Spontaneous symmetry breaking induced by complex fermion determinant --- yet another success of the complex Langevin method

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Y.I.-Nishimura, work in progress, arXiv:1608.XXXX[hep-lat]

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Introduction

In many interesting physical system, the fermion determinant becomes complex. And its phase plays an important role in the determination of the vacuum.

speculations

- QCD at low temperature and high density exotic fermion condensates
- superstring theory (type IIB matrix model)
- SSB of SO(10) down to SO(4)

wrong results

If one applies the complex Langevin method (CLM), [Parisi '83] [Klauder '83] the fermion determinant may cause the singular drift problem, which is associated with the appearance of small eigenvalues of Dirac op..

Our strategy

We propose to avoid this problem by introducing a fermion bilinear term in the action and modify the Dirac op.

> $S_{\rm f} \to S_{\rm f} + \Delta S_{\rm f} (M)$ $D \to D' (M)$

This modification deforms the eigenvalue distribution.

After that, we get the results for the original system by extrapolating the parameter M to zero using reliable data only.

new criterion (arXiv:1606.07627)

We test this idea in a simple matrix model with rotational SO(4) symmetry.

□ The phase of the fermion determinant induces the SSB of SO(4) down to SO(2).

cf) phase quenched model \longrightarrow no SSB



The obtained results agree well with the predictions obtained by the Gaussian expansion method.

[Nishimura-Okubo-Sugino '05]

Talk plan

- 1. Introduction
- 2. The SO(4) symmetric matrix model
- 3. The application of the CLM to the model
- 4. Results
- 5. Summary

The SO(4)-symmetric matrix model

Partition function

$$Z = \int dX \, \left(\det D\right)^N e^{-S_{\rm b}}$$
boson action

 $S_{\rm b} = \frac{1}{2}N\sum_{-1}^{4} \operatorname{tr}\left(X_{\mu}^{2}\right)$

fermion determinant

 X_{μ} : N imes N Hermitian matrices $\mu = 1 \sim 4$

 $\Gamma^{\mu}: 2 \times 2$ gamma matrices

ermion determinant

$$D = \sum_{i=1}^{4} \Gamma^{\mu} \otimes X_{\mu}$$

$$\Gamma^{\mu} = \begin{cases} i\sigma_{i} & \text{for } \mu = i = 1, 2, 3 \\ \mathbf{1}_{2 \times 2} & \text{for } \mu = 4 \end{cases}$$

 $\mu = 1$ This is exactly "massless" system

det D is complex

- SO(4) rotational symmetry
- The fermion determinant induces the SSB of SO(4)

suggested by the Gaussian expansion method

[Nishimura-Okubo-Sugino '05]

[Nishimura '02]

Order parameters for the SSB of SO(4)

In order to see the SSB

1. introduce a external field in the action to break SO(4) sym.

$$S_{\rm b} \to S_{\rm b+\epsilon} = S_{\rm b} + \frac{N}{2} \sum_{\mu=1}^{4} \epsilon m_{\mu} \operatorname{tr} \left(X_{\mu}^2 \right) \qquad \text{Here we chose} \\ m_{\mu} = (1, 2, 4, 8)$$

2.
$$N \rightarrow \infty$$

- 3. extrapolate ϵ to zero.
- define the order parameters

$$\langle \lambda_{\mu} \rangle = \lim_{\epsilon \to 0} \langle \lambda_{\mu} \rangle_{\epsilon}$$
 here $\langle \lambda_{\mu} \rangle_{\epsilon} = \left\langle \frac{1}{N} \operatorname{tr} \left(X_{\mu}^2 \right) \right\rangle$

In the Gaussian expansion method,

SO(2) case :
$$\langle \lambda_1 \rangle = \langle \lambda_2 \rangle \simeq 2.1, \ \langle \lambda_3 \rangle \simeq 1.0, \ \langle \lambda_4 \rangle \simeq 0.8$$

SO(3) case : $\langle \lambda_1 \rangle = \langle \lambda_2 \rangle = \langle \lambda_3 \rangle \sim 1.75, \ \langle \lambda_4 \rangle \sim 0.75$ comparing the free energy \bigvee
SO(3) case : $\langle \lambda_1 \rangle = \langle \lambda_2 \rangle = \langle \lambda_3 \rangle \sim 1.75, \ \langle \lambda_4 \rangle \sim 0.75$

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The application of the CLM to the model

action

$$S = S_{b+\epsilon} - N \ln \left(\det D\right)$$

complex Langevin equation

$$\frac{dX_{\mu}}{dt} = -\frac{\partial S}{\partial X_{\mu}} + \eta_{\mu} \left(t \right)$$

 $\eta_{\mu}(t): N \times N$ Hermitian matrices

 $\begin{array}{l} X_{\mu} \colon N \times N \text{ Hermitian matrices} \\ \to N \times N \text{ complex matrices} \end{array}$

drift term

 $\sim \sim$

$$\frac{\partial S}{\partial X_{\mu}} = N \left(1 + \epsilon m_{\mu} \right) X_{\mu} - N \operatorname{tr}_{\alpha} \left(\frac{D^{-1} \Gamma^{\mu}}{\mu} \right)$$

zero-eigenvalues of D correspond to a pole.

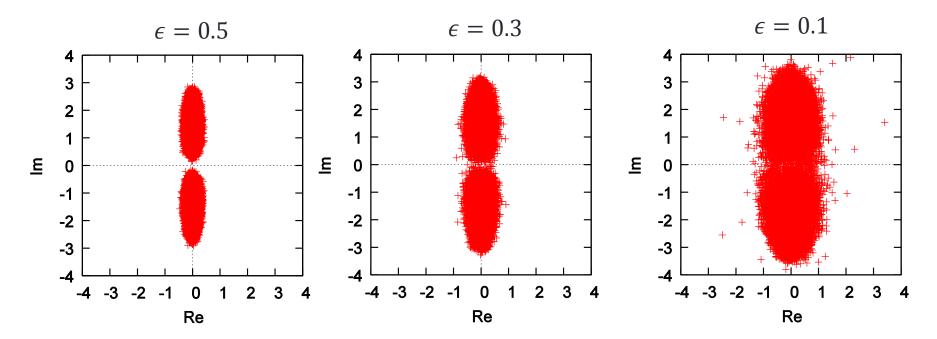
The eigenvalues of D close to zero cause the singular-drift problem.



P(X, t) needs to damp rapidly around the pole in order to justify the CLM.

The singular-drift problem

The eigenvalue distribution of the Dirac op.



Below some critical ϵ , the CLM does not work due to the singular-drift problem.

We try to solve this problem by introducing a fermion bilinear term in the action.

Introduction of a fermion bilinear term

• introduce a fermion bilinear term ΔS_f in the action

$$\Delta S_{\rm f} = -NM_{\mu} {\rm tr} \left(\bar{\psi} \Gamma^{\mu} \psi \right)$$

$$D o D' = \Gamma^\mu \otimes (X_\mu + M_\mu \mathbf{1})$$
 M_μ : parameter

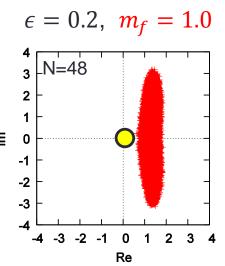
here we used

$$M_{\mu} = (0, 0, 0, m_{\rm f}) \quad \Gamma^{\mu} = \begin{cases} i\sigma_i & \text{for } \mu = i = 1, 2, 3 \, {}^{4} \, {}^{4} \, {}^{-3} \, {}^{-2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{3} \, {}^{-4} \, {}^{-4} \, {}^{-3} \, {}^{-2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{3} \, {}^{-4} \, {}^{-4} \, {}^{-3} \, {}^{-2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{3} \, {}^{2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{3} \, {}^{-4} \, {}^{-4} \, {}^{-3} \, {}^{-2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{-1} \, {}^{1} \, {}^{-2} \, {}^{-1} \, {}^{0} \, {}^{1} \, {}^{2} \, {}^{3} \, {}^{-1} \, {}^{$$

The eigenvalue distribution shifts to the real direction.

 \rightarrow The CLM becomes valid even at small ϵ .

This breaks the SO(4) symmetry minimally down to SO(3).
 First we investigate the SSB of this SO(3).



 $\epsilon = 0.2, \ m_f = 0$

N=48

3 2

-1 -2 -3

<u></u>

Validity of the CLM

n : the magnitude of the drift term

- From the histogram of the magnitude of the drift, the validity of the CLM can be judged.
 - Shimasaki's talk, Nagata's talk [Nagata-Nishimura-Shimasaki '16]

n

 $m_f = 1.0$ $n = \frac{1}{4} \sum_{\mu=1}^{4} \operatorname{tr} \left| \left(\frac{\partial S}{\partial X_{\mu}} \right)_{ij} \right|^{2}$ 0.001 ε=0.1 $\epsilon = 0.2$ 0.0001 $\varphi(n)$: probability distribution of 1e-005 the magnitude *n* (**u**)ф 1e-006 new criterion exponential 1e-007 $\varphi(n)$ should damp faster than exponential. 1e-008 power-law $\varphi(n)$ for $\epsilon \geq 0.2$ damps exponentially 1e-009 (but not for $\epsilon \leq 0.1$). 40000 50000 60000 70000 30000 80000 90000 10000

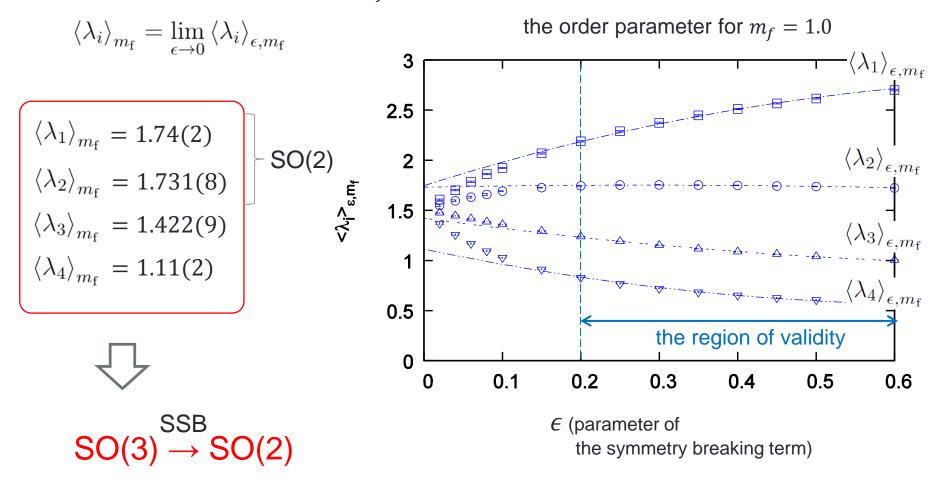
This enables us to extrapolate ϵ using correct data only.

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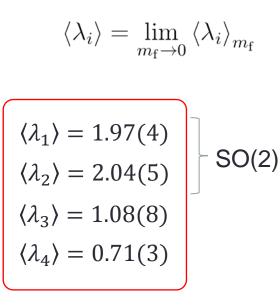
Extrapolation to $\epsilon = 0$ Y.I.-Nishimura, work in progress

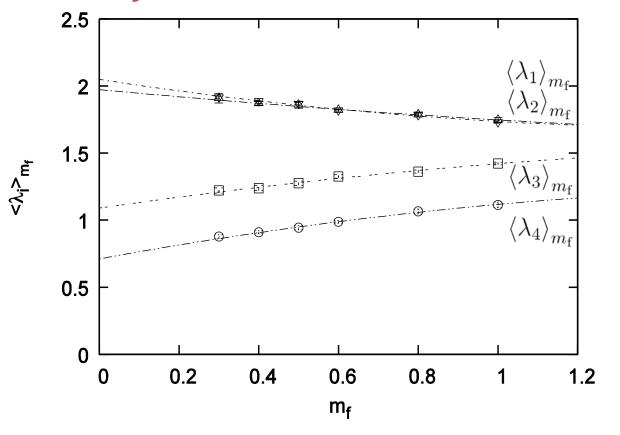
• order parameters for finite m_f



Extrapolation to $m_f = 0$ Y.I.-Nishimura, work in progress

order parameters





SSB from SO(4) to SO(2)

previous results (Gaussian expansion method)

$$\langle \lambda_1 \rangle = \langle \lambda_2 \rangle \simeq 2.1$$

 $\langle \lambda_3 \rangle \simeq 1.0 \quad \langle \lambda_4 \rangle \simeq 0.8$

Summary

For the success of the CLM in the present case, it was crucial to overcome the singular-drift problem.

Our strategy:

we deformed the Dirac op. while maintaining the qualitative feature of vacuum as much as possible.

Then, we extrapolated the deformation parameter to zero.

- In the SO(4)-sym. matrix model, we had to introduce infinitesimal symmetry breaking terms to probe the SSB.
- By using the criterion to justify the method, we can extrapolate the parameter using reliable data only.

> The results agree with the prediction obtained by the GEM.

- This strategy would be useful in finite density QCD at low temperature and high density.
- Since various exotic fermion condensates are expected there, deforming the Dirac op. with the corresponding bilinear term would not disturb the vacuum significantly.

Future Works

Application to the IKKT matrix model which has the SO(10) symmetry.

- The non-perturbative formulation of superstring theory.
- It is expected that 4d space emerges from compact 10d space.

It is suggested that the SO(10) breaks down to SO(3) by the Gaussian expansion method.

It is based on the systematic calculation, and using approximations.

[Nishimura, Okubo, Sugino '05]

We need to study this from first-principle calculation using the CLM.

The finite density QCD

In the high density low temperature phase, introducing an external source as we did here may help reduce the singular-drift problem.

Introduction

Complex action problem

$$Z = \int dx \, e^{-S(x)} \qquad S(x) \in \mathbb{C}$$

- e^{-S} is no longer regarded as the Boltzmann weight factor.
- It appears in finite density QCD, SYM theory, real time dynamics etc.
- The complex Langevin method
 - certain problems often appear in the CLM, which leads to wrong results.
 - The CLM works successfully in finite density QCD and Random matrix theory.

at deconfined phase $T > T_c$ with quark mass [Seiler, Sexty, Stamatescu '12][Sexty '13] [Mollgaard, Splittorff '14]

The complex Langevin method [Parisi '83] [Klauder '83]

For the partition function with a complex action

$$Z = \int dx \, e^{-S(x)} \qquad S(x) \in \mathbb{C}$$

complexify x and consider holomorphic extension of S

$$x \to z = x + iy$$
 $S(x) \to S(x, y)$

the complex Langevin equation

$$\begin{bmatrix} \frac{d}{dt}x = -\operatorname{Re}\partial_{z}S\left(z\right) + \frac{\eta\left(t\right)}{\operatorname{rea}}\\ \frac{d}{dt}y = -\operatorname{Im}\partial_{z}S\left(z\right) \end{bmatrix}$$

t : the Langevin time η : white noise

 $\begin{cases} \langle \eta (t) \rangle = 0 \\ \langle \eta (t) \eta (t') \rangle = 2\delta (t - t') \end{cases}$

The probability distribution satisfies the Fokker-Planck like equation

 $\partial_{t} P(x, y; t) = \partial_{x} \left(\partial_{x} + \operatorname{Re} \left[\partial_{z} S \right] \right) P(x, y; t) + \partial_{y} \operatorname{Im} \left[\partial_{z} S \right] P(x, y; t)$

Criteria for the CLM to be justified

For holomorphic observables O(x+iy), the expectation values is given in the CLM as

$$\langle O(x+iy)\rangle = \int dxdy P_{\rm eq}(x,y) O(x+iy)$$

• The crucial point is that there exists a real and non-negative weight P(x, y) such that

$$\int dx \,\rho(x) O(x) = \int dx dy \, P_{eq}(x, y) O(x + iy)$$

$$\sim e^{-S(x)} : \text{ complex}$$

$$P_{eq}(x, y) \text{ needs to damp rapidly} \quad \left[\begin{array}{c} \text{ in the imaginary direction. } \longrightarrow \text{ gauge cooling} \\ \text{ around singularities of the drift term.} \end{array} \right]$$

[Aarts, James, Seiler, Stamateseu '11] [Nishimura, Shimasaki '15]

• In particular, fermion determinant will cause the singular-drift problem.

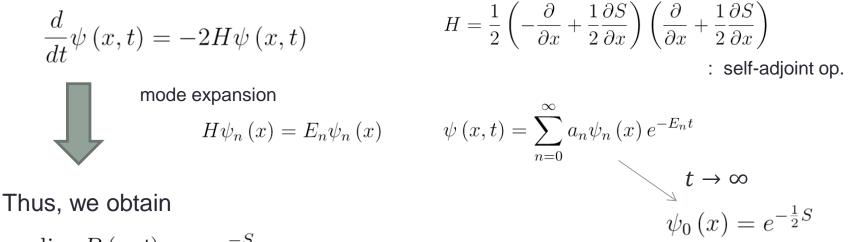
> We propose a new technique to resolve this problem in such a case.

Equivalence to the path integral

We assume

 $P(x,t) = \psi(x,t) e^{-S/2}$

Then, the FP eq. becomes



dominant

 $\lim_{t \to \infty} P\left(x, t\right) = a_0 e^{-S}$ up to normalization

The expectation value

$$\lim_{t \to \infty} \left\langle O\left(x^{(\eta)}\left(t\right)\right) \right\rangle_{\eta} = \lim_{t \to \infty} \int dx \, O\left(x\right) P\left(x;t\right) = \left\langle O\left(x\right) \right\rangle,$$

Proof of the relation

$$\int dx \,\rho\left(x;t\right) O\left(x\right) = \int dx dy \,P\left(x,y;t\right) O\left(x+iy\right)$$

• at t = 0, we can choose $P(x, y; 0) = \rho(x; 0) \delta(y) \longrightarrow$ The relation holds.

• for an arbitrary t, we need to show the below relations

$$\int dx O(x) \rho(x;t) = \int dx O(x;t) \rho(x;0)$$
$$\int dx dy O(x+iy) P(x,y;t) = \int dx dy O(x+iy;t) P(x,y;0)$$

the time-dependent observable is defined by

$$\frac{\partial}{\partial t}O(z;t) = \tilde{L}O(z;t)$$
$$\tilde{L} = \left(\frac{\partial}{\partial z} - \frac{\partial S}{\partial z}\right)\frac{\partial}{\partial z}$$

$$P_{eq}(x, y)$$
 damps rapidly
in the imaginary direction.
around singularities of the drift term.

consider time interpolating function

$$F(t,\tau) = \int dx dy O(x+iy;\tau) P(x,y;t-\tau)$$

• We show that $F(t, \tau)$ is independent of τ

$$\frac{\partial}{\partial \tau} F\left(t,\tau\right) = \int dx dy \frac{\partial}{\partial \tau} O\left(x+iy;\tau\right) P\left(x,y;t-\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x,y;t-\tau\right) \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x,y;t-\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x,y;\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x,y;\tau\right) + \int dx dy O\left(x+iy;\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x,y;\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x+iy;\tau\right) + \int dx dy O\left(x+iy;\tau\right) \frac{\partial}{\partial \tau} P\left(x+iy;\tau\right) + \int dx dy O\left(x+iy;\tau\right) + \int dx dy O$$

$$=\int dxdy\, ilde{L}O\left(x+iy; au
ight)P\left(x,y;t- au
ight)-\int dxdy\,O\left(x+iy; au
ight)L^{ op}P\left(x,y;t- au
ight),$$

partial derivative

 \checkmark

$$= \int dx dy \,\tilde{L}O(x+iy;\tau) P(x,y;t-\tau) - \int dx dy \,LO(x+iy;\tau) P(x,y;t-\tau),$$
$$\downarrow \quad LO(z) = \tilde{L}O(z) \quad \text{for a holomorphic function}$$
$$= 0$$

introduction of a fermion mass term in the 3rd direction

Introducing a mass term to fermions

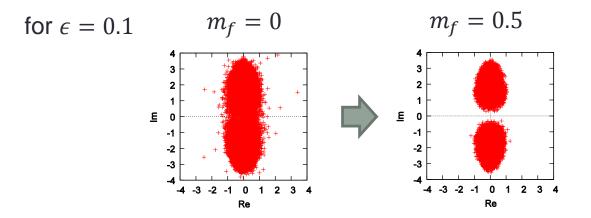
$$D \to D' = \sum_{\mu=1}^{4} \Gamma^{\mu} \otimes (X_{\mu} + \alpha_{\mu} \mathbf{1})$$

Here, we used

 $\alpha_{\mu} = (0,0,m_{\rm f},0)$

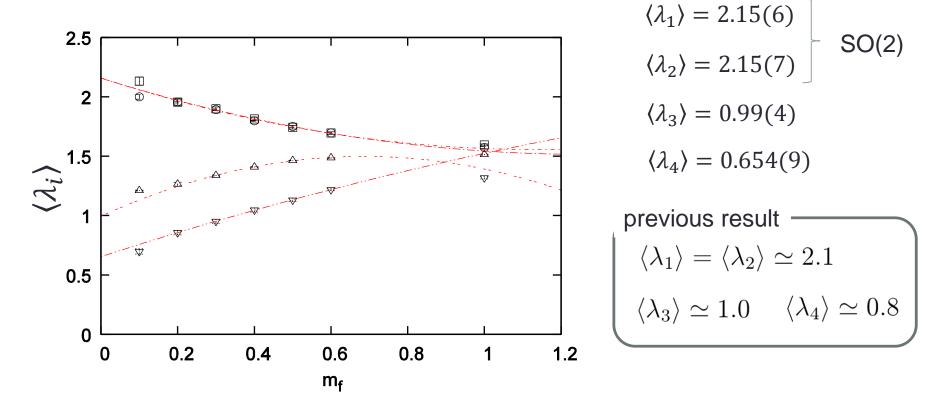
This term makes the eigenvalue distribution of *D* to avoid the pole.

 \rightarrow we can extrapolate the values of $\langle \lambda_i \rangle$ using much smaller ϵ .

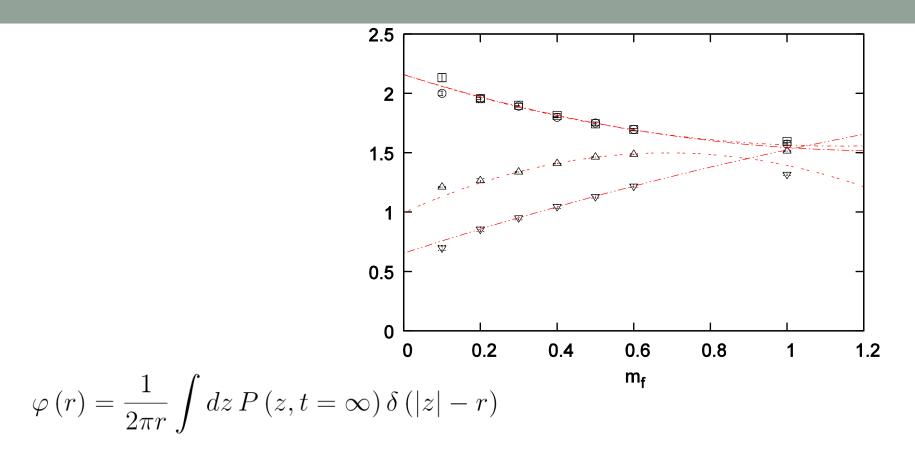


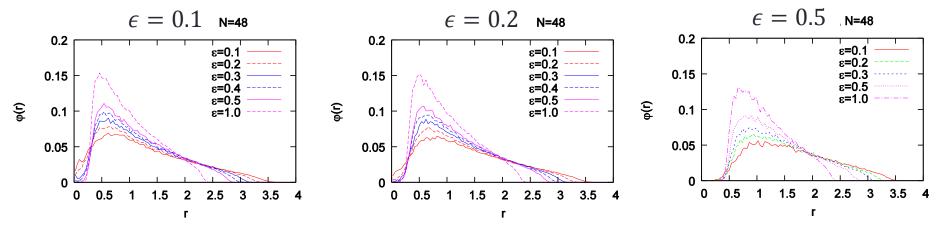
Improvement by means of fermion mass terms

- 1. taking the $\epsilon \rightarrow 0$ limit with m_f fixed.
- 2. taking the $m_f \rightarrow 0$ limit.



The result clearly shows the SSB from SO(4) to SO(2).





Idea of "gauge cooling"

For lattice gauge theory,

Link variables $U_{x,\mu}$

$$SU(N) \to SL(N,\mathbb{C})$$

Considering unitarity norm.

$$rac{1}{N} \mathrm{tr} \left(U U^{\dagger} - \mathbf{1}
ight)$$
 It is no longer zero.

• It is necessary to control the norm to be small.

→ "gauge cooling"

 $U_{x,\mu} \to \Omega_x U_{x,\mu} \Omega_{x+\hat{\mu}}^{-1} \qquad \Omega_x \in SL(N,\mathbb{C})$

• Gauge inv. observables are independent of the gauge cooling.