Monte Carlo for Boosted Dark Matter

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Outline

What is boosted dark matter?

What do we need from boosted dark matter MC?

Boosted DM using GENIE

What is Boosted DM?

Relic DM has soft recoils





Nucleus

The difference with boosted

Elastic Nucleon Scattering



 $v_{\chi_B} \approx 0.6 c$



The difference with boosted

Inelastic Nucleon Scattering

Pions & other Hadrons



 $v_{\chi_B} \gg 0.6 c$

Simple BDM Models Exist



Two component Dark annihilation



Where else can we look?

Neutrino detectors are huge

(c) Kamioka Observatory, ICRR(Institute for Cosmic Ray Research), The University of Tokyo,

Solar capture & Detection

Solar Capture

Proton scattering



Solar capture & Detection



Solar Capture & Detection

Re-scattering





Solar Capture & Detection



Proton scattering



Looking with Water Cherenkov

Threshold: $E_{\rm K,recoil} = 480 \,\,{\rm MeV}$

Hard to reconstruct inelastic



Experiments: Super-Kamiokande Hyper-Kamiokande

Super-Kamiokande: PRD79 (2009) 112010

Water Cherenkov Results



A Future in Liquid Argon TPC

Threshold: $E_{\rm K,recoil} \lesssim 50 \,\,{ m MeV}$

Inelastic reconstruction possible

Experiments: LArIAT ICARUS MicroBooNE DUNE



LArSoft

What do we need from Boosted Dark Matter Monte Carlo?

Overview of Requirements

- Dark matter is one of the beams
- Dark matter flux is not fixed direction in lab frame
- Fixed target for scattering
- Covers a range of energy scales: elastic nucleon recoil, baryon resonance dominated inelastic, deep inelastic
- Scattering off of nuclei, so need nuclear effects

Fixed target kinematics primer



X: p/n for elastic, mess of hadrons for inelastic $q^2 = -Q^2 = (p'-p)^2 \& W^2 = k'^2$ $0 \le Q^2 \le 4 p_{1,CM}^2 \& M_N \le W \le \sqrt{s} - M_\chi$ Inelastic can begin at $\gamma \ge 1 + M_\pi / M_N$

Three Different Processes







Elastic

Relatively easy

Needs a form factor near 1 GeV Dominated by Δ , N1 GeV < W < 2 GeV

Resonant

Need a model

Use standard parton

Deep Inelastic

model

DM beam?

Rein & Sehgal: Ann. of Phys. 133, 79-153 (1981)

All Processes Could Be Important



Nuclear effects are important!



Fermi Motion

Pauli Blocking

Think of large nucleus as a **Fermi gas** For large nucleus (say Ar): *p*_F ~ **250 MeV**

Other nuclear effects are important!

Nuclear Rescattering! Including: elastic, inelastic, charge exchange, pion production, absorption, ...

Boosted DM in GENIE

What is GENIE?

- Monte Carlo for neutrino interactions
- Used by FNAL-based neutrino experiments & beyond
- Includes all of the above requirements for neutrinos
- ✤ Also has *n-n̄* oscillation simulation (history of BSM!)

L. Alvarez-Ruso, C. Andreopoulos, C. Barry, F. Bench , S. Dennis, S. Dytman, H. Gallagher, R. Hatcher, L. Jiang, R. Jones, I. Kakorin, K. Kuzmin, A. Meregaglia, D. Naples, V. Naumov, G. Perdue, M. Roda, J. Wolcott, J. Tena Vidal, J. Yarba

Current Status of BDM in GENIE

2 models: fermion or scalar DM, axial Z' coupling

Elastic and Deep Inelastic scattering implemented

Framework mostly set for further models

Integrated into upcoming GENIE v3

What You Need to Put In

gevgen dm [-h] [-r run#] -n nev -e energy (or energy range) -m mass -t target pdg [-g zp coupling] [-z med ratio] [-f flux description] [-o outfile name] [-w] [--seed random number seed] [--cross-sections xml file] [--event-generator-list list name] [--message-thresholds xml file] [--unphysical-event-mask mask] [--event-record-print-level level] [--mc-job-status-refresh-rate rate] [--cache-file root file]

Sample event

| GENIE GHEP Event Record [print level: 3] | | | | | | | | | | |
|--|----------|---------|------------|-------------|----------|--------|---------|--------------|---------------------|---------------|
| Idx | Name | Ist | PDG | Mother | Daughter | Px | Ру | Pz | E m | -i |
| | | | | _1 _1 | · | | 0 000 1 | 17 301 I | 20 000 L 10 000 | -1 |
| | Ar40 | 0 | 1000180400 | -1 -1 | 2 3 | 0.000 | 0.000 | 0.000 | 37.216 37.216 | |
| 2 | proton | 11 | 2212 | 1 -1 | 5 5 5 | 0.109 | -0.063 | 0.028 | 0.926 **0.938 | M = 0.917 |
| 3 | C139 | 2 | 1000170390 | 1 -1 | 8 8 | -0.109 | 0.063 | -0.028 | 36.290 36.290 | 1 |
| 4 | chi_dm | 1 | 2000010000 | 0 -1 | -1 -1 | 1.084 | -0.475 | 16.318 | 19.175 10.000 | P = |
| 5 | proton | 14 | 2212 | 2 -1 | 6 7 | -0.975 | 0.412 | 1.031 | 1.750 0.938 | FSI = 4 |
| 6 | proton | 1 | 2212 | 5 -1 | -1 -1 | -0.012 | 0.056 | 0.053 | 0.941 0.938 | |
| 7 | neutron | 1 | 2112 | 5 -1 | -1 -1 | -1.073 | 0.139 | 0.884 | 1.684 0.940 | |
| | HadrBlob | 15 | 200000002 | 3 -1 | -1 -1 | 0.054 | 0.154 | -0.075 | 35.356 **0.000 | M = 35.356 |
| 9 | NucBindE | 1 | | -1 -1 | -1 -1 | -0.026 | 0.123 | 0.118 | 0.029 **0.000 | M = -0.170 |
| | NucBindE | 1 | 200000101 | -1 -1 | -1 -1 | -0.027 | 0.004 | 0.022 | 0.029 **0.000 | M = -0.020 |
| Fir | n-Init: | | | | | 0.000 | -0.000 | -0.000 | | - _! |
| Vertex: chi_dm @ (x = 0.00000 m, y = 0.00000 m, z = 0.00000 m, t = 0.000000e+00 s) | | | | | | | | | | _1 |
| Err flag [bits:15->0] : 000000000000000 1st set: Err mask [bits:15->0] : 11111111111111 Is unphysical: NO Accepted: YES | | | | | | | | | | |
| sig(Ev) = 1.02127e-34 cm^2 dsig(Q2;E)/dQ2 = 3.32007e-37 cm^2/GeV^2 Weight = 1.00000 | | | | | | | | | | |

Another Sample Event

| | GHEP EVENT RECOR | a [pr: | int level: 3] | | | | | | | | | | |
|--|-----------------------|------------|---|--------|------------|----------|-------------------|-----------|------|--------|-----------|---------|----------|
| Idx | Name | Ist | PDG | Mot | her | Daught | er | Px | I | Ру | Pz | Е | m |
| 0 | chi_dm | 0 | 2000010000 | -1 | 1- | 4 | 4 | 0.000 | | 0.000 | 17.321 | 20.000 | 10.000 |
| 1 | Ar40 | 0 | 1000180400 | -1 | -1 | 2 | 3 | 0.000 | | 0.000 | 0.000 | 37.216 | 37.216 |
| 2 | neutron | 11 | 2112 | 1 | -1 | 5 | 5 | -0.020 | | -0.071 | -0.205 | 0.929 | **0.940 |
| 3 | Ar39 | 2 | 1000180390 | 1 | -1 | 16 | 16 | 0.020 | | 0.071 | 0.205 | 36.286 | 36.286 |
| 4 | chi_dm | 1 | 2000010000 | 0 | -1 | -1 | -1 | -0.614 | | 0.353 | 15.958 | 18.846 | 10.000 |
| 5 | HadrSyst | 12 | 200000001 | 2 | -1 | 6 | 8 | 0.594 | Ι | -0.424 | 1.158 | 2.083 | **0.000 |
| 6 | neutron | 14 | 2112 | 5 | -1 | 9 | 9 | 0.273 | | -0.296 | 0.574 | 1.172 | 0.940 |
| 7 | pi+ | 14 | 211 | 5 | -1 | 13 | 14 | 0.148 | | 0.053 | -0.049 | 0.216 | 0.140 |
| 8 | pi- | 14 | -211 | 5 | -1 | 15 | 15 | 0.172 | | -0.181 | 0.633 | 0.695 | 0.140 |
| 9 | HadrClus | 16 | 200000300 | 6 | -1 | 10 | 12 | 0.273 | | -0.296 | 0.574 | 1.172 | **0.000 |
| 10 | proton | 1 | 2212 | 9 | -1 | -1 | -1 | -0.182 | | -0.362 | 0.153 | 1.033 | 0.938 |
| 11 | proton | 1 | 2212 | 9 | -1 | -1 | -1 | 0.353 | | -0.071 | 0.109 | 1.011 | 0.938 |
| 12 | neutron | 1 | 2112 | 9 | -1 | -1 | -1 | 0.102 | | 0.137 | 0.312 | 1.005 | 0.940 |
| 13 | pi+ | 1 | 211 | 7 | -1 | -1 | -1 | 0.038 | | -0.107 | 0.039 | 0.184 | 0.140 |
| 14 | neutron | 1 | 2112 | 7 | -1 | -1 | -1 | -0.080 | | 0.228 | -0.019 | 0.970 | 0.940 |
| 15 | pi- | 1 | -211 | 8 | -1 | -1 | -1 | 0.172 | | -0.181 | 0.633 | 0.695 | 0.140 |
| 16 | HadrBlob | 15 | 200000002 | 3 | -1 | -1 | -1 | 0.210 | I | 0.004 | 0.136 | 33.472 | **0.000 |
| | Fin-Init: | | | | | | ا | 0.000 | | -0.000 | I 0.000 I | 0.000 | |
| | Vertex: | chi_ | dm @ (x = 0 | 0.0000 | 0 m, y | / = | 0.000 | 00 m, z = | = | 0.000 | 00 m, t = | 0.00000 |)e+00 s) |
| Err | flag [bits:15->0] | : 00(| 000000000000000000000000000000000000000 | | 1st s | et: | _ | | | | | | none |
| Err | mask [bits:15->0] | : 11: | | | Is un | iphysica | .I: <u>-</u> - | NO 2 | Acc | epted: | YES | | |
| sig(Ev) = 5.68527e-35 cm^2 d2sig(x,y;E)/dxdy = 1.66546e-33 cm^2 Weight = 1.00000 | | | | | | | | 1.00000 | | | | | |

What do we need for DUNE?



Develop a Monte Carlo

Based on **GENIE** neutrino MC Does **inelastic** and **nuclear** effects Merged into GENIE v3

Integrate into LArSoft detector simulation

Develop analysis strategy

Projections for DUNE TDR

Theory: JB, Cui, Joglekar, Necib, Zhao Experiment: Petrillo, Russell, Tsai

Other Future Directions

- Include baryon resonance production
- More models: general interaction, inelastic DM, etc.
 See talk today by Doojin Kim
- Fluxes for various common sources
- Tweaks and debugging

Conclusions

Boosted DM Monte Carlo is here!

- Boosted dark matter is a new, well motivated dark matter scenario arising in several recent models
- Studying these models at large neutrino detectors requires new Monte Carlo tools
- These tools have now been developed within GENIE and should hopefully see official release shortly