



TAKE A SEAT

YOUNG PHYSICISTS

And mathematicians...

An Introduction to Higgspllosion

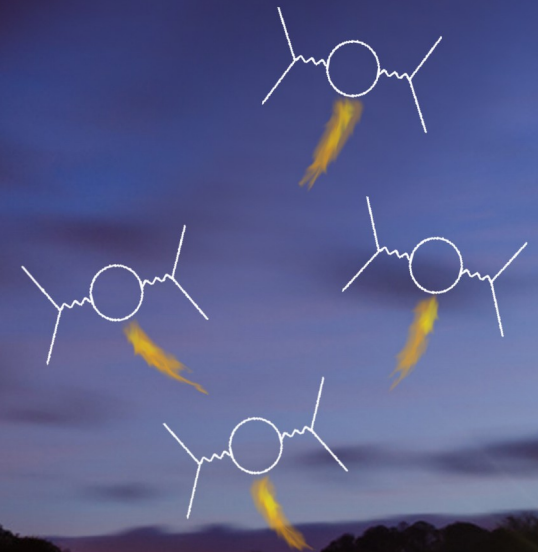
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YTF 10, January 2018

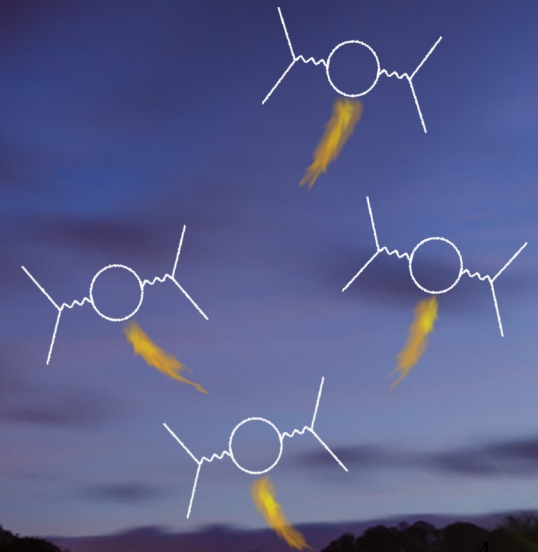
An Introduction to Higgspllosion

Joey Reiness
(Supervisor: Prof. V. Khoze)



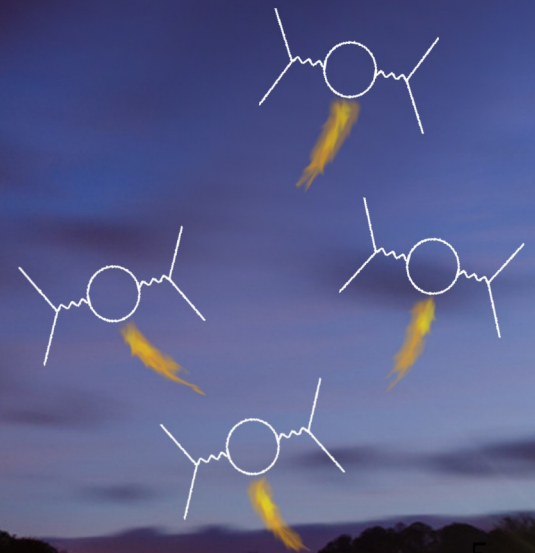
Split talk into three parts:

- 1) What is Higgspllosion?
- 2) Effect on RG running, with ϕ_4 example (1709.08655)
- 3) Higgsploding DM?



Higgsplosion a nutshell:

- Predict that highly energetic particles coupled to higgs develop exponentially growing decay rates for virtualities $p^2 > E_H^2$
- Just an expected consequence of scalar QFTs
- Nothing added to SM!



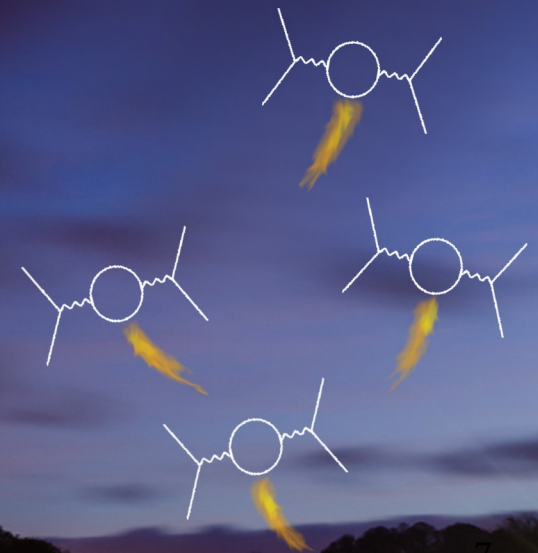
Basic Consequences:

- Distinct new phase of the theory: state of large number of soft quanta
- Loop integrals are cut off, regulating UV divs
- Sets a minimum resolvable distance $1/E_H$



Which means...

- Theory is UV finite: couplings freeze at scale E_H
- No Landau poles, asymptotically safe
- Hierarchy problem enormously reduced
- Easier to add heavy species





This is where the fun begins.

Scalars be crazy...

- In 90s, found that scalar QFTs such as phi⁴ theory have exponentially growing amplitudes for 1 → n threshold process
- Factorial growth of diagrams
- No destructive interference



Model (2.1) : $\mathcal{A}_{1 \rightarrow n}(p_1 \dots p_n) =$

Model (2.2) : $\mathcal{A}_{1 \rightarrow n}(p_1 \dots p_n) =$

$$n! \left(\frac{\lambda}{8m^2} \right)^{\frac{n-1}{2}} \exp \left[-\frac{5}{6} n \varepsilon \right],$$

$$n! \left(\frac{1}{2v} \right)^{n-1} \exp \left[-\frac{7}{6} n \varepsilon \right],$$

$$n \rightarrow \infty, \quad \varepsilon \rightarrow 0, \quad n\varepsilon = \text{fixed}$$



The higgs boson is a scalar...

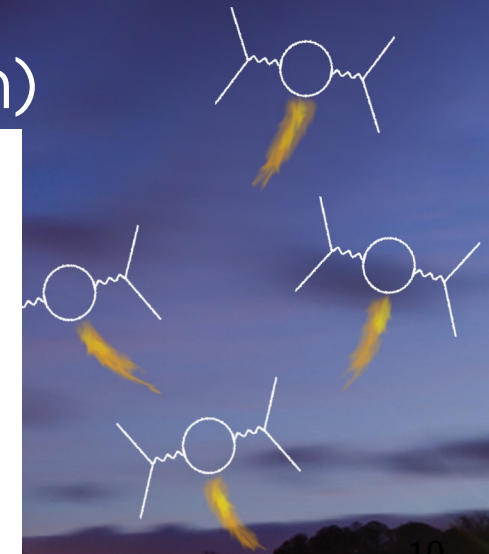
- Expect $h \rightarrow nh$ to have a large amplitude
- Decay rate grows exponentially

$$\Gamma_n(s) \propto \mathcal{R}(\lambda; n, \varepsilon) = \exp \left[n \left(\log \frac{\lambda n}{4} + \frac{3}{2} \log \frac{\varepsilon}{3\pi} + \frac{1}{2} - \frac{25}{12} \varepsilon + Q(\lambda n, \varepsilon) \right) \right].$$

- For real processes, intermediate higgs is dressed \rightarrow preserve unitarity (Higgspersion)

$$\sigma_{gg \rightarrow n \times h}^{\Delta} \sim y_t^2 m_t^2 \log^4 \left(\frac{m_t}{\sqrt{p^2}} \right) \times \frac{1}{p^4 + m_h^4 \mathcal{R}^2} \times \mathcal{R}_n,$$

$$\sigma_{gg \rightarrow n \times h} \sim \begin{cases} \mathcal{R} & : \text{for } \sqrt{s} \leq E_* \text{ where } \mathcal{R} \lesssim 1 \\ 1/\mathcal{R} \rightarrow 0 & : \text{for } \sqrt{s} \geq E_* \text{ where } \mathcal{R} \gg 1 \end{cases}.$$



Loop integrals and propagator:

- Can treat Γ as a going from $0 \rightarrow \text{inf}$ at $p^2 = E_H^2$
 \rightarrow propagator has a heaviside factor $\Theta(p^2 - E_H^2)$
 \rightarrow closing loop introduces p integral which is cut-off

$$\Delta_R(x, x) = \int_{p^2 \leq E_*^2} \frac{d^4 p}{(2\pi)^4} \frac{1}{p^2 + m^2}.$$



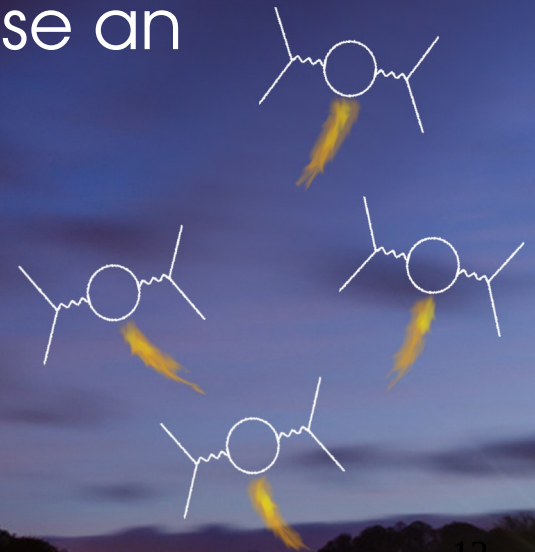
$$\Delta(x) := \langle 0 | T(\phi(x) \phi(0)) | 0 \rangle \sim \begin{cases} m^2 e^{-m|x|} & : \text{ for } |x| \gg 1/m \\ 1/|x|^2 & : \text{ for } 1/E_* \ll |x| \ll 1/m \\ E_*^2 & : \text{ for } |x| \lesssim 1/E_* \end{cases}$$

Higgsplosion scale:

- By dimensional grounds:

$$E_* = \frac{m_h}{f(\lambda)}, \quad \text{where } f(\lambda)|_{\lambda \rightarrow 0} \rightarrow 0.$$

- Similar to sphaleron where: $M_{\text{sph}} = \text{const} \frac{m_W}{\alpha_w}.$
- Both non-perturbative and semiclassical in nature
- Not in Lagrangian, but rather characterise an energy scale for a transition



Aim for this next part:

- Based on parts of (1709.08655)
- The exponential stuff done in 90s from a maths perspective: an interesting feature of scalar QFTs
- Now we have a scalar – can we apply to SM?
- Need to check if their 90s calcs are still consistent if the propagators/loops are affected as proposed



Review of Brown's method (1/2):

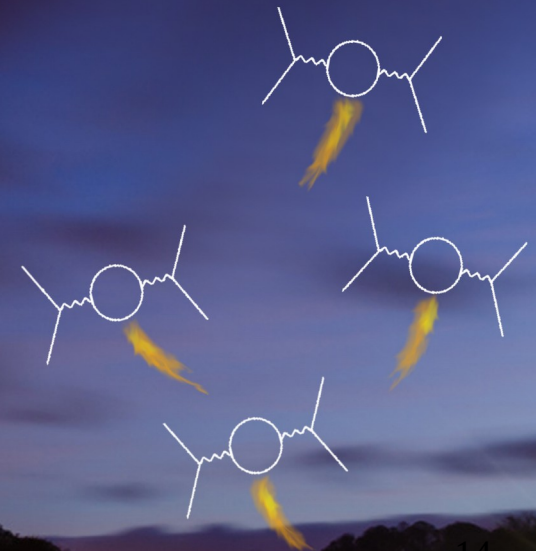
- Scalar ϕ^4 with source, LSZ reduction:

$$\mathcal{L} = -|\partial\phi|^2/2 - m^2\phi^2/2 - \lambda\phi^4/4! + \rho\phi$$

$$\langle n|\phi(x)|0\rangle = \prod_{a=1}^n \int (d^4x_a) e^{-ip_a x_a} (p_a^2 + m^2) \frac{\delta}{\delta\rho(x_a)} \langle 0 + |\phi(x)|0-\rangle^\rho |_{\rho=0}$$

- For tree level, generating function is classical solution:

$$\langle 0 + |\phi(x)|0-\rangle^\rho \rightarrow \phi_{cl}(x)$$



Review of Brown's method (2/2):

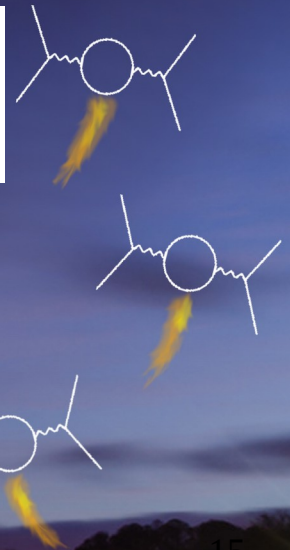
- Momenta vanish on threshold
- Can impose spatial uniformity in this limit:

$$\vec{p}_a = 0, \text{ so } \rho(x), \phi_{cl}(x) \rightarrow \rho(t) \equiv \rho_0 e^{i\omega t}, \phi_{cl}(t)$$

- Find that $\phi_{cl}(t) = \phi_{cl}(z(t))$ and therefore LSZ reduces to ODE:

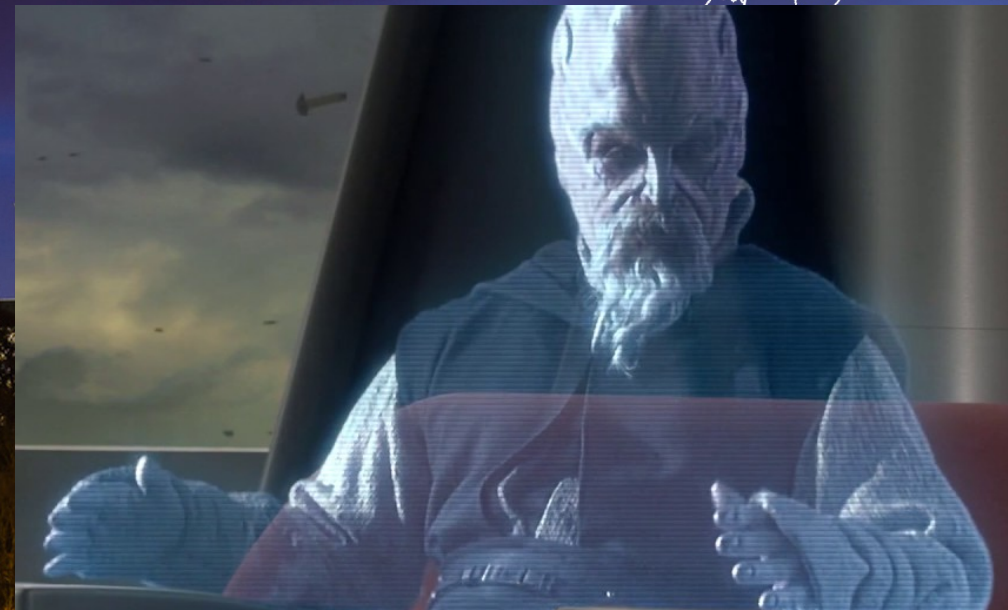
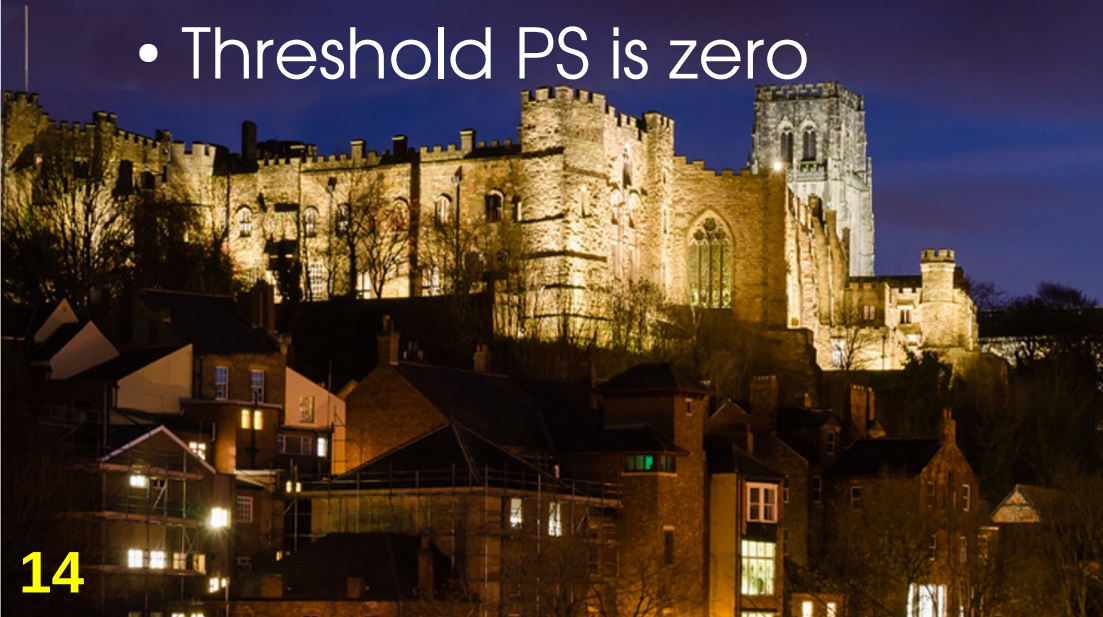
$$\int (d^4 x_a) e^{-ip_a x_a} (p_a^2 + m^2) \frac{\delta}{\delta \rho(x_a)} \phi_{cl}(t; [\rho]) = \frac{\partial}{\partial z_0} \phi_{cl}(z(t))$$

$$\langle n | \phi(0) | 0 \rangle_{threshold}^{tree} = \left(\frac{\partial}{\partial z} \right)^n \phi_{cl} |_{z=0}$$



Comments on Brown:

- Only tree level
- Higgspllosion boils down to cutting off loop momenta integrals \rightarrow no effect on tree level
- Only for simple scalar sector... what about the fermions' effect on the decay rate?
- Threshold PS is zero





“Oh I'm not brave enough for 1-loop calculations...”

Brief summary of Voloshin:

- Expand QM corrections around classical

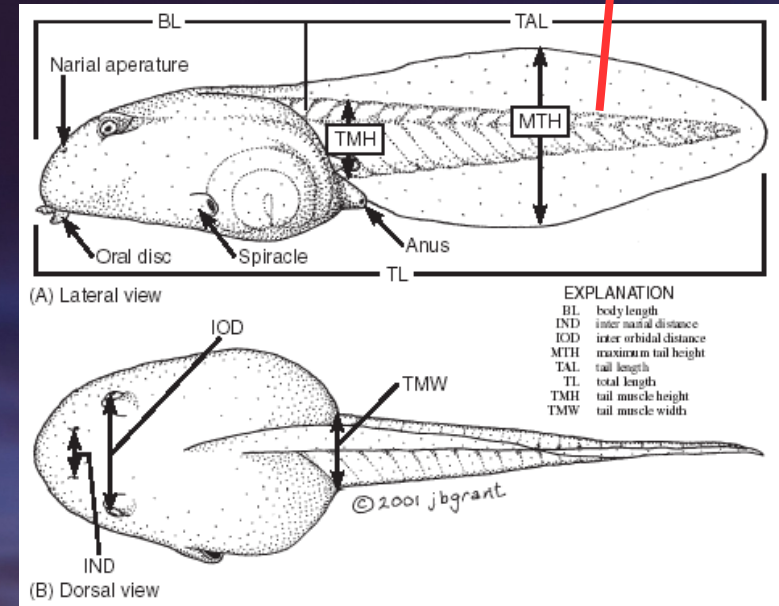
$$\langle \phi \rangle = \phi_0 + \langle \phi_q \rangle$$

- Use mixed space rep

$$\langle n | \phi | 0 \rangle = \left(\frac{\partial}{\partial z_0} \right)^n (\phi_0 + \langle \phi_q \rangle) \Big|_{z_0=0}.$$

- Clever time shift and rotation to make it look like a known QM example...

- Lots of dry maths



Enter Higgspllosion:

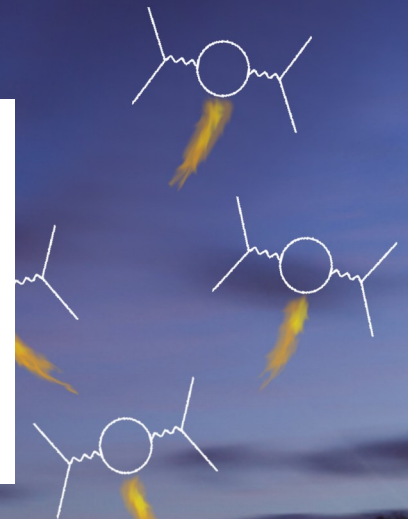
- Ultimately Voloshin finds:

$$G(x, x) = \frac{1}{16\pi^2} \left(E_*^2 - m^2 \log \frac{E_*^2}{m^2} \right) - \frac{1}{2\pi^2} \frac{3\lambda\phi_0^2}{8} \left(\log \frac{E_*^2}{m^2} + 1 \right) - \frac{3\lambda^2\phi_0^4}{32} F.$$

- Absorb divergences in renormalisation
- Higgspllosion scale enters as the cut-off to these divergences

$$\bar{\lambda}_* = \lambda - \frac{9\lambda^2}{8\pi^2} \left(\log \frac{E_*}{m} + \frac{1}{2} \right),$$

$$\bar{m}_*^2 = m^2 + \frac{3\lambda}{16\pi^2} \left(E_*^2 - m^2 \log \frac{E_*^2}{m^2} \right),$$





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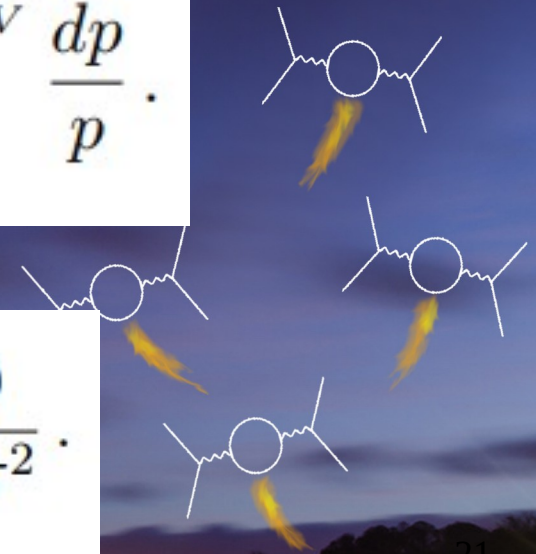
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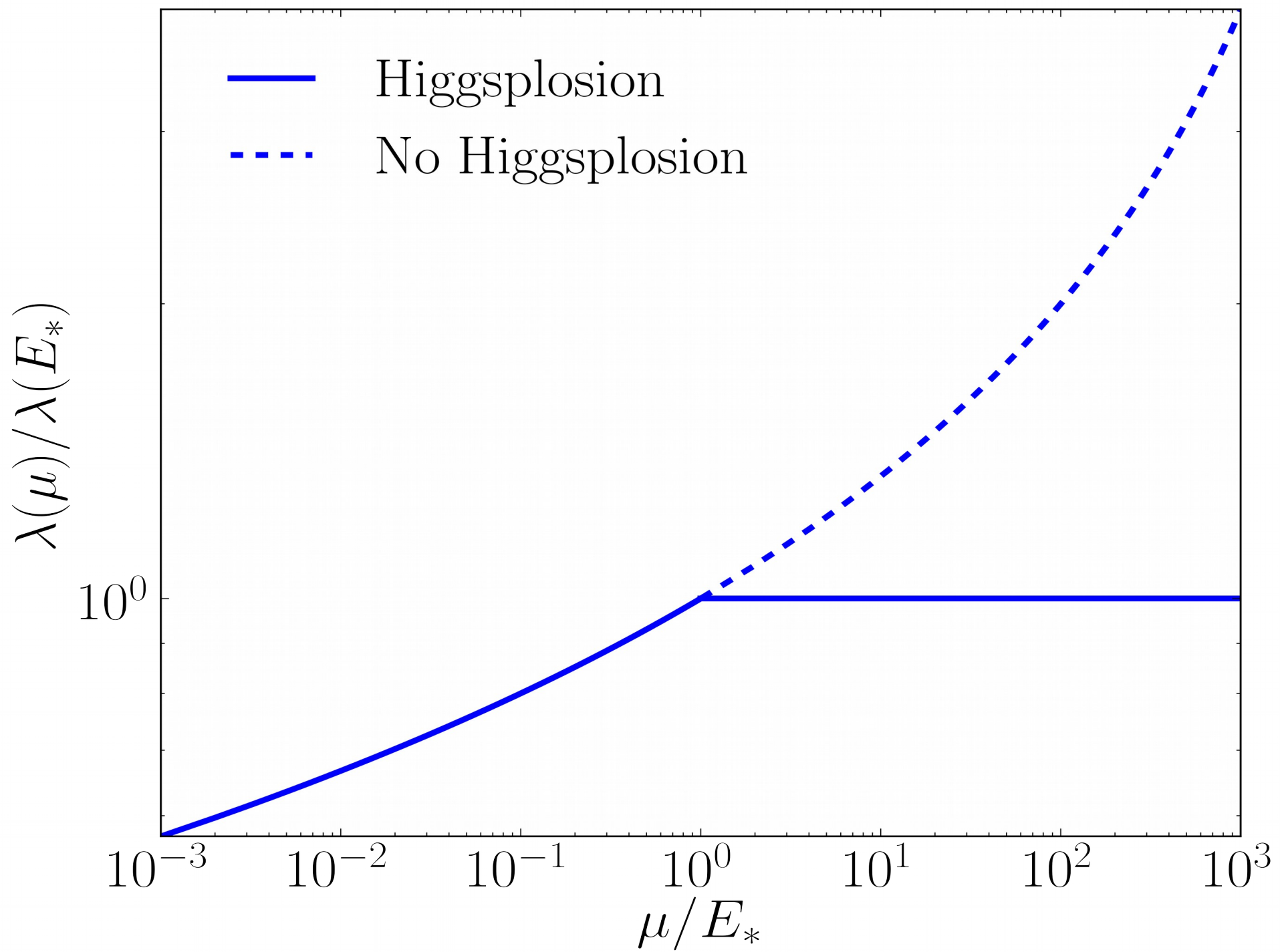
Running of coupling:

- Cut-off is now physically significant!
- Any corrections to finite terms sub-leading
- In RG, Higgspllosion freezes evolution at E_H

$$1/\lambda(\mu) = 1/\lambda(\Lambda_{UV}) + \frac{9}{8\pi^2} \int_{\mu}^{\Lambda_{UV}} \frac{dp}{p}.$$

$$\lambda(\mu) = \frac{1}{\beta_0 \log\left(\frac{\Lambda_{LP}}{\mu}\right)}, \quad \beta_0 = \frac{9}{8\pi^2}.$$





A surprise, to be sure, but a welcome one:

- Fixed UV point, conformal symmetry
- No Landau poles, asymptotic safety
- Probably not great for GUTs...



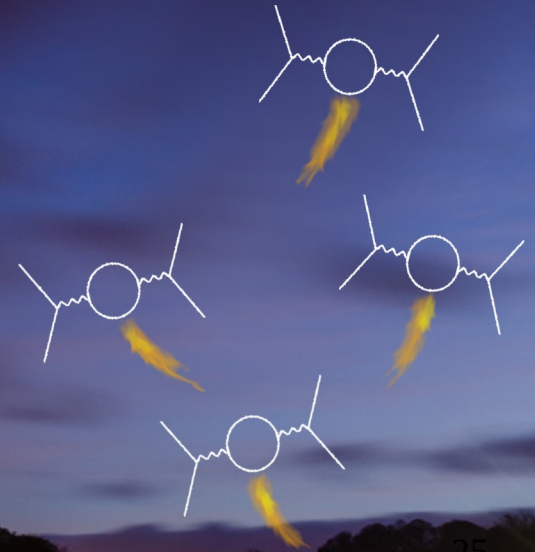
Summary & exponentiation

- Finite terms more-or-less unaffected in 1-loop correction
 - Libanov's 1994 exponentiation still valid
 - important for using **semi-classical** methods in non-perturbative regime
- This is just ϕ^4 , need to add vector & top loops etc
- But looks alright so far



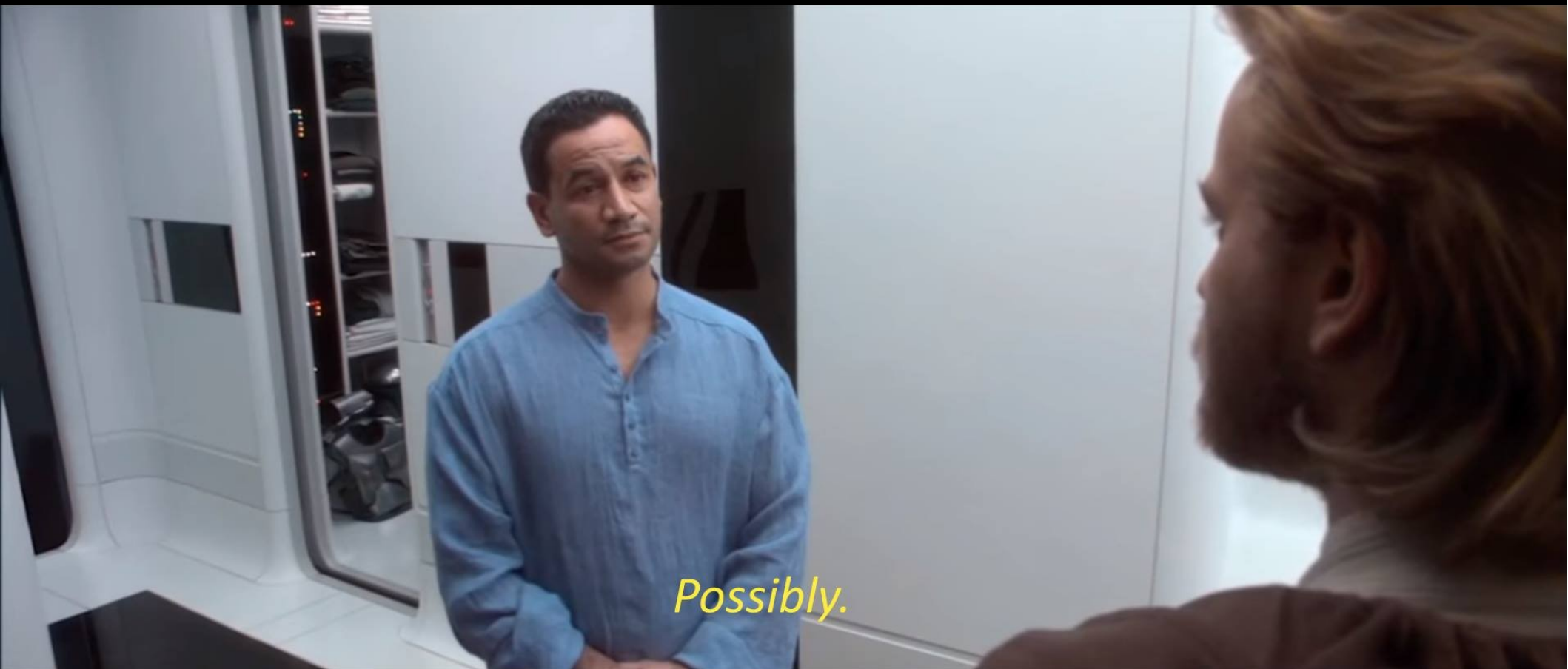
You're going down a path I can't follow.

Can we have a singlet scalar DM
that higgsplodes?





Yep



Possibly.

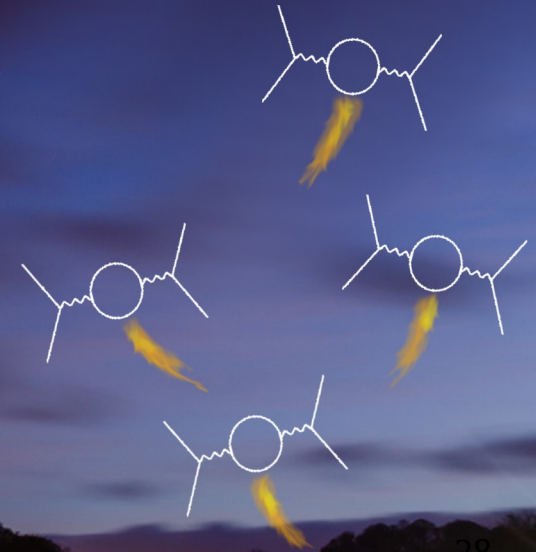
Just a singlet scalar & higgs portal...

- Keep it simple

$$\mathcal{L} = \mathcal{L}_{SM} + \partial_\mu X^\dagger \partial^\mu X - m_0^2 X^\dagger X - \frac{\lambda_X}{4} (X^\dagger X)^2 - \lambda_{HX} (X^\dagger X) (H^\dagger H).$$

- Assume $\lambda_X \ll \lambda_{HX}$ and m_0^2 small so that m_X is dominated by higgs bubble diagram $m_X \approx \sqrt{\lambda_{HX}} \frac{E_H}{4\pi}$.

- So now DM mass, portal coupling and Higgspllosion scale linked - only 2 free paramaters!



Standard freeze out stuff:

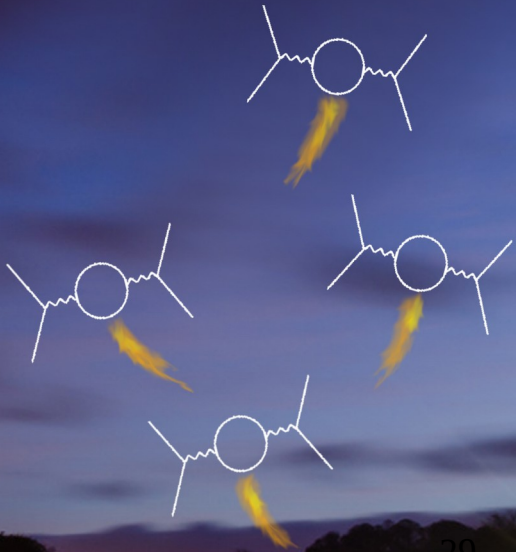
- Follow the standard recipe

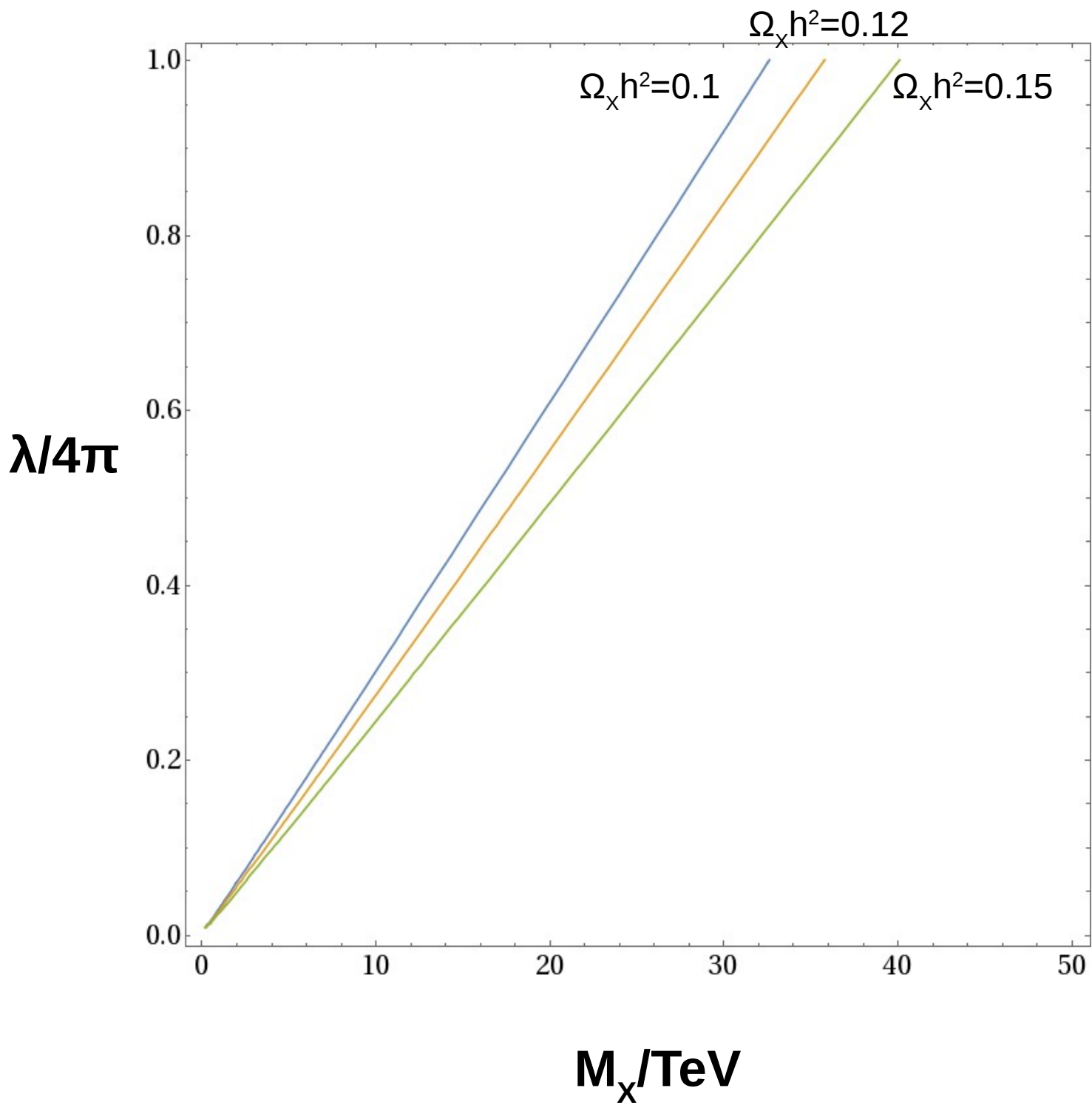
$$\Omega_X h^2 = \frac{(1.07 \times 10^9) x_f}{\sqrt{g_*} M_{PL} \text{GeV} \langle \sigma_{ann} v_{rel} \rangle}, \quad x_f \simeq \ln \left[\frac{0.038 M_{PL} m_X \langle \sigma_{ann} v_{rel} \rangle}{\sqrt{g_*} x_f} \right].$$

- At $m_X > m_{SM}$, annihilation dominated by $hh, W_L W_L, Z_L Z_L \rightarrow$ just $h_{1,2,3,4}$

$$\langle \sigma_{ann} v_{rel} \rangle \approx \frac{\lambda_{HX}^2}{16\pi m_X^2}.$$

- Demand Planck relic density $\Omega_X h^2 = 0.12$, reducing to 1 free parameter: m_X



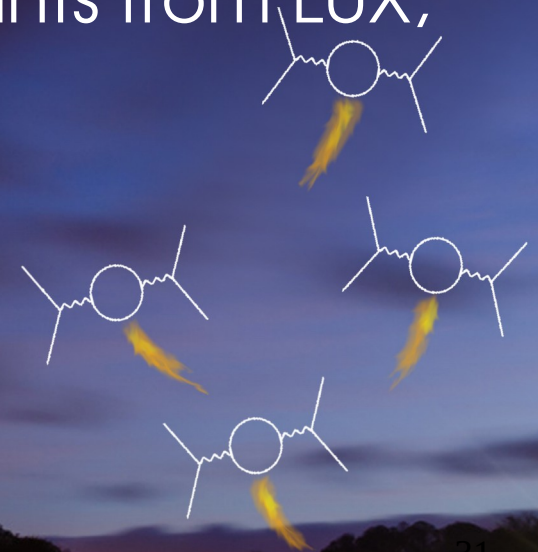


Direct detection:

- Elastic collision via higgs exchange:

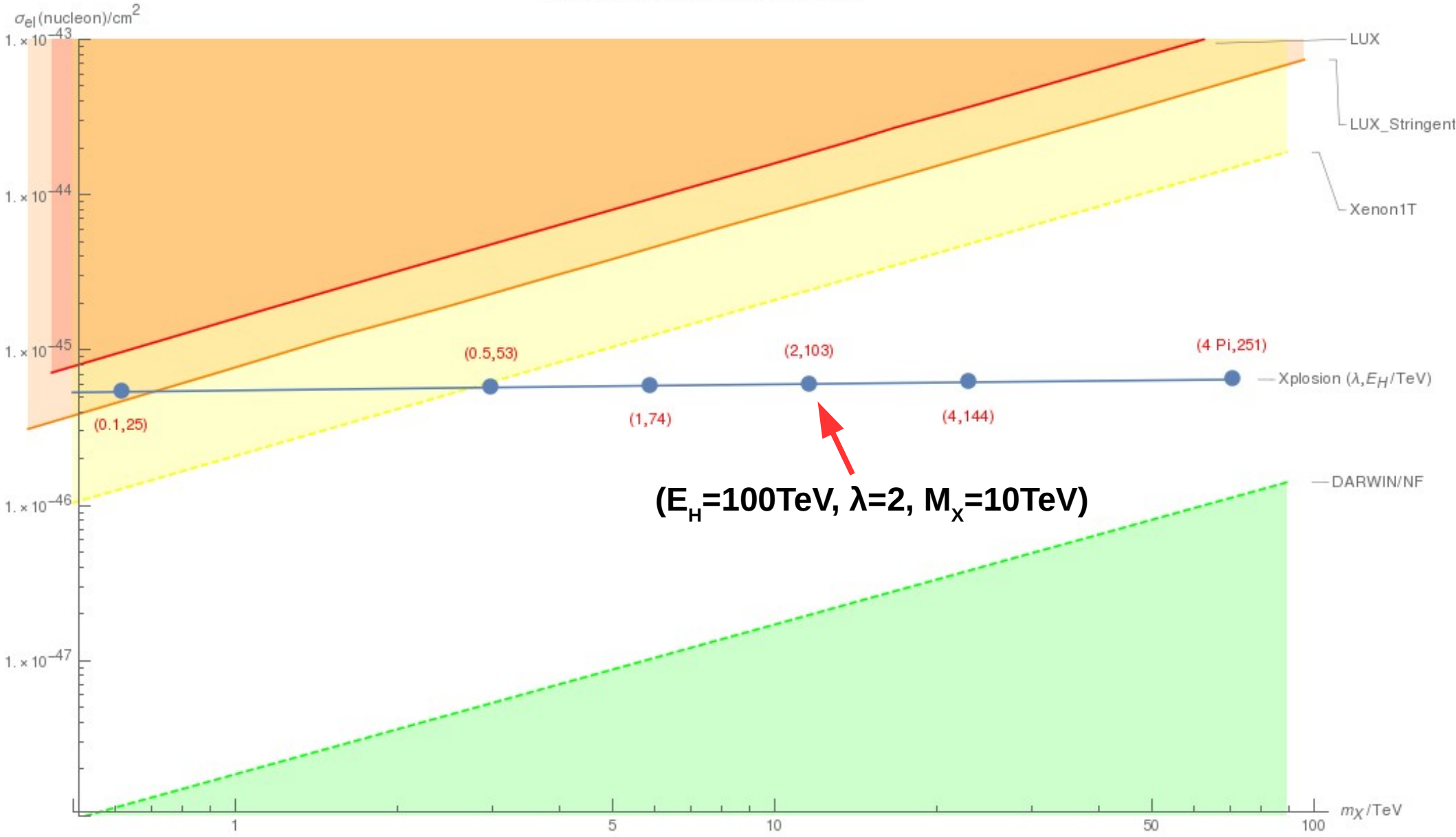
$$\sigma_{el} = 4\lambda_{HX}^2 \left(\frac{100\text{GeV}}{m_h = 125\text{GeV}} = 0.8 \right)^4 \left(\frac{50\text{GeV}}{m_X} \right)^2 (20 \times 10^{-42}\text{cm}^2).$$

- Demand 'perturbative' coupling, limiting solutions to a line segment!
- Plot with present and projected constraints from LUX, XENON-1T and DARWIN



$\langle\sigma v\rangle/\text{cm}^2$

Cross Section for different X masses



$(E_H=100\text{TeV}, \lambda=2, M_X=10\text{TeV})$

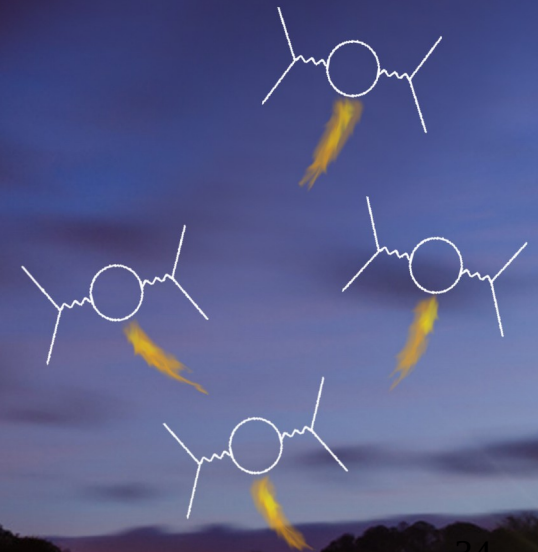
M_X/TeV



I don't like indirect detection...

Summary:

- Higgsploding DM can be fine, even at most primitive level
- In this case, still have small hierarchy problem
- Interesting to see: adding fermions, relaxing coupling assumption \rightarrow introducing important direct X self interactions





It's time for the talk to end...