

# Supercharging Superstrata

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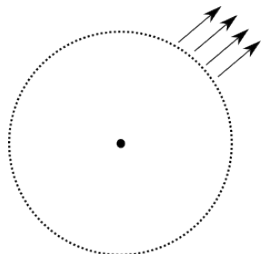
# Problems on the Horizon

[Bekenstein, Hawking]

- Information Paradox
  - The spectrum of emitted particles is thermal
- Entropy problem
  - Assign entropy to an object with a Horizon

$$S_{BH} = \frac{c^3 A_H}{4G_N \hbar}$$

- Microscopical description  $\leftrightarrow$  No-hair theorems

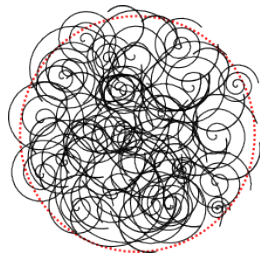


# Fuzzball Proposal

[Mathur, Lunin,...]

At the scale of the horizon, a new fuzz state appears.

- Black hole has  $e^{S_{BH}}$  microstates.
- Microstates constructed from String theory
- Some allow supergravity description  
→ **Microstate Geometries**



*Smooth, horizonless solutions of supergravity, valid within the supergravity approximation of string theory and have the same mass, angular momentum and charges as a given black hole*

# D1-D5-P Black Hole

- Supersymmetric 3-charge Black hole solution in on  $\mathbb{R}_t \times \mathbb{R}^4 \times S_y^1 \times T^4$
- Bekenstein–Hawking entropy

$$S_{BH} = \frac{2\pi^2(2\pi R_y)((2\pi)^4 V)\sqrt{Q_1 Q_5 Q_p}}{4G_{10}}$$

Consider the string theory setup

Object	Number	Charge	$\mathbb{R}_t$	$\mathbb{R}^4$	$S_y^1$	$T^4$
D1	$n_1$	$Q_1$	—	•	—	•
D5	$n_5$	$Q_5$	—	•	—	—
P	$n_p$	$Q_p$	—	•	—	•

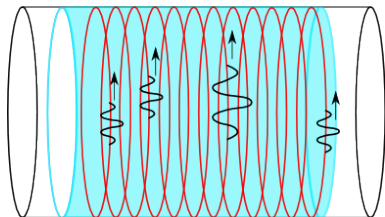
Use relations between  $Q$  and  $n$  to get

$$S_{BH} = 2\pi\sqrt{n_1 n_5 n_p}$$

# Entropy Matching

[Strominger, Vafa, Maldacena,...]

- Number of configurations  
= Partitions of  $n_1 n_5 n_p \rightarrow e^{2\pi\sqrt{n_1 n_5 n_p}}$
- Entropy  $S_{micro} = 2\pi\sqrt{n_1 n_5 n_p}$ ,



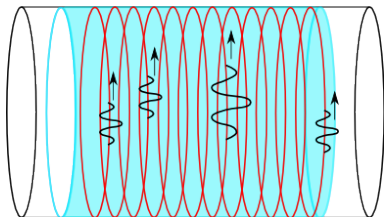
*Entropy calculations match in the two different regimes*

- What can we say about each individual configuration?

# Entropy Matching

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*Entropy calculations match in the two different regimes*

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Find new microstate geometries of the D1-D5-P black hole using the superstratum technology.

# Microstate Geometries

[Giusto, Martucci, Petrini, Russo]

- Supersymmetry constraints the form of the physical fields
- Use  $u = (t - y)/\sqrt{2}$ ,  $v = (t + y)/\sqrt{2}$ .

$$ds_{10}^2 = -2 \frac{\sqrt{Z_1 Z_2}}{\mathcal{P}} (dv + \beta) \left[ du + \omega + \frac{\mathcal{F}}{2} (dv + \beta) \right] + \sqrt{Z_1 Z_2} ds_4^2 + \sqrt{\frac{Z_1}{Z_2}} d\hat{s}_4^2,$$

$$C_0 = \frac{Z_4}{Z_1} \quad B = -\frac{Z_4}{\mathcal{P}} (du + \omega) \wedge (dv + \beta) + \mathbf{a}_4 \wedge (dv + \beta) + \delta_2,$$

where  $Z_1$ ,  $Z_2$ ,  $Z_4$ ,  $\mathcal{F}$  are scalars,  $\beta$ ,  $\omega$ ,  $\mathbf{a}_4$  one-forms, and  $\delta_2$  two-forms on  $\mathbb{R}^4$  with  $\mathcal{P} = Z_1 Z_2 - Z_4^2$ .

- Use  $\mathcal{D} \equiv d_4 - \beta \wedge \partial_v$ , and combine content into new two-forms, such as ( $\dot{\cdot} \equiv \partial_v$ )

$$\Theta_4 \equiv \mathcal{D} \mathbf{a}_4 + \dot{\delta}_2.$$



# BPS Equations

- Zeroth Layer:
  - The base space  $\mathbb{R}^4$  has a flat metric  $ds_4^2$ ,
  - $\beta$  is self-dual:  $d\beta = *_4 d\beta$ .
- First Layer: System of **linear** differential equations for pairs  $(Z_1, \Theta_2)$ ,  $(Z_2, \Theta_1)$ , and  $(Z_4, \Theta_4)$ , such as

$$*_4 \mathcal{D} \dot{Z}_4 = \mathcal{D} \Theta_4, \quad \mathcal{D} *_4 \mathcal{D} Z_4 = -\Theta_4 \wedge d\beta, \quad \Theta_4 = *_4 \Theta_4.$$

- Second Layer: Linear equations for  $\mathcal{F}$  and  $\omega$ .

$$\begin{aligned} \mathcal{D}\omega + *_4 \mathcal{D}\omega + \mathcal{F}d\beta &= Z_1 \Theta_1 + Z_2 \Theta_2 - 2Z_4 \Theta_4, \\ *_4 \mathcal{D} *_4 \left( \dot{\omega} - \frac{1}{2} \mathcal{D}\mathcal{F} \right) &= \partial_v^2 (Z_1 Z_2 - Z_4^2) - (\dot{Z}_1 \dot{Z}_2 \\ &\quad - (\dot{Z}_4)^2) - \frac{1}{2} *_4 (\Theta_1 \wedge \Theta_2 - \Theta_4 \wedge \Theta_4). \end{aligned}$$

# Seed Solution

[Lunin, Maldacena, Maoz, Kanitscheider, Skenderis, Taylor;  
Lunin, Mathur, Saxena]

- Two charge supertube solution with a perturbation:

$$Z_1 = \frac{Q_1}{\Sigma}, \quad Z_2 = \frac{Q_5}{\Sigma}, \quad Z_4 = R_y b a^k \frac{\sin^k \theta e^{-ik\phi}}{(r^2 + a^2)^{k/2} \Sigma} \quad \mathcal{F} = \Theta_4 = 0,$$

with  $\Sigma \equiv r^2 + a^2 \cos^2 \theta$ .

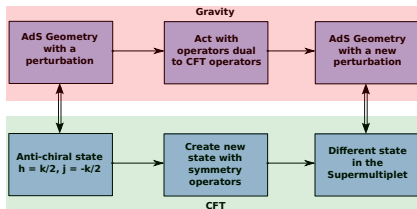
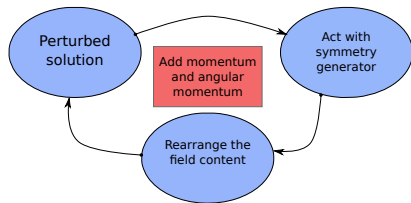
Asymptotically  $\text{AdS}_3 \times S^3 \times T^4 \Rightarrow$  use AdS/CFT duality:

- States of the CFT  $\Leftrightarrow$  Geometries in the bulk:  
Supertube with perturbation dual to an anti-chiral primary  $|k/2, -k/2\rangle$
- Symmetries of the Geometry  $\Leftrightarrow$  Symmetries of the CFT
  - $\text{AdS}_3$  symmetry  $\Leftrightarrow$  Conformal symmetry  
 $SO(2, 2) \simeq SL(2, \mathbb{R})_L \times SL(2, \mathbb{R})_R$
  - $S^3$  symmetry  $\Leftrightarrow$   $\mathcal{R}$ -symmetry  $SO(4) \simeq SU(2)_L \times SU(2)_L$

# Solution Generating technique

[Bena, Giusto, Martinec-Russo, Shigemori, Turton, Warner]

- Use symmetry generators  $L_{-1}$  of  $SL(2, \mathbb{R})_L$  and  $J_0^+$  of  $SU(2)_L$  to add momentum



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$$(J_0^+)^m (L_{-1})^n |k/2, -k/2\rangle \Leftrightarrow$$

$$Z_4 = b R_y \frac{\Delta_{k,m,n}}{\Sigma} \cos \hat{v}_{k,m,n},$$

$$\Theta_4 = -\sqrt{2} b \Delta_{k,m,n} \left[ \left( (m+n) r \sin \theta + n \left( \frac{m}{k} - 1 \right) \frac{\Sigma}{r \sin \theta} \right) \Omega^{(1)} \sin \hat{v}_{k,m,n} \right. \\ \left. + \left( m \left( \frac{n}{k} + 1 \right) \Omega^{(2)} + \left( \frac{m}{k} - 1 \right) n \Omega^{(3)} \right) \cos \hat{v}_{k,m,n} \right]$$

$$\Delta_{k,m,n} \equiv \left( \frac{a}{\sqrt{r^2 + a^2}} \right)^k \left( \frac{r}{\sqrt{r^2 + a^2}} \right)^n \cos^m \theta \sin^{k-m} \theta,$$

$$\hat{v}_{k,m,n} \equiv (m+n) \frac{\sqrt{2} v}{R_y} + (k-m) \phi - m \psi.$$

**What about supersymmetry?**

# Fermionic Solution Generating Technique

- CFT generator  $G_{\pm\frac{1}{2}}^{\alpha A}$  is dual to variations generated by Killing spinors of  $\text{AdS}_3 \times S^3 \times T^4$  (seed solution with  $b = 0$ )
- Procedure:
  - Solve the Killing spinor equations for empty  $\text{AdS}_3 \times S^3 \times T^4$  to find explicit form of these spinors
  - Identify the complex component  $\zeta_{\pm}^{\alpha A}$  corresponding to  $G_{\pm\frac{1}{2}}^{\alpha A}$
  - Act on the seed ( $b \neq 0$ ) to obtain a new solution using

$$\delta' \delta C^{(0)} = \frac{1}{2} e^{-\phi} \epsilon^T \Gamma^0 (i\sigma^2) \delta' \lambda,$$

$$\delta' \delta B_{\mu\nu} = 2\bar{\epsilon} \Gamma_{[\mu} \sigma^3 \delta' \psi_{\nu]}.$$

# Linear order in Perturbation

- The geometry dual to state  $\left(G_{-\frac{1}{2}}^{+1} G_{-\frac{1}{2}}^{+2} + \frac{1}{k} J_0^+ L_{-1}\right) |k/2, -k/2\rangle$  is

$$\begin{aligned} Z_4 &= 0, \\ \Theta_4 &= -\sqrt{2} b \Delta_{k,1,1} \left[ \frac{\Sigma}{r \sin \theta} \Omega^{(1)} \sin \hat{\nu}_{k,1,1} + \left( \Omega^{(2)} + \Omega^{(3)} \right) \cos \hat{\nu}_{k,1,1} \right], \end{aligned}$$

- Use this as a new seed to obtain geometry dual to  $(J_0^+)^m (L_{-1})^n \left(G_{-\frac{1}{2}}^{+1} G_{-\frac{1}{2}}^{+2} + \frac{1}{k} J_0^+ L_{-1}\right) |k/2, -k/2\rangle$

$$\begin{aligned} Z_4 &= 0, \\ \Theta_4 &= -\sqrt{2} b \Delta_{k,1+m,1+n} \\ &\quad \times \left[ \frac{\Sigma}{r \sin \theta} \Omega^{(1)} \sin \hat{\nu}_{k,1+m,1+n} + \left( \Omega^{(2)} + \Omega^{(3)} \right) \cos \hat{\nu}_{k,1+m,1+n} \right]. \end{aligned}$$

# Finite Deformation of $\text{AdS}_3 \times S^3 \times T^4$

- Make  $b$  finite  $\rightarrow$  turn on  $\mathcal{F}$  and  $\omega$ . For example for  $(k, m, n) = (2, 0, 0)$ :

$$\mathcal{F}_{2,0,0} = -\frac{b^2}{18(a^2 + r^2)^3} (8a^4 + 15a^2r^2 + 6r^4 + a^2(2a^2 + r^2)\cos 2\theta)$$

$$\omega_{2,0,0} = \frac{b^2 R_y}{18\sqrt{2}(a^2 + r^2)^2 \Sigma} ((4a^2 + 3r^2)\cos^2 \theta d\psi + (6a^4 + 8a^2r^2 + 3r^4)\sin^2 \theta d\phi) .$$

## Properties of these solutions

- Asymptotically flat extension is trivial
- Conserved charges match the CFT result

# Summary

- Found the explicit form of Killing spinors in  $\text{AdS}_3 \times S^3 \times T^4$
- New family of solutions with  $Z_4 = 0$  and a simpler  $\Theta_4$
- Each solution has a well defined CFT dual
- In agreement with spectrum analysis [Deger et al,...]
- The solutions are simpler ( $\nu$ -independent even in the Asymptotically flat case)