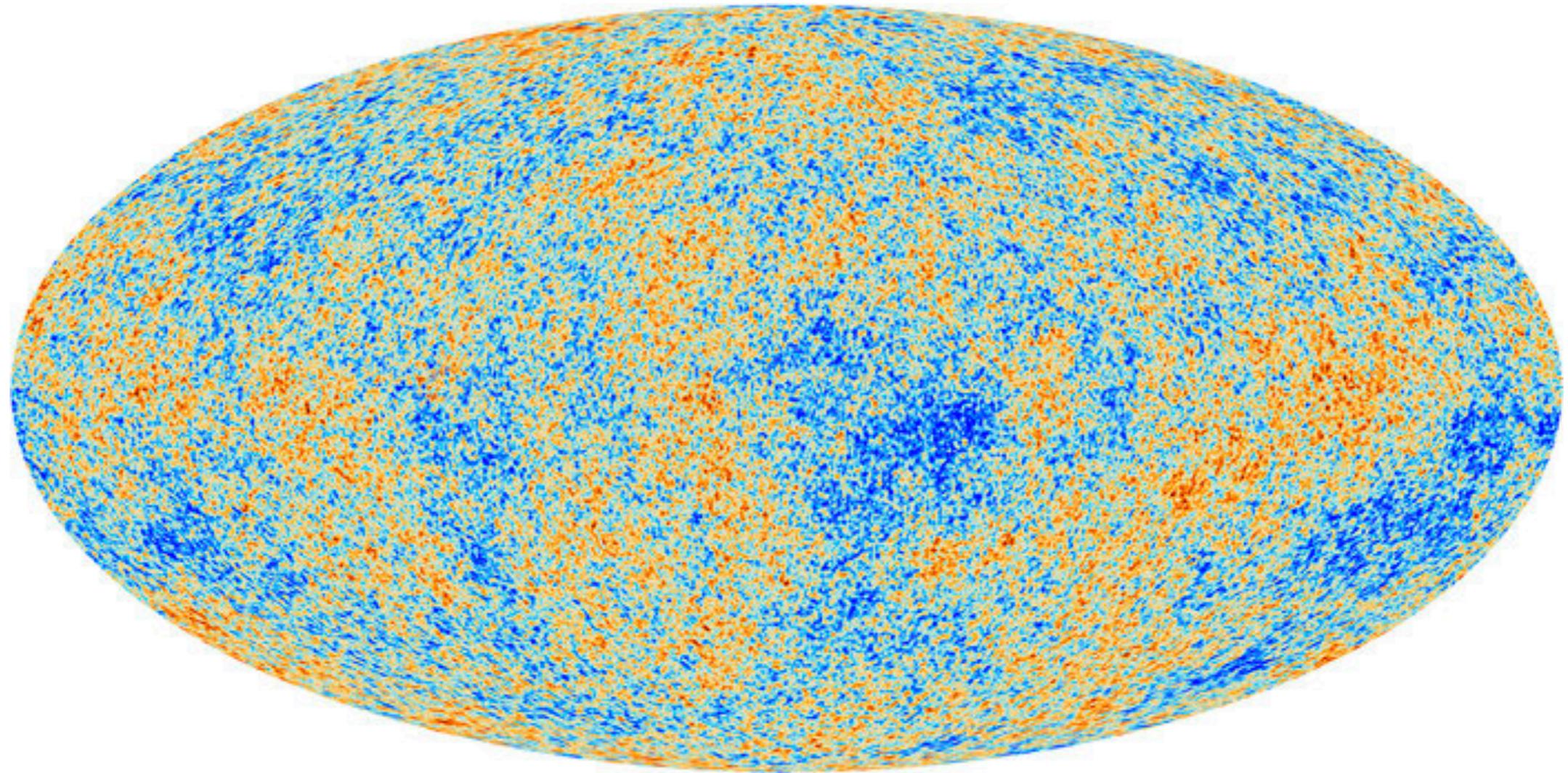
A visualization of the cosmic web, showing a complex network of dark matter filaments and galaxy clusters. The filaments are depicted as thin, greenish-blue lines, while the clusters are represented by dense, multi-colored (red, orange, yellow, and blue) regions. The background is a deep black, making the structures stand out.

# **DARK ENERGY & LARGE-SCALE GALAXY SURVEYS**

**Elisa Chisari**

*Royal Astronomical Society Research Fellow  
University of Oxford*

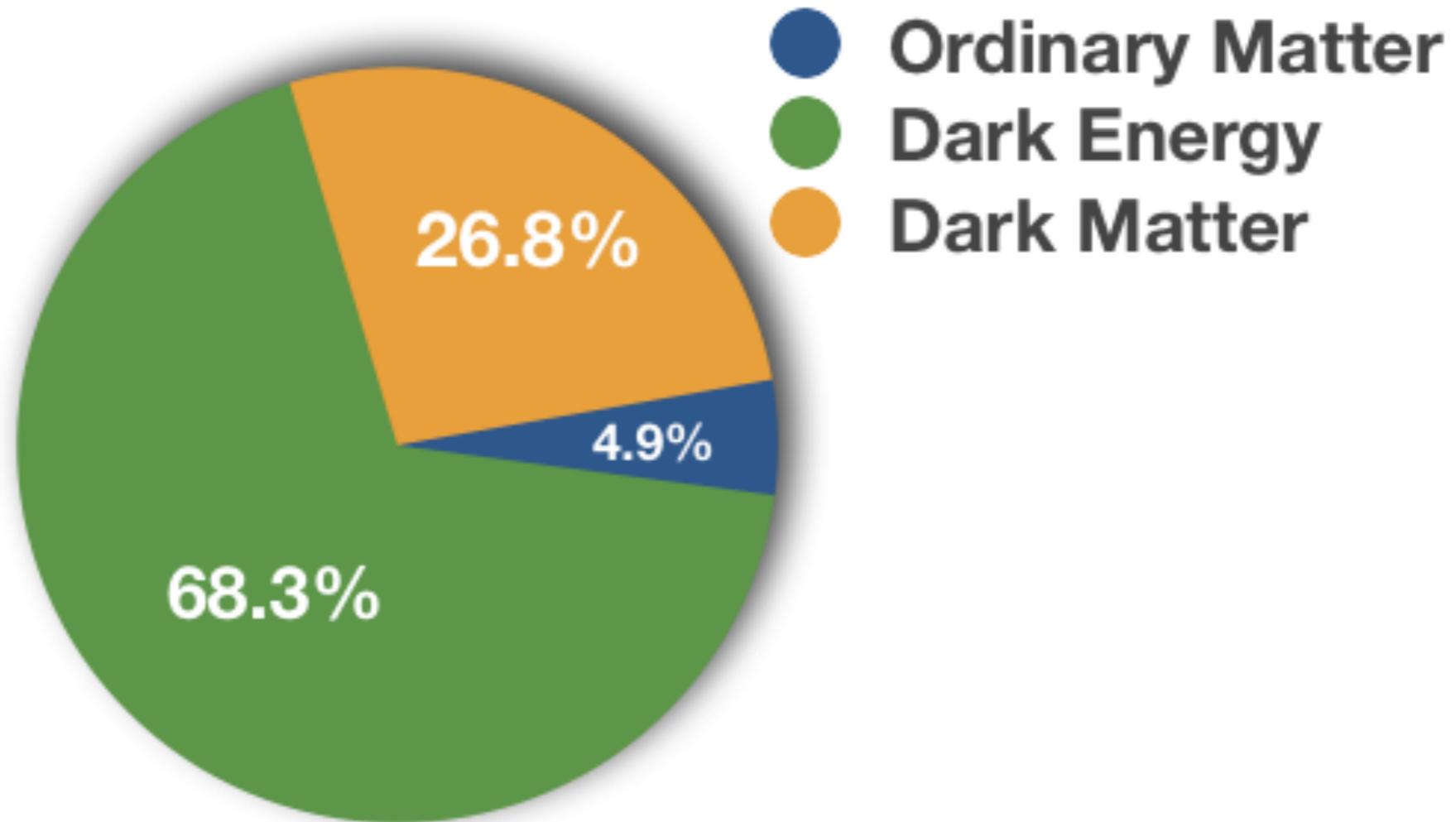
# PRECISION COSMOLOGY



*The cosmic microwave background (CMB)  
400,000 years after the Big Bang*

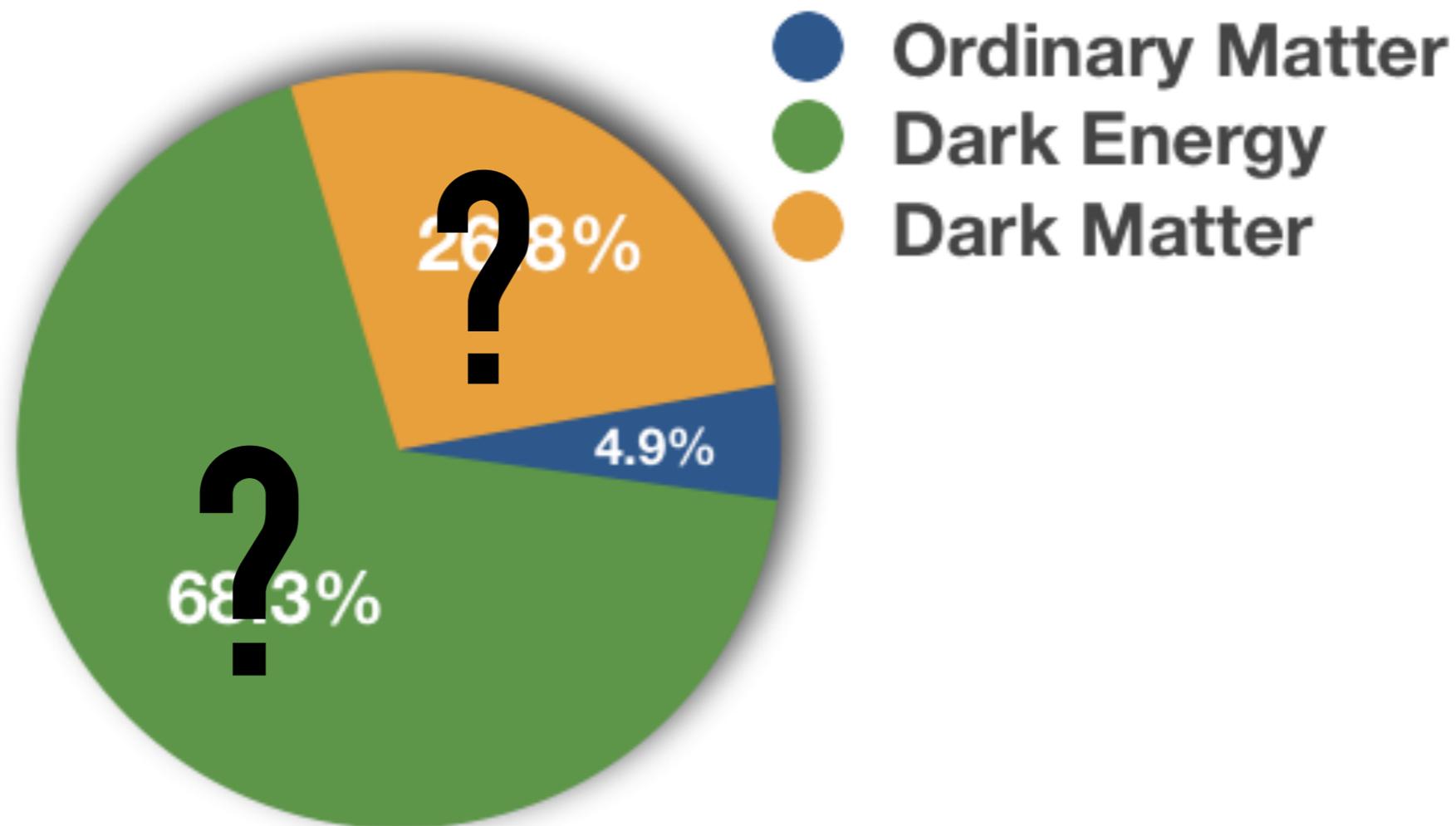
*Image: ESA & the Planck Collaboration*

# PRECISION COSMOLOGY



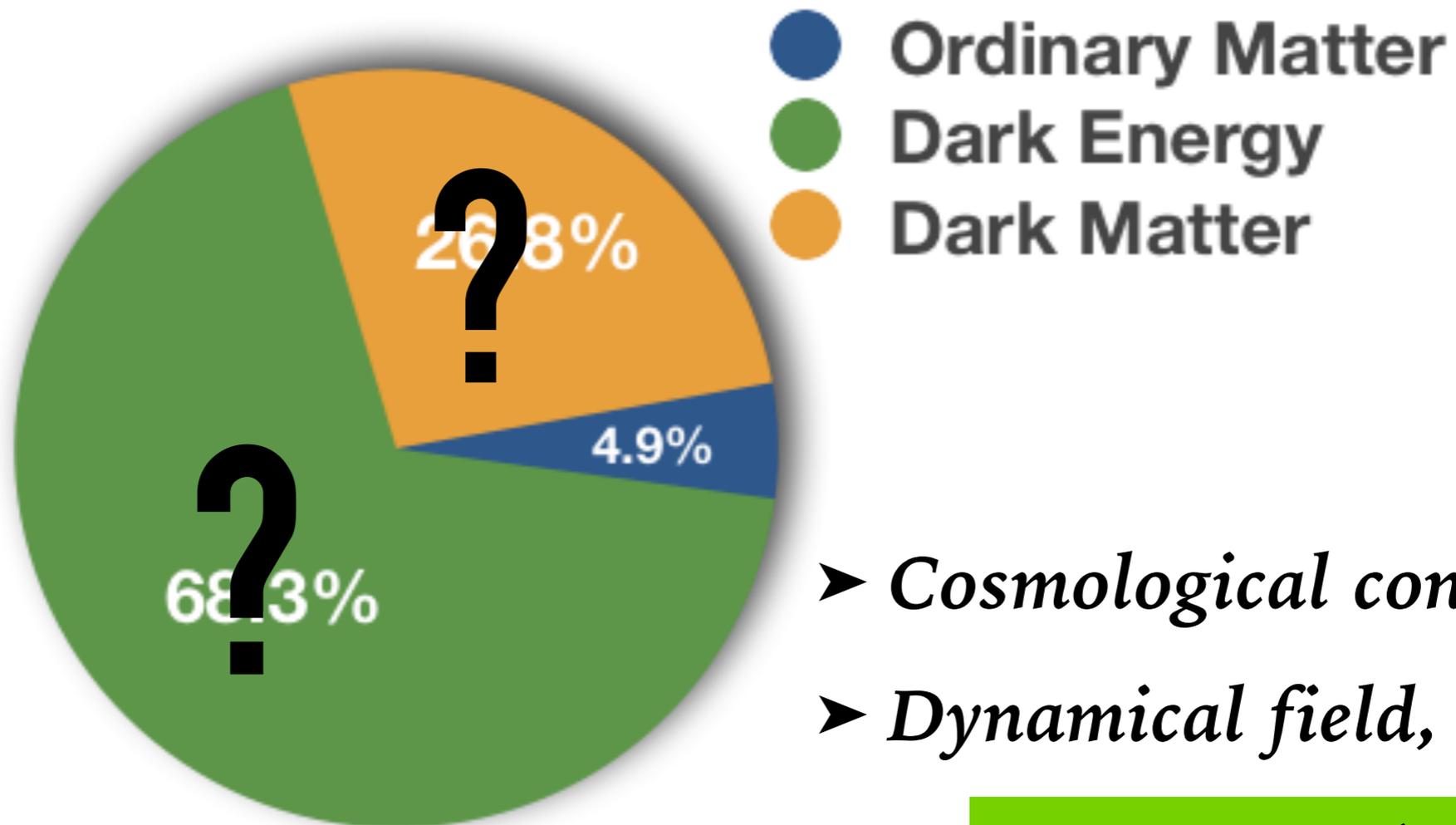
*Planck Collaboration*

# PRECISION COSMOLOGY



*Planck Collaboration*

# PRECISION COSMOLOGY



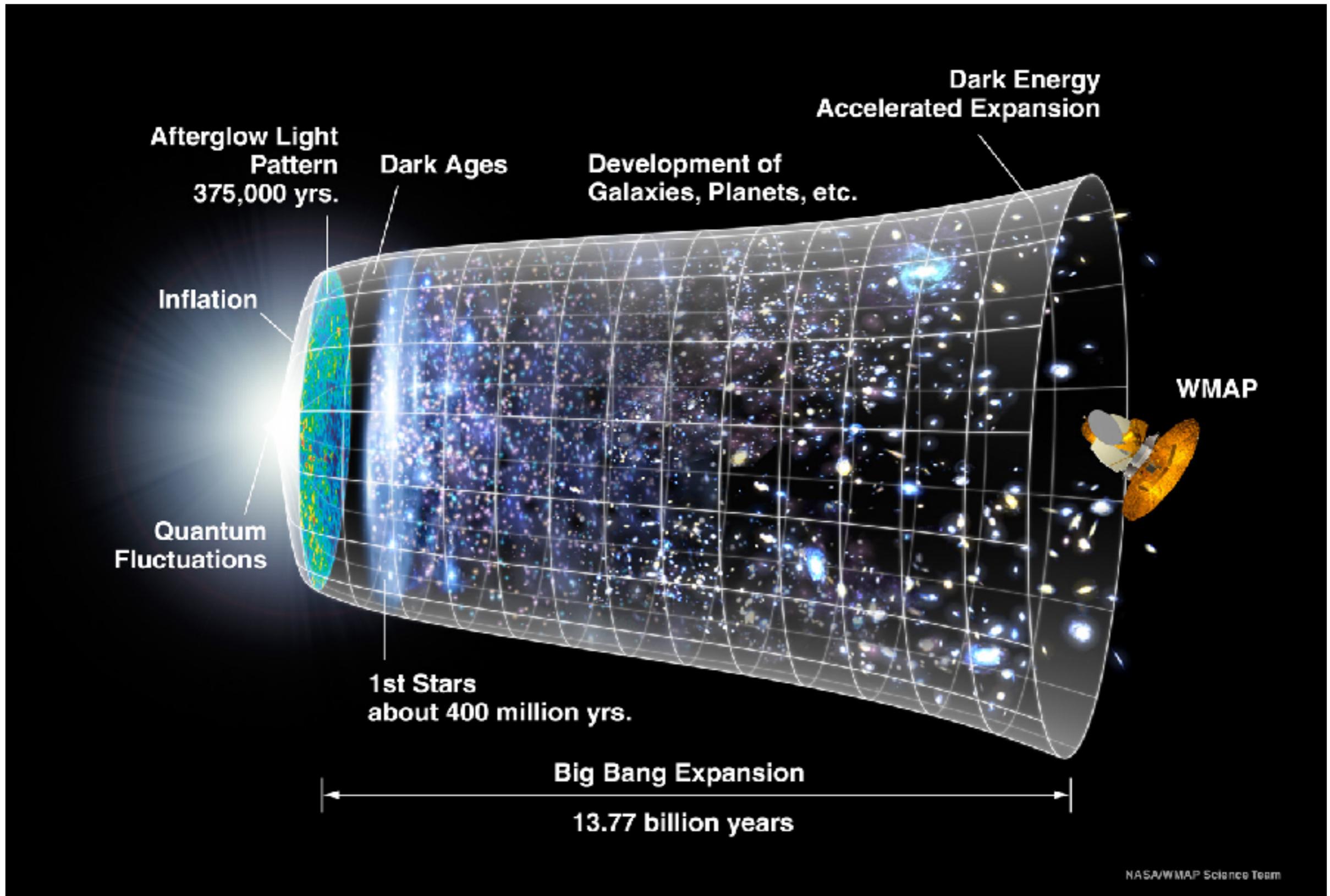
- *Cosmological constant,  $\Lambda$*
- *Dynamical field,  $p = w c^2 \rho$*

$$w = w_0 + w_a(1-a)$$

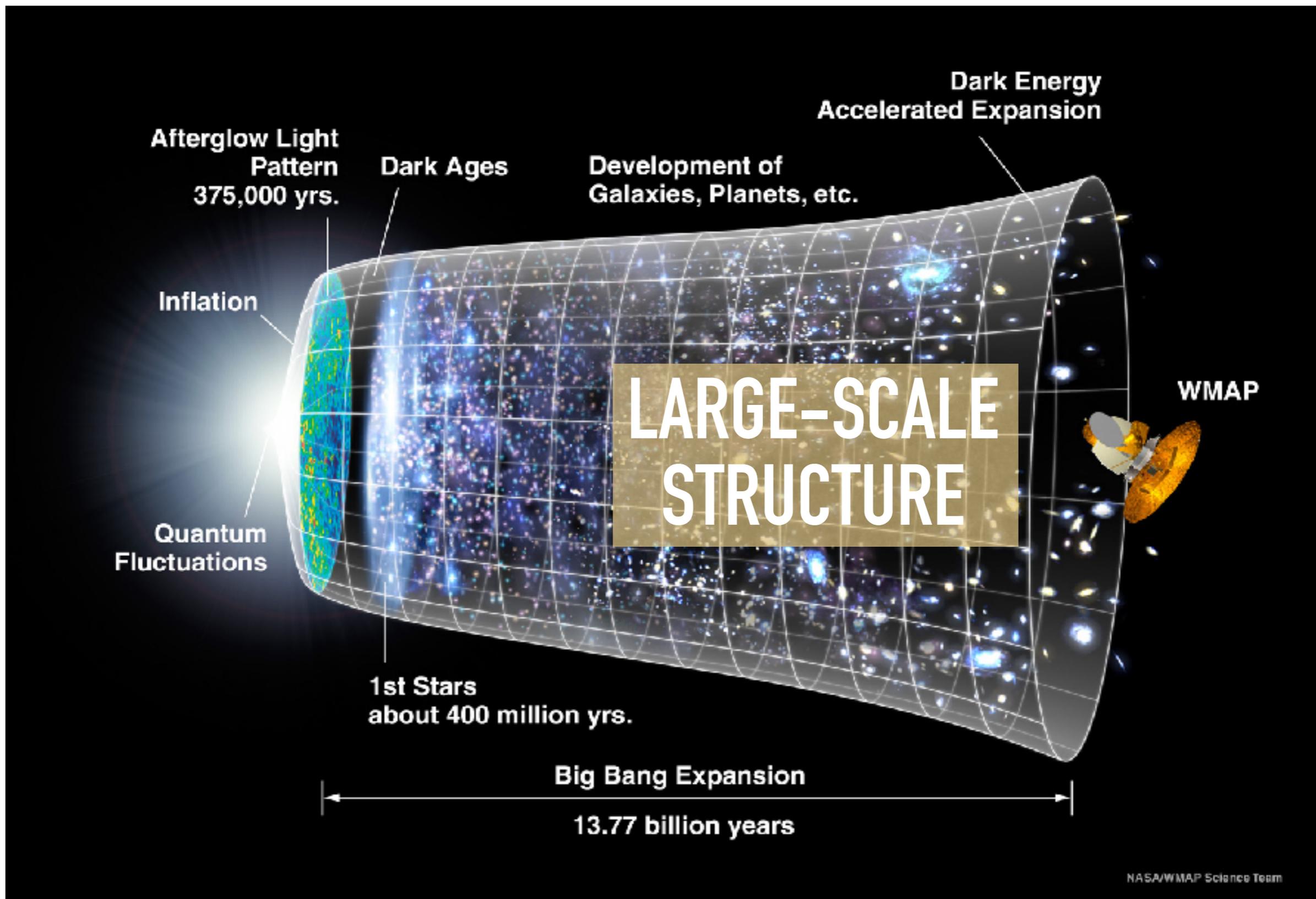
- *Modification of gravity.*

*Planck Collaboration*

# PRECISION COSMOLOGY



# PRECISION COSMOLOGY



# OUTLINE

Precision cosmology with galaxy surveys

*Gravitational lensing*

# OUTLINE

Precision cosmology with galaxy surveys

*Gravitational lensing*

Challenges

*I. Observational*

*II. Theoretical*

# OUTLINE

Precision cosmology with galaxy surveys

*Gravitational lensing*

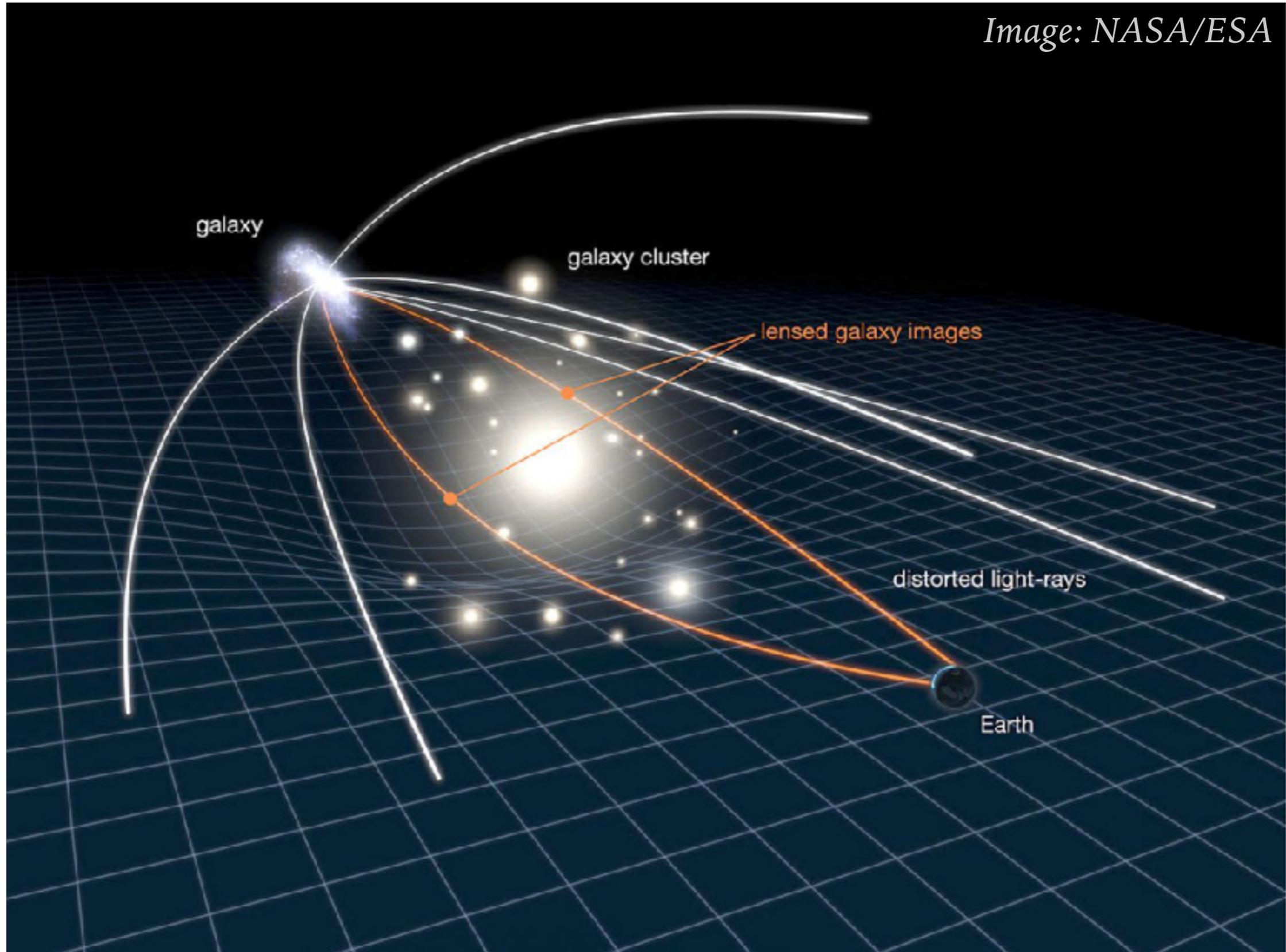
Challenges

*I. Observational*

*II. Theoretical*

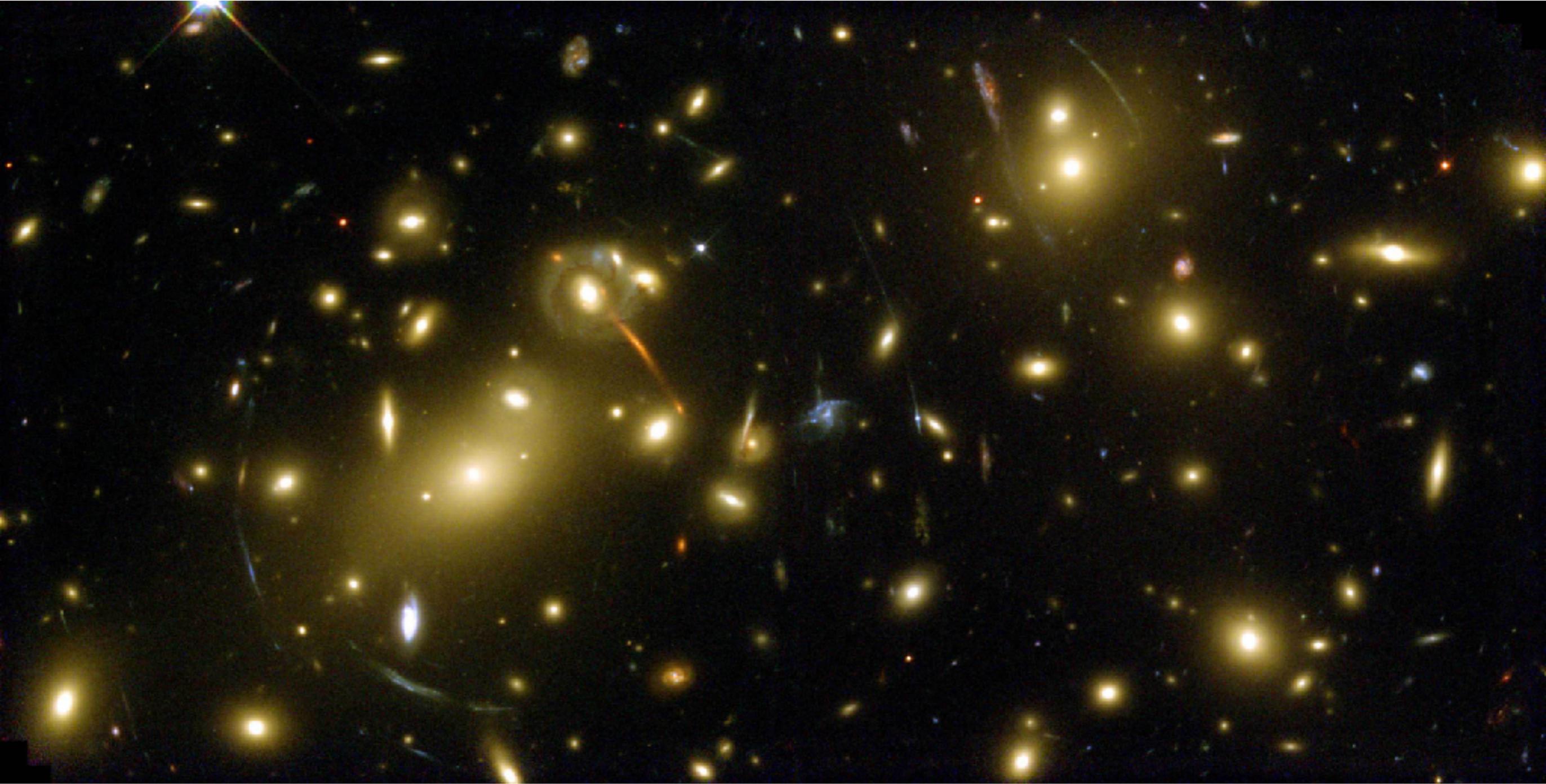
Current & future surveys

# GRAVITATIONAL LENSING



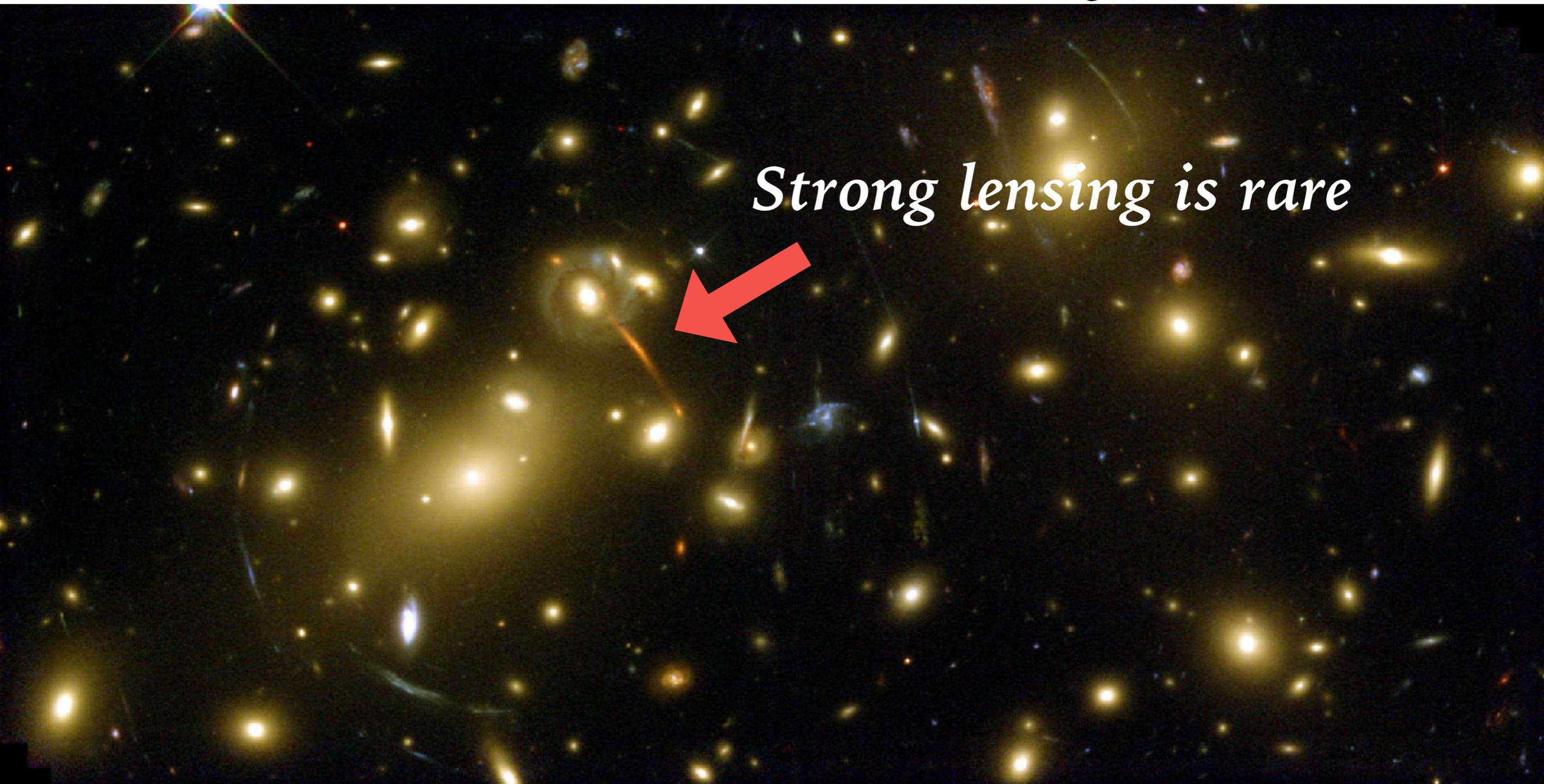
# GRAVITATIONAL LENSING

*Image: Abell 2218, NASA/ESA*



# GRAVITATIONAL LENSING

*Image: Abell 2218, NASA/ESA*



*Strong lensing is rare*

# GRAVITATIONAL LENSING

*Image: Abell 2218, NASA/ESA*

*Weak lensing can be  
measured statistically*



# CHALLENGES IN WEAK LENSING

*Image: Abell 2218, NASA/ESA*

*I. Observational: measuring galaxy shapes & distances*



# CHALLENGES IN WEAK LENSING

*Image: Abell 2218, NASA/ESA*

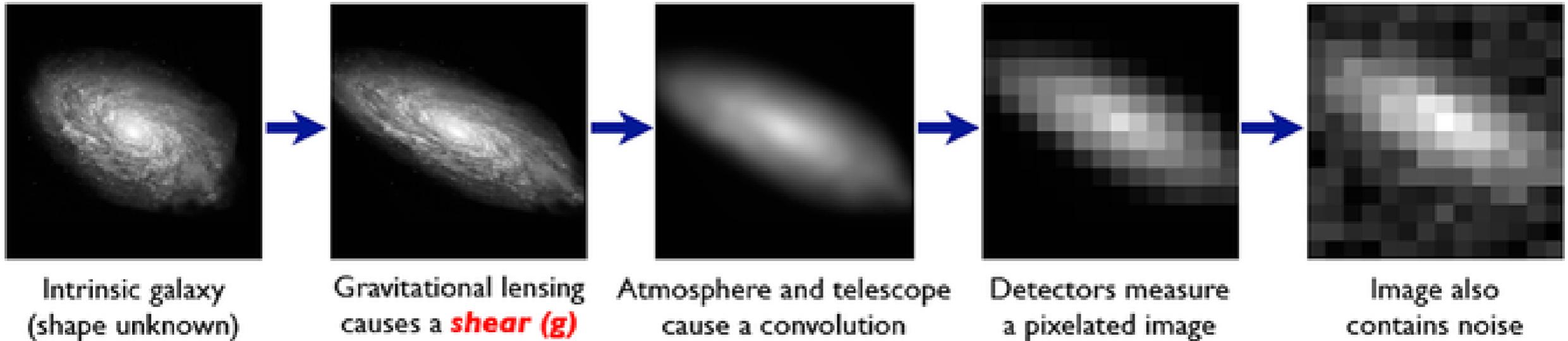
*I. Observational: measuring galaxy shapes & distances*

*II. Theoretical: modelling the large scale structure*

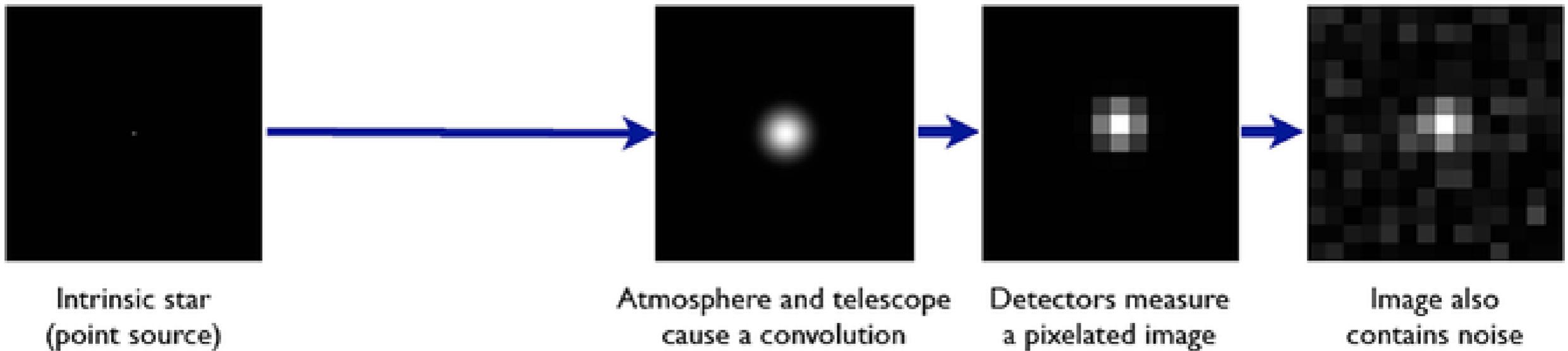
# OBSERVATIONAL CHALLENGES

## The Forward Process.

**Galaxies:** Intrinsic galaxy shapes to measured image:



**Stars:** Point sources to star images:



# OBSERVATIONAL CHALLENGES

**The Inverse Problem:**  
Measured images to *shear*

**Shear  
Field**



Set of galaxy images.  
Each contains:

- noise
- pixelisation
- convolution
- **shear**
- intrinsic shape

Set of star images.  
Each contains:

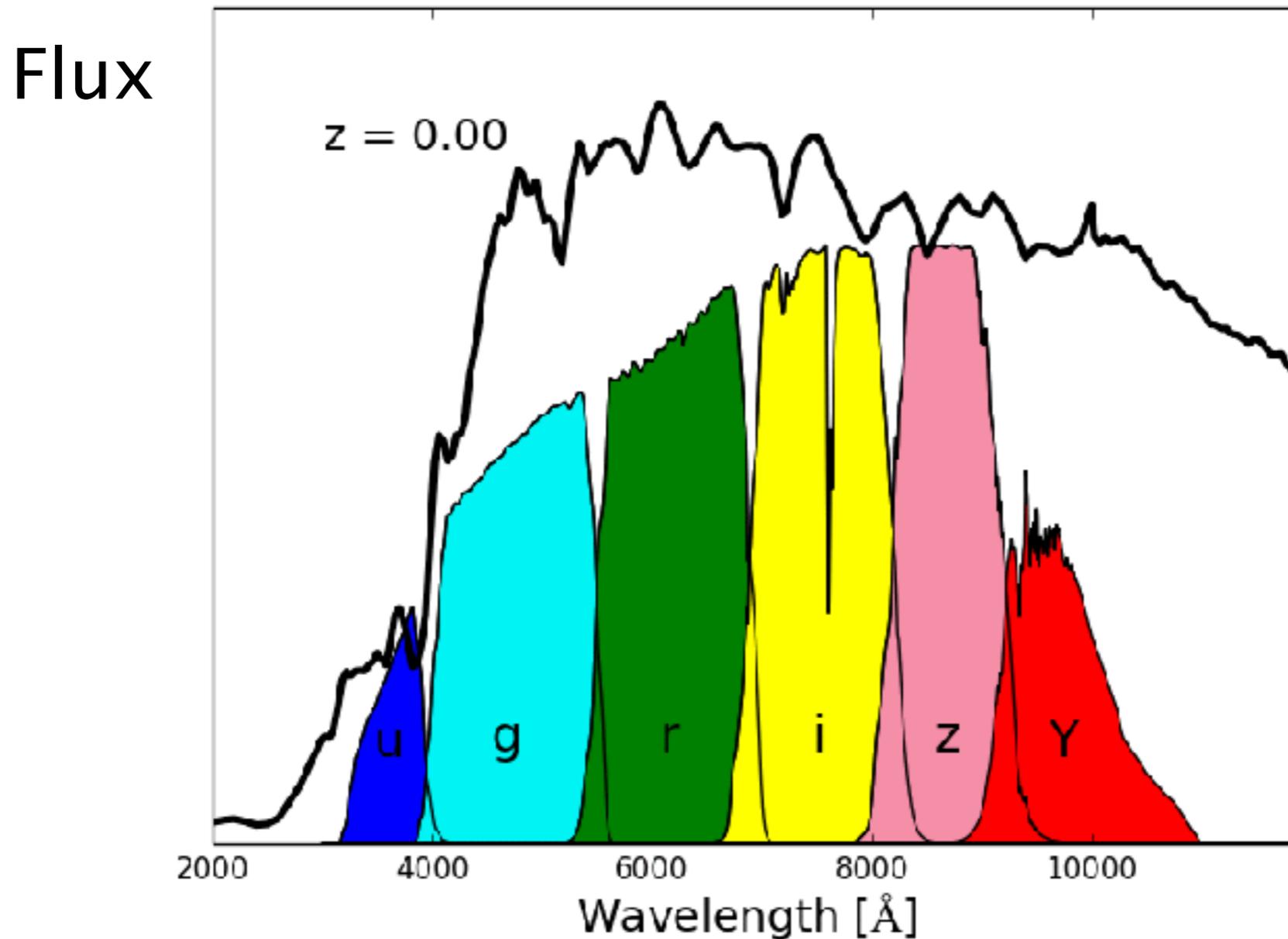
- noise
- pixelisation
- convolution

- **SHEAR CALIBRATION**
- **IMPACT OF BLENDING**
- **STAR-GALAXY SEPARATION**

*GREAT08 challenge handbook, Bridle+ (2008)*

# OBSERVATIONAL CHALLENGES

## PHOTOMETRIC REDSHIFTS



*Credit: LSST DESC Photo-z working group*

# CHALLENGES IN WEAK LENSING

*Image: Abell 2218, NASA/ESA*

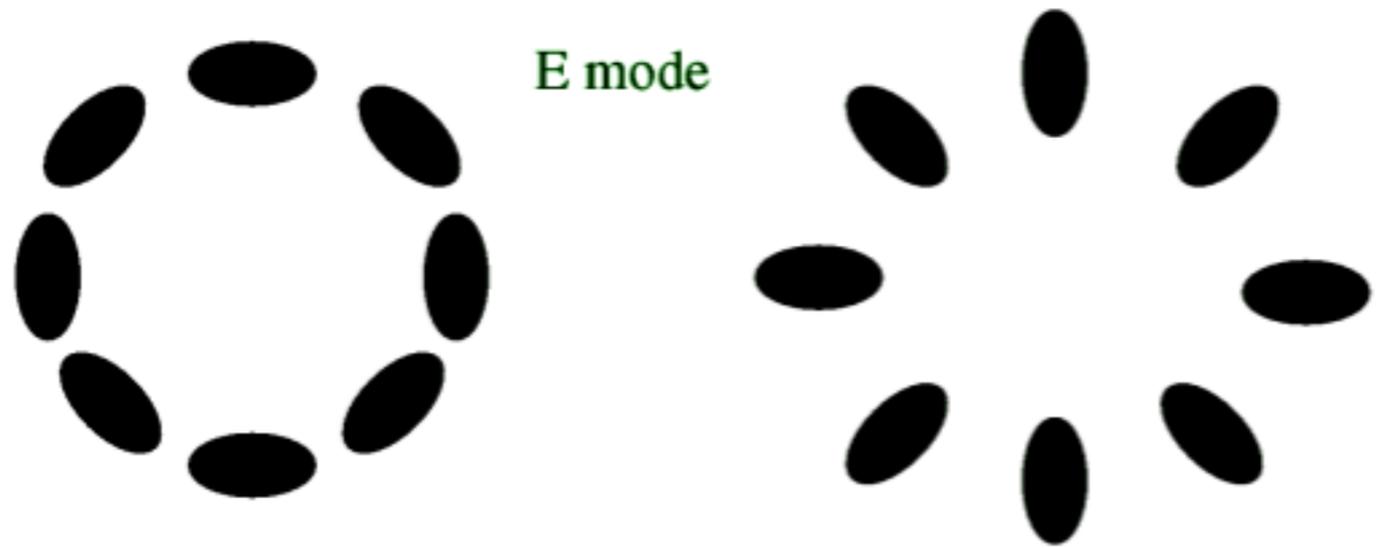
*I. Observational: measuring galaxy shapes & distances*

*II. Theoretical: modelling the large scale structure*

# THEORETICAL CHALLENGES

## OBSERVABLES:

CORRELATIONS OF GALAXY POSITIONS AND SHAPES

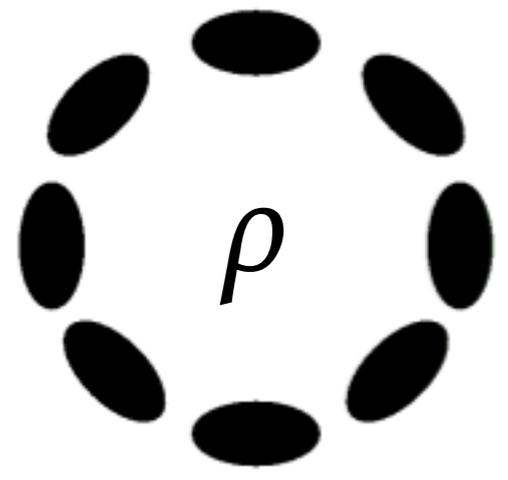
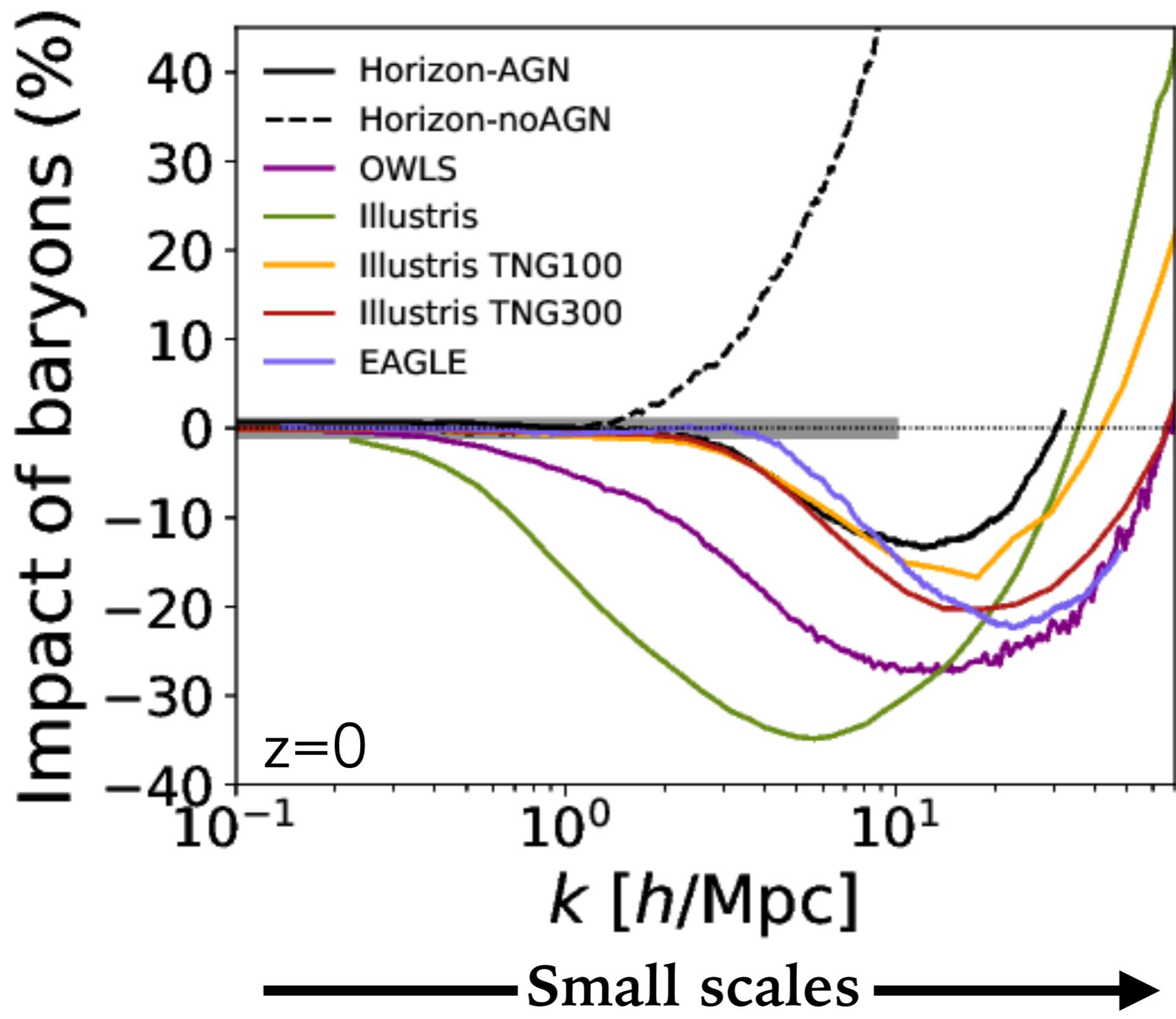


*van Waerbeke & Mellier (2003)*

## APPROACHES

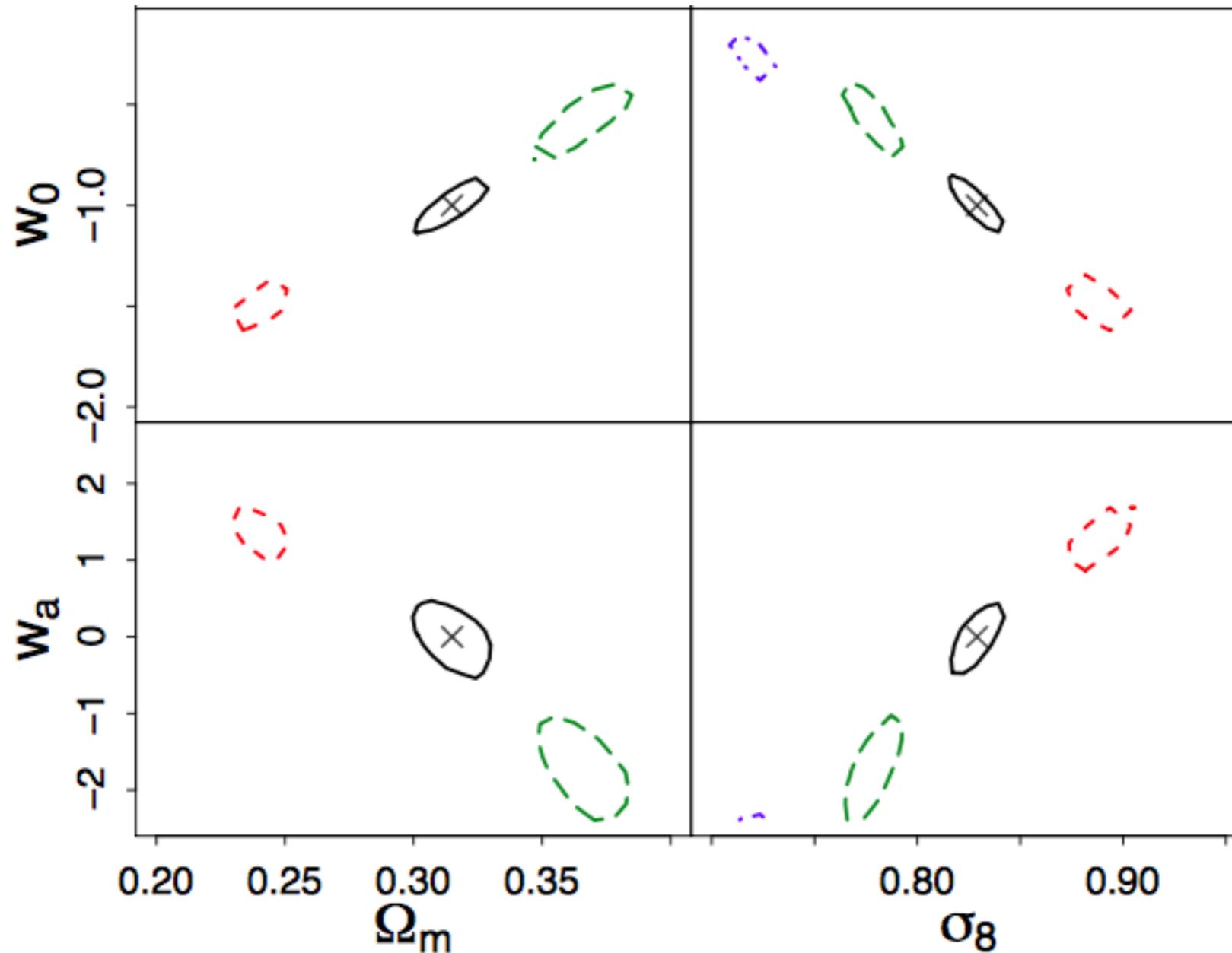
- ANALYTICAL MODELLING
- NUMERICAL SIMULATIONS

# THEORETICAL CHALLENGES



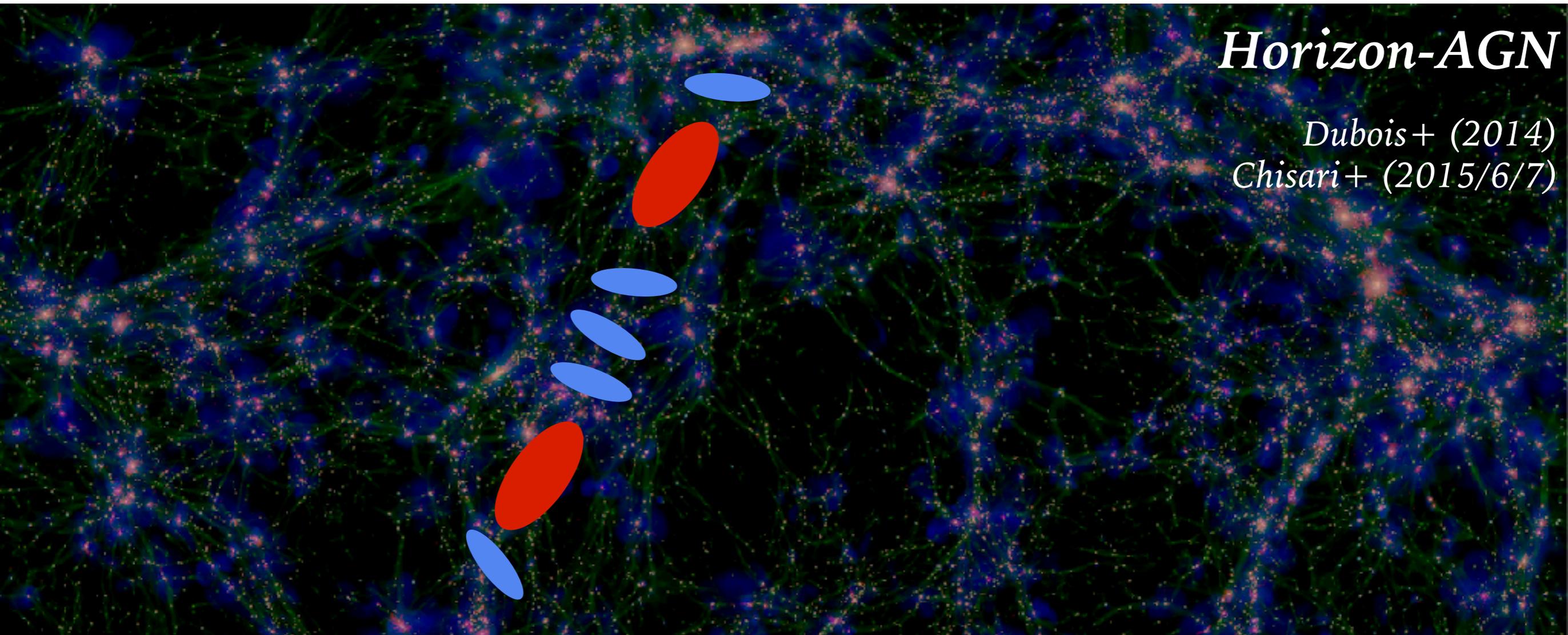
*Chisari+, in prep*  
*van Daalen+ (2011)*  
*Vogelsberger+ (2014)*  
*Hellwing+ (2016)*  
*Springel+ (2017)*

# THEORETICAL CHALLENGES



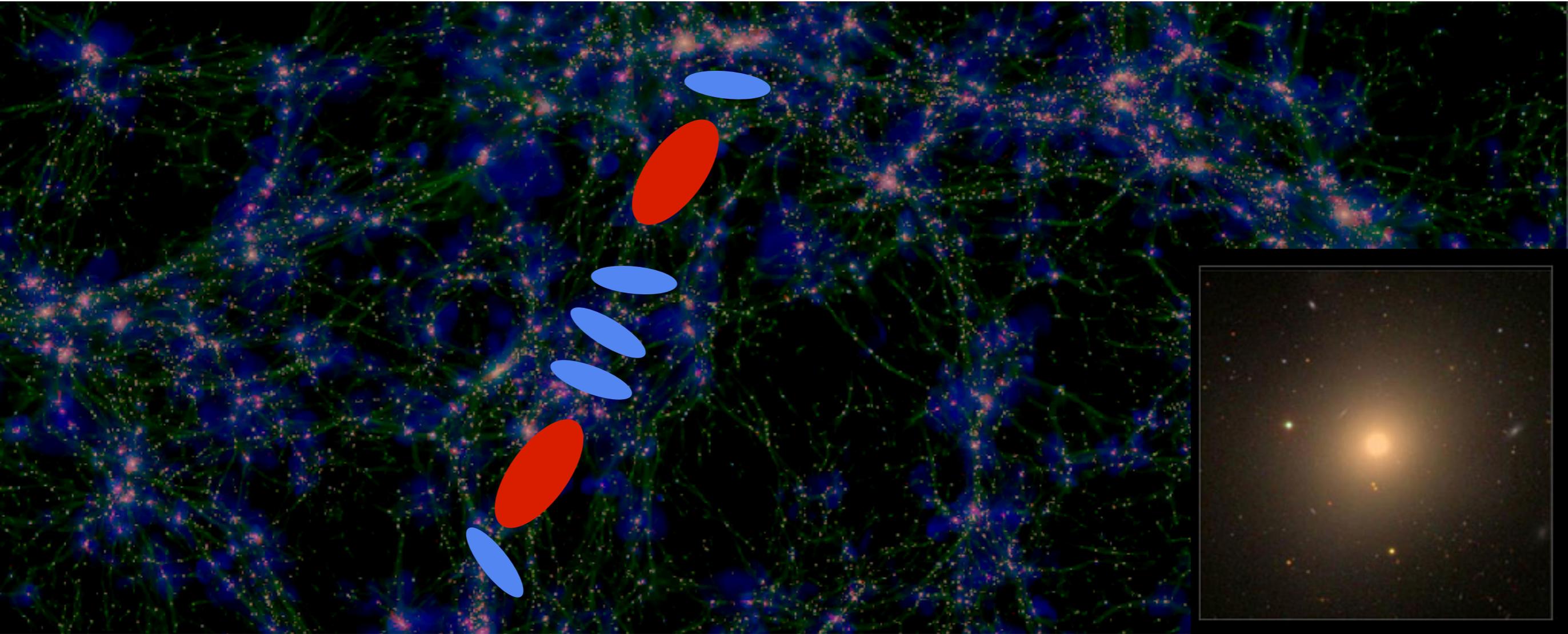
*Eifler + (2015)*

# THEORETICAL CHALLENGES



*Galaxy shapes = Lensing + Intrinsic alignment + Noise*

# THEORETICAL CHALLENGES

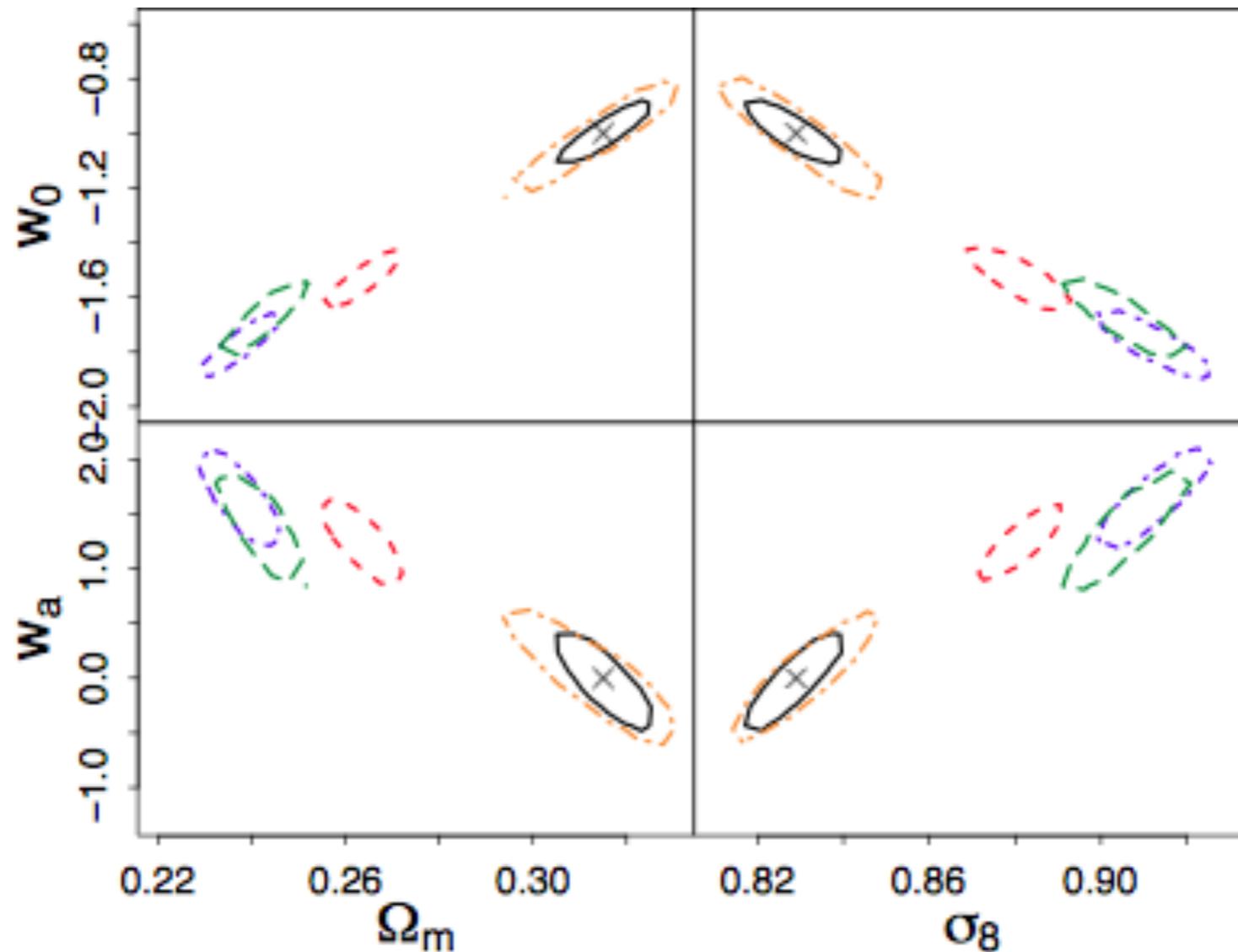


*Intrinsic shapes*  $\sim$  *Tidal field of the large-scale structure*

*Catelan + (2001)*

# THEORETICAL CHALLENGES

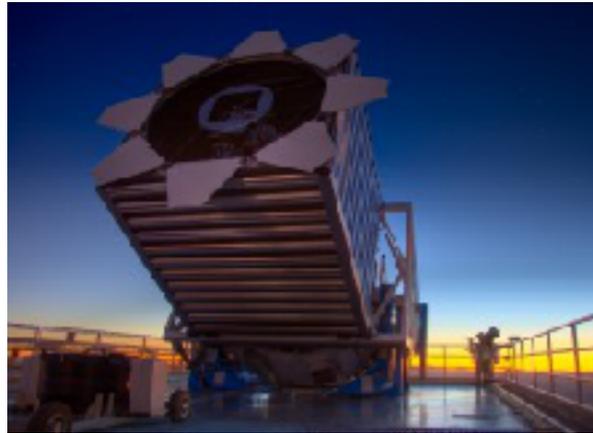
*Bias in cosmology due to galaxy alignments*



*Krause + (2015)*

# PAST, PRESENT & FUTURE GALAXY SURVEYS

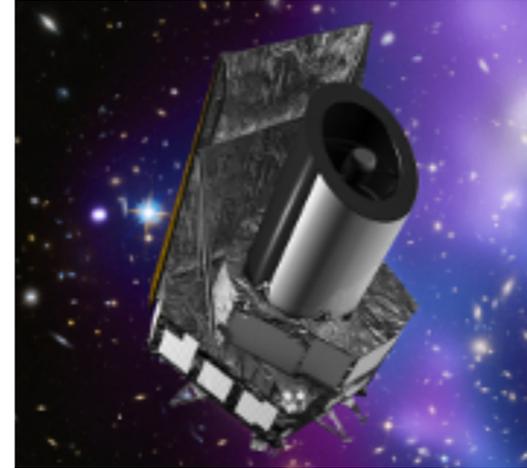
*SDSS*



*Kilo-Degree Survey*



*Euclid (2021)*



*LSST (2021)*



*CFHTLenS*



*DES*



*HSC*



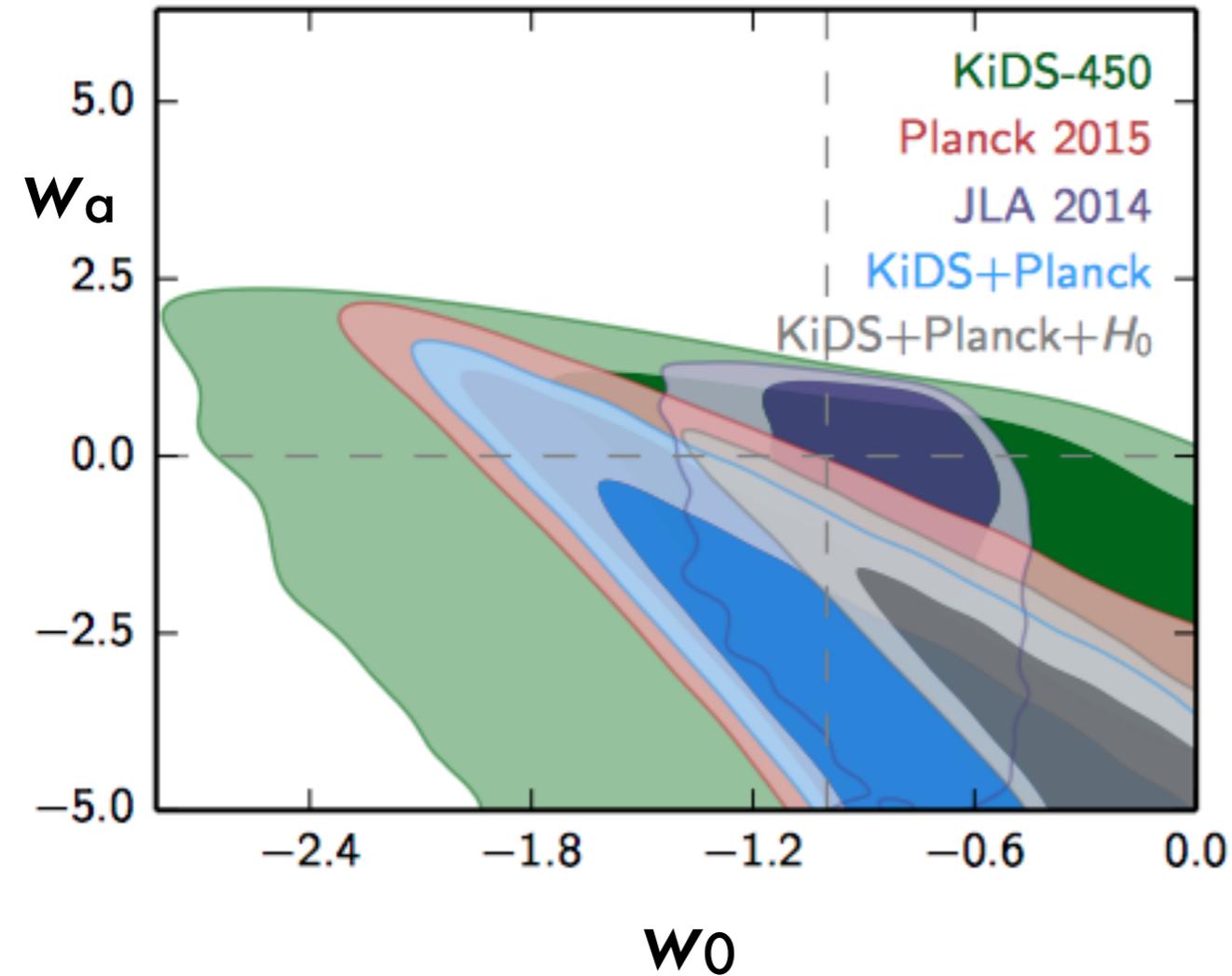
*WFIRST (2025)*



# COSMOLOGY WITH GALAXY SURVEYS

$$w = w_0 + w_a(1-a)$$

*KiDS-450*



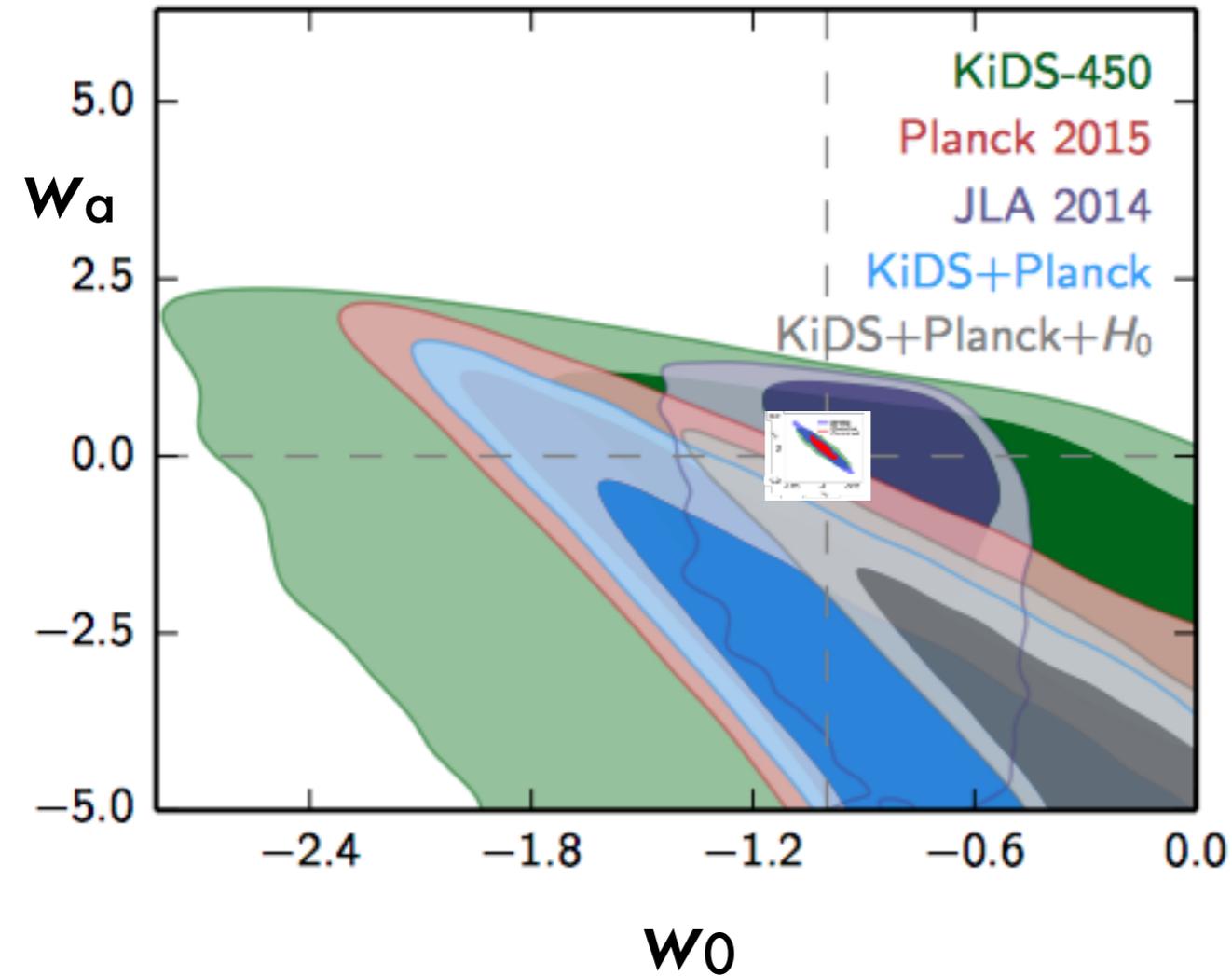
Joudaki+,  
KiDS collaboration, 2016

# COSMOLOGY WITH GALAXY SURVEYS

$$w = w_0 + w_a(1-a)$$

*KiDS-450*

*LSST expected*



Joudaki+,  
KiDS collaboration, 2016

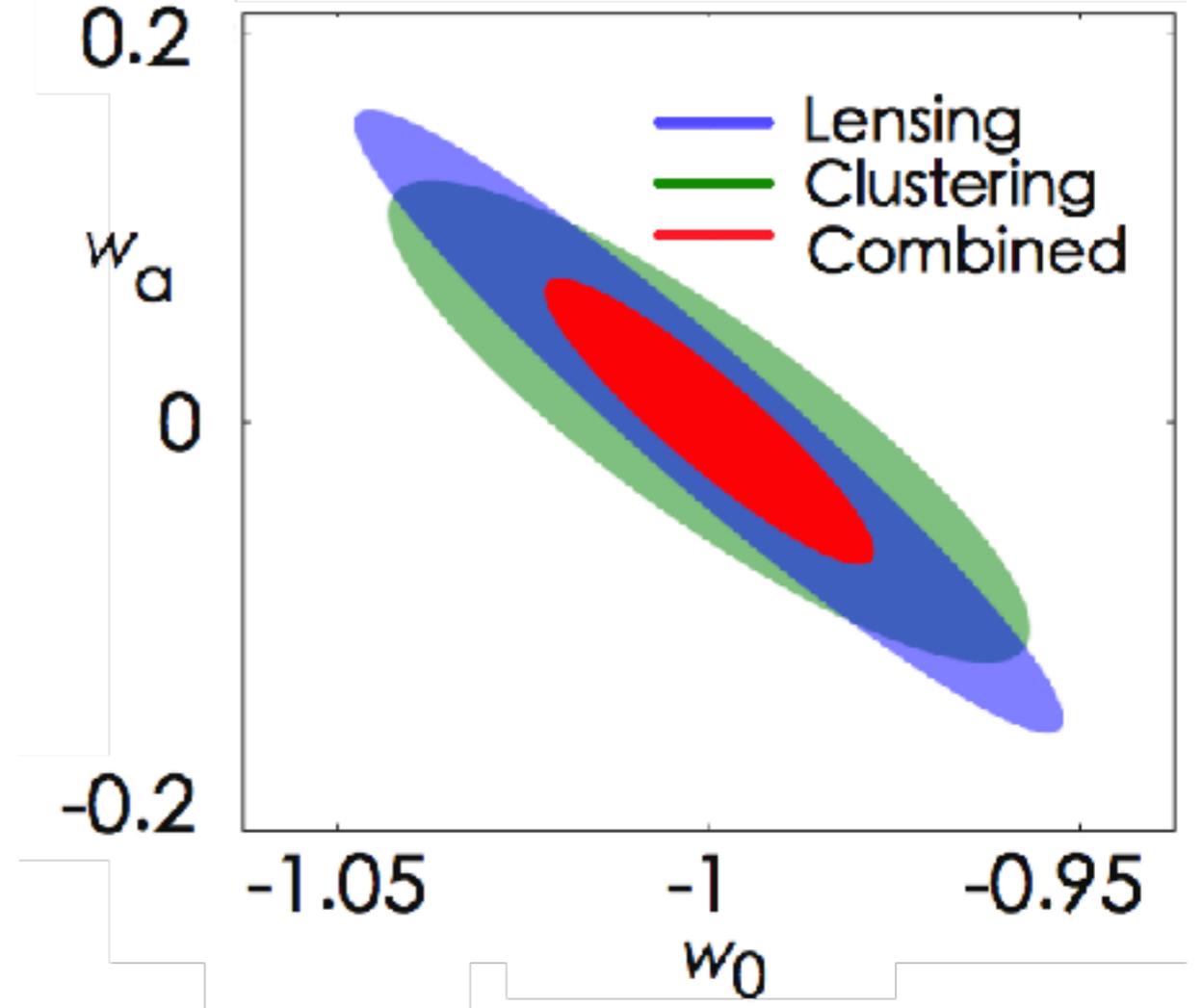
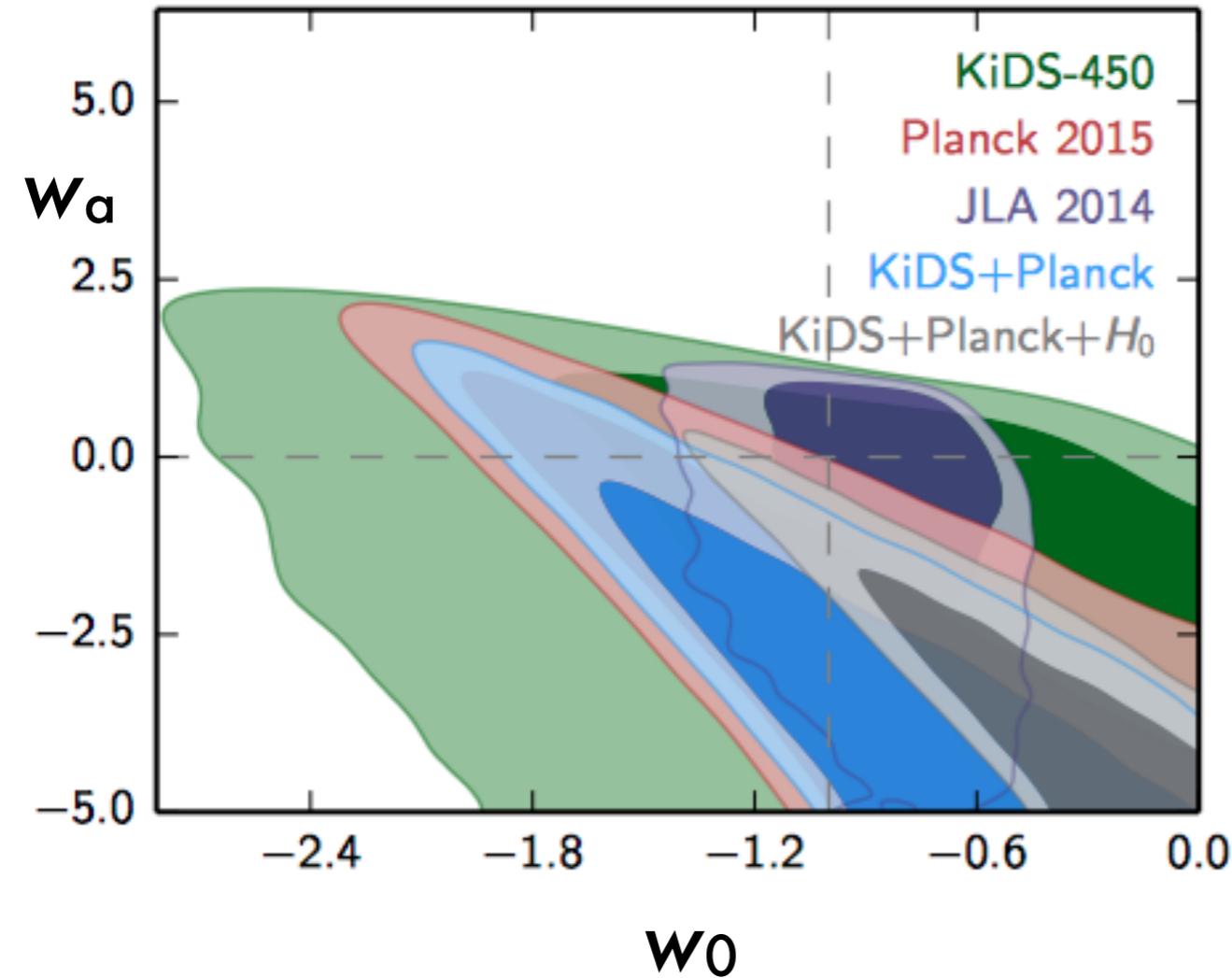
D. Alonso, E. Chisari+  
LSST-DESC Code Comparison

# COSMOLOGY WITH GALAXY SURVEYS

$$w = w_0 + w_a(1-a)$$

*KiDS-450*

*LSST expected*

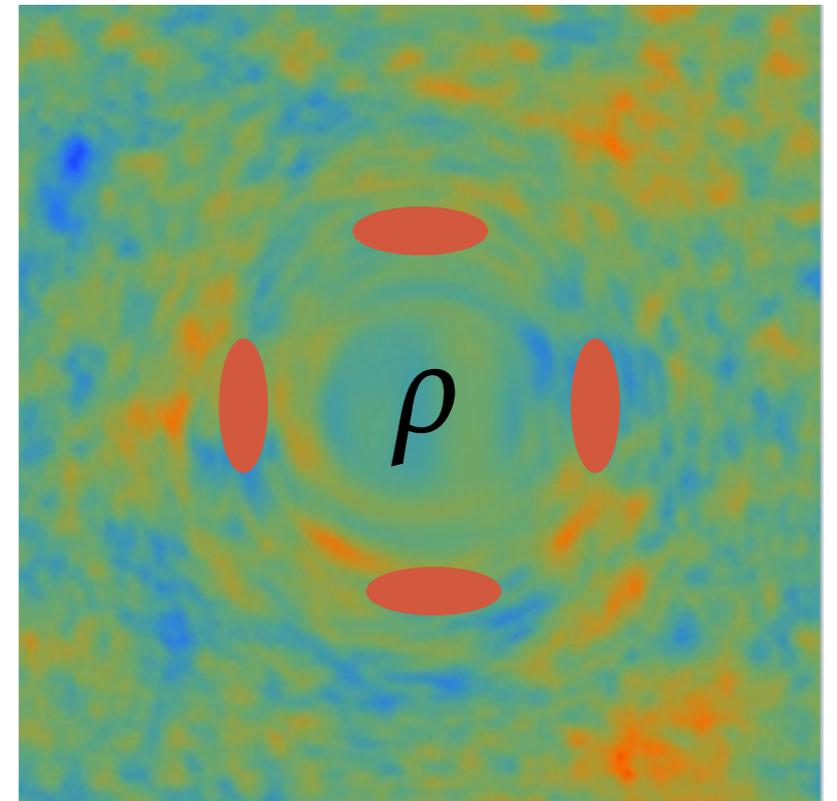
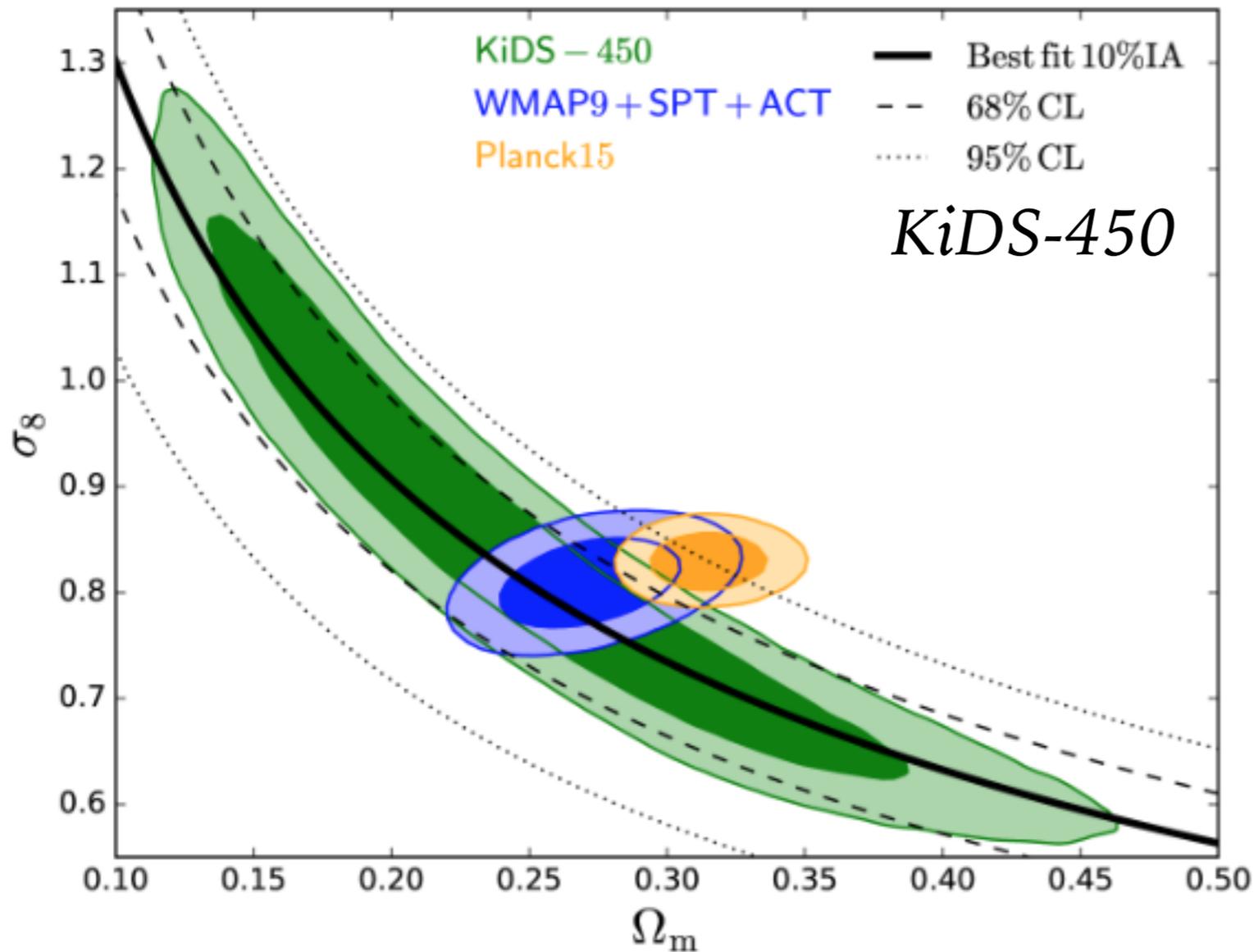


Joudaki+,  
KiDS collaboration, 2016

D. Alonso, E. Chisari+  
LSST-DESC Code Comparison

# COSMOLOGY WITH GALAXY SURVEYS

*Combining probes: cross-correlation with CMB lensing*



*Credit: W. Hu*

*Harnois-Déraps + (KiDS collaboration, 2017)*

# CONCLUSIONS

- *Gravitational lensing* is one of the most promising probes of dark energy in the next decade.
- The main observational challenges to cosmology with gravitational lensing are accurate estimation of *galaxy shapes and distances*.
- The main theoretical challenges in modelling the large-scale structure of the Universe are associated with the role of baryons and galaxy formation.
- Preparation is underway to enable the next generation of lensing experiments (LSST, Euclid) which will drive *a next era in precision cosmology*.