

# Moduli space of five-dimensional black holes

**Fred Tomlinson**  
f.tomlinson@ed.ac.uk

University of Edinburgh

YTF 2020

# Contents

## 1 Higher-Dimensional Gravity

- Motivation
- 5d Vacuum GR

## 2 Uniqueness Theorems

- 4d Vacuum Uniqueness
- Rod Structure
- 5d Vacuum Uniqueness

## 3 Integrability

- Linear Pair
- Inverse Scattering and New Results

# Motivation

- AdS/CFT correspondence
- Black hole microstate counting and string theory
- Better understanding of gravity by introducing a tuneable parameter (dimension)
- Richer structure

# Basic Features of 5d Vacuum GR

- 4d Kerr black hole ( $S^2$  horizon)  $\rightarrow$  5d (doubly rotating) Myers-Perry black hole ( $S^3$  horizon)
- In 5d there is the black ring solution with horizon topology  $S^2 \times S^1$  i.e. a non-spherical horizon
- There are also multi-black hole solutions: the black Saturn and di-ring solutions.

## 4d Vacuum Uniqueness

### 4d Vacuum Uniqueness [Carter '71, Robinson '75]

All stationary, axisymmetric, asymptotically flat solutions of the vacuum Einstein equations in 4d (subject to technical requirements) are uniquely determined by their mass and angular momentum.

- For a single horizon the unique solution is given by the Kerr solution ("no-hair")
- Can one extend this naively to 5d?

## 4d Vacuum Uniqueness

### 4d Vacuum Uniqueness [Carter '71, Robinson '75]

All stationary, axisymmetric, asymptotically flat solutions of the vacuum Einstein equations in 4d (subject to technical requirements) are uniquely determined by their mass and angular momentum.

- For a single horizon the unique solution is given by the Kerr solution ("no-hair")
- Can one extend this naively to 5d?
- **No!** The black ring is a counterexample (in fact displays 2-fold non-uniqueness)
- What extra information do we need for a uniqueness theorem? Rod structure

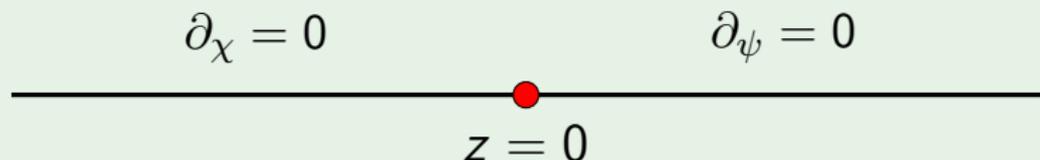
# Rod Structure

- By analogy with Kerr solution which is axisymmetric, we now consider "biaxisymmetric" solutions in 5d
- The orbit space (i.e. spacetime quotiented by the Killing vectors) can be identified with a half plane  $\{(\rho, z) | \rho > 0\}$  - Weyl-Papapetrou coordinates
- The axis  $\rho = 0$  divides into z-intervals (rods) which are either horizons or regions where integer linear combinations of the axial Killing vectors (rod vectors) vanish. This encodes topology
- The lengths of these rods together with these rod vectors gives the rod structure

# Rod Structure Examples 1

## Example (5d flat space)

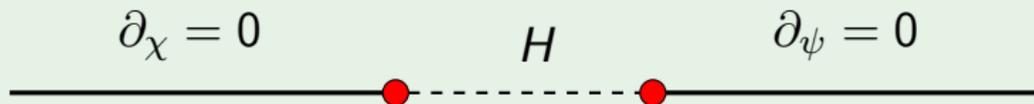
- $ds^2 = -dt^2 + \mu^2(\rho^2 + \mu^2)^{-1}(d\rho^2 + dz^2) + \mu d\psi^2 + \rho^2 \mu^{-1} d\chi^2$
- $\partial_t$  is the generator of time translations,  $\partial_\psi$  and  $\partial_\chi$  are the generators of rotation and  $\mu := \sqrt{\rho^2 + z^2} - z$
- $\rho = 0$  splits up into two (semi-infinite) rods:  $z \leq 0$  along which  $\partial_\chi$  vanishes and  $z \geq 0$  along which  $\partial_\psi$  vanishes
- Can be represented diagrammatically as:



## Rod Structure Examples 2

### Example (Myers-Perry)

- $\partial_t$ ,  $\partial_\psi$  and  $\partial_\chi$  generators of symmetry as before but now the metric is much more complicated
- Now there is a black hole horizon (the orbit space of which is written as  $H$ ):

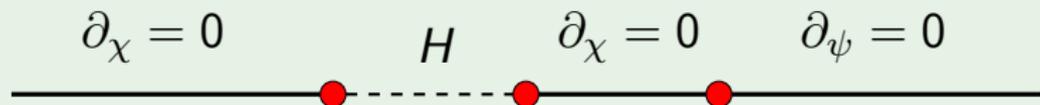


- The horizon rod length here is given by  $\frac{4}{3\pi} M$

## Rod Structure Examples 3

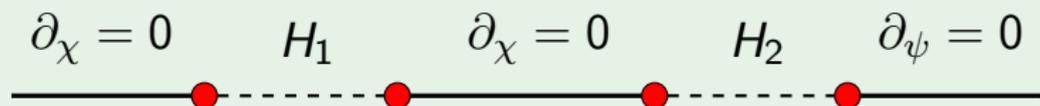
### Example (black ring)

- This adds in an extra finite rod:



### Example (black Saturn)

- In this case there are 2 horizon components,  $H_1$  and  $H_2$ :



# Uniqueness Theorem

## 5d Vacuum Uniqueness [Hollands & Yazadjiev '11]

All stationary, biaxially symmetric, asymptotically flat solutions of the vacuum Einstein equations in 5d are uniquely determined by their rod structure and angular momentum (subject to technical requirements).

# Uniqueness Theorem

## 5d Vacuum Uniqueness [Hollands & Yazadjiev '11]

All stationary, biaxisymmetric, asymptotically flat solutions of the vacuum Einstein equations in 5d are uniquely determined by their rod structure and angular momentum (subject to technical requirements).

- Note that this only gives **uniqueness** and not **existence**. In 4d we have Kerr (at least for single horizons), in 5d it's an open problem

# Einstein Equations As A Linear Pair

- Einstein equations:

$$\partial_\rho U + \partial_z V = 0, \quad (1)$$

$$U = \rho \partial_\rho g g^{-1}, \quad V = \rho \partial_z g g^{-1} \quad (2)$$

where  $g$  is the matrix of inner product of Killing vectors

- The equations are integrable, which means that they can be written as the integrability condition of a pair of linear equations:

$$\partial_\rho \Psi = A\Psi, \quad \partial_z \Psi = B\Psi \quad (3)$$

i.e.  $[\partial_\rho, \partial_z]\Psi = 0$  implies the Einstein equations

- $A, B, \Psi$  are matrix functions of  $\rho, z$  and  $\lambda$ , a new complex parameter.  $A, B$  are built out of the matrix  $g$

# How Does This Help?

- Inverse scattering method:
  - ▶ Seed solution + transformation parameters  $\rightarrow$  more complicated solution
  - ▶ Powerful - 5d solutions given above all generated using this method
- Integrating the LP along  $\rho = 0$  and at infinity:
  - ▶ In principle, gives solution for metric on  $\rho = 0$  for any rod structure subject to constraints on various moduli
  - ▶ Gives results for Kerr, Myers-Perry and black ring spacetimes
  - ▶ Can be used to rule out double Kerr in 4d and certain black lens solutions in 5d

# The End

Thanks for listening!

## Questions?

References:

- Review for higher-dimensional gravity: Emparan, Reall (2008), arXiv:0801.3471
- Reviews for uniqueness theorems: Hollands, Ishibashi (2012), arXiv:1206.1164; Chruściel, Costa, Heusler (2012) arxiv:1205.6112
- New 5d classification results: Lucietti, FT (2020), arXiv:2008.12761 and arXiv:2012.00381