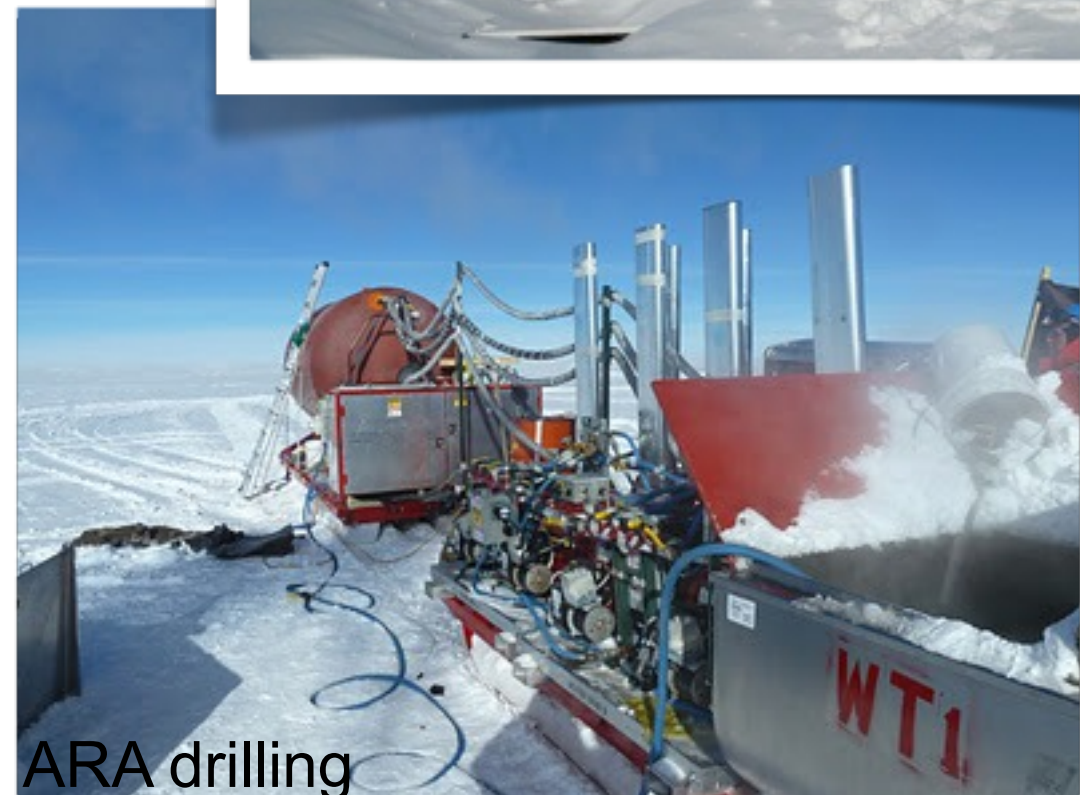


# UHE astrophysical neutrinos

- ANITA
  - Active project that will have new data in the next 18 months.
  - The third ANITA flight is scheduled for the Austral summer 2013/14
  - This flight will be the most sensitive to ultra-high energy neutrinos in the range  $10^{19}$  -  $10^{23}$  eV
- ARA (Askaryan Radio Array)
  - Will deploy the first three full prototype stations Jan 2013
  - One year of data comparable sensitivity to ANITA around  $10^{19}$  eV
  - If the stations prove successful would submit proposal for deploying the full 37 station array
  - Would seek wider UK/STFC support at that point



ANITA



ARA drilling

# Conclusions

- UK physicists are already working and assuming leadership roles in many small efforts that will grow in the medium term.
  - Even this list was not exhaustive!
- Several future experiments show good sensitivity to sterile  $\nu$  oscillation.
  - Expts that constitute steps in the global 3- $\nu$  paradigm must have priority.
  - IsoDAR and NuSTORM fit that bill.
  - Experiments that fulfil other important criteria should also be prioritised.
  - Twix grows out of nuclear non-proliferation efforts with significant KE potential. (And it's fast and inexpensive!)
- Lots of good work ongoing, with results and opportunities ahead.

*Not pushing an experiment, pushing an experimental programme!*





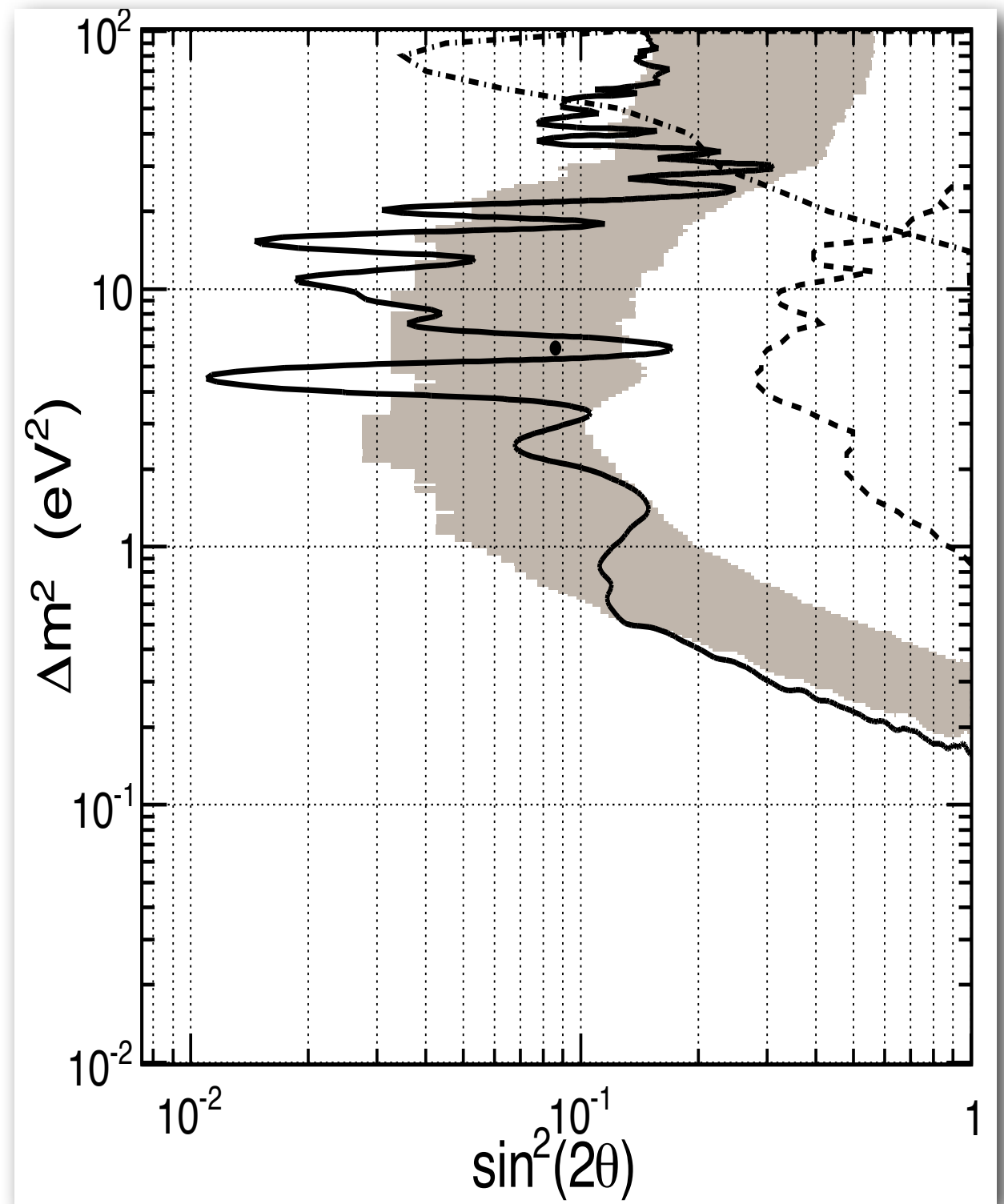
# THANK YOU!



# SciBooNE/MiniBooNE

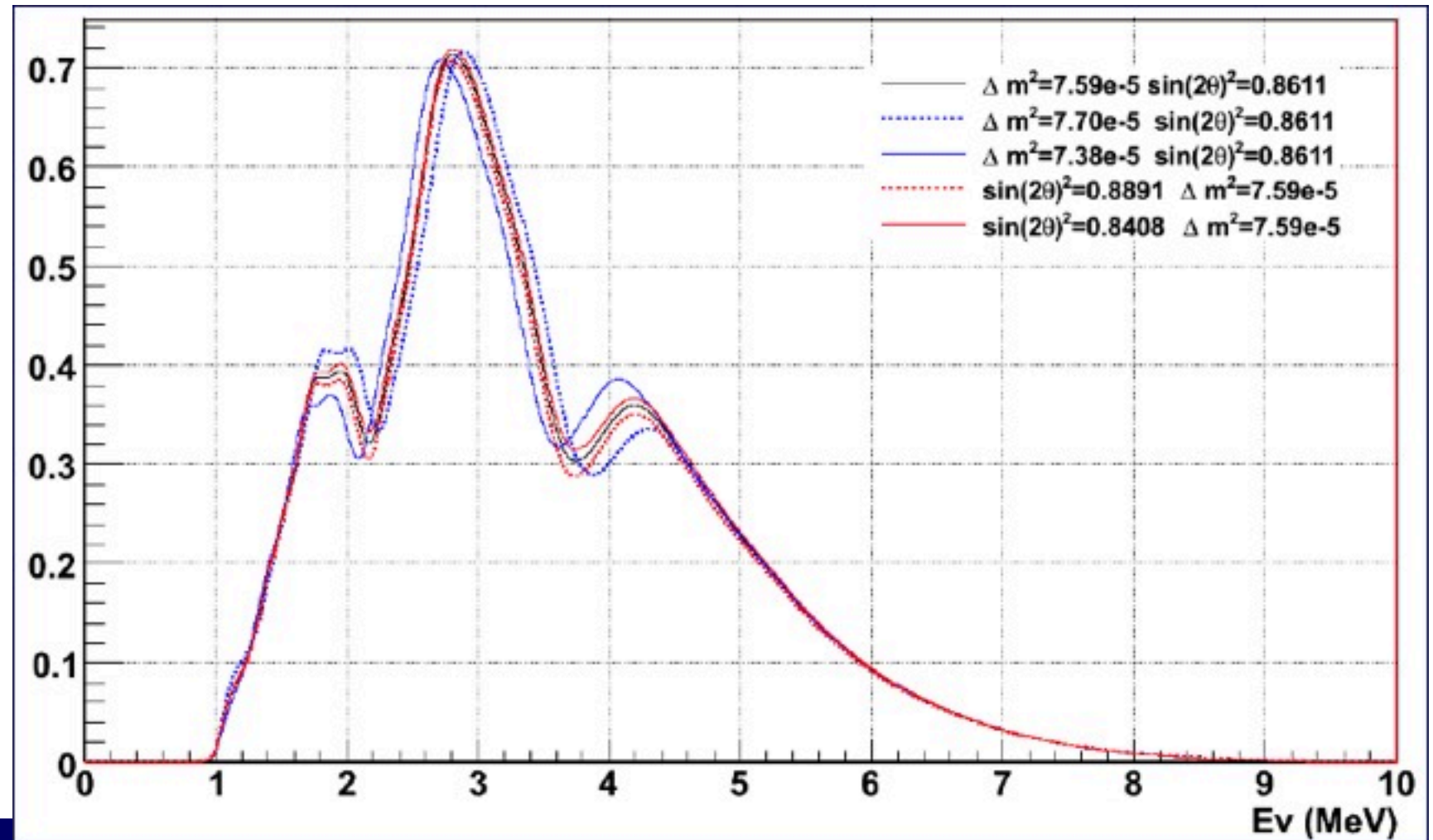
[arXiv:1208.0322 \[hep-ex\]](https://arxiv.org/abs/1208.0322)

- Muon antineutrino disappearance in MiniBooNE with SciBooNE as near detector
- World's best result over two decades of  $\Delta m^2$
- Data release available on the web:
- [http://www-sciboone.fnal.gov/data\\_release/joint\\_numubar\\_disap/](http://www-sciboone.fnal.gov/data_release/joint_numubar_disap/)



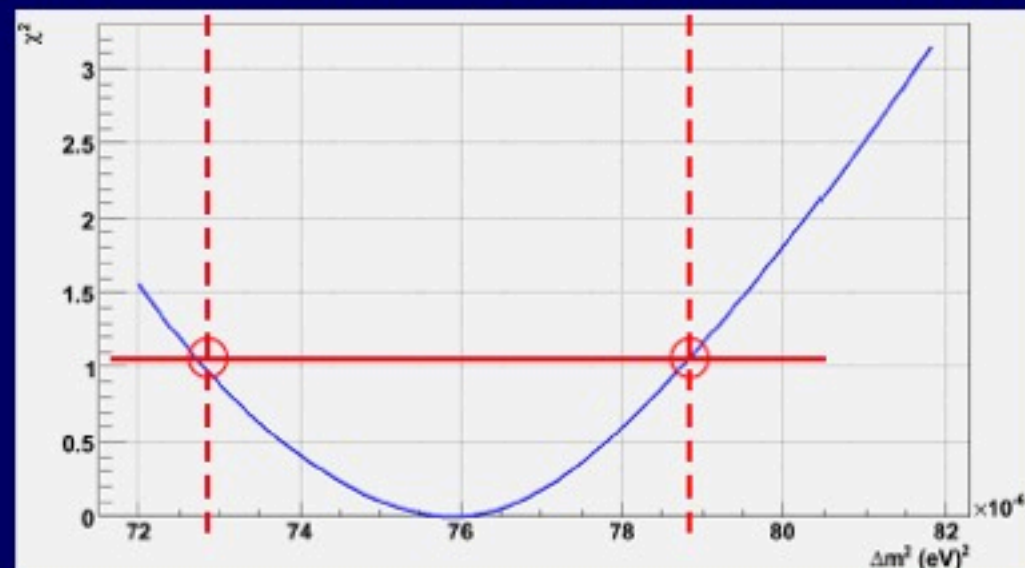


“Straw-man” calculation  
of statistics-limited  
precision as a function  
of assumed resolution



### • Study for $\Delta m^2$

Chi<sup>2</sup> test applied to the visible energy spectrum with a varying  $\Delta m^2$  around  $7.59 \times 10^{-5} \text{ eV}^2$  and a fixed value of  $\sin^2(2\theta)=0.8611$  (plot for 6% resolution).



Resolution	-1σ	+1σ
6%	$7.27 \times 10^{-5}$	$7.88 \times 10^{-5}$
3%	$7.32 \times 10^{-5}$	$7.82 \times 10^{-5}$

6% :  $\Delta m_{12}^2 = 7.59^{+0.29}_{-0.32} \times 10^{-5} (\text{eV})^2$

3% :  $\Delta m_{12}^2 = 7.59^{+0.23}_{-0.27} \times 10^{-5} (\text{eV})^2$

Relative  
Difference  
in errors  
~20%

(life-time= $1 \times 10^{32}$  proton-year, approx 1.8 years for the calculations)

# NuSTORM Assumptions

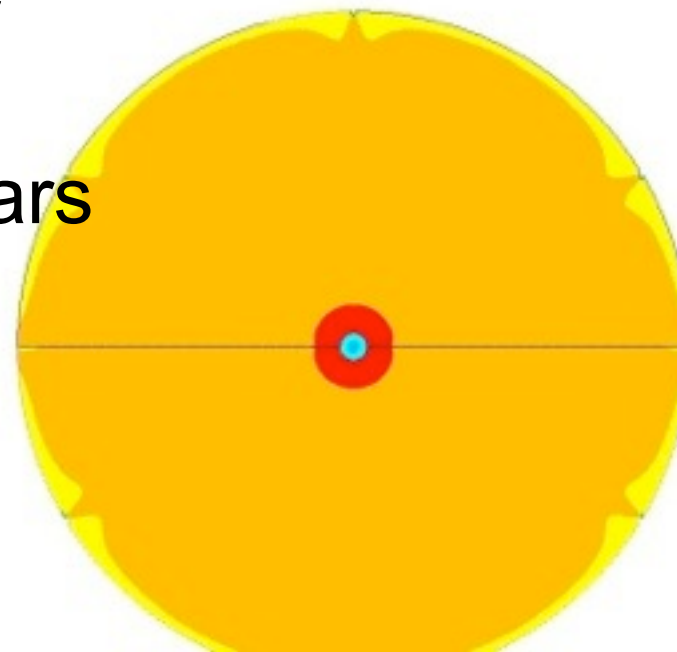
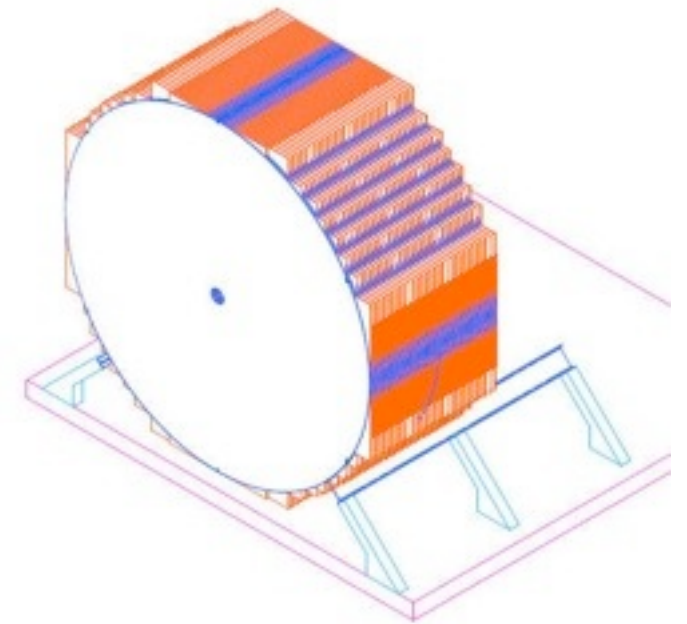
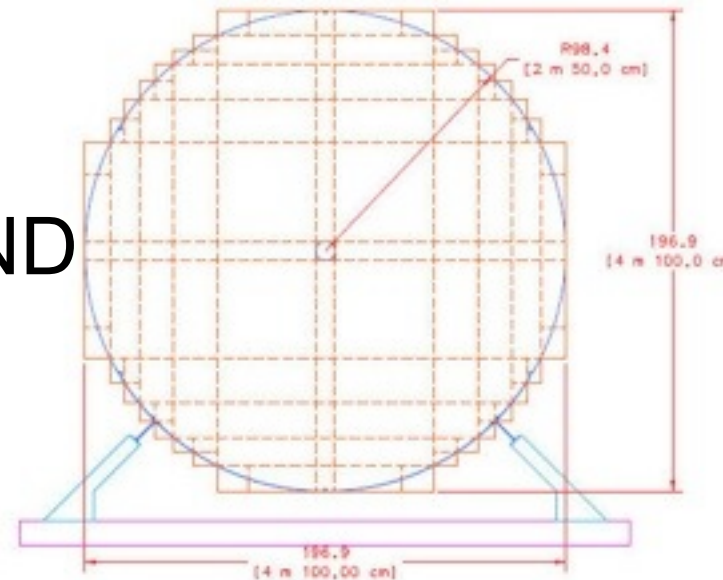
- $N_\mu = (\text{POT}) \times (\pi/\text{POT}) \times \epsilon_{\text{collection}} \times \epsilon_{\text{inj}} \times (\mu/\pi) \times A_{\text{dynamic}} \times \Omega$ 
  - $10^{21}$  POT in 5 years of running @ 60 GeV in Fermilab PIP era
  - 0.1  $\pi/\text{POT}$  (FODO)
  - $\epsilon_{\text{collection}} = 0.8$
  - $\epsilon_{\text{inj}} = 0.8$
  - $\mu/\pi = 0.08$  ( $\gamma_{\text{ct}} \times \mu$  capture in  $\pi \rightarrow \mu$  decay) [ $\pi$  decay in straight]
    - Might do better with a  $\pi \rightarrow \mu$  decay channel
  - $A_{\text{dynamic}} = 0.75$  (FODO)
  - $\Omega = \text{Straight/circumference ratio}$  (0.43) (FODO)
- This yields  $\approx 1.7 \times 10^{18}$  useful  $\mu$  decays



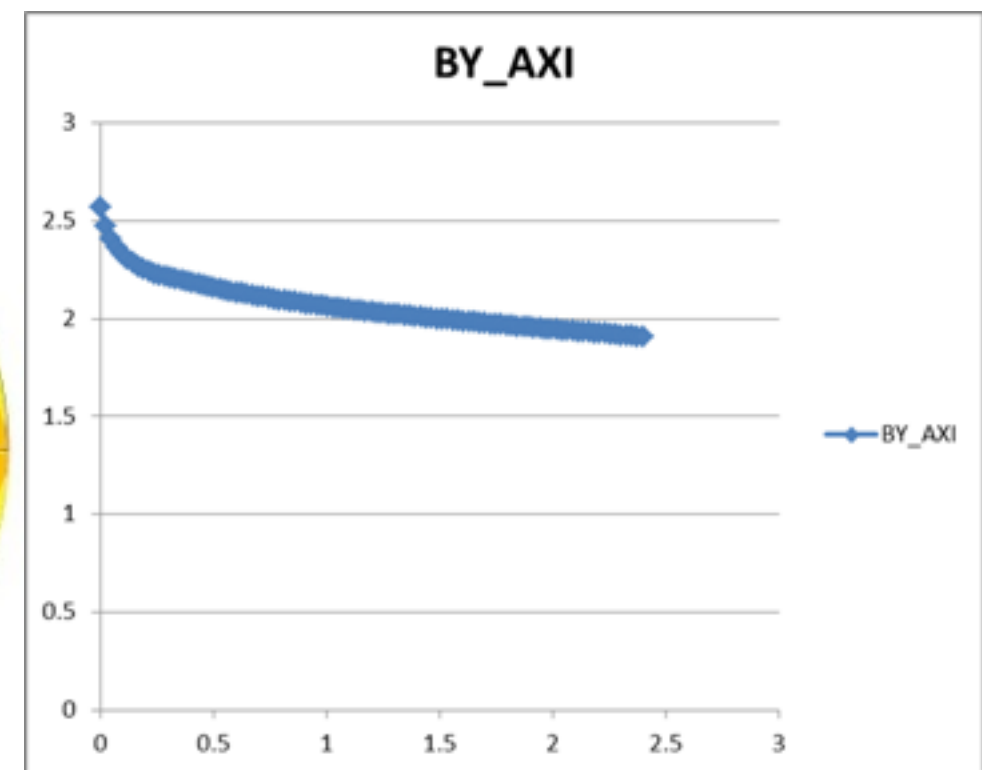
# Baseline Detector

## Super B Iron Neutrino Detector: SuperBIND

- Magnetized Iron
  - 1.3 kT
    - Following MINOS ND ME design
    - 1-2 cm Fe plate
    - 5 m diameter
- Utilize superconducting transmission line for excitation
  - Developed 10 years ago for VLHC
- Extruded scintillator + SiPM

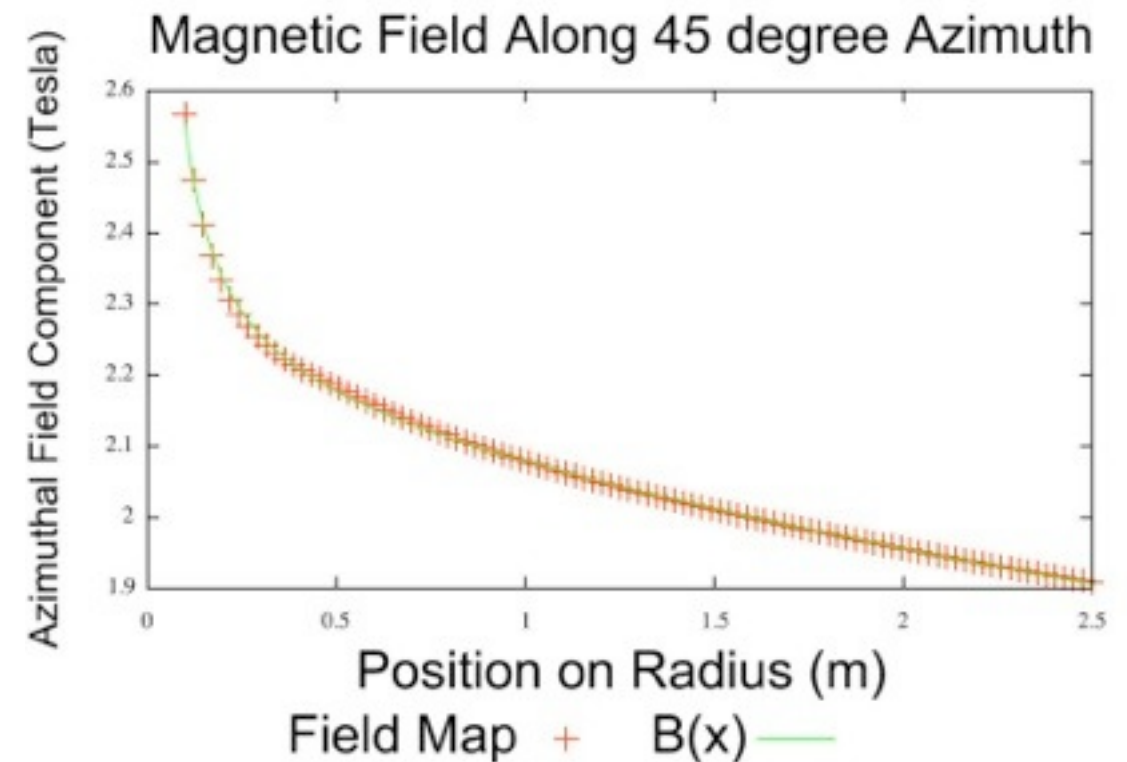


20 cm hole  
For 3 turns  
of STL

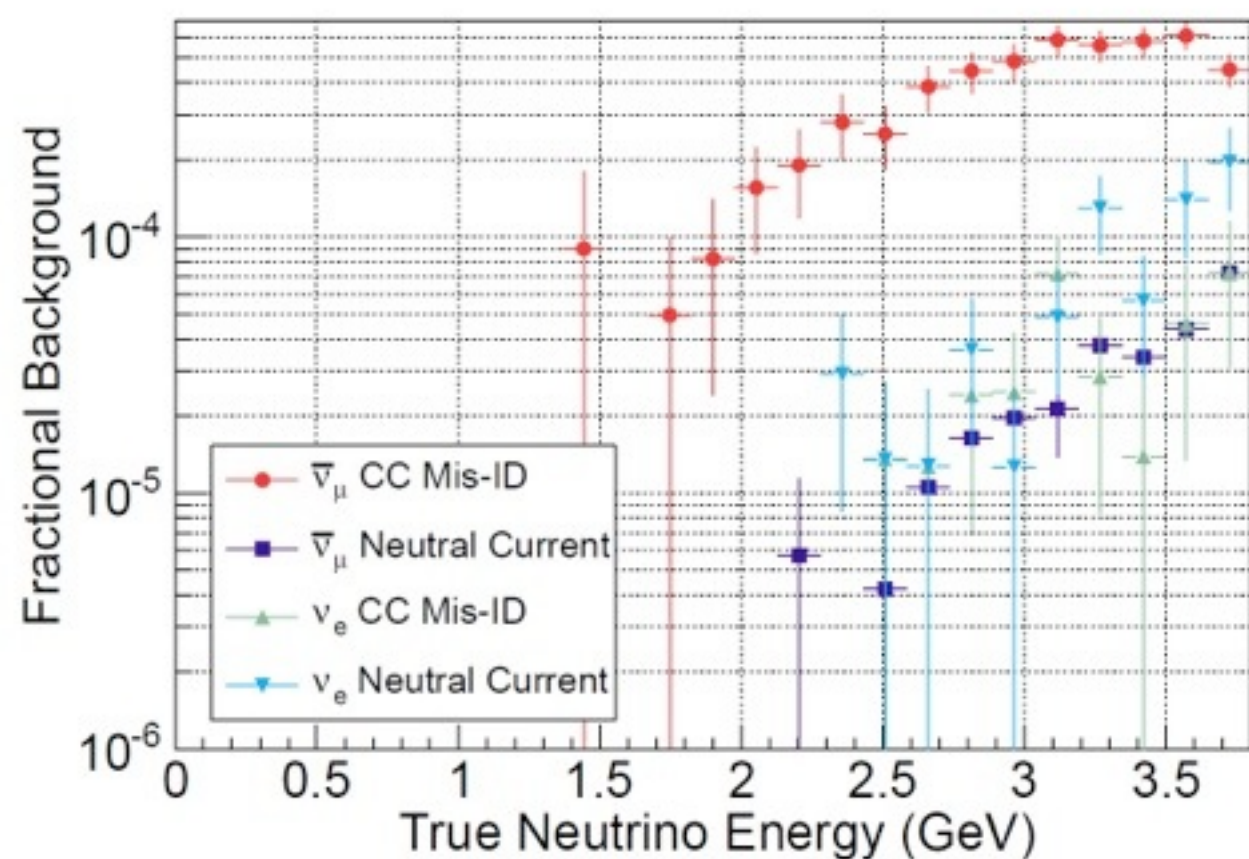
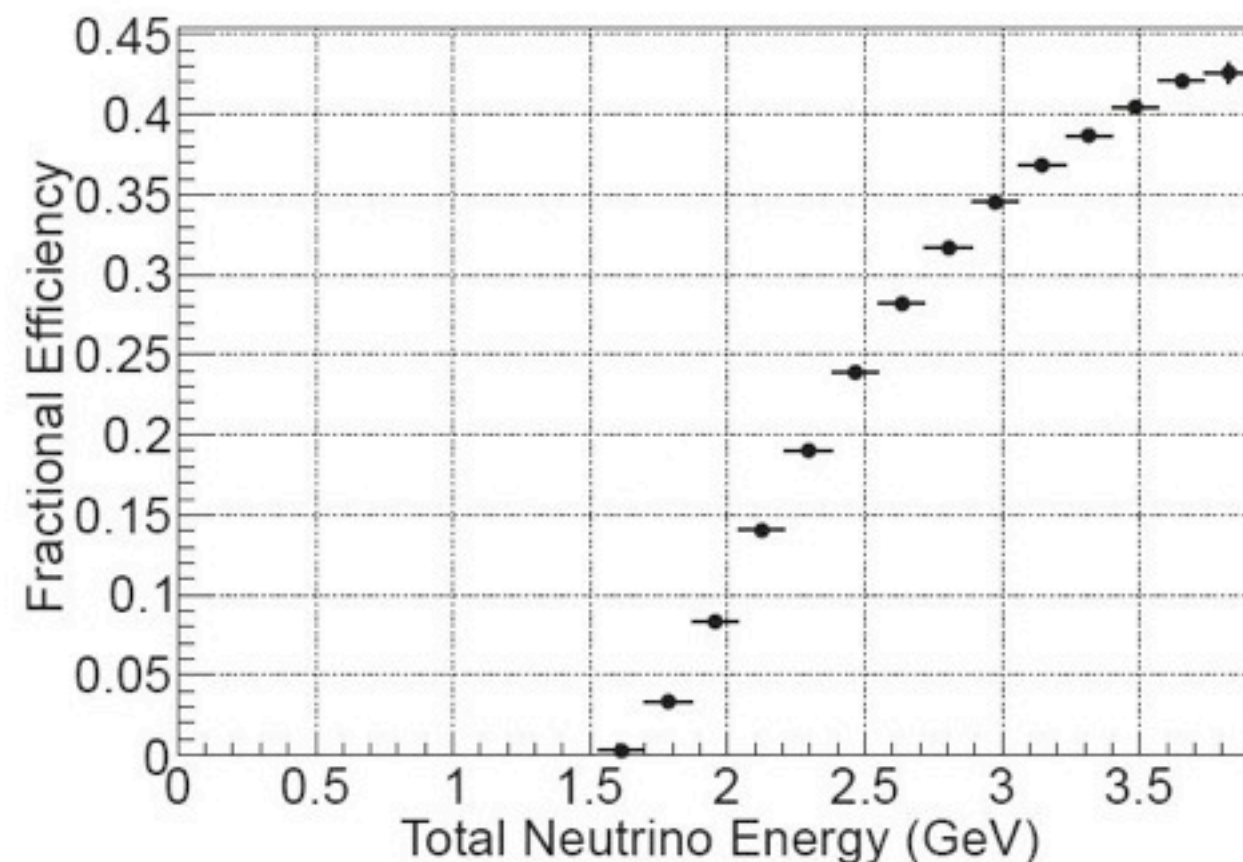
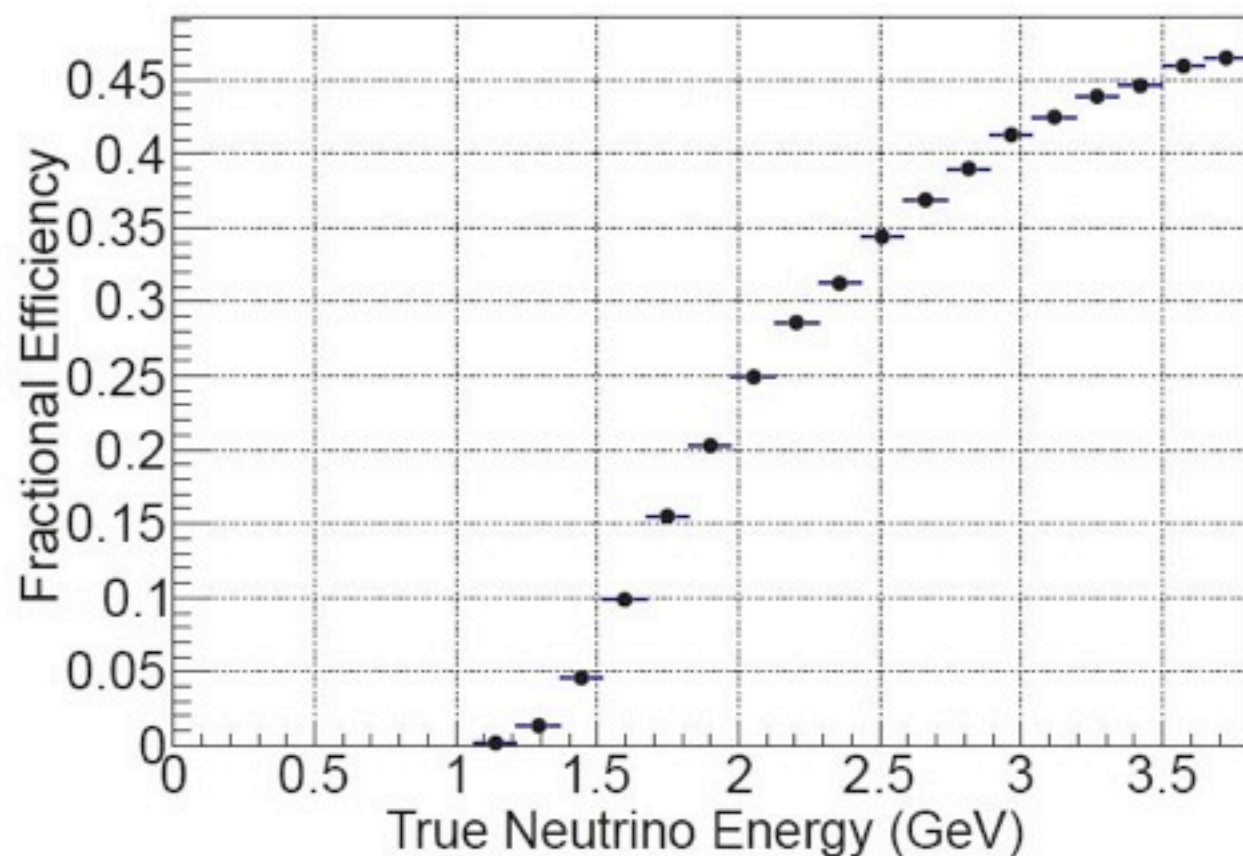


# Simulation – $\nu_\mu$ appearance

- Full GEANT4 Simulation
  - Extrapolation from ISS and IDS-NF studies for the MIND detector
  - Uses GENIE to generate the neutrino interactions.
  - Involves a flexible geometry that allows the dimensions of the detector to be altered easily (for optimization purposes, for example).
  - Does not yet have the detailed B field, but parameterized fit is very good
- Event selection/cuts
  - Cuts-based analysis
  - Multivariate to come later

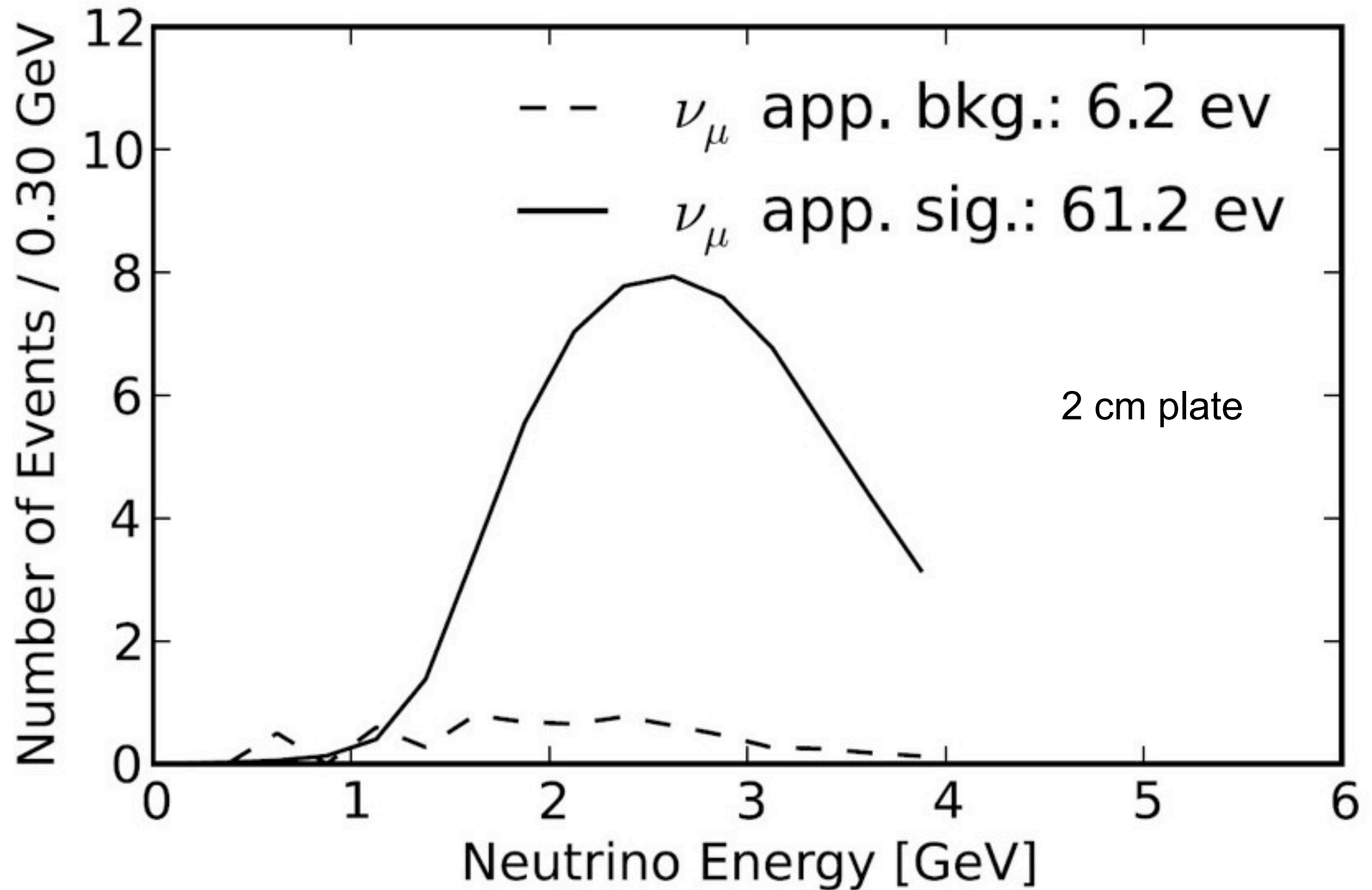








$\nu_e \rightarrow \nu_\mu$  appearance  
CPT invariant channel to MiniBooNE  $\bar{\nu}_e$



# Raw Event Rates

Neutrino mode with stored  $\mu^+$ .

Channel	$N_{\text{osc.}}$	$N_{\text{null}}$	Diff.	$(N_{\text{osc.}} - N_{\text{null}})/\sqrt{N_{\text{null}}}$
$\nu_e \rightarrow \nu_\mu$ CC	332	0	$\infty$	$\infty$
$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ NC	47679	50073	-4.8%	-10.7
$\nu_e \rightarrow \nu_e$ NC	73941	78805	-6.2%	-17.3
$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ CC	122322	128433	-4.8%	-17.1
$\nu_e \rightarrow \nu_e$ CC	216657	230766	-6.1%	-29.4

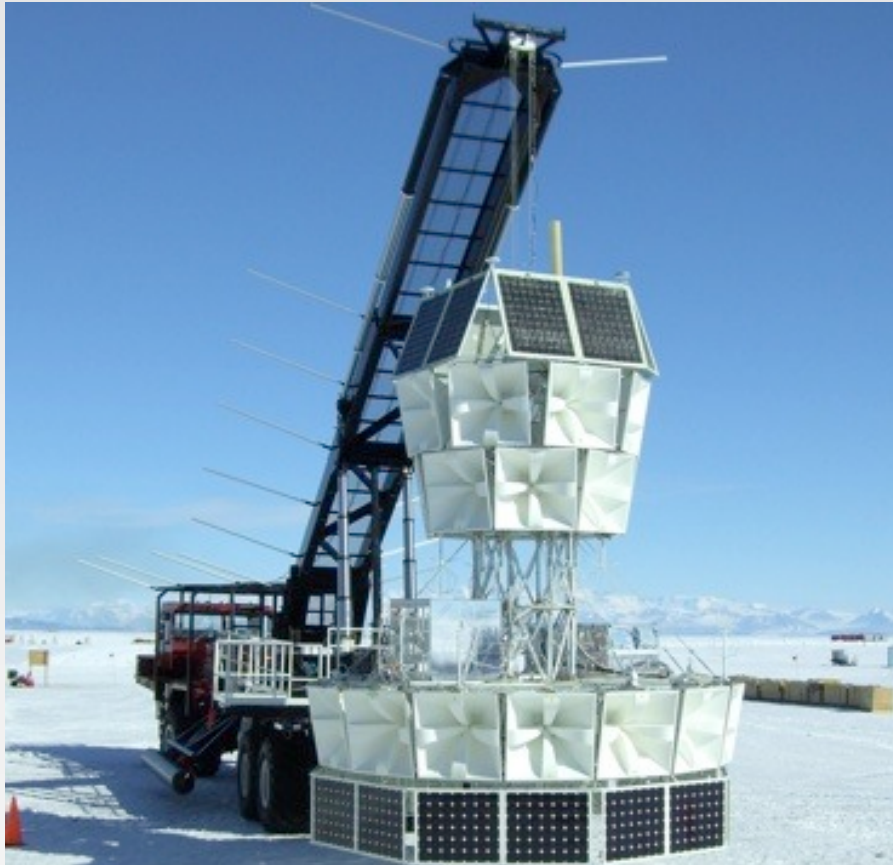
3+1  
Assumption

Anti-neutrino mode with stored  $\mu^-$ .

Channel	$N_{\text{osc.}}$	$N_{\text{null}}$	Diff.	$(N_{\text{osc.}} - N_{\text{null}})/\sqrt{N_{\text{null}}}$
$\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ CC	117	0	$\infty$	$\infty$
$\bar{\nu}_e \rightarrow \bar{\nu}_e$ NC	30511	32481	-6.1%	-10.9
$\nu_\mu \rightarrow \nu_\mu$ NC	66037	69420	-4.9%	-12.8
$\bar{\nu}_e \rightarrow \bar{\nu}_e$ CC	77600	82589	-6.0%	-17.4
$\nu_\mu \rightarrow \nu_\mu$ CC	197284	207274	-4.8%	-21.9

 Appearance channels

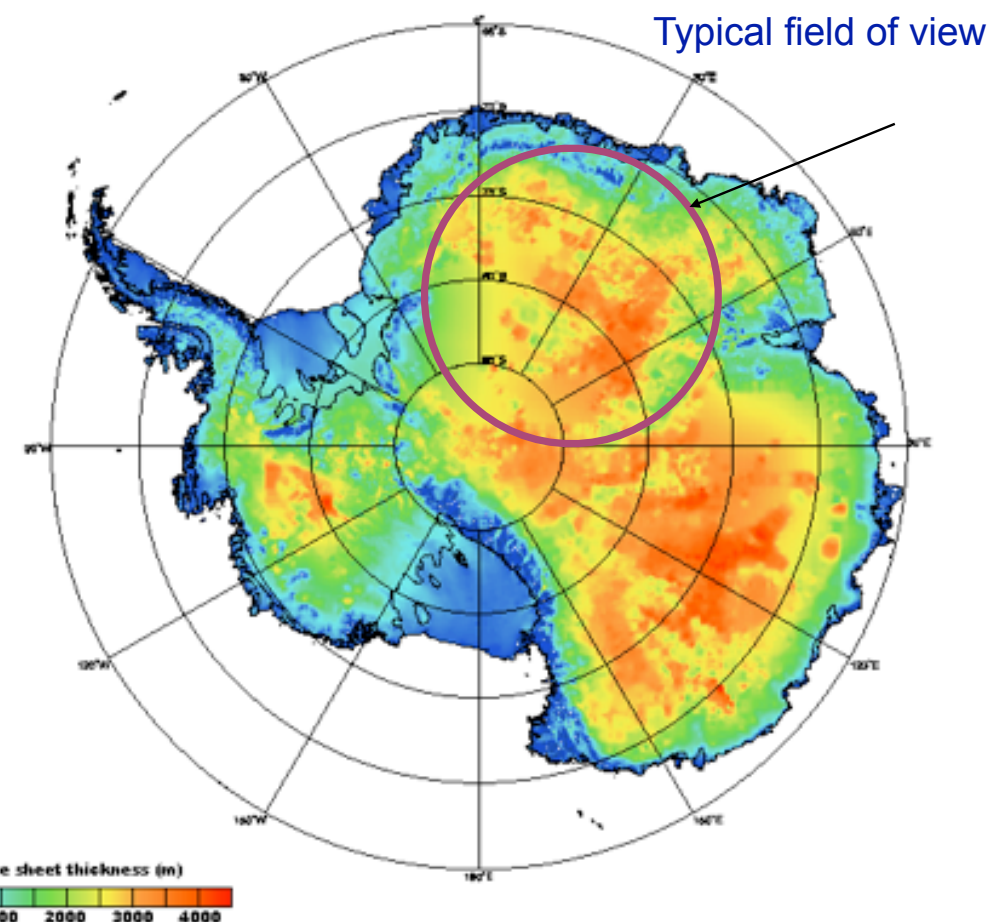
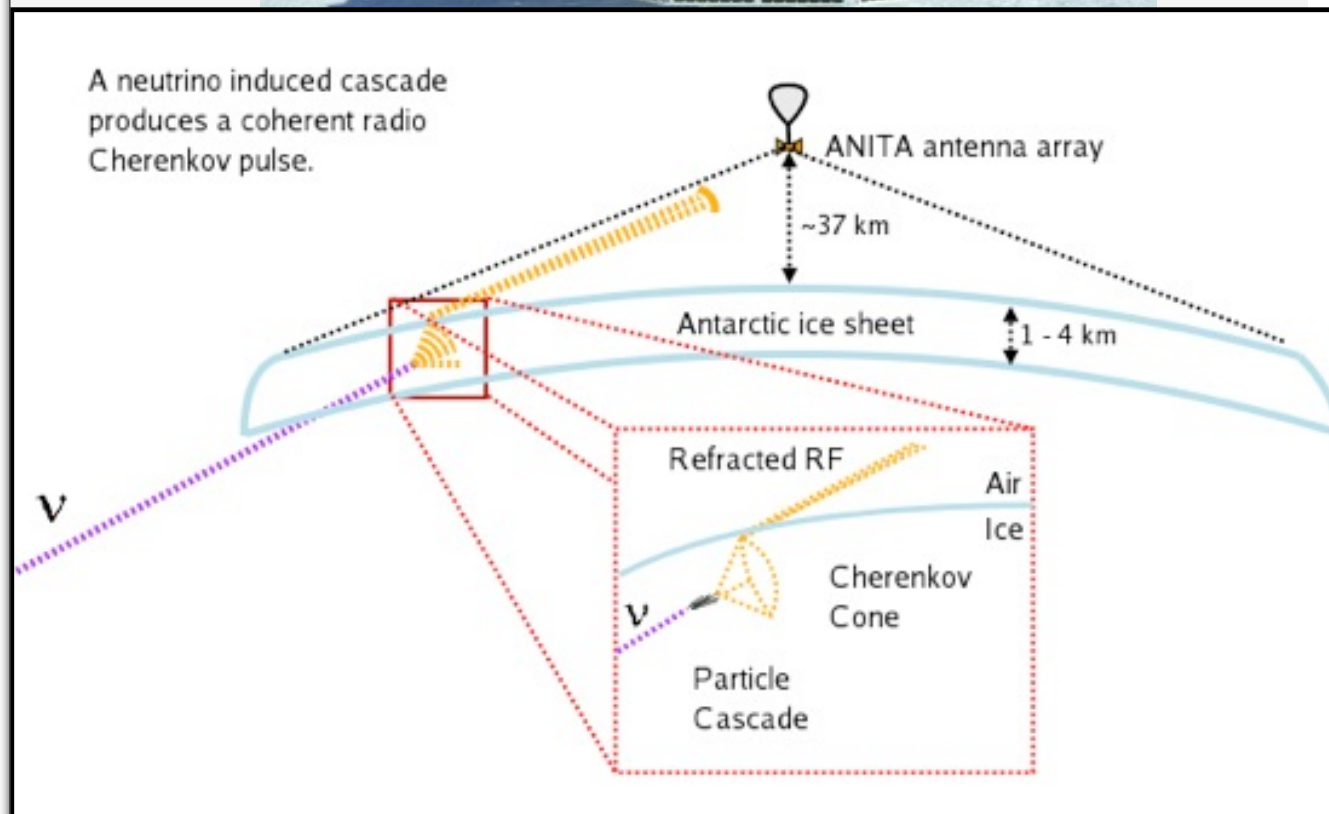
# ANITA



ANITA looks for neutrinos interacting in the antarctic ice, by dangling from a balloon 37km above the continent.

Over a million cubic kilometres of ice visible.

Third flight scheduled for December 2013



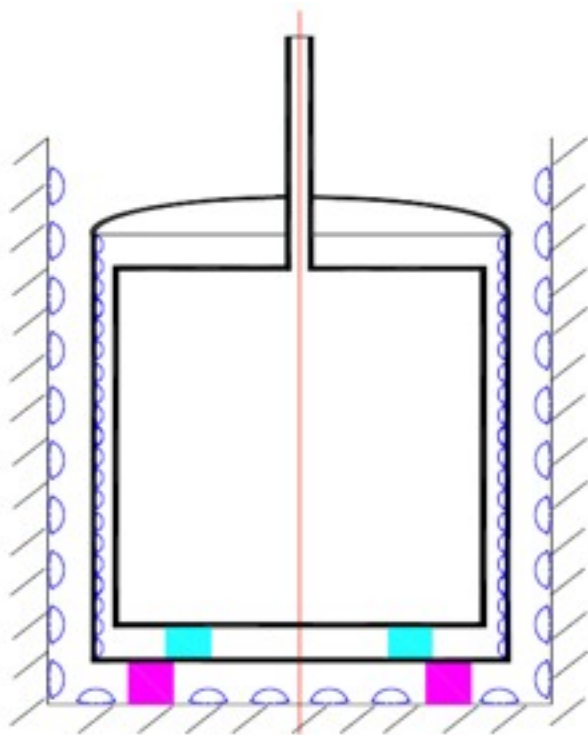


# Daya Bay 60km

Jenny Thomas <[jthomas@hep.ucl.ac.uk](mailto:jthomas@hep.ucl.ac.uk)>

>40GW

- Neutrino target: ~20kt  
LS, LAB based
- 30m(D)×30m(H)
- Oil buffer: 6kt
- Water buffer: 10kt
- PMT: 15000 20"



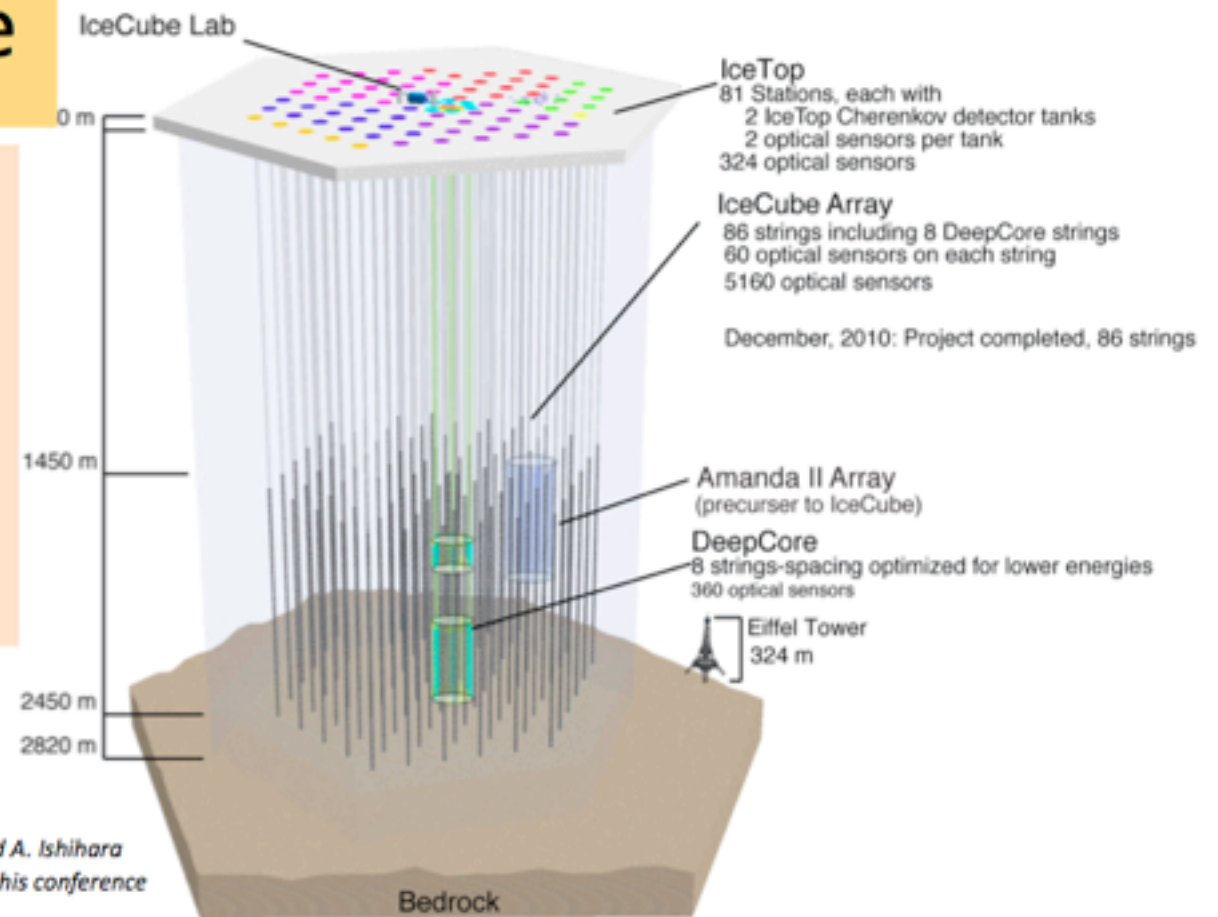
2012-2014 R&D  
2015 proposal to government  
2016-2020 construction  
2020 start operation

- Deep Core has proved the method of reduction of energy threshold with higher density strings
- Pingu will push it further... and they will still have 1 MegaTon!!!
- 2 years running will measure MH to between  $3-11\sigma$

## Water/ice Cherenkov detectors: IceCube

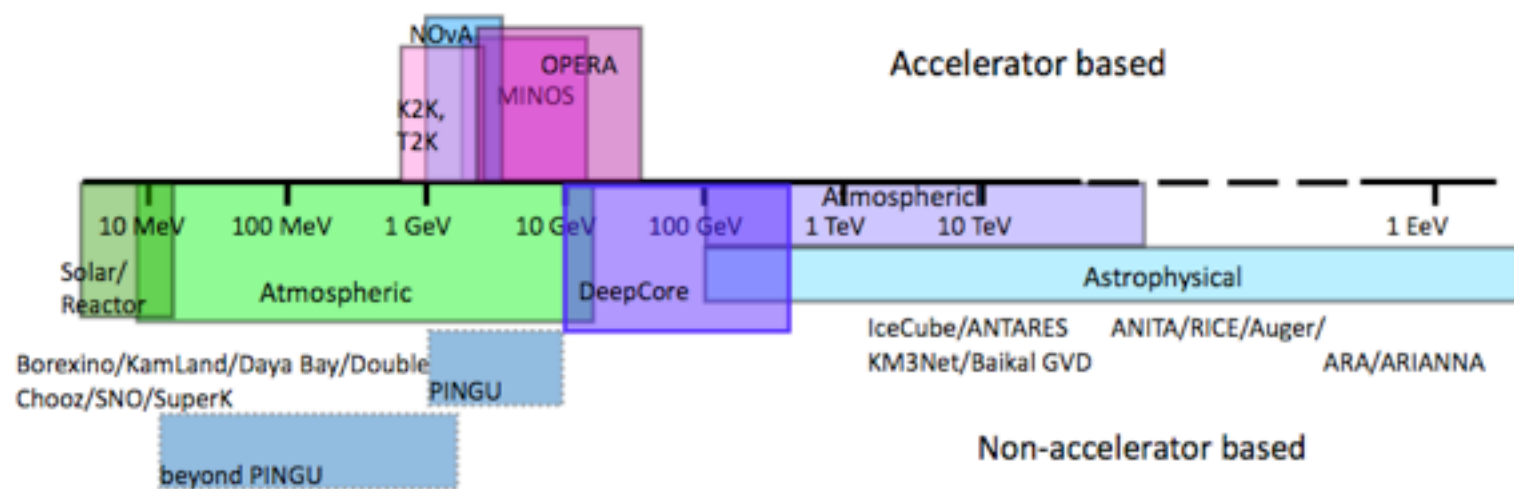
### IceCube

- Total of 86 strings and 162 IceTop tanks;
- Completion with 86 strings: December 2010
- Full operation with all strings since May 2011.



A. Karle, UW-Madison

## The Neutrino Detector Spectrum



~70 active members in feasibility studies:  
IceCube, KM3Net, Several neutrino experiments  
Photon detector developers  
Theorists

Slide after:  
Darren Grant  
NNN 2011



- Further exploitation of the NuMI beam is a good idea in the short term
  - The beam is the highest intensity beam in the world
  - It exists
- FNAL could consider:
  - Off-axis at Ash River (810km, 14mr)
  - On-axis on surface at Soudan or beyond (735km, LE)
  - On-axis underground at Soudan Laboratory (735km, LE)
  - Cherenkov detector in Lake Superior (655km, 20mr off axis)
- There is room for thinking out of the box, and time to get results
  - NuMI will run for 6-10 years







水戸の梅の花