# Effective Potentials and Radiative Mass Generation

Dumitru Dan Smaranda

11th of January 2017

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CW Mechanisms

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#### Outline

The Effective Action

- Thermodynamics Background
- QFT equivalent
- Computing  $\Gamma_{\text{eff}}$

2 Coleman Weinberg mechanism

- Simple massless scalar model
- SSB induced via loop corrections
- Incorporating the CW mechanism
  - The SM and fine tuning
  - CW in the SM
  - CW in extended models

### The Partition Function

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- The generating field functional is the QFT equivalent to the partition function of a thermal system
- Thermal fluctuations are replaced by quantum fluctuations

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#### Magnetic system in an external field H

Lets look at a magnetic system at a non zero temperature  $T \neq 0$ . The preferred state will be given by that which minimises the Gibbs free energy.

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Lets look at a magnetic system at a non zero temperature  $T \neq 0$ . The preferred state will be given by that which minimises the Gibbs free energy.

For a magnetic system one defines the Helmholtz free energy F(H):

$$Z(H) = \exp(-\beta F(H)) = \int \mathscr{D}s \exp\left(-\beta \int dx (\mathscr{H}(s) - Hs(x))\right)$$

where H is the exterior magnetic field,  $\mathscr{H}(s)$  is the spin energy density, and  $\beta=1/k_BT$  .

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## ${\sf Magnetisation}\ M$

The magnetisation M of a system is given by the 1st moment of the spins across the defined spatial region, and can be found via differentiation of the Helmholtz free energy:

$$-\left.\frac{\partial F}{\partial H}\right|_{\beta=ct} = \frac{1}{\beta}\frac{\partial}{\partial H}(\log Z)$$

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# The Gibbs free Energy G

In turn the Gibbs free energy G is defined as a *Legendre transform* of the Helmholtz free energy:

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If H = 0, G reaches an extremum at a corresponding value of M, and the thermodynamic stable state is the **minimum of** G(M), i.e. G(M) gives a picture of the favoured thermodynamic state that includes all effects of thermal fluctuations.

We now construct the analogous in QFT. Considering a real scalar field  $\phi$  in the presence of an external source J(x), we define a vacuum energy function E(J):

$$Z(J) = \exp(-iE(J)) = \int \mathscr{D}\phi \exp\left(i\int d^4x \mathcal{L}(\phi) + J(x)\phi\right)$$

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Analogous to the thermodynamic case , we now take the functional derivative of E(J) w.r.t.  $J(x)\colon$ 

$$\frac{\delta}{\delta J(x)} E(J) = i \frac{\delta}{\delta J(x)} \log Z = -\frac{\int \mathscr{D}\phi \cdot \phi(x) \exp\left(i \int \mathcal{L} + J\phi\right)}{\int \mathscr{D}\phi \exp\left(i \int \mathcal{L} + J\phi\right)}$$

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#### QFT equivalent

# Constructing the QFT equivalent 2

We treat this value as the thermodynamic variable conjugate to J(x), and through it define the **classical field**  $\phi_{cl}(x)$  :

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$$\Gamma(\phi_{\rm cl}) \equiv -E(J) - \int d^4y J(y) \phi_{\rm cl}$$

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Again, if we have J(x) = 0, the effective action satisfies :

$$\frac{\delta}{\delta\phi_{\rm cl}}\Gamma_{\rm eff}(\phi_{\rm cl})=0$$

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i.e. the solutions, are values for the VEV of  $\phi(x)$  in the stable quantum states of the theory. Therefore by extremising the effective action one finds the exact vacuum state of the QFT, including all effects of quantum corrections

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## The Effective potential $V_{\rm \scriptscriptstyle eff}$

Moreover, the effective action can be written as a derivative expansion around the classical field:

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$$\Gamma_{\rm eff}(\phi_{\rm cl}) = \int d^4x \left[-V_{\rm eff}(\phi_{\rm cl}) + \frac{1}{2} (\partial_\mu \phi_{\rm cl})^2 Z(\phi_{\rm cl}) + \ldots\right]$$

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Furthermore, since we're only interested in cases in which the VEV is translational invariant, the minimisation of  $\Gamma_{\rm eff}$  reduces to:

$$\Gamma_{\rm eff}(\phi_{\rm cl}) = -(VT) \cdot V_{\rm eff}(\phi_{\rm cl}) \qquad \Rightarrow \qquad \frac{\partial}{\partial \phi_{\rm cl}} V_{\rm eff}(\phi_{\rm cl}) = 0$$

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# Finding the effective action

Using the concept of functional derivatives, and by expanding around the classical field by defining  $\phi(x)=\phi_{\rm cl}(x)+\eta(x)$ , one can find the effective action:

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$$\begin{split} \Gamma_{\rm eff} &= \int d^4 x \mathcal{L}_1(\phi_{\rm cl}) + \\ &+ \frac{i}{2} \log \! \left( \det \! \left( - \frac{\delta \mathcal{L}_1}{\delta \phi \delta \phi} \right) \right) - i \cdot (\text{connected diagrams}) + \end{split}$$

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#### Computing $\Gamma_{eff}$

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Implicitly, if we know the effective action we know the effective potential

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## Toy model

Let's now try and look at a simple model and actually try and compute the effective potential, consider the following **massless** self interacting scalar theory:

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$$\mathcal{L} = \frac{1}{2} (\partial_{\mu} \phi)^2 - \frac{\lambda}{4} \phi^4 + \frac{1}{2} A (\partial_{\mu} \phi)^2 - \frac{1}{2} B \phi^2 - \frac{1}{4!} C \phi^4$$

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where,  $A = Z_{\phi} - 1$ ,  $B = Z_m - 1$  and  $C = Z_{\lambda} - 1$  are the usual counterterms.

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#### 1 Loop Order - I

It turns out that computing the functional determinant in the effective action  $\log\left(\det\left(-\frac{\delta \mathcal{L}_1}{\delta\phi\delta\phi}\right)\right)$ , is equivalent with finding the 1-loop approximation to the effective potential.

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This in turn can be calculated for our  $\phi^4$  theory via the polygon graph series:



Figure: Polygon Graph Corrections to the potential

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### 1 Loop Order - II

Introducing the counter-terms as well , the 1LO effective potential is:

$$V_{\rm eff}(\phi_{\rm cl}) = \frac{\lambda}{4!}\phi_{\rm cl}^4 - \frac{1}{2}B\phi_{\rm cl}^2 - \frac{1}{4!}C\phi_{\rm cl}^4 + i\int \frac{d^dk}{(2\pi)^d}\sum_{n=1}^{\infty}\frac{1}{2n}\left(\frac{\lambda\phi_{\rm cl}^2/2}{k^2 + i\epsilon}\right)^n$$

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Which, after summing the series provides:

$$V_{\rm eff}(\phi_{\rm cl}) = \frac{\lambda}{4!}\phi_{\rm cl}^4 - \frac{1}{2}B\phi_{\rm cl}^2 - \frac{1}{4!}C\phi_{\rm cl}^4 + \frac{1}{2}\int \frac{d^dk}{(2\pi)^d}\log\left(1 + \frac{\lambda\phi_{\rm cl}^2/2}{k^2 + i\epsilon}\right)$$

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## 1 Loop Order - III

Using the usual bag of tricks for renormalization, be it dimReg or momentum cut-off schemes, we impose the renormalisation conditions at an arbitrary **mass scale** M:

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We find:

$$V_{\rm eff}(\phi_{\rm cl}) = \frac{\lambda}{4!} \phi_{\rm cl}^4 + \frac{\lambda^2 \phi_{\rm cl}^4}{256\pi^2} (\log \frac{\phi_{\rm cl}^2}{M^2} - \frac{25}{6})$$

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### Coleman - Weinberg Mechanism I

Looking at the potential more closely :

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Image: A matrix

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#### Coleman - Weinberg Mechanism II



Our loop corrections have made our potential unstable, i.e. the quantum corrections induce SSB.

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#### But why do we care?

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From S.M. Higgs potential:

$$V(H) = \mu |H|^2 + \lambda |H|^4$$

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after applying the quantum corrections, mainly the dominant top quark correction:



we find the corrected Higgs mass squared (up to one loop):

$$m_H^2=\mu^2-\frac{y_t^2}{8\pi^2}\Lambda_{\rm UV}^2$$

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#### What can we do?

This is where the CW mechanism comes in. It can provide a naturally small mass scale via radiative generation!

## CW mechanism in the Standard Model I

In their original paper, Coleman and Weinberg assumed a Weinberg-Salam theory of leptons, and based on the W, Z masses it predicted a Higgs Mass:

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## CW mechanism in the Standard Model II

It gets even worse when you add in the top corrections , which make the Higgs mass squared negative!

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So, What can we do?

As per usual in BSM physics, you add in more stuff!

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#### Neutrino Scalar extension to the SM - I

For example, following Meissner and Nicolai (arXiv:hep-th/0612165) we can look at a scalar extension to the SM which couples with neutrinos via a Yukawa interaction, and to the Higgs field via a quadratic coupling:

$$\begin{split} \mathcal{L} = & (\bar{L}^i \Phi Y^E_{ij} E^j + \bar{Q}^i \epsilon \Phi^* Y^D_{ij} D^j + \bar{Q}^i \epsilon \Phi^* Y^U_{ij} U^j + \\ & + \bar{L}^i \epsilon \Phi^* Y^\nu_{ij} \nu^j_R + \phi(\nu^i_R)^T \mathcal{C} Y^M_{ij} \nu^j_R + h.c.) - \\ & - \frac{\lambda_1}{4} (\Phi^\dagger \Phi)^2 - \frac{\lambda_2}{2} \phi(\Phi^\dagger \Phi) - \frac{\lambda_3}{4} \phi^4 \end{split}$$

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#### Neutrino Scalar extension to the SM - II

Assuming small neutrino masses, discarding the  $SU(2)_W \times U(1)_Y$  and  $SU(3)_C$  and keeping the top quark as the dominant fermion contribution, one can find the effective potential:

$$\begin{split} V_{\text{eff}}(H,\phi) &= \frac{\lambda_1 H^4}{4} + \frac{\lambda_2 H^2 \phi^2}{2} + \frac{\lambda_3 \phi^4}{4} + \\ &+ \frac{3}{256\pi^2} (\lambda_1 H^2 + \lambda_2 \phi^2)^2 \ln\left(\frac{\lambda_1 H^2 + \lambda_2 \phi^2}{\nu^2}\right) + \\ &+ \frac{1}{64\pi^2} F_+^2 \ln\left(\frac{F_+}{\nu^2}\right) + \frac{1}{64\pi^2} F_-^2 \ln\left(\frac{F_-}{\nu^2}\right) - \\ &- \frac{3}{16\pi^2} g_t^4 H^4 \ln\left(\frac{H^2}{\nu^2}\right) - \frac{1}{32\pi^2} g_M^4 \phi^4 \ln\left(\frac{\phi^2}{\nu^2}\right) \end{split}$$

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## Neutrino Scalar extension to the SM - III

Since minimising the potential to obtain the VEV of the fields analytically isn't really feasible we turn to a numeric search. Redefining our fields via the auxiliary dimensionless fields  $h, \varphi$ , we find the numerical minimum of the potential:



Figure: The dimensionless effective potential as a function of  $h, \varphi$ 

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we can scan over the free parameter range and find a set of suitable values!

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## Stability and Landau poles - I

After finding the respective set of suitable bare couplings we still need to check the effective couplings for Landau poles or instabilities up to a high scale.

Based on where we consider the cut off scale (e.g. the Planck Scale) the couplings will behave as :


### Stability and Landau poles - II



#### Predicted $\phi$ mass

If the high scale is dictated by the Planck scale, then we end up with a real scalar with a mass of the order:

 $m_{\phi} \approx \mathcal{O}(500 \, \text{GeV})$ 

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No Fine Tuning!

And note that we didn't have to fine tune anything!

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To sum it up:

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- The effective potential incorporates the quantum corrections
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- The CW mechanism failing in the SM implies extra fields
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Future Work will be centred around incorporating the CW mechanism into a  $SO(10)~{\rm GUT}$  model.

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