

**Particle Physics
beyond the LHC**

Higgs-Maxwell
Workshop



UNIVERSITY OF
LIVERPOOL

Muons

Themis Bowcock

THE ROYAL
SOCIETY
OF EDINBURGH

Experimental Particle Physics

- ◇ A little history
- ◇ What experiments & phenomenology?
 - ◇ $g-2$
 - ◇ $\text{Mu}3e$
 - ◇ $\text{Mu}2e$
 - ◇ (MuonE)
 - ◇ Theory
- ◇ UK effort
- ◇ Other muon measurements
- ◇ Why are we doing this now? Commentary
- ◇ Summary

Experimental Portfolio

2009

Experiment	Facility	Driver
ATLAS	CERN	Higgs/SUSY
LHCb	CERN	SUSY
T2K	J-PARC	ν Mixing

10 years preparing!



UNIVERSITY OF
LIVERPOOL

2019

Experiment	Facility	Driver
ATLAS	CERN	Higgs/SUSY
LHCb	CERN	Top/LFV
T2K	J-PARC	ν Mixing
HK	Kamiokande	CP
DUNE	FNAL/Sanford	CP
LZ	Sanford	Direct DM
g-2	FNAL	a, d
Mu2e	FNAL	LFV
Mu3e	PSI	LFV
CTA/LSST		DM/DE
MAGIS	FNAL	DM/Gravity



Over 10 years

Two European Strategies worth

What has changed for PP?

- 1) Discovery of Higgs
- 2) No evidence to date of SUSY
- 3) Neutrino Physics
- 4) New technologies (Quantum)

Barometer from undergraduate teaching in
UK – Zeitgeist

“what do PhD applicants find attractive?”



IT COULD HAVE BEEN
DIFFERENT !

Disclaimer: Particle Physics Beyond the LHC

- ◊ We are in a European Strategy consultation

Title does not imply *without* the LHC

- ◊ CERN remains a key part of our UK strategy delivering crucial physics, technology, R&D capability and training for fundamental physics
- ◊ Not discussing Future Colliders or CERNs strategy here
 - ◊ Prioritising and funding a diverse programme
 - ◊ Introducing new and dynamic ideas ;)
 - ◊ Informing future research at CERN

Experimental Portfolio

"SANS
CERN"
ACTIVITY

2009

Experiment	Facility	Driver
T2K	J-PARC	ν Mixing



UNIVERSITY OF
LIVERPOOL

2019 Particle Physics
beyond the LHC

Experiment	Facility	Driver
T2K	J-PARC	ν Mixing
HK	Kamiokande	CP
DUNE	FNAL/Sanford	CP
LZ	Sanford	Direct DM
g-2	FNAL	a, d
Mu2e	FNAL	LFV
Mu3e	PSI	LFV
CTA/LSST		DM/DE
MAGIS	FNAL	DM/Gravity

+Pheno



Physics Drivers 2019

Junior Staff

Dark Matter

Neutrino Nature

SM Tests Collider

Higgs Properties



**SM Tests Precision
(LFV, g-2)**

SUSY

Dark Energy

Axions

Senior Staff

Dark Matter (SUSY)

Higgs Properties

SM Tests Collider

Neutrino Nature

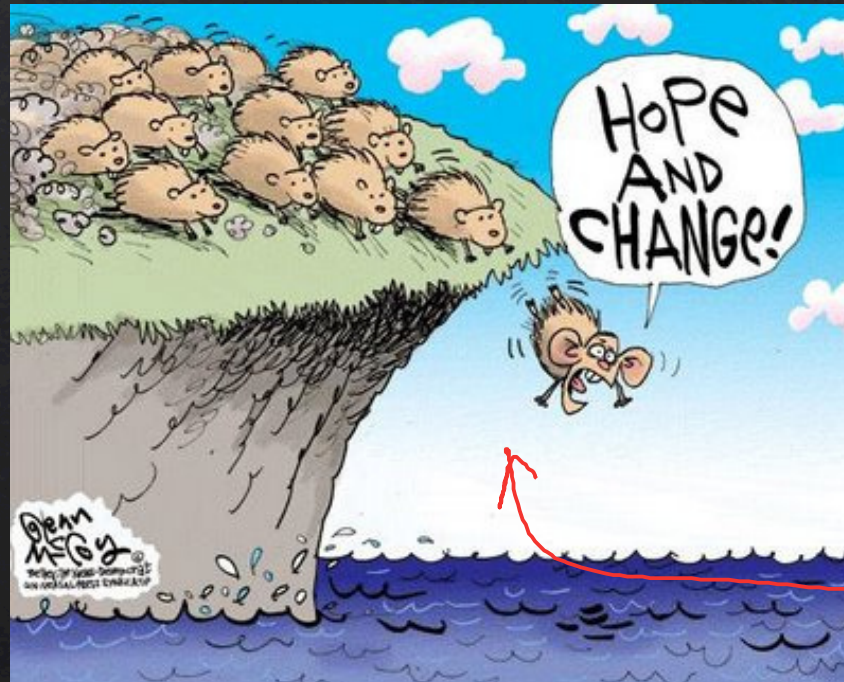
Quark CP

**SM Tests Precision
(LFV, g-2)**

Dark Energy

Axions

The “cliff” ...



2010 2015 2020 2025 2030

Physics Beyond Colliders



UNIVERSITY OF
LIVERPOOL

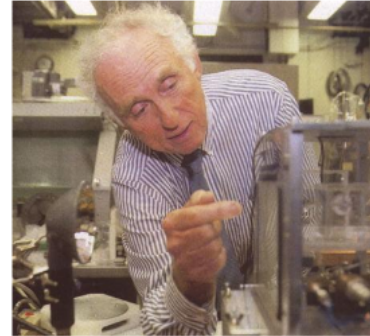
QUANTUM SENSORS
FOR FUNDAMENTAL
PHYSICS

Themis Bowcock

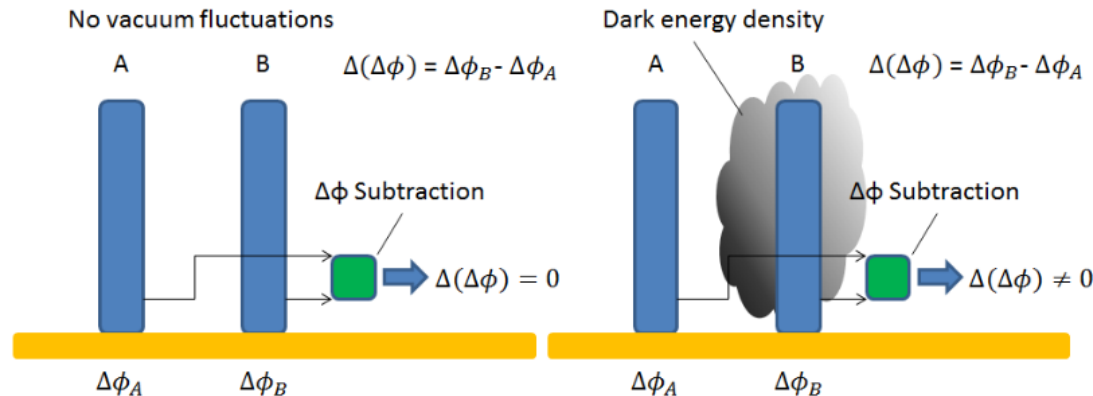
DEAI

Dark Energy

Work in Collaboration
with late Martin Perl
(Nobel Laureate,
visiting professor @
U. Of Liverpool)



A terrestrial search for dark contents of the vacuum, such as dark energy, using atom interferometry (Ronald J. Adler, Holger Mueller and Martin L. Perl)

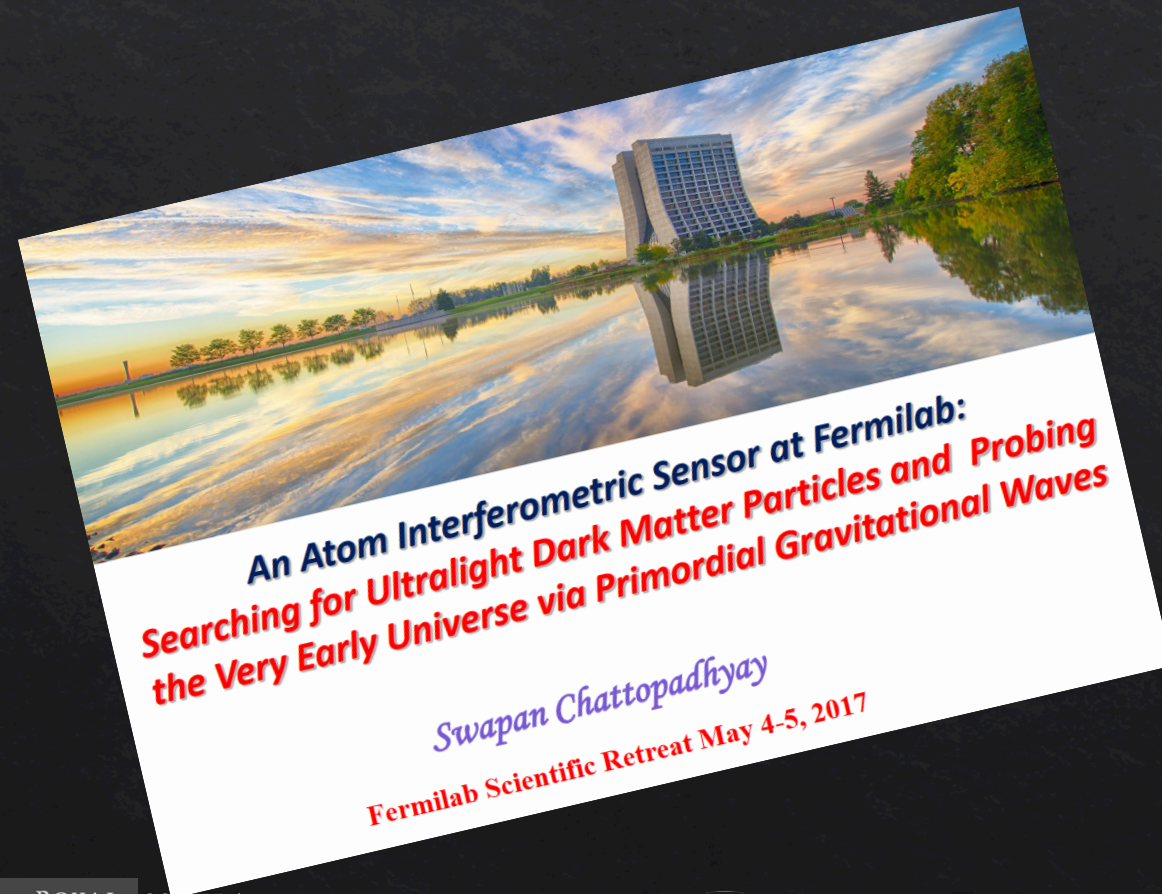


Liverpool
Experiment

Since 2008

Led to
QSFP
Initiative

New Initiative-MAGIS-100

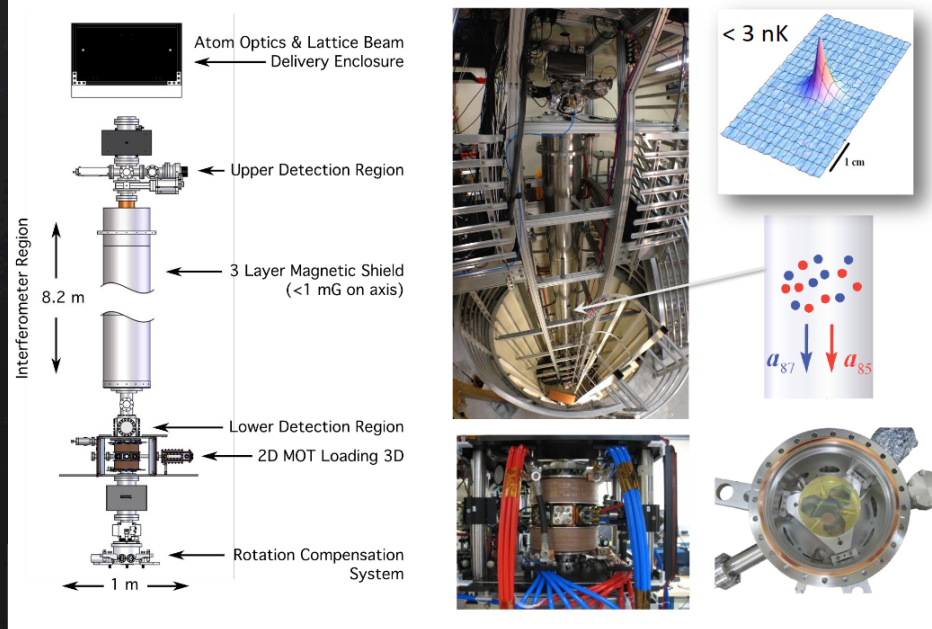


Stanford
Berkeley
FNAL
Liverpool

Magis-100



10 meter scale atomic fountain at Stanford



cosmic inflation in
frequency range
inaccessible to LIGO/LISA

Atomic interferometers
can also be sensitive detectors
of “dark” energy.

Magis-100

Mid-band Gravitational Wave Detection

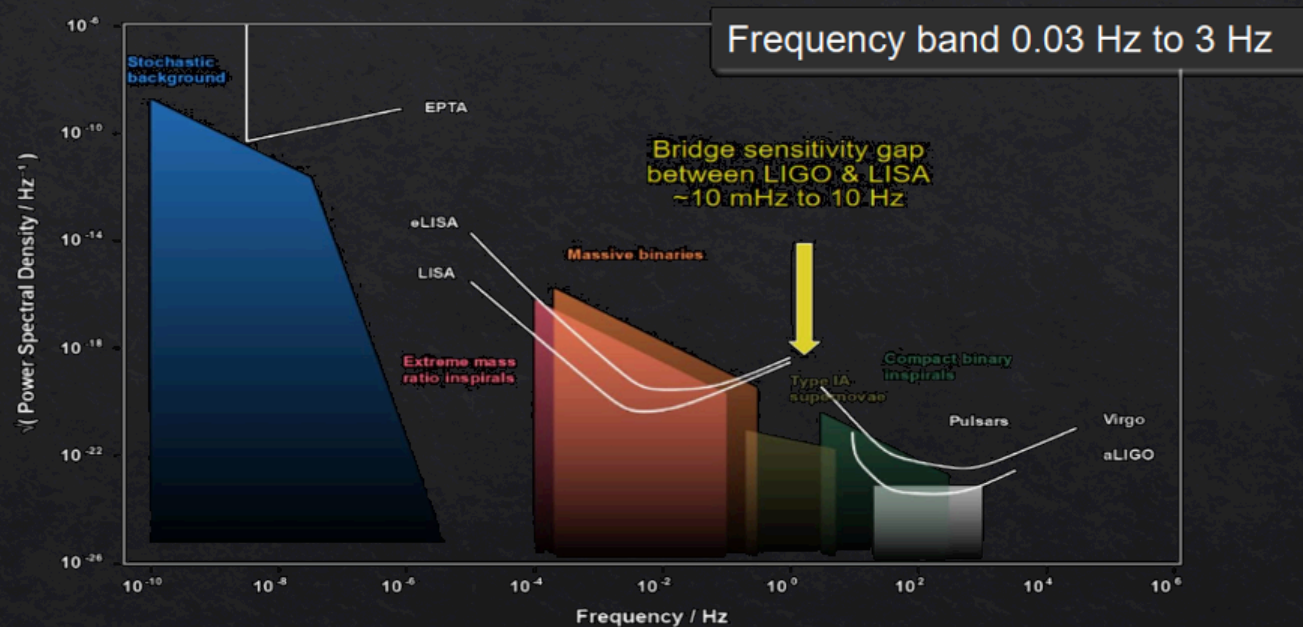


Figure: C. J. Moore et al., Class. Quantum Grav. **32**, 015014 (2015)

Many other things in QSFP

CONTACT ME

IAN SHARPEY..

See Martin Bauer's talk...

But note ... with QSFP style techniques we can probe new physics even up to Planck Scale (Beckenstein)

GREAT ENGAGEMENT WITH THEORY COMMUNITY...

EXCITEMENT & DYNAMISM

2013+ charged lepton programme

With **Mark Lancaster** (now spokesperson of g-2)

Joined and funded

g-2

Then (2016)

Mu2e/Mu3e

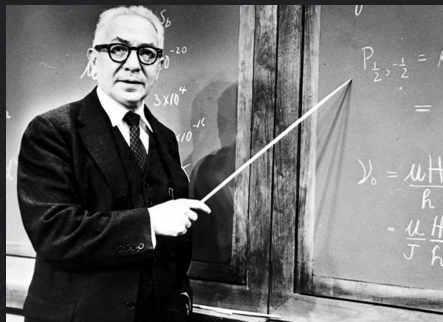
Future possibilities



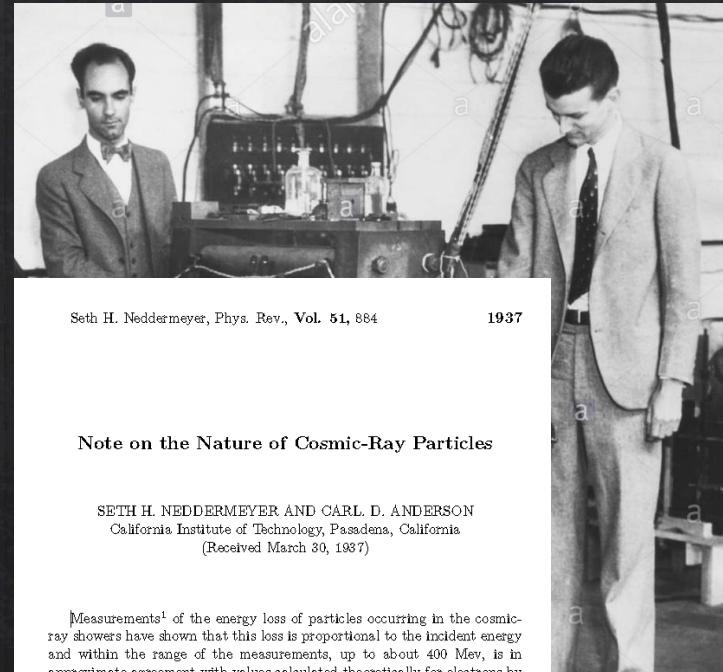
Why study them?

2019-1937 = 82 years

Positrons and then muons discovered in cosmic rays ...



"Who ordered that?"



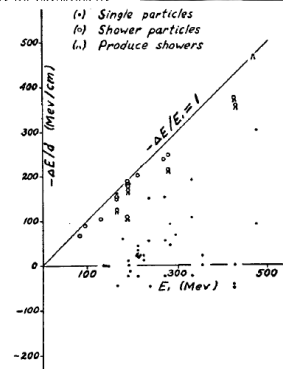
Seth H. Neddermeyer, Phys. Rev., Vol. 51, 884

1937

Note on the Nature of Cosmic-Ray Particles

SETH H. NEDDERMEYER AND CARL D. ANDERSON
California Institute of Technology, Pasadena, California
(Received March 30, 1937)

Measurements¹ of the energy loss of particles occurring in the cosmic-ray showers have shown that this loss is proportional to the incident energy and within the range of the measurements, up to about 400 Mev, is in approximate agreement with values calculated theoretically for electrons by Bethe and Heitler. These measurements were taken using (0.35 cm), and the observed individual losses were four amount below experimental detection up to the whole particle, with a mean fractional loss of about 0.5. If these correct it is evident that in a much thicker layer of heavy losses should become much more important, and the probability of a particle loss less than a large fraction of its initial energy is small. For the purpose of testing this inference and all previous measurements² which had shown the presence of less massive than protons but more penetrating than electrons, the Bethe-Heitler theory, we have taken about 6000 counter-tubes with a 1 cm plate of platinum placed across the center of this plate is equivalent in electron thickness to 1.96 or 1.86 cm of lead for a Z^2 absorption. The results of 58 particles in the range below 500 Mev are given in Fig. 1. The distribution of particles is shown as a function of the fractional energy loss. The shaded part of the diagram represents particles which



¹Anderson and Neddermeyer, Phys. Rev. 50, 263 (1936).

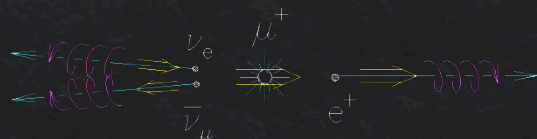
²Anderson and Neddermeyer, Report of London Conference, Vol. 1, p. 100 (1936).

Properties

Well known!

Mass	(2016)	$105.6583745 \pm 0.0000024 \text{ MeV}$
τ	(2013)	$2.1969811 \pm 0.0000022 \mu\text{s}$
$\tau \downarrow \mu \uparrow + / \tau \downarrow \mu \uparrow -$	(1984)	1.000024 ± 0.000078
a_μ	(2006)	$11659208.9 \pm 5.4 \pm 3.3 (\times 10^{-10})$
d_μ	(2009)	$11659208.9 \pm 5.4 \pm 3.3 (\times 10^{-19} \text{ e cm})$

Decays



	Mode	Fraction (Γ_i/Γ)
Γ_1	$e^- \bar{\nu}_e \nu_\mu$	$\approx 100\%$
Γ_2	$e^- \bar{\nu}_e \nu_\mu \gamma$	[a] $(6.0 \pm 0.5) \times 10^{-8}$
Γ_3	$e^- \bar{\nu}_e \nu_\mu e^+ e^-$	[b] $(3.4 \pm 0.4) \times 10^{-5}$

Citation: M. Tanaka et al. (Particle Data Group), Phys. Rev. D 98, 030001 (2018)



$$J = \frac{1}{2}$$

μ MASS (atomic mass units u)

The muon's mass is obtained from the muon-electron mass ratio as determined from the measurement of Zeeman transition frequencies in muonium (μ^+e^- atoms). Since the electron's mass is most accurately known in u , the muon's mass is also most accurately known in u . The conversion factor to MeV has approximately the same relative uncertainty as the mass of the muon in u . In this datablock we give the result in u , and in the following datablock in MeV.

VALUE (u)	DOCUMENT ID	TECH	COMMENT
0.113960037 ± 0.000000008	MOHR	16	RVUE 2014 CODATA value
*** We do not use the following data for averages, fits, limits, etc. ***			
0.113928267 ± 0.000000008	MOHR	12	RVUE 2010 CODATA value
0.113928256 ± 0.000000008	MOHR	08	RVUE 2006 CODATA value
0.113928264 ± 0.0000000030	MOHR	05	RVUE 2002 CODATA value
0.113928168 ± 0.0000000034	¹ MOHR	39	RVUE 1998 CODATA value
0.11392813 ± 0.0000000017	² COHEN	87	RVUE 1986 CODATA value

¹ MOHR 39 make use of other 1998 CODATA entries below.
² COHEN 87 make use of other 1986 CODATA entries below.

μ MASS

2010 CODATA (MOHR 12) gives the conversion factor from u (atomic mass units, see the above datablock) to MeV as 931.494 061(24). Earlier values use the then-current conversion factor. The conversion error contributes significantly to the uncertainty of the masses given below.

VALUE (MeV)	DOCUMENT ID	TECH	CHG	COMMENT
105.6583745 ± 0.0000004	MOHR	16	RVUE	2014 CODATA value
*** We do not use the following data for averages, fits, limits, etc. ***				
105.6583715 ± 0.00000035	MOHR	12	RVUE	2010 CODATA value
105.6583668 ± 0.00000038	MOHR	08	RVUE	2006 CODATA value
105.6583692 ± 0.00000094	MOHR	05	RVUE	2002 CODATA value
105.6583598 ± 0.00000052	MOHR	39	RVUE	1998 CODATA value
105.658353 ± 0.0000016	¹ COHEN	87	RVUE	1986 CODATA value
105.658396 ± 0.000044	² MARIAM	82	CNTR	+
105.65836 ± 0.00006	³ CROWE	72	CNTR	
105.65865 ± 0.000044	⁴ CRANE	71	CNTR	

¹ Converted to MeV using the 1986 CODATA value of the conversion constant, 931.494013 ± 0.000037 MeV/ u .

² MARIAM 82 give $m_\mu/m_0 = 206.768253(82)$.

³ CROWE 72 give $m_\mu/m_0 = 206.7682(5)$.

⁴ CRANE 71 give $m_\mu/m_0 = 206.76876(85)$.

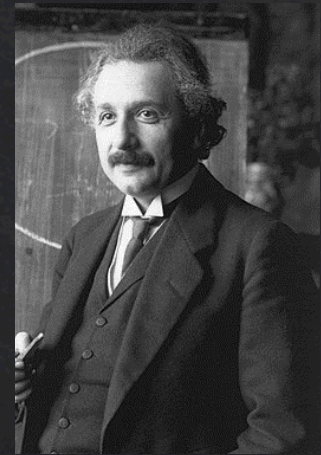
HTTP://PDG.LBL.GOV

Page 1

Created: 6/5/2018 19:00

Spin $\frac{1}{2}$ particle

$$i\hbar \partial \psi / \partial t = [p^2 / 2m - e / 2m (L + 2S) \cdot B]$$



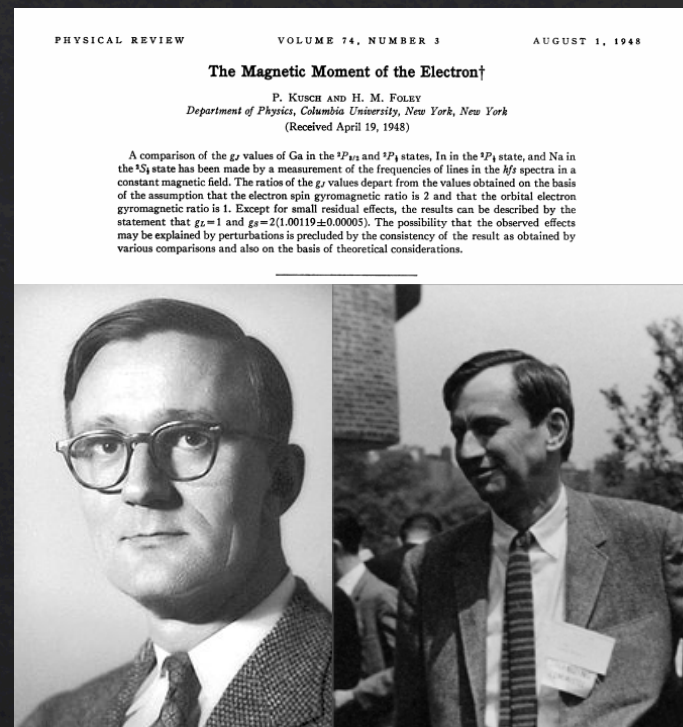
1948: Precise Measurement and Calculation (e)

Kusch and Foley measure g_e

$$g_e = 2.00238 \pm 0.00006$$

Anomalous Magnetic Moment

$$a \downarrow e = g - 2 / 2 = 0.00119 \pm 0.00003$$



1947: QED



$$g_e \approx 2\left(1 + \frac{\alpha}{2\pi}\right) \approx 2.00232$$



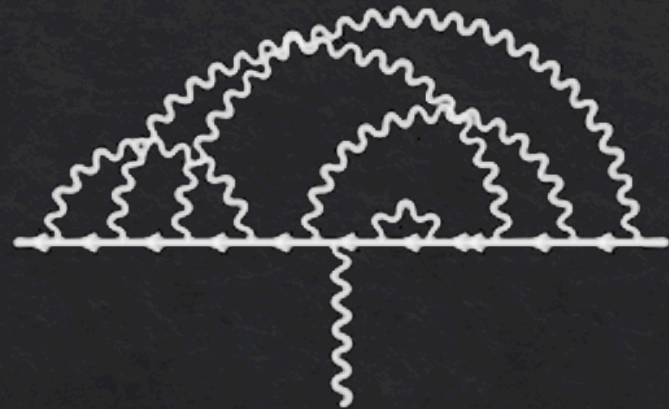
... and
Feynman and
Tomonaga



electron g-2 recently

Predicted: $\mu/\mu_B = -1.001\,159\,652\,181\,78\,(77)$

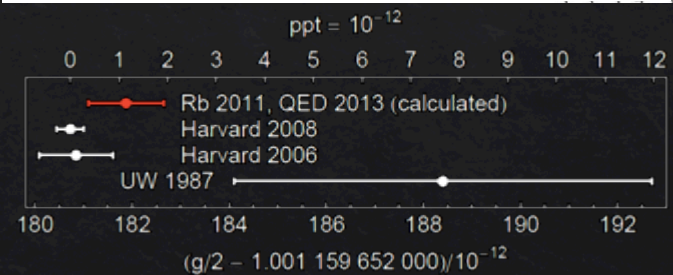
Measured: $\mu/\mu_B = -1.001\,159\,652\,180\,73\,(28)$



The standard model's greatest triumph

Gerald Gabrielse

December 2013 Physics Today



e g-2 status

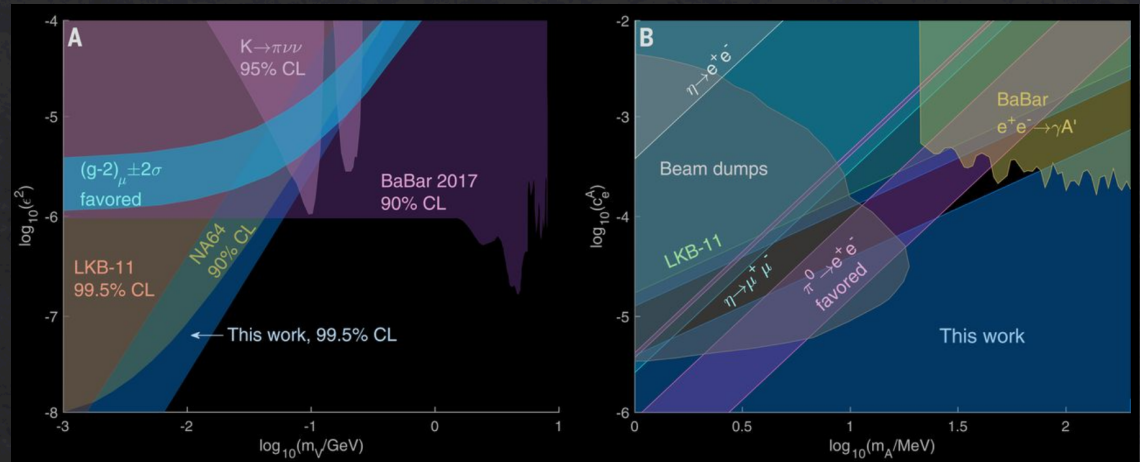
$$1/\alpha = 137.035999046(27)$$

Science, 13 Apr 2018: Vol. 360,
Issue 6385, pp. 191-195

2.5 σ difference

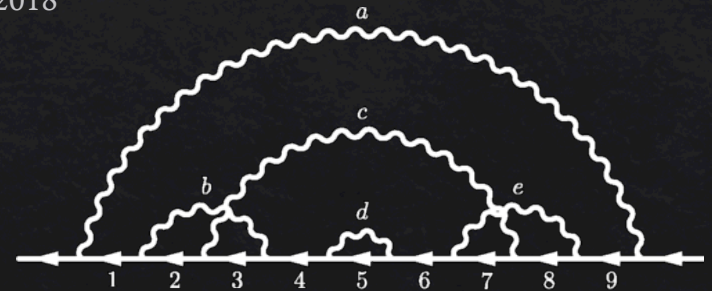
$a_e = 0.00115965218161(23)$
used matter-wave interferometry with a cloud of
cesium atoms to make the most accurate
measurement of α to date. QFSP

ATOM
INTERFEROMETRY



Revised and improved value of the QED tenth-order
electron anomalous magnetic moment
Tatsumi Aoyama, Toichiro Kinoshita, and Makiko
Nio

Phys. Rev. D 97, 036001 – Published 8 February
2018



Garwin, Lederman, Weinrich
2.00+/-0.10
Phys Rev 105, 1415 (Jan 57) @
Columbia

...how about muons?

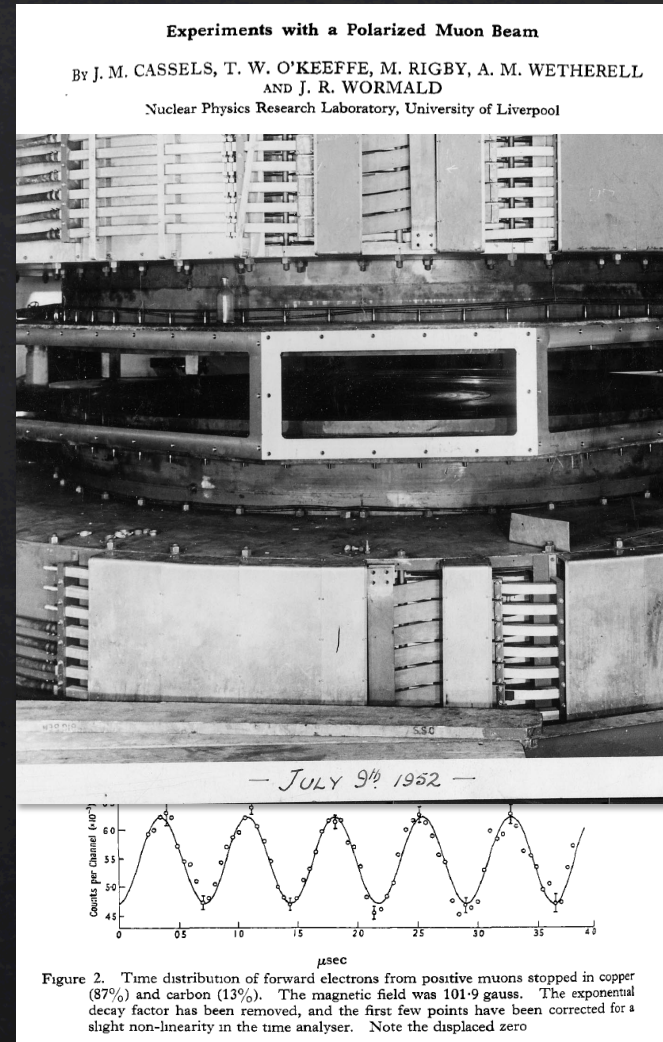
1933: Stern and Esterman $g_p=5.6$
Rabi $g_n=-3.8$

New heavy particle....

$$g=2.004 \pm 0.014 \text{ (0.6\%)}$$

$$a=0.002 \pm 0.007$$

1957 Proc. Phys. Soc. A 70 543

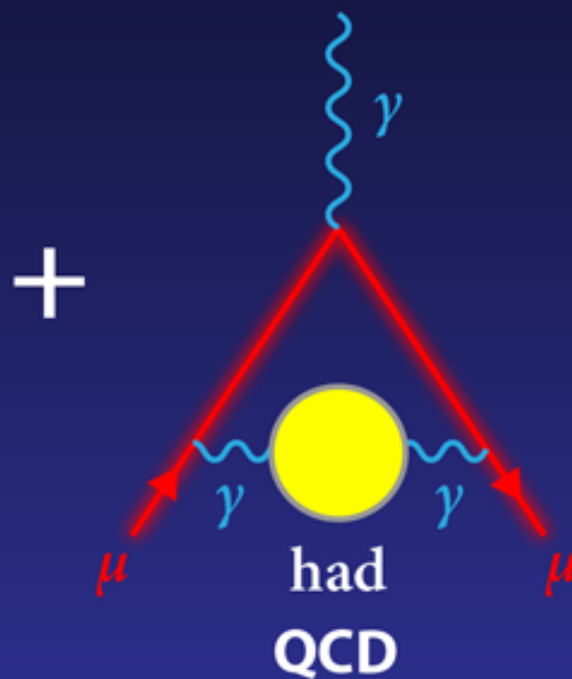
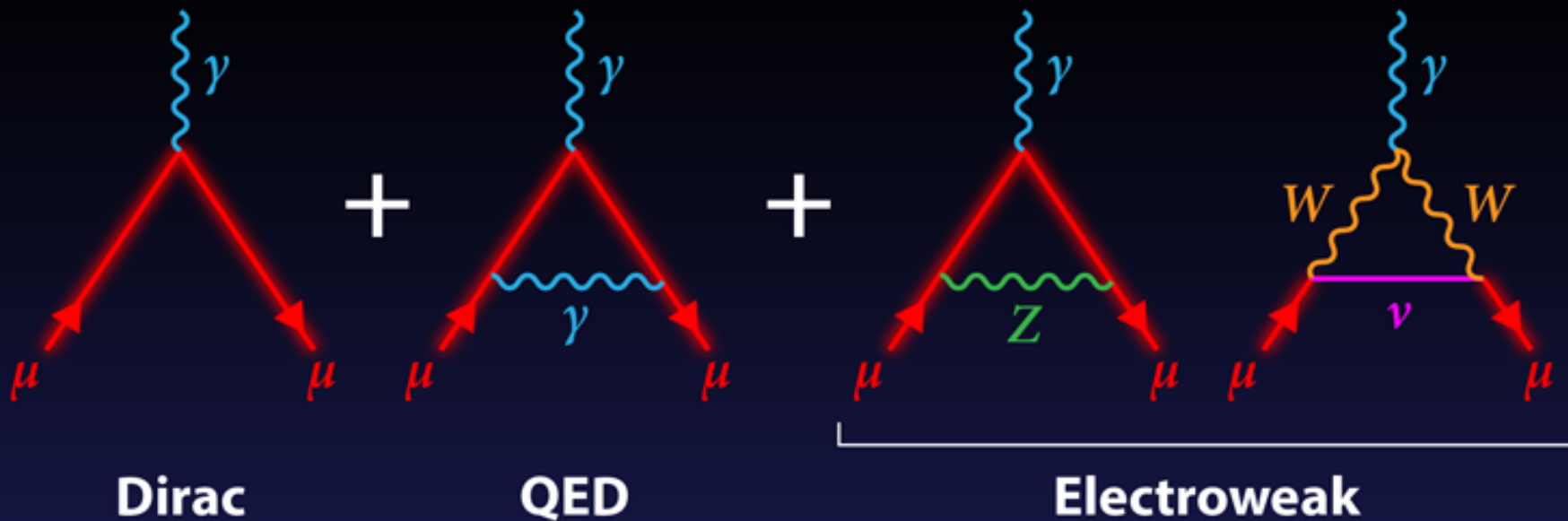


Today

June 25, 2018 • *Physics* 11, 65

“The muon anomalous magnetic moment is an important and unique quantity in subatomic physics, since its value represents a sum over all known standard model physics. This wide sensitivity exists because the anomalous moment depends on all particles in nature that can couple to the muon, including as-yet-undiscovered ones.”





W. Bennett *et al.*, “Final Report of the E821
Muon Anomalous Magnetic Moment
Measurement at BNL,”

[Phys. Rev. D **73**, 072003 \(2006\).](#)²¹

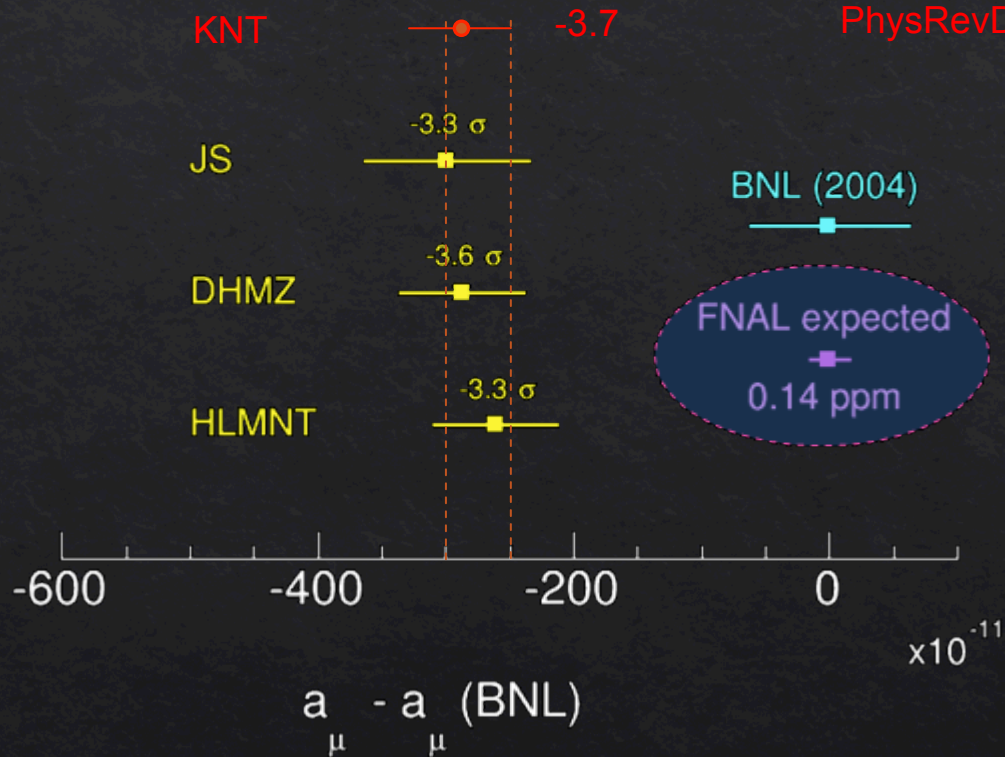
$$a_{\mu}(\text{Expt}) = 11659208.0(5.4) \\ (3.3) \times 10^{-10}$$

$$\Delta a_{\mu} = (27.06 \pm 7.26) \times 10^{-10}$$

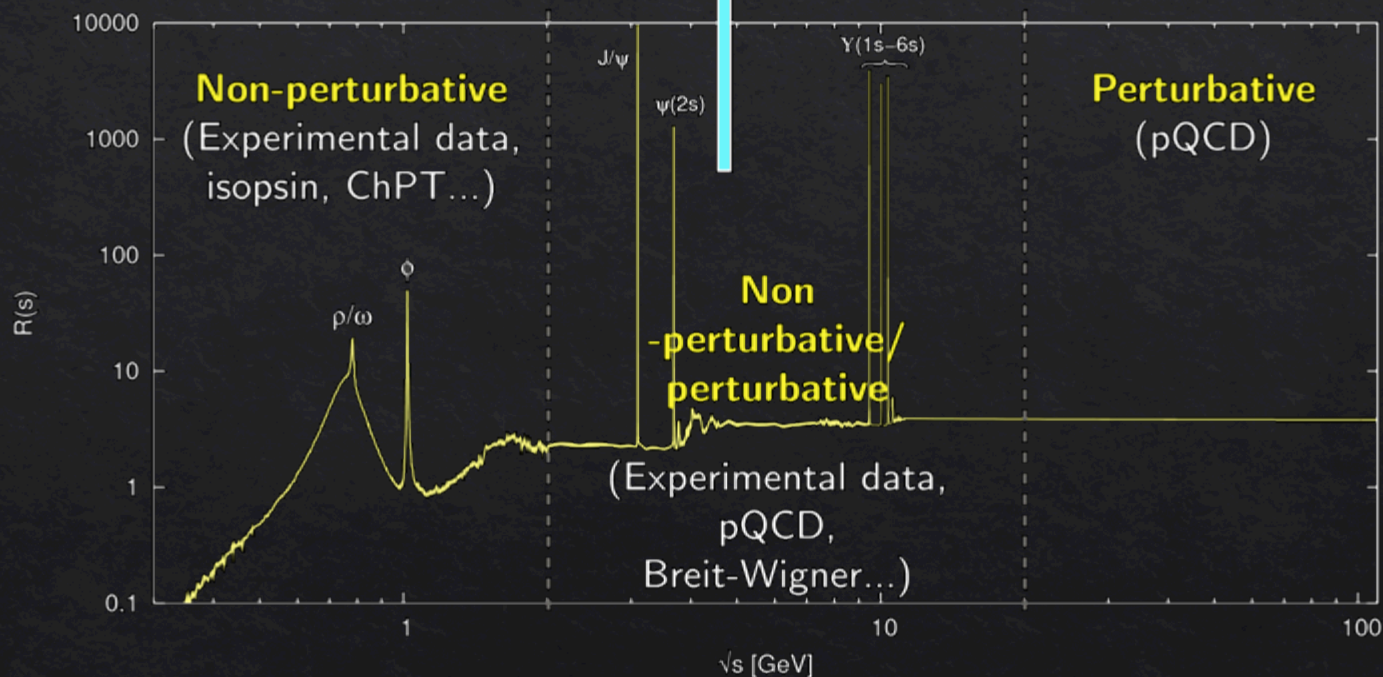


Comparison of SM & BNL Measurement

Keshavarzi, Nomura, Teubner
PhysRevD.97.114025

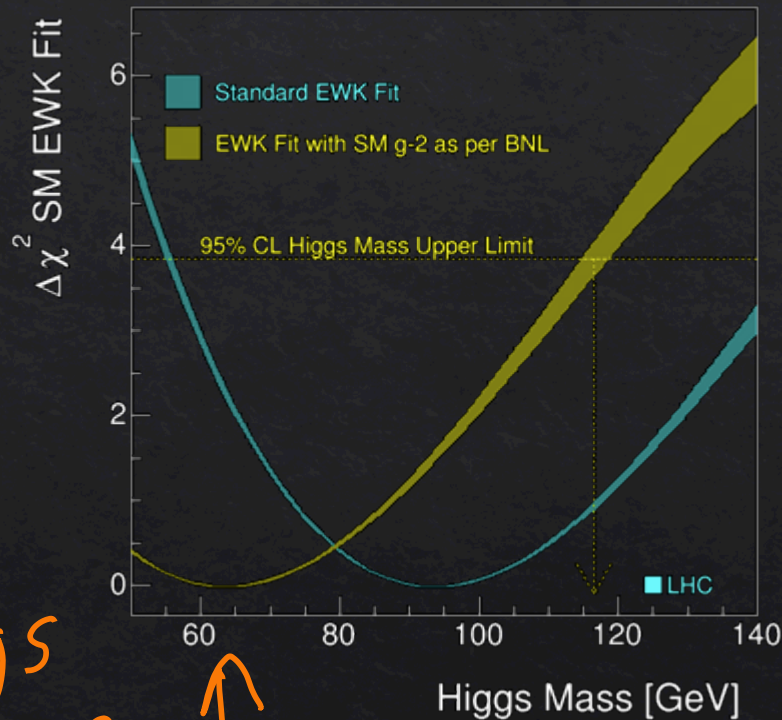


$$a_{\mu}^{\text{had, LO VP}} = \frac{\alpha^2}{3\pi^2} \int_{s_{th}}^{\infty} \frac{ds}{s} R(s) K(s), \text{ where } R(s) = \frac{\sigma_{\text{had},\gamma}^0(s)}{4\pi\alpha^2/3s}$$



Hadronic Corrections

For the BNL result to match the SM prediction then the SM hadronic estimate would need to be wrong by 6σ



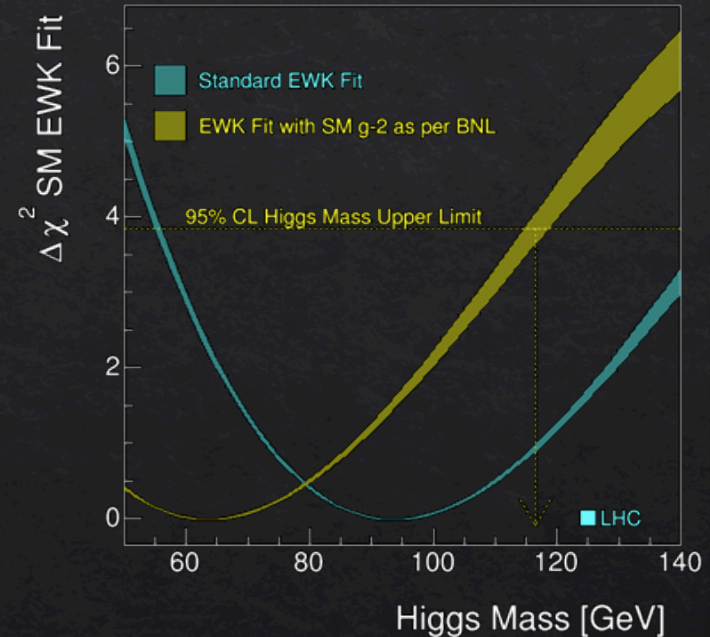
The beauty of the SM is that everything is related

"You cannot cook-up a zero $g-2$ SM anomaly and be consistent with the LHC Higgs mass!"

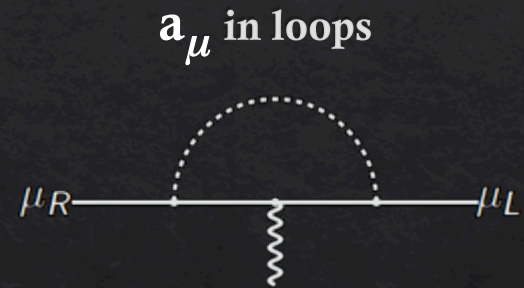
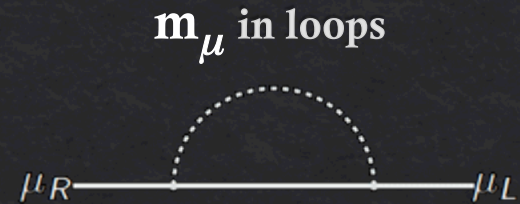
Higgs
with $g-2$ ↑

SM in action...

For the BNL result to match the SM prediction then the SM hadronic estimate would need to be wrong by 6σ



Any new physics
that contributes to
the muon mass can
contribute to a_μ



Why μ not e ?

Electron g-2 is presently measured x 2,000 better than muon g-2

But $\left(\frac{m_\mu}{m_e}\right)^2$ is 44,000.

2nd Generation Leptons v. useful.

Muon has sensitivity to new physics from < MeV to TeV.

New physics contributes as:

$$\left(\frac{m_\ell}{M_{\text{NEW}}}\right)^2$$

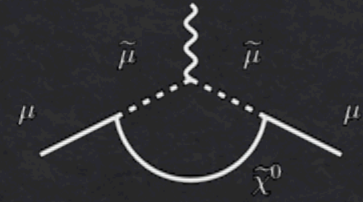
Difficult to use taus!

Any ideas for
Z's ???

Many BSM studies use g-2 as constraint or even motivation

SUSY could easily explain g-2

- Main 1-loop contributions:



- Simplest case:

$$a_{\mu}^{\text{SUSY}} \simeq \text{sgn}(\mu) 130 \times 10^{-11} \tan \beta \left(\frac{100 \text{ GeV}}{\Lambda_{\text{SUSY}}} \right)^2$$

- Needs $\mu > 0$, 'light' SUSY-scale Λ and/or large $\tan \beta$ to explain 281×10^{-11}

- This is already excluded by LHC searches in the simplest SUSY scenarios (like CMSSM); causes large χ^2 in simultaneous SUSY-fits with LHC data and g-2

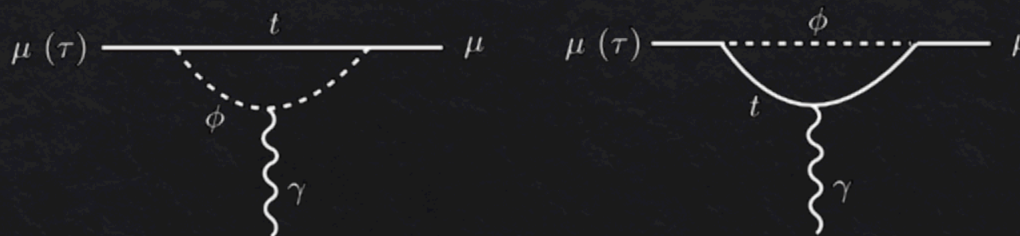
- However:

- * SUSY does not have to be minimal (w.r.t. Higgs),
- * could have large mass splittings (with lighter sleptons),
- * be hadrophobic/leptophilic,
- * or not be there at all, but don't write it off yet...

MANY THEORETICAL STUDIES

- Don't have to have full MSSM (like coded in GM2Calc [by Athron, ..., Stockinger et al., EPJC 76 (2016) 62], which includes all latest two-loop contributions), and
 - **extended Higgs sector** could do, see, e.g. Stockinger et al., JHEP 1701 (2017) 007, 'The muon magnetic moment in the 2HDM: complete two-loop result'
- lesson: 2-loop contributions can be highly relevant in both cases; one-loop analyses can be misleading
- **1 TeV Leptoquark** Bauer + Neubert, PRL 116 (2016) 141802

one new scalar could explain several anomalies seen by BaBar, Belle and LHC in the flavour sector (e.g. **violation of lepton universality** in $B \rightarrow K\ell\ell$, enhanced $B \rightarrow D\tau\nu$) and solve $g-2$, while satisfying all bounds from LEP and LHC



How to measure the anomaly

- ◇ Store longitudinally polarised muons in a dipole field
- ◇ Measure 2 quantities:
 - ◇ ω_a the precession frequency
 - ◇ $\langle B \rangle$ the average magnetic field sampled by the muon distribution

$$\boxed{\omega_a} = \omega_s - \omega_c = a_\mu \frac{e \langle B \rangle}{m_\mu c}$$

Larmor Precession Thomas Precession

$$\omega_s = \frac{geB}{2mc} + (1 - \gamma) \frac{eB}{\gamma mc}$$

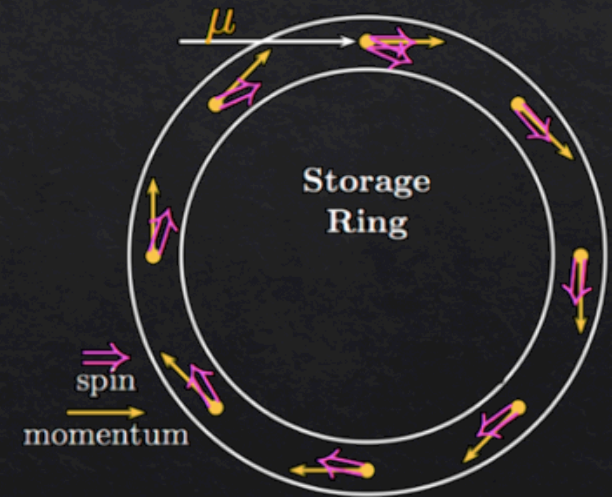
Spin Precession frequency

$\sim 140\text{ns}$

$$\omega_c = \frac{eB}{\gamma mc}$$

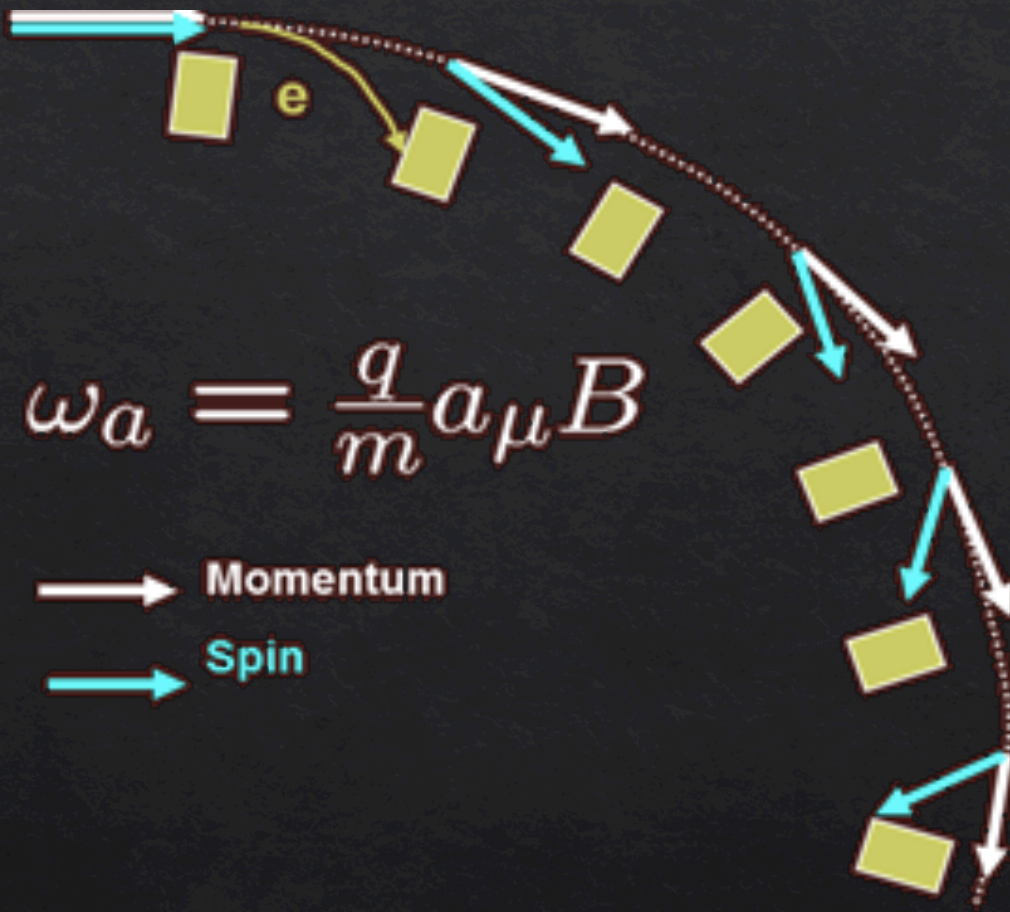
Cyclotron frequency

$\sim 149\text{ns}$

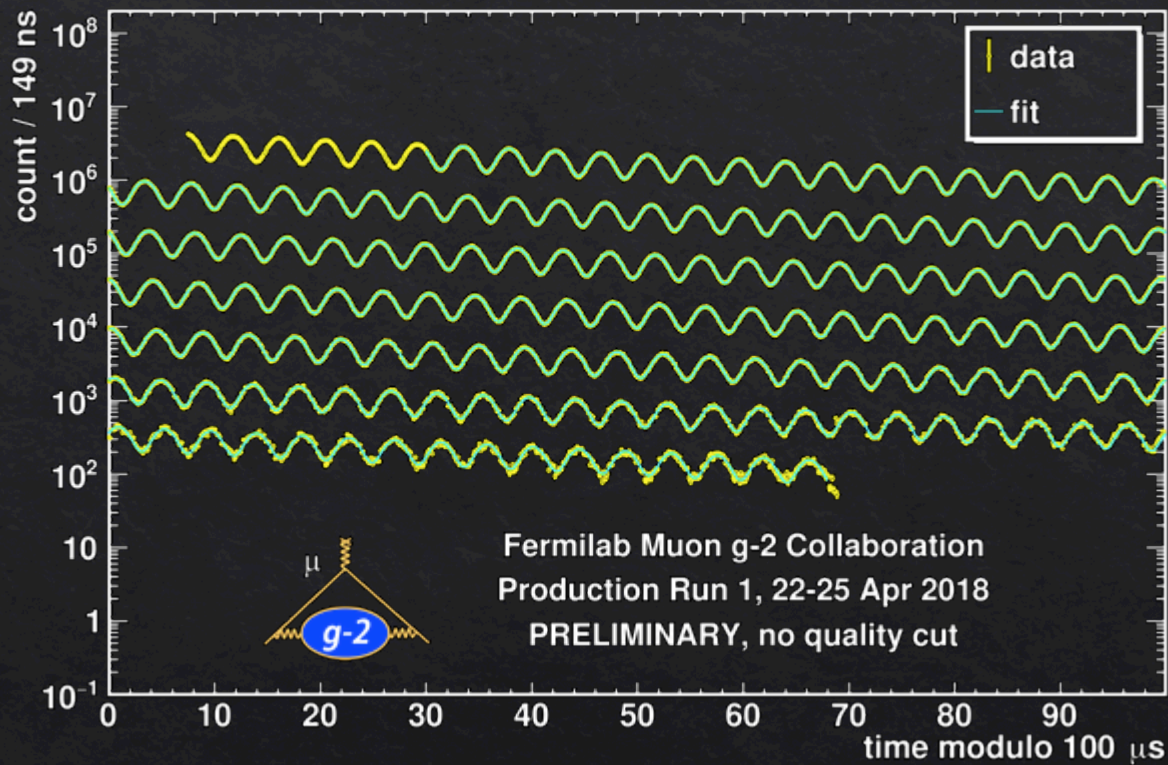


actual precession $\times 2$

"LIGHTHOUSE ON A CAROUSEL"

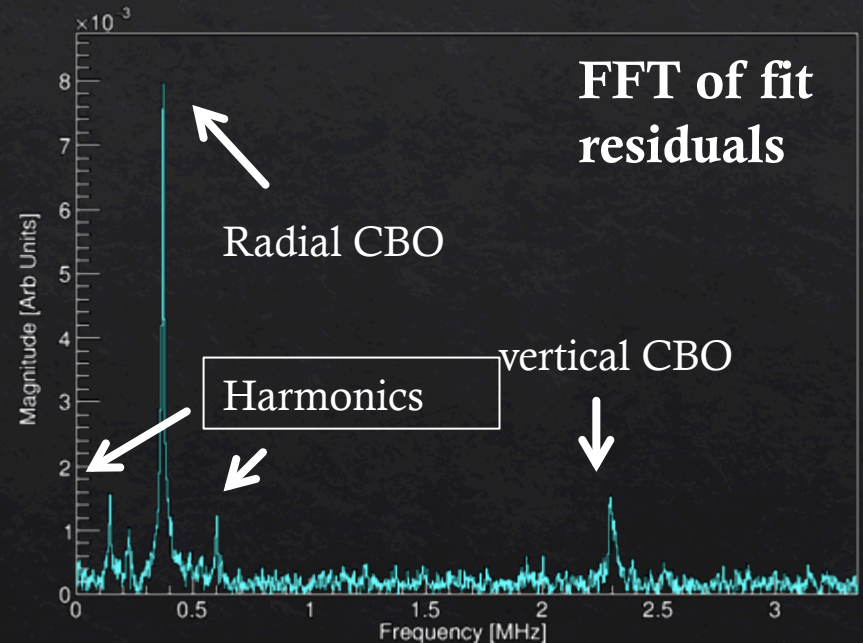


Measuring the precession



Beam oscillations

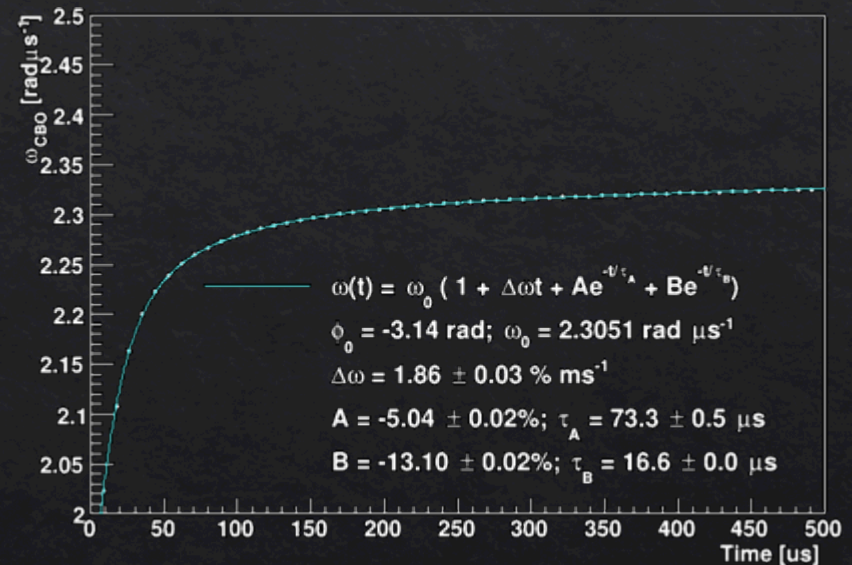
- ◊ Beam oscillations introduce additional fluctuations in the e^+ arrival spectrum that need to be accounted for
- ◊ Measured using the Liverpool-built tracking detectors



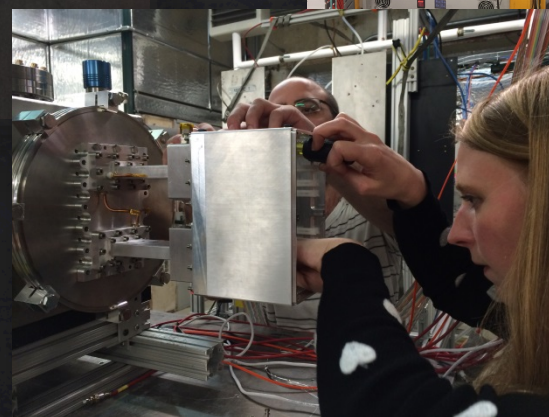
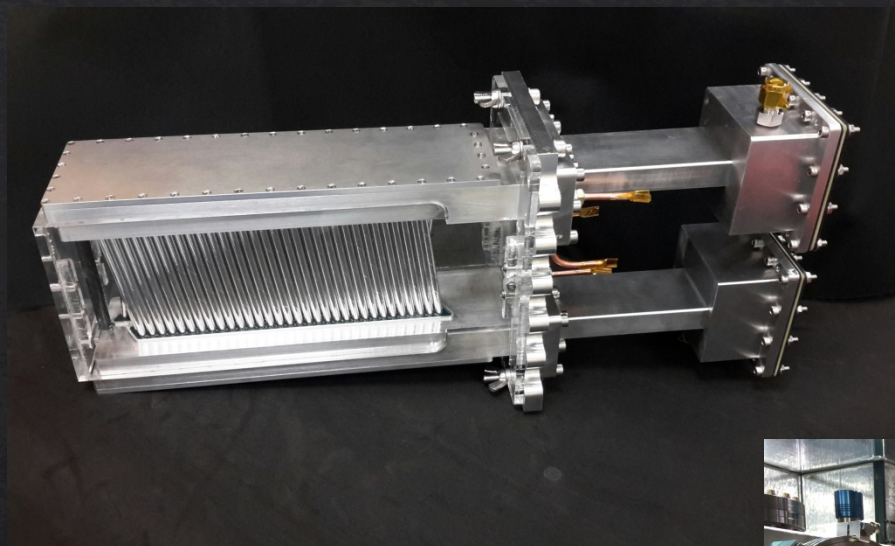
Changing frequency

UK CONTRIBUTIONS CRITICAL

- ◊ Frequency observed to change during fill
- ◊ Must be accounted for in fits
- ◊ Crucial measurement from the trackers!

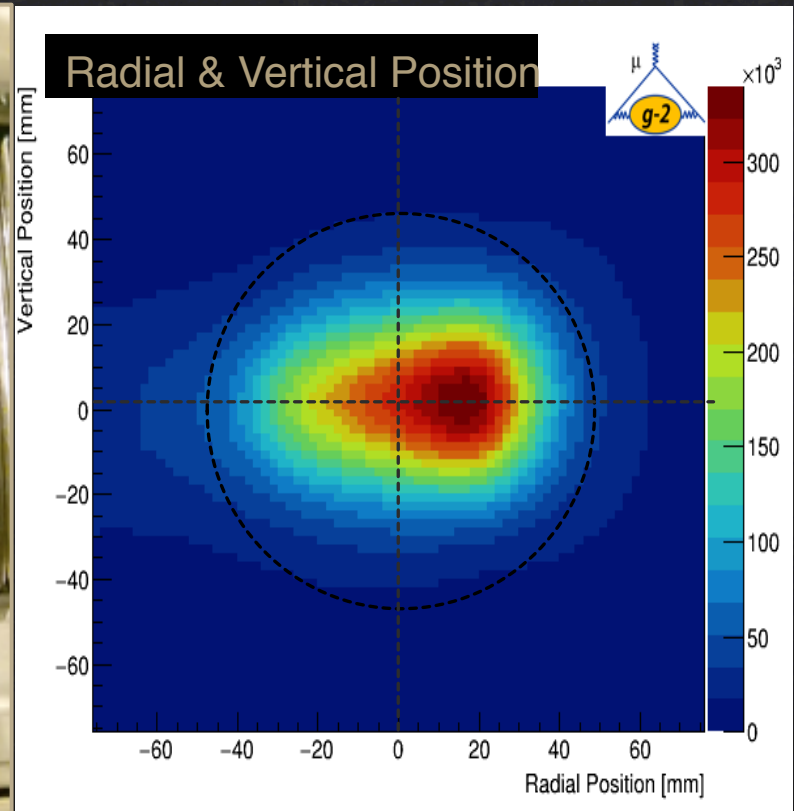
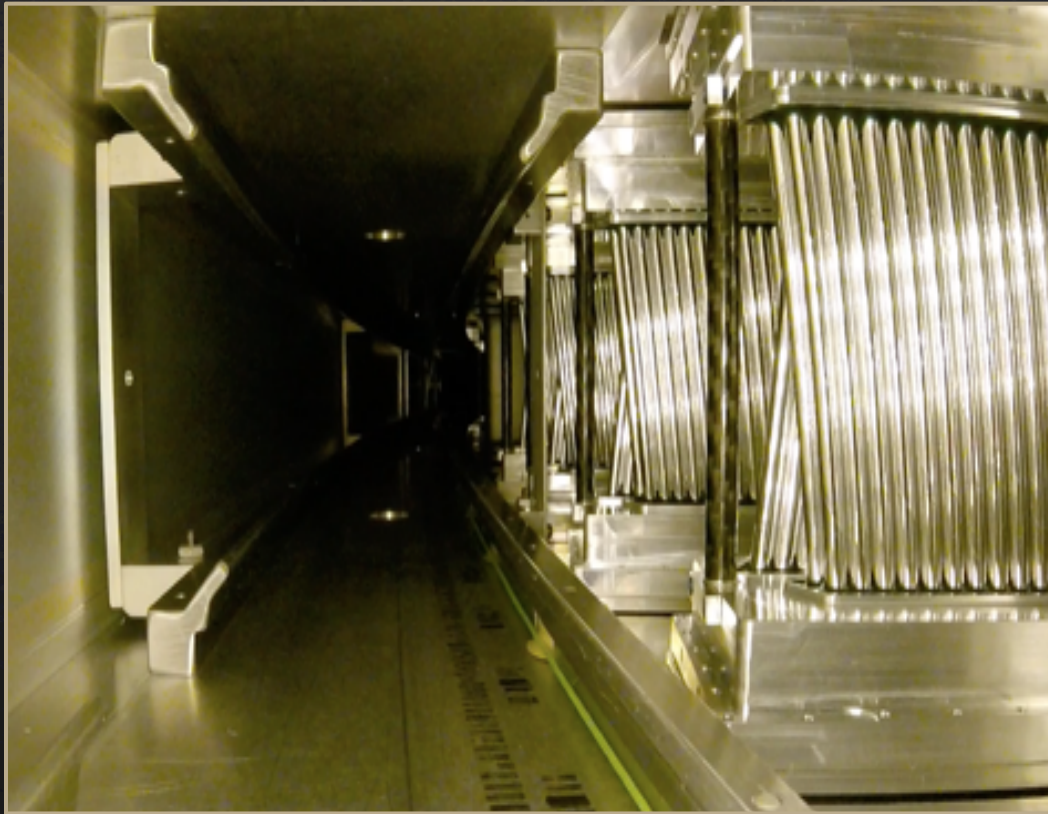


Straw trackers

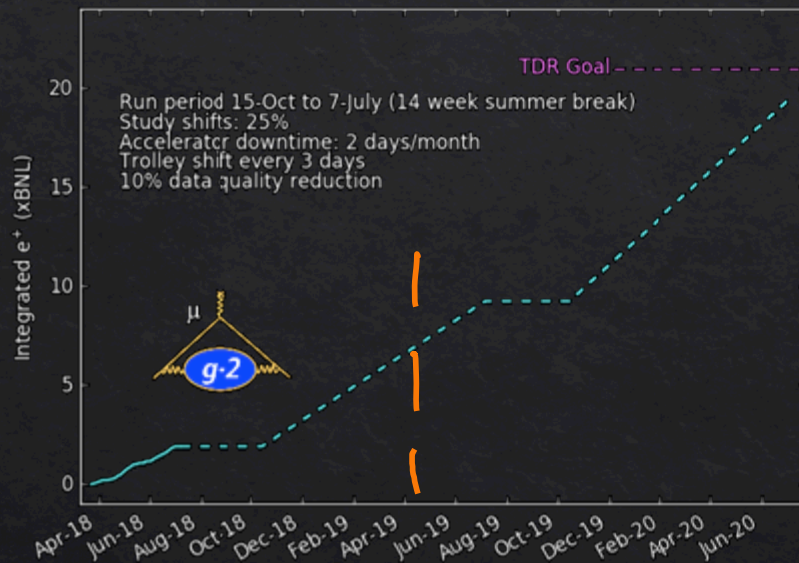


**100 μm radial
resolution achieved**

Tracking detectors

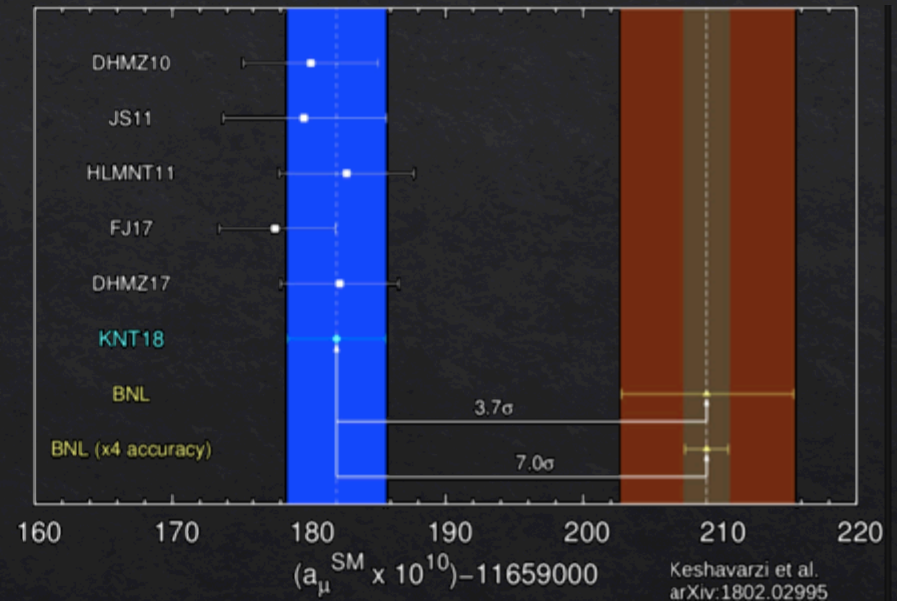


Data Collection



FIRST RESULTS

Significance gap



... a lepton-flavour violating dark photon..?

...a model with a large muon EDM..?

arXiv:1807.1148

Combined explanations of $(g-2)_\mu$, a_μ and implications for a large muon EDM

Andreas Crivellin, Martin Hoferichter, Philipp Schmidt-Wellenburg

Fortunately UK using $g-2$ to make a 1-2 order magnitude improvement in μ EDM!

Explaining electron and muon anomalies

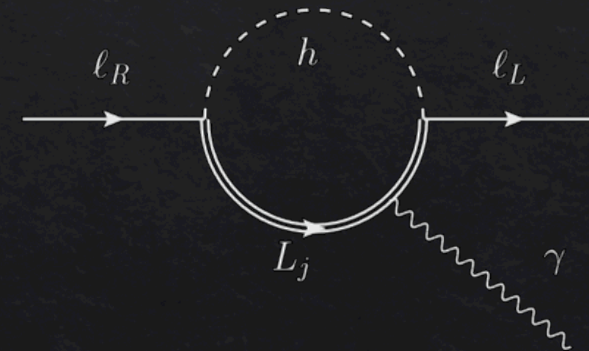
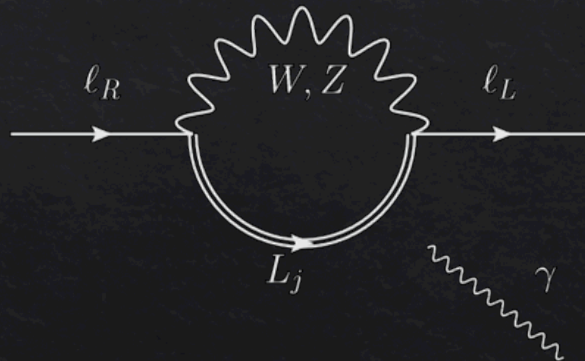
New

Crivellin + Hoferichter + Schmidt-Wellenburg, arXiv:1807.11484

'Combined explanation of $(g-2)_{\mu,e}$ and implications for a large muon EDM'

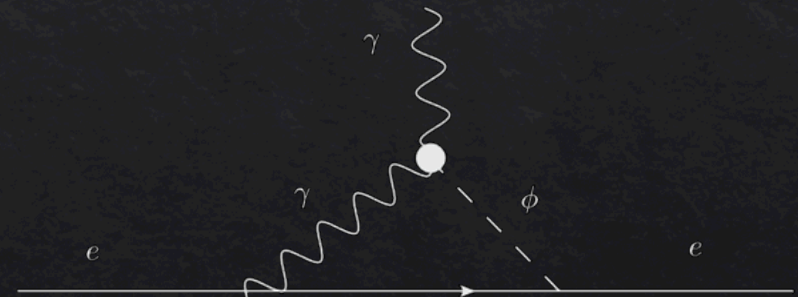
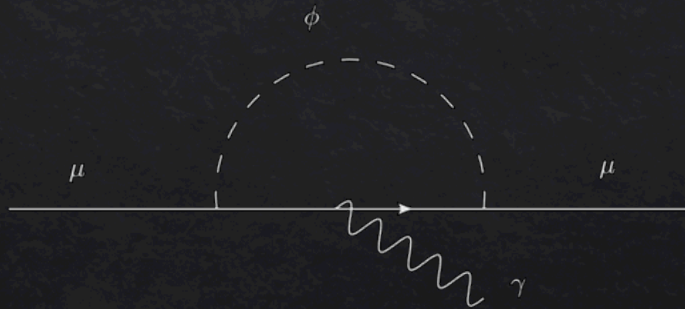
discuss UV complete scenarios with vector-like fermions (not minimally flavor violating) which solve both puzzles and at the same time give sizeable muon EDM contributions,

$|d_\mu| \sim 10^{-23}-10^{-21}$,
but escaping
constraints from
 $\mu \rightarrow e \gamma$.

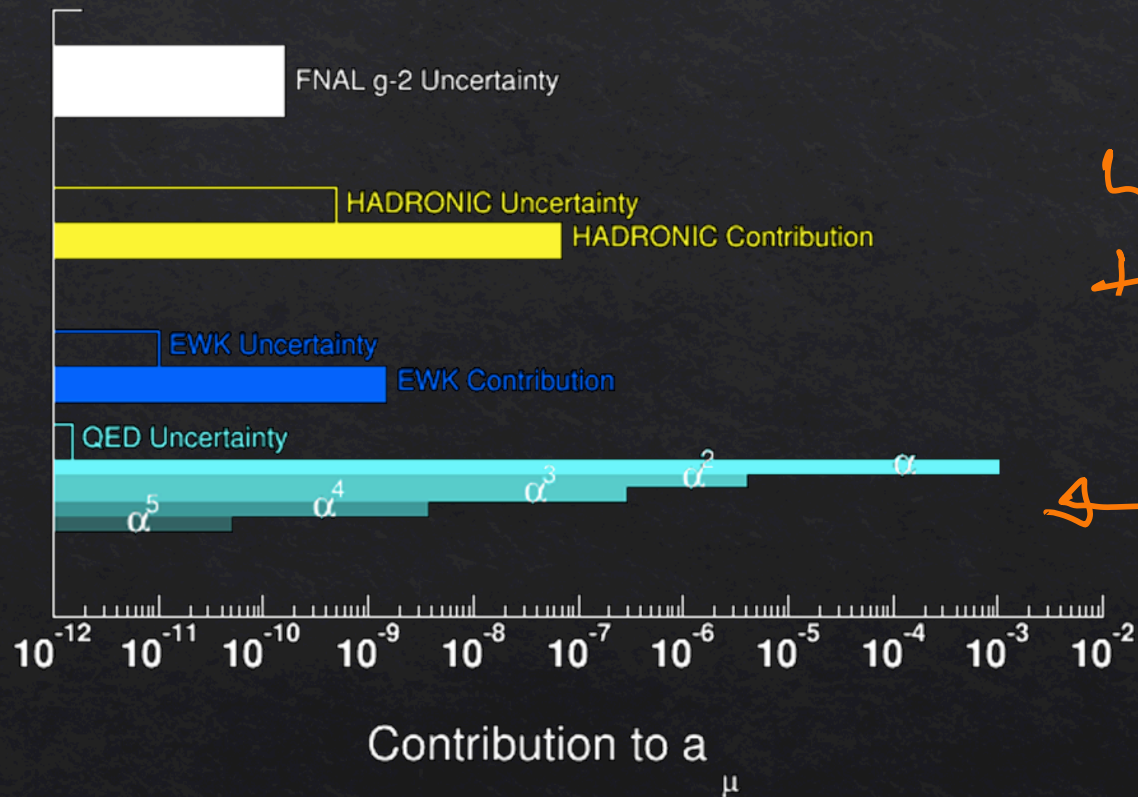


Explaining electron and muon anomalies

- **Davoudiasl + Marciano, 'A Tale of Two Anomalies'**, PRD96(2018)096018
use one singlet real scalar ϕ with mass ~ 250 -1000 MeV and couplings $\sim 10^{-3}$ and $\sim 10^{-4}$ for μ and e , in one- and two-loop diagrams



Summary of g-2 SM corrections/uncertainties



LATTICE
+ ...

→ LIFETIMES
OF WORK!

Purcell and Ramsey

On the Possibility of Electric Dipole Moments for Elementary Particles and Nuclei

E. M. PURCELL AND N. F. RAMSEY
Department of Physics, Harvard University, Cambridge, Massachusetts
 April 27, 1950

IT is generally assumed on the basis of some suggestive theoretical symmetry arguments¹ that nuclei and elementary particles can have no electric dipole moments. It is the purpose of this note to point out that although these theoretical arguments are valid when applied to molecular and atomic moments whose electromagnetic origin is well understood, their extension to nuclei and elementary particles rests on assumptions not yet tested.

One form of argument against the possibility of an electric dipole moment of a nucleon or similar particle is that the dipole's orientation must be completely specified by the orientation of the angular momentum which, however, is an axial vector specifying a direction of circulation, not a direction of displacement as would be required to obtain an electric dipole moment from electrical charges. On the other hand, if the nucleon should spend part of its time asymmetrically dissociated into opposite magnetic poles of the type that Dirac³ has shown to be theoretically possible, a circulation of these magnetic poles could give rise to an electric dipole moment. This argument, however, may be countered by noting that this electric dipole would be a polar vector, being the product of the angular momentum (an axial vector) and the magnetic pole strength, which is a pseudoscalar in conformity with the usual convention that electric charge is a simple scalar.

The argument against electric dipoles, in another form, raises directly the question of parity. A nucleus with an electric dipole moment would show an asymmetry between left- and right-handed coordinate systems; in one system the dipole moment would be parallel to the angular momentum and in the other, antiparallel. But there is no compelling reason for excluding this possibility. It would not be the only asymmetry of particles of ordinary experience, which already exhibit conspicuous asymmetry in respect to electric charge. Although magnetic poles were used above as an illustration of a particular mechanism by which a nuclear electric dipole could arise, this is, of course, not the only possibility.

The question of the possible existence of an electric dipole moment of a nucleus or of an elementary particle in view of the above becomes a purely experimental matter. The evidence from most past experiments on molecules, atoms, nucleons, and elementary particles is not as conclusive as one might suppose. Most past experiments are in fact very insensitive to the effects of a nuclear electric dipole, because of the smallness of the electric field at the position of a charged nucleus or the antisymmetric nature of the electric dipole potential. We have analyzed a number of experiments including conversion of ortho- to para-hydrogen, depolarization of neutron beams, ionization by neutrons, relaxation times of nuclei in liquids, nuclear scattering of neutrons, hyperfine structure studies, the Lamb-shift-type experiment and the scattering of neutrons on the interaction of electrons and neutrons. Non-scattering experiments on charged nuclei are particularly insensitive to the existence of an electric dipole moment and even the most favorable would not have revealed an electric dipole moment smaller than the charge of the electron multiplied by a distance D less than 10^{-19} cm. The scattering experiments^{1,2} to detect an electron-neutron interaction are by far the most sensitive; the results of Havens, Rabi, and Rainwater³ would correspond to a D of 3×10^{-18} cm if they were due to an electric dipole moment.

We are now undertaking, in collaboration with Mr. James H. Smith, an experiment which should directly measure the electric dipole moment of the neutron if it has a value of D of approximately the above magnitude. The experiment will utilize a neutron beam magnetic resonance⁴ apparatus of high resolution⁵ to detect a possible shift of the neutron precession frequency upon the application of a strong electric field.

The authors wish to thank Mr. Smith for suggesting an important correction to our original calculation on the neutron-electron interaction experiment.

¹ A typical argument is given by H. A. Bethe, *Elementary Nuclear Theory* (John Wiley and Sons, Inc., New York).

² P. A. M. Dirac, *Phys. Rev.* **74**, 817 (1948).

* Havens, Rabi, and Rainwater, *Phys. Rev.* **72**, 634 (1947).

⁴ E. Fermi and L. Marshall, Phys. Rev. **72**, 1139 (1947).
⁵ L. W. Alvarez and F. Bloch, Phys. Rev. **57**, 111 (1940).
⁶ N. F. Ramsey, Phys. Rev. **76**, 996 (1949).

¹ N. F. Ramsey, Phys. Rev. **70**, 990 (1949).

Supernovae®

L. B. BORST

Brookhaven National Laboratory, Upton, Long Isl.
April 27, 1950

SUPERNOVAE of type I are character

SUPERNOVAE of type I are characterized by an intensity maximum of 20 to 30 days, an exponential tail to the light curve of half-life 1.5 ± 0.0012 magnitudes per day;¹ (c) a maximum emission of nearly 10^{46} ergs;² (d) a maximum hydrogen content expanding at a rate of 10^4 km/sec; (e) a maximum radiating 10^{36} ergs/sec. visible

These characteristics may support the proposed mechanism. The sun, e.g., $15M_{\odot}$, undergoes a contraction of its hydrogen. As the temperature will rise up of 2 to 3×10^9 °C between alpha-

This re
 rapid r
 by m
 is a h

HARVARD UN
 James H. Sm
 is shown
 trace of s
 trans.
 Harv

where Z is the number of particles, α the alpha particle, τ the reaction threshold; & k is the temperature. It may be noted that the volume increases as the square of the radius exponentially with the temperature. The star may collapse under conditions of gravitation in a time approaching

The reaction will proceed until there are sufficient quantities of the reaction products to produce the equilibrium constant. The expression at equilibrium may be given by

$$K = [\text{Be}^7]/n/[\text{He}^7]$$

where the entries denote atomic concentrations per unit. Since neutrons will be absorbed rapidly in a system co-

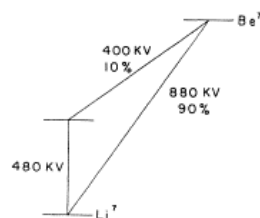
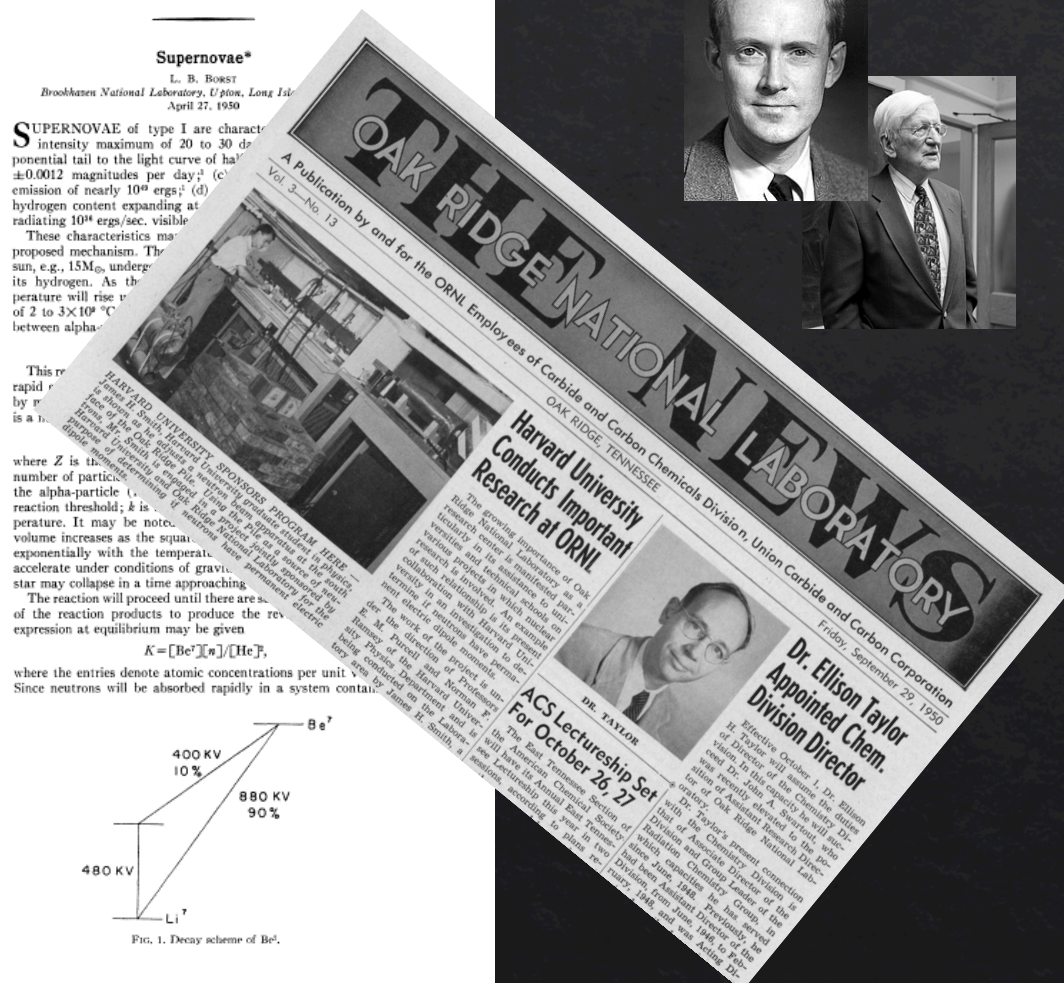
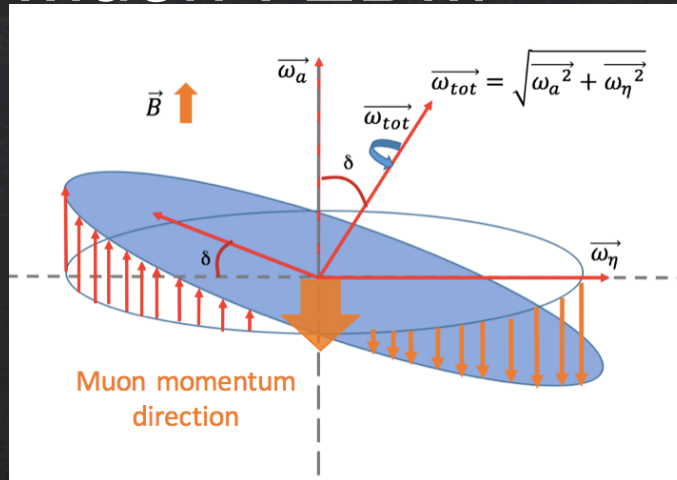


FIG. 1. Decay scheme of Be^3

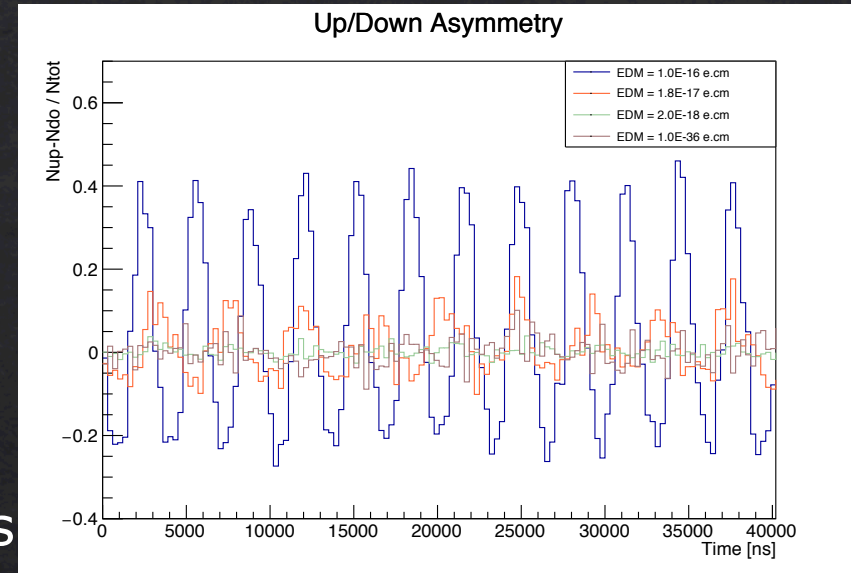


Muon : EDM



O(1M) events in trackers
(few weeks)
--> sensitivity at 10^{-19}
[BNL]

Expect several billion
events in the
trackers and so reach
 10^{-21}



- Precession plane tilts towards center of ring
- Causes an increase in muon precession frequency
- Oscillation is 90° out of phase with the a_μ oscillation

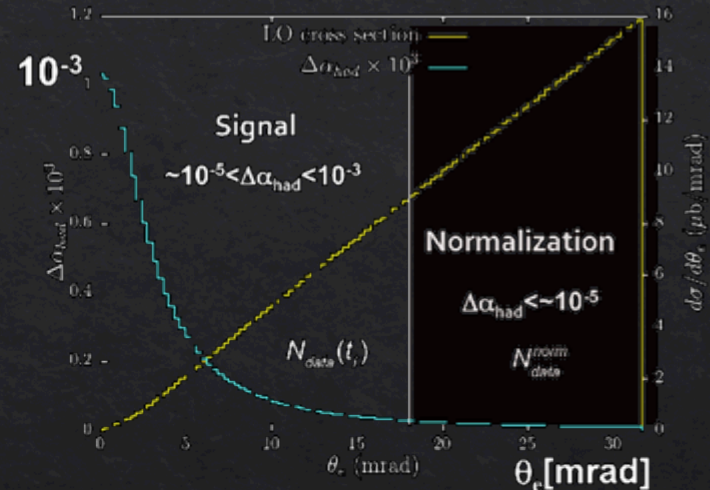
MuonE

WE COULDN'T GET
THIS FUNDED (YET.)

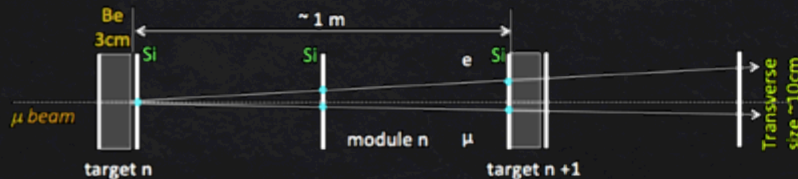
Theory limited by hadronic LO corrections, a_μ^{HNLO}
Traditional calculation from $ee \rightarrow \text{hadrons}$
→ need x2 improvement to keep up with g-2

MUonE will measure space-like region:
→ scattering of high energy mu (150 GeV) on e

$$a_\mu^{\text{HLO}} = \frac{\alpha}{\pi} \int_0^1 (1-x) \Delta\alpha_{\text{had}}(t(x)) dx$$



Up to 20 Be targets + Si detectors
downstream calorimeters + muon PID



Schedule:

- 2017: test beam at CERN H8 Beam Line
- 2019: LOI to SPSC
- 2020/1: construction & installation
- 2022/4: (after LHC LS2) start data taking

Outlook/Conclusions

- ◇ ~1.4 x BNL dataset taken during run 1 after quality cuts
- ◇ Liverpool-built trackers crucial component of measurement
- ◇ Currently analysing - hardware and software blinded in both frequency and field measurements
- ◇ Expect unblinding in early summer 2019

- ◇ Run 2 underway - aiming for 3 times more data this year
- ◇ Additional tracking station being added summer 2019

Scientific breakthrough could be as simple as measuring the wobble of a muon

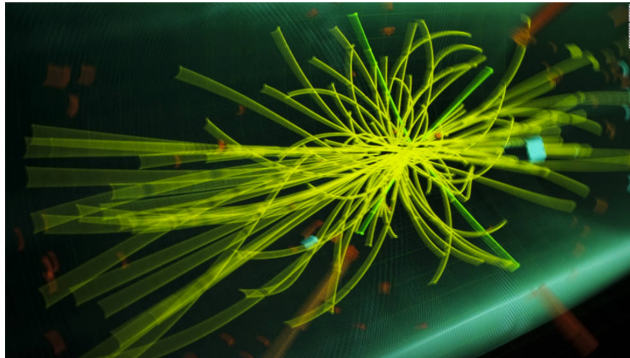
By Don Lincoln

🕒 Updated 1648 GMT (0048 HKT) February 13, 2018



Fermilab  @Fermilab · 3h

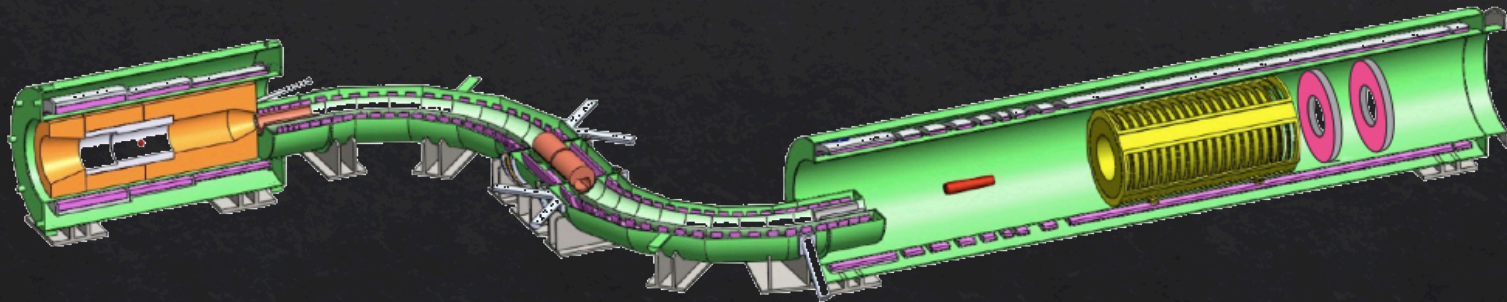
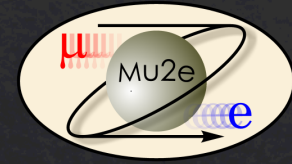
"If I were to put my money on something that would signal new physics, it's the g-2 experiment at Fermilab."



After some 21 years of $(g-2)$ measurements on the muon at CERN, a great deal of territory has been brought within the civilized domain of QED theory, and the precision of the most recent result defines the limits within which that domain is secure against any future theoretical excursions. As we have stressed above, any modification to the photon propagator or new coupling common to both muons and electrons would imply a perturbation of a_μ by a factor $(m_\mu/m_e)^2$ larger than for a_e . Thus in the absence of possible coupling particular to the electron, the present muon result ensures that a_e is a “pure QED quantity” down to the level of three parts in 10^{10} .

However, all the effort expended in this activity has brought us no nearer to understanding the mystery of the muon mass. No evidence of a special coupling to the muon has been found. On more general observational grounds it is known that the neutrinos distinguish between the charged leptons. The neutrinos clearly know the difference in the sense that the electron, the muon and the new lepton of mass $1.8 \text{ GeV}/c^2$, discovered by Perl et al. [68], each have their own associated neutral massless fermion; perhaps it is in this area that enquiry should be made for an answer to the charged lepton mass splittings.

For the present, however, the thread which has linked many experimenters together in the common cause of measuring the muon $(g-2)$ factor at CERN is now broken and those who have shared this experience have gone their separate ways. It remains to be seen whether or not future refinement of the theory of the weak, electromagnetic, and strong interactions will call for the discerning scrutiny of further measurements of even greater precision.



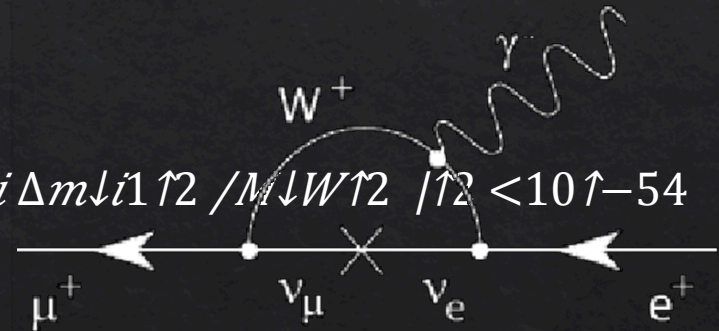
Lepton Flavour Violation

15 year programme ...

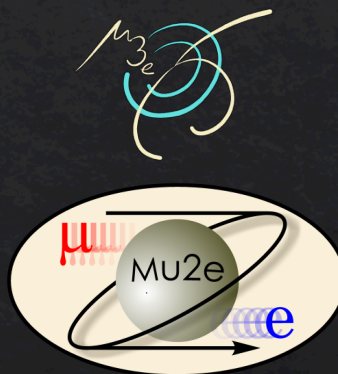
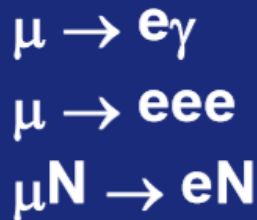
CFLV

In SM

$$Br(\mu \rightarrow e \gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i} U_{ei}^* \right|^2 \frac{\Delta m_{i1}^2}{M_W^2} < 10^{-54}$$



Other possible decays



point.¹ I
centric
immeas
The hig
gas from
In vie
appears
could be

* The w
Research
sota by gr
1 Sep 10
and Sons
Pulford 15
Rev. 70, 5
1.1 G. 15
Nier, Phys
1.1 G. 15
759 (1947)
1.1 G. 15
Rev. 73, 2



Search for Gamma-Radiation in the 2.2-Microsecond Meson Decay Process

E. P. HINCKS AND B. PONTICORVO
National Research Council, Chalk River Laboratory,
Chalk River, Ontario, Canada
December 9, 1947

THE meson decay process which is identified by a mean life of 2.2 microseconds¹ has been usually thought of as consisting of the emission of an electron and a single neutrino, as suggested by the well-known Yukawa explanation of the ordinary beta-process in nuclei. However, the Yukawa theory is at variance with the results of the experiment of Conversi, Pancini, and Piccioni,² and since there remains no strong justification for the electron-neutrino hypothesis,³ a direct experiment to test an alternative hypothesis—that the decay process consists of the emission of an electron and a photon, each of about 50 Mev—has been performed.

The apparatus, illustrated in Fig. 1, consists of three rows of Geiger-Müller counters, A, B, and C, each having an effective area of approximately 38 cm² × 20 cm. Above A there are 15 cm of lead, and between A and B, 1.5 cm of lead. Mesons traversing A and B, and stopped in a graphite absorber 38 cm × 19 cm × 5 cm thick, produce decay electrons which may be detected in either B or C. Decay photons, if present, could also be detected in B or C, whose efficiency for gamma-radiation was increased by introducing 2.1 mm of lead between the graphite and both B and C. The twofold function of B—first, detection of the passage of a meson by a coincidence with A (event “(A, B)”), and second, detection of a decay electron (or photon) following “(A, B)” —is permitted by the circuit design. Although one of the eight counters of B (that through which the meson passed) is insensitive to the decay particle because of the long counter dead time, the use of B in this manner allows an advantageous geometry. The outputs of the three rows are mixed by circuits whose function is schematically shown in the diagram, and the following delayed events are finally recorded:

RS TO THE EDITOR

257

TABLE I. Delayed single and coincidence counting rates.

	(B) _{del} (Counts/hr.)	(C) _{del} (Counts/hr.)	(B) _{del} +(C) _{del} (Counts/hr.)	(B, C) _{del} (Counts/hr.)
With graphite plus lead—108.5 hours of observation	11.98±0.34	12.26±0.34	24.19±0.48	0.21±0.05
Without graphite plus lead—97.3 hours of observation	6.48±0.29	6.64±0.25	11.12±0.38	0.49±0.08
Net effect due to decay electrons from graphite plus lead	5.45±0.45	7.82±0.42	13.07±0.62	

1. “(B)_{del}” discharges of B occurring between 0.6 and 5.3 microseconds after “(A, B).”
2. “(C)_{del}” discharges of C occurring between 0.6 and 5.3 microseconds after “(A, B).”
3. “(B, C)_{del}” coincidences of B and C occurring between 0.6 and 5.3 microseconds after “(A, B).”

Runs were made with and without the graphite plus lead between B and C, and the results are presented in Table I. Other runs with graphite only, with lead only, and with other thicknesses of graphite and lead, were performed and these will be reported in a more complete account of the experiment. Check runs with a 1.6- to 6.3-microsecond delay gave results consistent with a mean life of 2.2 microseconds.

The observed rate (B, C)_{del} could be due to the following causes:

- (i) genuine electron-photon coincidences from the meson decay,
- (ii) single decay electrons which traverse both B and C,
- (iii) casual events.

The casual rate (iii), which is due essentially to mesons traversing B and C between 0.6 and 5.3 microseconds after an event “(A, B),” has been estimated from the measured double and triple coincidence rates and from the characteristics of the circuits to be 0.22±0.02 counts per hour. It is independent of the presence or absence of graphite plus lead. Effect (ii) should be detected only in absence of graphite plus lead, since otherwise the total thickness of material between B and C is of the order of the expected range of the electrons. We observe, in fact, that (B, C)_{del} increases appreciably when the graphite plus lead is removed. The presence of this effect was verified by a sub-

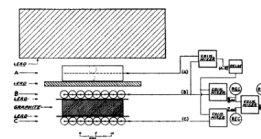
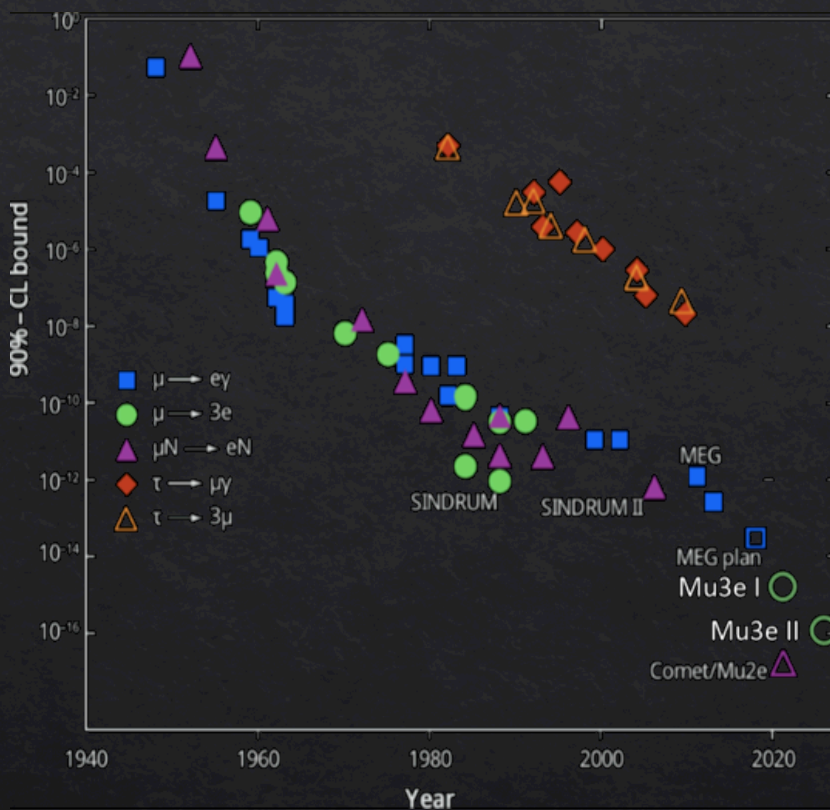


FIG. 1. Arrangement of apparatus.



Other CLFV searches

At e^+e^- or pp GPDs:

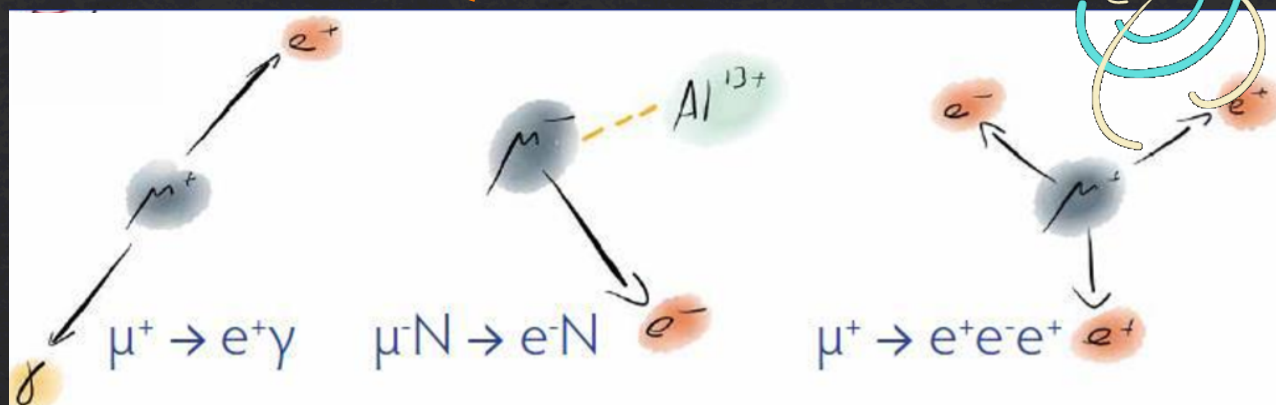
- $Z \rightarrow e\mu, \tau\mu, e\tau,$
- $H \rightarrow e\mu, \tau\mu, e\tau$

In flavour experiments:

- LFV in hadron decays

↑ Now

MEG

 $\mu 2e$
(COMET) $\mu 3e$ 

back-to-back electron
and photon

$$E_\gamma = E_e = \frac{1}{2} m_\mu$$

Muon decay from
muonic atom.

Monochromatic
electron

$$E_e = m_\mu - E_{\text{binding}} - E_{\text{recoil}}$$

3 co-planar electrons

$$\Sigma P_e = 0, \Sigma E_e = m_\mu$$

Radiative decay:

$$\mu \rightarrow e\nu\nu\gamma$$

Accidental backgrounds:

$\mu \rightarrow e\nu\nu$ + conversion or
Bhabha electrons

Muon Decay in orbit

beam related:

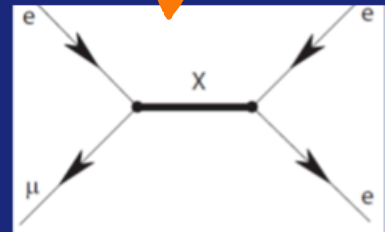
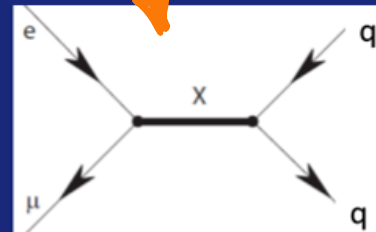
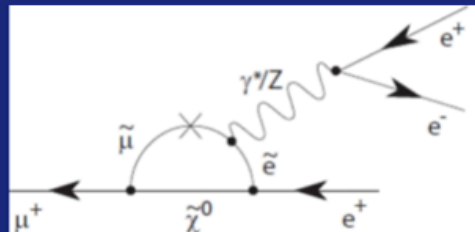
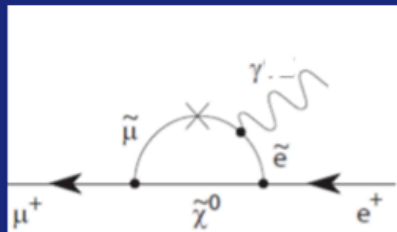
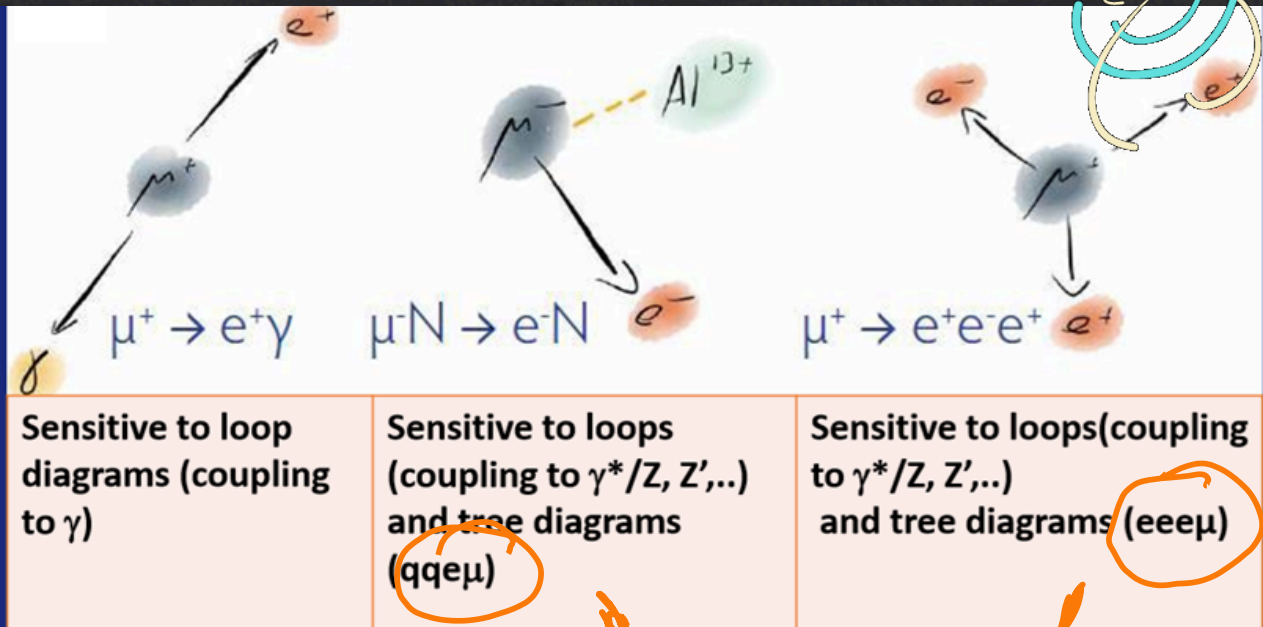
prompt antiprotons,
pions,...

Radiative decay

$$(\mu \rightarrow eee\nu\nu);$$

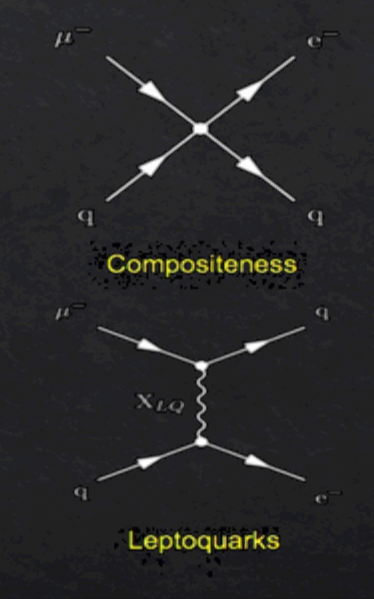
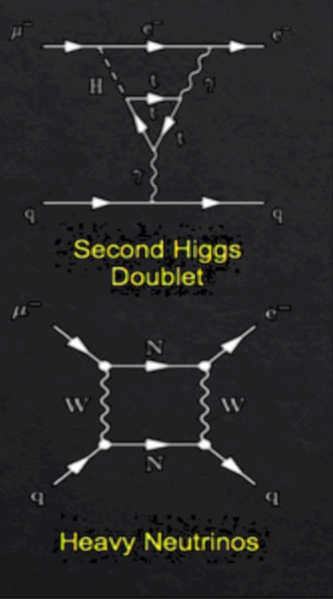
Accidental backgrounds
 $\mu \rightarrow e\nu\nu$ + conversion or
Bhabha pairs

LIVERPOOL DOES BOTH

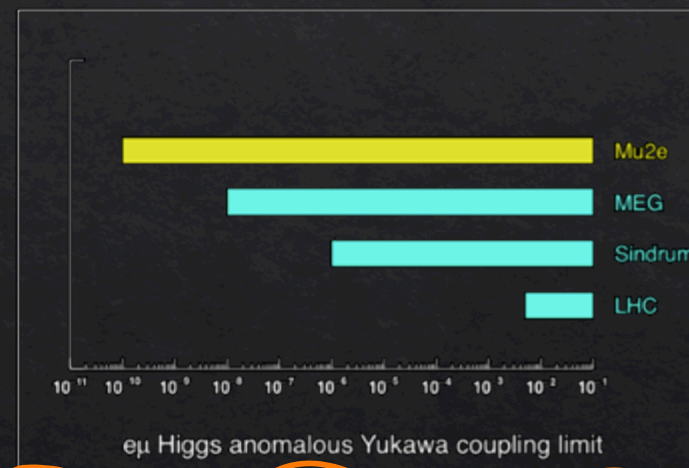


Many BSM models include charged lepton flavour violation

- leptoquarks, compositeness, Higgs doublets, heavy neutrinos...
...or invoke it for leptogenesis of matter-antimatter asymmetry



Probe LQ masses up to 300 TeV
cf 1 (120) TeV at HL-LHC (LHCb)

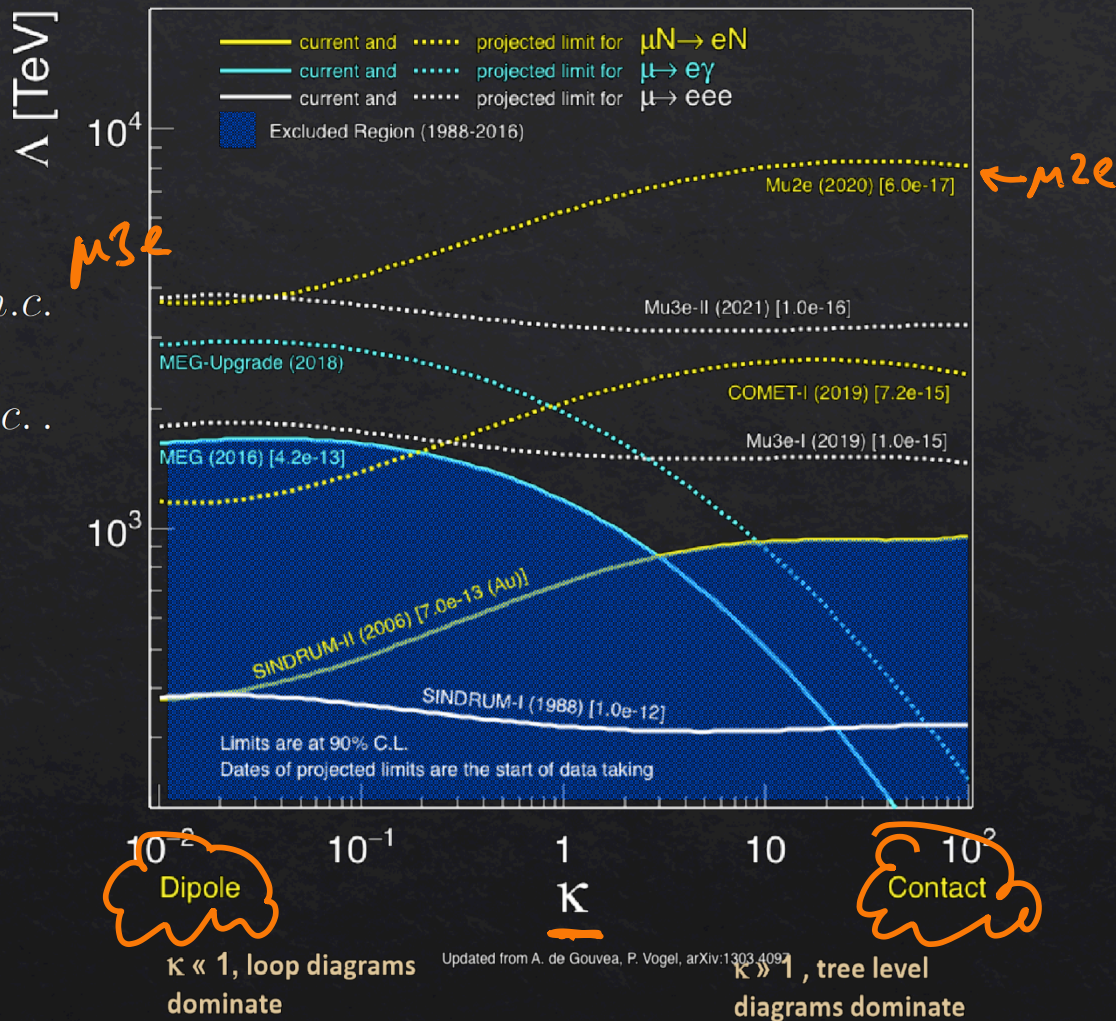


Sensitivity to flavour-violating Higgs couplings

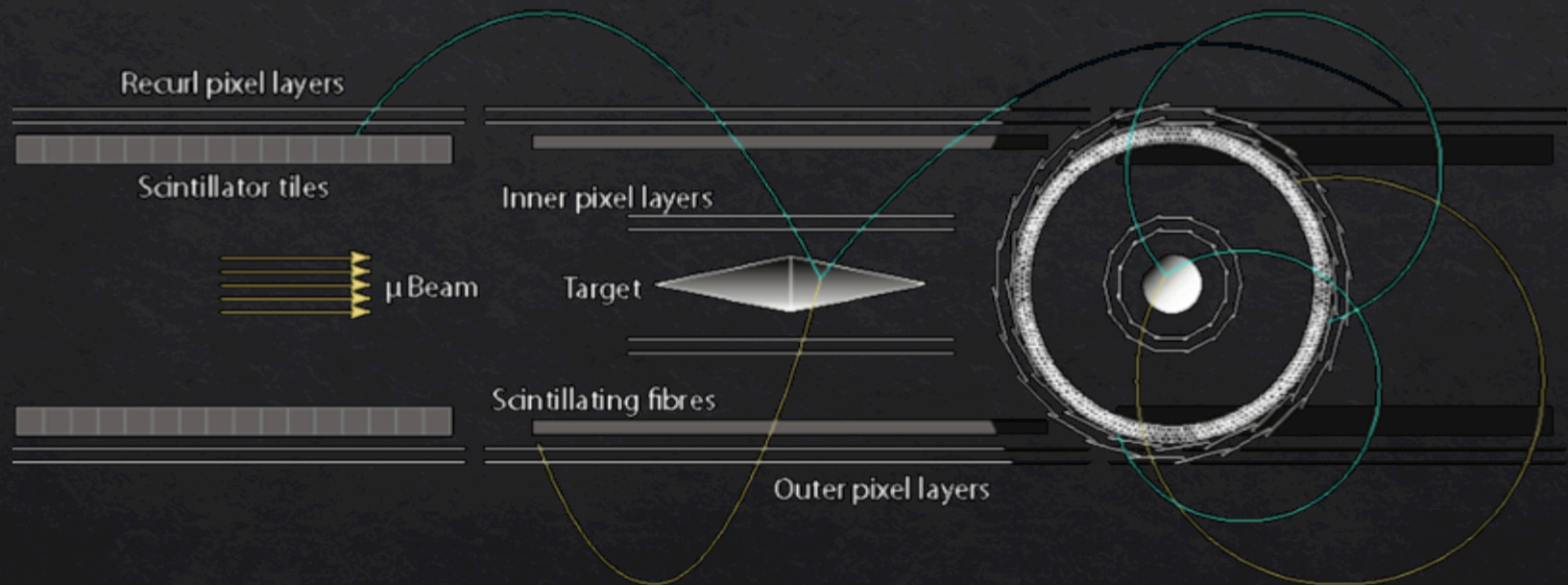


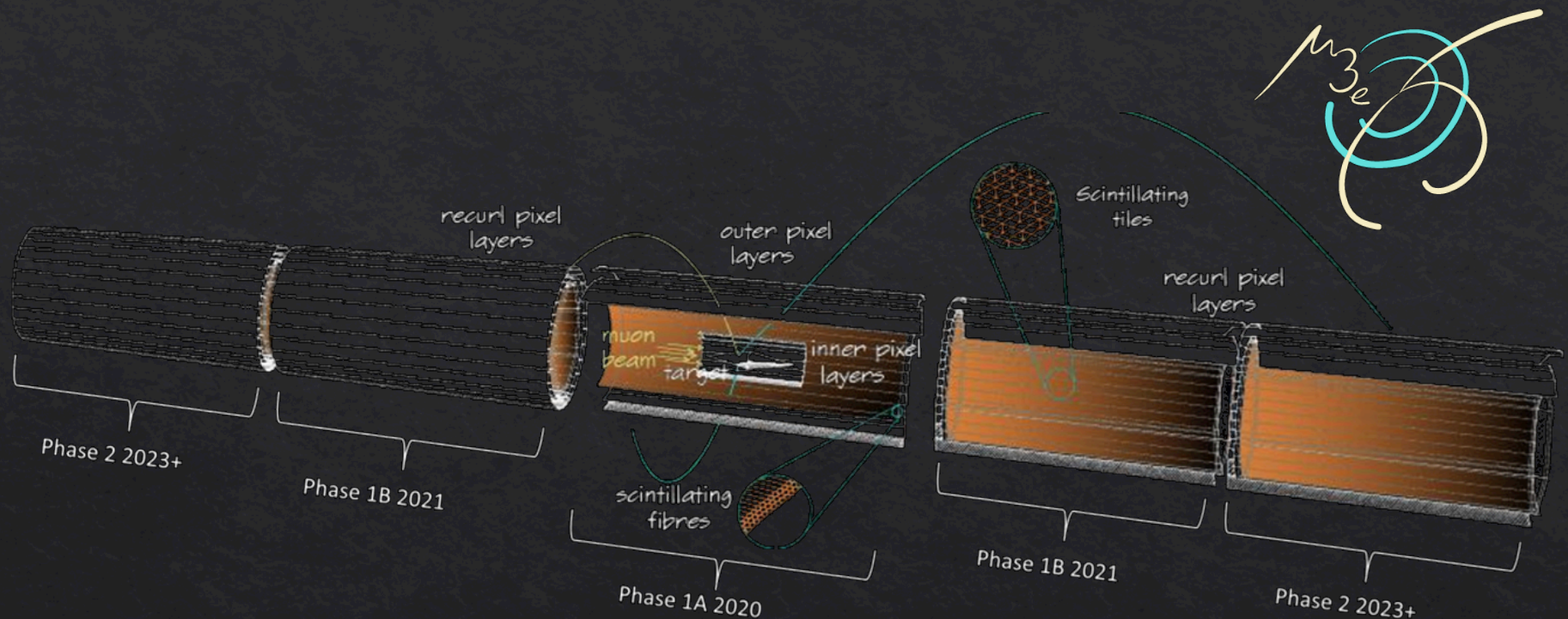
Update from de Gouvea & Vogel, Prog. in Part. and Nucl. Phys. 71 (2013).

$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + h.c. \\ \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{e} \gamma^\mu e) + h.c..$$



Mu3e





Phase 1A/1B: $BR(\mu \rightarrow eee) < 4 \times 10^{-15}$, ~3 years running at $10^8 \mu/s$ from PSI $\pi e5$ Compact Muon Beam Line

Phase 2: $BR(\mu \rightarrow eee) < 10^{-16}$, ~3 years running with extended acceptance detector at $2 \times 10^9 \mu/s$ from planned High Intensity Muon Beam (HIMB)

UK deliverables (Phase 1)

- Assembly off all outer pixel layers of the MuPix tracker
- Mu3e clock-and-control system for the time-slice based DAQ

HV-MAPS sensors

Adaptation from CMOS-MAPS using high-voltage compliant CMOS processes.

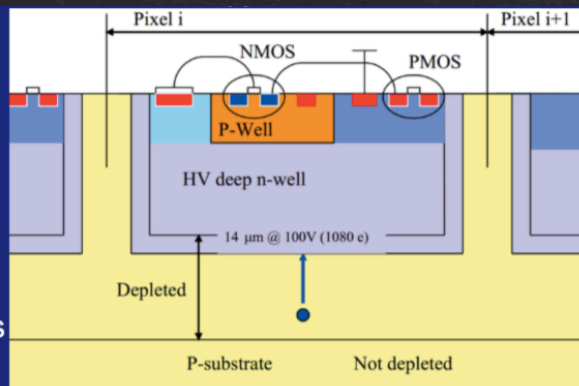
- Specific is deep N-well that collects charge and includes analogue and digital circuits. (no parasitic collection)
- N-well is biased to $> 80 - 200$ V giving $10 - 30$ μm depletion in bulk.
- High signal and fast charge collection, *combining compactness of CMOS with performance of hybrid planar silicon sensors.*

Critical properties for Mu3e:

- Sensors can be thinned to 50 μm without signal loss.
- Sensors can operate in a high rate environment ($\sigma(t) < 25$ ns)

Mu3e is the first PP experiment to employ HV-MAPS in a tracker

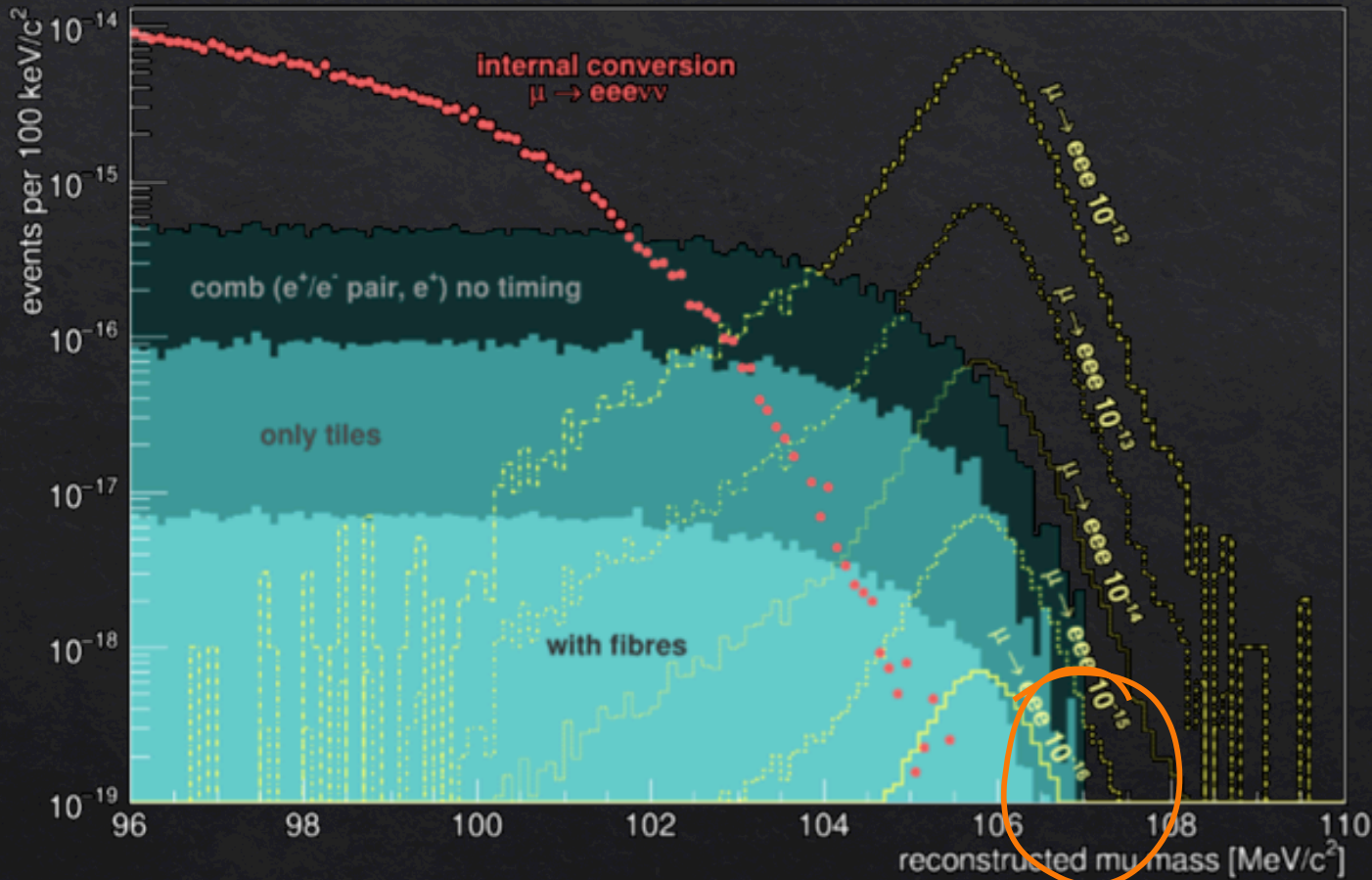
Mu3e would not be possible without this new technology! (sensitivity $\sim (X/X_0)^3$)



GREAT TECHNOLOGY
— MORE ADVANCED
THAN LHC

Mu3e

Events per stopped μ^+



TAIL END OF SPECTRA

One CLFV interaction in 10^{16} muon decays is like...
looking for one specific grain of sand




CLFV

- ◇ If $g-2$ confirmed CLFV checks whether NP has a lepton mixing angle
- ◇ Particular sensitivity to 4 fermion contact terms
- ◇ Mu3e can search for dark photons
- ◇ UK on mu2e (with upgrade options) and mu3e

Other muon experiments

- ◇ muEDM – parasitically at g-2 and upgrade
- ◇ Dedicated experiment for muEDM
- ◇ J-PARC experiments
- ◇ Proton radius
- ◇

A few comments

- ◇ There is a great richness and elegance in these experiments
- ◇ Creativity and training 
- ◇ Whatever the outcome of ES this diversity should be preserved
- ◇ Enthusiasm of students and teams working on this
- ◇ Cheap (UK resource CLFV minimal)
- ◇ But major contributions to all experiments so far
 - ◇ g-2: Theory, Hardware, Analysis, Leadership
 - ◇ Mu3e: building vertex detector
 - ◇ Mu2e: STM

Moments ...

"If you enjoy doing difficult experiments, you can do them, but it is a waste of time and effort because the result is already known" :

Pauli



*"No experiment is so dumb, that it should not be tried" : **Gerlach***

*"the Muon obeys QED.
g-2 is correct to 0.5%.
In my opinion, it will be
right to any accuracy. So it's not worth
doing the experiment"*

**Head of CERN Theory at time of CERN
EDMs**

*"would you like to predict
the result ?" : **F. Farley FRS***

Muons

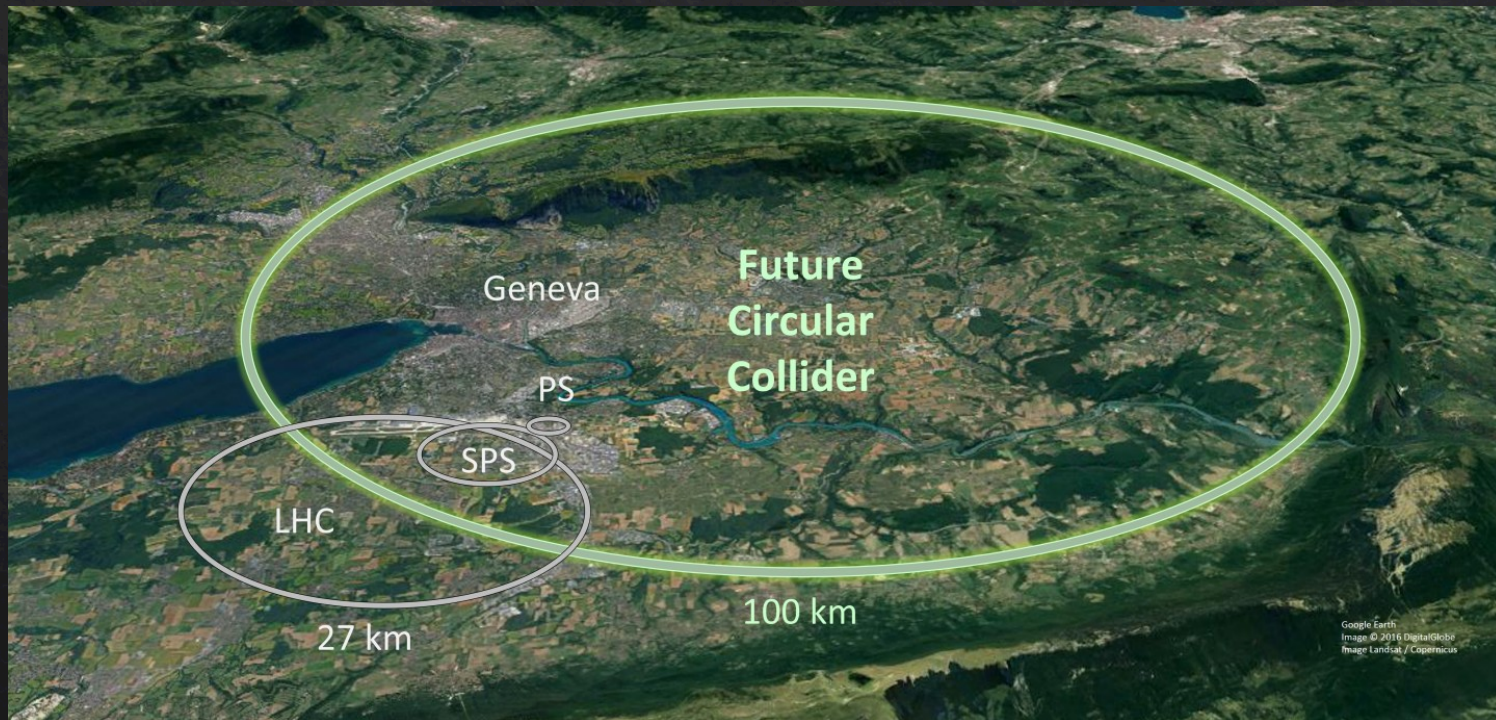
- Why Liverpool and UK joined
- Part of UK strategy ...
- Partial snapshot
- Missed out muons at LHCb
 - Universality etc



Physics Beyond the LHC is exciting!

Access to PeV in the next decade

Can we do physics with loops? Existential question....



Theory

CRITICAL TIME FOR SUBJECT

— LEADERSHIP & IDEAS

BUILD 95 km AND INVEST IN THEORY?
(AND PBC?) !