Durham seminar, 4th October 2018

Implications of Decoupling New Physics

Tevong You



The Branco Weiss Fellowship Society in Science Durham seminar, 4th October 2018

SMEXIT Implications of decoupling new physics



Introduction

- **Soft** exit from the SM: New physics around the corner
 - Usual *low-scale SUSY/compositeness/extra-dimensions*... just a bit more fine-tuned
 - Neutral naturalness/Twin Higgs... hidden naturalising sector
- Hard exit from the SM: New physics decoupled
 - Accept fine-tuning while *SUSY/compositeness/extra-dimensions* resolve other problems at heavier scales
 - Anthropic landscape, censorship-type approaches...
 - Cosmological relaxation, clockwork...
- Phenomenological framework: **SM EFT**

Outline

- Part I: SM EFT
 - A phenomenological framework for decoupled new physics
- Part 2: B Anomalies
 - Signs of a non-zero Wilson coefficient?
- Part 3: Cosmological Relaxation
 - A new approach to decoupling new physics without finetuning

- Part I: SM EFT
- Part 2: B Anomalies
- Part 3: Cosmological Relaxation

Why SM EFT?

Include only experimentally discovered degrees of freedom in our theory

The TeV Scale

What effective theory captures everything we know experimentally about weak interactions?

1933–1982 4-fermion interactions

$$\checkmark \qquad \sim G_{\rm F} E^2 \qquad \Rightarrow \Lambda \sim {\rm TeV}$$

1982–2011 SM without Higgs



2012-now SM + higher-dimension operators?

$$\Rightarrow \Lambda \lesssim M_{\rm P}?$$

Why SM EFT?

Take a step back: recall the situation before 2012

The TeV Scale

What effective theory captures everything we know experimentally about weak interactions?

1933–1982 4-fermion interactions

$$\frown \qquad \qquad \Rightarrow \Lambda \sim \text{TeV}$$

1982–2011 SM without Higgs $\begin{array}{c}
\overset{\sim}{\sim} & \overset{\sim}{\sim} & \overset{\sim}{\sim} & \frac{g^2 E^2}{m_W^2} \Rightarrow \Lambda \sim \text{TeV}
\end{array}$ 2012 now SM is bighter dimension ended

2012-now SM + higher-dimension operators?

$$\Rightarrow \Lambda \lesssim M_{\rm P}?$$

Beyond the Standard Model?

A priori many ways to break electroweak symmetry!



New scalars could also be something other than a Higgs

EFT for weak bosons

- **1980s-2012**: Discovery of weak bosons. Non-linear effective Lagrangian for spontaneously-broken global symmetry (*breaking mechanism unknown!*)
- Global symmetry-breaking pattern gives low-energy effective theory regardless of UV mechanism responsible for it

 $SU(2) \times SU(2) \rightarrow SU(2)_V \qquad (\rho \equiv M_W/M_Z \cos \theta_w \sim 1)$

$$\mathcal{L} = rac{v^2}{4} \mathrm{Tr} D_\mu \Sigma^\dagger D^\mu \Sigma - m_i \bar{\psi}_L^i \Sigma \psi_R^i + \mathrm{h.c.}$$

$$\Sigma = \exp\left(i\frac{\sigma^a\pi^a}{v}\right)$$

EFT for weak bosons + scalar

• 2012: Non-linear electroweak Lagrangian + general couplings to singlet scalar

$$\begin{split} \mathcal{L} &= \frac{v^2}{4} \mathrm{Tr} D_{\mu} \Sigma^{\dagger} D^{\mu} \Sigma \left(1 + 2 \frac{a}{v} \frac{h}{v} + \frac{b}{v^2} \frac{h^2}{v^2} + ... \right) - m_i \bar{\psi}_L^i \Sigma \left(1 + \frac{c}{v} \frac{h}{v} + ... \right) \psi_R^i + \mathrm{h.c.} \\ &+ \frac{1}{2} (\partial_{\mu} h)^2 + \frac{1}{2} m_h^2 h^2 + \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 + \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 + ... \quad , \end{split}$$

$$\Sigma &= \exp\left(i \frac{\sigma^a \pi^a}{v} \right)$$

Fit experimental data to couplings

Could have had very different coupling





Fit experimental data to couplings

Could have had very different coupling



July 2012 post-discovery J. Ellis and T.Y. [arXiv:1207.1693]

Fit experimental data to couplings

Could have had very different coupling



Moriond 2013 J. Ellis and T.Y. [arXiv:1303.1879]

Why SM EFT?

Assuming a SM Higgs and decoupled new physics at higher energies, the SM EFT is the next phenomenological framework

The TeV Scale

What effective theory captures everything we know experimentally about weak interactions?

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$$\frown \qquad \sim G_{\rm F} E^2 \qquad \Rightarrow \Lambda \sim {\rm TeV}$$

1982–2011 SM without Higgs



2012-now SM + higher-dimension operators? $\Rightarrow \Lambda \lesssim M_P$?

- New physics appear to be decoupled at higher energies
- Given particle content, write down *all* terms allowed by symmetries...



- ...Including **higher-dimensional** operators!
 - $\mathbf{\mathcal{L}}_{\mathrm{SM}}^{\mathrm{dim-6}} = \sum_{i} \frac{c_i}{\Lambda^2} \mathcal{O}_i$
- Generated by new physics at scale $\Lambda \gg v$

- Unique dim-5 Weinberg operator, violates lepton number, predicts neutrino masses
- Modulo flavour structure, there are 59 dim-6 (CP-even) operators in a non-redundant basis

Buchmuller and Wyler (Nucl. Phys. B 268 (1986) 621) Gradkowski et al [arXiv:1008.4884]

• ~19 operators relevant for EWPT, TGC, and Higgs physics

EWPTs	Higgs Physics	TGCs				
$\mathcal{O}_W = rac{ig}{2} \left(H^{\dagger} \sigma^a \stackrel{\leftrightarrow}{D^{\mu}} H \right) D^{\nu} W^a_{\mu\nu}$						
$\mathcal{O}_B = rac{ig'}{2} \left(H^\dagger D^{\downarrow} D^{\downarrow} \right)$	$\mathcal{O}_{3W} = g \frac{\epsilon_{abc}}{3!} W^a_\mu {}^\nu W^b_{\nu\rho} W^{c\rho\mu}$					
$\mathcal{O}_T = rac{1}{2} \left(H^\dagger \overleftrightarrow{D}_\mu H ight)^2$	$\mathcal{O}_{HW} = ig(D^{\mu}H$	$)^{\dagger}\sigma^{a}(D^{\nu}H)W^{a}_{\mu\nu}$				
$\mathcal{O}_{LL}^{(3)l} = (\bar{L}_L \sigma^a \gamma^\mu L_L) \left(\bar{L}_L \sigma^a \gamma_\mu L_L \right)$	$\mathcal{O}_{HB} = ig'(D^{\mu})$	$(D^{\nu}H)^{\dagger}(D^{\nu}H)B_{\mu\nu}$				
$\mathcal{O}_R^e = (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H)(\bar{e}_R \gamma^{\mu} e_R)$	$\mathcal{O}_g = g_s^2 H ^2 G^A_{\mu u} G^{A\mu u}$					
$\mathcal{O}_R^u = (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H)(\bar{u}_R \gamma^{\mu} u_R)$	$\mathcal{O}_{\gamma} = g^{\prime 2} H ^2 B_{\mu\nu} B^{\mu\nu}$					
$\mathcal{O}_R^d = (iH^\dagger \stackrel{\leftrightarrow}{D_\mu} H)(\bar{d}_R \gamma^\mu d_R)$	$\mathcal{O}_H = rac{1}{2} (\partial^\mu H ^2)^2$					
$\mathcal{O}_L^{(3)q} = (iH^{\dagger}\sigma^a D^{\leftrightarrow}_{\mu}H)(\bar{Q}_L \sigma^a \gamma^{\mu} Q_L)$	$\mathcal{O}_f = y_f H ^2 \bar{F}_L H^{(c)} f_R + \text{h.c.}$					
$\mathcal{O}_L^q = (iH^\dagger \overset{\leftrightarrow}{D_\mu} H)(\bar{Q}_L \gamma^\mu Q_L)$	$\mathcal{O}_6 = \lambda H ^6$					

In SILH basis (Giudice et al. hepph/0703164), adopted from Pomarol and Riva (1308.1426)

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Combinations of operators constrained in EWPT more easily set to zero in Higgs and TGCs

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$\mathcal{O}_B = \frac{ig'}{2} \left(H^{\dagger} D \right)$	$\left(^{\mu}H ight) \partial^{ u}B_{\mu u}$	$\mathcal{O}_{3W} = g \frac{\epsilon_{abc}}{3!} W^a_\mu {}^\nu W^b_{\nu\rho} W^{c\rho\mu}$	
$\mathcal{O}_T = \frac{1}{2} \left(H^{\dagger} \stackrel{\leftrightarrow}{D}_{\mu} H \right)^2$	$\mathcal{O}_{HW} = ig(D^{\mu}H$	$)^{\dagger}\sigma^{a}(D^{\nu}H)W^{a}_{\mu\nu}$	r
$\mathcal{O}_{LL}^{(3)l} = (\bar{L}_L \sigma^a \gamma^\mu L_L) \left(\bar{L}_L \sigma^a \gamma_\mu L_L \right)$	$\mathcal{O}_{HB} = ig'(D^{\mu})$	$H)^{\dagger}(D^{\nu}H)B_{\mu\nu}$	F
$\mathcal{O}_R^e = (iH^\dagger \overset{\leftrightarrow}{D_\mu} H)(\bar{e}_R \gamma^\mu e_R)$	$\mathcal{O}_g = g_s^2 H ^2 G^A_{\mu\nu} G^{A\mu\nu}$		١
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$\mathcal{O}_L^{(3)q} = (iH^{\dagger}\sigma^a \overset{\leftrightarrow}{D_{\mu}}H)(\bar{Q}_L\sigma^a\gamma^{\mu}Q_L)$	$\mathcal{O}_f = y_f H ^2 \bar{F}_L H^{(c)} f_R + \text{h.c.}$		ph,
$\mathcal{O}_L^q = (iH^\dagger \overset{\leftrightarrow}{D_\mu} H)(\bar{Q}_L \gamma^\mu Q_L)$	$\mathcal{O}_6 = \lambda H ^6$		POI

Operators constrained by measurements at per cent level or worse

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$\mathcal{O}_T = \frac{1}{2} \left(H^{\dagger} \stackrel{\leftrightarrow}{D}_{\mu} H \right)^2$	$\mathcal{O}_{HW} = ig(D^{\mu}H)$	$)^{\dagger}\sigma^{a}(D^{\nu}H)W^{a}_{\mu u}$				
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Operators benefit from per mille precision at LEP

In SILH basis (Giudice et al. hepph/0703164), adopted from Pomarol and Riva (1308.1426)

LEP EWPT Example

• (Pseudo-)Observables

$$T_{2}^{*} = T_{had} + 3T_{2}^{*} + 3T_{2}^{*} \quad R_{\ell} = \frac{T_{had}}{T_{\ell}} \quad \mathcal{O}_{had} = 12\pi \frac{\Gamma_{e} T_{had}}{\mathcal{O}_{2}^{*}} \quad \mathcal{A}_{FB}^{\dagger} = \frac{3}{4} \mathcal{A}_{e} \mathcal{A}_{f} \quad M_{W} = c_{W} M_{2}$$

$$R_{q} = \frac{\Gamma_{q}}{T_{had}}$$

• Depends on

$$\Gamma_{L}^{L} = \frac{52}{52} \frac{G_{F}}{G_{F}} \frac{M_{E}^{2} M_{E}}{G_{R}} \left[(g_{L}^{f})^{2} + (g_{R}^{f})^{2} \right] \qquad A_{f} = \frac{(g_{L}^{f})^{2} - (g_{R}^{f})^{2}}{(g_{L}^{f})^{2} + (g_{R}^{f})^{2}} \qquad B_{f}^{f} = \frac{(g_{L}^{f})^{2} - (g_{R}^{f})^{2}}{(g_{L}^{f})^{2} + (g_{R}^{f})^{2}}$$

 Dim-6 operators can modify observables directly through Zff couplings contributions or indirectly through redefinitions of input observables

 $m_{t}^{2} = (m_{z}^{2})^{\circ} (1 + \pi_{t}) \qquad G_{f} = G_{f}^{\circ} (1 - \pi_{uw}^{\circ}) \qquad \propto (m_{t}) = \alpha^{\circ}(m_{z}) (1 + \pi_{y})$

LEP EWPT Example

• Individual (green) and marginalised (red) 95% CL limits



Ellis, Sanz and T.Y. 1410.7703

• 8 (combinations of) operators probed by EWPT

- Assume SM Z and W couplings to fermions in diboson measurements
- Interpret in anomalous TGC framework:



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Updated Global SMEFT Fit

J. Ellis, C. Murphy, V. Sanz and TY, 1803.03252

- Combine EWPT, diboson, Higgs data
- Fit to 20 dim-6 CP-even operators simultaneously
- Present results in Warsaw and SILH basis
- Match to **simplified models**

Updated Global SMEFT Fit

• SILH basis

$$\begin{aligned} \mathcal{L}_{\text{SMEFT}}^{\text{SILH}} &\supset \frac{\bar{c}_{W}}{m_{W}^{2}} \frac{ig}{2} \left(H^{\dagger} \sigma^{a} \vec{D}^{\mu} H \right) D^{\nu} W_{\mu\nu}^{a} + \frac{\bar{c}_{B}}{m_{W}^{2}} \frac{ig'}{2} \left(H^{\dagger} \vec{D}^{\mu} H \right) \partial^{\nu} B_{\mu\nu} + \frac{\bar{c}_{T}}{v^{2}} \frac{1}{2} \left(H^{\dagger} \vec{D}_{\mu} H \right)^{2} \\ &+ \frac{\bar{c}_{ll}}{v^{2}} 2(\bar{L}\gamma_{\mu}L)(\bar{L}\gamma^{\mu}L) + \frac{\bar{c}_{He}}{v^{2}} (iH^{\dagger} \vec{D}_{\mu}H)(\bar{e}_{R}\gamma^{\mu}e_{R}) + \frac{\bar{c}_{Hu}}{v^{2}} (iH^{\dagger} \vec{D}_{\mu}H)(\bar{u}_{R}\gamma^{\mu}u_{R}) \\ &+ \frac{\bar{c}_{Hd}}{v^{2}} (iH^{\dagger} \vec{D}_{\mu}H)(\bar{d}_{R}\gamma^{\mu}d_{R}) + \frac{\bar{c}'_{Hq}}{v^{2}} (iH^{\dagger}\sigma^{a} \vec{D}_{\mu}H)(\bar{Q}_{L}\sigma^{a}\gamma^{\mu}Q_{L}) \\ &+ \frac{\bar{c}_{Hq}}{v^{2}} (iH^{\dagger} \vec{D}_{\mu}H)(\bar{Q}_{L}\gamma^{\mu}Q_{L}) + \frac{\bar{c}_{HW}}{m_{W}^{2}} ig(D^{\mu}H)^{\dagger}\sigma^{a}(D^{\nu}H)W_{\mu\nu}^{a} + \frac{\bar{c}_{HB}}{m_{W}^{2}} ig'(D^{\mu}H)^{\dagger}(D^{\nu}H)B_{\mu\nu} \\ &+ \frac{\bar{c}_{3W}}{m_{W}^{2}}g^{3}\epsilon_{abc}W_{\mu}^{a\nu}W_{\nu\rho}^{b}W^{c\rho\mu} + \frac{\bar{c}_{g}}{m_{W}^{2}}g_{s}^{2}|H|^{2}G_{\mu\nu}^{A}G^{A\mu\nu} + \frac{\bar{c}_{\gamma}}{m_{W}^{2}}g'^{2}|H|^{2}B_{\mu\nu}B^{\mu\nu} \\ &+ \frac{\bar{c}_{H}}{v^{2}}\frac{1}{2}(\partial^{\mu}|H|^{2})^{2} - \sum_{f=e,u,d}\frac{\bar{c}_{f}}{m_{W}^{2}}y_{f}|H|^{2}\bar{F}_{L}H^{(c)}f_{R} \\ &+ \frac{\bar{c}_{3G}}{m_{W}^{2}}g_{s}^{3}f_{ABC}G_{\mu}^{A\nu}G_{\nu}^{A\mu}G_{\rho}^{C\mu} - \frac{\bar{c}_{uG}}{m_{W}^{2}}4g_{s}y_{u}H^{\dagger} \cdot \bar{Q}_{L}\gamma^{\mu\nu}T_{a}u_{R}G_{\mu\nu}^{A}. \end{aligned}$$

Updated Global SMEFT Fit

• Warsaw basis

$$\begin{split} \mathcal{L}_{\mathrm{SMEFT}}^{\mathrm{Warsaw}} \supset \frac{\bar{C}_{Hl}^{(3)}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}^{I}H)(\bar{l}\tau^{I}\gamma^{\mu}l) + \frac{\bar{C}_{Hl}^{(1)}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{l}\gamma^{\mu}l) + \frac{\bar{C}_{ll}}{v^2}(\bar{l}\gamma_{\mu}l)(\bar{l}\gamma^{\mu}l) \\ &+ \frac{\bar{C}_{HD}}{v^2} \left| H^{\dagger}D_{\mu}H \right|^2 + \frac{\bar{C}_{HWB}}{v^2} H^{\dagger}\tau^{I}H W_{\mu\nu}^{I}B^{\mu\nu} \\ &+ \frac{\bar{C}_{He}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{e}\gamma^{\mu}e) + \frac{\bar{C}_{Hu}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{u}\gamma^{\mu}u) + \frac{\bar{C}_{Hd}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{d}\gamma^{\mu}d) \\ &+ \frac{\bar{C}_{Hq}^{(3)}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{q}\tau^{I}\gamma^{\mu}q) + \frac{\bar{C}_{Hq}^{(1)}}{v^2} (H^{\dagger}i\overleftrightarrow{D}_{\mu}H)(\bar{q}\gamma^{\mu}q) + \frac{\bar{C}_{W}}{v^2} \epsilon^{IJK} W_{\mu}^{I\nu} W_{\nu}^{J\rho} W_{\rho}^{K\mu} \end{split}$$

$$\begin{split} \mathcal{L}_{\text{SMEFT}}^{\text{Warsaw}} &\supset \frac{\bar{C}_{eH}}{v^2} (H^{\dagger}H) (\bar{l}eH) + \frac{\bar{C}_{dH}}{v^2} (H^{\dagger}H) (\bar{q}dH) + \frac{\bar{C}_{uH}}{v^2} (H^{\dagger}H) (\bar{q}u\widetilde{H}) \\ &+ \frac{\bar{C}_G}{v^2} f^{ABC} G^{A\nu}_{\mu} G^{B\rho}_{\nu} G^{C\mu}_{\rho} + \frac{\bar{C}_{H\square}}{v^2} (H^{\dagger}H) \Box (H^{\dagger}H) + \frac{\bar{C}_{uG}}{v^2} (\bar{q}\sigma^{\mu\nu}T^A u) \widetilde{H} G^A_{\mu\nu} \\ &+ \frac{\bar{C}_{HW}}{v^2} H^{\dagger}H W^I_{\mu\nu} W^{I\mu\nu} + \frac{\bar{C}_{HB}}{v^2} H^{\dagger}H B_{\mu\nu} B^{\mu\nu} + \frac{\bar{C}_{HG}}{v^2} H^{\dagger}H G^A_{\mu\nu} G^{A\mu\nu} \,. \end{split}$$

Observables

• LEP and SLC EWPTs, M_W from ATLAS, Tevatron

Observable	Measurement	Ref.	SM Prediction	Ref.
$\Gamma_Z \; [\text{GeV}]$	2.4952 ± 0.0023	[39]	2.4943 ± 0.0005	[38]
$\sigma_{\rm had}^0$ [nb]	41.540 ± 0.037	[39]	41.488 ± 0.006	[38]
R^0_ℓ	20.767 ± 0.025	[39]	20.752 ± 0.005	[38]
$A^{0,\ell}_{ m FB}$	0.0171 ± 0.0010	[39]	0.01622 ± 0.00009	[114]
$\mathcal{A}_{\ell}\left(P_{\tau}\right)$	0.1465 ± 0.0033	[39]	0.1470 ± 0.0004	[114]
$\mathcal{A}_{\ell}(\mathrm{SLD})$	0.1513 ± 0.0021	[39]	0.1470 ± 0.0004	[114]
R_b^0	0.021629 ± 0.00066	[39]	0.2158 ± 0.00015	[38]
R_c^0	0.1721 ± 0.0030	[39]	0.17223 ± 0.00005	[38]
$A^{0,b}_{ m FB}$	0.0992 ± 0.0016	[39]	0.1031 ± 0.0003	[114]
$A^{0,c}_{ m FB}$	0.0707 ± 0.0035	[39]	0.0736 ± 0.0002	[114]
\mathcal{A}_b	0.923 ± 0.020	[39]	0.9347	[114]
\mathcal{A}_{c}	0.670 ± 0.027	[39]	0.6678 ± 0.0002	[114]
M_W [GeV]	80.387 ± 0.016	[40]	80.361 ± 0.006	[114]
M_W [GeV]	80.370 ± 0.019	[94]	80.361 ± 0.006	[114]

- LEP WW measurements
- ATLAS WW high pT overflow bin

Tevong You (Cambridge)

Observables

• ATLAS+CMS Higgs Run 1

Production	Decay	Signal Strength	Production	Decay	Signal Strength
ggF	$\gamma\gamma$	$1.10^{+0.23}_{-0.22}$	Wh	$\tau \tau$	-1.4 ± 1.4
ggF	ZZ	$1.13_{-0.31}^{+0.34}$	Wh	bb	1.0 ± 0.5
ggF	WW	0.84 ± 0.17	Zh	$\gamma\gamma$	$0.5^{+3.0}_{-2.5}$
ggF	$\tau \tau$	1.0 ± 0.6	Zh	WW	$5.9^{+2.6}_{-2.2}$
VBF	$\gamma\gamma$	1.3 ± 0.5	Zh	$\tau \tau$	$2.2^{+2.2}_{-1.8}$
VBF	ZZ	$0.1^{+1.1}_{-0.6}$	Zh	bb	0.4 ± 0.4
VBF	WW	1.2 ± 0.4	tth	$\gamma\gamma$	$2.2^{+1.6}_{-1.3}$
VBF	$\tau \tau$	1.3 ± 0.4	tth	WW	$5.0^{+1.8}_{-1.7}$
Wh	$\gamma\gamma$	$0.5^{+1.3}_{-1.2}$	tth	$\tau \tau$	$-1.9^{+3.7}_{-3.3}$
Wh	WW	$1.6^{+1.2}_{-1.0}$	tth	bb	1.1 ± 1.0
pp	$Z\gamma$	$2.7^{+4.6}_{-4.5}$	pp	$\mu\mu$	0.1 ± 2.5

Observables

• ATLAS+CMS Higgs Run 2

	Production	Decay	Sig. Stren.		Production	Decay	Sig. Stren.
[96]	1-jet, $p_T > 450$	$b\bar{b}$	$2.3^{+1.8}_{-1.6}$	[105]	pp	$\mu\mu$	-0.1 ± 1.4
[97]	Zh	$b\bar{b}$	0.9 ± 0.5	[106]	Zh	$b\bar{b}$	$0.69^{+0.35}_{-0.31}$
[97]	Wh	$b\bar{b}$	1.7 ± 0.7	[106]	Wh	$b\bar{b}$	$1.21_{-0.42}^{+0.45}$
[98]	$t\bar{t}h$	$b\bar{b}$	$-0.19^{+0.80}_{-0.81}$	[107]	$t\bar{t}h$	$b\bar{b}$	$0.84^{+0.64}_{-0.61}$
[99]	$t\bar{t}h$	$1\ell + 2\tau_h$	$-1.20^{+1.50}_{-1.47}$	[108]	$t\bar{t}h$	$2\ell os + 1\tau_h$	$1.7^{+2.1}_{-1.9}$
[99]	$t\bar{t}h$	$2\ell ss + 1\tau_h$	$0.86^{+0.79}_{-0.66}$	[108]	$t\bar{t}h$	$1\ell + 2\tau_h$	$-0.6^{+1.6}_{-1.5}$
[99]	$t\bar{t}h$	$3\ell + 1\tau_h$	$1.22^{+1.34}_{-1.00}$	[108]	$t\bar{t}h$	$3\ell + 1\tau_h$	$1.6^{+1.8}_{-1.3}$
[100]	$t\bar{t}h$	$2\ell ss$	$1.7^{+0.6}_{-0.5}$	[108]	$t\bar{t}h$	$2\ell ss + 1\tau_h$	$3.5^{+1.7}_{-1.3}$
[100]	$t\bar{t}h$	3ℓ	$1.0^{+0.8}_{-0.7}$	[108]	$t\bar{t}h$	3ℓ	$1.8^{+0.9}_{-0.7}$
[100]	$t\bar{t}h$	4ℓ	$0.9^{+2.3}_{-1.6}$	[108]	$t\bar{t}h$	$2\ell ss$	$1.5^{+0.7}_{-0.6}$
[101]	0-jet	WW	$0.9^{+0.4}_{-0.3}$	[109]	VBF	WW	$1.7^{+1.1}_{-0.9}$
[101]	1-jet	WW	1.1 ± 0.4	[109]	Wh	WW	$3.2^{+4.4}_{-4.2}$
[101]	2-jet	WW	1.3 ± 1.0	[110]	$B(h \to \gamma \gamma) / B(h \to 4\ell)$		$0.69^{+0.15}_{-0.13}$
[101]	VBF 2-jet	WW	1.4 ± 0.8	[110]	0-jet	4ℓ	$1.07^{+0.27}_{-0.25}$
[101]	Vh 2-jet	WW	$2.1^{+2.3}_{-2.2}$	[110]	1-jet, $p_T < 60$	4ℓ	$0.67^{+0.72}_{-0.68}$
[101]	Wh 3-lep	WW	-1.4 ± 1.5	[110]	1-jet, $p_T \in (60, 120)$	4ℓ	$1.00^{+0.63}_{-0.55}$
[102]	ggF	$\gamma\gamma$	$1.11_{-0.18}^{+0.19}$	[110]	1-jet, $p_T \in (120, 200)$	4ℓ	$2.1^{+1.5}_{-1.3}$
[102]	VBF	$\gamma\gamma$	$0.5^{+0.6}_{-0.5}$	[110]	2-jet	4ℓ	$2.2^{+1.1}_{-1.0}$
[102]	$t\bar{t}h$	$\gamma\gamma$	2.2 ± 0.9	[110]	"BSM-like"	4ℓ	$2.3^{+1.2}_{-1.0}$
[102]	Vh	$\gamma\gamma$	$2.3^{+1.1}_{-1.0}$	[110]	VBF, $p_T < 200$	4ℓ	$2.14^{+0.94}_{-0.77}$
[103]	ggF	4ℓ	$1.20^{+0.22}_{-0.21}$	[110]	Vh lep	4ℓ	$0.3^{+1.3}_{-1.2}$
[104]	0-jet	$\tau \tau$	0.84 ± 0.89	[110]	$t\bar{t}h$	4ℓ	$0.51^{+0.86}_{-0.70}$
[104]	boosted	$\tau \tau$	$1.17\substack{+0.47\\-0.40}$				
[104]	VBF	$\tau \tau$	$1.11_{-0.35}^{+0.34}$				

Observables

Including kinematical information facilitated by **STXS**

Sig. Stren. Production Sig. Stren. Production Decay Decay $2.3^{+1.8}_{-1.6}$ [96] 1-jet, $p_T > 450$ $b\bar{b}$ 105 -0.1 ± 1.4 pp $\mu\mu$ $0.69^{+0.35}_{-0.31}$ [97 Zh $b\bar{b}$ Zh $b\bar{b}$ 0.9 ± 0.5 106 $1.21_{-0.42}^{+0.45}$ Wh $b\bar{b}$ Wh $b\bar{b}$ 97 1.7 ± 0.7 [106] $t\bar{t}h$ $b\bar{b}$ $-0.19^{+0.80}_{-0.81}$ $t\bar{t}h$ $b\bar{b}$ $0.84_{-0.61}^{+0.64}$ [98] [107] $-1.20^{+1.50}_{-1.47}$ $1.7^{+2.1}_{-1.9}$ [99] $t\bar{t}h$ [108] $t\bar{t}h$ $1\ell + 2\tau_h$ $2\ell os + 1\tau_h$ $0.86^{+0.79}_{-0.66}$ $-0.6^{+1.6}_{-1.5}$ [99] $t\bar{t}h$ $2\ell ss + 1\tau_h$ [108] $t\bar{t}h$ $1\ell + 2\tau_h$ $1.22^{+1.34}_{-1.00}$ $1.6^{+1.8}_{-1.3}$ $t\bar{t}h$ $t\bar{t}h$ [99] [108] $3\ell + 1\tau_h$ $3\ell + 1\tau_h$ $1.7^{+0.6}_{-0.5}$ $3.5^{+1.7}_{-1.3}$ $t\bar{t}h$ $t\bar{t}h$ [100] $2\ell ss$ [108] $2\ell ss + 1\tau_h$ $1.0^{+0.8}_{-0.7}$ $1.8\substack{+0.9 \\ -0.7}$ [100] $t\bar{t}h$ [108] $t\bar{t}h$ 3ℓ 3ℓ $0.9^{+2.3}_{-1.6}$ $1.5^{+0.7}_{-0.6}$ $t\bar{t}h$ $t\bar{t}h$ [100] 4ℓ [108] $2\ell ss$ $0.9^{+0.4}_{-0.3}$ $1.7^{+1.1}_{-0.9}$ [101]0-jet WW[109]VBF WW $3.2^{+4.4}_{-4.2}$ [101]1-jet WW 1.1 ± 0.4 [109]WhWW $0.69^{+0.15}_{-0.13}$ $B(h \to \gamma \gamma) / B(h \to 4\ell)$ [101]2-jet WW 1.3 ± 1.0 [110] $1.07^{+0.27}_{-0.25}$ [101]VBF 2-jet WW 1.4 ± 0.8 [110]0-iet 4ℓ $0.67^{+0.72}_{-0.68}$ $2.1^{+2.3}_{-2.2}$ [101]Vh 2-jet WW[110]1-jet, $p_T < 60$ 4ℓ $1.00^{+0.63}_{-0.55}$ [101]Wh 3-lep WW -1.4 ± 1.5 [110] 1-jet, $p_T \in (60, 120)$ 4ℓ $1.11_{-0.18}^{+0.19}$ $2.1^{+1.5}_{-1.3}$ [102]ggF[110]1-jet, $p_T \in (120, 200)$ 4ℓ $\gamma\gamma$ $0.5\substack{+0.6 \\ -0.5}$ $2.2^{+1.1}_{-1.0}$ [102]VBF [110]2-jet 4ℓ $\gamma\gamma$ $2.3^{+1.2}_{-1.0}$ [102] $t\bar{t}h$ 2.2 ± 0.9 [110]"BSM-like" 4ℓ $\gamma\gamma$ $2.14_{-0.77}^{+0.94}$ [102] $2.3^{+1.1}_{-1.0}$ Vh[110]VBF, $p_T < 200$ 4ℓ $\gamma\gamma$ $1.20^{+0.22}_{-0.21}$ $0.3^{+1.3}_{-1.2}$ [103][110]Vh lep ggF 4ℓ 4ℓ $0.51_{-0.70}^{+0.86}$ $t\bar{t}h$ [104]0-jet 0.84 ± 0.89 [110] 4ℓ $\tau\tau$ $1.17\substack{+0.47 \\ -0.40}$ [104]boosted $\tau \tau$ $1.11_{-0.35}^{+0.34}$ [104]VBF $\tau \tau$

• ATLAS+CMS Higgs Run 2

• SILH basis, fit each operator individually



• SILH basis, fit each operator individually



• SILH basis, fit all operators simultaneously



• Warsaw basis, fit each operator individually


• Warsaw basis, fit all operators simultaneously

Marginalised 0.10.050. -0.05 pre-LHC Run 2 only All data ٠ -0.1 $\begin{array}{c|c} 10^{-1}\bar{C}_{dH} \\ 10^{-1}\bar{C}_{eH} \\ \bar{C}_{HB} \\ \bar{C}_{HB} \\ \bar{C}_{Hd} \\ \bar{C}_{Hd} \\ \bar{C}_{He} \\ \bar{C}_{Hu} \\ \bar{C}_{H$ \bar{C}_{HWB} $\bar{C}_{\ell\ell}$ $^{-1}\bar{C}_{uG}$ $^{-2}\bar{C}_{uH}$ $)^{-1}\bar{C}_W$

• Warsaw basis, improvement from Run 1 to 2 (lower is better) for individual fit



• Warsaw basis, improvement from Run 1 to 2 (lower is better) for marginalised fit



• Warsaw basis, summary







• Simplified models: stops (Run 1)





• Simplified models: stops (Run 2)

 $\tan\beta=20$



• Simplified models: renormalisable SM extensions

Name	Spin	SU(3)	SU(2)	U(1)	Name	Spin	SU(3)	SU(2)	U(1)
S	0	1	1	0	Δ_1	$\frac{1}{2}$	1	2	$-\frac{1}{2}$
\mathcal{S}_1	0	1	1	1	Δ_3	$\frac{1}{2}$	1	2	$-\frac{1}{2}$
φ	0	1	2	$\frac{1}{2}$	Σ	$\frac{1}{2}$	1	3	0
Ξ	0	1	3	0	Σ_1	$\frac{1}{2}$	1	3	-1
Ξ_1	0	1	3	1	U	$\frac{1}{2}$	3	1	$\frac{2}{3}$
B	1	1	1	0	D	$\frac{1}{2}$	3	1	$-\frac{1}{3}$
\mathcal{B}_1	1	1	1	1	Q_1	$\frac{1}{2}$	3	2	$\frac{1}{6}$
\mathcal{W}	1	1	3	0	Q_5	$\frac{1}{2}$	3	2	$-\frac{5}{6}$
\mathcal{W}_1	1	1	3	1	Q_7	$\frac{1}{2}$	3	2	$\frac{7}{6}$
N	$\frac{1}{2}$	1	1	0	T_1	$\frac{1}{2}$	3	3	$-\frac{1}{3}$
E	$\frac{1}{2}$	1	1	-1	T_2	$\frac{1}{2}$	3	3	$\frac{2}{3}$

Classification and tree-level matching dictionary

De Blas, Criado, Perez-Victoria, Santiago [1711.10391]

• Simplified models: renormalisable SM extensions



Model	χ^2	$\chi^2/n_{ m d}$	Coupling	Mass / TeV
SM	157	0.987	-	-
\mathcal{S}_1	156	0.986	$ y_{\mathcal{S}_1} ^2 = (6.3 \pm 5.9) \cdot 10^{-3}$	$M_{\mathcal{S}_1} = (9.0, 49)$
φ , Type I	156	0.986	$Z_6 \cdot \cos\beta = -0.64 \pm 0.59$	$M_{\varphi} = (0.9, 4.3)$
Ξ	155	0.984	$ \kappa_{\Xi} ^2 = (4.2 \pm 3.4) \cdot 10^{-3}$	$M_{\Xi} = (12, 35)$
N	155	0.978	$ \lambda_N ^2 = (1.8 \pm 1.2) \cdot 10^{-2}$	$M_N = (5.8, 13)$
\mathcal{W}_1	155	0.984	$\left \hat{g}_{\mathcal{W}_1}^{\phi} \right ^2 = (3.3 \pm 2.7) \cdot 10^{-3}$	$M_{W_1} = (4.1, 13)$
E	157	0.993	$ \lambda_E ^2 < 1.2 \cdot 10^{-2}$	$M_E > 9.2$
Δ_3	156	0.990	$ \lambda_{\Delta_3} ^2 < 1.9 \cdot 10^{-2}$	$M_{\Delta_3} > 7.3$
Σ	157	0.992	$ \lambda_{\Sigma} ^2 < 2.9 \cdot 10^{-2}$	$M_{\Sigma} > 5.9$
Q_5	156	0.990	$ \lambda_{Q_5} ^2 < 0.18$	$M_{Q_5} > 2.4$
T_2	157	0.992	$\left \lambda_{T_2}\right ^2 < 7.1 \cdot 10^{-2}$	$M_{T_2} > 3.8$
S	157	0.993	$\left y_{\mathcal{S}}\right ^2 < 0.32$	$M_S > 1.8$
Δ_1	157	0.993	$ \lambda_{\Delta_1} ^2 < 5.7 \cdot 10^{-3}$	$M_{\Delta_1} > 13$
Σ_1	157	0.993	$ \lambda_{\Sigma_1} ^2 < 7.3 \cdot 10^{-3}$	$M_{\Sigma_1} > 12$
U	157	0.993	$ \lambda_U ^2 < 2.8 \cdot 10^{-2}$	$M_U > 6.0$
D	157	0.993	$ \lambda_D ^2 < 1.4 \cdot 10^{-2}$	$M_D > 8.4$
Q_7	157	0.993	$ \lambda_{Q_7} ^2 < 7.7 \cdot 10^{-2}$	$M_{Q_7} > 3.6$
T_1	157	0.993	$ \lambda_{T_1} ^2 < 0.13$	$M_{T_1} > 3.0$
\mathcal{B}_1	157	0.993	$\left \hat{g}_{\mathcal{B}_{1}}^{\phi}\right ^{2} < 2.4 \cdot 10^{-3}$	$M_{\mathcal{B}_1} > 21$

Streamlines process of interpreting limits on BSM parameter space

Future e+e- Constraints



- Future precision sensitive to TeV scale, even for loop-induced operators
- One-loop matching simplified by a Universal One-Loop Effective Action

Henning, Lu, Murayama, 1412.1837; Drozd, J. Ellis, Quevillon, TY, 1512.03003; S.A.R. Ellis, Quevillon, TY, Z. Zhang, 1604.02445, 1706.07765.

- Part I: SM EFT
- Part 2: B Anomalies
- Part 3: Cosmological Relaxation

B anomalies

$$\mathcal{O}_{ij}^l = (\bar{s}\gamma^\mu P_i b)(\bar{l}\gamma_\mu P_j l)$$

- Anomalies in processes involving $b \rightarrow s \mu^+ \mu^-$ transitions:
- LHCb 3.4 σ in P5' angular distribution of $B \rightarrow K^* \mu^+ \mu^-$ (2 σ for Belle)
- Various other kinematic observables in $b \rightarrow s \ \mu^+\mu^-$
- 3.2 σ in $B_s \rightarrow \varphi \ \mu^+ \mu^-$
- ~4 σ non-zero Wilson coefficient in global fit to these "messy" observables
- 2.5 σ in "clean" observable R_K
- 2.5 σ in "clean" observable R_K^*
- ~4 σ non-zero Wilson coefficient in combined fit to just these two clean observables
- Consistency of all these various anomalies is non-trivial

Motivation for future colliders

- If $b \rightarrow s\mu^+\mu^-$ anomalies are confirmed, can we *definitely* discover directly the source (i.e. LQ/Z') at higher energies? (80 TeV unitarity limit = **no general no-lose theorem** at FCC-hh) Di Luzio, Nardecchia [1706.01868]
- Consider sensitivity to most **pessimistic** scenario: only include minimal couplings required to explain $b \rightarrow s\mu^+\mu^-$ anomalies



• More realistic models will only be *easier* to discover

• Extrapolate current 13 TeV di-muon search:







51

• Extrapolate current 13 TeV di-muon search:



Z' Sensitivity





52

• Extrapolate current 13 TeV di-muon search:







• Extrapolate current 13 TeV di-muon search:



• Actual limits depend on Z' couplings in signal x-section





 μ^+

• Extrapolate current 13 TeV di-muon search:





 \overline{s}

Z' Sensitivity



• Extrapolate current 13 TeV di-muon search:



• Actual limits depend on $Z'_{\text{Teyong You}}$ in signal x-section

• Extrapolate current 13 TeV di-muon search:



• 100 TeV can cover **all** parameter space of most *pessimistic* scenario



• Extrapolate current 13 TeV di-muon search:



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Allanach, Corbett, Dolan, TY [1810.tomorrow]

• **Indirect effects** from effective operators in LHC di-muon tail would point towards **fat Z'** at higher energies



Allanach, Corbett, Dolan, TY [1810.tomorrow]

• Improved study including **large widths**:



Leptoquark Sensitivity

• Extrapolate current 8 TeV LQ di-muon+di-jet search: group



- Pair production for scalar LQ depends only on QCD coupling
- Upper limit from Bs mixing constraint

Tevong You

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Take-home message

- Complete coverage of Z' models at 100 TeV FCC-hh
- Contrived LQ models may still survive FCC-hh
- Future studies: consider backgrounds, other channels, more realistic benchmark models, etc.
- Even if anomalies vanish, motivates **direct** discovery potential of future hadron colliders and interplay with **indirect** sensitivity from B physics

- Part I: SM EFT
- Part 2: B Anomalies
- Part 3: Cosmological Relaxation

Beyond the Standard Model?

• Hierarchy problem is still a problem: $(m_h)^2_{tree} + (m_h)^2_{radiative} = (m_h)^2_v$

$$\delta m_{\phi}^2 \propto m_{
m heavy}^2, \quad \delta m_{\psi} \propto m_{\psi} \log\left(rac{m_{
m heavy}}{\mu}
ight)$$

• Earliest example of an unnatural, arbitrary feature of a fundamental theory:

 $m_{inertial} = q_{gravity}$

• Classical electromagnetism fine-tuning:

$$(m_e c^2)_{
m obs} = (m_e c^2)_{
m bare} + \Delta E_{
m coulomb},$$

$$\Delta E_{\rm coulomb} = \frac{e^2}{4\pi\epsilon_0 r_e}$$

- Pions cut-off also at natural scale
- Higgs? Expect underlying explanation => fine-tuned unless new physics close to weak scale (just try writing down a model with a calculable Higgs potential...)

Beyond the Standard Model?



http://resonaances.blogspot.com.es/2016/01/do-or-die-year.html

• Maybe Nature is trying to tell us we are missing something in the way we think about the hierarchy problem

P. W. Graham, D. E. Kaplan and S. Rajendran, [arXiv:1504.07551]

L. F. Abbott, Phys. Lett. B 150 (1985) 427

• Higgs mass is naturally at large cut-off M

$$V_{\text{soft}}(a) \simeq (ga - \underline{M}^2)|h|^2 + gM^2a + \dots$$

- Axion-like particle *a* protected by shift symmetry, explicitly broken through technically-small parameter *g*
- Scans an effective Higgs mass
- Barriers switch on after EWSB

$$V_{\cos}(a) = \Lambda_G^4 \cos(a/f) \qquad \Lambda_G^4 \equiv \Lambda_G^{4-n} v^n$$



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• Trapped when barrier height = slow-roll slope



P. W. Graham, D. E. Kaplan and S. Rajendran, [arXiv:1504.07551]

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Constraints: H < v, classical rolling vs quantum, inflaton energy density dominates relaxion, etc.

Very small g and natural scanning range lead to super-planckian field excursions, exponential e-foldings...

• Trapped when barrier height = slow-roll slope



Relaxation Models

(apologies for lack of references)

$$V_{\cos}(a) = \Lambda_G^4 \cos(a/f) \quad \Lambda_G^4 \equiv \Lambda_G^{4-n} v^n \qquad gM^2 \sim \frac{\Lambda_G^{4-n} v^n}{f_\phi}$$

• **n=1 models** Graham et al [arXiv:1504.07551]

- G=QCD: Need additional ingredients to overcome strong-CP problem
- New gauge group G: new physics at weak scale + coincidence problem

• **n=2 models** Espinosa et al [arXiv:1506.09217]

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- Requires second scalar to relax relaxion barriers: doublescanning mechanism

• **n=0 models** Hook and Marques-Tavares [arXiv:1607.01786], **TY** [arXiv:1701.09167]

• More promising, make use of axial gauge coupling $\mathcal{L} = \frac{1}{32\pi^2} \frac{a}{f} \epsilon^{\mu\nu\rho\sigma} \text{Tr}G_{\mu\nu}G_{\rho\sigma}$

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Relaxation backreaction on inflation

TY [arXiv:1701.09167]

• Minimal relaxion setup, **no v-dependence in relaxion sector**

$$\mathcal{L} \supset \left(M^2 - g\phi\right)|h|^2 + gM^2\phi + ... + \Lambda_G^4 \cos\left(rac{\phi}{f_\phi}
ight) - rac{lpha_D}{f_D}\phi F_{\mu
u} ilde{F}^{\mu
u},$$

• Backreaction instead ends inflation



Inflating with electroweak dissipation: $\mathcal{L} \supset -\frac{\alpha}{f} \sigma F_{\mu\nu} \tilde{F}^{\mu\nu}$ $\ddot{\sigma} + 3H\dot{\sigma} + V'_{\sigma}(\sigma) = -I\frac{\alpha}{f} \left(\frac{H}{\xi}\right)^4 e^{2\pi\xi}$ $\xi \equiv \frac{\alpha}{2f} \frac{\dot{\sigma}}{H}$ See e.g. Anber and Sorbo 0908.4089

- Hubble falls
- Dark dissipation increases
- Relaxion loses KE and is trapped

	M	g	H_I	H_c	N_e	Λ_G	f_{ϕ}	f_D/α_D
$\sim [\text{GeV}]$	10^{8}	10^{-11}	10^{-2}	10^{-5}	10^{18}	$10^{3.5}$	10^{9}	10^{15}

Relaxation Models

(apologies for lack of references)

$$V_{\cos}(a) = \Lambda_G^4 \cos(a/f) \quad \Lambda_G^4 \equiv \Lambda_G^{4-n} v^n \qquad gM^2 \sim \frac{\Lambda_G^{4-n} v^n}{f_\phi}$$

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Relaxation backreaction on particle production

Hook and Marques-Tavares [arXiv:1607.01786]

• v-dependence in gauge particle production

- For M ~ 10-100 TeV sub-Planckian field excursions, no tiny parameters
- Model can be realised before, during, or after inflation

Relaxation backreaction on particle production

- Relaxation after inflation: **relaxion can reheat universe** but low T
- **Leptogenesis during reheating**: L and CP violation by higherdimensional operators parametrising decoupled new physics

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{\Lambda_1} \lambda_{1,ij} H H \bar{L}_j^c L_i + \frac{1}{\Lambda_2^2} \lambda_{2,ijkl} (\bar{L}_i \gamma^\mu L_j) (\bar{L}_k \gamma_\mu L_l) + \frac{1}{\Lambda_3^2} \lambda_{3,ijkl} (\bar{L}_i \gamma^\mu L_j) (\bar{E}_k \gamma_\mu E_l) + h.c.$$



Hamada & Kawana [arXiv:1510.05186]

• Attractive features in reheating leptogenesis for cosmological relaxation with particle production

Minho Son, Fang Ye, **TY** [1804.06599]

• **Minimal EFT setup** for naturally decoupled new physics

Conclusion

- A SM-like Higgs boson and no direct signs of new physics may turn out to be a significant experimental **null result**
- Null results may still lead to deeper understanding



- No new physics at the TeV scale could be our "Michelson-Morley" moment
- Experiment will *always* play a key role: **need future colliders**!

Lose theorem: If we *don't* go to higher energies we definitely *won't* have any direct knowledge of what fundamental physics may lie at 10 TeV

Conclusion

- Decoupled new physics motivates an SM EFT approach to phenomenology
- Future precision may probe even loop-induced operators at the TeV scale
- B anomalies could be the first indirect signs of new physics at accessible energy scales
- A desert above the weak scale has interesting implications for naturalness and model-building
- Cosmological relaxation mechanisms one possible avenue to explore

Observables

Including kinematical information facilitated by **STXS**

Sig. Stren. Production Sig. Stren. Production Decay Decay $2.3^{+1.8}_{-1.6}$ [96] 1-jet, $p_T > 450$ $b\bar{b}$ 105 -0.1 ± 1.4 pp $\mu\mu$ $0.69^{+0.35}_{-0.31}$ [97 Zh $b\bar{b}$ Zh $b\bar{b}$ 0.9 ± 0.5 106 $1.21_{-0.42}^{+0.45}$ Wh $b\bar{b}$ Wh $b\bar{b}$ 97 1.7 ± 0.7 [106] $t\bar{t}h$ $b\bar{b}$ $-0.19^{+0.80}_{-0.81}$ $t\bar{t}h$ $b\bar{b}$ $0.84_{-0.61}^{+0.64}$ [98] [107] $-1.20^{+1.50}_{-1.47}$ $1.7^{+2.1}_{-1.9}$ [99] $t\bar{t}h$ [108] $t\bar{t}h$ $1\ell + 2\tau_h$ $2\ell os + 1\tau_h$ $0.86^{+0.79}_{-0.66}$ $-0.6^{+1.6}_{-1.5}$ [99] $t\bar{t}h$ $2\ell ss + 1\tau_h$ [108] $t\bar{t}h$ $1\ell + 2\tau_h$ $1.22^{+1.34}_{-1.00}$ $1.6^{+1.8}_{-1.3}$ $t\bar{t}h$ $t\bar{t}h$ [99] [108] $3\ell + 1\tau_h$ $3\ell + 1\tau_h$ $1.7^{+0.6}_{-0.5}$ $3.5^{+1.7}_{-1.3}$ $t\bar{t}h$ $t\bar{t}h$ [100] $2\ell ss$ [108] $2\ell ss + 1\tau_h$ $1.0^{+0.8}_{-0.7}$ $1.8\substack{+0.9 \\ -0.7}$ [100] $t\bar{t}h$ [108] $t\bar{t}h$ 3ℓ 3ℓ $0.9^{+2.3}_{-1.6}$ $1.5^{+0.7}_{-0.6}$ $t\bar{t}h$ $t\bar{t}h$ [100] 4ℓ [108] $2\ell ss$ $0.9^{+0.4}_{-0.3}$ $1.7^{+1.1}_{-0.9}$ [101]0-jet WW[109]VBF WW $3.2^{+4.4}_{-4.2}$ [101]1-jet WW 1.1 ± 0.4 [109]WhWW $0.69^{+0.15}_{-0.13}$ $B(h \to \gamma \gamma) / B(h \to 4\ell)$ [101]2-jet WW 1.3 ± 1.0 [110] $1.07^{+0.27}_{-0.25}$ [101]VBF 2-jet WW 1.4 ± 0.8 [110]0-iet 4ℓ $0.67_{-0.68}^{+0.72}$ $2.1^{+2.3}_{-2.2}$ [101]Vh 2-jet WW[110]1-jet, $p_T < 60$ 4ℓ $1.00^{+0.63}_{-0.55}$ [101]Wh 3-lep WW -1.4 ± 1.5 [110] 1-jet, $p_T \in (60, 120)$ 4ℓ $1.11_{-0.18}^{+0.19}$ $2.1^{+1.5}_{-1.3}$ [102]ggF[110]1-jet, $p_T \in (120, 200)$ 4ℓ $\gamma\gamma$ $0.5\substack{+0.6 \\ -0.5}$ $2.2^{+1.1}_{-1.0}$ [102]VBF [110]2-jet 4ℓ $\gamma\gamma$ $2.3^{+1.2}_{-1.0}$ [102] $t\bar{t}h$ 2.2 ± 0.9 [110]"BSM-like" 4ℓ $\gamma\gamma$ $2.14_{-0.77}^{+0.94}$ [102] $2.3^{+1.1}_{-1.0}$ Vh[110]VBF, $p_T < 200$ 4ℓ $\gamma\gamma$ $1.20^{+0.22}_{-0.21}$ $0.3^{+1.3}_{-1.2}$ [103][110]Vh lep ggF 4ℓ 4ℓ $0.51_{-0.70}^{+0.86}$ $t\bar{t}h$ [104]0-jet 0.84 ± 0.89 [110] 4ℓ $\tau\tau$ $1.17\substack{+0.47 \\ -0.40}$ [104]boosted $\tau \tau$ $1.11_{-0.35}^{+0.34}$ [104]VBF $\tau \tau$

• ATLAS+CMS Higgs Run 2

Simplified Template Cross-Sections

ATLAS preliminary

Sub-division into kinematic regions for production processes



• Facilitates combination and interpretation

STXS measurements



• STXS dim-6 predictions

Cross-section region	$\sum_i A_i c_i$	
$gg \to H \ (0\text{-jet})$		
$gg \to H \ (1\text{-jet}, \ p_T^H < 60 \ \text{GeV})$	$56c'_a$	
$gg \rightarrow H \ (1\text{-jet}, \ 60 \leq p_T^H < 120 \ \text{GeV})$	5	
$gg \rightarrow H$ (1-jet, $120 \le p_T^H < 200 \text{ GeV}$)	$56c'_{g} + 18c3G + 11c2G$	
$gg \to H \ (1\text{-jet}, \ p_T^H \ge 200 \ \text{GeV})$	$56c'_{g} + 52$ c3G + 34c2G	
$gg \rightarrow H \ (\geq 2\text{-jet}, \ p_T^H < 60 \ { m GeV})$	$56c'_a$	
$gg \rightarrow H \ (\geq 2\text{-jet}, \ 60 \leq p_T^H < 120 \ \text{GeV})$	$56c'_q + 8c3G + 7c2G$	
$gg \rightarrow H \ (\geq 2\text{-jet}, \ 120 \leq p_T^H < 200 \ \text{GeV})$	$56c'_{g} + 23$ c3G $+ 18$ c2G	
$gg \rightarrow H \ (\geq 2\text{-jet}, \ p_T^H \geq 200 \ \text{GeV})$	$56c'_g + 90$ c3G + 68 c2G	
$gg ightarrow H~(\geq 2 ext{-jet VBF-like},~p_T^{j_3} < 25~ ext{GeV})$	$56c'_g$	
$gg ightarrow H~(\geq 2 ext{-jet VBF-like},~p_T^{j_3} \geq 25~ ext{GeV})$	$56c'_g + 9$ c3G $+ 8$ c2G	
$qq ightarrow Hqq~({ m VBF-like},~p_T^{j_3} < 25~{ m GeV})$	-1.0 cH - 1.0 cT + 1.3 cWW - 0.023 cB - 4.3 cHW	
	-0.29 cHB + 0.092 cHQ - 5.3 cpHQ - 0.33 cHu + 0.12 cHd	
$qq ightarrow Hqq~({ m VBF-like},~p_T^{j_3} \ge 25~{ m GeV})$	$-1.0 {\tt cH} - 1.1 {\tt cT} + 1.2 {\tt cWW} - 0.027 {\tt cB} - 5.8 {\tt cHW}$	
	-0.41 cHB + 0.13 cHQ - 6.9 cpHQ - 0.45 cHu + 0.15 cHd	
$qq ightarrow Hqq \; (p_T^j \geq 200 { m GeV})$	$-1.0 {\tt cH} - 0.95 {\tt cT} + 1.5 {\tt cWW} - 0.025 {\tt cB} - 3.6 {\tt cHW}$	
	-0.24 cHB + 0.084 cHQ - 4.5 cpHQ - 0.25 cHu + 0.1 cHd	
$qq \rightarrow Hqq \ (60 \le m_{jj} < 120 \text{ GeV})$	-0.99 cH - 1.2 cT + 7.8 cWW - 0.19 cB - 31 cHW	
	-2.4 cHB + 0.9 cHQ - 38 cpHQ - 2.8 cHu + 0.9 cHd	
$qq \rightarrow Hqq ~({ m rest})$	$-1.0 \mathtt{cH} - 1.0 \mathtt{cT} + 1.4 \mathtt{cWW} - 0.028 \mathtt{cB} - 6.2 \mathtt{cHW}$	
	-0.42 cHB + 0.14 cHQ - 6.9 cpHQ - 0.42 cHu + 0.16 cHd	
$aa/a\bar{a} \rightarrow ttH$	$-0.98 { m cH}+2.9 { m cu}+0.93 { m cG}+310 { m cuG}$	Hays, Sanz, Zemaityte
$gg/qq \rightarrow \iota \iota m$	+27c3G -13 c2G	

Tevong You (Cambridge)

STXS

• STXS dim-6 predictions



Good agreement with optimised non-STXS fit



• Though more information lost in VH case

De Blas, Lohwasser, Musella, Mimasu [in progress]