

### Hartmut Grote (PI) on behalf of the QI consortium



Quantum Technologies for Fundamental Physics UK HEP forum 10/Nov/2020



### Our consortium



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### Two Themes

### Theme 1: The nature of dark matter (UKRI/STFC C4)

### Theme 2: The nature of space-time (UKRI/STFC C3)

## Frontier Physics At the limits of energy, space and time



- Baryonic dark matter inconsistent with CMB and micro-lensing
- Non-baryonic dark matter

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- WIMPs: the detectors will face neutrino background soon
- Axions (and scalar and vector fields): large unexplored parameter space

Can be tested using interferometric measurements

• Sterile neutrinos: new experiments coming soon





## Work packages

Work package 1: Axions in the galactic halo
axions with masses from 10<sup>-16</sup> eV up to 10<sup>-8</sup> eV

#### Work package 2: Light-shining-through-wall

• transition edge sensor with background <10<sup>-6</sup>/s







## WP1: Interferometry Haloscope

• Axion field behaves classically

$$\mathbf{a}(\mathbf{t}) = \mathbf{a}_0 \cos(\Omega_a(t) + \Phi(t))$$

- Phase difference between left- and right-polarized light
- Can observe the phase interferometrically

Phys. Rev. D 98, 035021, Phys. Rev. Lett. 121, 161301, Phys. Rev. D 100, 023548



## WP1: Interferometry Haloscope

- Resonate the pump and signal fields in the main cavity
- Choose the axion mass by tuning the auxiliary cavity
- Experimental support from MIT (Prof. M. Evans)





- The readout is limited by quantum shot noise
- Sources of squeezed states (developed by AEI Hannover) will reduce the noise (increase SNR) by a factor of 3





## WP1: Sensitivity

- Table top setups can provide new limits
- The layout can be potentially scaled to km lengths



## Light-shining-through-wall (LSW)

Lab "Light-Shining Through wall"





# ALPS @ DESY, Germany



arXiv:2009.14294

- International collaboration (DESY, U Florida, AEI Hannover, Mainz, Cardiff)
- Under construction in former HERA accelerator tunnel at DESY
- WP2 contribution: improved TES and ALPS commissioning





- Transition Edge Sensor: a micro-calorimeter
- Heat absorption brings superconductor to resistance increase
- Measure resulting current with SQUID
- Energy resolution to discriminate large green photon flux

# TES readout for ALPS

NIST chips (W, Tc = 170 mK) with optical resonator and metallic mirrors reaching 98% quantum efficiency for 1064 nm





- Provide improved TES at 1064nm
- 30mK dilution refrigerator platform
- Next generation TES with SQUID readout (NIST/Magnicon)
- Target QE > 80%
- Target dark count < 10<sup>-6</sup>/s
- Novel muon veto scheme with second on-chip detector
- Collaboration with NIST and PTB

#### **References:**

R. H. Hadfield Nat. Photon 3 696 (2009)A. Lita et al. Optics Express 16 3032 (2008)A. Lita et al. Proc. SPIE 7681 (2010)



## WP1/WP2: Sensitivity

- "JURA" could be a next generation LSW experiment
- Km-long arms and higher power



## ALPS: final magnet installation



# Theme 2: Quantum aspects of space-time

- Quantum mechanics and general relativity passed all experimental tests to date.
  - Google's Sycamore quantum computer
  - Observation of gravitational waves by LIGO / Virgo
- How can gravity be united with the other fundamental forces?
- We pursuit the directions of search for quantization of space-time and quantum optomechanical tests

Phys. Rev. Lett. 110, 170401, Phys. Rev. A 101, 063804, Phys. Rev. D 100, 066020



## Work packages

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#### Work package 2: Light-shining-through-the-wall

• transition edge sensor with background <10<sup>-6</sup>/s

### Work package 3: Quantisation of space-time

• sensitivity of 2x10<sup>-19</sup> m/rt(Hz) above 1 MHz

Work package 4: Semiclassical gravity

• confirm or rule out with 5 sigma precision



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## WP3: Quantization of space-time



Space-time foam?

- Quantization of space-time at Planck scale of 10<sup>-35</sup> m?
- Holographic principle may make this accessible to interferometry
- Flexible table-top to test different predictions

## QI

## Before: Fermilab "Holometer"



PRL 117, 111102 (2016) CQG 34, 6 (2017)

### WP3: co-located table-top interferometers



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- 10 kW in interferometer
- Output modecleaner
- 500MHz data sampling
- Squeezed light
- Can test new squeezing configurations
- Theoretical support Fermilab





## Squeezing shot noise

Two world records with squeezers from our AEI (Hannover, Germany) partners:







# Aiming for World-best table-top sensitivity in interferometry



A side remark:

Also sensitive to very high frequency gravitational waves and scalar field dark matter



## Scalar field DM with interferometry

- Scalar field DM (variation of fundamental constants) is searched for with atmic clocks and atom interferometry (in QTFP: Atomic clock network, AION)
- Also possible with precision interferometry (GW detectors and others) PRR 1, 033187, 2019



# WP4: Quantum-optomechanics



- Measured phase is  $\phi_{out} = \phi_{in} + \phi_{shot} + \frac{8F}{\lambda}(x_{cl} + x_{rad})$
- x<sub>rad</sub> is the quantum back action signal from the laser field to the mechanical oscillator
- $x_{rad}$  may show gravity features in the spectrum
- Theoretical support from Prof Y. Chen (Caltech)



# WP4: Test of semi-classical gravity (Schrödinger-Newton equation)

- Two cryogenic silicon cavities to suppress laser noise  $\phi_{in}$
- Observe a splitting of the resonance due to semiclassical gravity
- The mirror should be made out of a crystal for a test of coherent self gravitation of atoms in a lattice



# WP4: Cryogenic silicon

- $\ensuremath{^\circ}$  Low thermal noises at 123 K and below 20 K
- Primary candidate for mirror substrate in future gravitational wave detectors
- The mirrors will be coated with custom coatings



## WP4: Vibration isolation

- The tabletop will be isolated using commercial dampers
- Motion of the cryostat plate will be actively stabilised
- Vibrations will be further passively filtered via a suspension system







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## Work packages

Two themes are united by quantum technologies

Optical cavities at their quantum limits

Sources of non-classical states of light

Transition edge sensors

Extreme-performance optical coatings





## Mirror Coating Requirements Driven by GW detectors

- Absorption <0.5 ppm
- Scatter <2 ppm
- ITM transmission:  $(5 \pm 0.25) \times 10^{-3}$
- ETM transmission: <10 ppm
- Mechanical loss: 3x10<sup>-5</sup> Goal (1x10<sup>-4</sup>)
- Uniformity +/- 0.5% over 34 cm







### Ion beam deposition:





### *Current state-of-the-art*

Industry best radio frequency (RF) ion beam deposition - second largest system in the world, hosted in SCAPA national laser/plasma facility in Strathclyde.



up to 62 cm diameter / 200 kg optics

### Next-generation

Electron-cyclotron resonance (ECR) ion beam deposition for optical coatings - first demonstration in the world.







### Quantum-enhanced Interferometry

