#### A Photon Collider at TESLA

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- Luminosity optimisation
- Detector and backgrounds
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#### Introduction

#### Physics Motivation:

Want to study  $\gamma \gamma \to H \to b\bar{b}$ ,  $\gamma \gamma \to W^+W^-$ ,  $\gamma \gamma \to SUSY \to need \sqrt{s} = 120 GeV - maximum possible with high luminosity.$ 

TESLA bunch structure: bunch trains with 2800 bunches/train and 337 ns bunch crossing time

laser completely driven by time structure, study only partially valid for warm technology

#### Detector design:

- Large disruption angle requires crab crossing with  $\alpha \approx 35 \text{mrad}$
- forward part of detector completely driven by laser and crossing angle
- outer part kept identical to e<sup>+</sup>e<sup>-</sup> TDR-detector

# Beam parameters for $\sqrt{s_{\rm ee}} = 500 \, {\rm GeV}$

	e <sup>+</sup> e <sup>-</sup>	$\gamma\gamma$	$\gamma\gamma$
			(optimistic)
$N/10^{10}$	2	2	2
$\sigma_z \; [\mathrm{mm}]$	0.3	0.3	0.3
pulses/train	2820	2820	2820
Repetition rate [Hz]	5	5	5
$\gamma \epsilon_{x/y} / 10^{-6} \text{ [m·rad]}$	10./0.03	3./0.03	2.5/0.03
$\beta_{x/y}$ [mm] at IP	15/0.4	4/0.4	1.5/0.3
$\sigma_{x/y}$ [nm]	553/5	157/5	88/4.3
$\mathcal{L}(z > 0.8z_m)$	3.4	0.6	1.1
$[10^{34} \text{cm}^{-2} \text{s}^{-1}]$			

#### The Laser

Wavelength of powerful solid state lasers is in the 1 $\mu$ m range, e.g. Nd:YAG  $\lambda = 1.06 \mu$ m

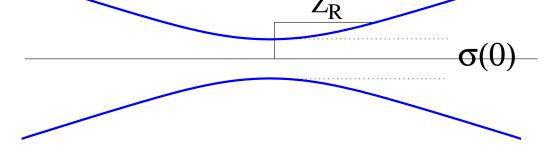
$$(x = 4.5 \text{ for } \sqrt{s} = 500 \,\text{GeV})$$

(If really needed can double or triple frequency)

Laser focusing in diffraction limited region:

$$\sigma_{L,r}(z) = \sigma_{L,r}(0)\sqrt{1+z^2/Z_R^2} \quad \sigma_{L,r}(0) = \sqrt{\frac{\lambda Z_R}{2\pi}}$$

 $Z_R$ : Rayleigh length



 $\rightarrow$  cannot vary length and diameter of laser spot simultaneously

Optimum around  $Z_R \approx \sigma_z \implies$  half opening angle of  $\mathcal{O}(1^\circ)$ 

Fraction of converted electrons:

$$k = N_{\gamma}/N_e \approx 1 - \exp(-A/A_0)$$

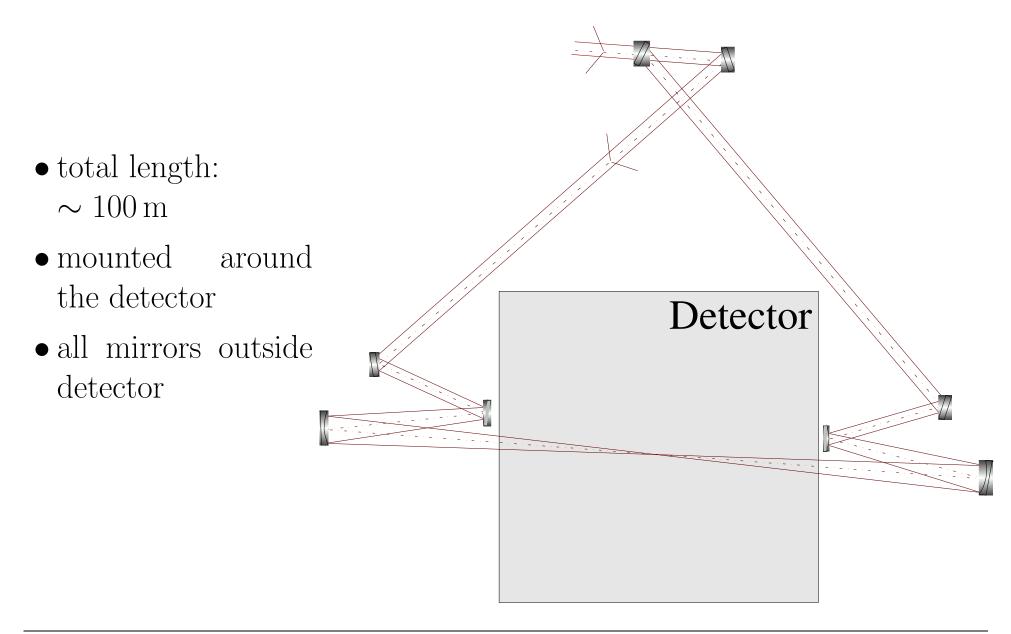
A: pulse energy of laser

For  $Z_R \approx \sigma_z$  and head on laser-beam collisions:

$$A_0 \approx \frac{\pi \hbar c \sigma_z}{\sigma_c} \approx 1.5 \text{J}$$

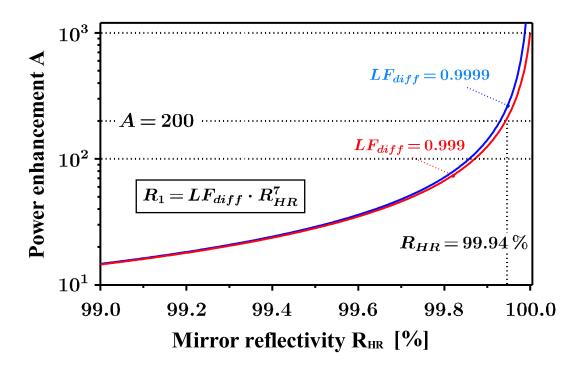
- $\Rightarrow$  need  $A \approx 2J$  (corresponds to  $\xi^2 \approx 0.2$ ) (for head on  $e^-$ -laser collisions)
- $\Rightarrow$  total laser power of  $\sim 2 \times 30 \,\mathrm{kW}$  needed
- $\sim$  60 Mercury lasers from the Livermore fusion program

# TESLA solution: recycle photons in resonant ring cavity:



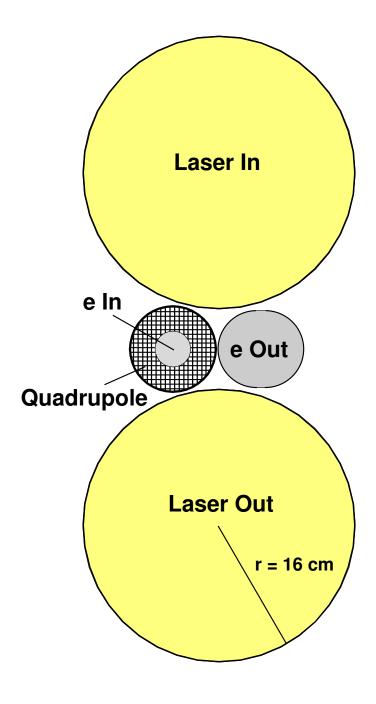
# Principles of a cavity

- $\bullet$  cavity with N mirrors with reflectivity  $R_i$
- loss per round trip  $V = R_2 \cdot R_3 \dots \cdot R_N \cdot L$  (L = other losses)
- power enhancement of cavity  $A = \frac{1 R_1}{(1 \sqrt{R_1 V})^2}$  ( $R_1$  =coupling mirror)
- maximal for  $R_1 = V$



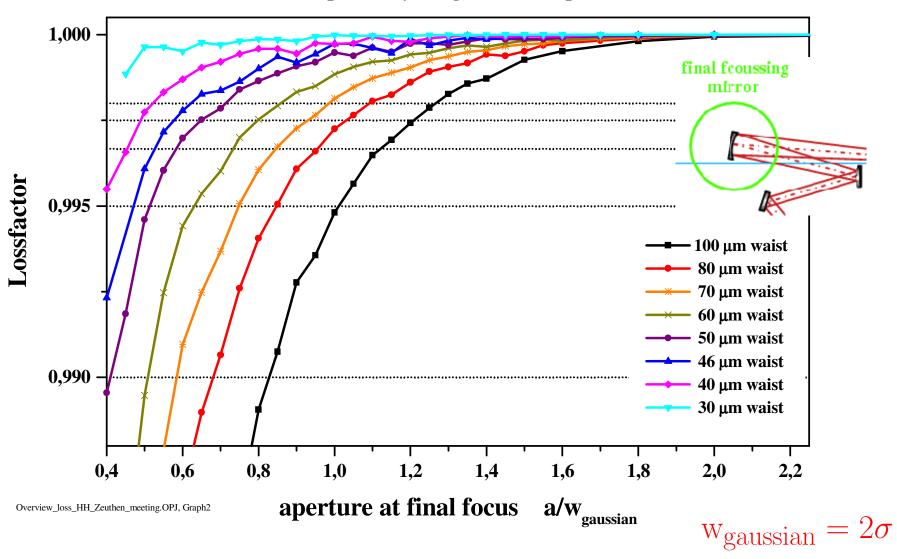
Power enhancement > 100 possible for realistic reflectivities

- To have highly efficient mirrors need crossing angle beam-laser
- crossing angle results in smaller conversion probability
- laser divergence and therefore mirror size depends on Rayleigh length
- finite mirrors result in diffraction losses and broadening of the focus
- have to find optimum crossing angle/Rayleigh length
- $\Rightarrow$  even higher laser power needed

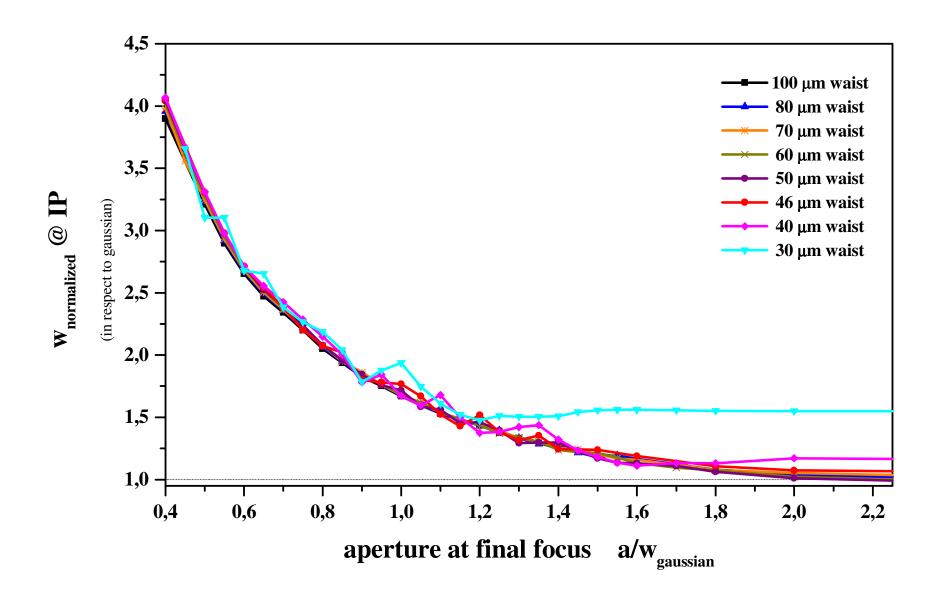


#### Diffraction losses are small even for small mirrors

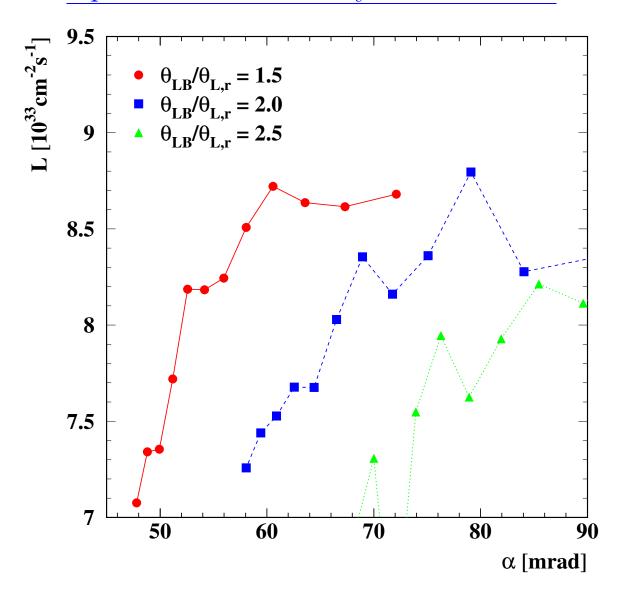
#### telescopic cavity, magnification sqrt(3)



#### However diffraction broadening is serious



# Optimum for relatively small mirrors



#### Optimum parameters

Laser parameters	TDR PT. VI	This study
Rayleigh length $Z_R$	0.35  mm	0.63 mm
Collision angle $\alpha_0$		55.1 mrad
Laser energy $A$	5 J	9.0 J
pulse duration $\sigma_{L,z}$	1.5 ps	1.5 ps
nonlinearity parameter $\xi^2$	0.30	0.30
Total Luminosity $[10^{34} \text{cm}^{-2} \text{s}^{-1}]$	1.10	1.05

# TDR parameters can be reproduced

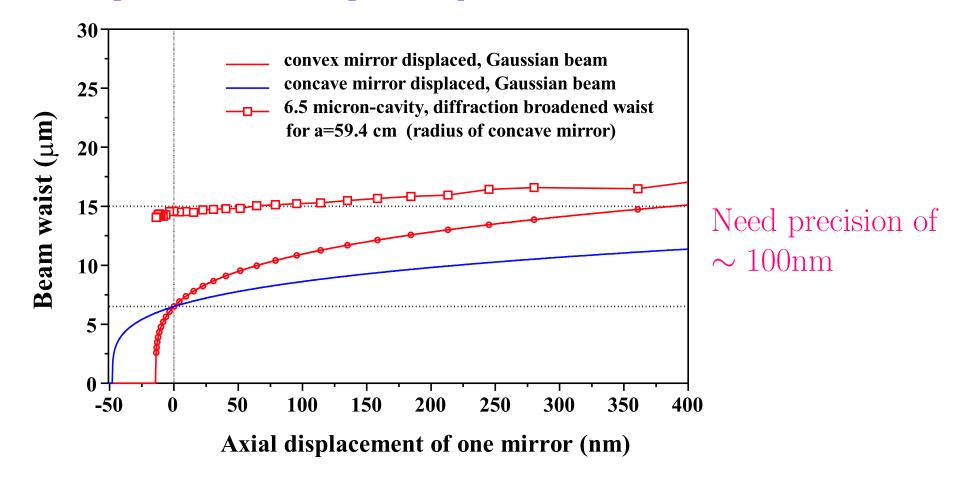
#### However larger laser pulse-power needed

(Collision angle chosen as compromise between luminosity and dead region at low angles)

# Alignment tolerances

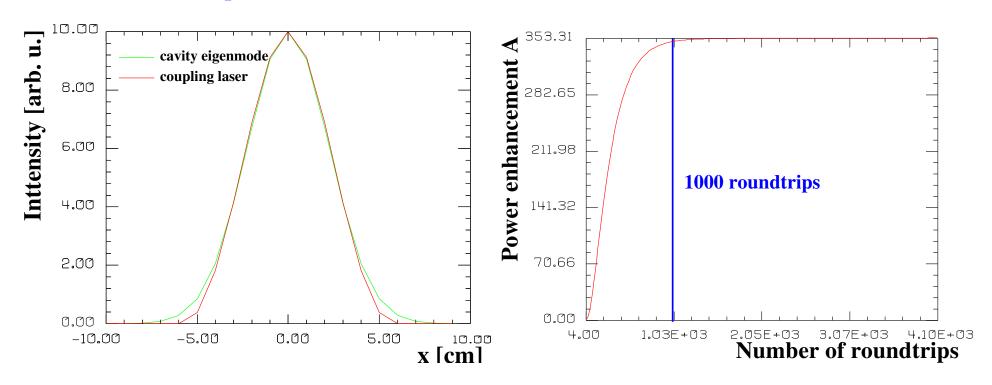
Total length of cavity:  $\Delta L \sim 0.3 \text{nm}$ 

Correction procedure understood e.g. from gravitational wave antennas Misalignment of focusing telescope:

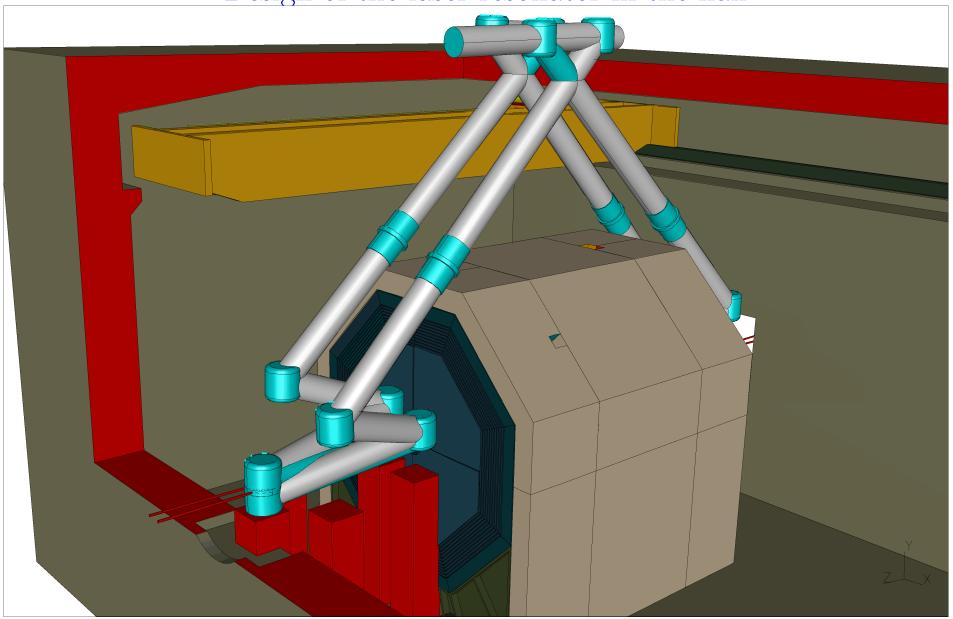


# Filling of the cavity

- Eigenmode in cavity is non-Gaussian due to diffraction broadening
- However filling of cavity with Gaussian mode works well
- need  $\sim 1000$  pulses for A = 350



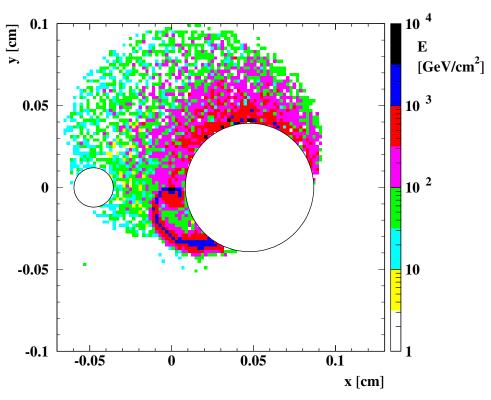
Design of the laser resonator in the hall



# Detector and backgrounds

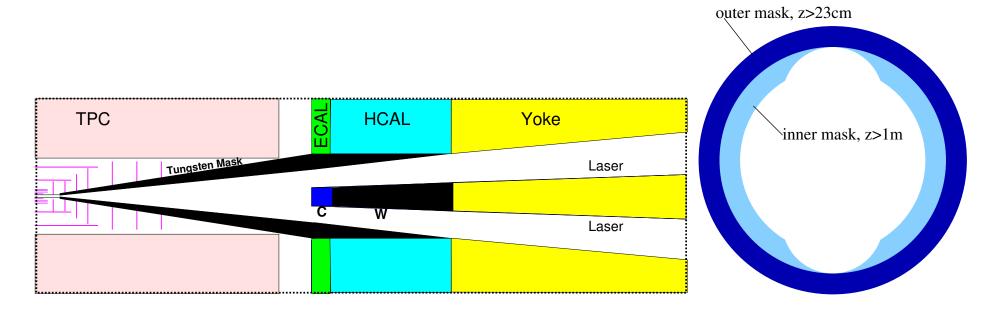
#### Background in the detector driven by

- large disruption angle
- angle between outgoing beam and B-field

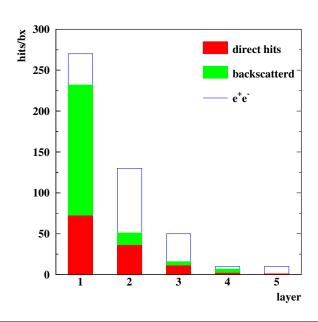


- direct background from pair production smaller than in e<sup>+</sup>e<sup>-</sup> due to anti-pinch effect
- large potential background from backscattering at detector exit

#### Background can be suppressed by masks and choice of material



- Backgrounds are similar to e<sup>+</sup>e<sup>-</sup> and should thus be manageable
- However detector is dead for  $\theta < 7.5^{\circ}$

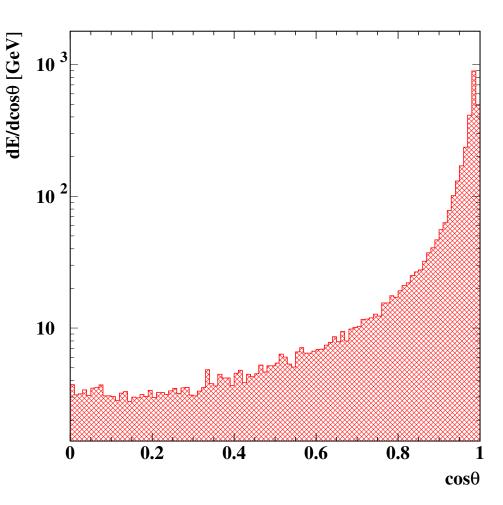


# Low energy qq background

- Large luminosity and large cross section  $\gamma \gamma \to q\bar{q}$  at low  $\sqrt{s}$
- $\mathcal{O}(1)$  event/bx overlaid to physics events (pileup)

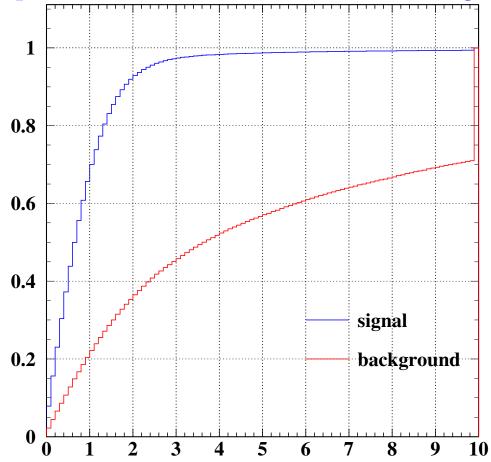
Due to large boost pileup tracks are forward peaked
 Can be largely.

• Can be largely rejected if physics in not forward peaked (like  $\gamma \gamma \to W^+W^-$ )



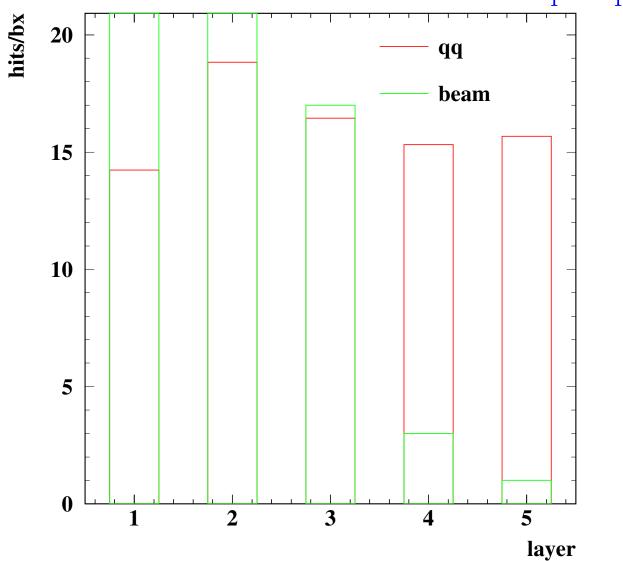
- Additional help/complication: beamspot length  $\sim 200 \mu \text{m}$
- $\Rightarrow$  signal and pileup separated in z
  - microvertex detector can help to separate
  - -can screw up b-tagging, e.g. in Higgs analysis

Integrated Impact Parameter distribution for signal and pileup

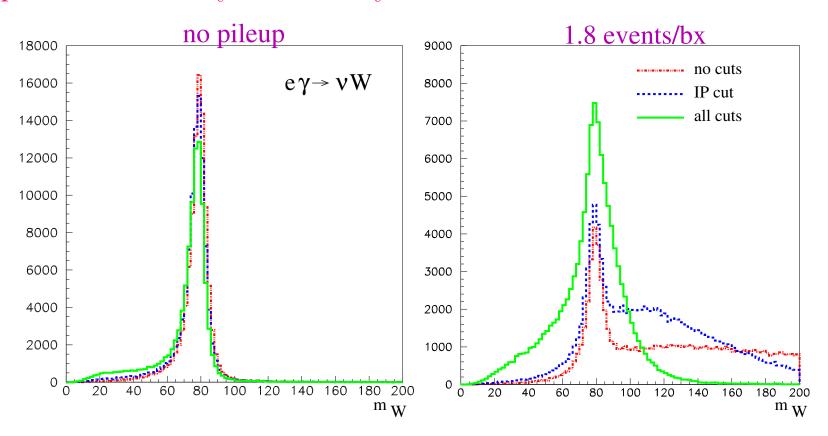


# Pileup gives also non negligible background in detector

Hits in vertex detector from beam and pileup



#### Pileup affects seriously some analyses

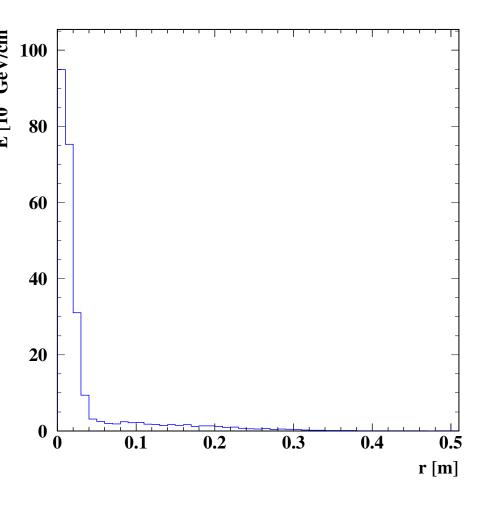


#### Conclusions pileup

- $\bullet$  Pileup is a serious issue at a  $\gamma\gamma$ -collider
- Very good time stamping is a must (no problem at TESLA)
- the long bunches at TESLA help additionally

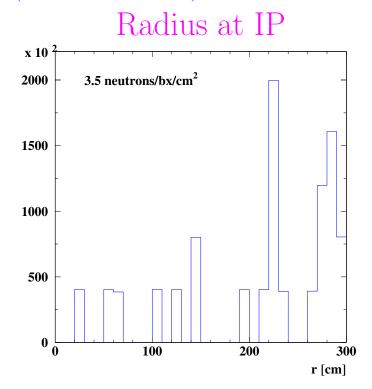
# Some thoughts about the dump

- photons need straight line to the dump neutrons
- assume electrons are swept to another dump
- radial distribution of photons
  - most contained within  $r < 5 \,\mathrm{cm}$
  - -however tail of 15% up to  $r = 30 \,\mathrm{cm}$
- assume a plain water dump (at least up to 30 cm)
- complicated structures (like Valery's noses) only possible at larger radii



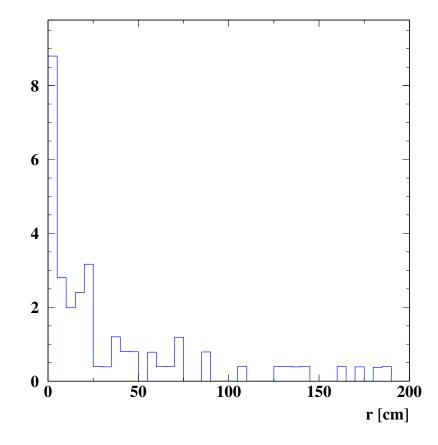
# Neutrons from the dump

- assume dump at  $d = 100 \,\mathrm{m}$
- calculate neutrons with Geant 4.6.2
- physics list QGSP\_HP, cross section bias 100
- cross checks with LHEP\_GN, QGSP\_GN, cross section bias 1
- results  $(E_{\nu} > 15 \text{ keV})$ :



# Kinetic energy 4 2 0 0 0.1 0.2 0.3 0.4 0.5 E<sub>kin</sub> [GeV]

- 3.5 neutrons/bx/cm<sup>2</sup> from  $\gamma$ s
- $\sim 5 \cdot 10^{11} \text{ neutrons/cm}^2/\text{a}$ 
  - similar amount from electrons, if sent to the same dump
  - most neutrons leave dump at small radius



Can reduce at most a factor two with Valery's nose

#### Conclusions

- No showstoppers found so far
- The laser-cavity seems difficult but possible
- Backgrounds are under control
- However the price to pay is a dead detector below 7.5°
- Neutrons may be a problem for the vertex detector
- If you want the photon collider to become a reality you have to work on the technical issues