

Curry Parke-Taylor's

Atul Sharma

Mathematical Institute, Oxford

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[Based on work with T. Adamo, L. Mason]

Motivations

Why study curved backgrounds?

- ▶ Test domain of validity of reformulations of perturbative QFT.
- ▶ Many future applications:
 - Strong field QED/QCD
 - Gravitational waves
 - AdS/CFT
 - Cosmology
 - Black hole physics
 - Instanton backgrounds
 - \vdots

Amplitude for n -Gluon Scattering

Stephen J. Parke and T. R. Taylor

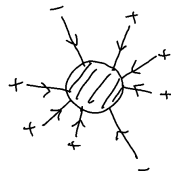
Fermi National Accelerator Laboratory, Batavia, Illinois 60510

(Received 17 March 1986)

A nontrivial squared helicity amplitude is given for the scattering of an arbitrary number of gluons to lowest order in the coupling constant and to leading order in the number of colors.

Maximally helicity-violating (MHV) tree-level gluon amplitude

$$\mathcal{A}(1^+ \dots r^- \dots s^- \dots n^+) = \frac{\langle r s \rangle^4}{\langle 1 2 \rangle \langle 2 3 \rangle \dots \langle n-1 n \rangle \langle n 1 \rangle} \delta^4 \left(\sum_{i=1}^n k_i \right)$$



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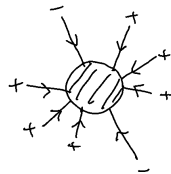
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$$k_i^{\alpha\dot{\alpha}} := \frac{1}{\sqrt{2}} \sigma_{\mu}^{\alpha\dot{\alpha}} k_i^{\mu}, \quad \sigma_{\mu} = (\mathbb{1}_2, \vec{\sigma}), \quad \det(k_i^{\alpha\dot{\alpha}}) = k_i^2 = 0,$$

$$k_i^{\alpha\dot{\alpha}} = \kappa_i^{\alpha} \bar{\kappa}_i^{\dot{\alpha}}, \quad \langle i j \rangle := \kappa_i^{\alpha} \kappa_{j\alpha}, \quad [i j] = \bar{\kappa}_i^{\dot{\alpha}} \bar{\kappa}_{j\dot{\alpha}}$$

$$\alpha, \beta = 0, 1, \quad \dot{\alpha}, \dot{\beta} = \dot{0}, \dot{1}, \quad \epsilon_{\alpha\beta} = \epsilon_{\dot{\alpha}\dot{\beta}} = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}.$$



Cartan-valued SD radiative backgrounds

We study gauge field backgrounds $A \in \Omega^1(\mathbb{R}^4, \mathfrak{g}_{\mathbb{C}})$.

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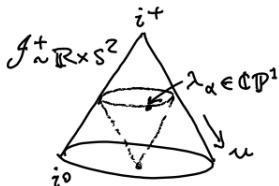
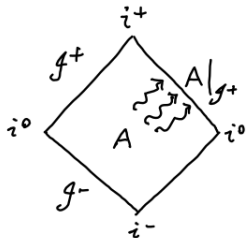
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► **Radiative** A fixed by radiation data $A|_{\mathcal{I}^+}$.

$$A|_{\mathcal{I}^+} = \mathcal{A}(u, \lambda, \bar{\lambda}) D\lambda + \tilde{\mathcal{A}}(u, \lambda, \bar{\lambda}) D\bar{\lambda}.$$

$(u, \lambda_{\dot{\alpha}}, \bar{\lambda}_{\dot{\alpha}})$ homog. coordinates on $\mathcal{I}^+ \simeq \mathbb{R} \times \mathbb{CP}^1$.

$D\lambda = \langle \lambda d\lambda \rangle$, $D\bar{\lambda} = [\bar{\lambda} d\bar{\lambda}]$ basis of $T^*\mathbb{CP}^1$.



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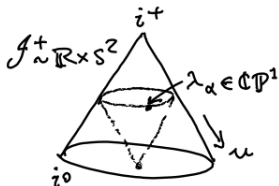
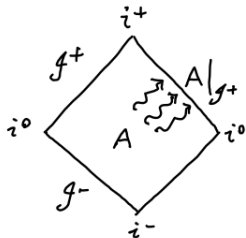
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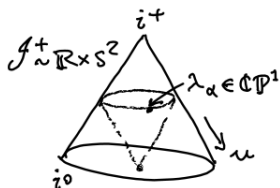
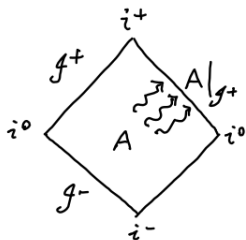
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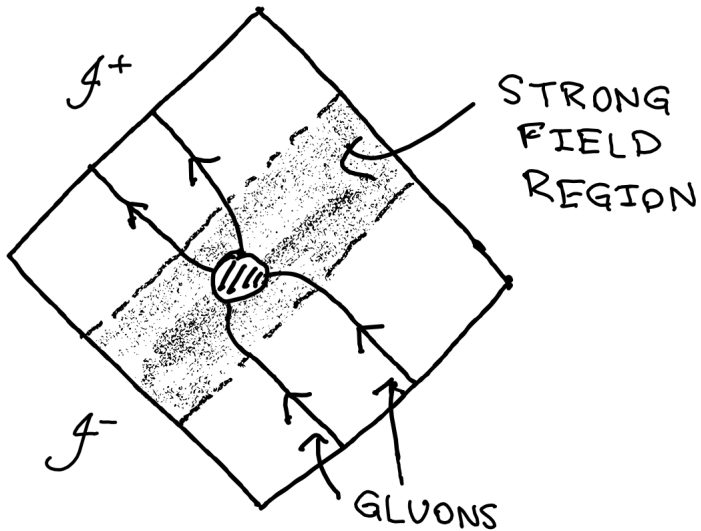
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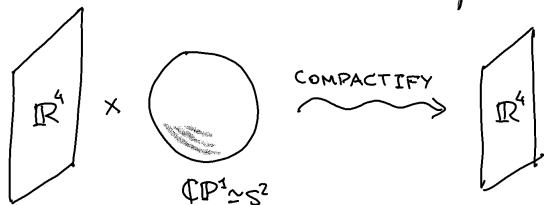




A weak-weak duality [Mason '05] [Boels, Mason, Skinner '07]

$$\mathbb{P}T \simeq \mathbb{R}^4 \times \mathbb{P}^1 \simeq \mathcal{O}_{\mathbb{P}^1}(1)^{\oplus 2}$$

Spacetime



Twistor action

$$\int_{\mathbb{P}T} b \bar{\partial} a + \sum_{n=2}^{\infty} \int_{\mathbb{R}^4} \int_{(\mathbb{P}^1)^n} b^2 a^{n-2}$$



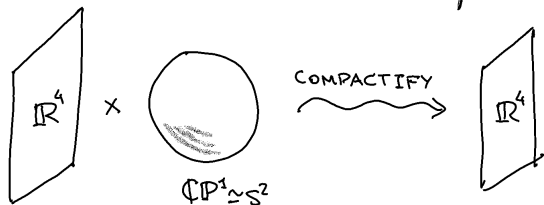
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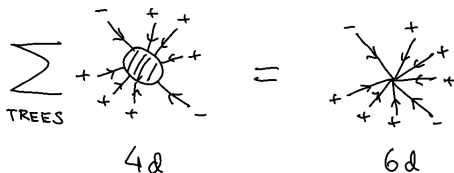
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A MAGIC TRICK :



MHV gluon amplitudes on SD radiative $A = A_\mu dx^\mu \in \text{Cartan}$
 with gluons r, s negative helicity : [\[Adamo, Mason, AS '19\]](#)

$$\frac{\langle r s \rangle^4}{\langle 1 2 \rangle \langle 2 3 \rangle \cdots \langle n-1 n \rangle \langle n 1 \rangle} \int d^4x \exp \sum_{i=1}^n (i k_i \cdot x + e_i g(x, \kappa_i)) .$$

- e_i gluon charge w.r.t. Cartan, $\sum_i e_i = 0$; $k_i^{\alpha\dot{\alpha}} = \kappa_i^\alpha \bar{\kappa}_i^{\dot{\alpha}}$.

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- e_i gluon charge w.r.t. Cartan, $\sum_i e_i = 0$; $k_i^{\alpha\dot{\alpha}} = \kappa_i^\alpha \bar{\kappa}_i^{\dot{\alpha}}$.
- $g(x, \kappa)$ gives holographic imprint of A : $A|_{\mathcal{I}^+} = \tilde{\mathcal{A}}(u, \lambda, \bar{\lambda}) D\bar{\lambda}$.

$$g(x, \kappa) = \frac{1}{2\pi i} \int_{\mathbb{CP}^1} \frac{\langle \kappa \xi \rangle}{\langle \kappa \lambda \rangle \langle \lambda \xi \rangle} D\lambda \wedge D\bar{\lambda} \tilde{\mathcal{A}}|_{u=x^{\alpha\dot{\alpha}}\lambda_\alpha \bar{\lambda}_{\dot{\alpha}}} .$$

$\lambda_\alpha \in \mathbb{CP}^1$ (celestial sphere), ξ_α a reference spinor, $D\lambda = \langle \lambda d\lambda \rangle$.

Features of our formula

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- ▶ Simplicity without a shred of momentum conservation!
- ▶ Lot less integrals in comparison to Feynman rules.
For n -gluon scattering, $(n-2)$ cubic vertices should give $(n-2)$ $d^4 x$ integrals, but we found only one.
Behind the scenes: remarkable IBP relations collapse most integrals.

Conclusions

- ▶ A similar story holds for graviton scattering in SD radiative spacetimes (to appear “soon”).

See also [\[Adamo, Mason, AS '19\]](#).

- ▶ Many important future directions:
 - SD backgrounds with colour/gravitational sources and black holes.
 - Double copy of amplitudes on curved backgrounds.
 - Non-self-dual backgrounds using ambitwistor strings.
 - Generalisation of other modern methods like unitarity, on-shell recursion, polytopes...

