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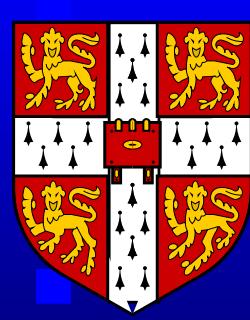
# Directions for New Physics



Ben Allanach<sup>a</sup> (University of Cambridge)

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<sup>a</sup>BCA, Gripaios, 1202.6616; BCA, Sridhar 1205.5170; BCA,  
Parker 1211.3231.



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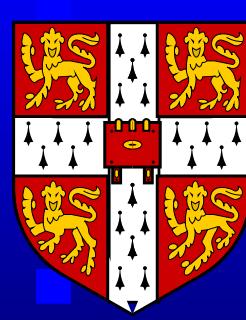


# Hierarchy problem solved

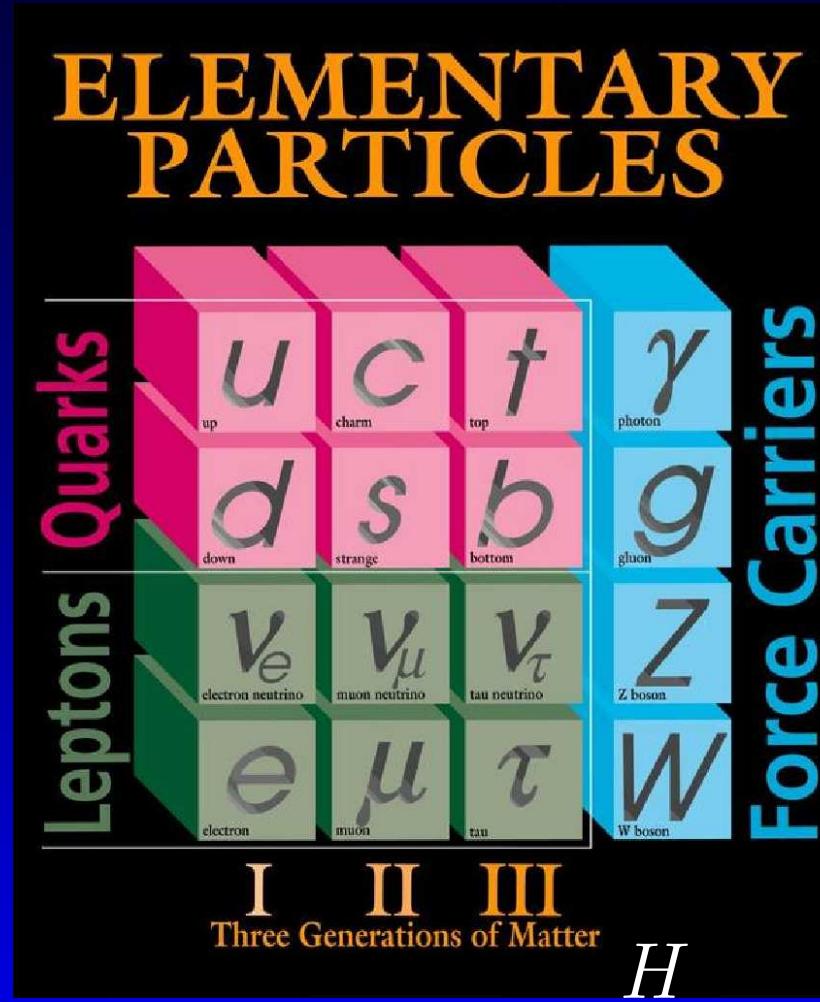
The supersymmetric copies *cancel* the quantum fluctuations. (This is a *very* difficult problem to solve). Not only that, the model contains a dark matter candidate the neutralino  $\chi_1^0$ :

- Weakly interacting
- Stable
- Massive

To predict how much of it is around today, we must make assumptions on our *cosmology*, and on the *particle physics*. If we can measure the particle physics, we have a handle on the cosmology.

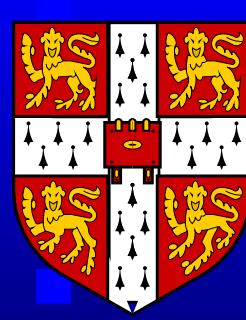


# Supersymmetric Copies

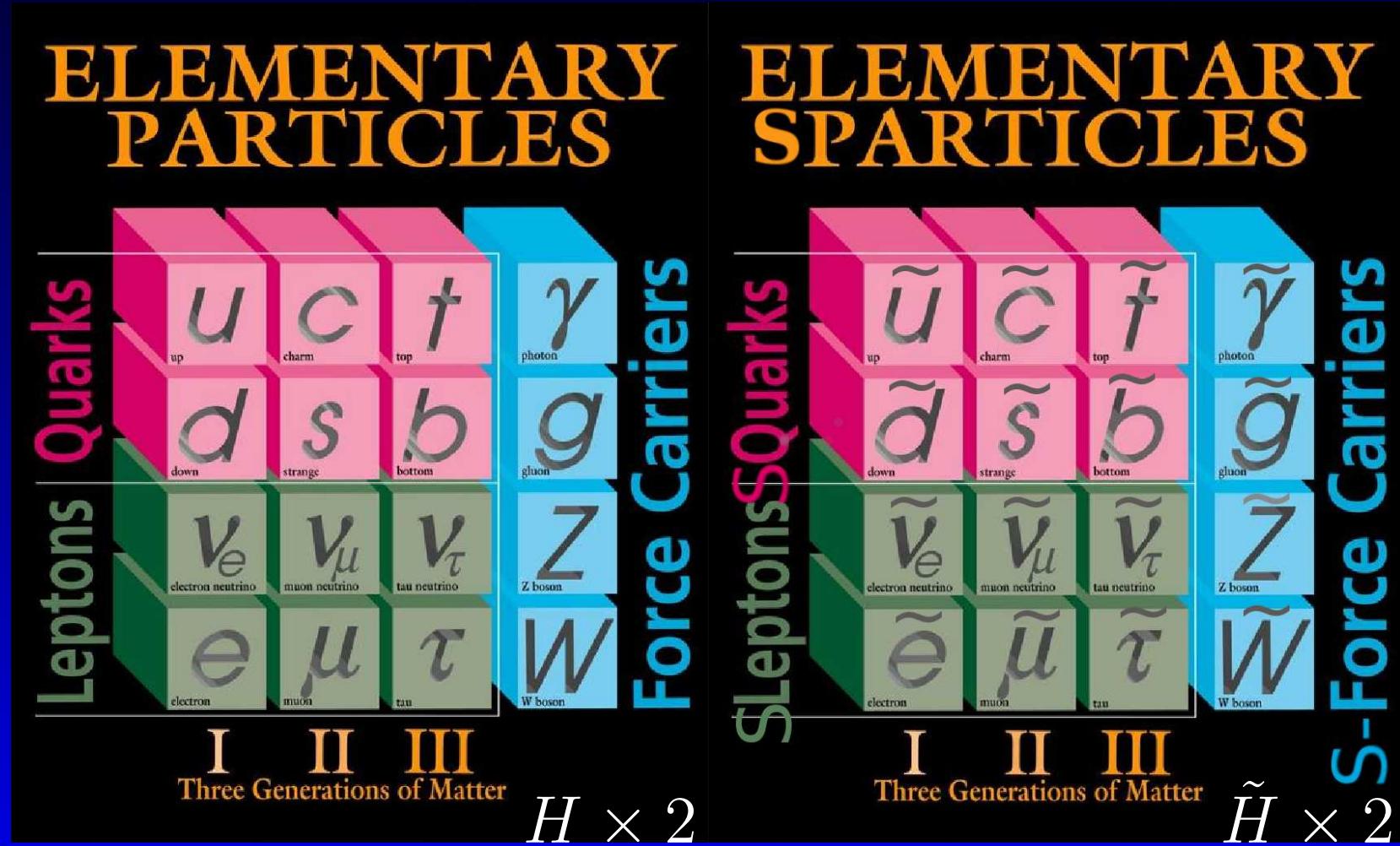


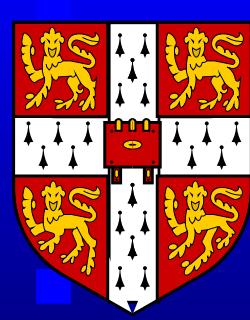
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# Supersymmetric Copies





# Review of R-Parity

The superpotential of the MSSM can be separated into two parts:

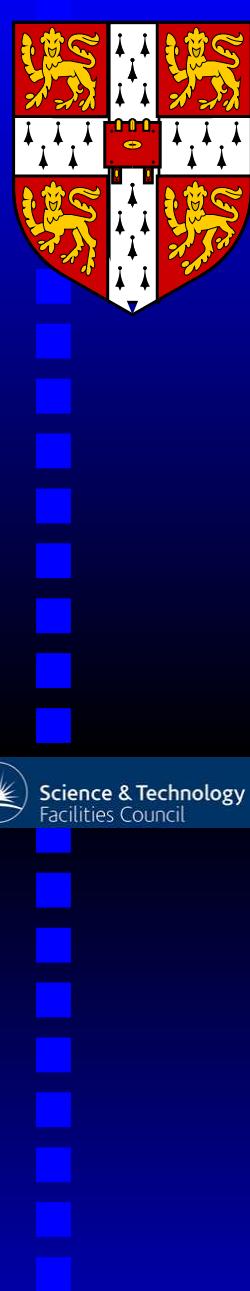
$$W_{R_p} = h_{ij}^e L_i H_1 \bar{E}_j + h_{ij}^d Q_i H_1 \bar{D}_j + h_{ij}^u Q_i H_2 \bar{U}_j + \mu H_1 H_2,$$

$$W_{R_P} = \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k + \kappa_i L_i H_2.$$

$W_{R_p}$  is what is usually meant by the MSSM.

$\mathcal{Q}$ : Why ban  $W_{R_P}$ ?

$\mathcal{A}$ : “Proton decay”



# Definition of R-Parity

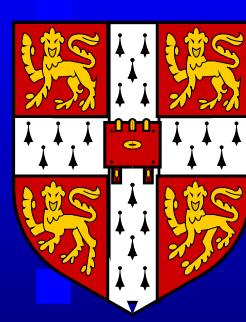
$\mathcal{Q}$ : How is  $W_{R_P}$  normally banned?  
 $\mathcal{A}$ : By defining discrete symmetry  $R_p$

$$R_p = (-1)^{3B+L+2S}.$$

→ SM fields have  $R_p = +1$  and superpartners have  $R_p = -1$ . There are two important consequences:

- Because initial states in colliders are  $R_p$  EVEN, we can only pair produce SUSY particles
- The *lightest superpartner is stable*

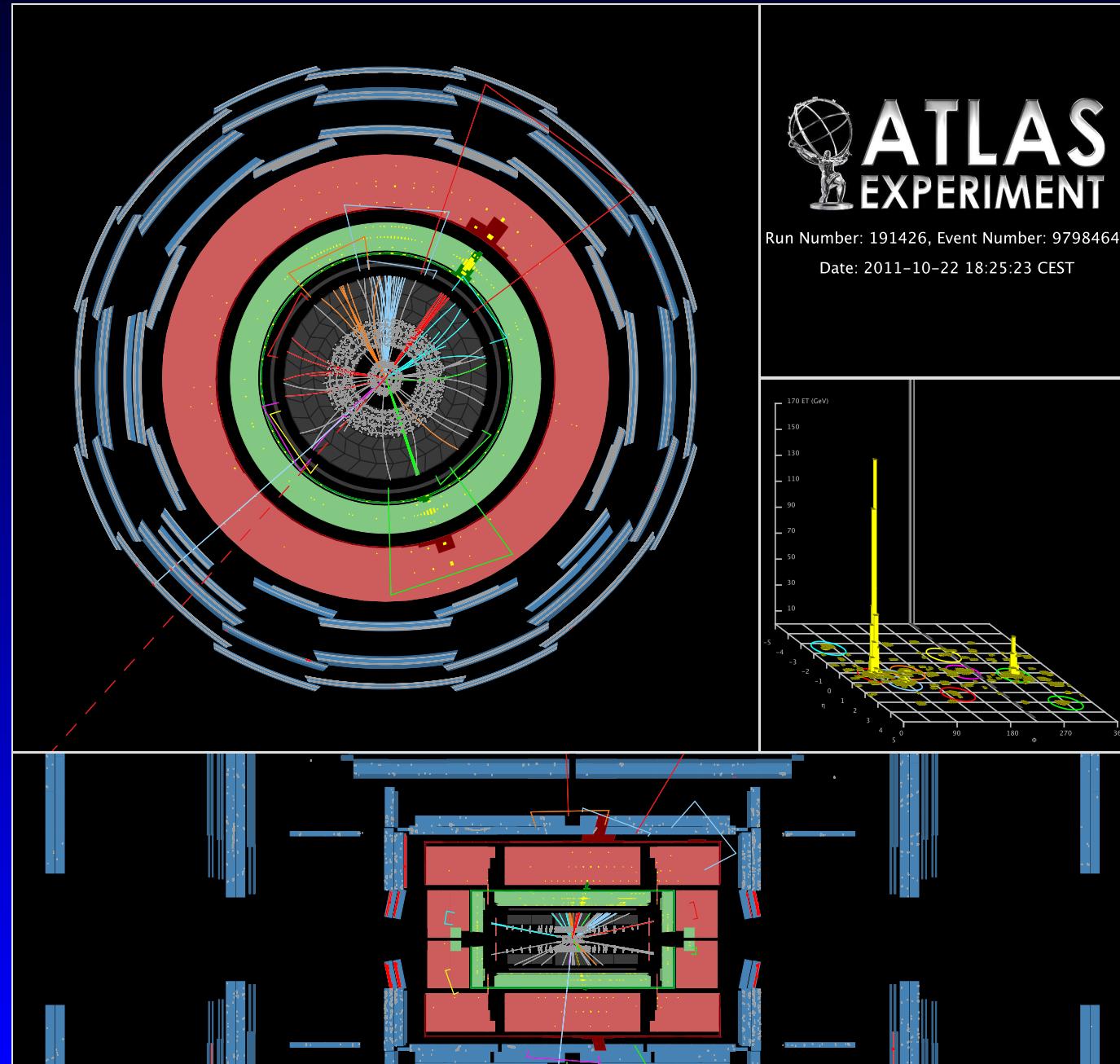


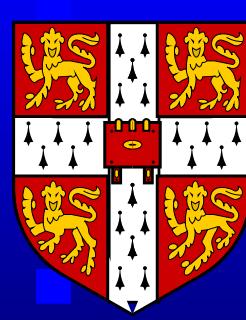


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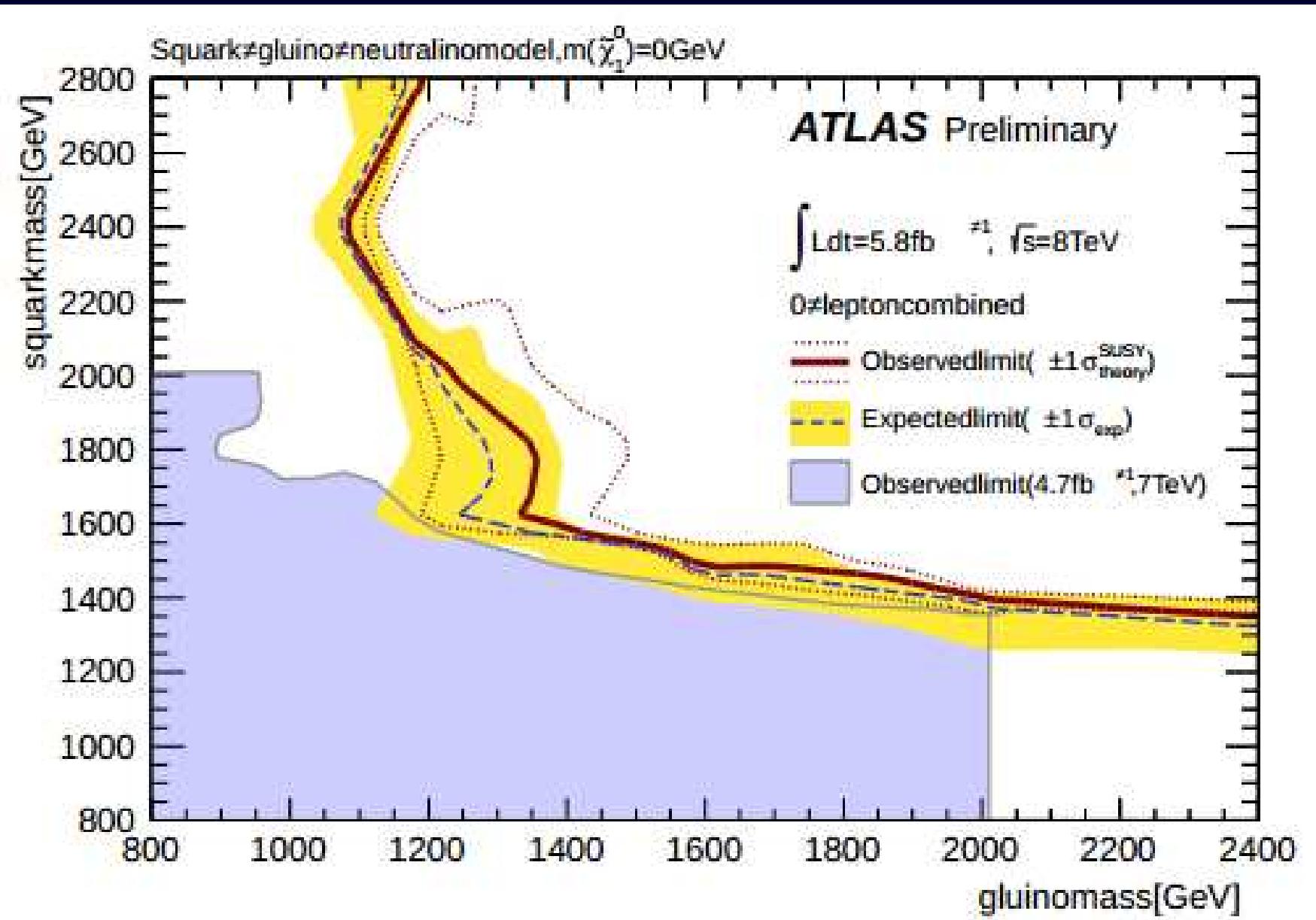
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# Candidate Event: High $E_T(j)$



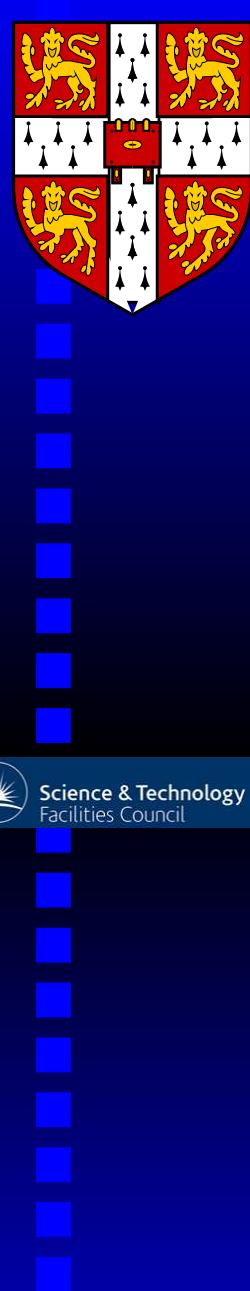


# Jets Plus $E_T$ Search



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# EWSB in the MSSM

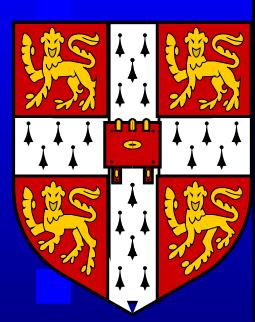
Minimisation of MSSM Higgs potential yields the constraint:

$$\mu^2 = \frac{\tan 2\beta}{2} [m_{H_2}^2 \tan \beta - m_{H_1}^2 \cot \beta] - \frac{M_Z^2}{2}.$$

For successful EWSB,  $\mu^2 > 0$ , i.e.

$$[m_{H_2}^2(M_{SUSY}) \tan \beta - m_{H_1}^2(M_{SUSY}) \cot \beta] < \frac{M_Z^2}{\tan 2\beta}.$$





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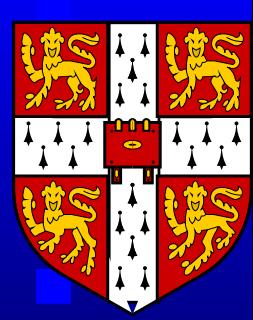


# RGEs of EWSB

$$16\pi^2 \frac{\partial m_{H_2}^2}{dt} = 6 \left[ (m_{H_2}^2 + m_{\tilde{Q}_3}^2 + m_{\tilde{u}_3}^2 + A_t^2) h_t^2 \right] - 6g_2^2 M_2^2 \\ - \frac{6}{5} g_1^2 M_1^2 + \frac{3}{5} g_1^2 (m_{H_2}^2 - m_{H_1}^2 + \text{Tr}[m_{\tilde{Q}}^2 - m_{\tilde{L}}^2 - 2m_{\tilde{u}}^2 + m_{\tilde{d}}^2 + m_{\tilde{e}}^2]) ,$$

where  $t = \ln Q$ .

$$h_t(M_Z) = \frac{\sqrt{2} m_t(M_Z)}{v(M_Z) \sin \beta(M_Z)},$$

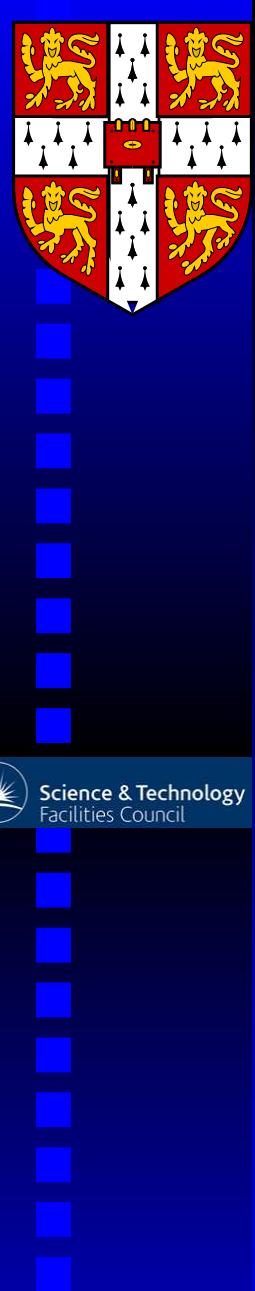


# Approximate Solution

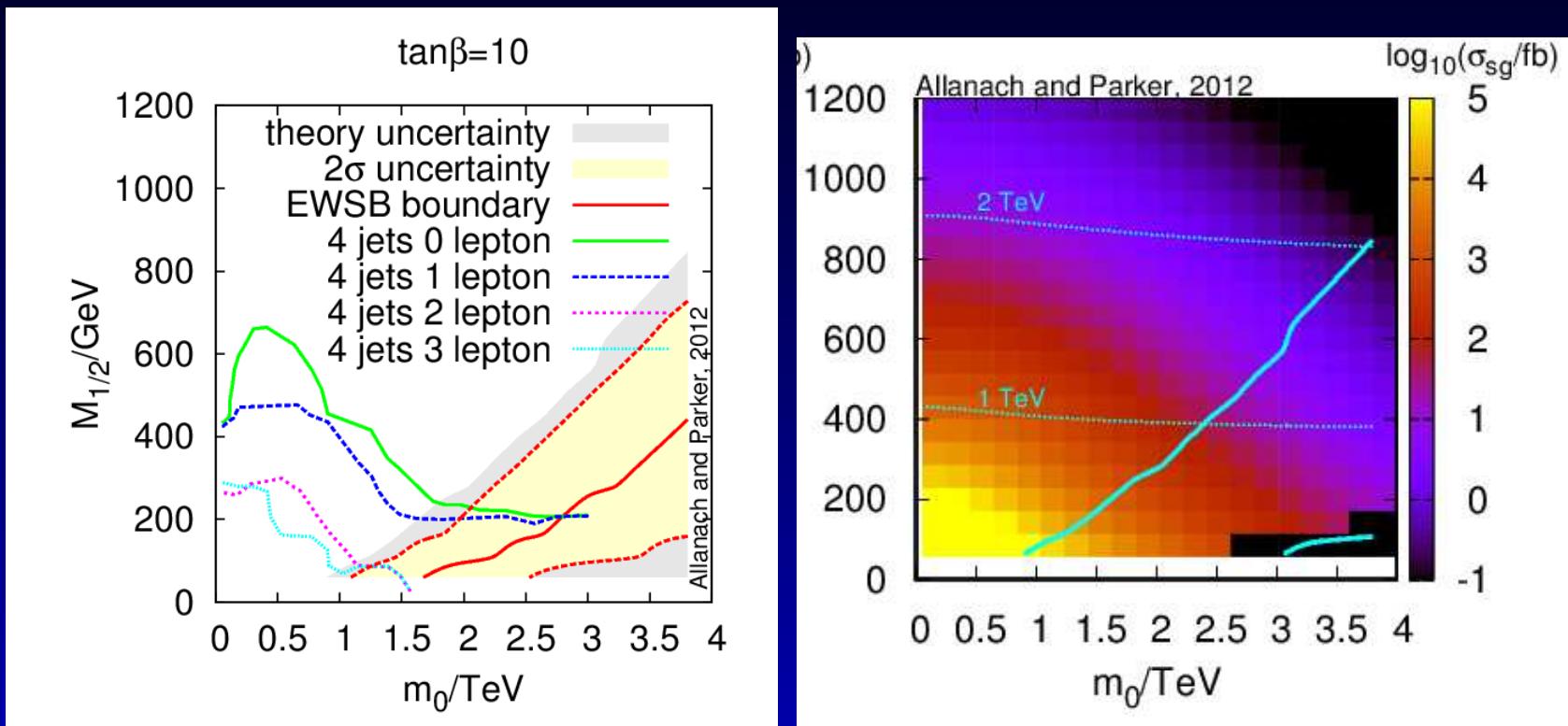
$$\begin{aligned} m_{H_2}^2(M_{SUSY}) = & \ m_{H_2}^2(M_X) - A \left( 6 \left[ m_{H_2}^2(M_X) \right. \right. \\ & + m_{\tilde{Q}_3}^2(M_X) + m_{\tilde{u}_3}^2(M_X) + \\ & \left. \left. A_t^2(M_X) \right] h_t^2(M_X) \right), \end{aligned}$$

where  $A = \log(M_X/M_{SUSY})/(16\pi^2)$ . Plug into our  $\mu^2 > 0$  condition, then at  $M_X$ :

$$\begin{aligned} m_{H_2}^2 < & \frac{1}{1 - 6Ah_t^2} \left\{ 6Ah_t^2 \left( m_{\tilde{Q}_3}^2 + m_{\tilde{d}_3}^2 + A_b^2 \right) + \frac{M_Z^2}{\tan \beta \tan \alpha} \right. \\ & \left. \frac{1}{\tan^2 \beta} \left[ m_{H_1}^2 - 6Ah_b^2(m_{\tilde{Q}_3}^2 + m_{\tilde{d}_3}^2 + A_b^2) \right] \right\}, \end{aligned}$$

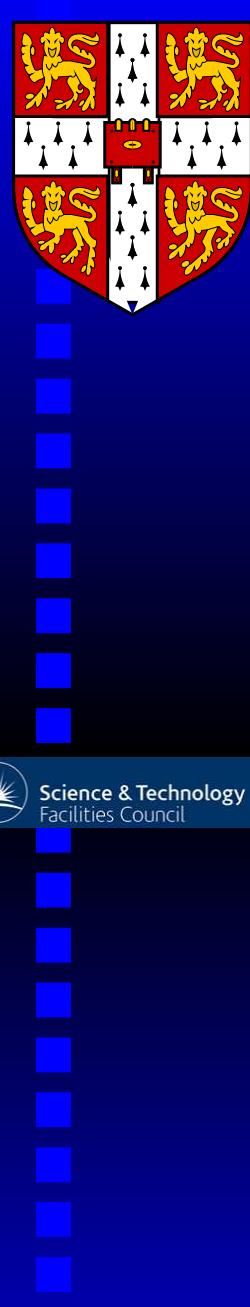


# EWSB in the CMSSM at 14 TeV, $1 \text{ fb}^{-1}, 5 \sigma$



The position of the EWSB boundary is extremely uncertain due to the error in  $m_t = 172.5 \pm 1 \text{ GeV}$ .  
*Experiments should use the higher value (plus theory error) when making grids of MSSM spectra<sup>a</sup>*

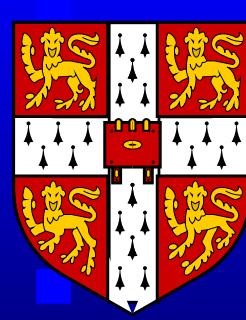
<sup>a</sup>BCA, Parker 1211.3231



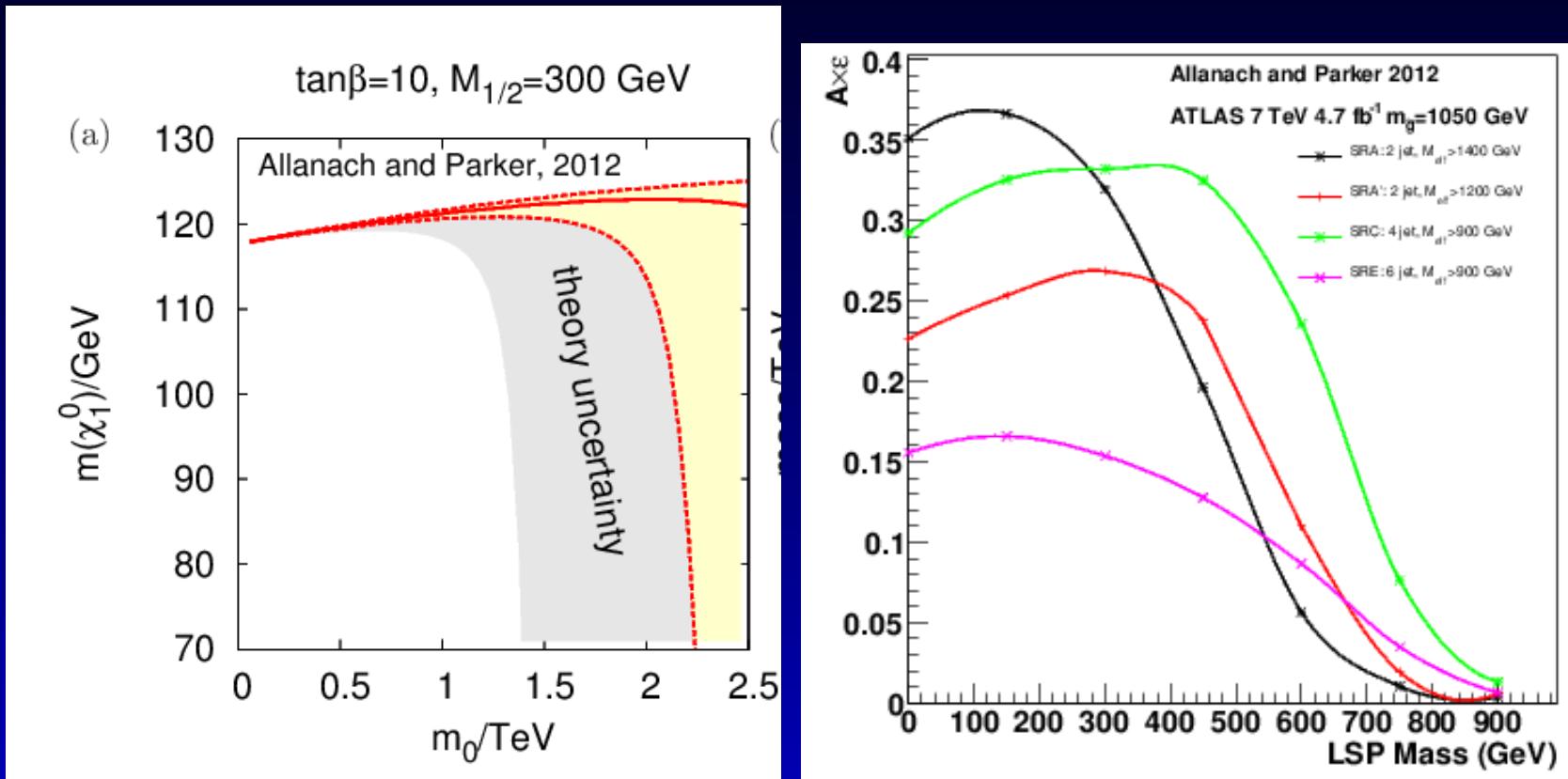
# Extreme Dependence on $m_t$

- This extreme dependence on  $m_t$  can happen in *models with high scale boundary conditions* - not just the CMSSM.
- It affects particle masses, and hence theoretical uncertainties etc.
- However, this uncertainty does not affect low-scale models such as the pMSSM (or simplified models).





# Neutralino Mass

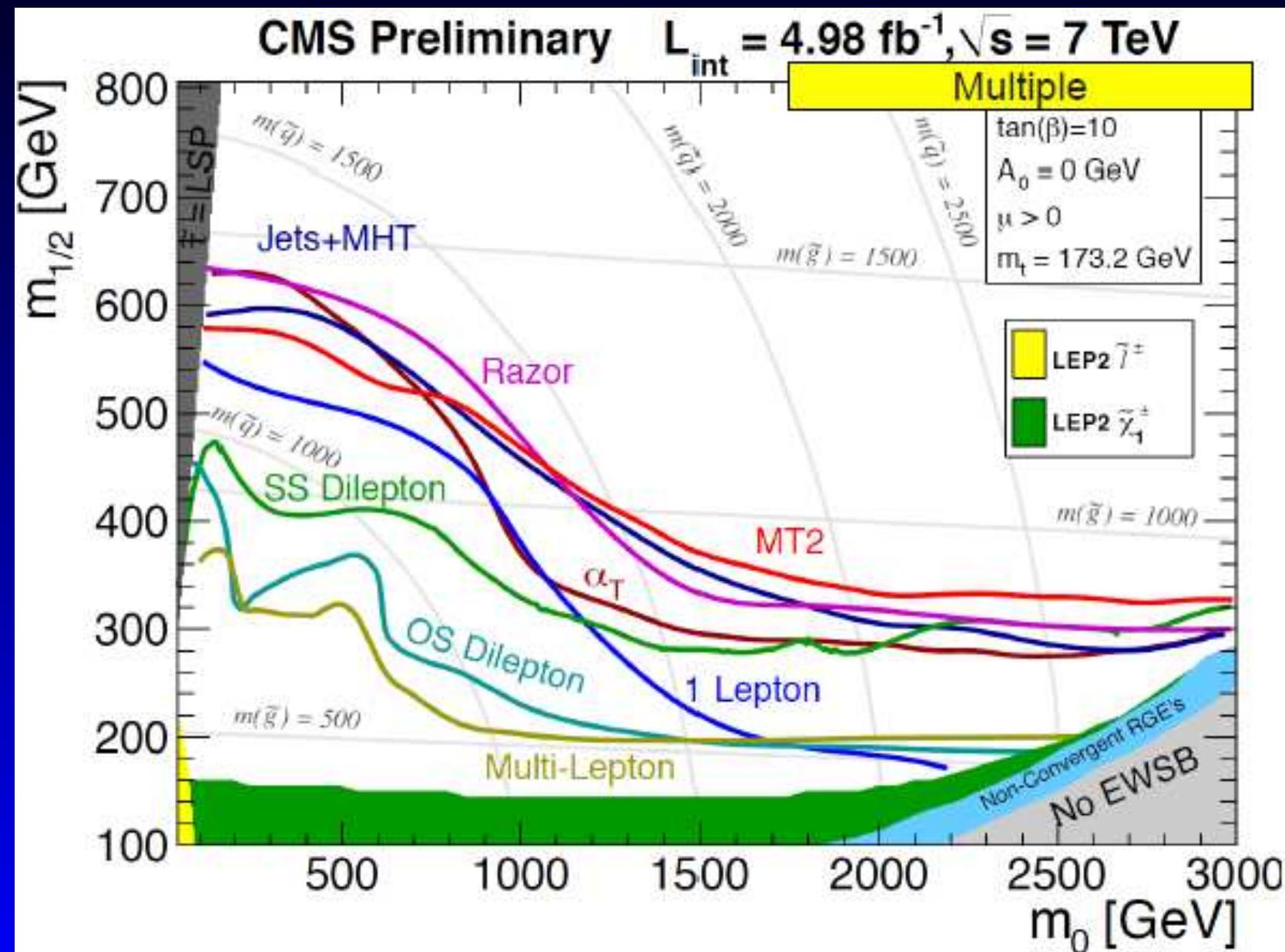


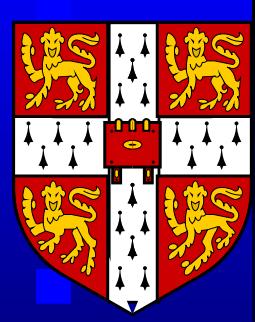
For a given value of  $m_t$ , on the boundary the LSP mass goes to 0. Therefore, for a given point in parameter space in the dodgy region, there is a 100% uncertainty in its mass  $\Rightarrow$ a theoretical error.





# Bad convergence?

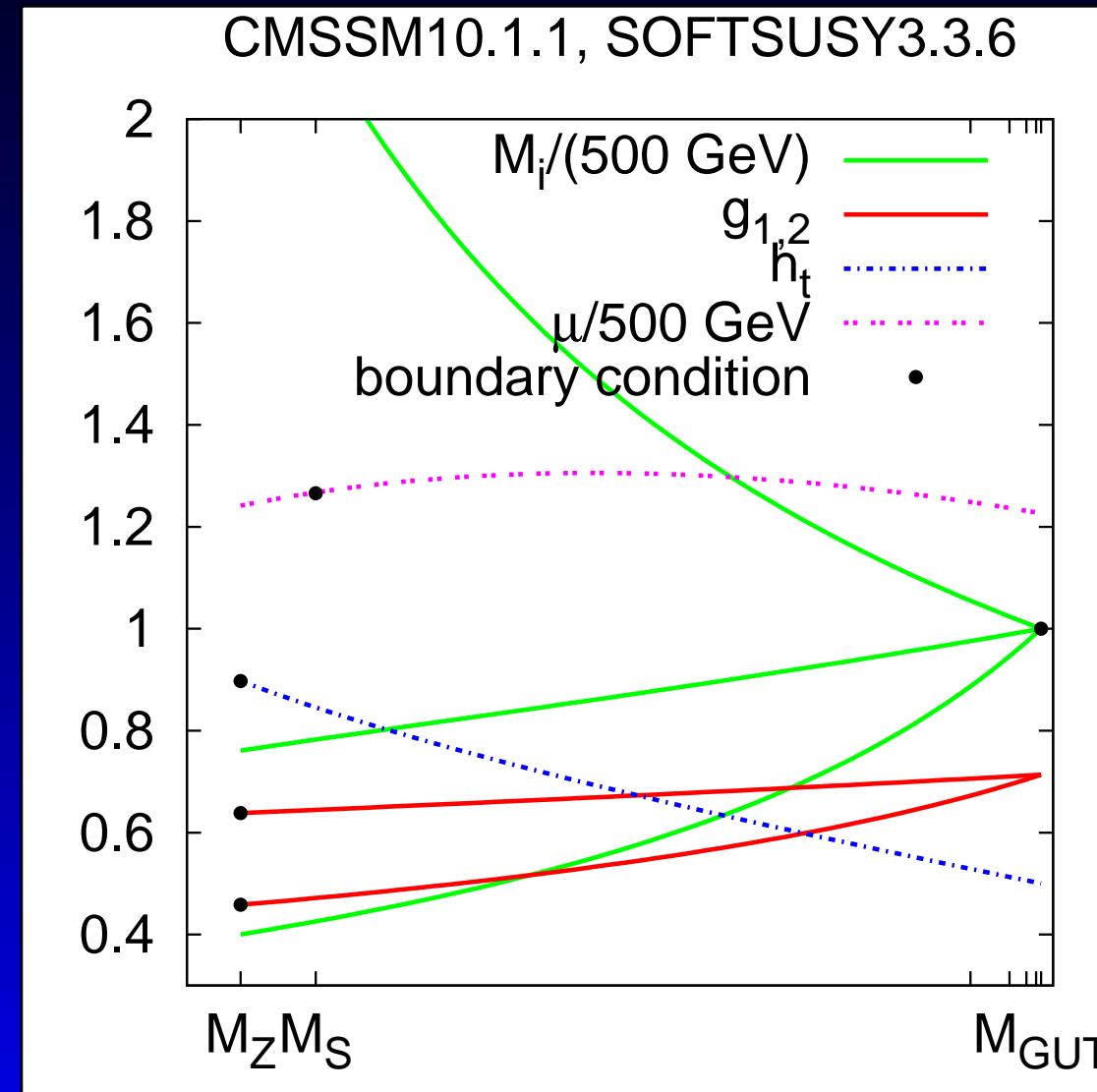


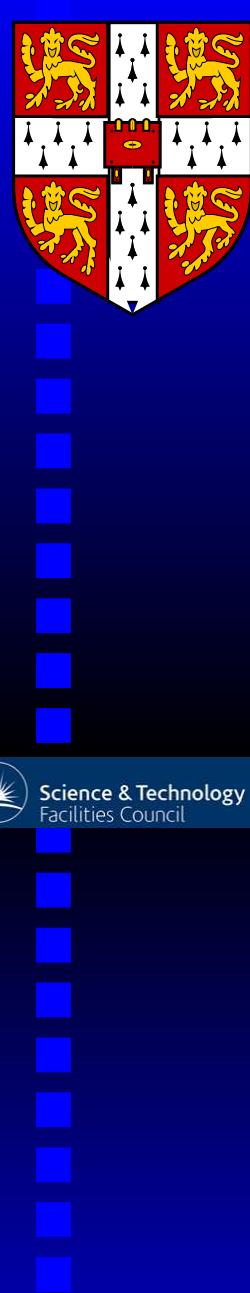


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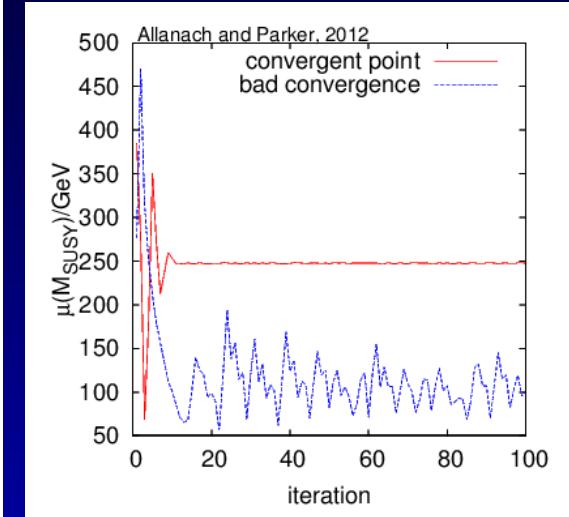
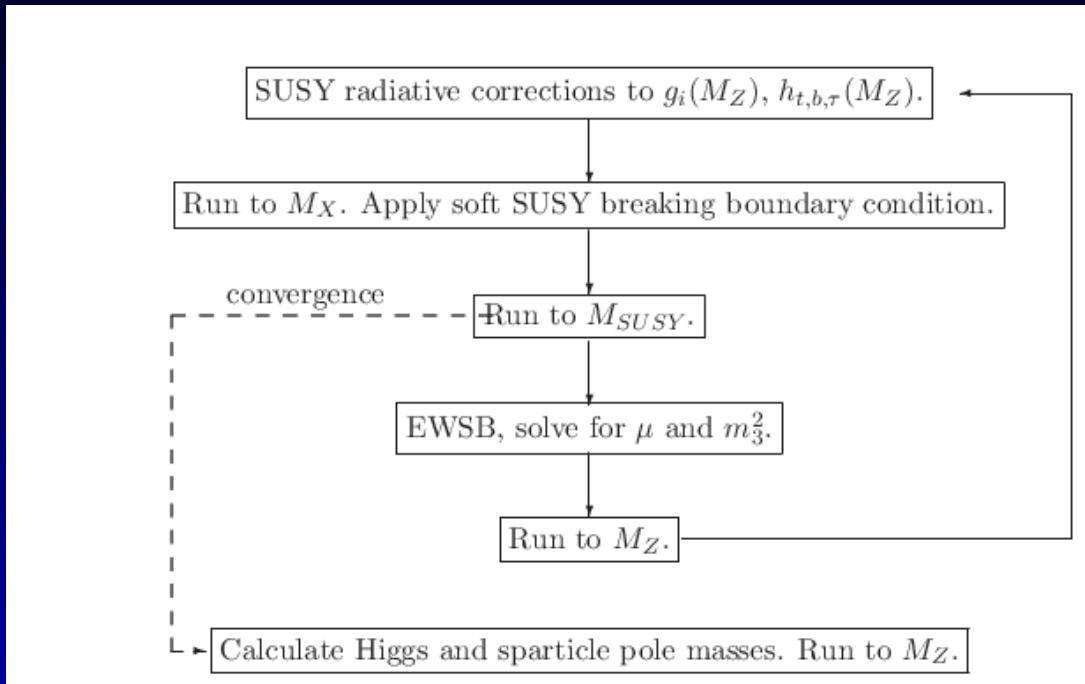
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# 2 Boundary Problem

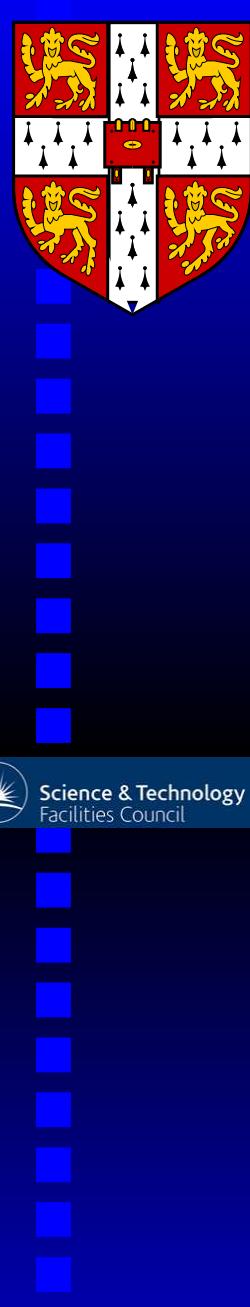




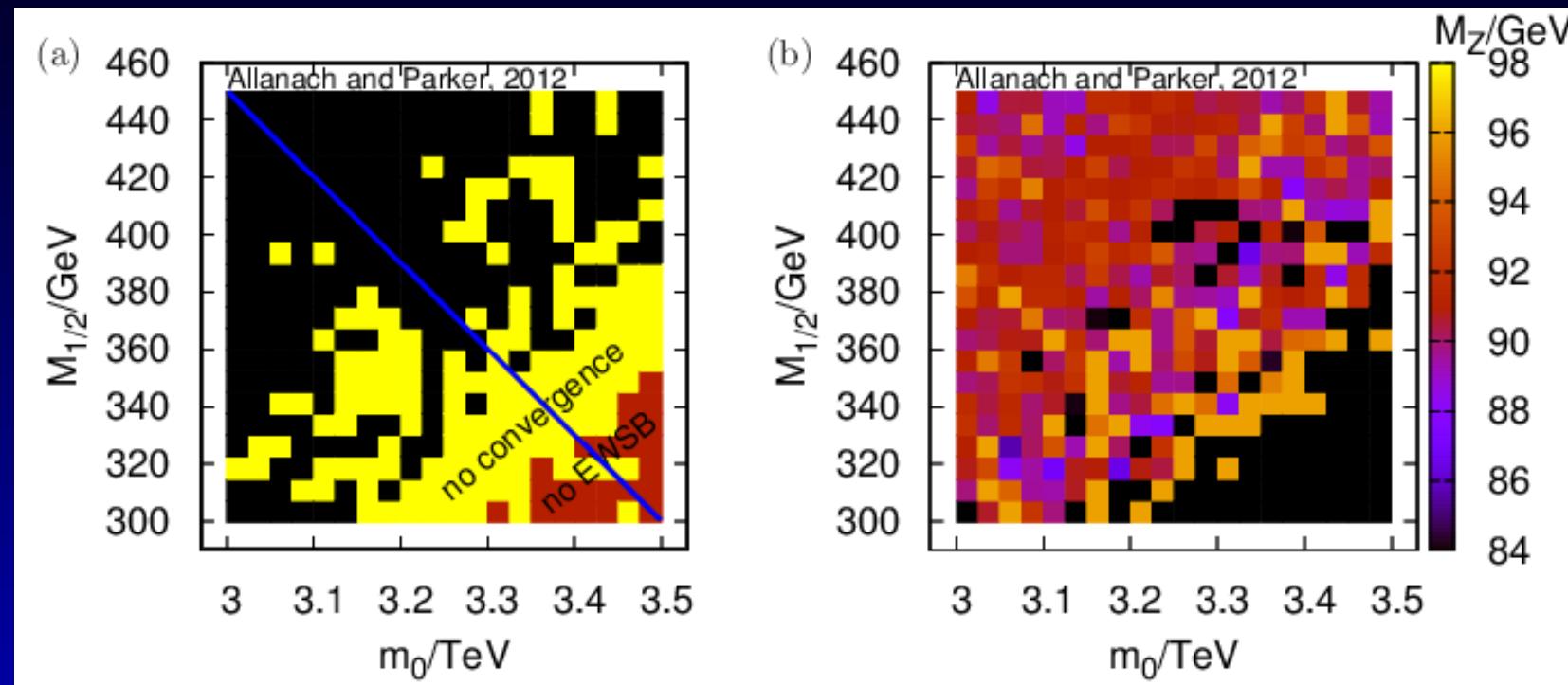
# Bad convergence

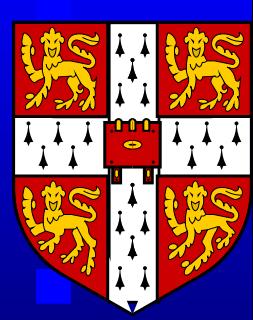


Previously, convergence criteria based on SUSY parameters converging iteration by iteration. But there is an absolute criterion:  $M_Z^2$ .



# Bad convergence region

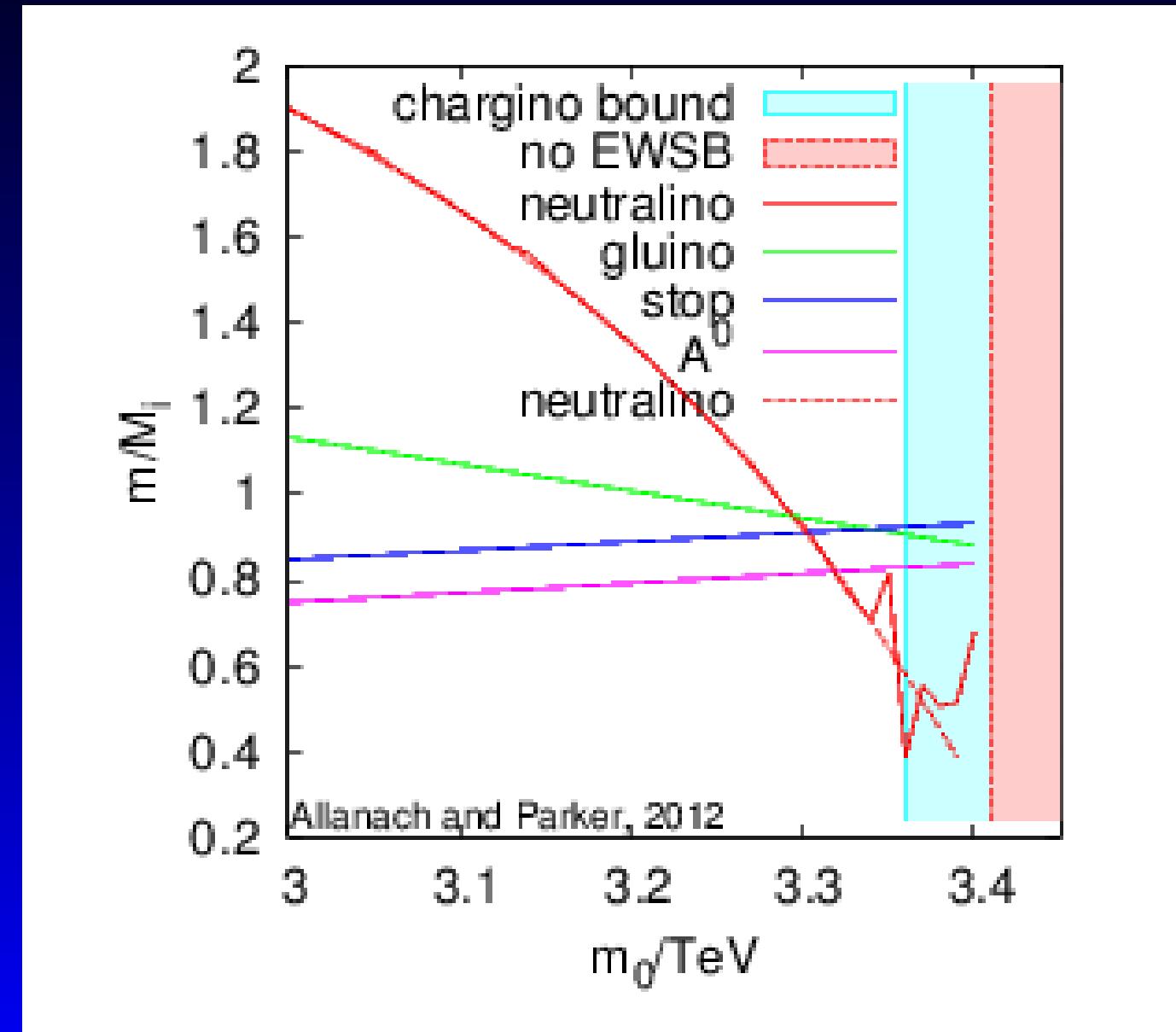


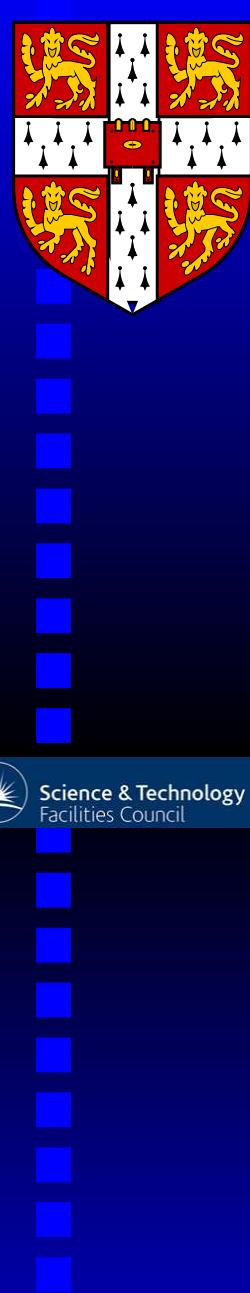


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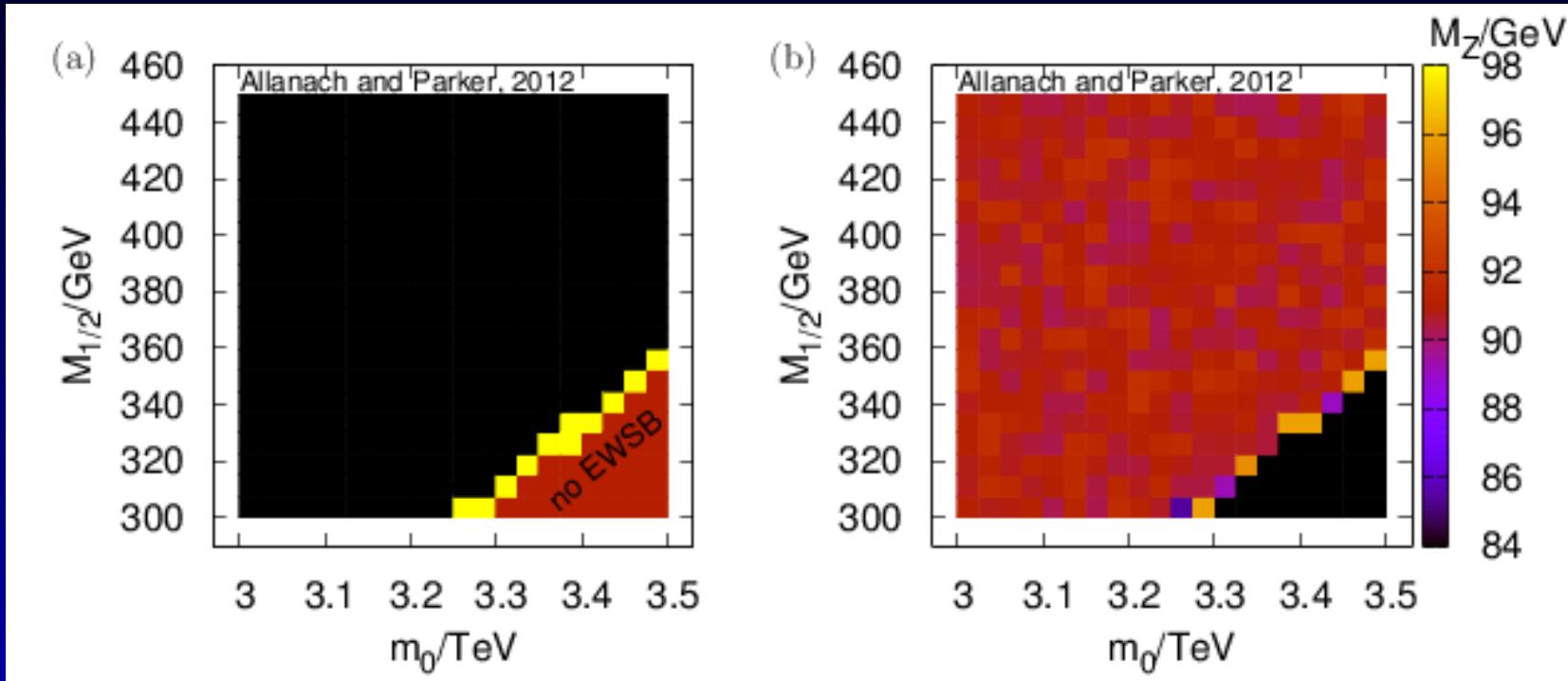
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# Masses On The Blue Line

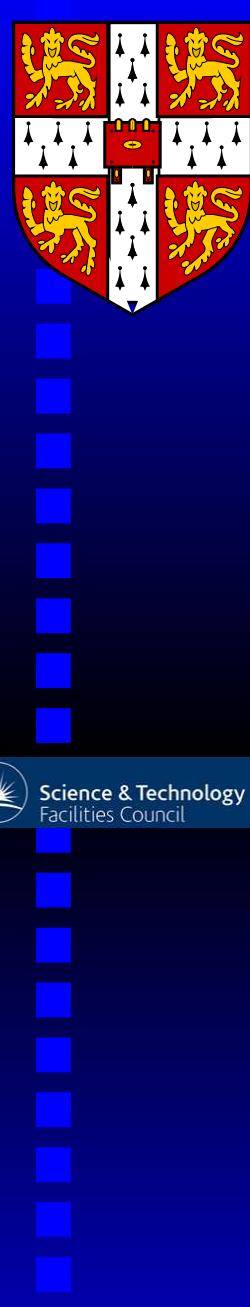




# Tweaked up convergence



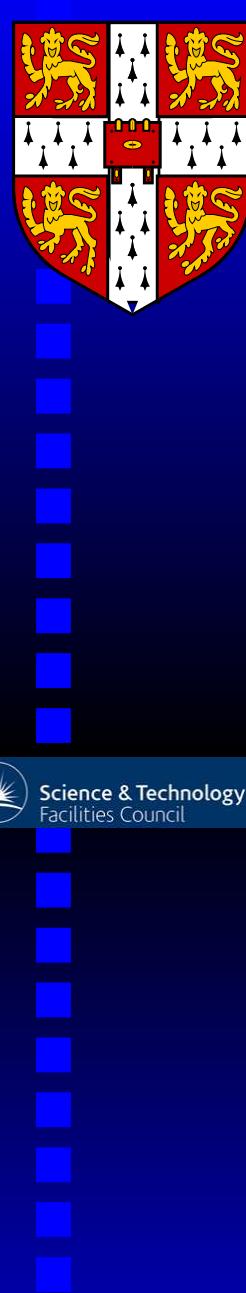
Iteration proceeds also until  $M_Z$  converges.  
Sub-iterations made more accurate, and iteration of  $\mu$  is damped.



# Recommendations for searches

- When devising grids of SUSY particle masses for exclusions, generate the masses at  $m_t + 2\sigma$ .
- Include a parameteric variation of  $m_t$  when estimating theory errors in SUSY in (or near) the uncertain region.
- Use the more recent version of SOFTSUSY with improved convergence criteria.



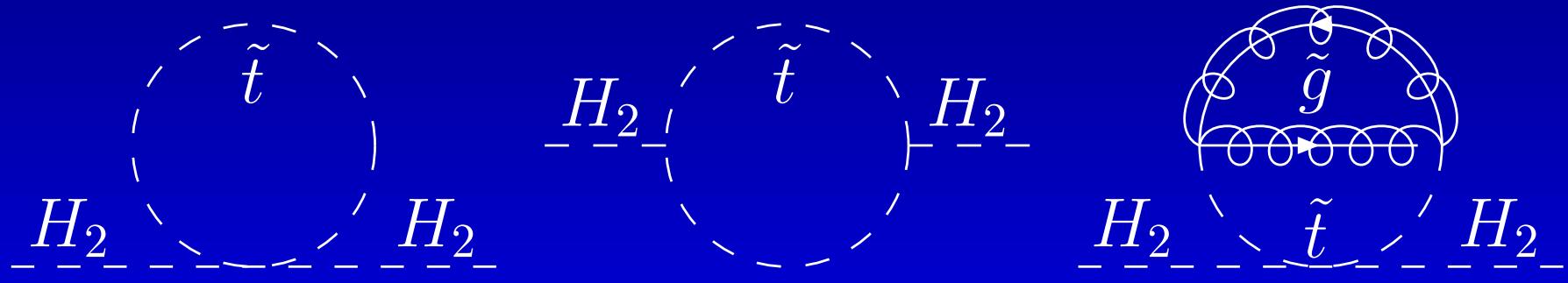


# Natural SUSY

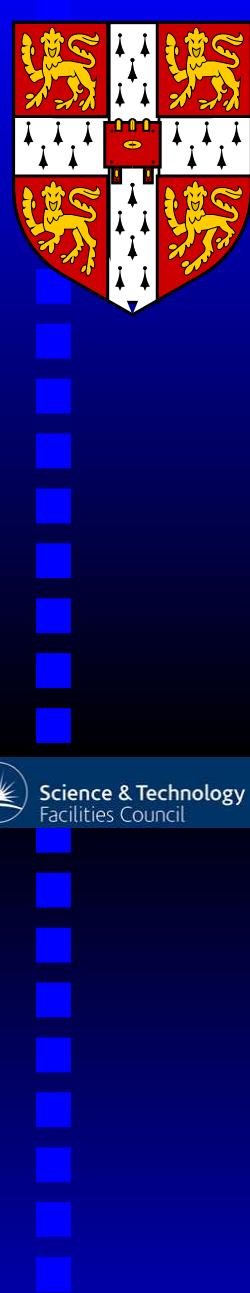
The particles coupling the most strongly to the higgs are the *stops*<sup>a</sup>. Minimising the MSSM Higgs potential,

$$-\frac{M_Z^2}{2} \approx |\mu|^2 + m_{H_2}^2,$$

$$\delta m_{H_2}^2 \approx \frac{-3h_t^2}{4\pi^2} m_{\tilde{t}}^2 \ln\left(\frac{\Lambda_{UV}}{m_{\tilde{t}}}\right)$$



<sup>a</sup>M. Papucci, J. T. Ruderman and A. Weiler, arXiv:1110.6926;  
C. Brust, A. Katz, S. Lawrence and R. Sundrum, arXiv:1110.6670



# Natural SUSY

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No cancellation  $\Rightarrow$

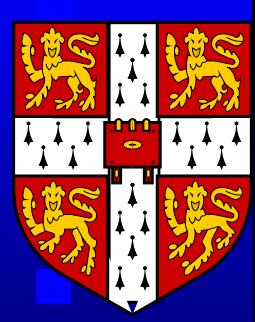
$$m_{\tilde{t}} \lesssim 700 \text{ GeV}, \quad m_{\tilde{g}} \lesssim 1000 \text{ GeV}.$$

Experimental  $E_T$  searches are probing this parameter space!

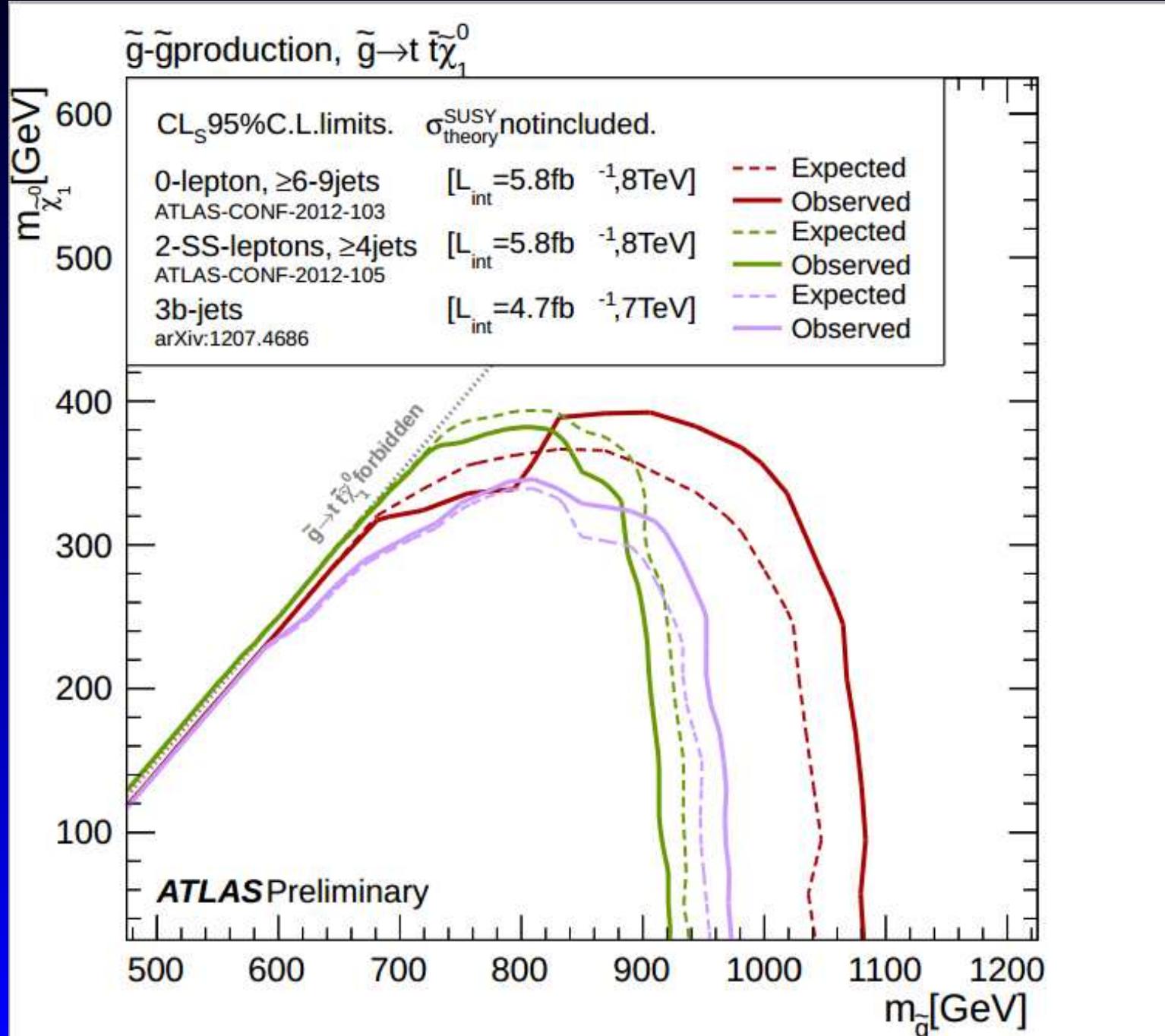
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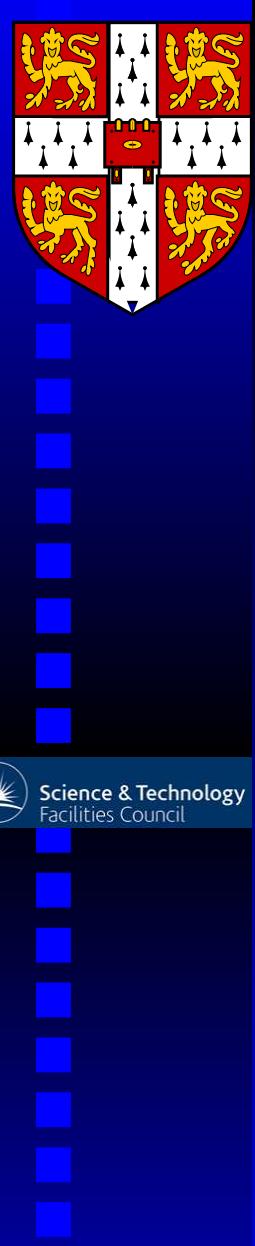
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C. Brust, A. Katz, S. Lawrence and R. Sundrum, arXiv:1110.6670





gt:  $E_T > 50/120 \text{ GeV}$ ,  $N_b \geq 2$ ,  
 $\#j \geq 2$ ,  $H_T > 320 \text{ GeV}$





# Bottom Up Implications of 2012 Data

Naturalness is under **pressure**. Ways to get around it:

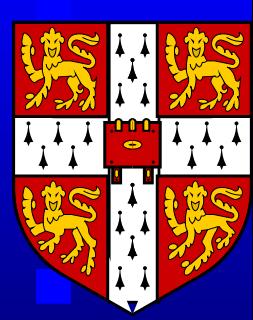
- First two generation squarks  $> 1.8 \text{ TeV}$ ,  
 $m_{\tilde{t}} = 0.5 - 0.7 \text{ TeV}$ ,  $m_{\tilde{g}} = 0.9 - 1 \text{ TeV}$ ,  
 $m_{\chi_1^0} < 0.5 \text{ TeV}$ . Ruled out soon?
- Compressed spectra<sup>a</sup>: ISR monojets/ $\cancel{E}_T$  searches give you  $m_{\tilde{g}} > 500 \text{ GeV}$  only.
- RPV decreases/removes the  $\cancel{E}_T$ <sup>b</sup>.
  - Explains why natural SUSY hasn't been found yet
  - Like-sign dileptons is a generic signature, as we'll see

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<sup>a</sup>Dreiner, Kramer, Tattersall arxiv:1207.1613

<sup>b</sup>BCA, Gripaios arXiv:1202.6616



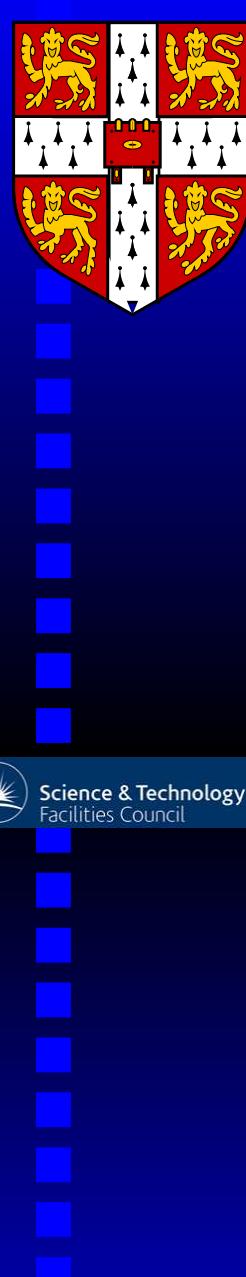


# RPV and Dark Matter

If one gives up  $R$ -parity,  $\chi_1^0$  is no longer a good dark matter candidate, since it **decays**. One then has to have something else, eg:

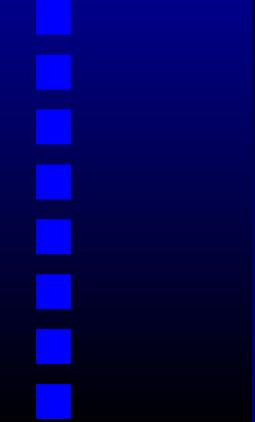
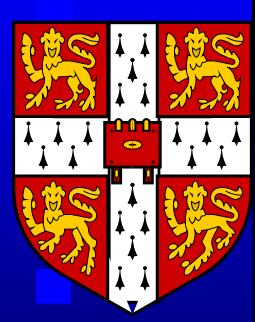
- *Gravitino* - still decays, but lifetime may be much longer than the age of the universe
- *Hidden sector matter*
- Axion/axino

The implications of each of these is that (in-)direct dark matter searches shouldn't find anything.



# Upper Bounds on $\lambda''_{ijk}$

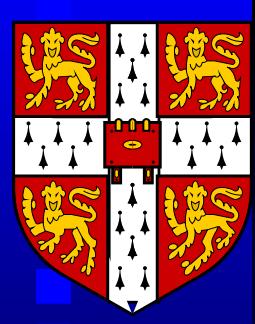
$\lambda''_{11k}$			$(10^{-8} - 10^{-7})(10^8 s/\tau_{osc}) \tilde{m}^{5/2}$ $[n\bar{n}]$ (6.128)
$\lambda''_{112}$			$10^{-6} [NN] (\tilde{m} = 300 \text{ GeV})$ (6.131) $6 \times 10^{-17} \tilde{s}_R^2 (m_{3/2}/1 \text{ eV})$ $[p \rightarrow K^+ \tilde{G}]$ (6.121) $8 \times 10^{-17} C_q^{-1} \tilde{s}_R^2 (F_a/10^{10} \text{ GeV})$ $[p \rightarrow K^+ \tilde{a}]$ (6.122)
$\lambda''_{113}$			$10^{-3} [NN] (\tilde{m} = 300 \text{ GeV})$ (6.131)
$\lambda''_{123}$			1.25 [RG]
$\lambda''_{212}$			1.25 [RG]
$\lambda''_{213}$			1.25 [RG]
$\lambda''_{223}$			1.25 [RG]
$\lambda''_{312}$		$1.45 [R_l] (6.41)$ $(\tilde{m} = 100 \text{ GeV})$	$4.28 [RG]$ $2.1 \times 10^{-3} [n\bar{n}]$ (6.129)
$\lambda''_{313}$		$1.46 [R_l] (6.41)$ $(\tilde{m} = 100 \text{ GeV})$	$1.12 [RG]$ $2.6 \times 10^{-3} [n\bar{n}]$ (6.129)
$\lambda''_{323}$		$1.46 [R_l] (6.41)$ $(\tilde{m} = 100 \text{ GeV})$	$1.12 [RG]$
$\lambda''_{ijk}$			$(10^{-11} \tilde{m}^3 - 10^{-8} \tilde{m}^2)$ $\times (m_{3/2}/1 \text{ eV}) [p \rightarrow K^+ \tilde{G}]$ (6.123) $\times (F_a/10^{10} \text{ GeV}) [p \rightarrow K^+ \tilde{a}]$ (6.124)



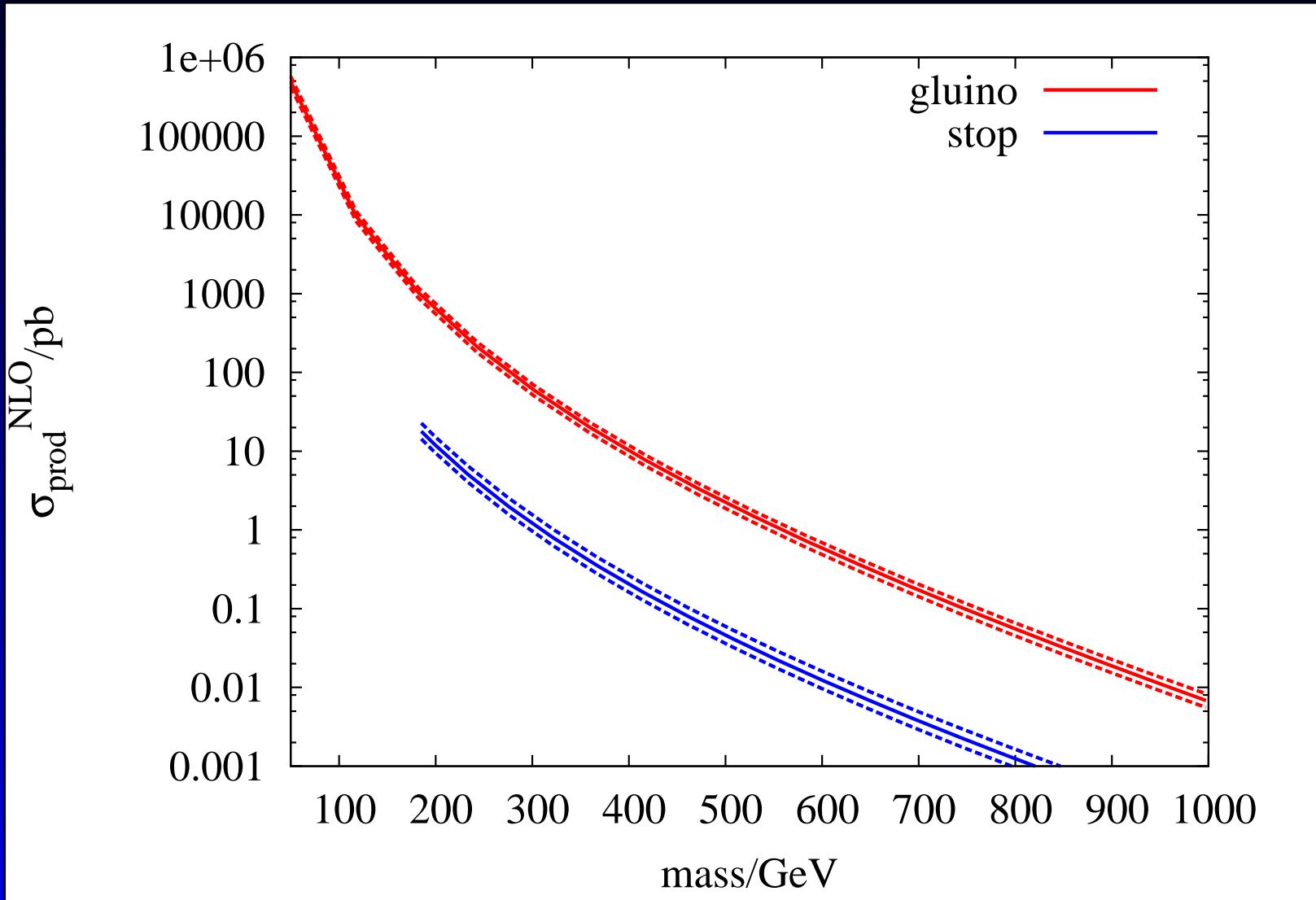
# Bottom up thinking

We are **not** assuming

- MSSM: don't put higgs constraints - there might be higher dimension operators coming from heavier fields.
- Some specific (or even generic) string model.
- We only put in the particles important for our analysis. Others may be decoupled, or light: it shouldn't matter.

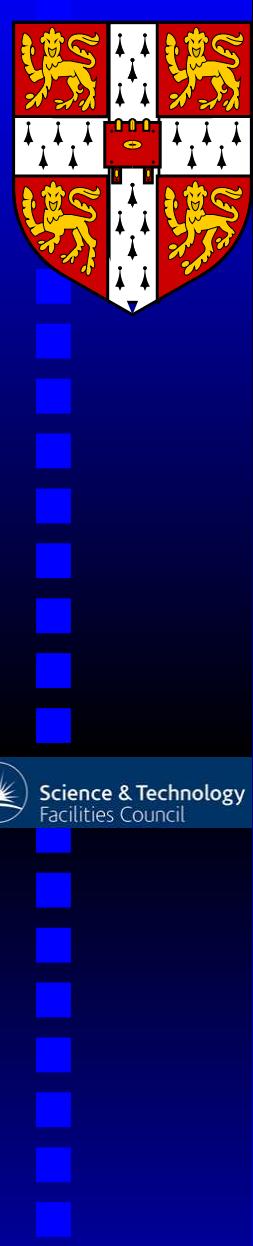


# Gluino/stop production at LHC7



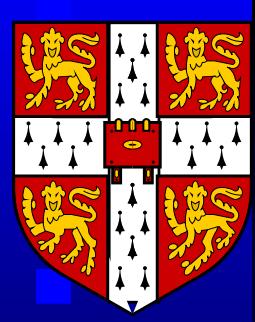
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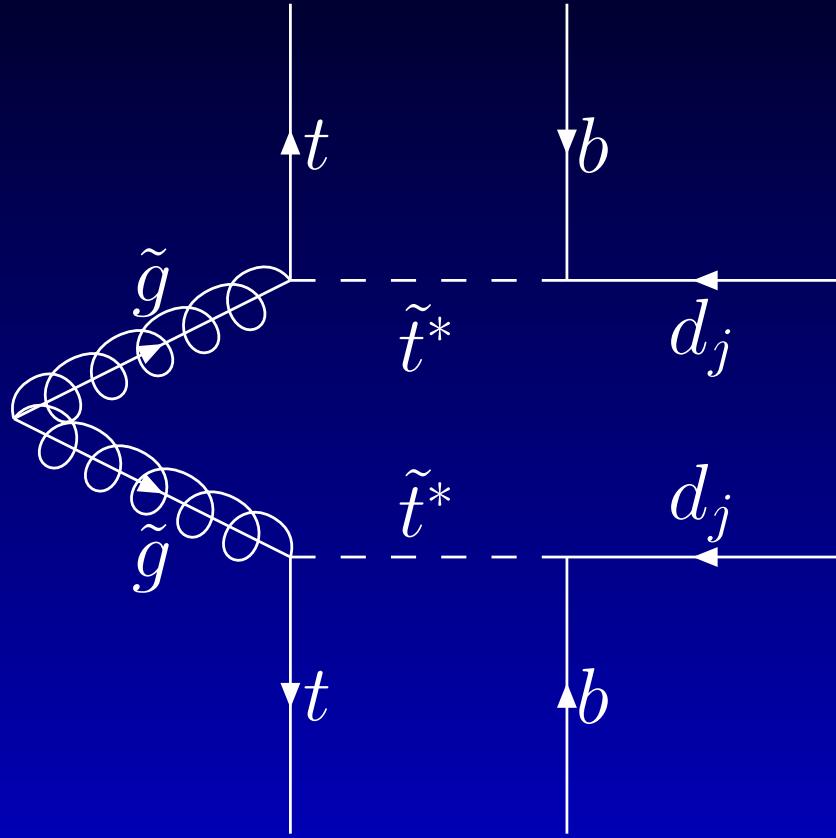


# Gluinos With $R_p$ Violation

- We assume lightish  $\tilde{g}$ ,  $\tilde{t}_R$ . If one has lepton number violating  $LLE$  or  $LH_1$  operators, the gluinos decay producing various leptons. These cases ought to be easy to find, and are good candidates for searches. **Get same-sign leptons.**
- With  $LQD$  operators, the stops will again decay into leptons, easy to see. Flavour constraints imply that  $L_3 QD$  operators are likely to be the largest. **Get same-sign leptons** in  $\sim \frac{7}{9}$  of  $\tilde{g}\tilde{g}$  events.
- With  $UDD$  operators, the (right-handed) top decays directly into jets.



# Baryon Number violating Example

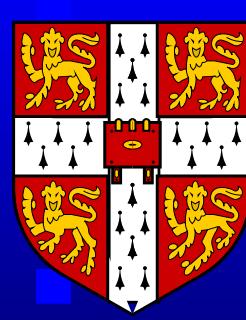


$\tilde{g}\tilde{g}$  production dominates. Here, you can look for like-sign di-leptons since gluinos decay into  $t$  and  $\bar{t}$  with equal branching ratios.

Can lead to natural SUSY with light stops and gluinos that hasn't been excluded yet. *A difficult case<sup>a</sup>:*  $W \supset \lambda''_{ijk} U_i D_j D_k$ .

---

<sup>a</sup>BCA and Ben Gripaios,  
arXiv:1202.6616



# Other Light States

How robust is the **same-sign dilepton** signature in the case that other states are also light?

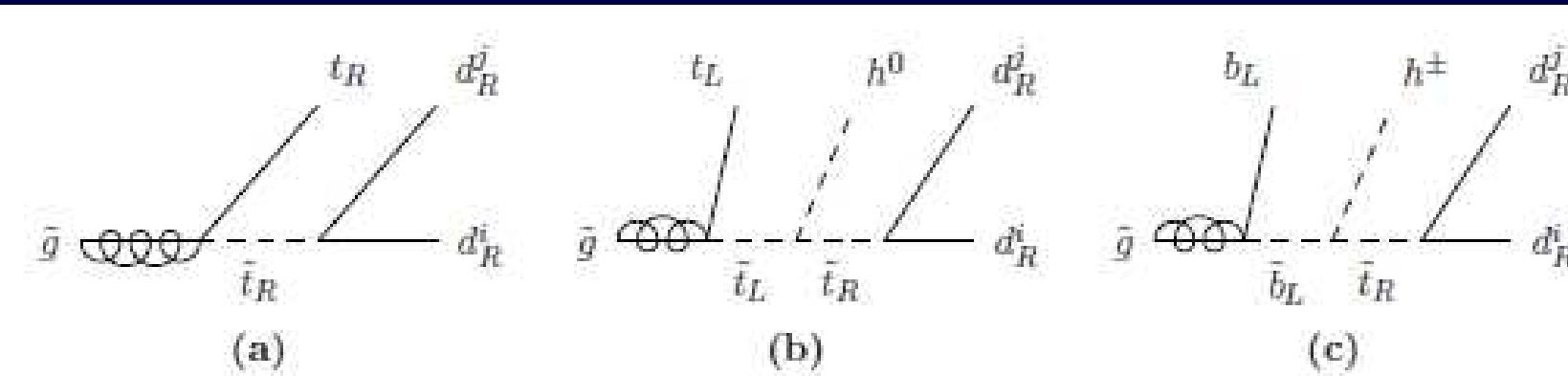


Figure 1. Gluino decays without right-handed bottom squarks in the presence of  $W \supset U_3 D_i D_j$ , via (a) right-handed top, (b) left-handed top, and (c) left-handed bottom. Same sign leptons are obtained in (c) only if the charged Higgs subsequently decays to  $t\bar{b}$  or to leptons.

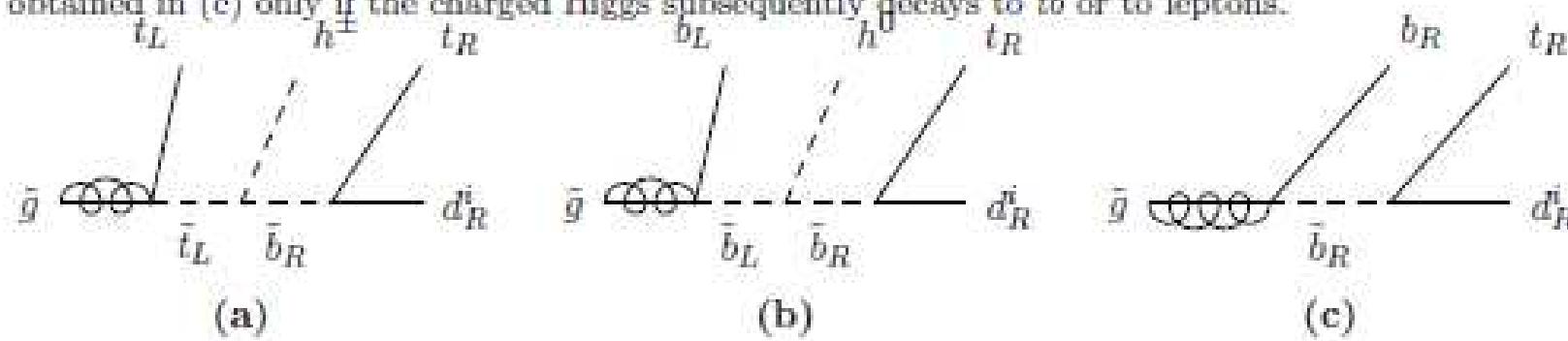
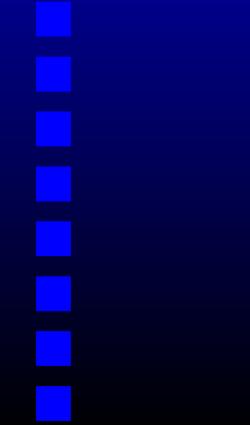
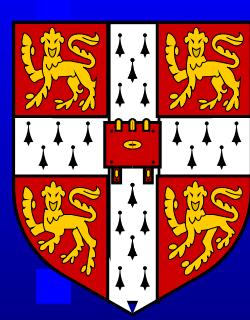
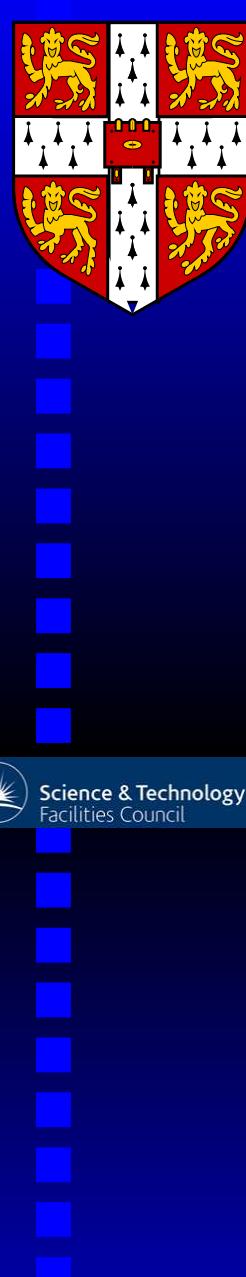


Figure 2. Gluino decays with right-handed bottom squarks in the presence of  $W \supset U_3 D_i D_3$ , via (a) left-handed top squark, (b) left-handed bottom squark, and (c) right-handed bottom squark.

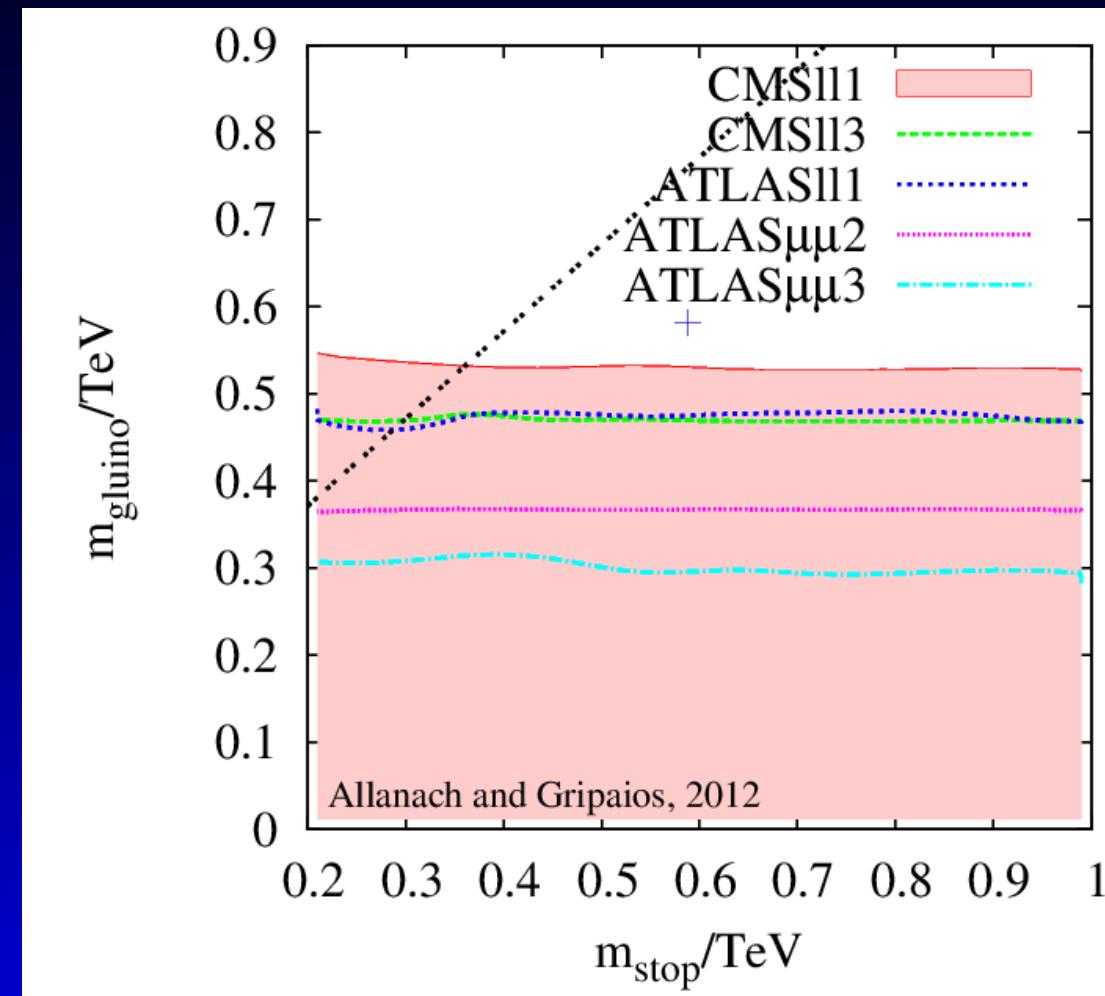


# Can we avoid SS dileptons?

- For  $m_{H^\pm} > m_t + m_b$ ,  $H^+ \rightarrow t\bar{b}$  dominates, which again will yield same-sign dileptons.
- $H^+ \rightarrow \tau^+\nu_\tau$  is also OK, since we'll get like-sign di-taus.
- Only fly in the ointment comes from Fig. 1c: when  $H^+ \rightarrow c\bar{b}$  (but only happens when  $\tan\beta \ll m_t V_{cb}/m_\tau \sim 3$ , which seems unlikely).

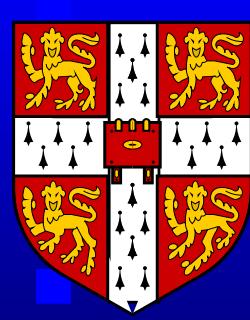


# Same-sign $E_T$ Limits

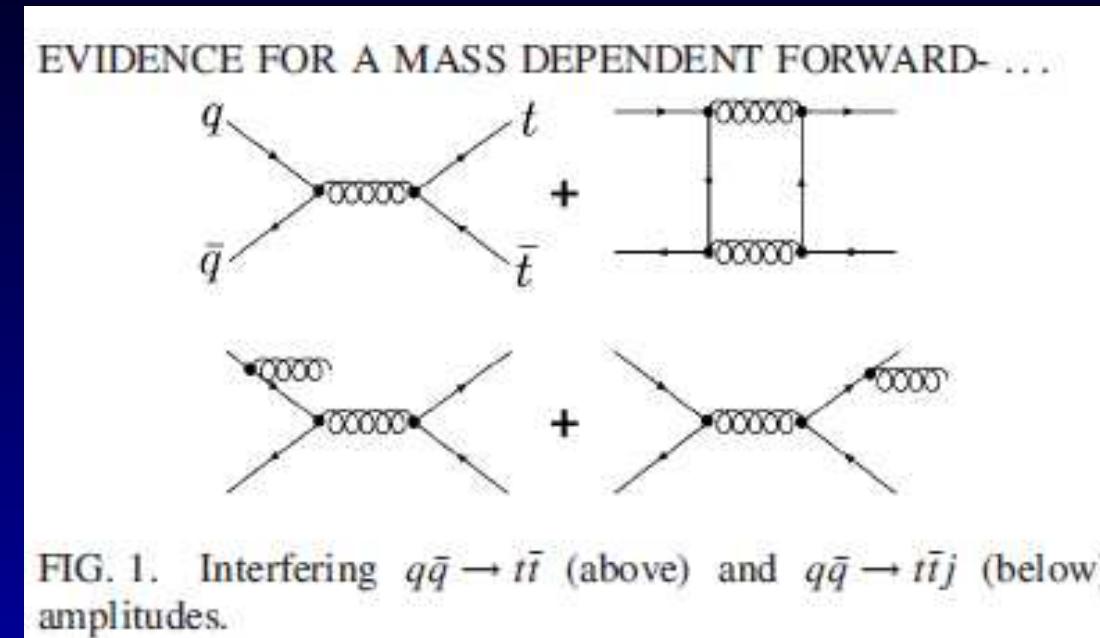


CMSSll1<sup>a</sup>:  $H_T > 400 \text{ GeV}, |E_T| > 120 \text{ GeV}$

<sup>a</sup>CMS-PAS-SUS-010; ATLAS-CONF-2012-004



# $A_{FB}$ in the Standard Model

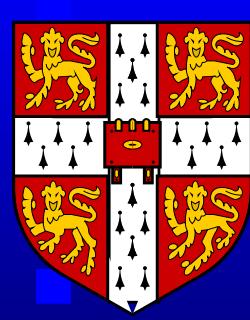


1.96 TeV  $p\bar{p}$  collisions at the Tevatron.

$$A_{FB} = \frac{N(\Delta y > 0) - N(\Delta y < 0)}{N(\Delta y > 0) + N(\Delta y < 0)}, c = \cos \theta.$$

NLO QCD+NLNLL SM Prediction<sup>a</sup>:  $7.2 \pm 1.0\%$ .

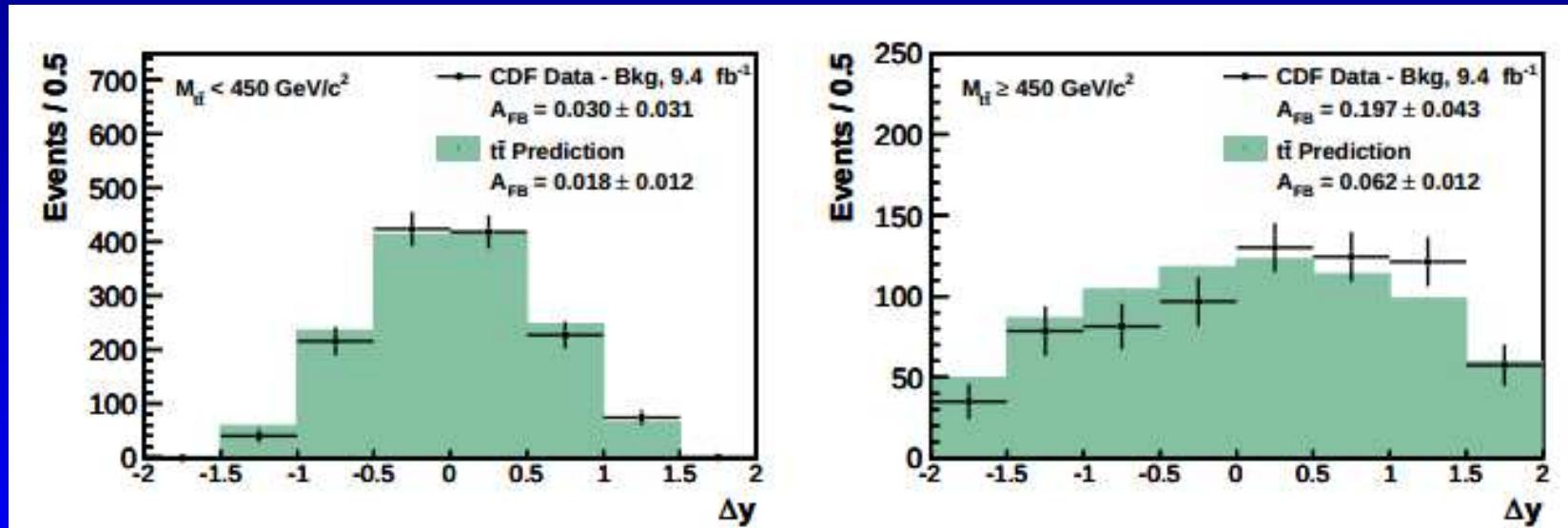
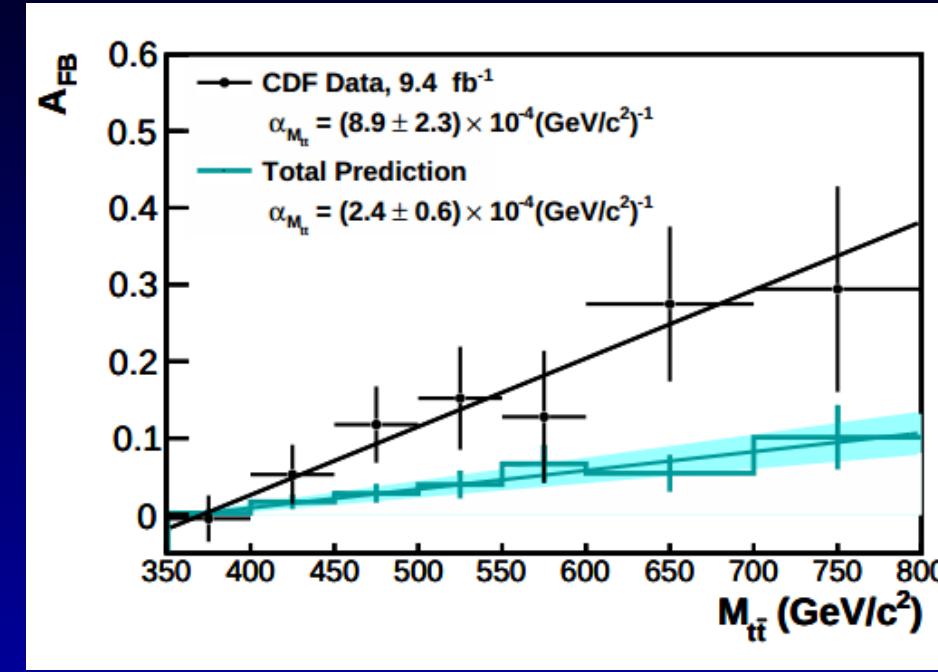
<sup>a</sup>Ahrens et al, PRD84 (2011) 074004

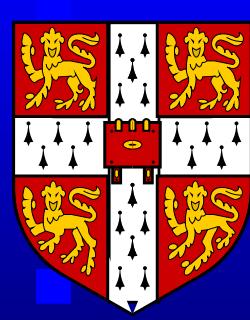


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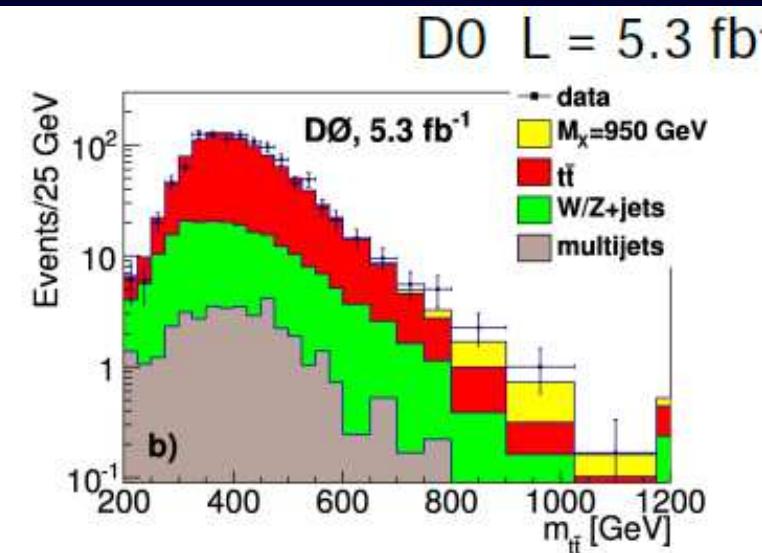
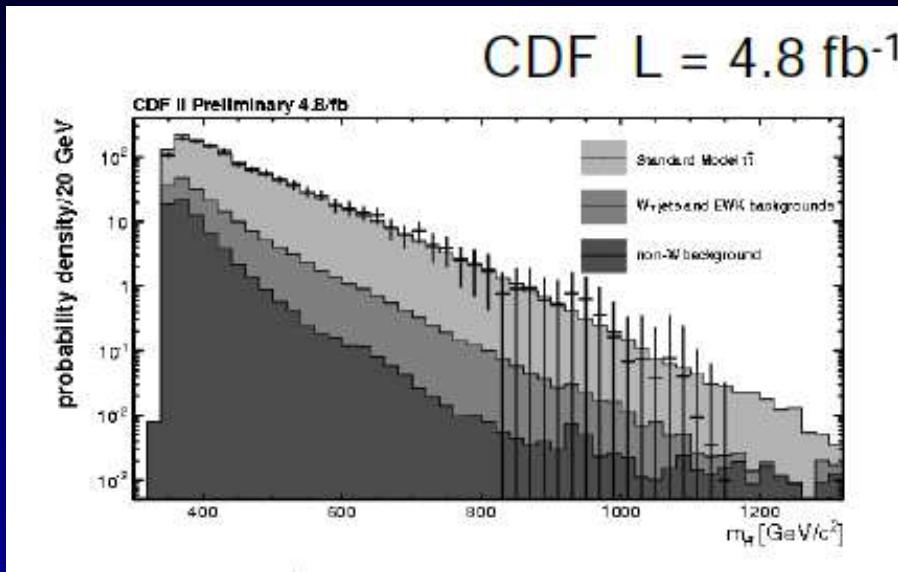
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# Measurements of $A_{FB}$





# Other Constraints

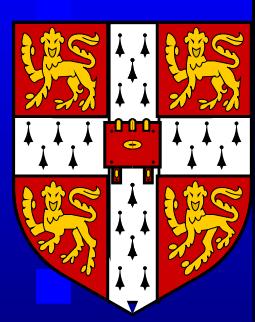


$$\sigma_{t\bar{t}}^{LHC7} = 173.4 \pm 10.6 \text{ pb},$$

$$\sigma_{t\bar{t}}^{SM} = 163 \pm 10 \text{ pb}$$

$$A_C^y = \frac{N(|y_t| > |y_{\bar{t}}|) - N(|y_{\bar{t}}| > |y_t|)}{N(|y_t| > |y_{\bar{t}}|) + N(|y_{\bar{t}}| > |y_t|)} = -0.015 \pm 0.04,$$

where  $y_i = 1/2 \ln(E_i - p_{iz})/(E_i + p_{iz})$  is the rapidity of particle  $i$ .  $A_C^{y,SM} = 0.006 \pm 0.002$ .



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# $\Delta A_{FB}$

Many models of random particles have been proposed, but most don't fit all the data. However, this one does:

$$W = \frac{\lambda''_{313}}{2} t_R d_R b_R$$

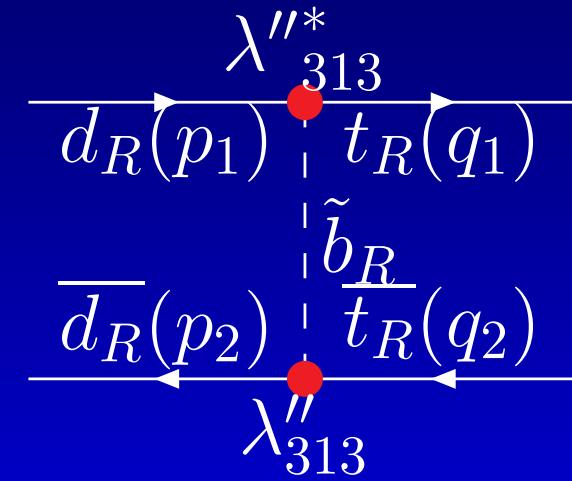
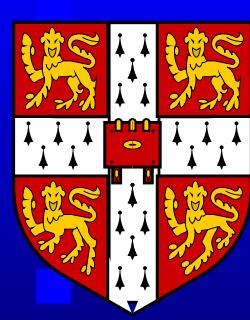


Figure 1: SUSY contribution to  $A_{FB}$ <sup>a</sup>

<sup>a</sup>BCA, Sridhar arXiv:1205.5170



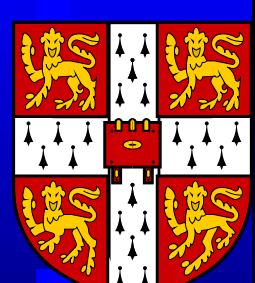
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# New Physics Contribution

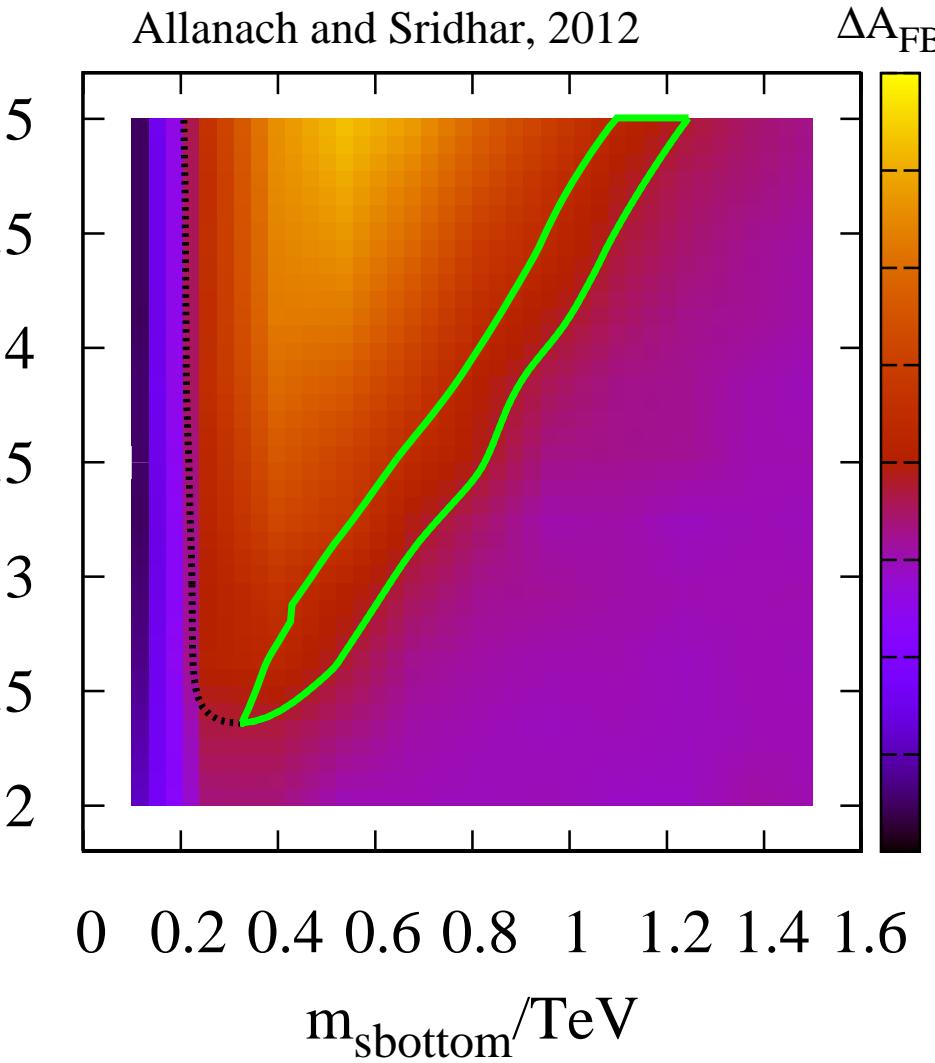
$$\frac{d\Delta\sigma}{dc} = \frac{|\lambda''_{313}|^4 \beta \hat{s}}{384\pi} \left[ \frac{(\beta c - 1)}{\hat{s}(\beta c - 1) + 2m_t^2 - 2m_{\tilde{b}_R}^2} \right]^2 + \frac{\alpha_s |\lambda''_{313}|^2 \beta}{72\hat{s}} \frac{4m_t^2 + \hat{s}(\beta c - 1)^2}{\hat{s}(\beta c - 1) + 2m_t^2 - 2m_{\tilde{b}_R}^2},$$

where  $\beta = \sqrt{1 - 4m_t^2/\hat{s}}$ ,  $\hat{s} = (p_1 + p_2)^2$ ,  $\alpha_s$  is the strong coupling constant and  $m_t$  is the top quark mass.

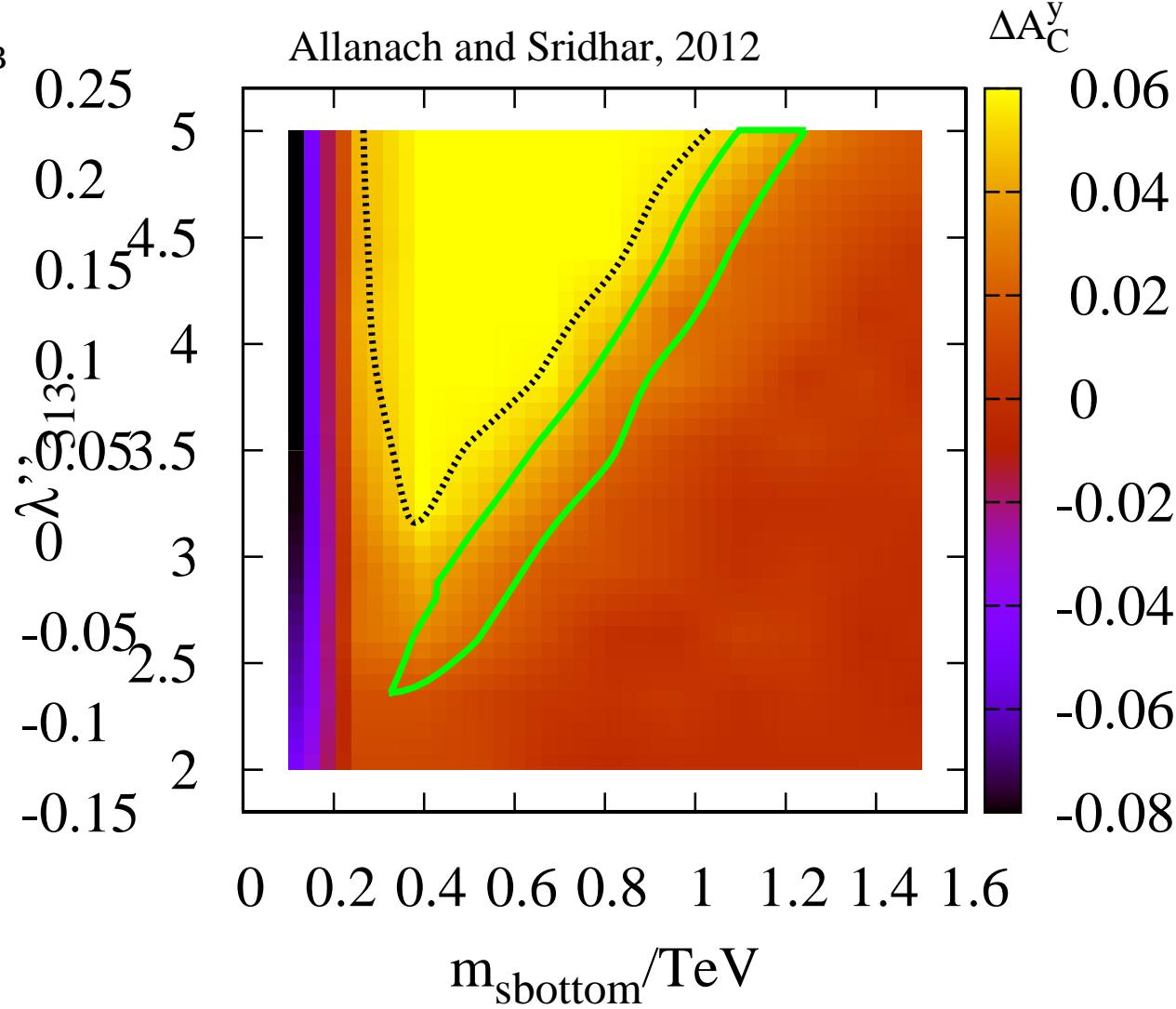


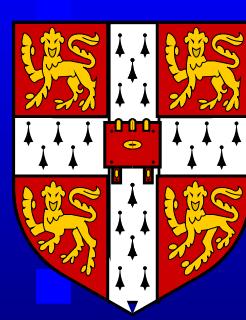
# Calculate observables with MadGraph arXiv:1205.5170

Allanach and Sridhar, 2012



Allanach and Sridhar, 2012





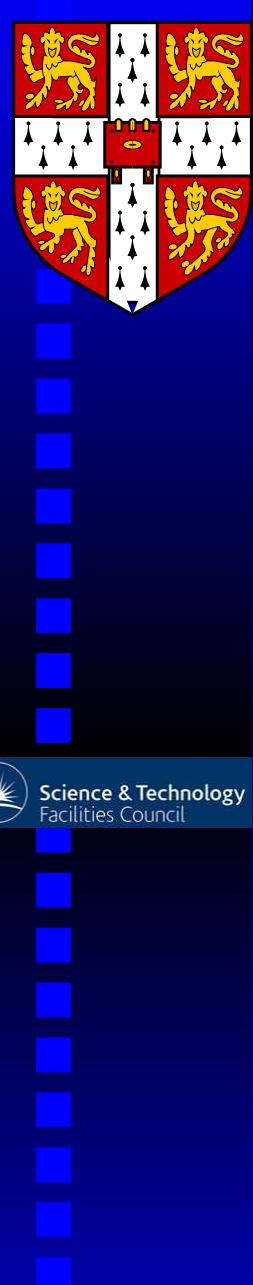
# Constraints and Predictions

$0.037 < \Delta A_{FB} < 0.205$	$-0.079 < \Delta A_C^y < 0.061$
$-0.65 < \Delta \sigma_{t\bar{t}}^{TEV}/\text{pb} < 1.51$	$4 < \Delta \sigma_{t\bar{t}}^{TEV}(\text{bin})/\text{fb} < 156$
$-0.38 < \Delta A_{FB}^l < 0.23$	$0.062 < \Delta A_{FB}^h < 0.33$
$-19.2 < \Delta \sigma_{t\bar{t}}^{LHC7}/\text{pb} < 39.2$	

Table 1: 95% CL constraints

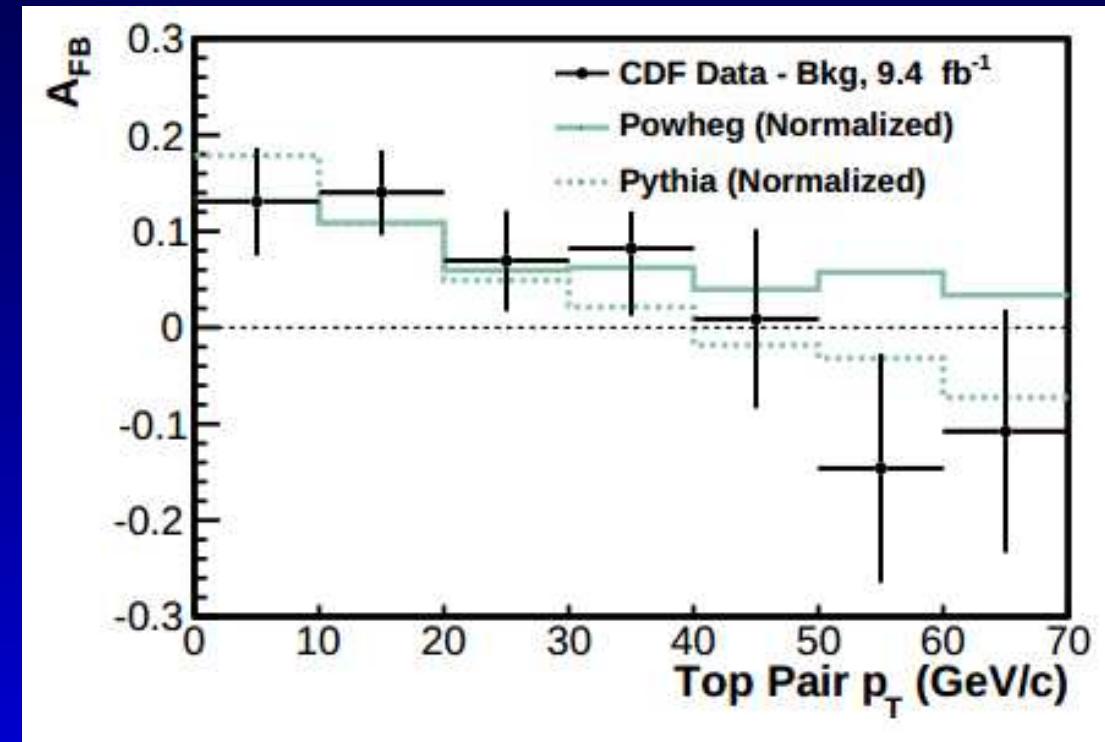
$0.037 < \Delta A_{FB} < 0.09$	$0.02 < \Delta A_C^y < 0.06$
$0 < \Delta \sigma_{t\bar{t}}^{TEV}/\text{pb} < 1.8$	$110 < \Delta \sigma_{t\bar{t}}^{TEV}(\text{bin})/\text{fb} < 156$
$0.005 < \Delta A_{FB}^l < 0.04$	$0.062 < \Delta A_{FB}^h < 0.14$
$8 < \Delta \sigma_{t\bar{t}}^{LHC7}/\text{pb} < 25$	$13 < \Delta \sigma_{t\bar{t}}^{LHC8}/\text{pb} < 33$

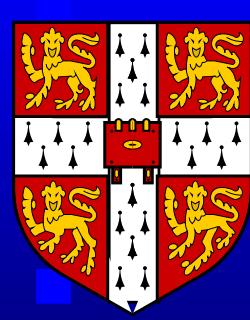
Table 2: Predicted values in good fit region



# More Tests

We are now checking to see if the most recent data from the LHC (eg  $d\sigma_{t\bar{t}}/dm_{t\bar{t}}$ ) rule our model out.  
Also,



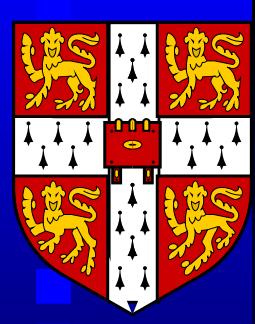


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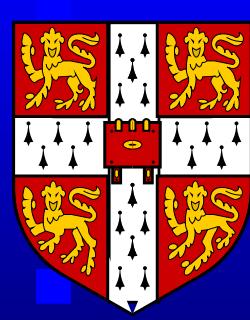
# Summary

- Need to scan over  $m_t$  when doing CMSSM searches!
- $R_p$  violation allows natural SUSY with lightish gluinos and squarks that have not yet been ruled out by searches.
- **Same-sign dilepton** searches *without* huge  $\cancel{E}_T$  cut will be interesting. It covers almost all possible cases of RPV operator.
- In case of  $U_i D_j D_k$  operators, current searches<sup>a</sup>  $\Rightarrow m_{\tilde{g}} > 800$  GeV.
- Anomalous  $A_{FB}$  measurements can *also* be explained by  $U_i D_j D_k$  type operator - fits data as of May 2012



# Backup





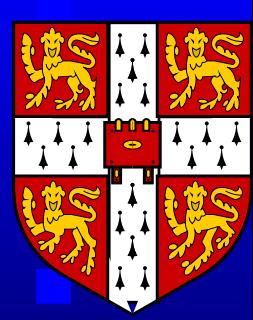
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# Sets of Cuts

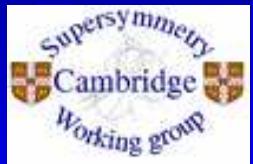
Signal Region	$m_{\mu\mu}/\text{GeV}$	$\sigma_{SS\mu\mu}^{test}/\text{fb}$	$A/10^{-3}$	$\sigma_{SS\mu\mu}^{95}/\text{fb}$
ATLAS $\mu\mu 1$	$>15$	12	1.3	58
ATLAS $\mu\mu 2$	$>100$	7.5	0.86	16
ATLAS $\mu\mu 3$	$>200$	2.1	0.29	8.4
ATLAS $\mu\mu 4$	$>300$	0.41	0.077	5.3

Table 3: The ATLAS same-sign di-muon analysis search regions.





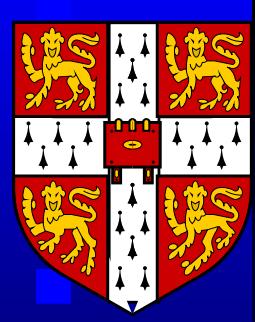
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# Sets of Cuts

Signal Region	$ \vec{p}_T^{\text{miss}} /\text{GeV}$	$m_T(l_1)/\text{GeV}$	$A/10^{-3}$	$\sigma_{SSll}^{95}/\text{fb}$
ATLASll1	$> 150$	$> 0$	1.0	1.6
ATLASll2	$> 150$	$> 100$	0.6	1.5

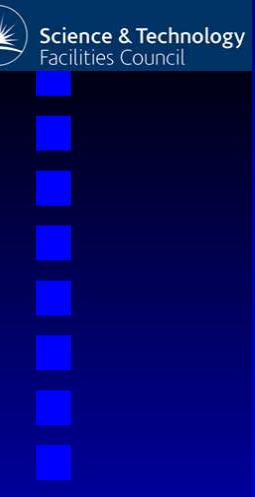
Table 4: ATLAS same sign-di lepton analysis search regions.

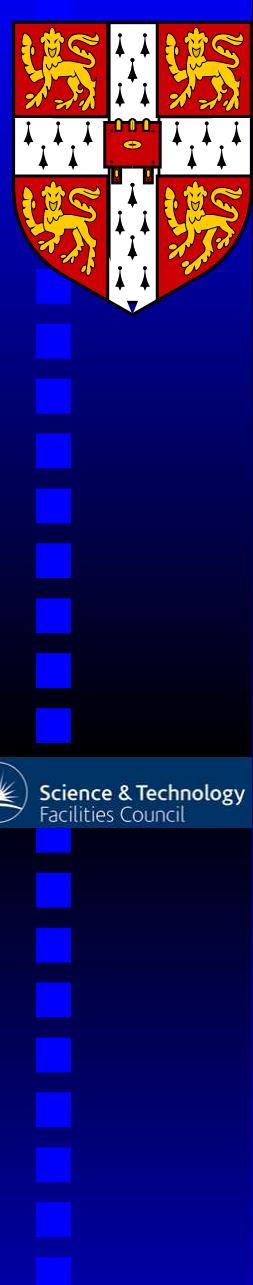


# Sets of Cuts

Signal Region	$H_T/\text{GeV}$	$ \vec{p}_T^{\text{miss}} /\text{GeV}$	$N_{ll}/\text{fb}$	$A \times \epsilon / 10^{-3}$	$N_{ll}^{95}$
CMSll1	$>400$	$>120$	2.4	3.5	$<3.7$
CMSll2	$>400$	$>50$	4.6	6.8	$<8.9$
CMSll3	$>200$	$>120$	2.5	3.7	$<7.3$

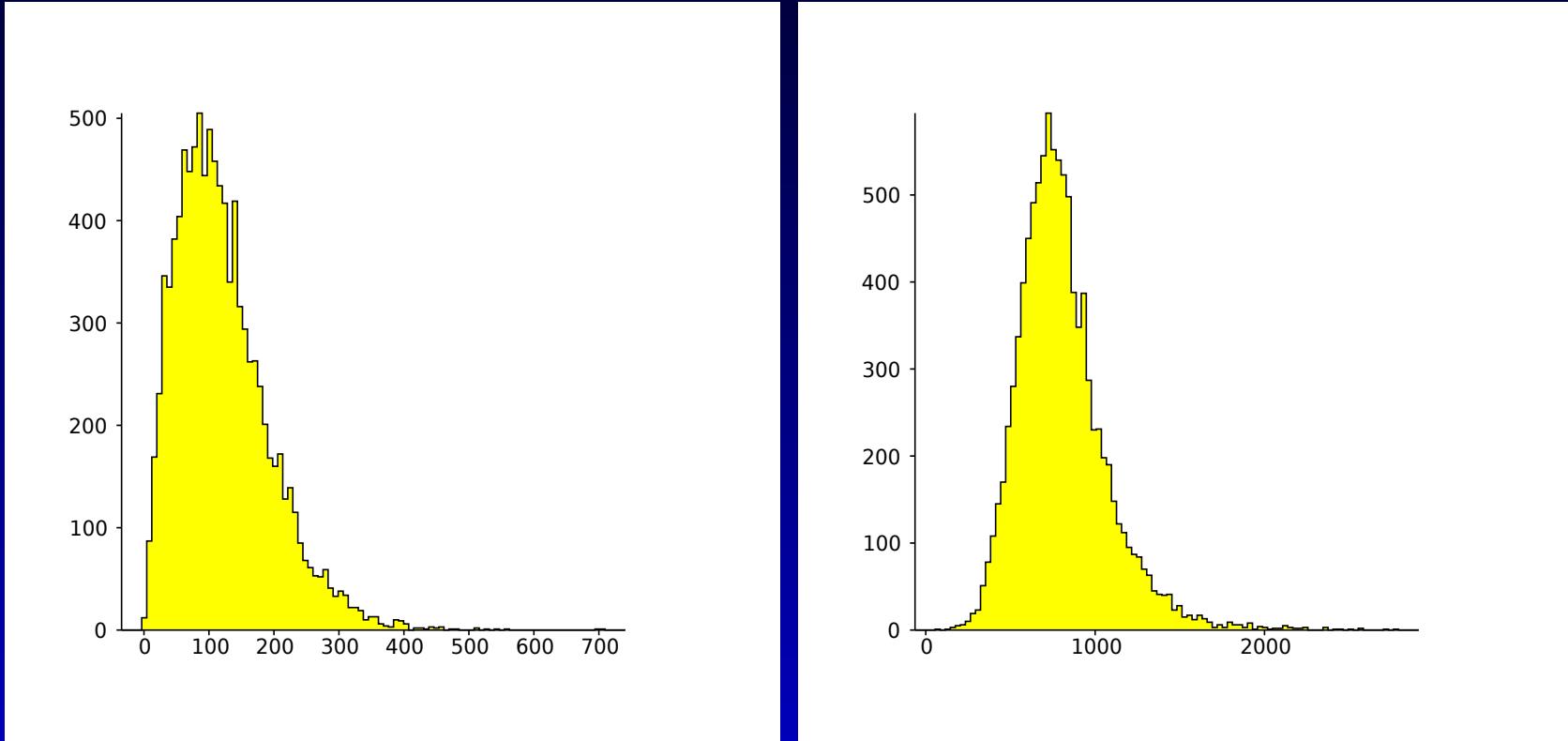
Table 5: Number of events past cuts for the CMS same sign-di lepton analysis  $N_{ll}$  predicted by our test point over SM backgrounds, and acceptance  $A$  times efficiency  $\epsilon$  of the signal selection, for the test point.





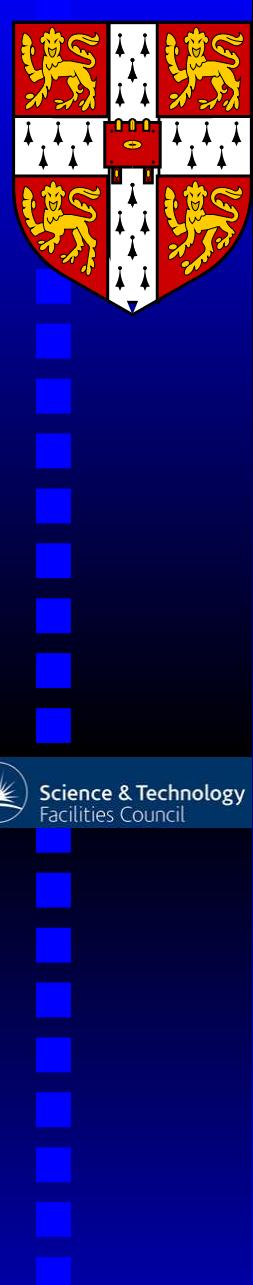
# Test Point

$$m_{\tilde{g}} = 588 \text{ GeV}, m_{\tilde{t}} = 581 \text{ GeV}.$$



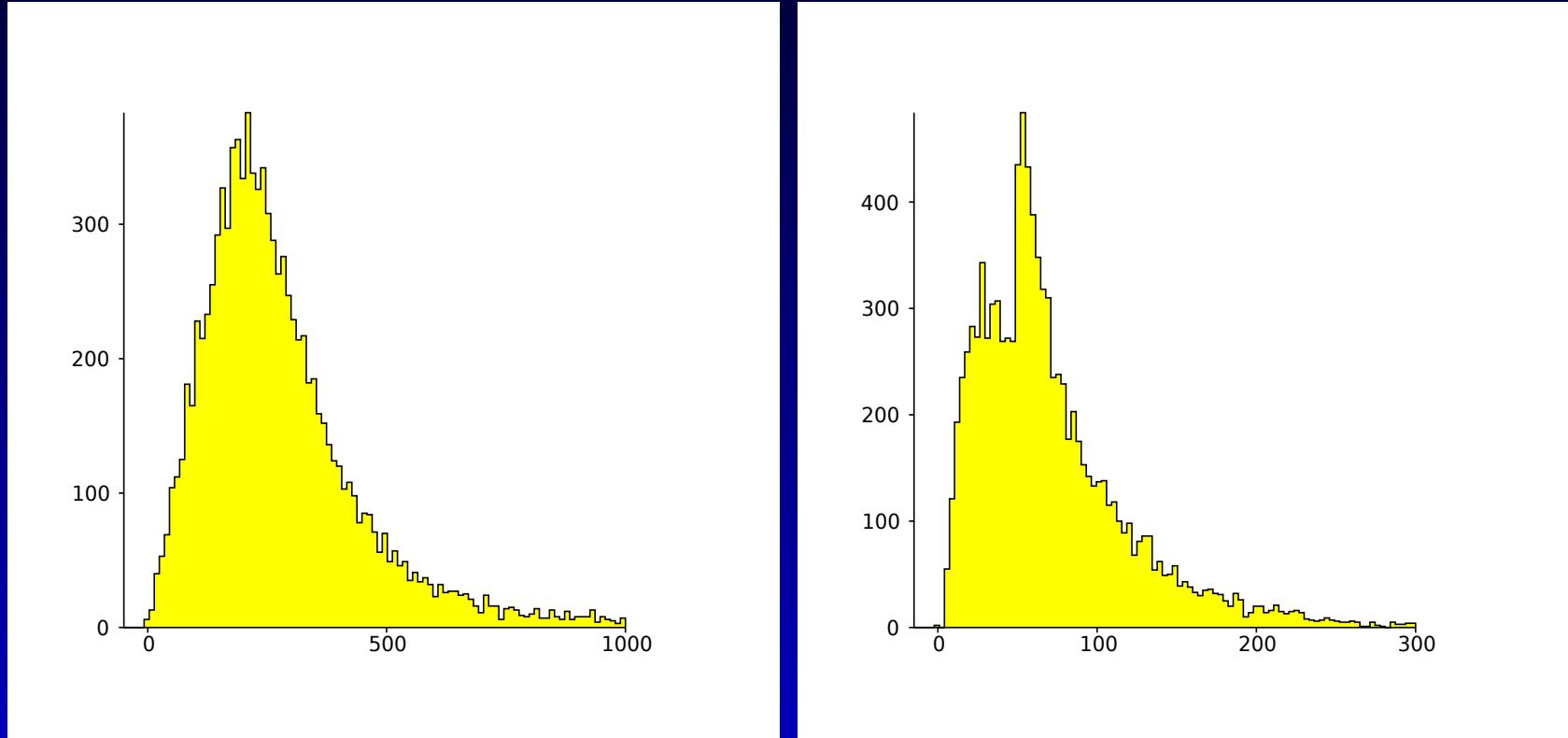
LH panel:  $\mathcal{E}_T/\text{GeV}$ , RH panel:  $H_T/\text{GeV}$





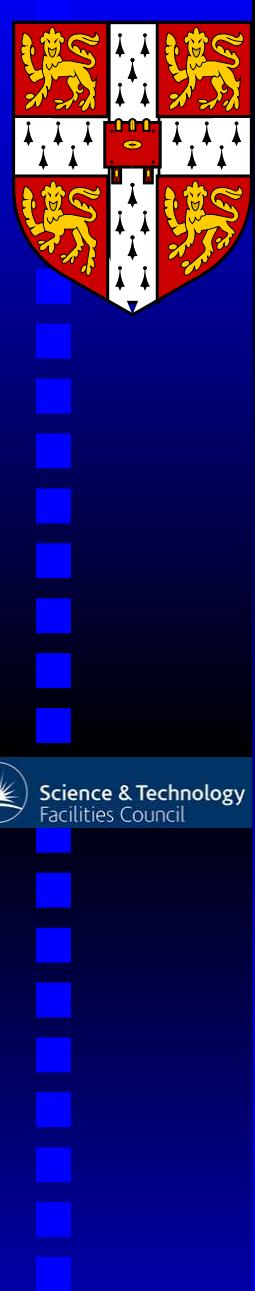
# Test Point

$$m_{\tilde{g}} = 588 \text{ GeV}, m_{\tilde{t}} = 581 \text{ GeV}.$$



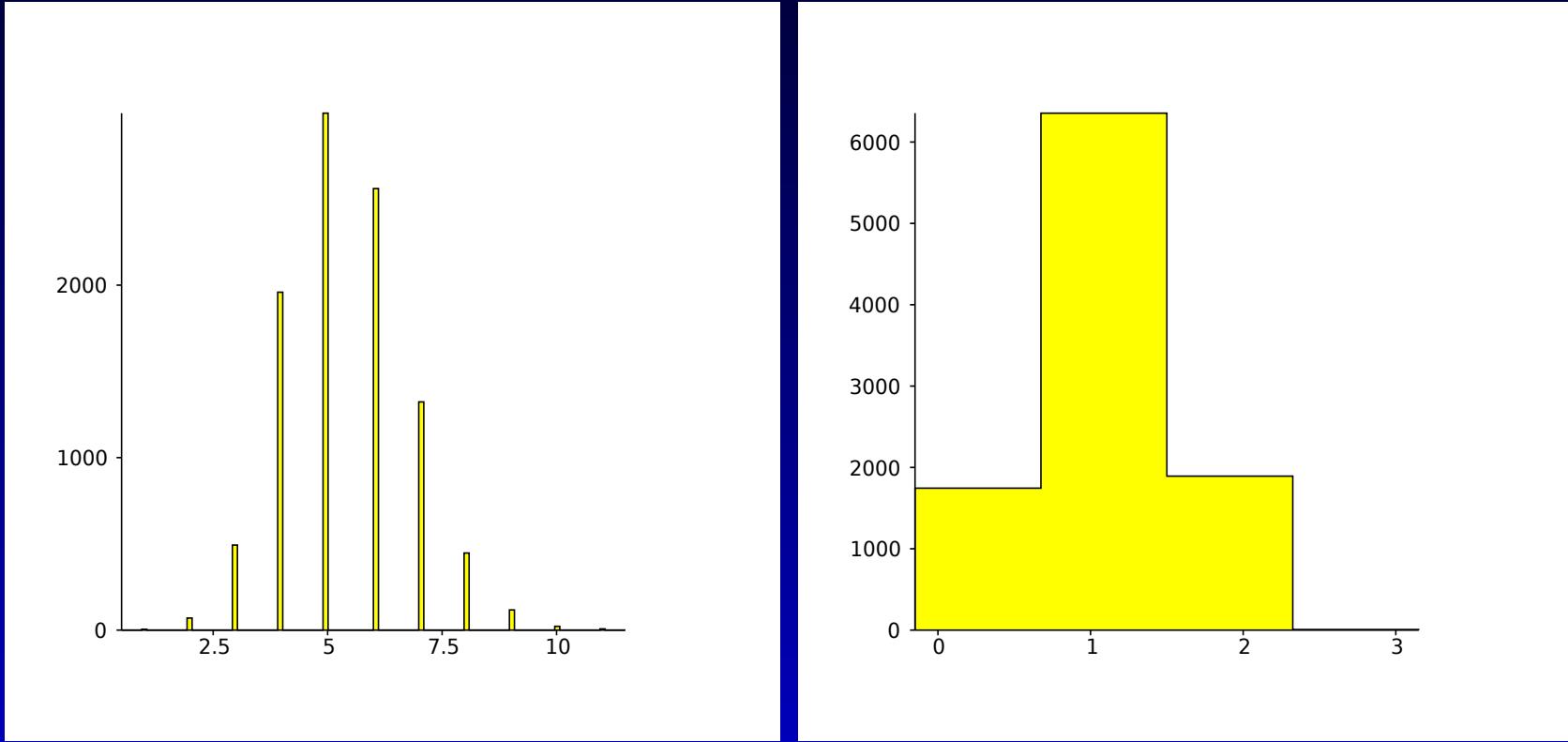
LH panel:  $p_T(j_1)/GeV$ , RH panel:  $p_T(l_1)/GeV$



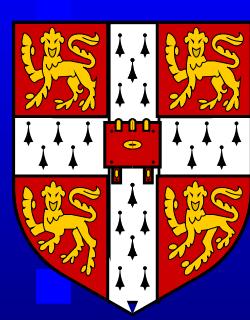


# Test Point

$$m_{\tilde{g}} = 588 \text{ GeV}, m_{\tilde{t}} = 581 \text{ GeV}.$$



LH panel:  $N_J$ , RH panel:  $N_{isol\ e,\mu}$



# Efficiencies of CMS $ll E_T$

$$\epsilon = \frac{\text{SUSY events past cuts}}{\text{SUSY events}}$$

You pay for the di-leptonic  $t\bar{t}$  branching ratio.

