Neutrino Astrophysics

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Ryan Nichol UK Input to European Strategy



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Highly cited papers

- Higgs Discovery (2012)
 8000+ citations
- Atmospheric neutrino oscillations (1998)
 –5000+ citations
- Top quark discovery (1995)
 - -3000+ citations
- Solar neutrino oscillations (2002)

-3000+ citations

Kaon CP violations (1964)

-3000+ citations

- Reactor antineutrino [KamLAND] (2003)
 –2000+ citations
- Gravitational Waves (2016)

-2000+ citations

- c-quark discovery (1974)
 –2000+ citations
- Solar neutrino [Homestake] (1968-1998)
 –2000+

- Z discovery (1983) -2000+
- Positron excess [PAMELA] (2008) –1900+
- Reactor neutrino theta_13
 (2012)

-1900+

- b-quark discovery (1977)
 –1900+
- W-discovery (1983) -1800+
- Z width (2005) -1700+
- Proton spin crisis (1989)
 –1700+
- LSND anomaly (2001) -1700+
- Parity non conservation (1957) –1600+
- LUX Dark Matter (2013) –1500+
- g-2 (2006)

-1500+

- SN1987a neutrinos (1987)
 –1500+
- Weak neutral current (1973) –1500+
- Charmomium (2003) -1400+
- B-Bbar oscillation (1987)
 –1300+
- nu_mu -> nu_e (2011) -1300+
- Accelerator neutrino oscillation (2002) –1100+
- Muon neutrino discovery (1962) –1100+
- DAMA/Libra (2008-) -1000+
- Pentaquark [LEPS] (2003)
 –1000+
- Neutron EDM (2006)₃ -1000+

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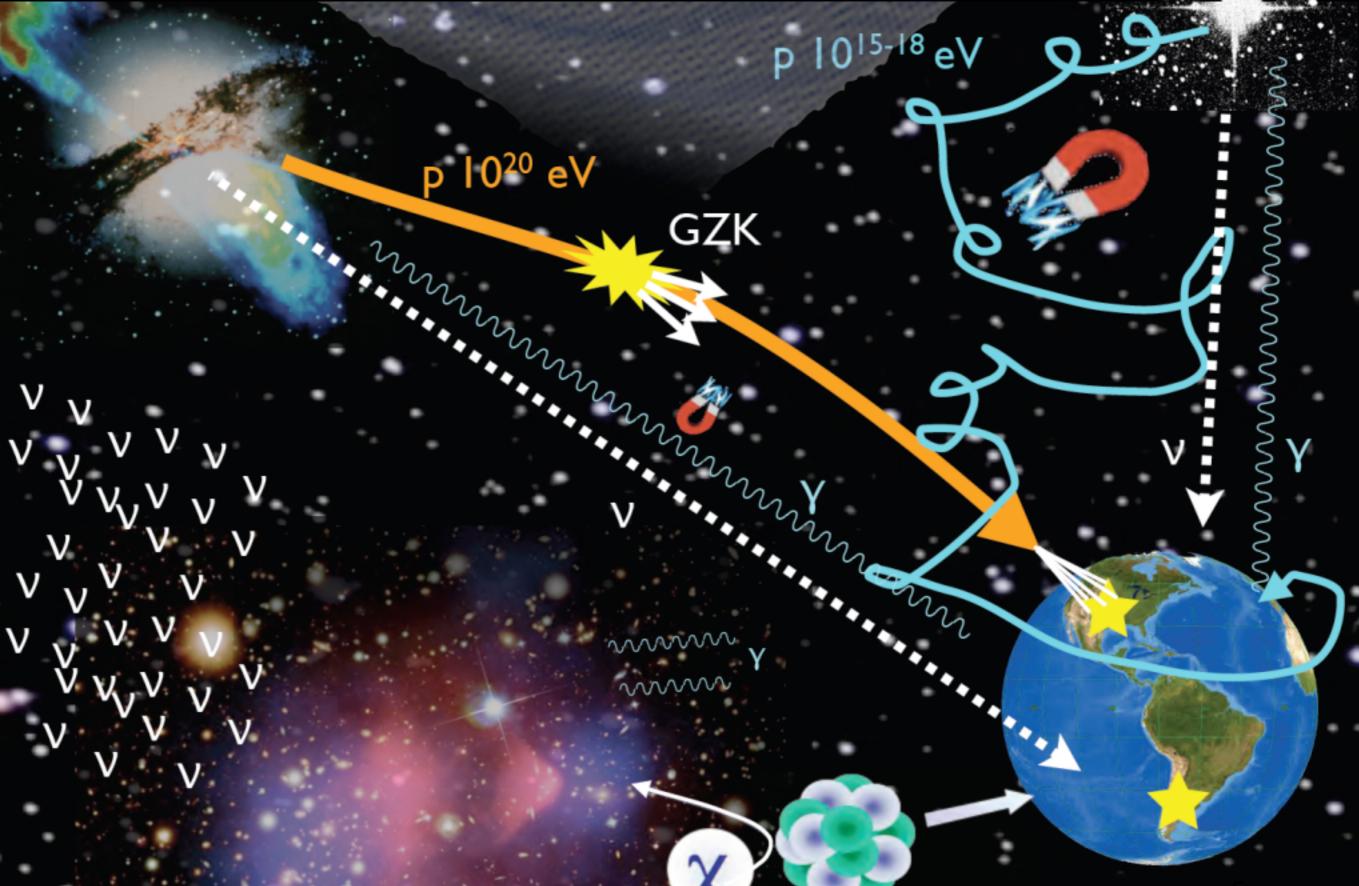
-1500+

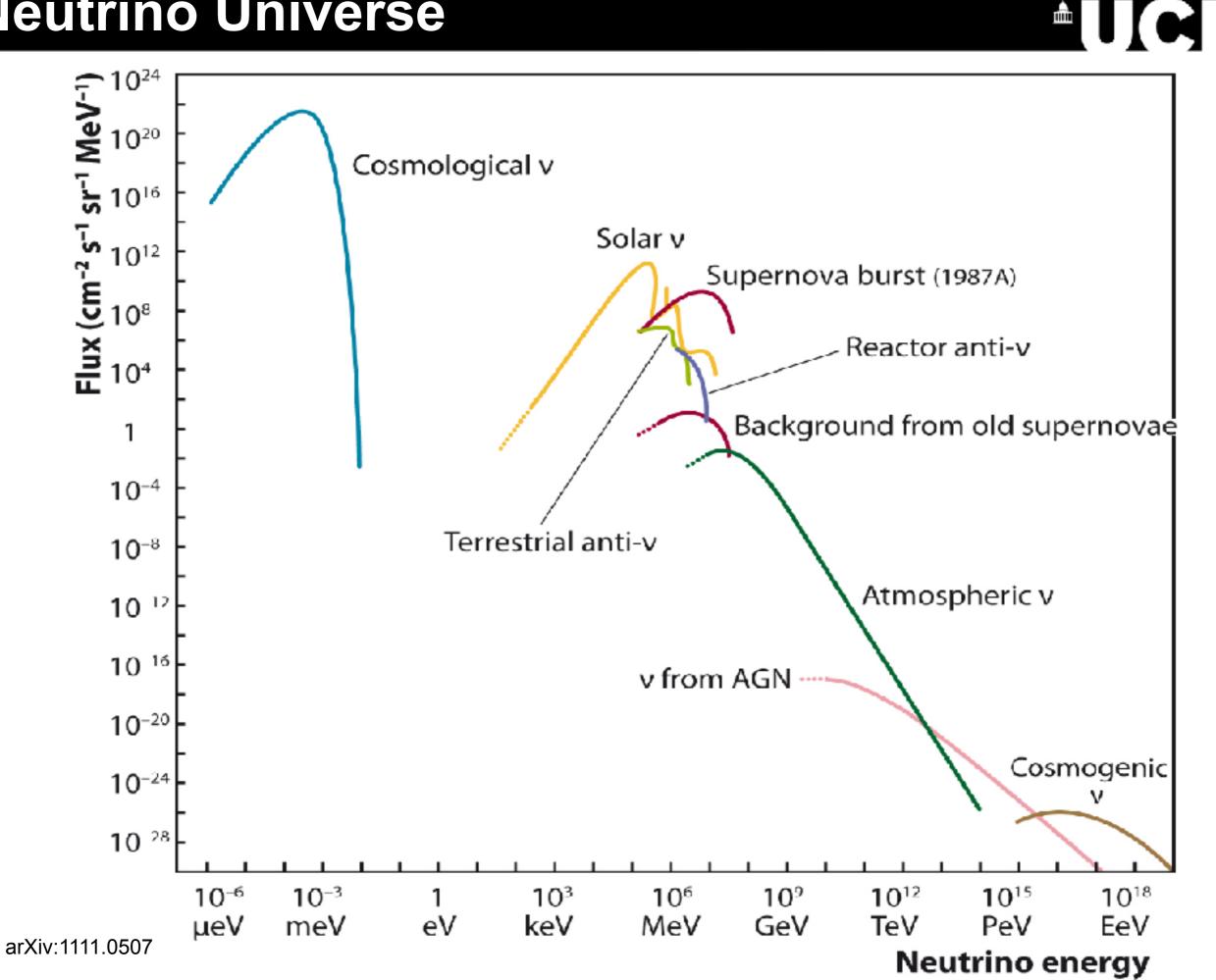
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-1100+

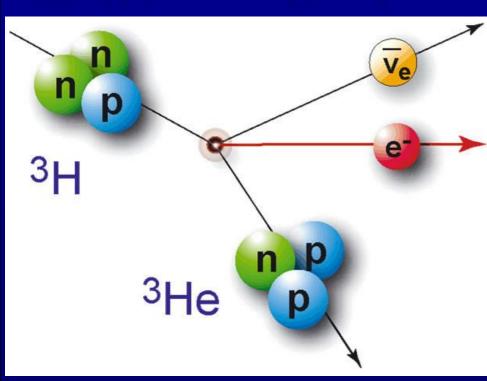
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- Neutron EDM (2006)₄ -1000+

Big Picture (from Johannes Blümer)





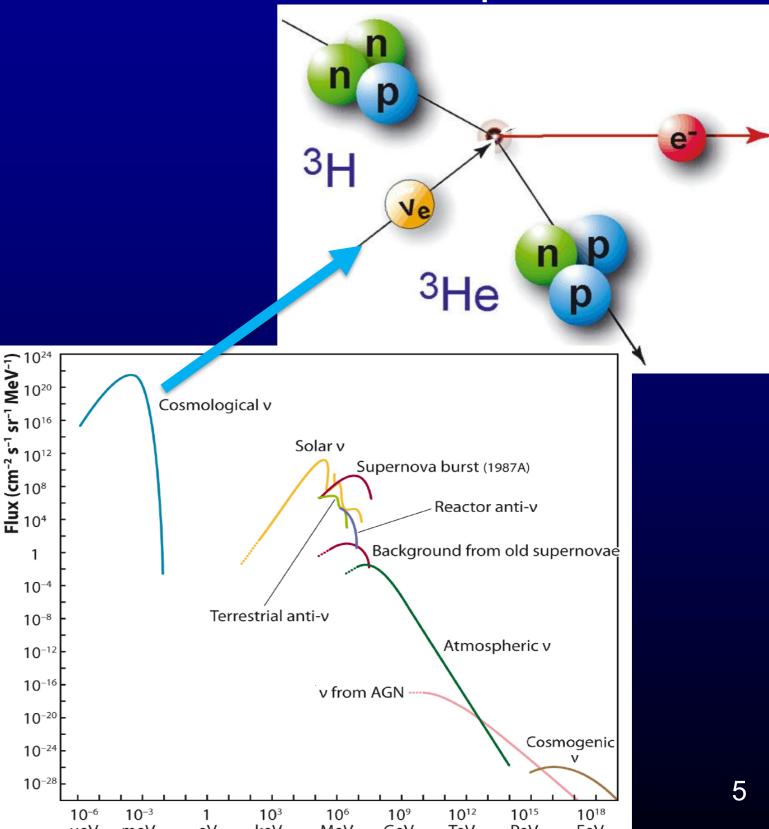
Tritium β-decay and Neutrino Capture

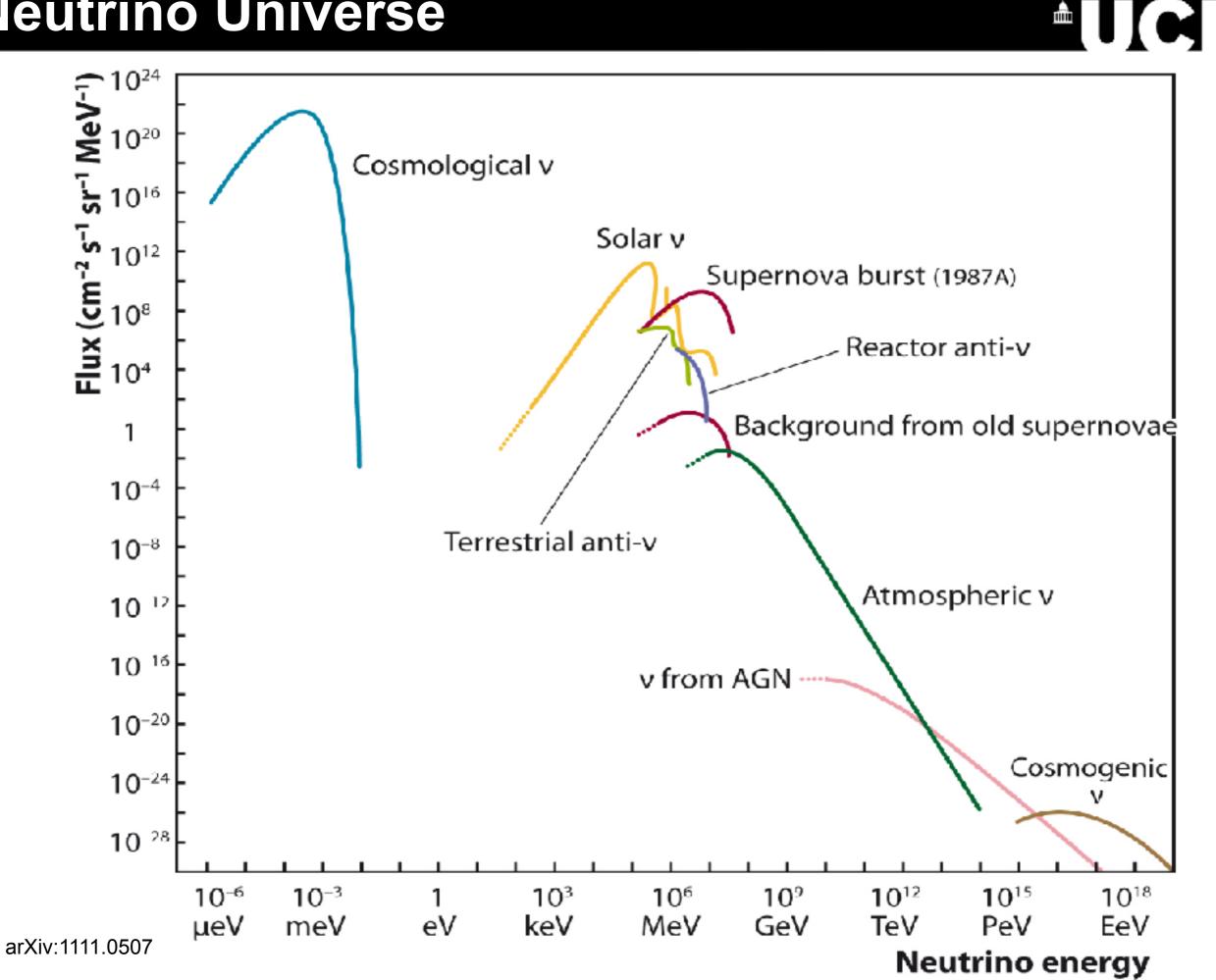


Tritium β-decay (12.3 yr half-life)

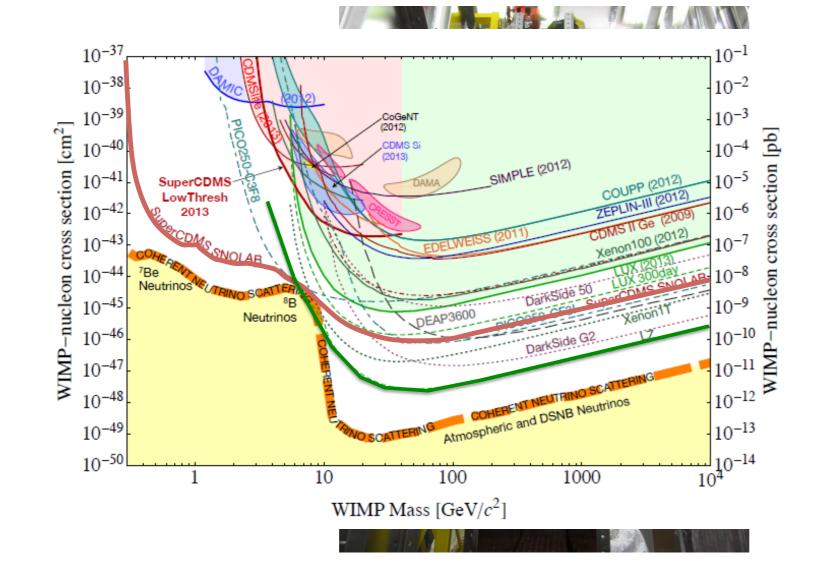
PTOLEMY, Chris Tully https://indico.cern.ch/event/640340

Neutrino capture on Tritium

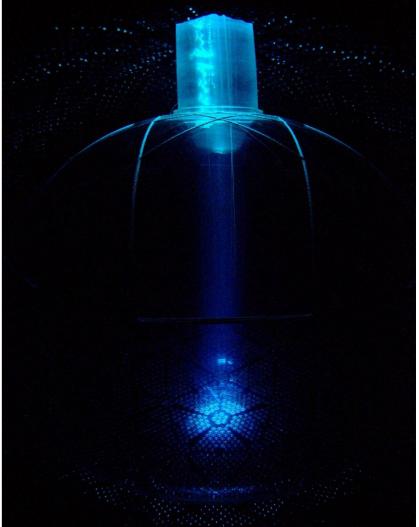




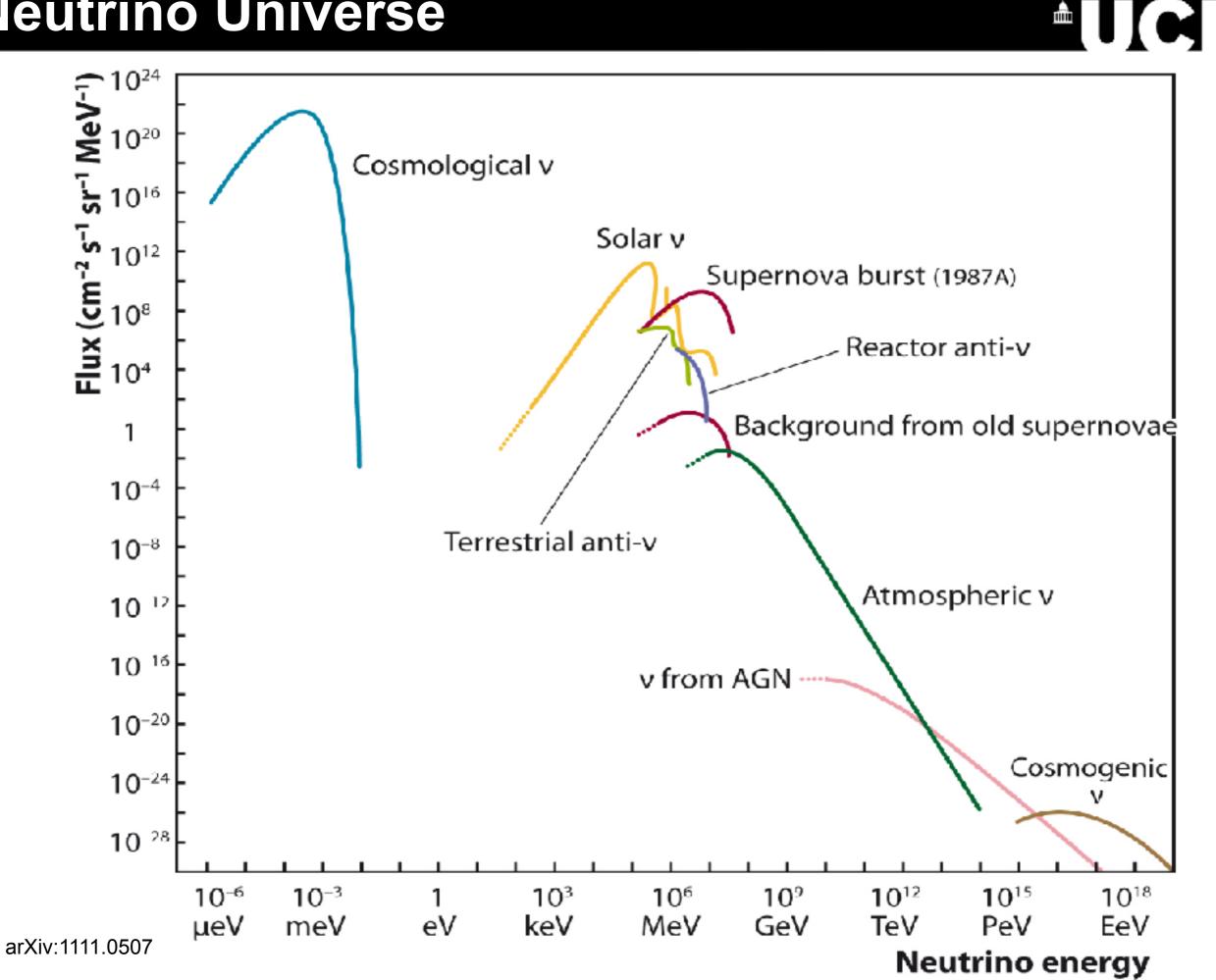
Solar Neutrinos



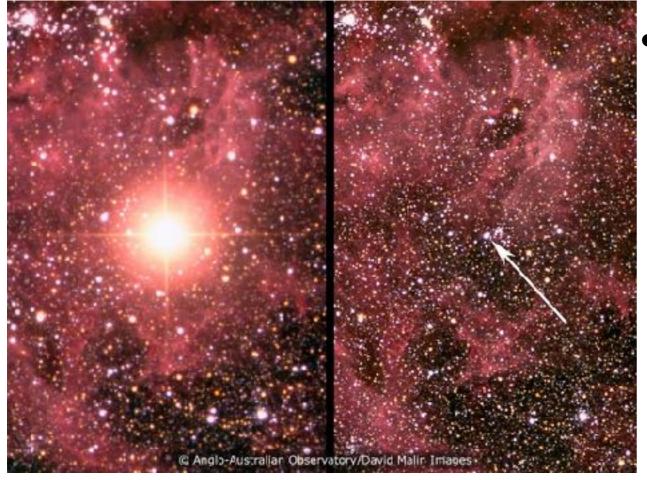


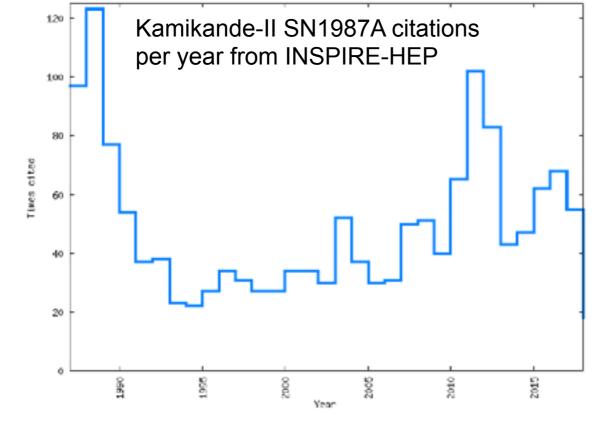






Aside: Super Nova Neutrinos



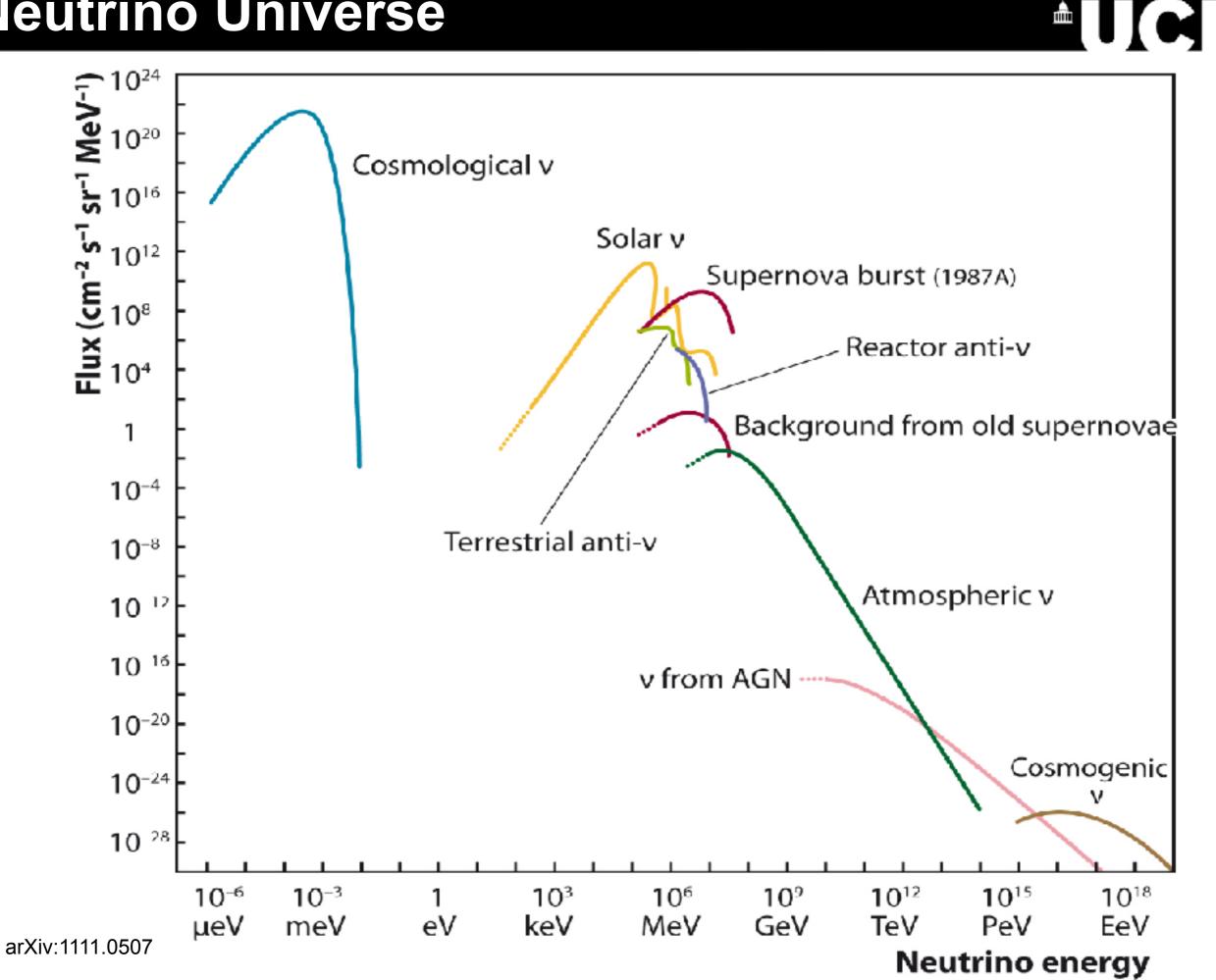


• SN1987A

–24 neutrino events
 detected by Kamikande-II,
 IMB and Baksan

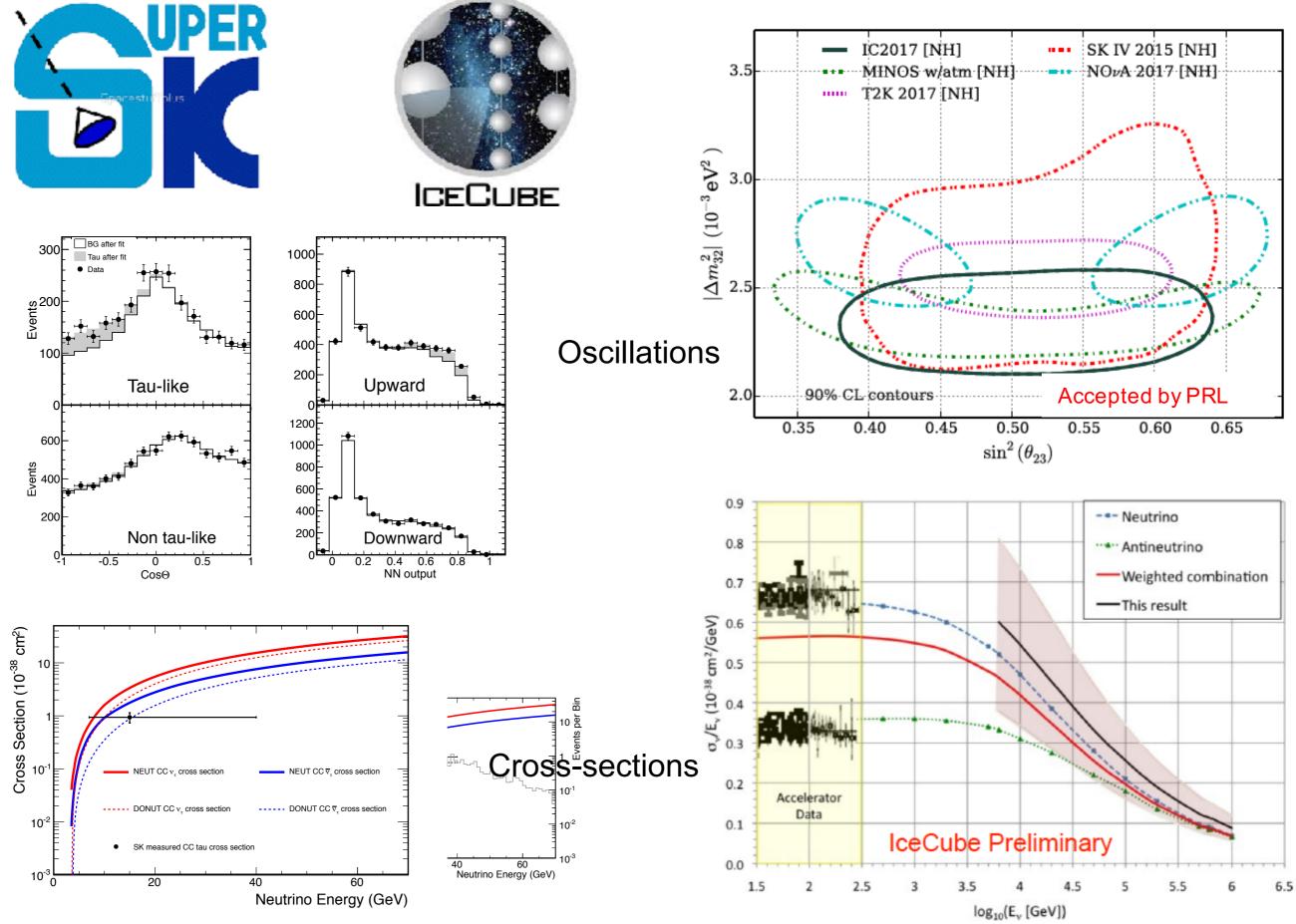
- -Learned about
 - Supernova collapse mechanisms
 - Neutrinos feel gravity (similarly to photons)
 - Neutrino mass < 23eV from time of flight dispersion
 - Neutrinos are not charged
 - Limits on non-neutrino weakly interacting particles
 - Axion bounds
 - Neutrino mixing and oscillations
 - Exotic neutrino disappearahce

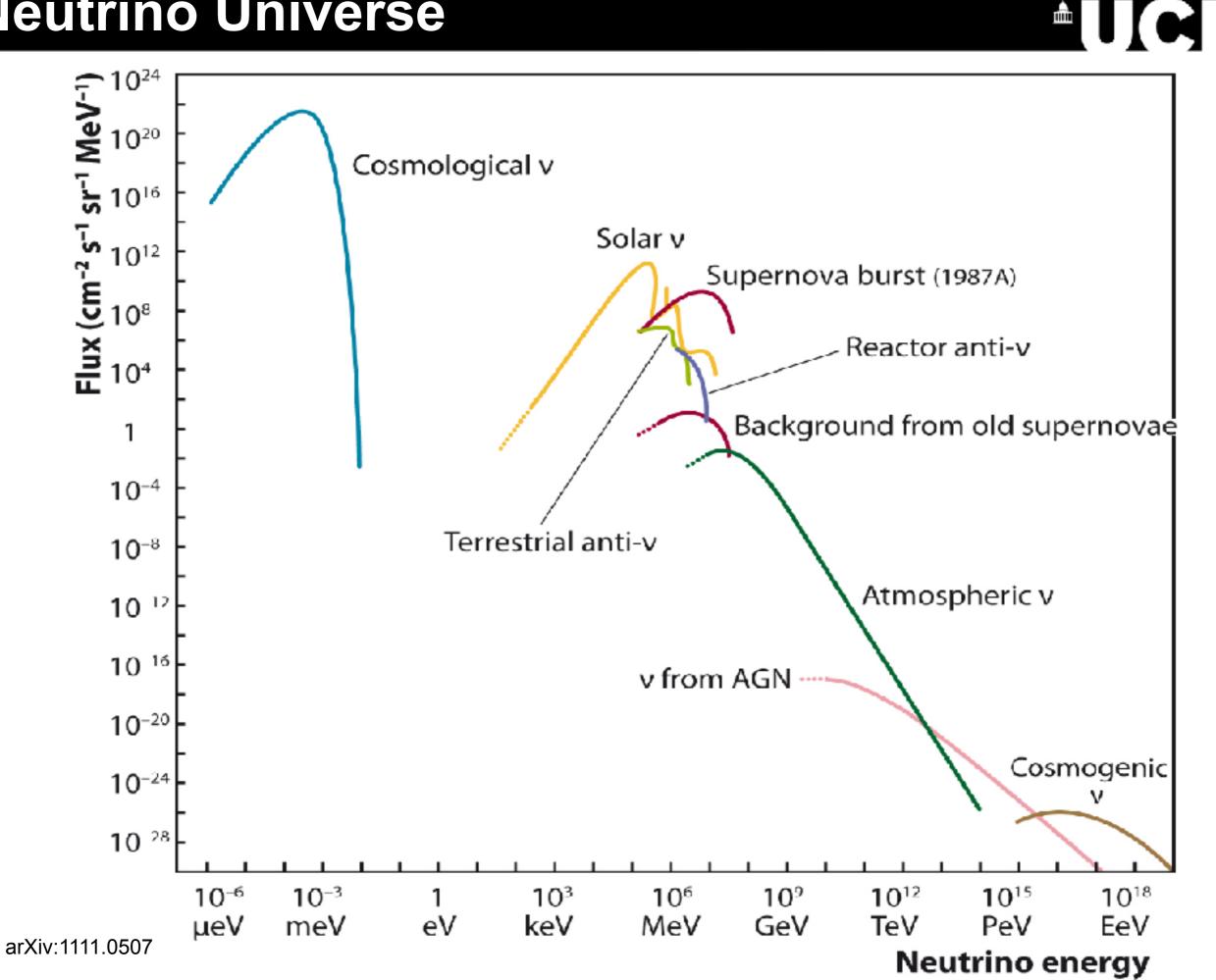
* before 2011, excluding solar

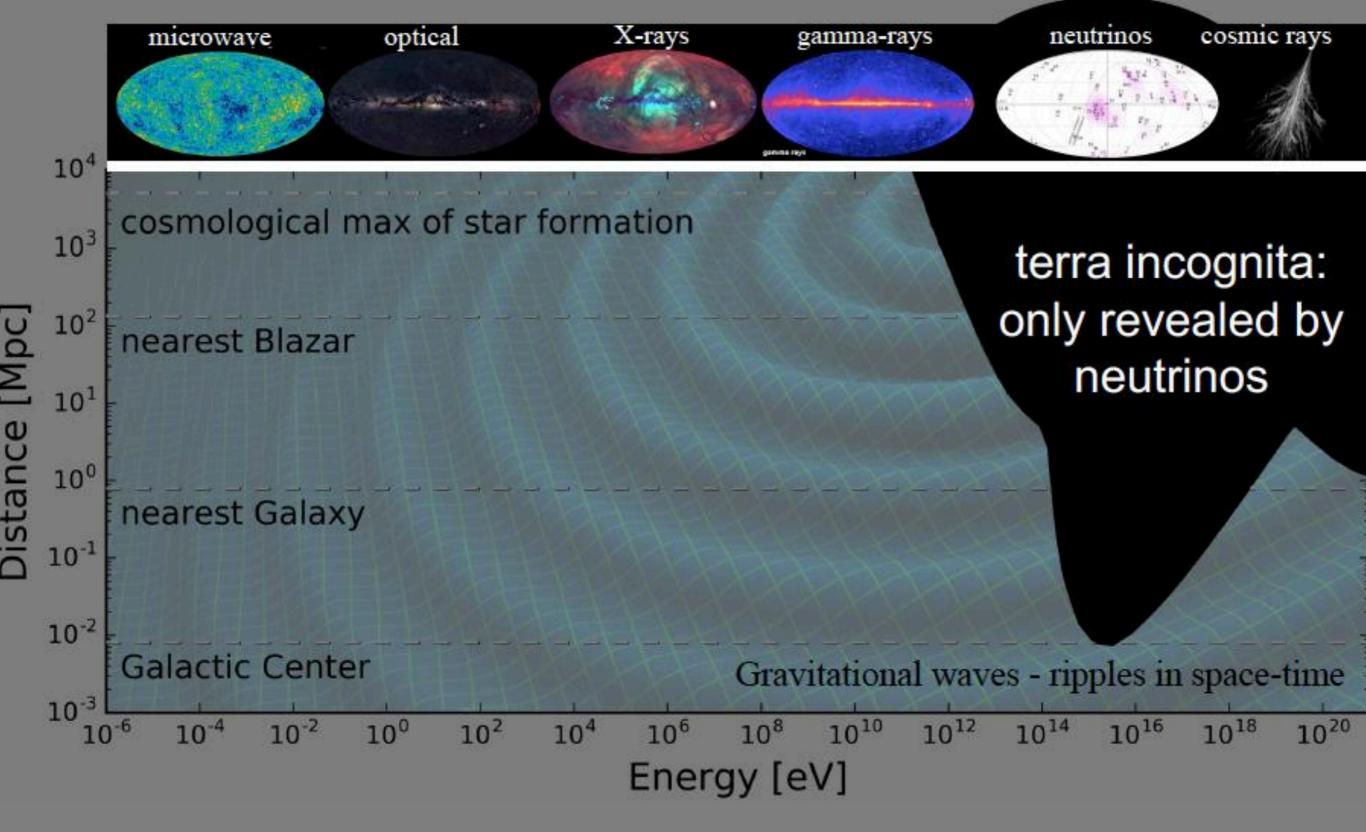


Atmospheric Neutrinos







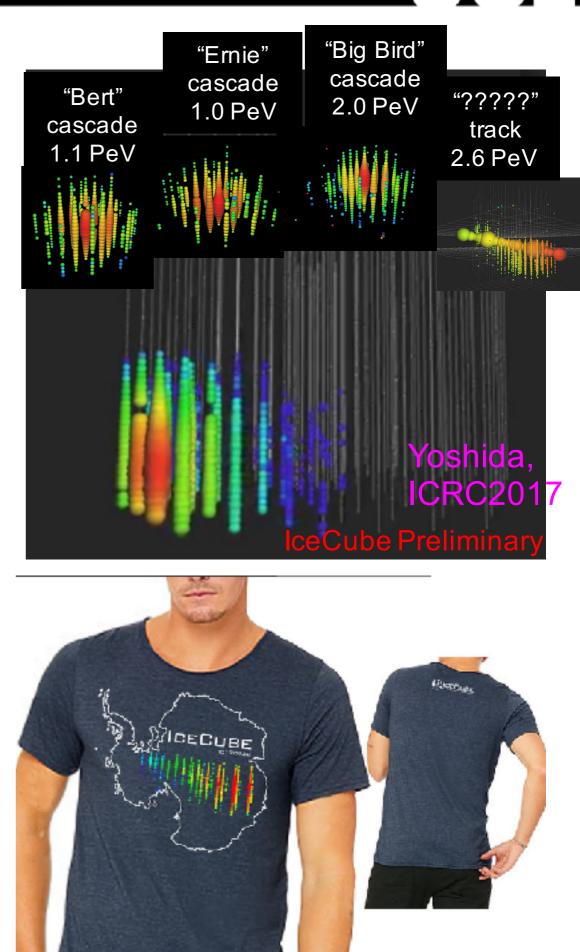


- 20% of the Universe is opaque to the EM spectrum
- non-thermal Universe powered by cosmic accelerators
- probed by gravity waves, neutrinos and cosmic rays

IceCube

- IceCube detected 5th PeV neutrino (name: ????)
 - -Cascade type
 - -~2.7 PeV (the highest energy)
 - –unlikely to be cosmogenic neutrino

- They are selling new IC170922 t-shirts
 –Point source candidate?
 - -Blazar coincidence?



Future: IceCube-Gen2

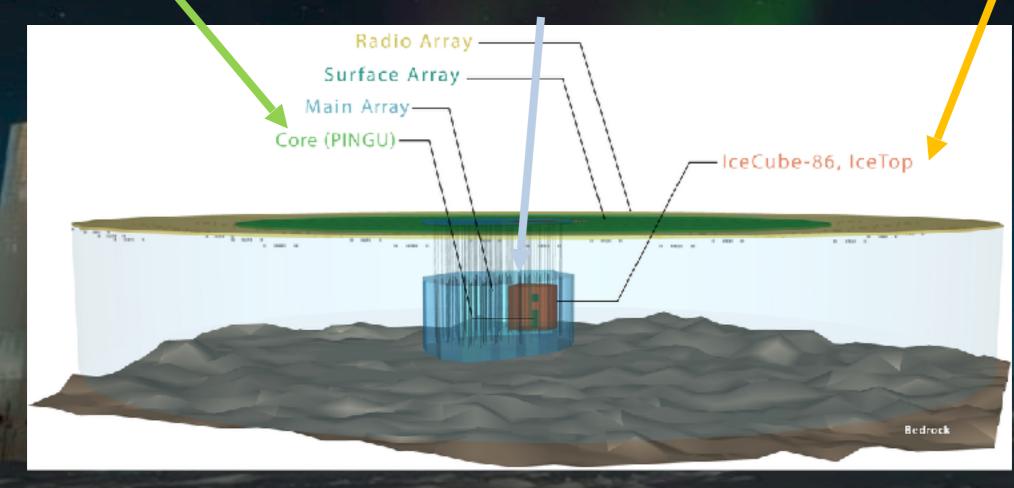
IceCube-Gen2 covers particle physics from MeV to EeV with real discovery potential

PINGU (GeV)

- ~20m spacing dense array
- neutrino mass ordering

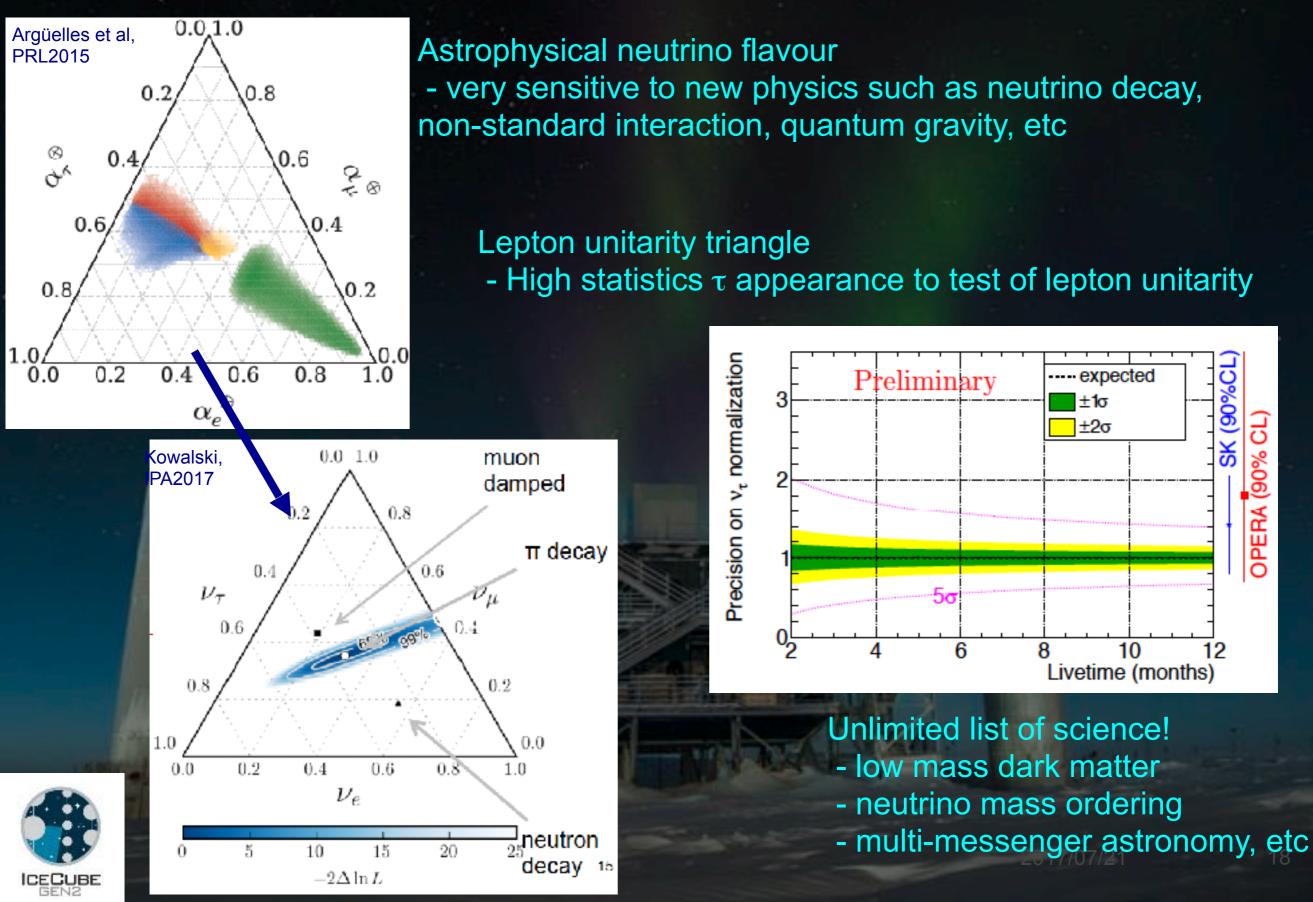
Askaryan Radio Array (EeV) - GZK neutrinos

- Main array (TeV-PeV)
- 120 new strings, 80 DOMs per string
- 240m separation to cover x8 volume
- x2 QE PMTs, and/or new photo-sensors





Physics of IceCube-Gen2



IceCube-Gen2 in UK

IceCube data analysis

- Mass ordering analysis on DeepCore
- Test of quantum gravity

Software development

- Atmospheric flux systematics Evans et al, PRD95(2017)023012
- Hadronization systematics Katori et al, JPhysG42(2015)115004
 PINGU fast oscillation analysis code paper in preparation

Hardware

- FEB firmware development
- DOM Fermilab beam test

paper in preparation

Analysis coordination

- Gen2 low E convener (Justin Evans)

On top of these, there is a large theory contribution from Oxford (Subir Sarkar)

LHCb for prompt-v production, JHEP02(2016)130 High energy neutrino cross section, JHEP08(2011)042 etc

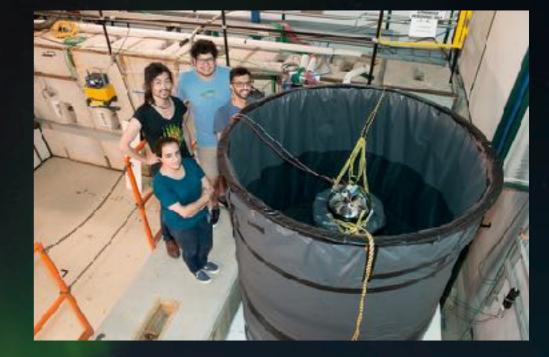


Manchester: J. Evans, S. Söldner-Rembold, S. Wren Oxford: S. Sarkar Queen Mary: T. Katori, S. Mandalia









The IceCube-Gen2 Collaboration

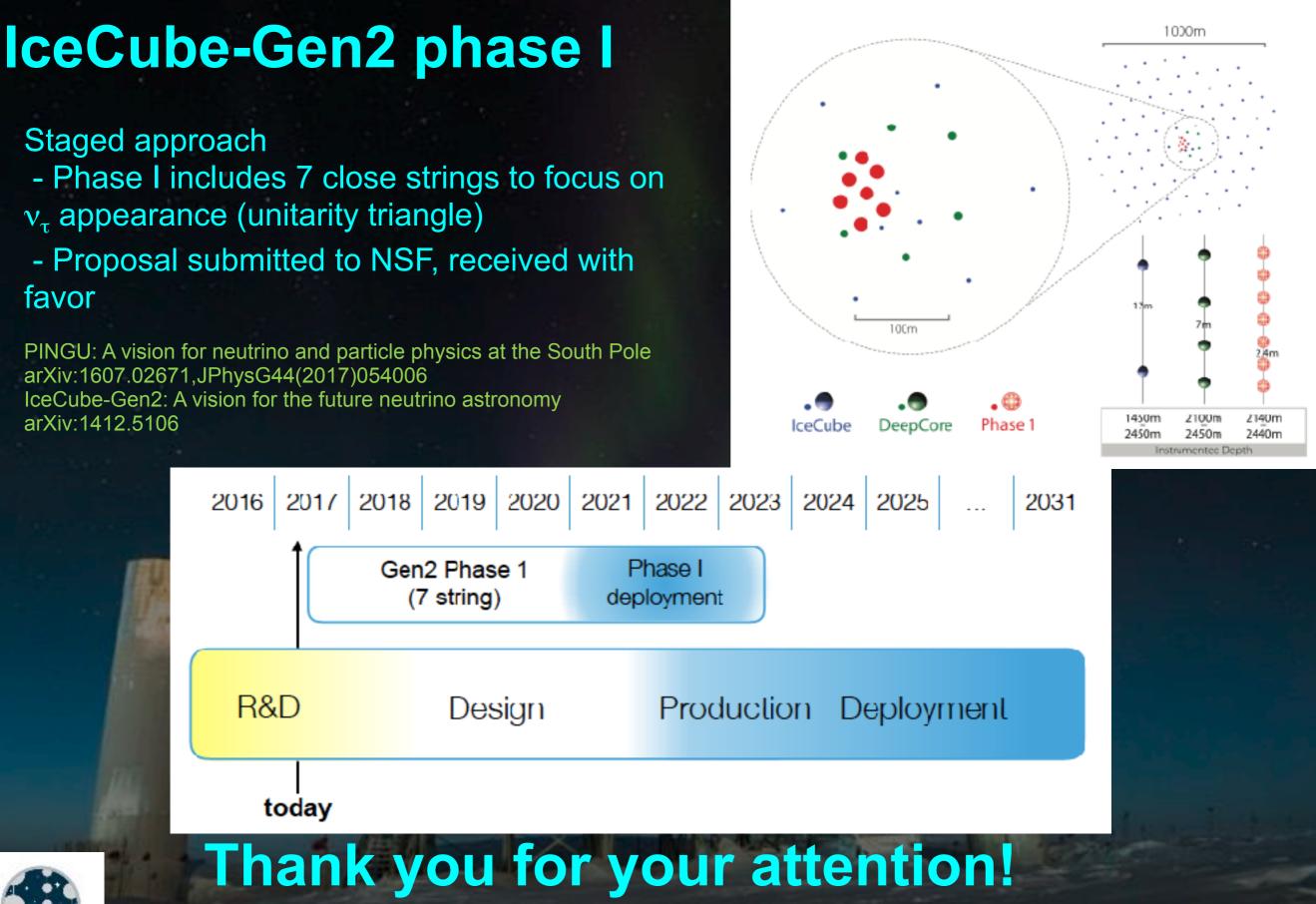


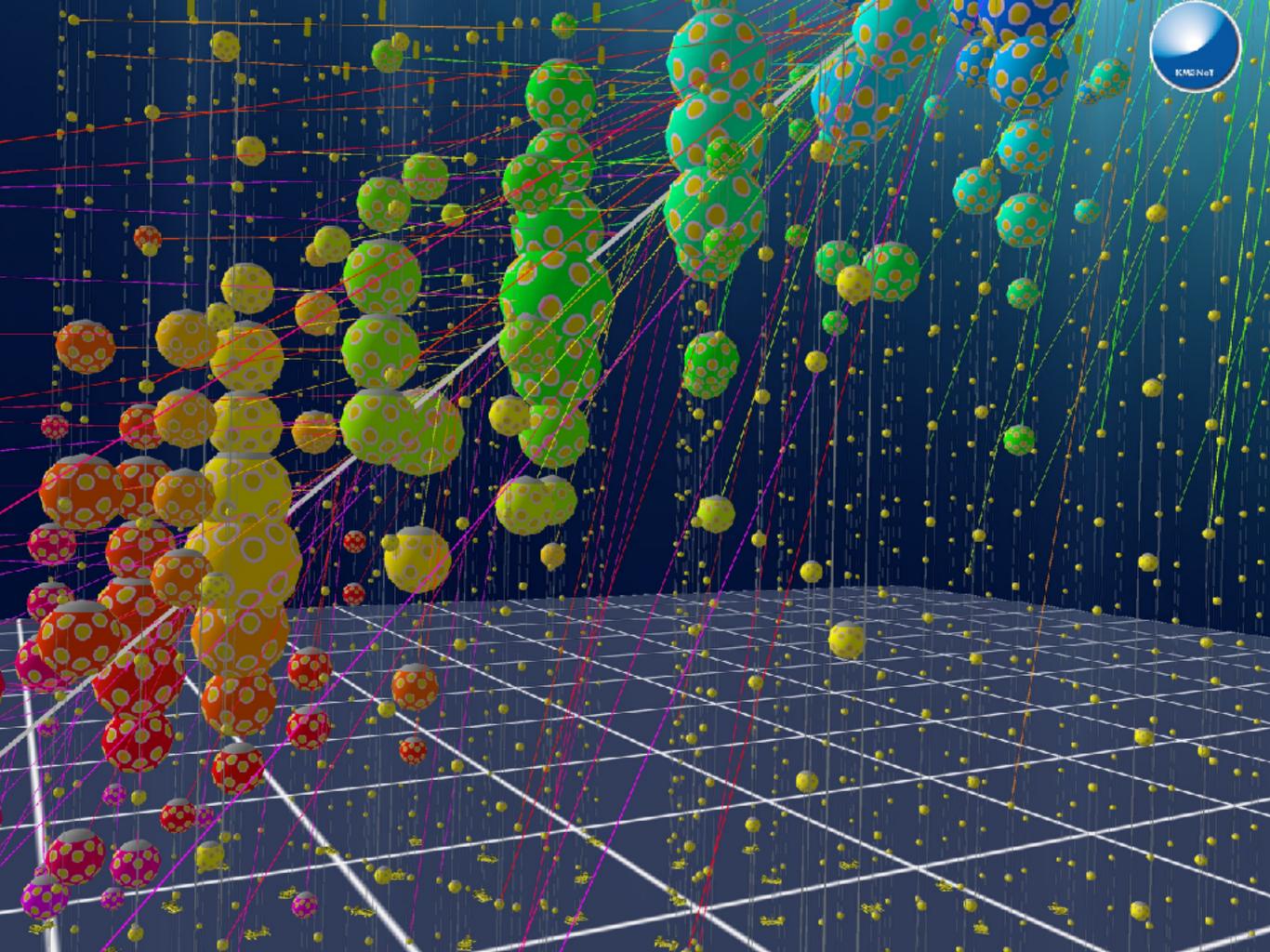
Factor San Balance State (1997) Factor Weinerschappel (in Onderson-Visa science (FWD-Visascience) Federal History of Education & America (BHSF) German Research Federation & (BPG) Deutschen Biolanzmein Synchronismi (DEST) lande Feutralation für Science, Japan Krist and Allice Walkenerg Feuerdation NSF-officer i Polar Programs NSF-officer i Polar Programs

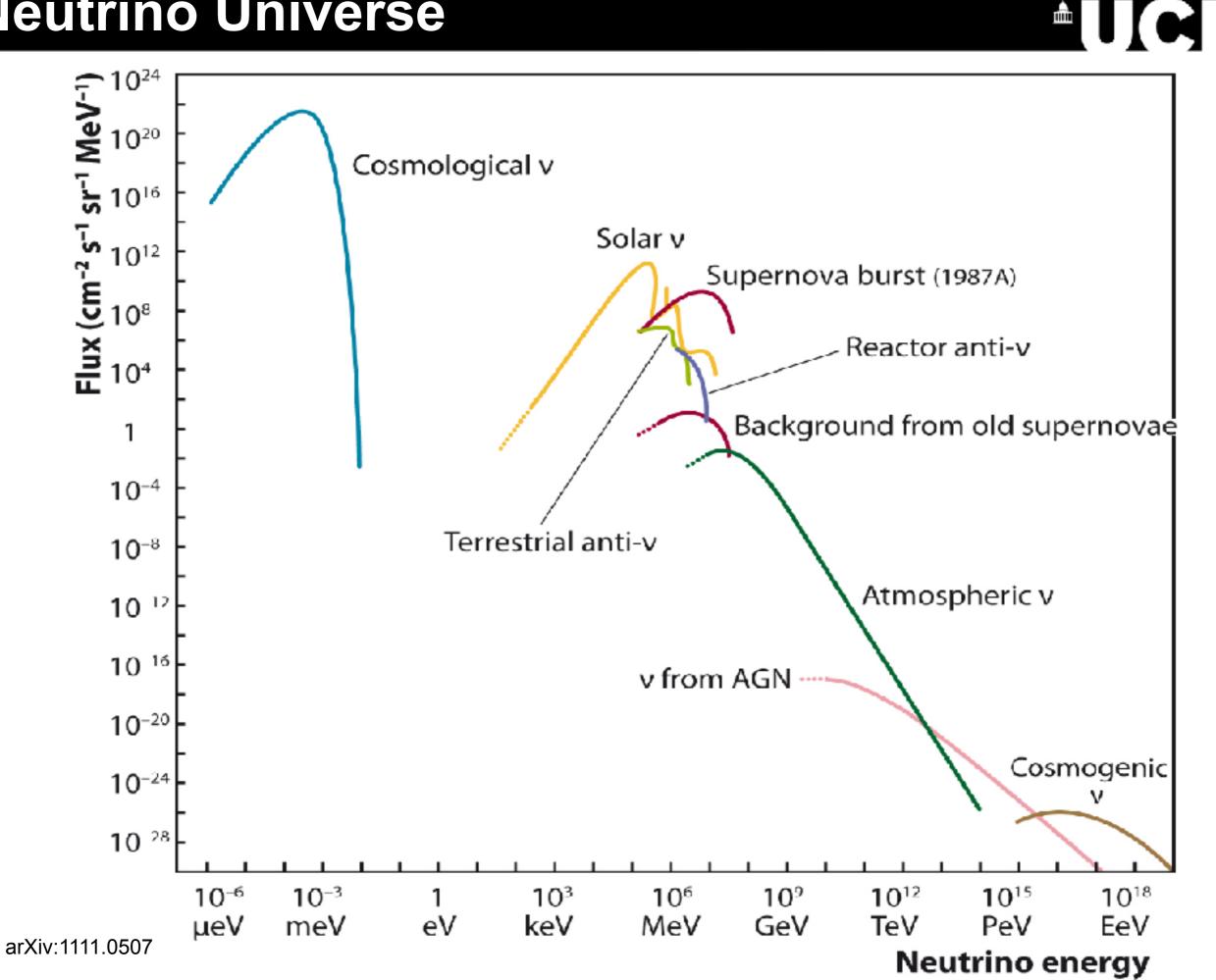
versity of Wisconso Ali

S National Science Foundation (NSE

For nutrition (WARF)

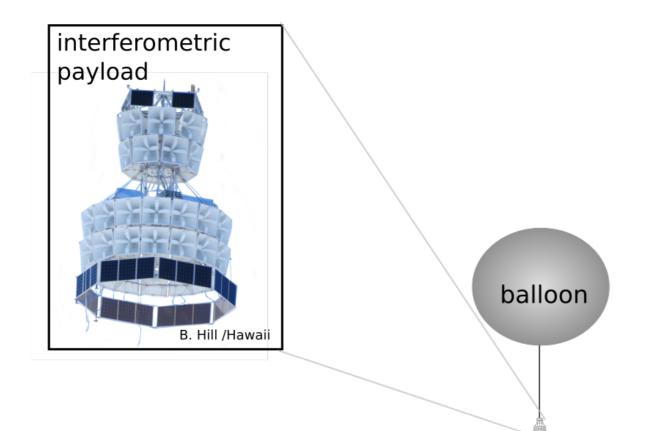




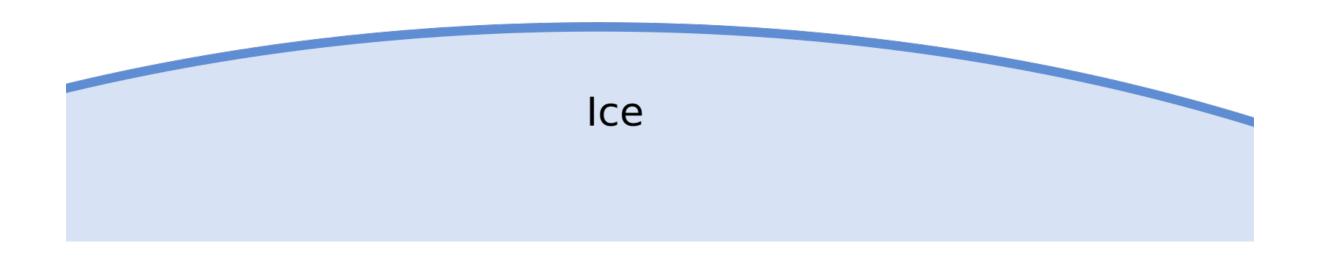


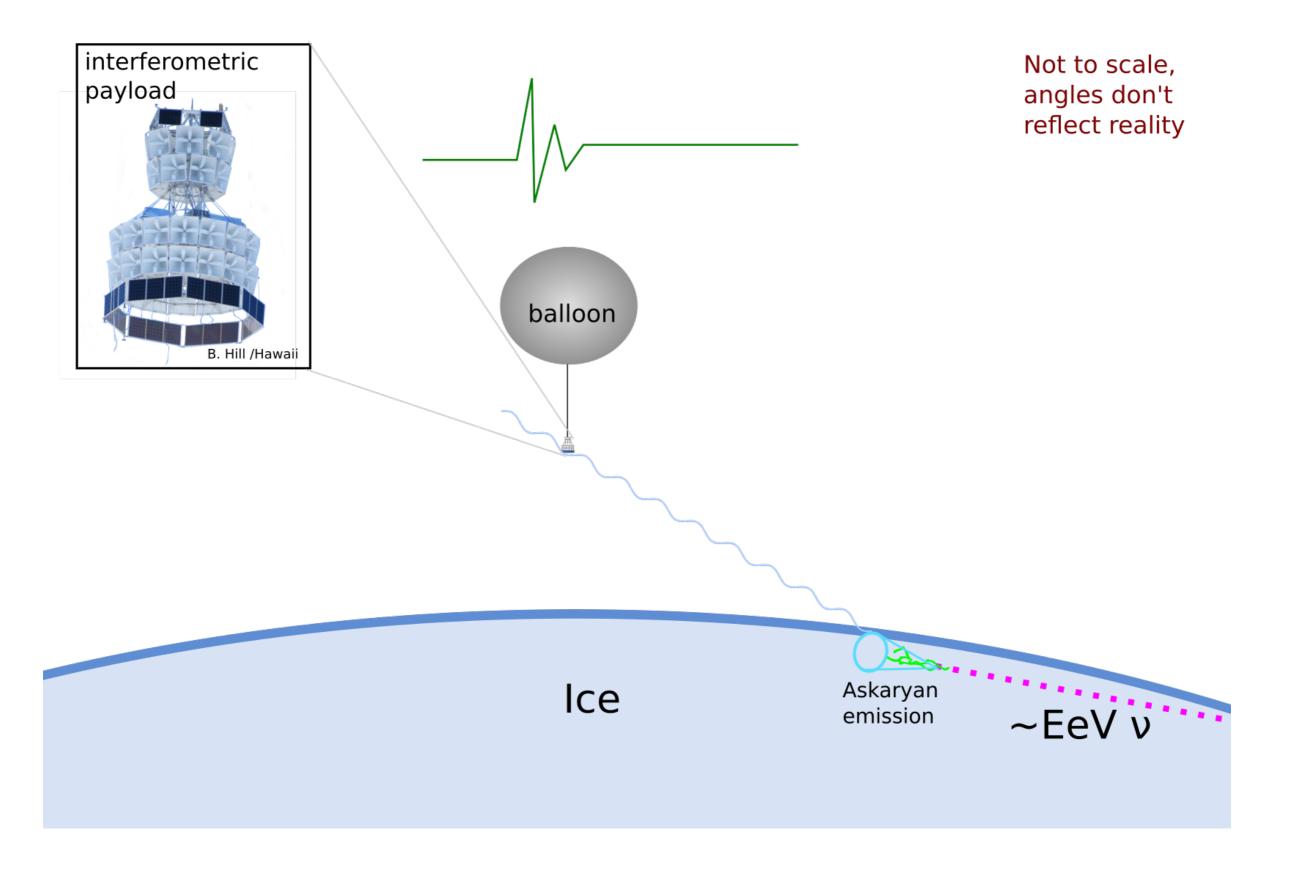
ANITA

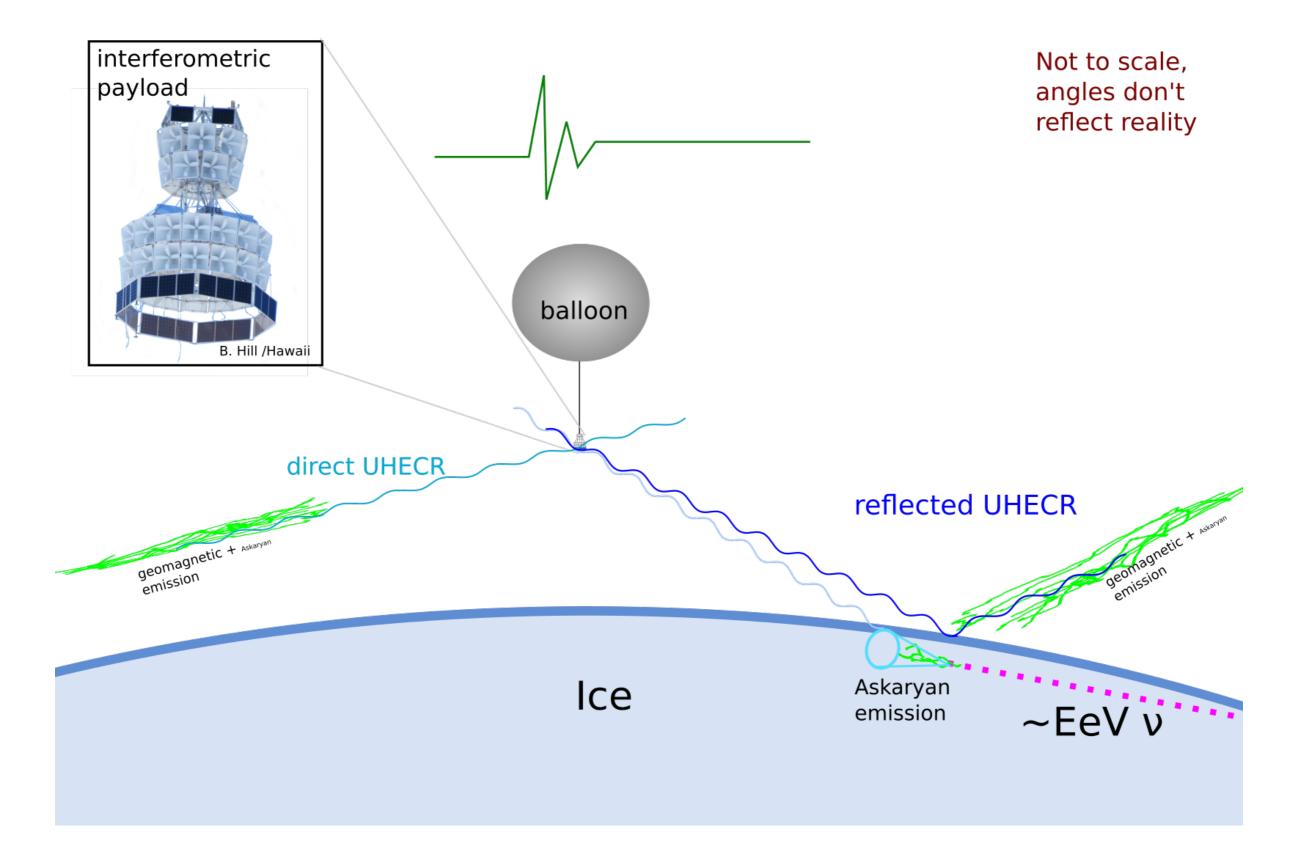
Photo: H. Schoorlemmer, University of Hawaii

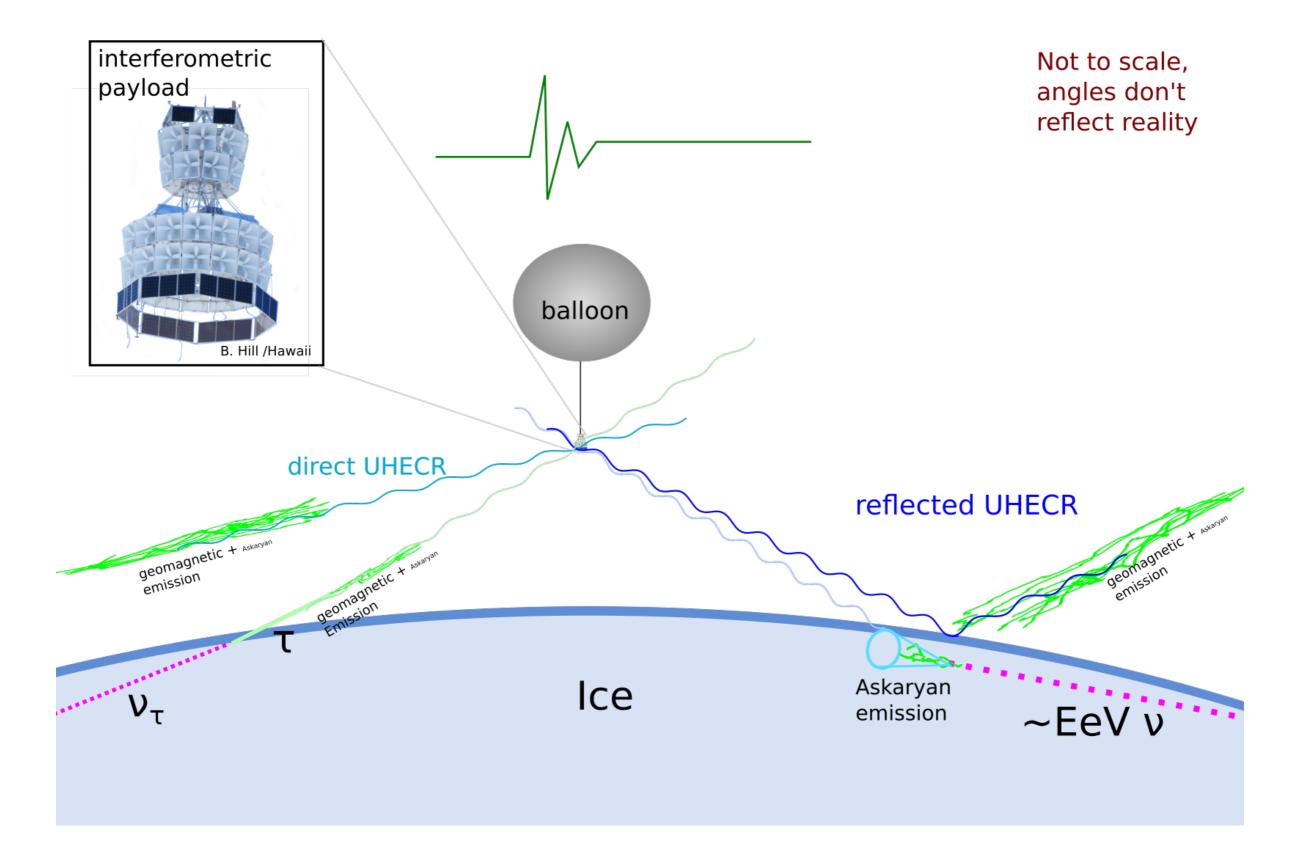


Not to scale, angles don't reflect reality

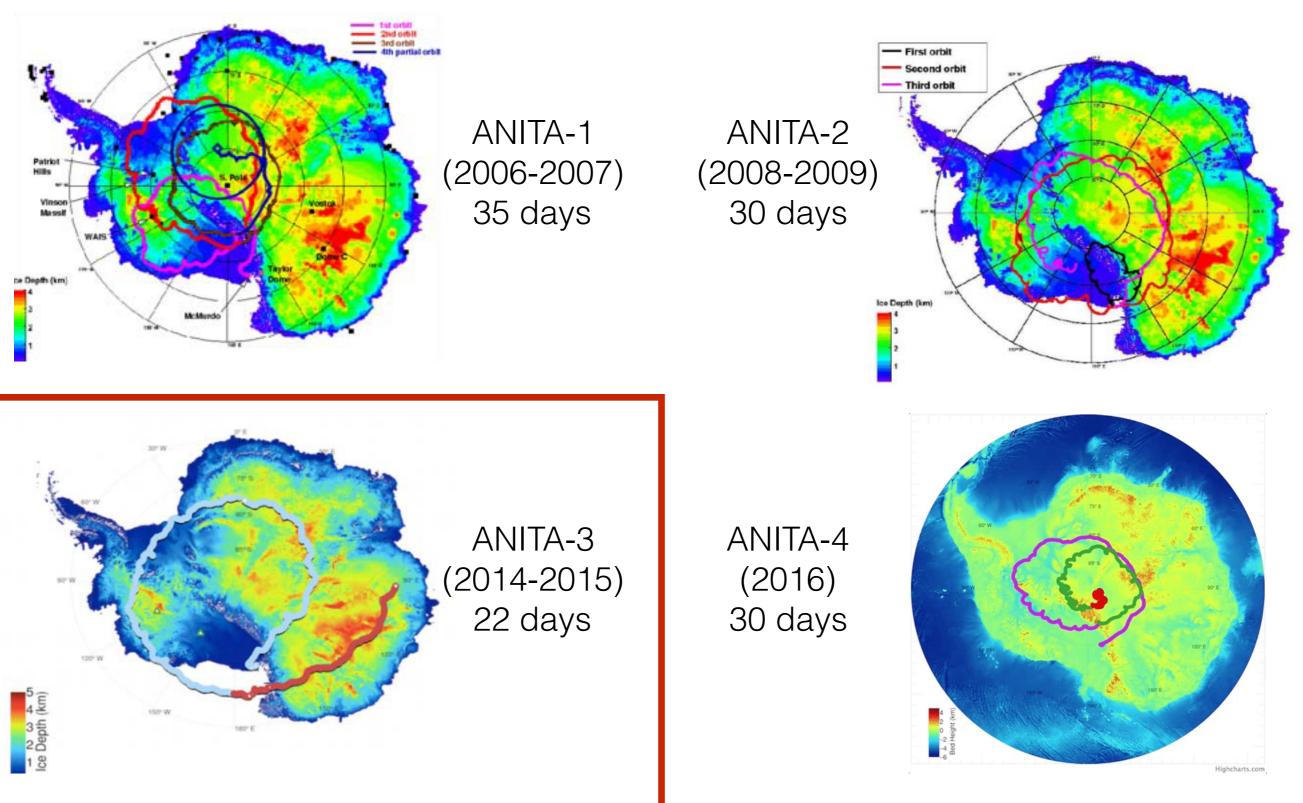








ANITA Flights



L. Cremonesi

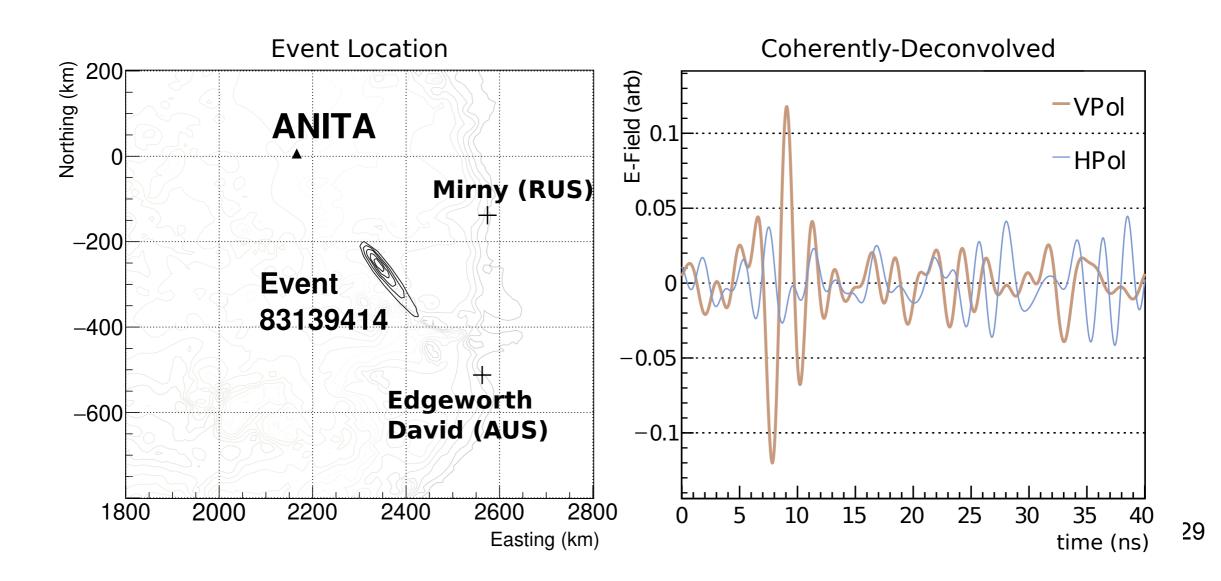
"UHE neutrinos and ANITA"

ANITA-3 Results

 New diffuse neutrinos search from ANITA-3 –arXiv 1803.02719

Constraints on the diffuse high-energy neutrino flux from the third flight of ANITA

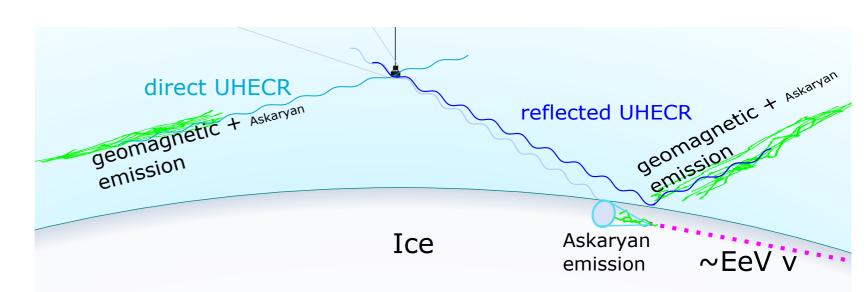
P. Allison, O. Banerjee, L. Batten, J. J. Beatty, K. Bechtol, K. Belov, D. Z. Besson, W. R. Binns, V. Bugaev, P. Cao, C. C. Chen, C. H. Chen, P. Chen, J. M. Clem, A. Connolly, L. Cremonesi, B. Dailey, C. Deaconu, P. F. Dowkontt, B. D. Fox, J. W. H. Gordon, P. W. Gorham, C. Hast, B. Hill, S. Y. Hsu, J. J. Huang, K. Hughes, R. Hupe, M. H. Israel, K. M. Liewer, T. C. Liu, A. B. Ludwig, L. Macchiarulo, S. Matsuno, C. Miki, K. Mulrey, J. Nam, C. Naudet, R. J. Nichol, A. Novikov, E. Oberla, S. Prohira, B. F. Rauch, J. M. Roberts, A. Romero-Wolf, B. Rotter, J. W. Russell, D. Saltzberg, D. Seckel, H. Schoorlemmer, J. Shiao, S. Stafford, J. Stockham, M. Stockham, B. Strutt, M. S. Sutherland, G. S. Varner, A. G. Vieregg, S. H. Wang, S. A. Wissel

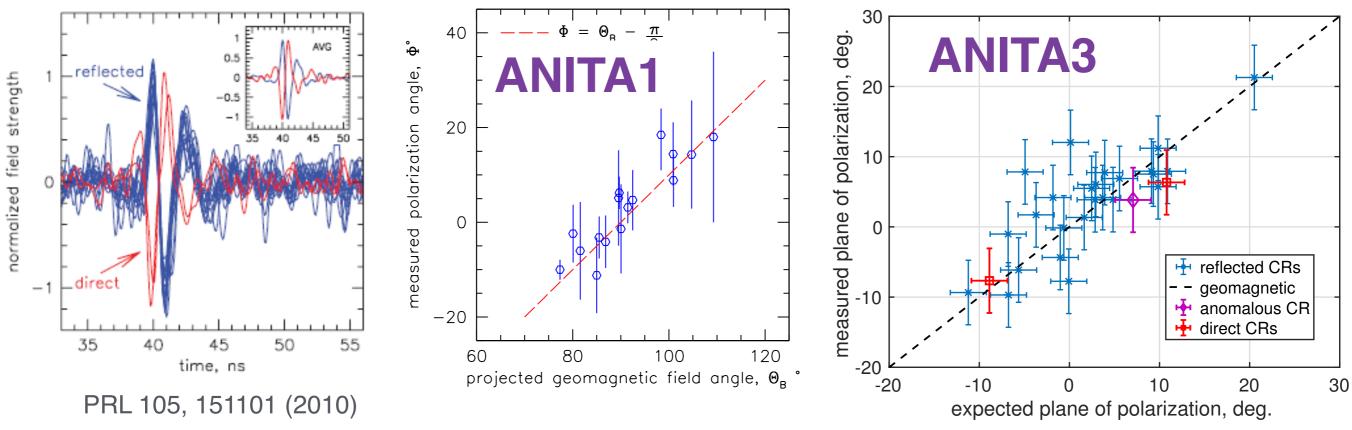


UHECR

ANITA1: 16 UHECR 14 reflected + 2 direct ANITA-2: 2 UHECR H-pol trigger was off ANITA-3: 25 UHECR ANITA-4: analysis in progress

L. Cremonesi



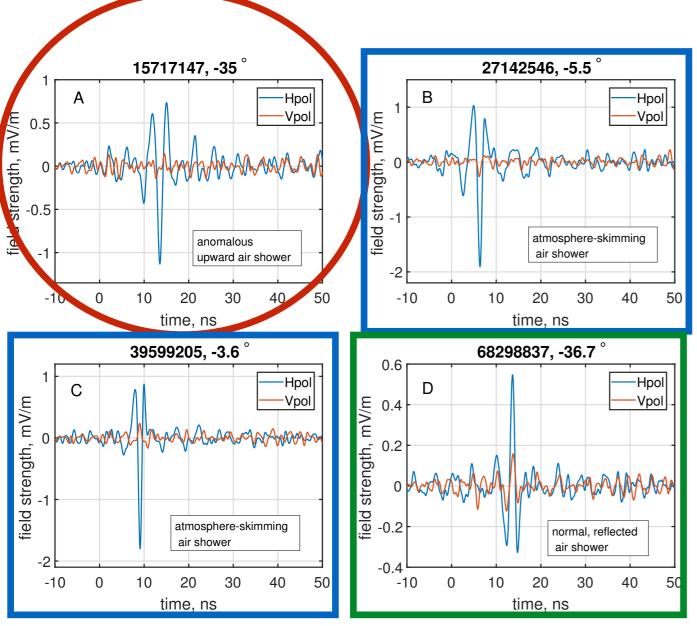


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arXiv:1803.05088 [astro-ph.HE]

"UHE neutrinos and ANITA"

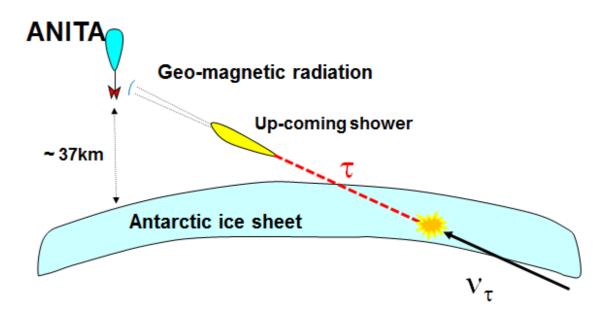
And ANITA-3 mystery event



Chord length: 5500-7000 km (20-30,000km water equivalent) 1600km SM interaction length @ 1 EeV

Background estimate < 10⁻²

L. Cremonesi



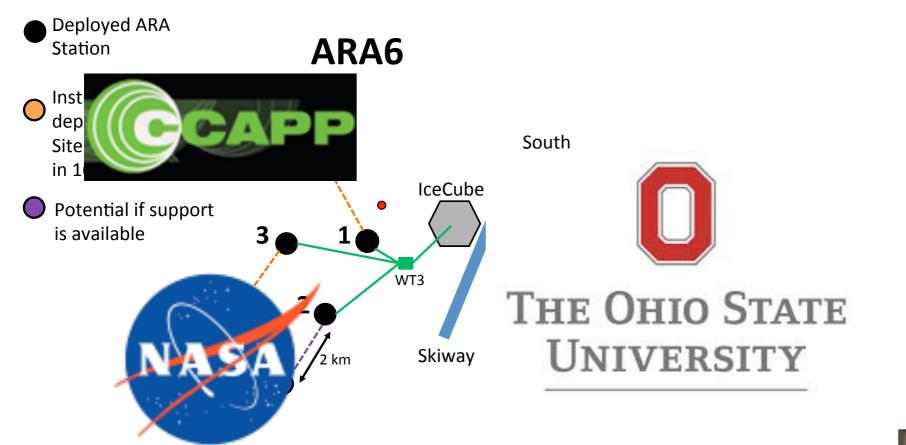
Direct Cosmic Rays

Reflected Cosmic Rays

"UHE neutrinos and ANITA"

Future Radio



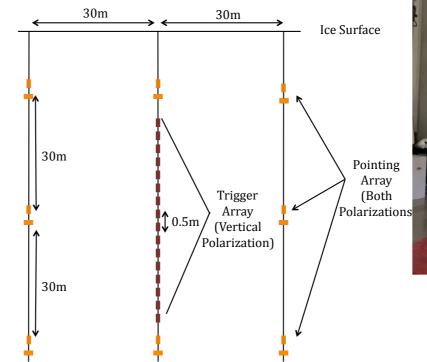


EVA



Greenland Neutrino Observatory







Questions?

- What consensus / conflicts (on what should be done in longer term European HEP) are there in this area?
- What are the experimental possibilities? Are different scenarios already envisaged?
 - IceCube is a tremendous success, the science case for IceCube-Gen2 (both the high and low energy extensions) are clear
 - -Some version of KM3NeT will exist
 - -Small experiments (i.e. ANITA) have discovery potential
 - -Hard to disentangle politics
- What are the choices for the strategy? What can the UK agree to input?

-Astroparticle physics should be mentioned

• What are the potential developments in this field? How do they relate to fundamental physics questions? ³³