

# COMET/PRISM and cLFV

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## PPAP Community Meeting

Ajit Kurup

22<sup>nd</sup> July 2014

**Imperial College**  
London

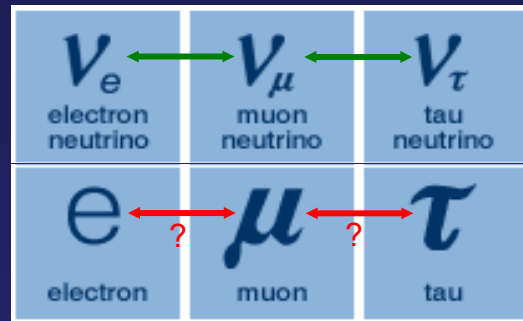
# Introduction

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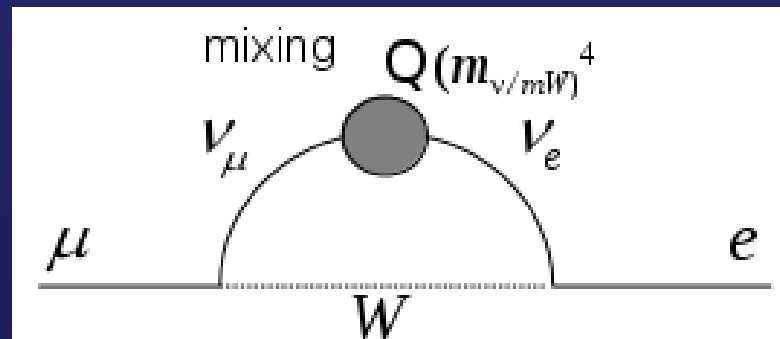
- Physics motivation for charged lepton flavour violation searches.
- Why muon to electron conversion is promising.
- COMET
- PRISM
- On behalf of COMET-UK

# Physics Motivation

- Neutrinos in the SM are massless but observation of neutrino oscillations is direct evidence that neutrinos have mass.
  - Proof that neutral lepton flavour number is not conserved.



- Simplest extension to SM to include neutrino mixing.

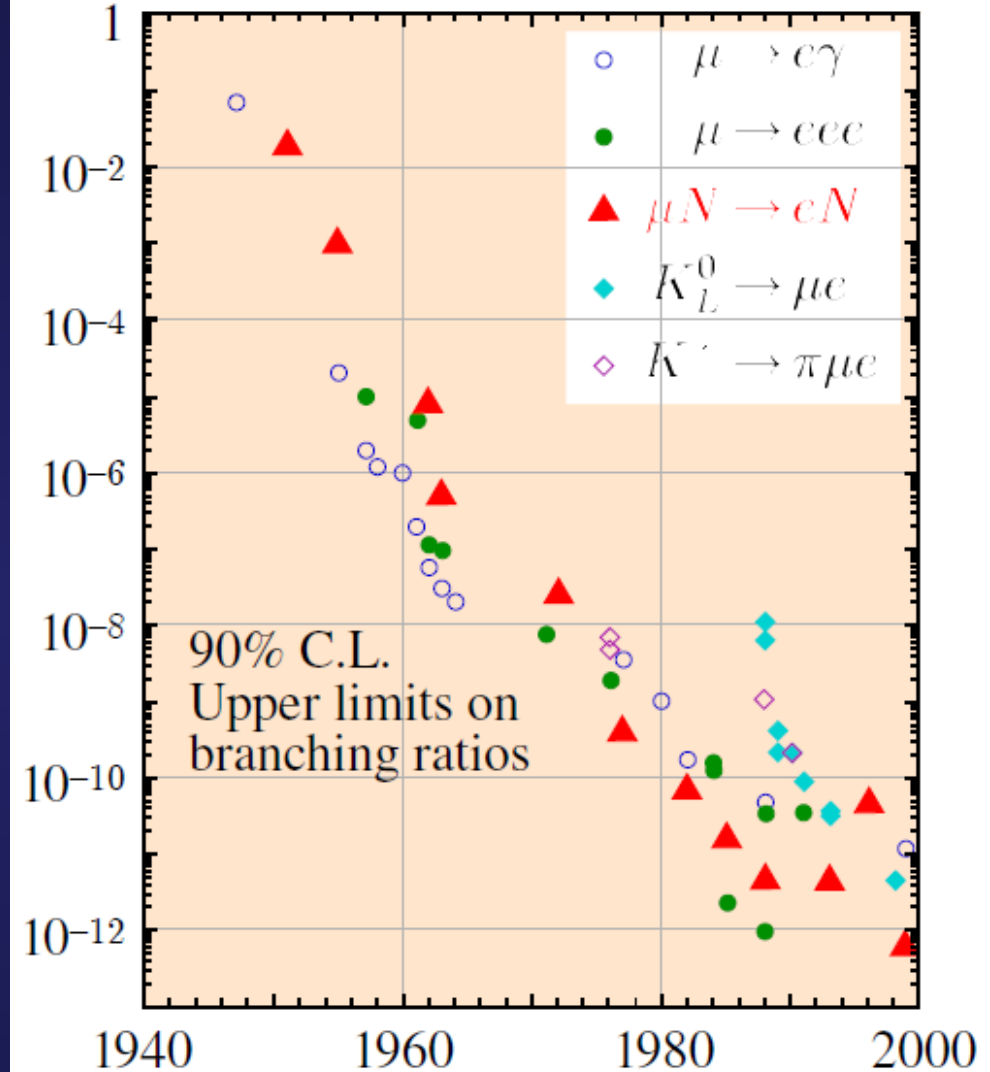


Rate  $< 10^{-52}$

- A number of BSM theories predict charged lepton flavour number non-conservation.
  - Rates much higher, e.g. muon to electron conversion could be as large as  $10^{-13}$  –  $10^{-15}$

# cLFV Processes

- Tau based, e.g.
  - $\tau \rightarrow \mu \gamma$
  - $\tau \rightarrow 3\mu$
  - $\tau \rightarrow p \mu \mu$
  - ...
- Muon based, e.g.
  - $\mu^+ \rightarrow e^+ \gamma$
  - $\mu^+ \rightarrow e^+ e^+ e^-$
  - $\mu^-(A, Z) \rightarrow e^-(A, Z)$



First CLFV search by Hincks  
and Pontecorvo in 1947 for  
 $\mu^+ \rightarrow e^+ + \gamma$



# cLFV Processes

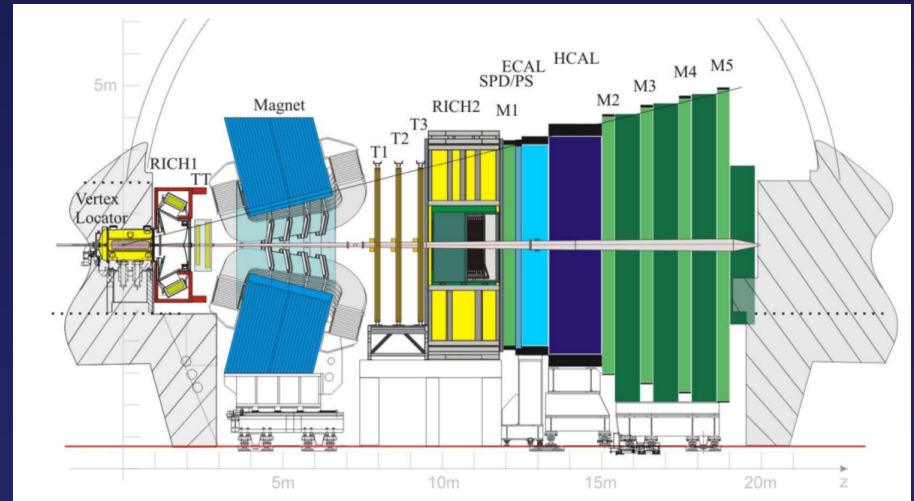
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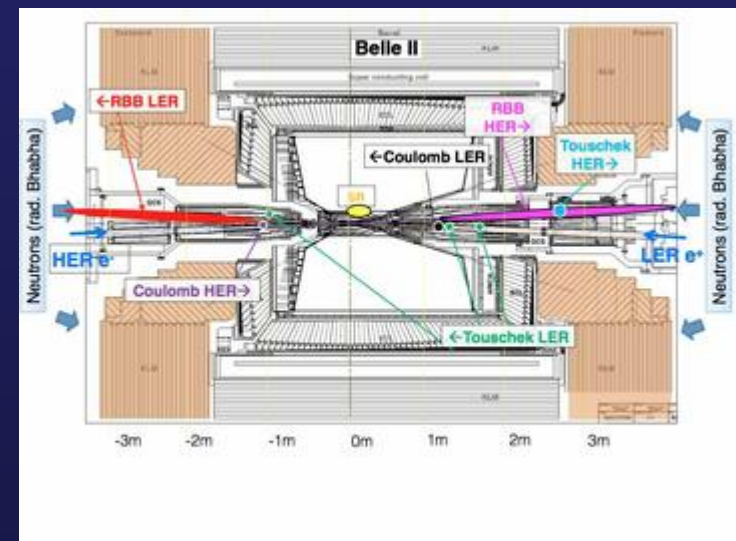
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## LHCb



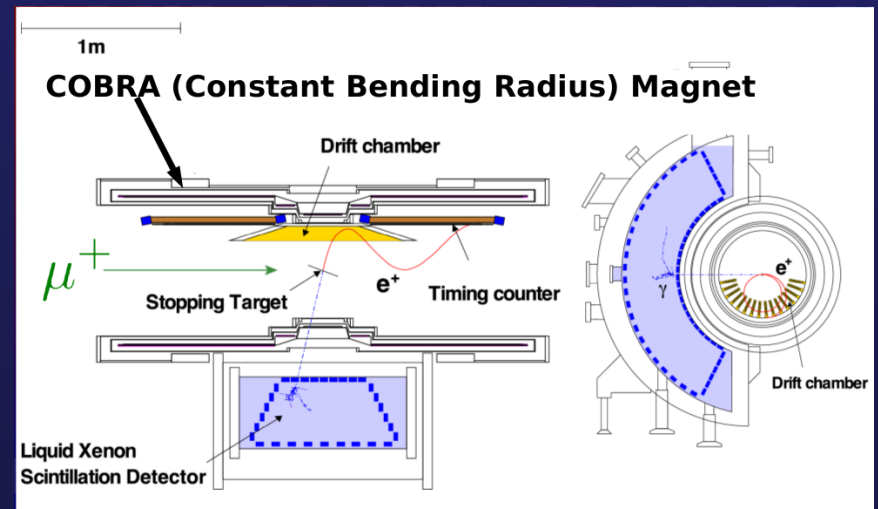
## BELLE-II



# cLFV Processes

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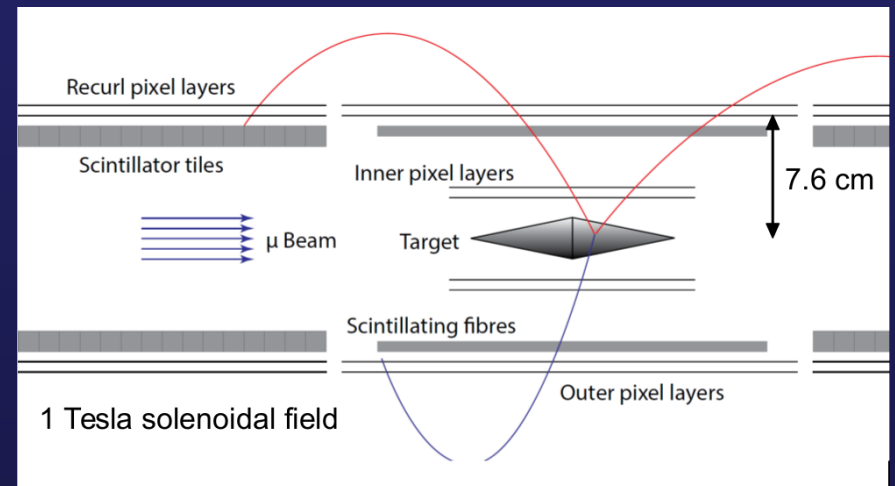
## MEG @PSI



# cLFV Processes

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## Mu3e @PSI



# cLFV Processes

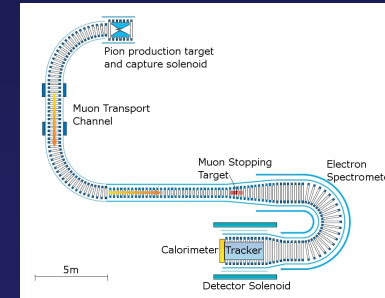
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- ...

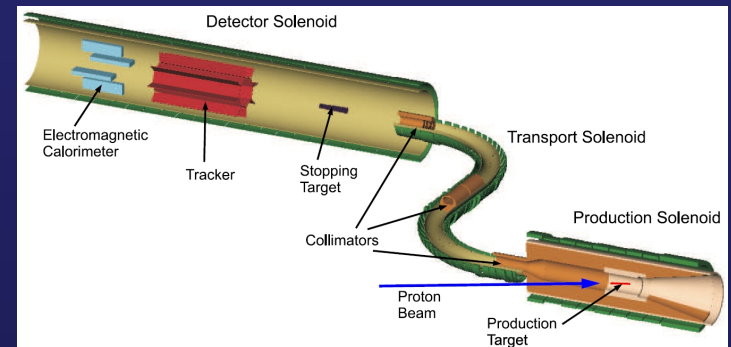
- Muon based, e.g.

- $\mu^+ \rightarrow e^+ \gamma$
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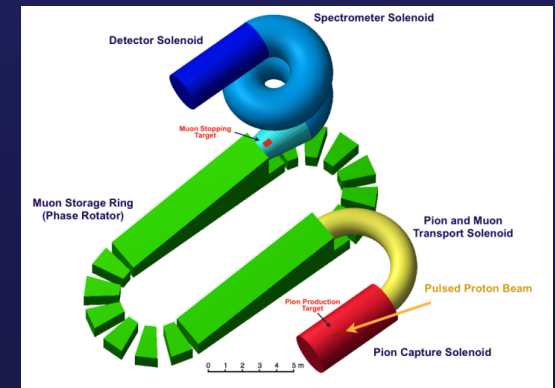
## COMET @J-PARC



## Mu2e @FNAL



## PRISM

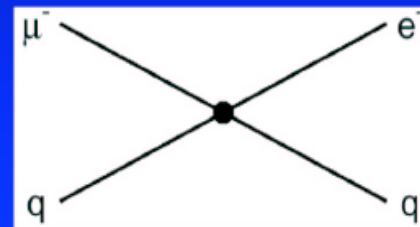
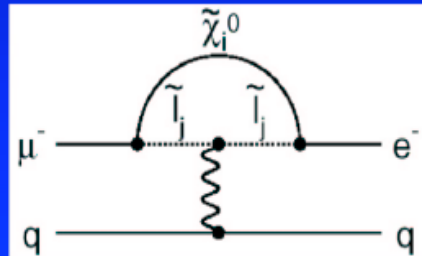


# Muon to Electron Conversion

- Many BSM models give predictions  $10^{-13}$ - $10^{-15}$ .

Supersymmetry

Predictions at  $10^{-15}$



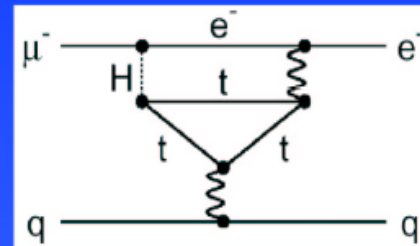
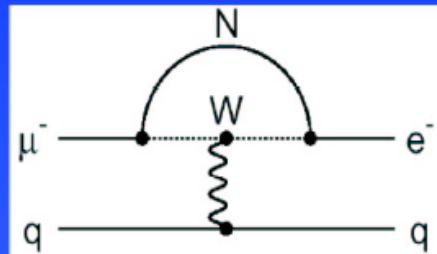
Compositeness

$$\Lambda_c = 3000 \text{ TeV}$$

Heavy Neutrinos

$$|U_{\mu N}^* U_{eN}|^2 =$$

$$8 \times 10^{-13}$$



Second Higgs doublet

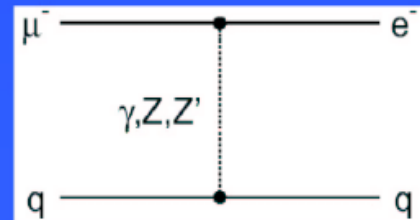
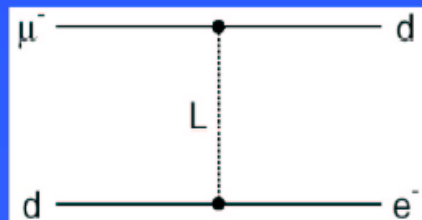
$$g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$$

Leptoquarks

$$M_L =$$

$$3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$$

After W. Marciano



Heavy  $Z'$ ,  
Anomalous  $Z$   
coupling

$$M_{Z'} = 3000 \text{ TeV}/c^2$$

$$B(Z \rightarrow \mu e) < 10^{-17}$$

# Muon to Electron Conversion

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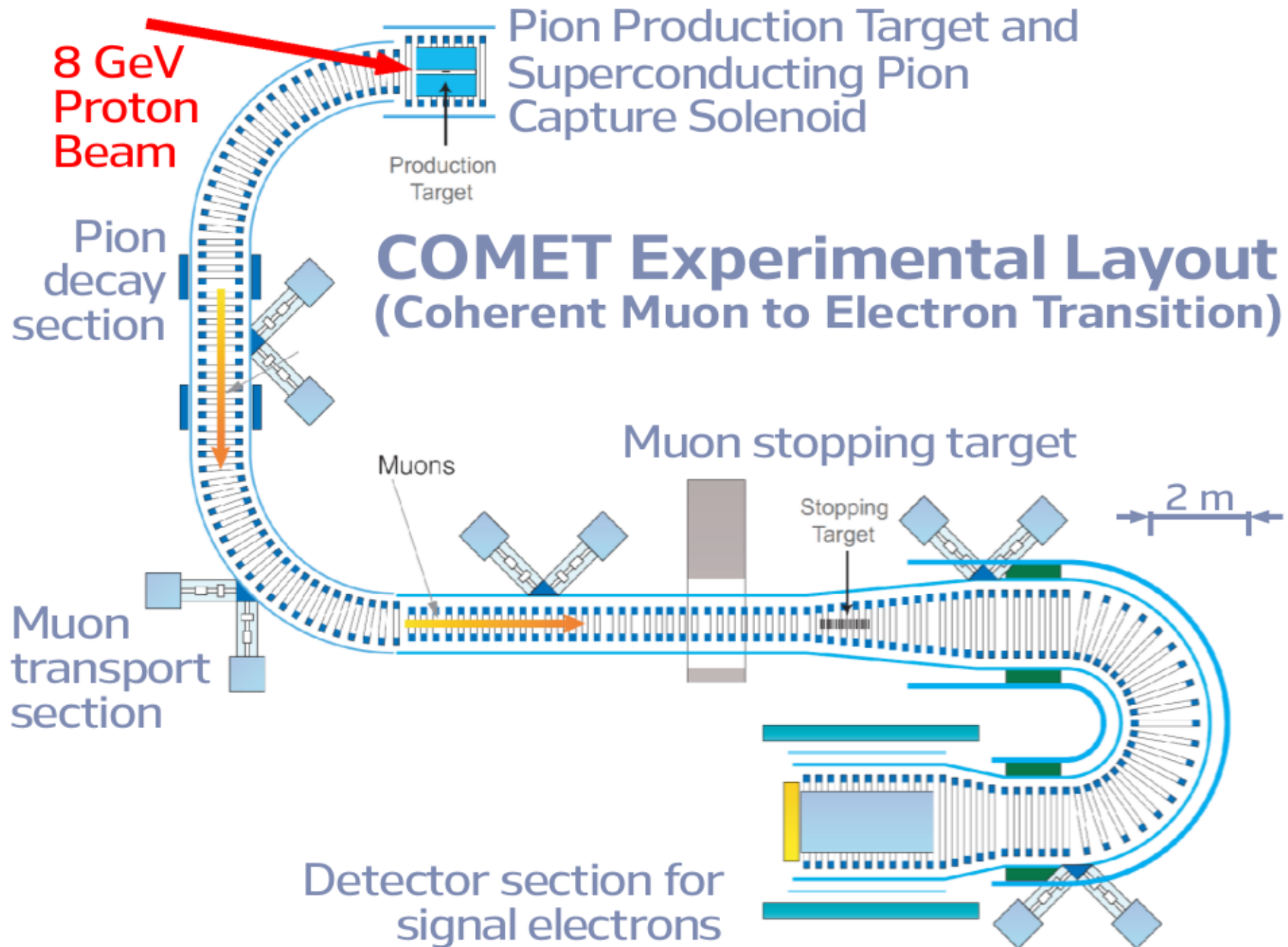
- Many BSM models give predictions  $10^{-13}$ - $10^{-15}$ .
  - Current limit by SINDRUM II  $7 \times 10^{-13}$
- Developments in accelerator concepts allows significantly higher intensity muon beams.
- Expected  $>10^4$  improvement in sensitivity utilising current accelerator technology.
  - COMET and Mu2e.
- Expected  $>10^6$  improvement in sensitivity utilising future accelerator technology.
  - PRISM.

# COMET Status

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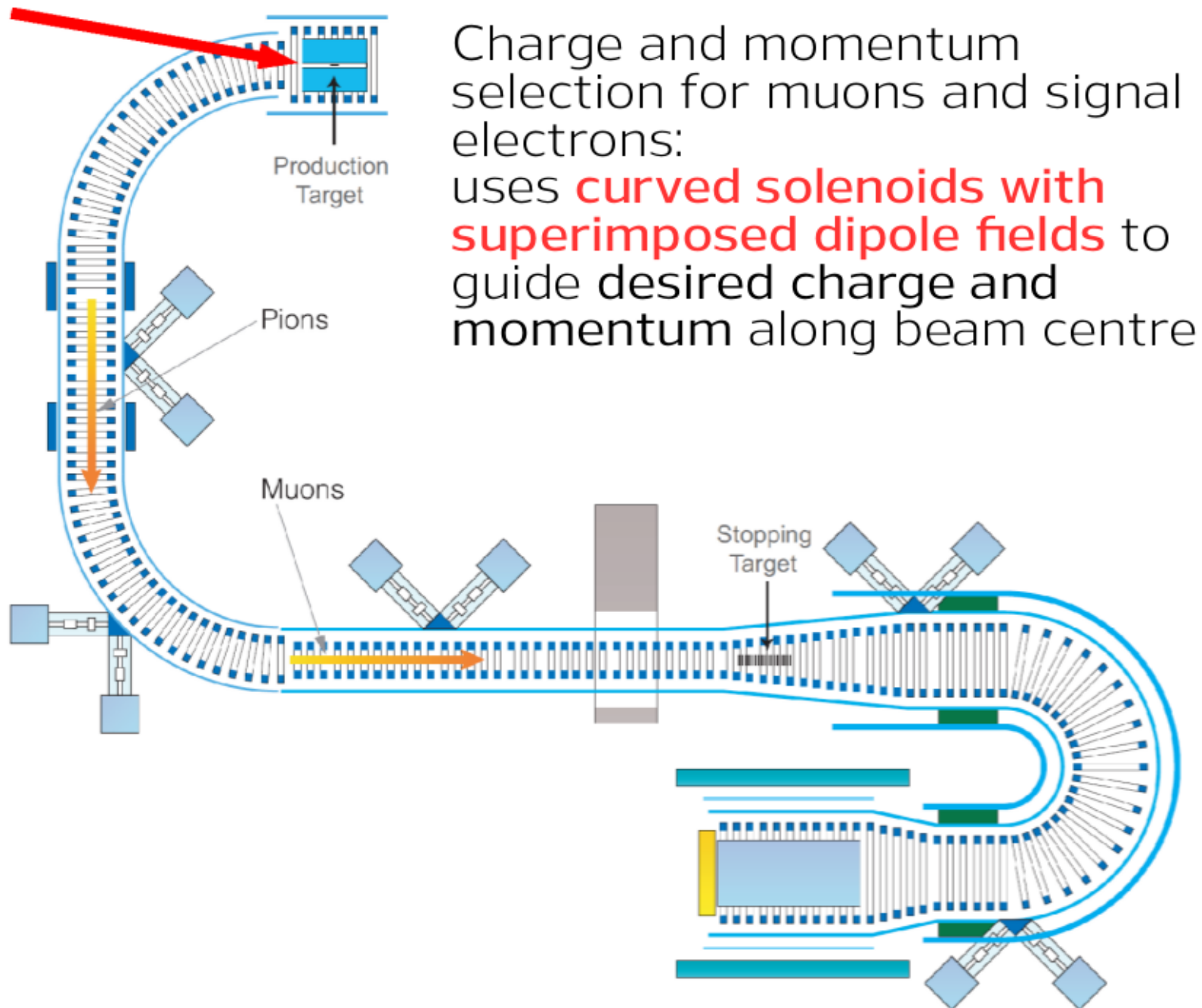
- Currently under construction.
  - Proton beam line magnets being installed.
  - Experimental hall being excavated.
  - Superconducting magnets being manufactured.
  - Drift chamber construction started, tracker and ECAL designs being finalised.
- Phase-I data in JFY 2015/16.
- Next flagship experiment for J-PARC.
  - Included in the Japanese Research Masterplan.
- Strong recommendation for muon physics in P5.

# The Coherent Muon to Electron Transition (COMET) Experiment





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## The Effect of Solenoid and Dipole Fields

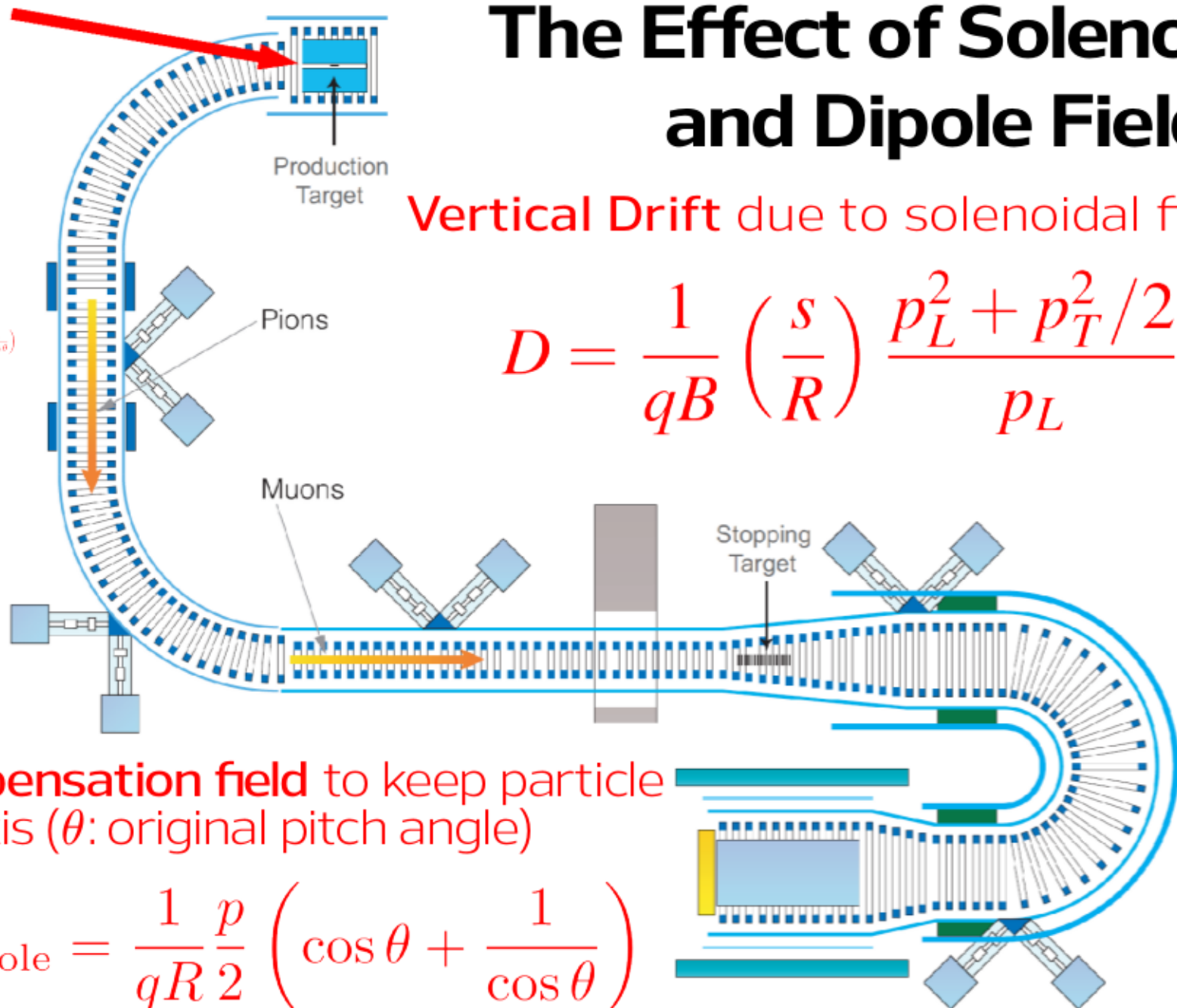
Vertical Drift due to solenoidal field:

$$D = \frac{1}{qB} \left( \frac{s}{R} \right) \frac{p_L^2 + p_T^2/2}{p_L}$$

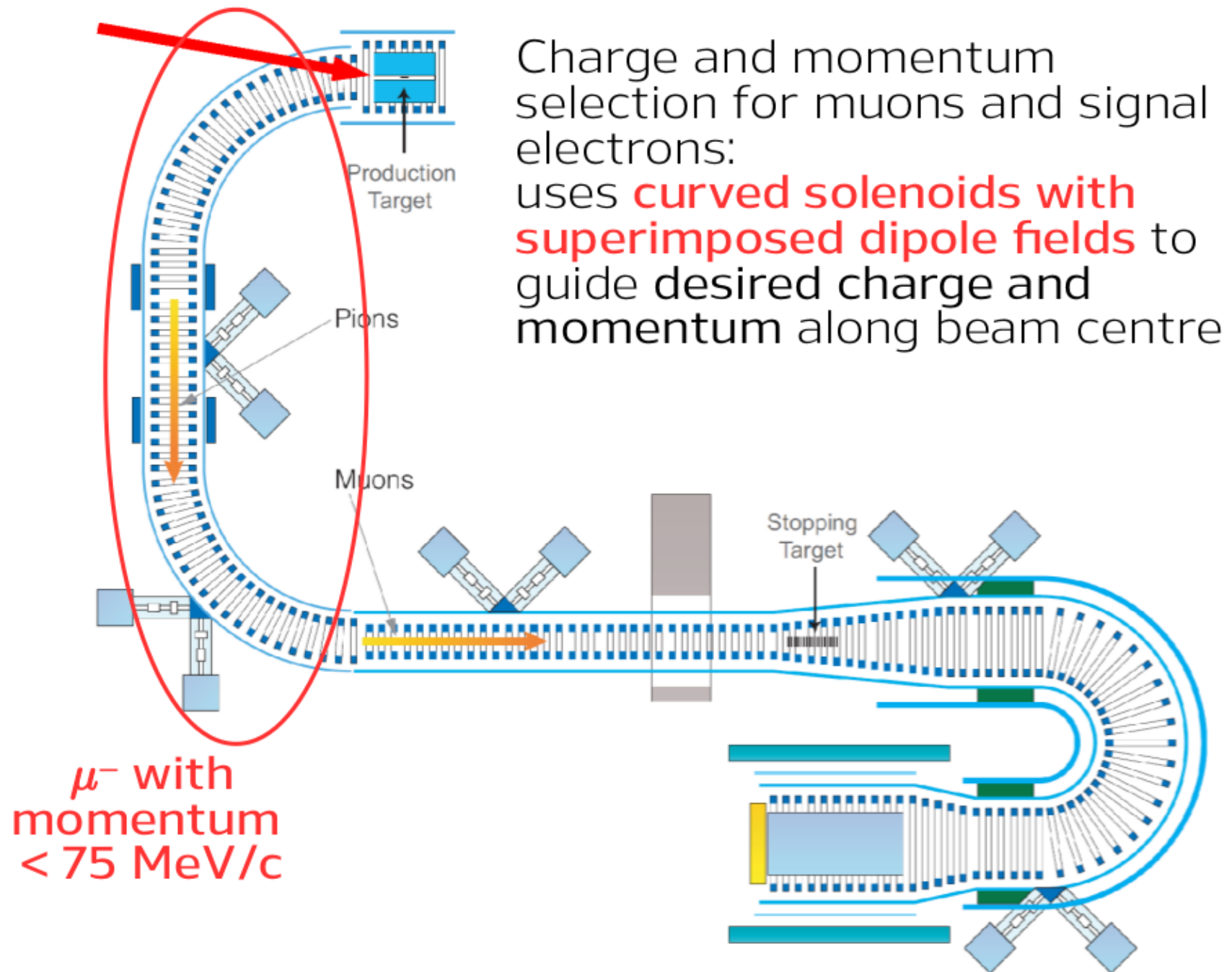
$$B_{\text{dipole}} = \frac{1}{qR} \frac{p}{2} \left( \cos \theta + \frac{1}{\cos \theta} \right)$$

Compensation field to keep particle on-axis ( $\theta$ : original pitch angle)

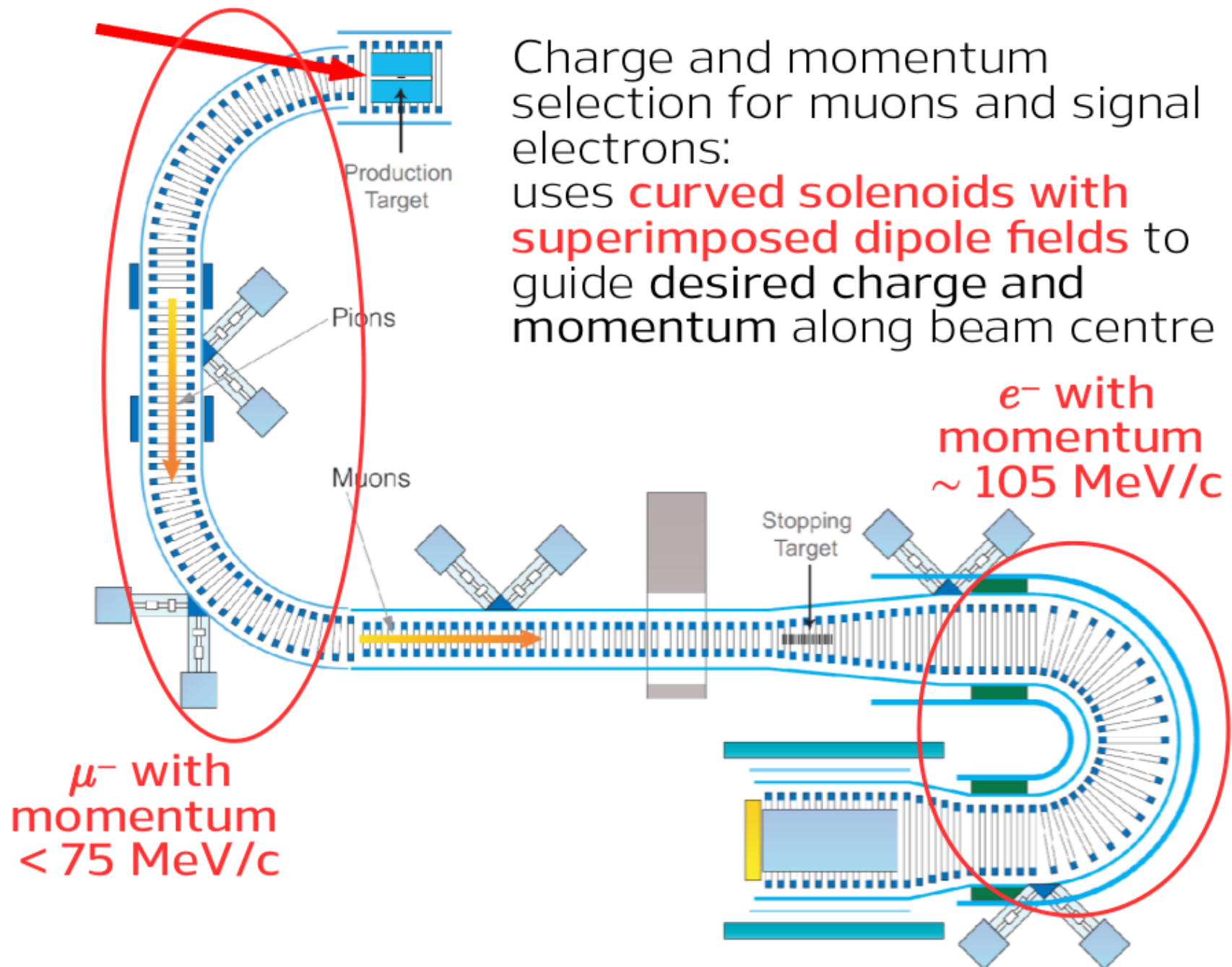
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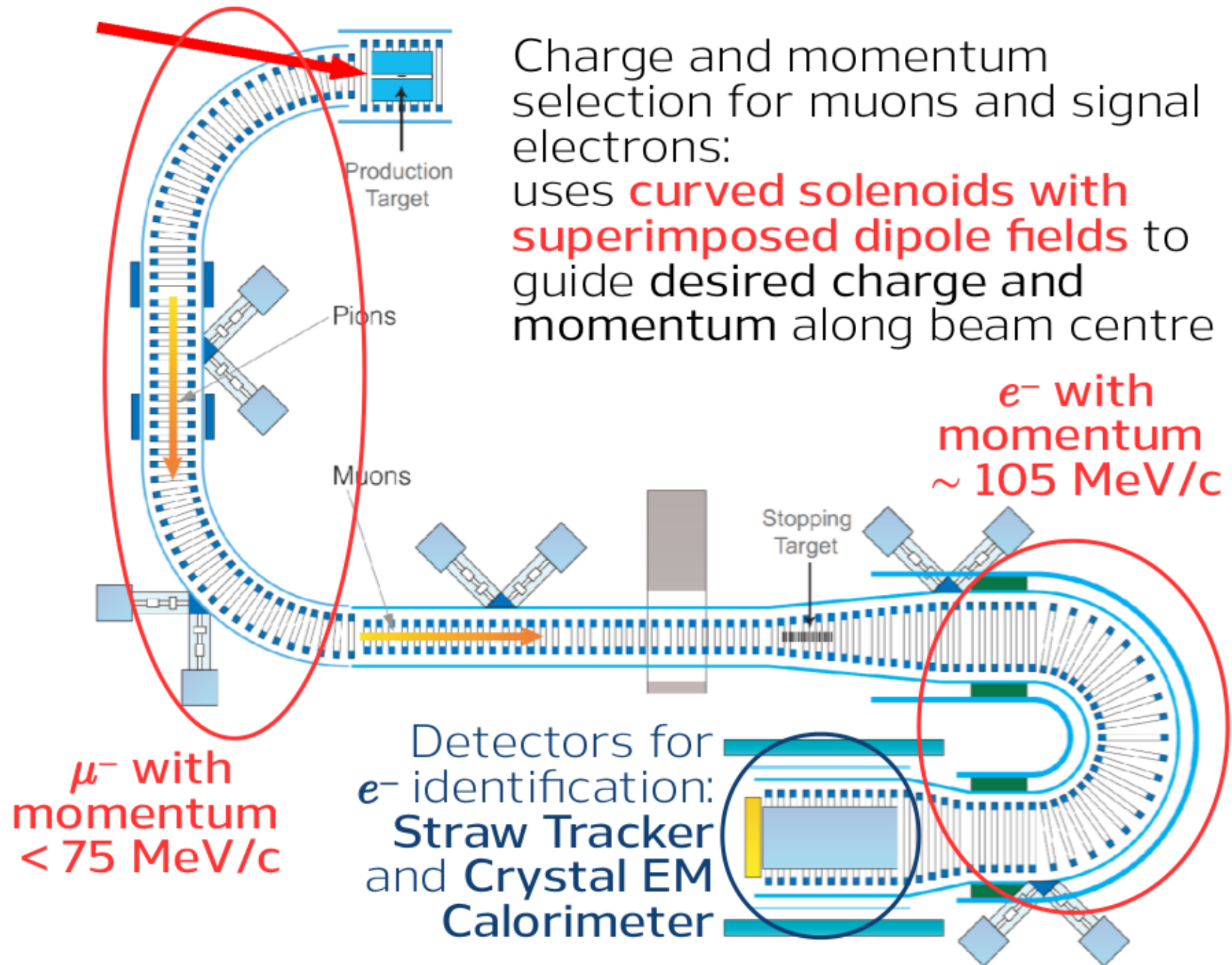
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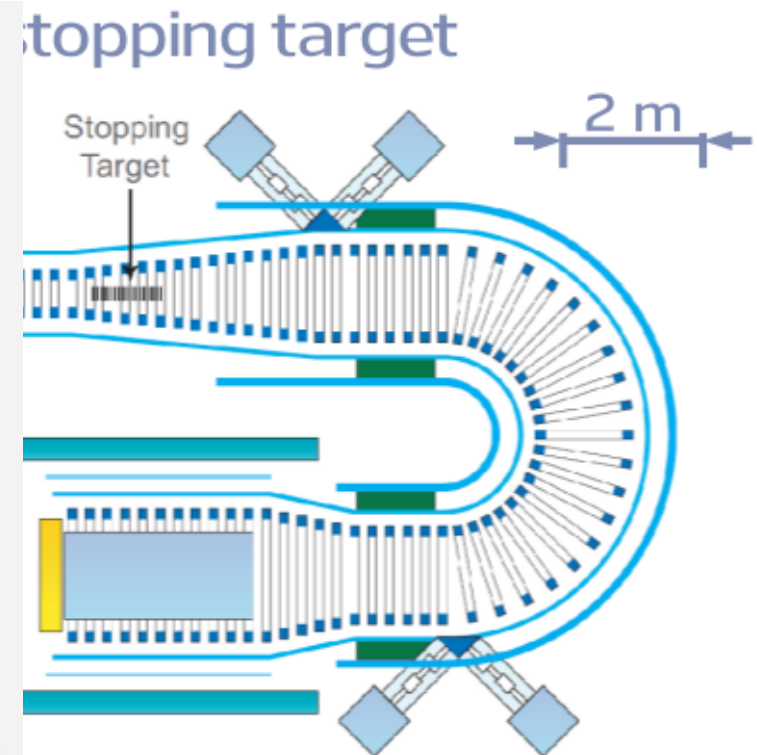
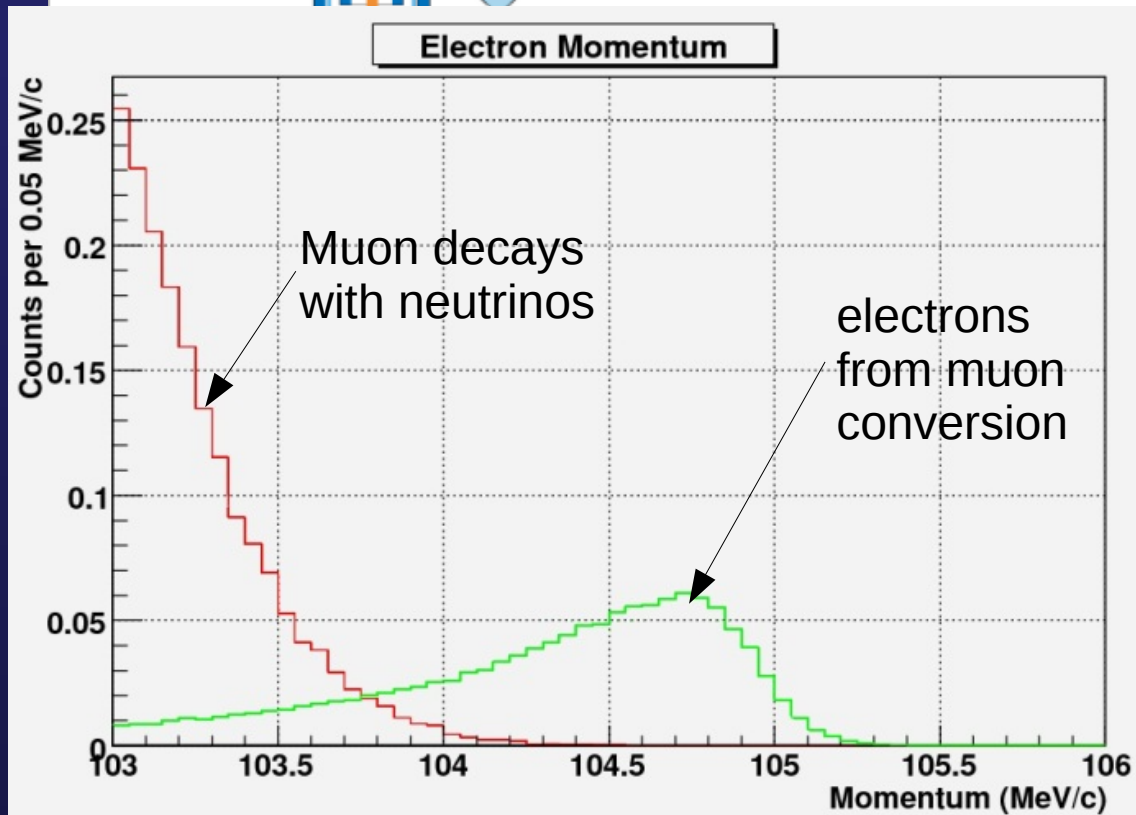
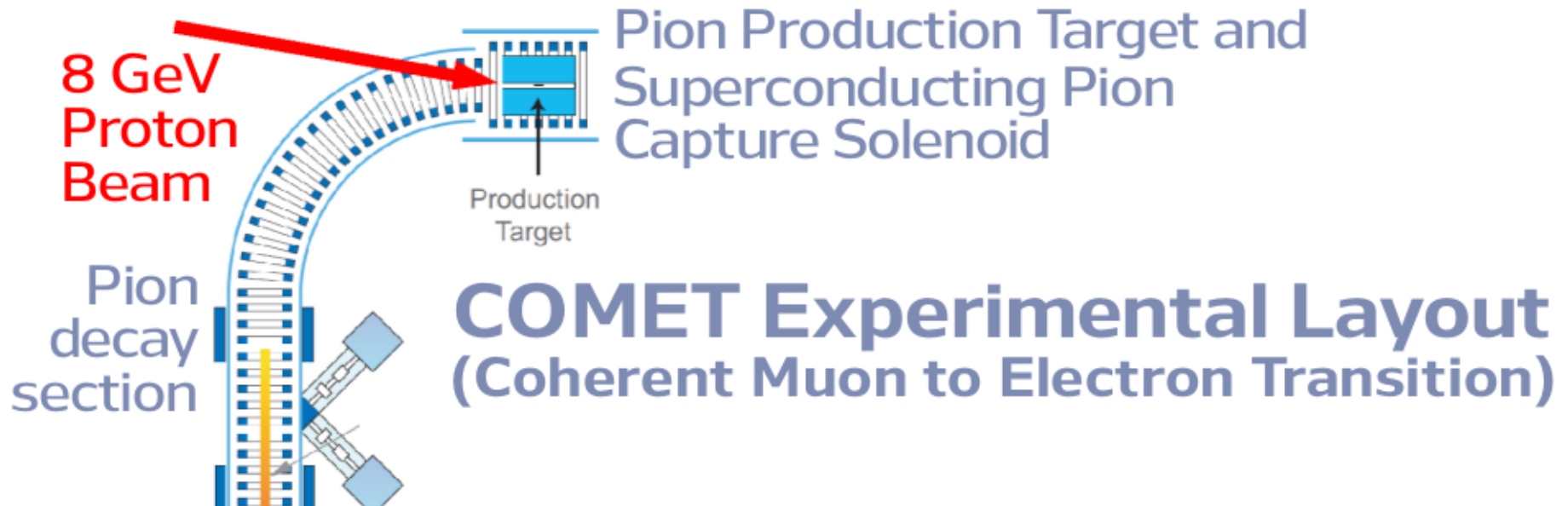


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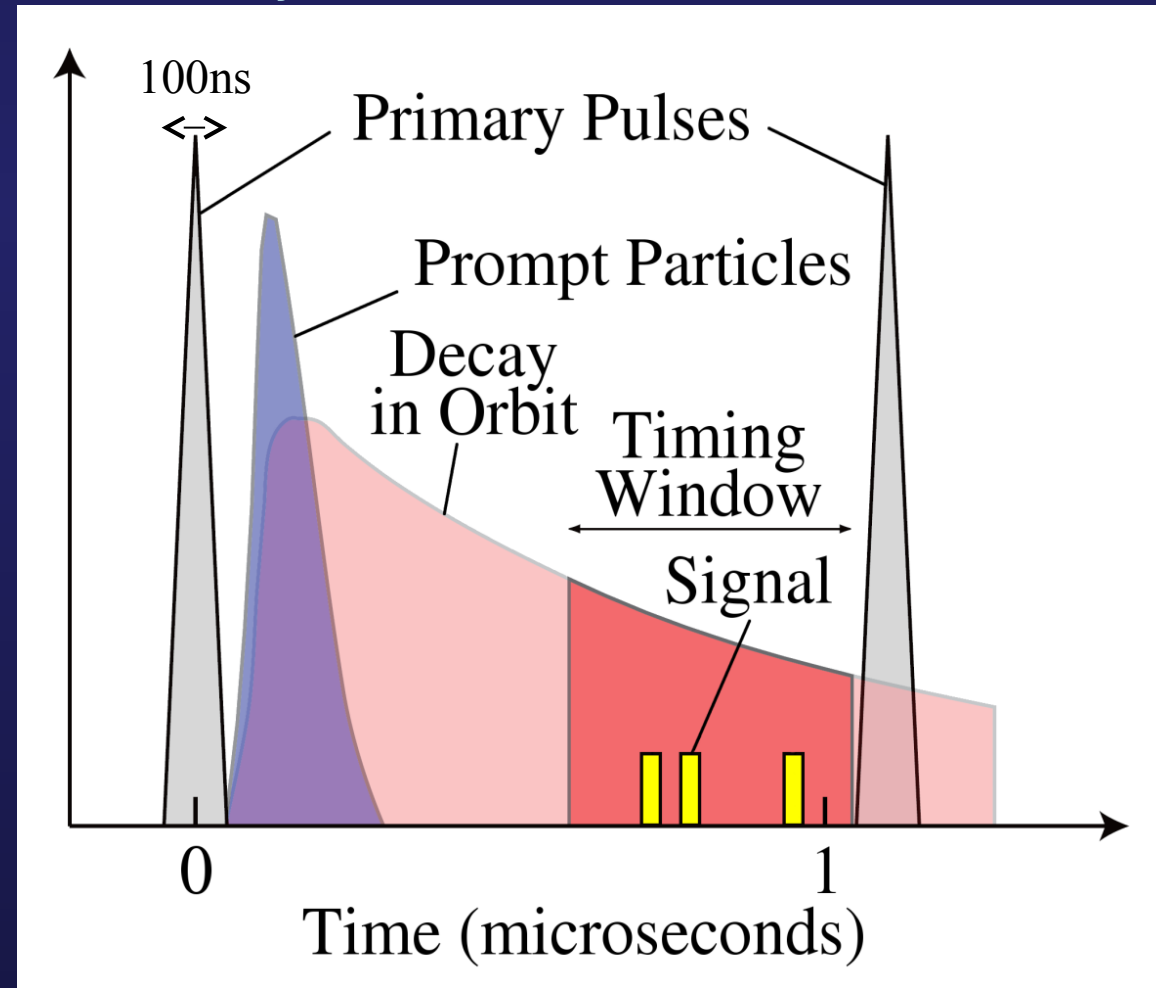


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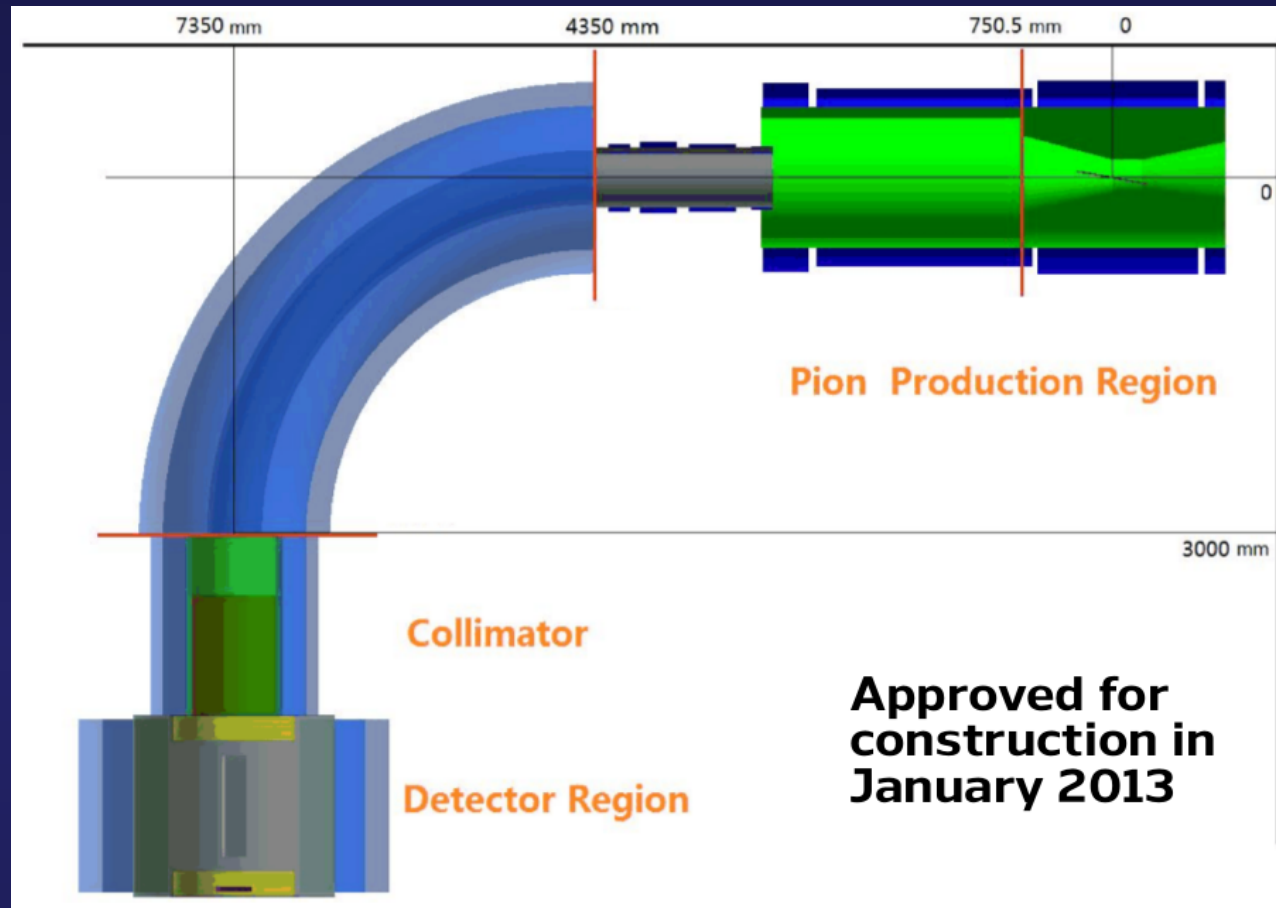


# Proton Beam and Backgrounds

- Proton beam requires specific pulse structure to reduce backgrounds.
- Cleanliness quantified by extinction factor.
  - Require  $10^{-9}$ .
  - Important to measure this.
- Understanding background modes is the key.
  - Decay in orbit.
  - Prompt beam related.
    - Decay in flight.
    - Radiative capture.
  - Scattered electrons.
  - Cosmic rays.
  - Neutrons.



# COMET Phase-I



- Several advantages to staging.
  - Allows measurement of backgrounds.
  - Test beam line and detector technology.
- Sensitivity 100 times better than World's best limit.
  - 5% of the proton beam intensity of Phase-II and 90 days running.



# COMET-UK Background

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- Participated in COMET since first collaboration meeting (2008).
  - Imperial College, Manchester/Cockcroft, RAL, UCL.
- UK involvement instrumental in the preparation of the COMET CDR, the Phase-I EOI, LOI and TDR, which culminated in approval of Phase-I.
- Sol submitted in 2012.
- Consolidated Grant funding from PPGP.
- In 2013, submitted to STFC a three-year proposal (2014–17) for participation in COMET Phase-I.
  - PPRP recommended it but was not approved by Science Board due to lack of funds.
- Continued participation through other means.
  - Consolidated Grant, fellowships, University funds, bridging funds and PhD studentships.
- Targeting full participation in Phase-II including R&D.

# COMET Management Structure

## COMET Phase-I Management Structure (July 2014)

### Executive Board (Elected)

Project Manager:  
S. Mihara (KEK)

Spokesperson:  
Y. Kuno (Osaka)

Collaboration  
Board Chair:  
Y. Uchida (Imperial)

Institute 1

• One member  
• per institute

Institute N

Collaboration  
Board

**33 Institutions**

Proton beam-line:  
Y. Fukao (KEK)

Proton target:  
C. Densham (RAL)  
S. Mihara (KEK)

Solenoid magnets:  
M. Yoshida (KEK)

Muon target and  
beam collimator:  
TBD

Straw Chambers:  
H. Nishiguchi (KEK)  
P. Evtoukhovitch  
(JINR)

ECAL:  
J. Tojo (Kyushu)

Proton beam  
monitor:  
TBD

Muon beam  
monitor:  
TBD

CyDet:  
A. Sato (Osaka)

CyDet physics  
and software  
Y. Yuan (IHEP)

X-ray calibration:  
TBD

Cosmic veto:  
A. Drutskoy (ITEP)  
TBD (KEK)

DAQ-trigger:  
P. Dauncey  
(Imperial)

DAQ-readout:  
M. Wing (UCL)

Software:  
A. Kurup (Imperial)

Offline  
infrastructure:  
Y. Igarashi (KEK)

Infrastructure:  
S. Mihara (KEK)

Safety:  
TBD

**Sub-Project  
Coordinators**

(also UK leadership of **Internal Review Panel, CDR/TDR Editorial boards & Speakers Bureau**)

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Muon target and  
beam collimator:  
TBD

Straw Chambers:  
H. Nishiguchi (KEK)  
P. Evtoukhovitch  
(JINR)

ECAL:  
J. Tojo (Kyushu)

Proton beam  
monitor:  
TBD

Muon beam  
monitor:  
TBD

CyDet:  
A. Sato (Osaka)

CyDet physics  
and software  
Y. Yuan (IHEP)

X-ray calibration:  
TBD

Cosmic veto:  
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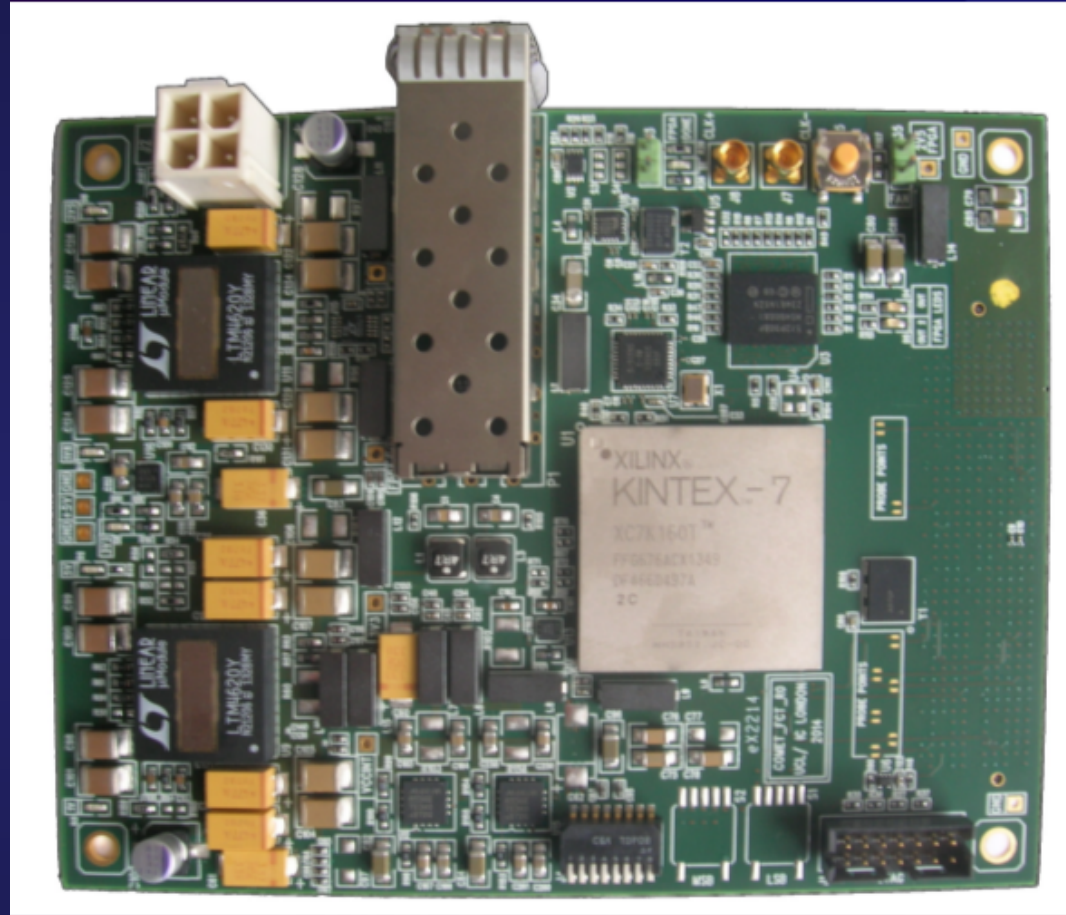
# COMET-UK: Physics and Software

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- UK leads the COMET Software and Analysis groups.
- COMET software adapted from T2K ND280 software.
  - Exploiting UK leadership in ND280 software (online and offline).
- Software and analysis methods need to be fully established before data-taking.
- Other regions contribute fully to software effort, but transfer of software leadership is not feasible.
- Combining RA effort with PhD and undergraduate student effort to meet obligations.
- Critical for COMET that CG effort continues into full exploitation stage for Phase-I.

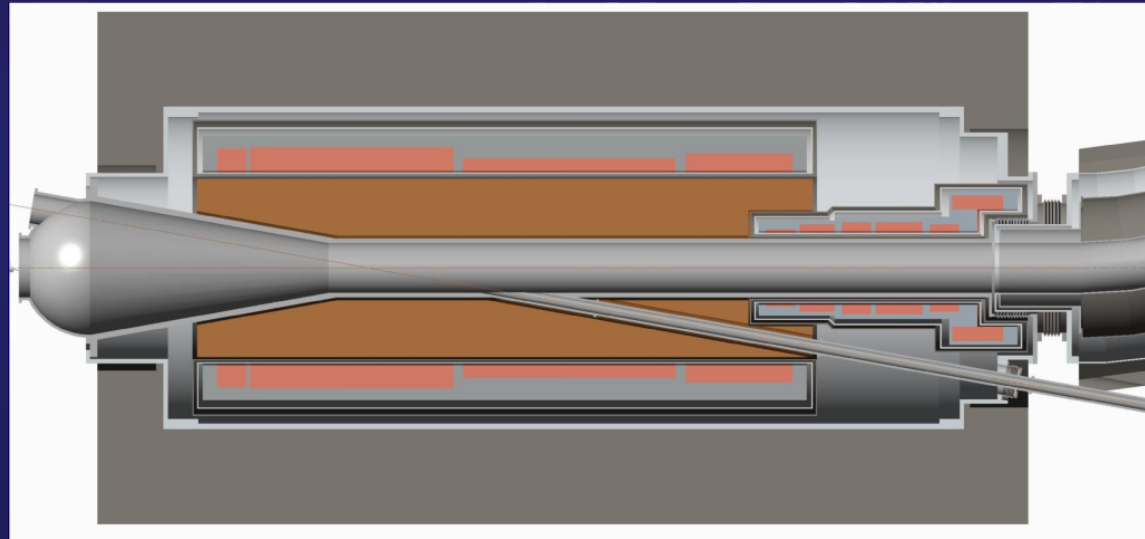
# COMET-UK: Fast Control & DAQ

- Full system designed by UK for Phase-I physics use and serves as critical Phase-II R&D.
- Produced FCT boards.
- Central processing to use FC7 boards designed in UK for CMS Upgrade.
- Need new dedicated RA effort for software and debugging.
- Now transferring trigger work and most hardware and subdetector-specific obligations to international partners.
- But FC & DAQ central deployment and debugging effort must remain UK-led if further delays are to be avoided.



# COMET-UK: Beam Line

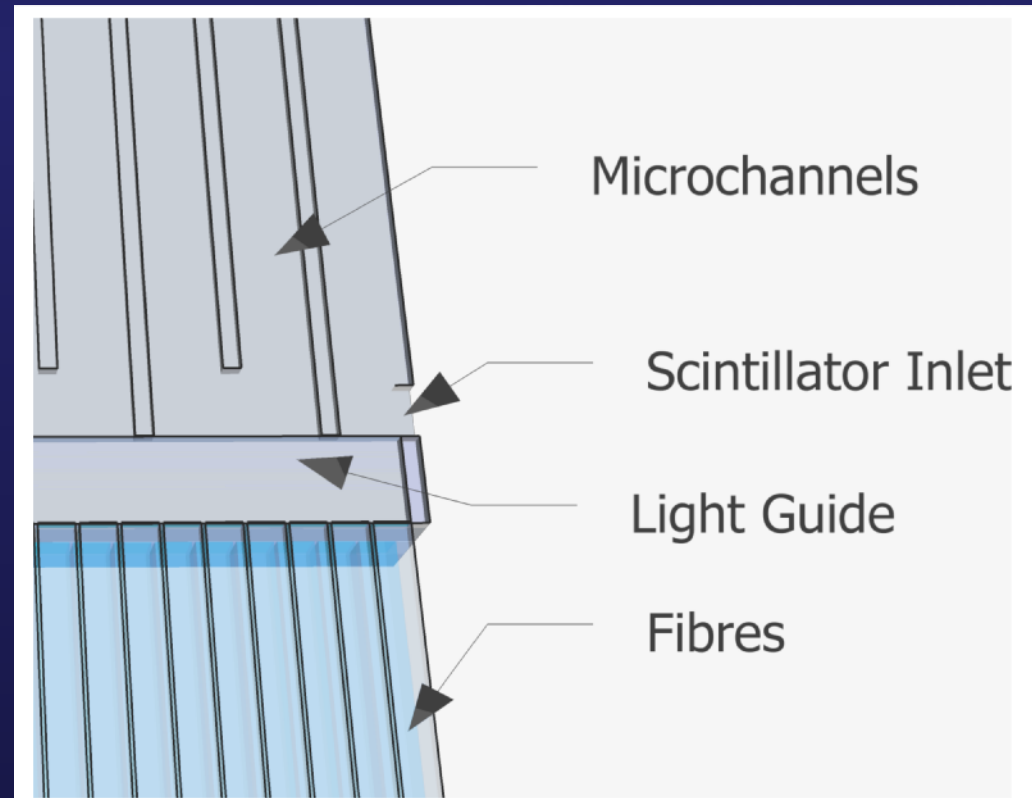
- RAL TD high-power targets group ideally positioned for proton beam target and infrastructure work required by COMET.
  - Strong synergies with other ongoing work.
  - Highly-valued contributions to T2K at J-PARC.
- Contributed to significant design decisions.
  - New choice of target material for Phase-I.
  - Common shielding design for Phase-I and Phase-II.
  - Remote handling and infrastructure layout.
- Input is recognised as critically important by KEK/J-PARC.
- Now providing conceptual designs for Phase-I instead of full technical designs and target hardware.





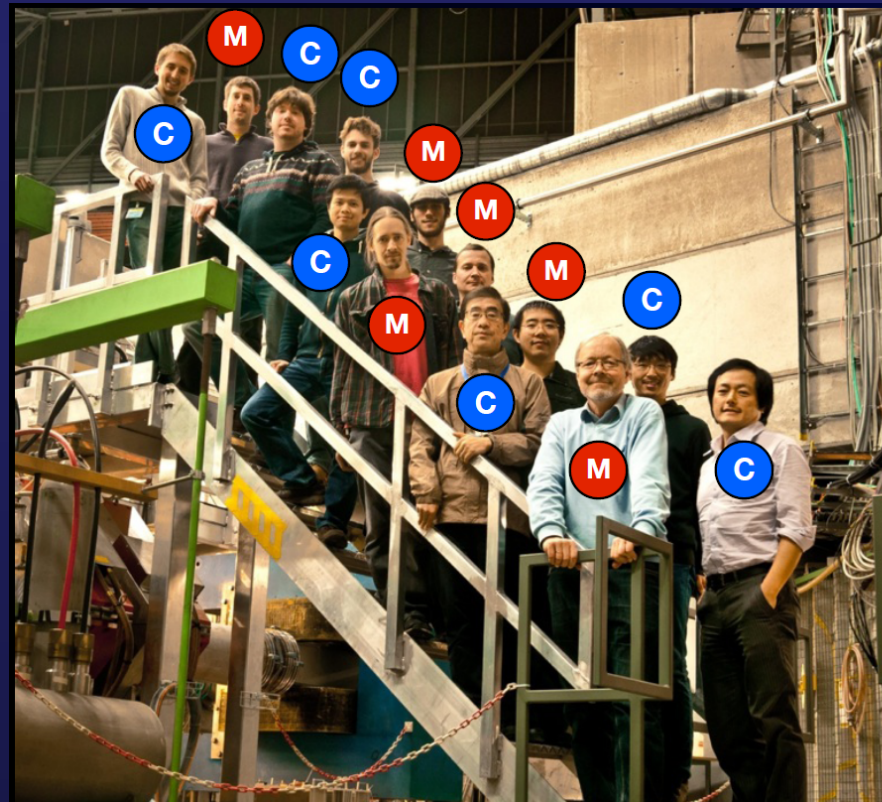
# COMET-UK: Beam Monitoring

- Monitoring of high-intensity beams is critical.
  - Large emittance muon beams are particularly challenging.
  - Highlighted by International External Review panel (Jan 2014) as needing addressing by the collaboration.
- Perform R&D using Phase-I, while contributing to physics (i.e. background characterisation).
- Pursuing novel silicon micro-channel liquid scintillator technology.
  - Led by Manchester group.
  - Seeking funding in UK and Europe.
- Utilising UK strengths in accelerator and detector technologies.



# COMET-UK: AICap Experiment

- Joint experiment at PSI between COMET and Mu2E collaborations.
- Important measurements for high-intensity muon physics.
  - (Z-dependent particle emission from muonic atoms).
- First run in December 2013.
- Strong UK involvement (Imperial & UCL).
- Physics publications expected later this year and UK PhD theses.
- Further data-taking with enhanced set-up and continued physics exploitation planned for 2015.



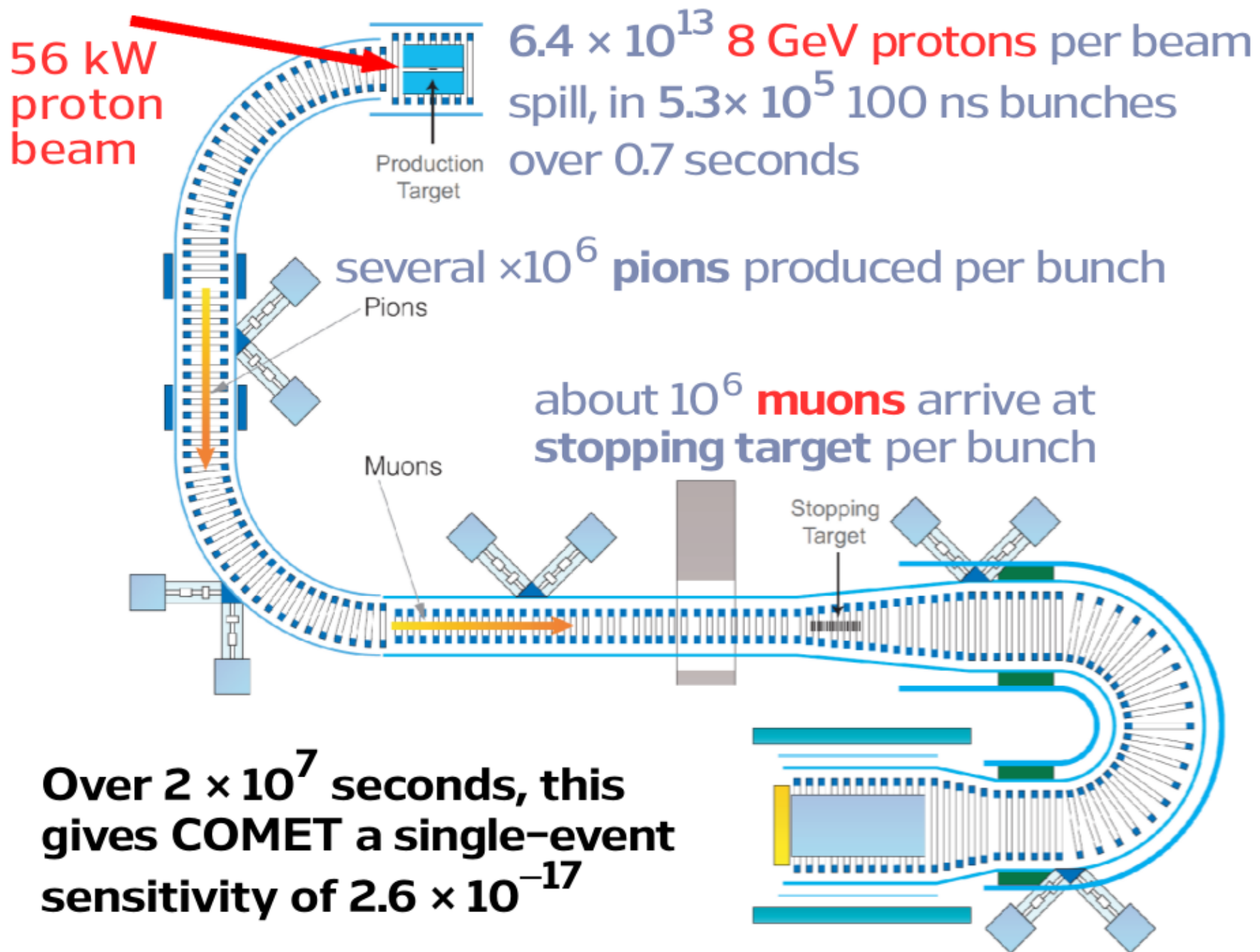
AICap@PSI

Dec. 13

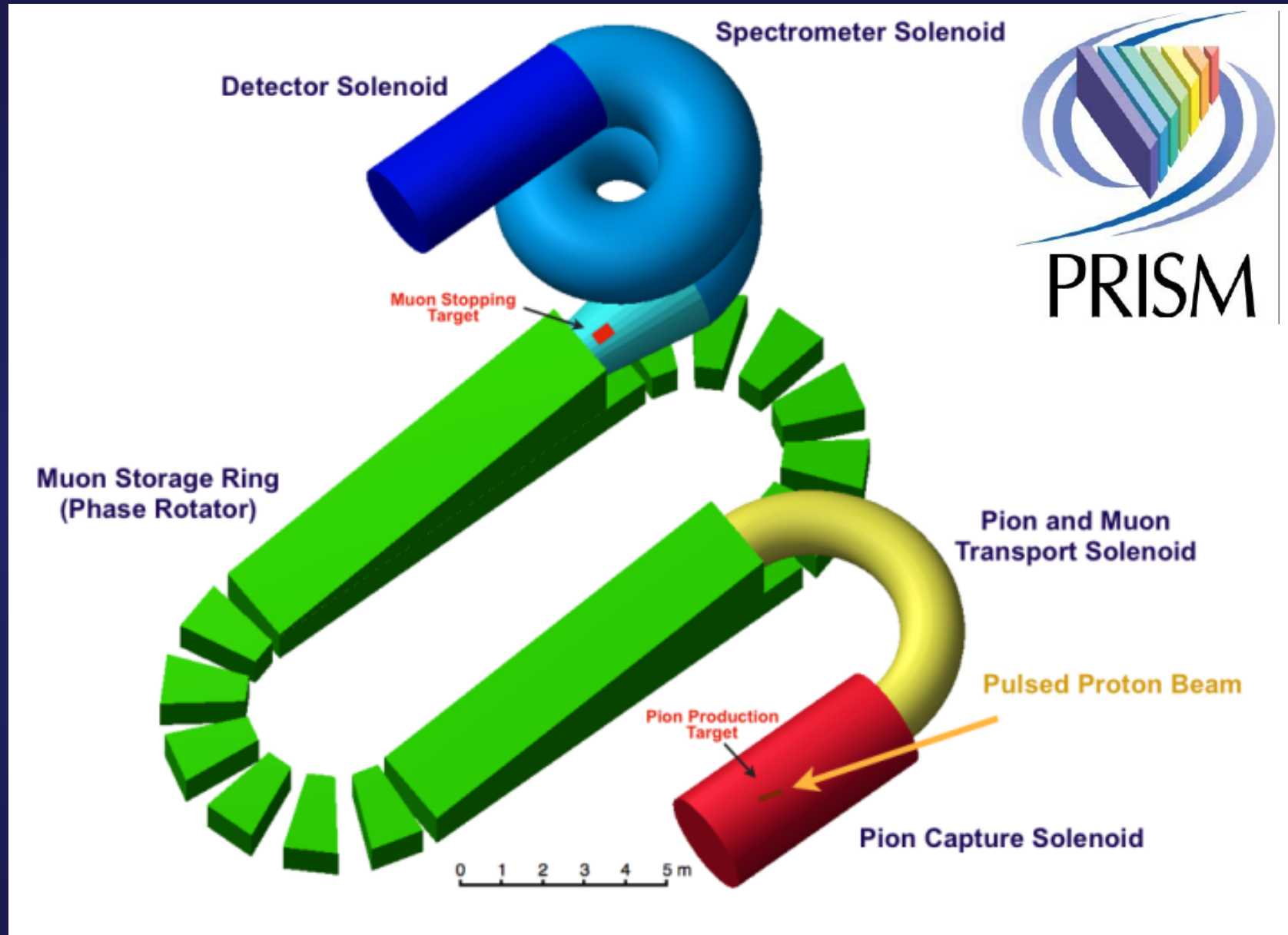
**C** COMET  
**M** Mu2E



# COMET Phase-II



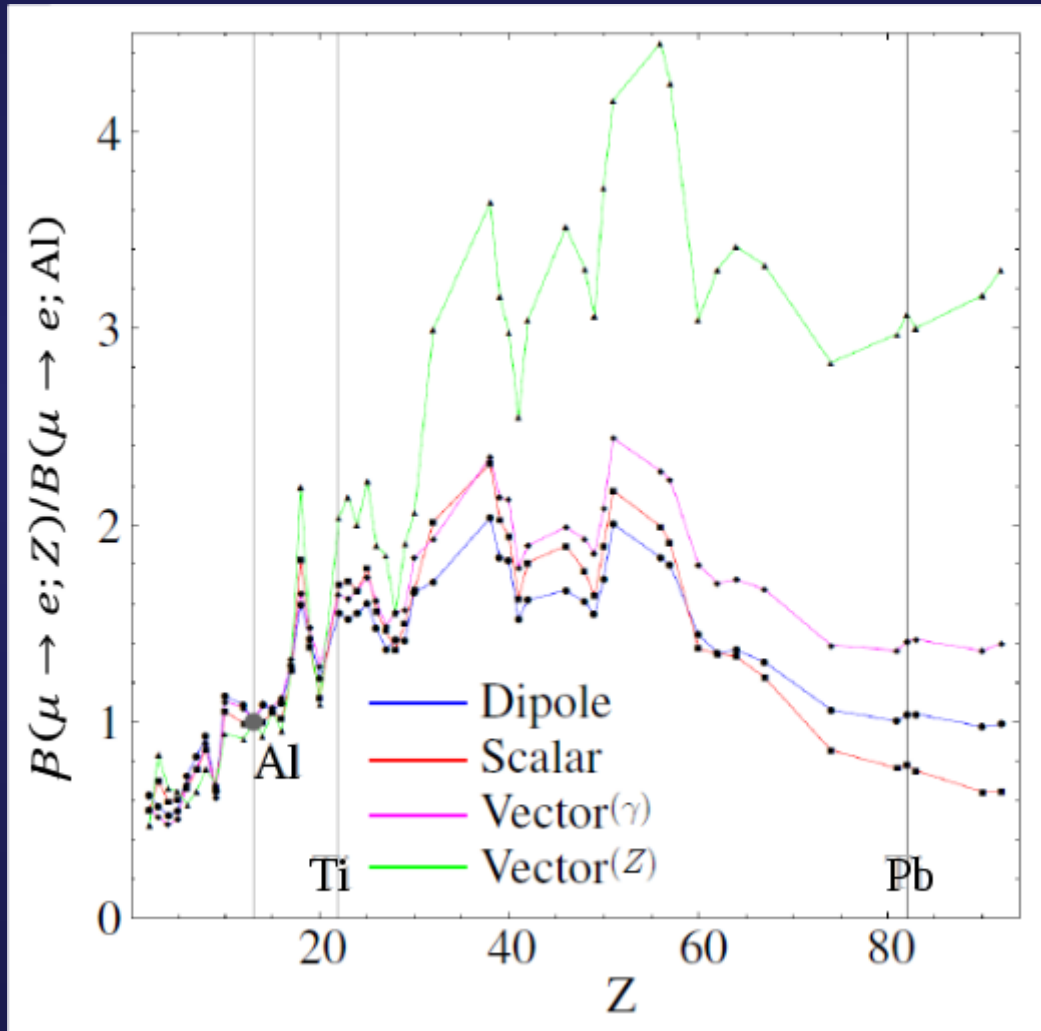
# Phase Rotated Intense Source of Muons (PRISM)



- PRISM task force led by Jaroslaw Pasternak, Imperial College London.

# PRISM

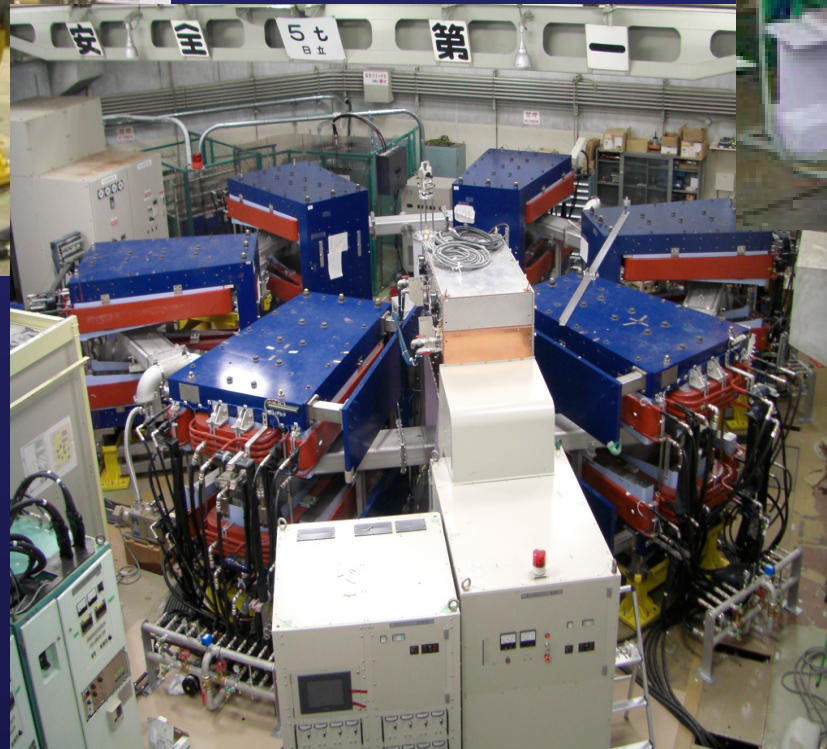
- Sensitivity down to  $10^{-19}$ .
- Branching ratio as a function of  $Z$ .



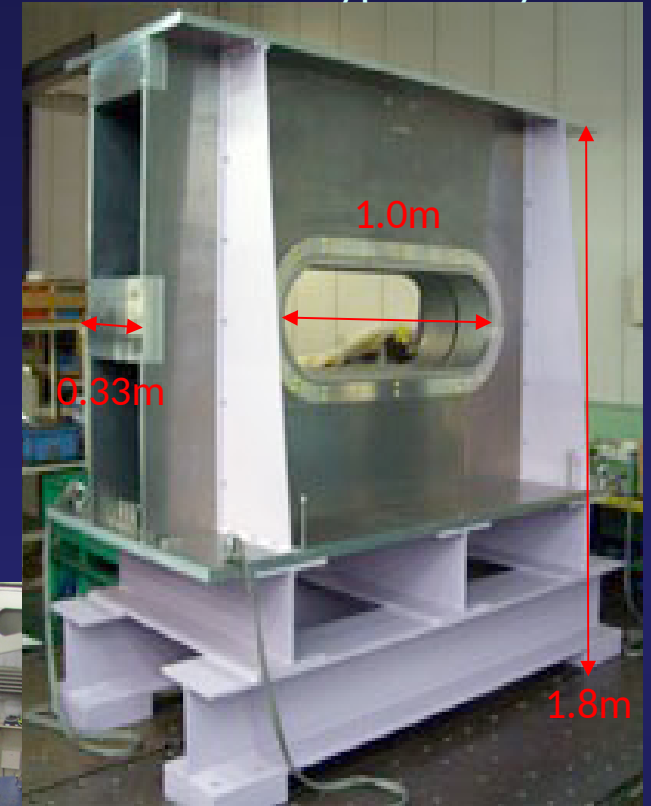
Cirigliano, Kitano,  
Okada and Tuzon,  
arXiv:0904.0957.

- Requires novel accelerator technology.

# PRISM R&D

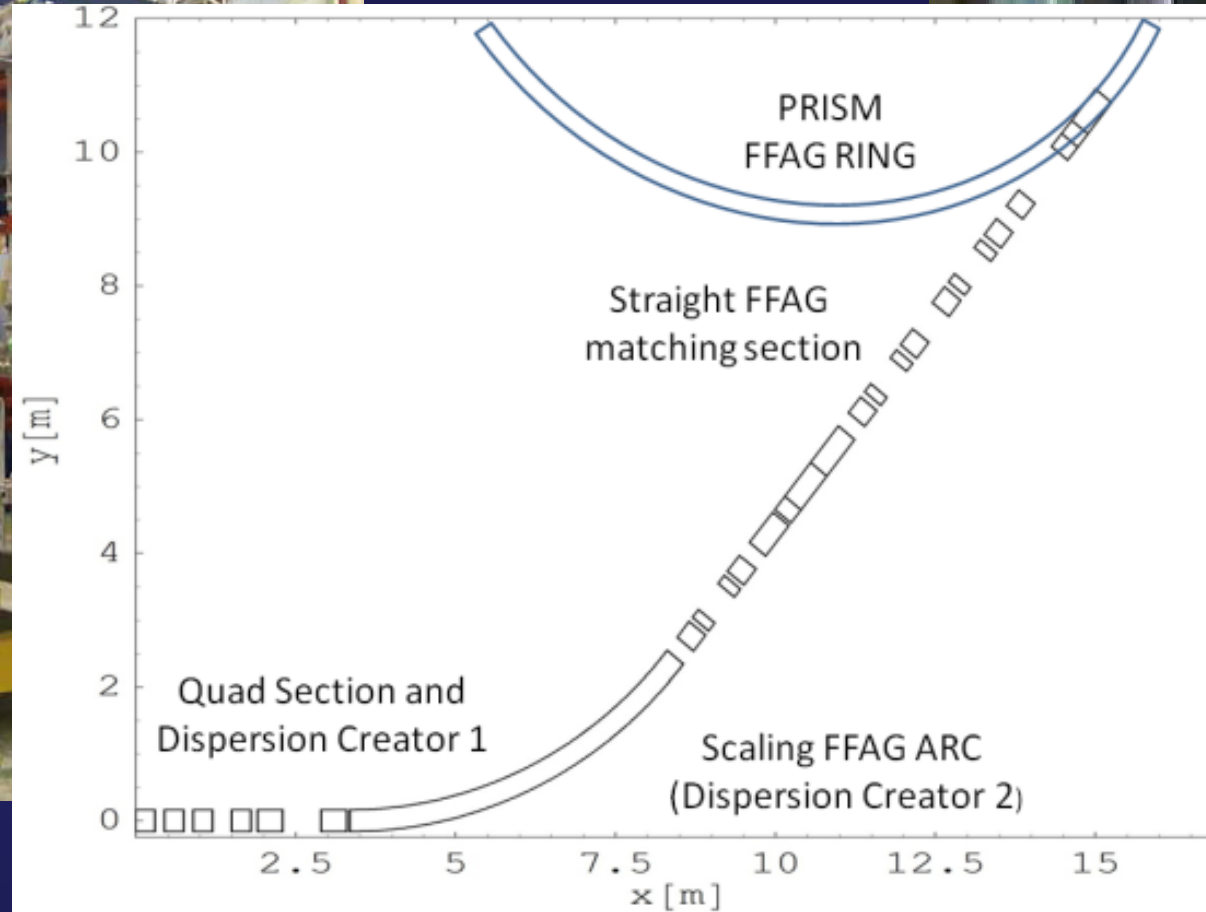
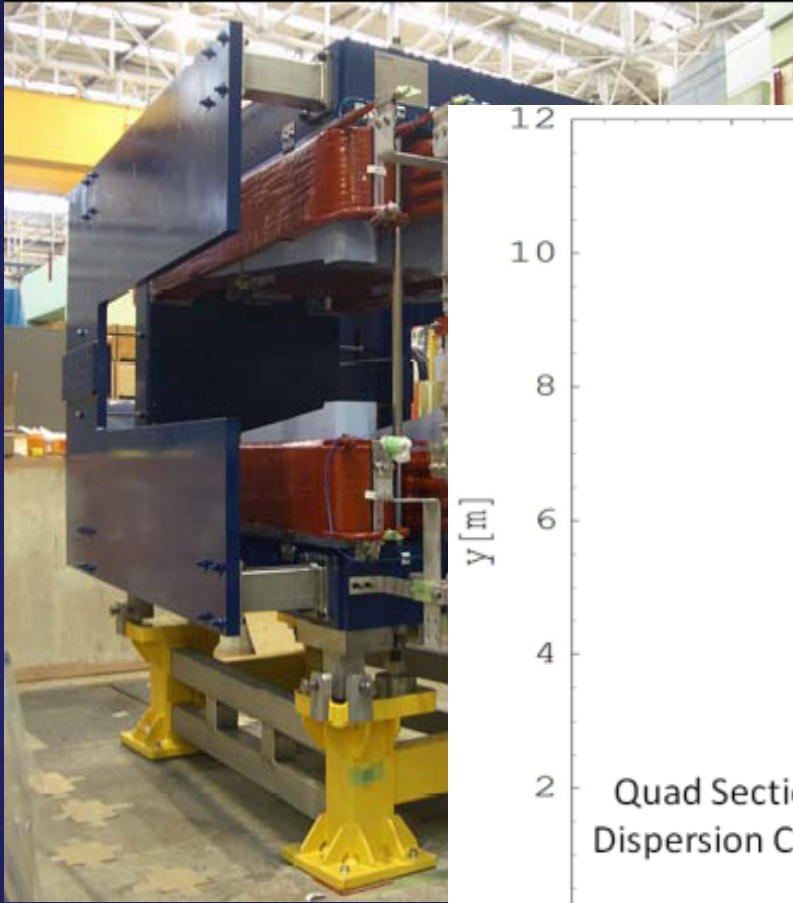


PRISM Prototype Cavity

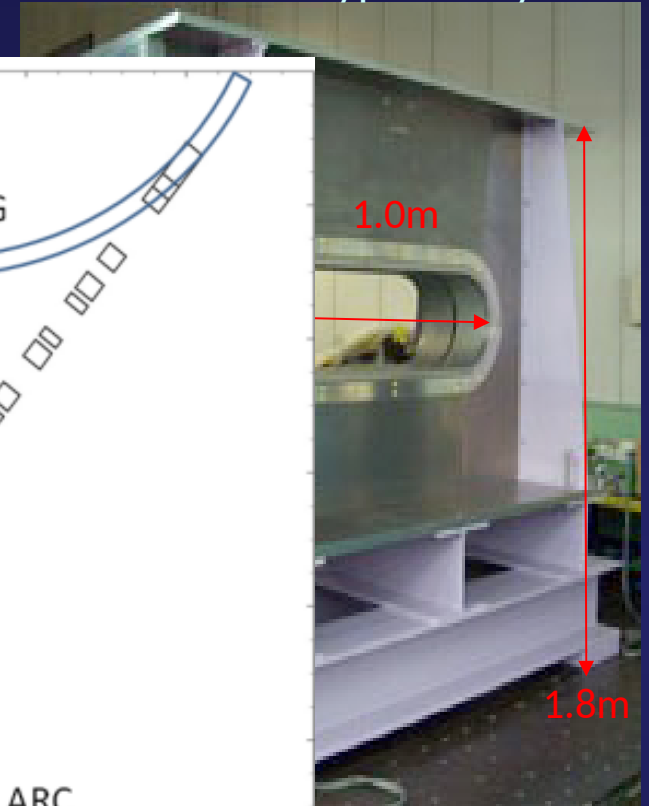




# PRISM R&D



PRISM Prototype Cavity



- Accelerator challenges being addressed by PRISM task force.

# Summary

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- Physics case for muon to electron is strong.
  - $> 10^4$  to  $10^6$  improvement in sensitivity.
  - Strong discovery potential.
- COMET is happening!
  - Commissioning runs / physics data in 2015/16.
- Continuing our strong involvement in driving COMET Phase-I forward.
  - Fully exploit investment for physics from Phase-I.
- Targeting major participation in COMET Phase-II.

Backup slides

# Schedule of COMET Phase-I and Phase-II

	JFY	2013	2014	2015	2016	2017	2018	2019	2020	2021
COMET Phase-I	construction									
	data taking									
COMET Phase-II	construction									
	data taking									

COMET Phase-I :  
2016 ~  
S.E.S.  $\sim 3 \times 10^{-15}$   
(for 1~3 months  
with 3,2 kW proton beam)

COMET Phase-II :  
2020~  
S.E.S.  $\sim 3 \times 10^{-17}$   
(for  $2 \times 10^7$  sec  
with 56 kW proton beam)



# Mu2E Schedule (from May 2013)

