Applied String Theory

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Holography (AdS/CFT)

Strongly interacting quantum field theory



Gravity in (at least) one dimension higher



Many Explicit Examples

d=3+1 SU(N) Yang-Mills
 + scalars and fermions
(with lots of supersymmetry)



d=9+1 Type IIB string theory on AdS₅ x S⁵

d=2+1 *SU(N)* x *SU(N)* Chern-Simons + matter (with lots of supersymmetry)



d=9+1 Type IIA string theory on $AdS_4 \ge CP^3$

e.g

Two Major Directions

• Explore details of individual supersymmetric theories

• Study physical properties of generic theories

Physics in the Unrealistic Regime

Why you shouldn't care

- Complicated theories
- Large N
- (Supersymmetric)

Why you should

- Strongly coupled
- Completely solvable
- Universal behaviour

Universal Behaviour

The Vacuum State



My attempt at drawing Anti de-Sitter spacetime

 $ds^2 = rac{L^2}{r^2} \left(dr^2 + \eta_{\mu
u} dx^\mu dx^
u
ight)$

Heating up the Boundary Theory



Boundary field theory

- Black hole
- Hawking radiation = finite temperature

Finite Density Matter



Boundary field theory

- Reissner-Nordstrom black hole
- Hawking radiation = finite temperature
- Electric field = chemical potential

Three Stories

Story 1: Fluid Mechanics

Ultimate Low Energy Effective Theory



 $\rho(\vec{x},t) \qquad T(\vec{x},t) \qquad \vec{u}(\vec{x},t)$

Shear Viscosity



Kovtun, Policastro, Son, Starinets (2003)

Fluid-Gravity Correspondence



Navier-Stokes Equations

Einstein Equations

Battacharya, Hubeny, Minwalla, Rangamani (2009)

Happy Corollary

Extra terms in hydrodynamics

Minwalla et al. (2011) Son and Surowka (2009)

For the Future

Singularities? Turbulence?

Story 2: Conductivity

Unconventional Superconductors



Cleanest examples in d=2+1 dimensional planes

Ohm's Law



Boundary field theory d=2+1

 $j(\omega) = \sigma(\omega)E(\omega)$

Optical Conductivity in d=2

Herzog, Kovtun, Sachdev and Son (2007); Hartnoll (2008)

$$j(\omega) = \sigma(\omega) E(\omega)$$





Optical Conductivity in d=2

Herzog, Kovtun, Sachdev and Son (2007); Hartnoll (2008)

 $\operatorname{Re}\sigma(\omega) \sim K\,\delta(\omega)$





Because our theory is missing an underlying lattice

The lattice black hole

Horowitz, Santos, Tong (2012)



Spatially dependent chemical potential

 $\mu = \mu [1 + A\cos(k_L x)]$

<u>Parameters:</u> T, μ, k_L, A



Optical Conductivity

 $j(\omega) = \sigma(\omega) E(\omega)$



Delta function spreads out

Low Frequency Behaviour



Classical Drude model from general relativity

Mid-Frequency Behaviour



$$|\sigma(\omega)| = \frac{B}{\omega^{2/3}} + C$$

Surprising power-law behaviour

Robust Power-Law





$$|\sigma(\omega)| = \frac{B}{\omega^{2/3}} + C$$

B temperature independent

Scaling in High T_c Superconductors



 $Bi_2Sr_2Ca_{0.92}Y_{0.08}Cu_2O_{8+\delta}$

van der Marel et al. (2003)

For the Future

Why?



 $|\sigma(\omega)| \sim \omega^{-2/3}$

Story 3: Non-Fermi Liquids

Landau Fermi Liquid Theory

Interacting Fermions at density



Free fermions with Fermi surface

Except...



No Quasi-Particles



Major open problem in theoretical physics

Fermions in AdS/CFT



Solve the Dirac equation in this background: $D \psi - m \psi = 0$

Sung-Sik Lee (2009)

Non-Fermi Liquids

- Fermi Surface
 - Zero energy excitations at non-zero momentum
- Excitation spectrum depends on various parameters
 - Fermion mass, background geometry



Cubrovic et al. (Leiden) Liu et al. (MIT) (2009)

But there's an instability...

Hartnoll, Polchinski, Siverstein and Tong (2009)



...and a star forms in the bulk.

This is an *electron star*. The low energy physics appears to be that of a Landau Fermi liquid. ???

Explored in detail by Hartnoll and collaborators.

For the Future

Understand the electron star in more detail

Summary

Holography very useful tool to explore aspects of strongly interacting matter

Much Much More...

Jet Quenching Chesler and Yaffe



Quantum Turbulence Adams, Chesler and Liu



Crystal Formation Donos, Gauntlett + many others



Lowest Landau Level Physics Blake, Bolognesi, Tong, Wong



The End

Thank you for your attention

Extra Material

DC Resistivity

$$\rho \sim T^{2\nu - 1}$$

• Exponent depends on lattice spacing: $\nu = \frac{1}{2}\sqrt{5 + 2(k/\mu)^2 - 4\sqrt{1 + (k/\mu)^2}}$

• This is characteristic of a *locally critical* theory i.e. $z \to \infty$

Hartnoll, Hofman (2012)

• More recently, a mechanism suggested to drive this to linear resistivity.

Donos, Hartnoll (2012)

Navier-Stokes Equations

$$\rho(\vec{x},t) \qquad T(\vec{x},t) \qquad \vec{u}(\vec{x},t)$$

$$\partial^{\mu}T_{\mu\nu} = 0 \qquad \qquad T_{\mu\nu} = \dots$$
$$\partial^{\mu}J_{\mu} = 0 \qquad \qquad J_{\mu} = \dots$$

Evolution of Exact Results at Strong Coupling



- '70s and '80s: Quantities protected by symmetries and anomalies
 - Chiral Lagrangian
- '90s and '00s: Quantities protected by supersymmetry
 - BPS quantities, superpotentials
- '00s and '10s: Anything you like
 - But only in very particular, large N, theories